Machine Detector Interface Summary

Donatella Lucchesi, Sergo Jindariani, Nikolai Mokhov Christian Carli, **Nadia Pastrone** *Yunhai Cai, Francesco Collamati, Matthias Mentink, Yuri Nosochkov, Physics&Detector Group*

1st Community Meeting - May 21, 2021

MDI Session

Is the investment into a full CDR scientifically justified?

Is the 10+ TeV energy region feasible demonstrating the physics potential with sufficient instantaneous **luminosity?**

2

The physics goals of a Muon Collider can only be achieved with self-consistent designs of the ring, IR, high-field SC magnets, MDI and detector

suppress by several orders of magnitude the beam **induced background (BIB) rates in the detector volume**

MDI studies at present

Studies 1994-2014 (MAP 2011-2018) +++

- coherent dedicated efforts on the lattice design (led by Yuri Alexahin)
- high-field large-aperture SC magnet design (led by Sasha Zlobin)
- mitigating BIB on IR and detector optimized in MARS (led by Nikolai Mokhov)
- First tracker detector studies in MARS15/ILCRoot/Geant4 (led by Vito Di Benedetto & C. Gatto)

Physics&Detector Group studies 2018-on going

- MAP lattice and BIB in FLUKA \rightarrow new tool for physics&detector full simulation effort ready to implement new optics design (led by Donatella Lucchesi)
- New full detector with improved capabilities in FLUKA/ILCsoft/Geant4 to be redesigned for higher energies (led by Donatella Lucchesi)

MDI studies @1.5 TeV

with $2\times10^{12}\mu$ /bunch $\rightarrow 2\times10^5$ decays per meter of lattice single pass

Three sources of beam-induced backgrounds (BIB) and radiation loads in MC: **incoherent e+e- pair production at the IP beam halo loss on limiting apertures muon beam decays Nikolai** Mokhov

BIB Characteristics comparison 125 GeV/1.5 TeV

Comparison between $\sqrt{s} = 1.5$ TeV and $\sqrt{s} = 125$ GeV

- BIB absolute fluxes very similar
- Momentum distribution quite different
- Time distribution as expected and Z distribution very similar $\mathbb{E} \left[\mathbb{E} \left[\mathbb{$ Would be possible to do it also at high energy? The IR has been designed to obtain that.

6

interaction point for the cases of a 1.5 TeV (left) and a 125 GeV (right) collider.

 $F = \frac{P}{\sqrt{2}}$

Simulation tool structure

7

 \overline{a}

MDI Layout Description @1.5 TeV

Francesco Collamati

The 1.5TeV case benchmark **MARS-FLUKA Results Comparison**

9

Francesco Collamati

Residual discrepancies in particles time and energy distribution:

- Minor layout differences (passive elements, absorbers) \bullet
- Intrinsic differences between codes \bullet

The 1.5TeV case benchmark **MARS-FLUKA Results Comparison**

The role of the Nozzle:

10

Francesco Collamati

New detector configuration @ 1.5 TeV

Tracker detector considerations perar fah the beam-induced background backgro ● Double-sensor layers ● Cluster shape analysis using

- Timing window applied to reduce hits from out-of-time BIB
- Granularity optimized to ensure $\leq 1\%$ occupancy
- Realistic digitization in progress \rightarrow BIB suppression based on cluster shape pe reconstruction time re
- **If primary vertex could be known before** \rightarrow **effective angular matching of hit doublets**

12

 $\overline{12}$ $\overline{20}$

To be tuned in presence of secondary vertices or long-lived particles

Other detector considerations <u>Contra di Geronio</u>

BIB deposits large amount of energy in both ECAL and HCA

International
UON Collider ollaboration

Calorimeters Timing and shower profile should be used in clusters reconstructions

First look at a Muon Collider detector magnet design

CLIC-like Superconducting Solenoid, with 3.6 T at the interaction point

- 3.6 T at IP and with return yoke \rightarrow 1.8 GJ stored magnetic energy
- For reference, Compact Muon Solenoid has stored energy of 2.6 GJ

