GHENT 24-26 NOV. Muographers 2021

Muography developments within the MuAR project: advances in simulations and new detectors designs

H. Asorey[1,2] and A. Almela[1] for the MuAr group

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Muography developments within the MuAR project: advances in simulations and new detectors designs

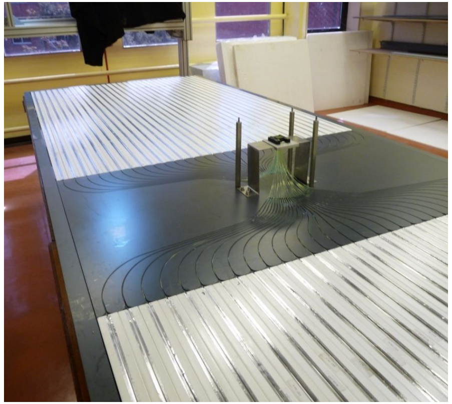
Background