M. Mentink, A. Dudarev, and B. Cure

Similar to Compact Muon Solenoid First-order cost estimate: ~80-100 MCHF (2008)

Organizational challenges requires a long-term (15-20 years) schedule and support from multiple institutes

Conductor challenge

Consideration of HTS-based detector magnets for potential future cost savings

IR Lattice & Magnet Design @ 1.5 TeV

International
UON Collider llaboration

Important role is played by the absorber materials

Beam sizes and aperture of the FF magnets $a \geq 5\sigma_{\text{max}} + 1$ cm, B=8T in dipole. L_B = 1.5 to 1.7 m in quads and 6 m in B1

Cross-sections and good-field regions $(|\delta B/B|<10^{-4})$ of Q1, Q2 and Q3-Q5 quads with $G = 250$, 187 and -130 T/m, respectively

15

Deposited power density in $Q1$ (mW/g)

Nikolai Mokhov, Fermilab, May 20, 2021

IR Linear Optics @ 6 TeV Design

N COLLIDER. 6 TEV CMS $With 32 version 8.51/15$ 21/11/14 17.02.48 400. 4.0 $\beta^{^{n2}}$ (m^{n3}) $\widehat{\epsilon}$ 3.5 350. 3.0 2.5 300. 2.0 250. 200. 1.0 0.5 150. $0.\theta$ 100. -0.5 -1.0 50. -1.5 0.0 -2.0 tóo. 200. 400. 0.0 300. 500. $s(m)$

Yunhai Cai and Yuri Nosochkov

Field@poletip: <15 T in quads <20 T in bends

 β^{\star} _{x,y} = 1 cm $L^* = 6$ m

Y. I. Alexahin, E. Gianfelice-Wendt, V.V. Kashikhin, N.V. Mokhov, A.V. Zlobin, V.Y. Alexakhin, PRSTAB 14, 061001 (2011)

16 M.-H. Wang, Y. Nosochkov, Y. Cai, M. Palmer, 2016 JINST 11 P09003

6 TeV IR and MDI studies

17

Several LOIs were submitted to Snowmass, a kind of a pre-feasibility study to address design of a lattice for the average luminosity of 10^{35} cm⁻²s⁻¹ with β^* =3 mm, high-field / high-gradient superconducting magnets with the coil apertures as large as 25 cm in IR and 15 cm in the arcs, $abs(G) = 78$ to 200 T/m, 8-T dipoles with hybrid multilayer coils made of HTS and LTS with the stress management (AZ), optimized MDI with tungsten liners, masks, nozzles etc, detector BIB mitigation techniques as well as appropriate measures to mitigate neutrino hazard.

IR Linear Optics @ 10 TeV scaled

Yunhai Cai and Yuri Nosochkov

18

- Scaled from 6 TeV design with factor: $F = 10/6$
	- All: Length->Length*F
	- Bending angle kept same
	- $K_1 > K_1/F^2$
	- $K_2 > K_2/F^3$
- Keep phase advances same
- No change of beam size

 $\beta^*_{x,y}$ = 1.7 cm, L* = 10 m, Circumference= 10.5 km

Yunhai Cai and Yuri Nosochkov

Next steps: simulation studies
Results are in publication *larXiv:2105.09116)* Donatella Lucchesi

Results are in publication (arXiv:2105.09116)

- Study Beam-Induced Background at $\sqrt{s} = 3$ TeV, use MAP IR and the nozzle of $\sqrt{s} = 1.5$ TeV, then
	- Optimize nozzle
	- Optimize IR
- Detector studies are just at the first step, a lot of room for improvements!
- Physics objects performance are very good even if not optimize, room for improvements in particular with ML techniques
- Dedicated studies and optimization is needed for the forward region, covered by the nozzle

A. Mereghetti

Strong collaboration between accelerator and detector physicists is mandatory for the proper MDI design

> **Thanks to the speakers and** *everybody* who contributed to *the work and discussion*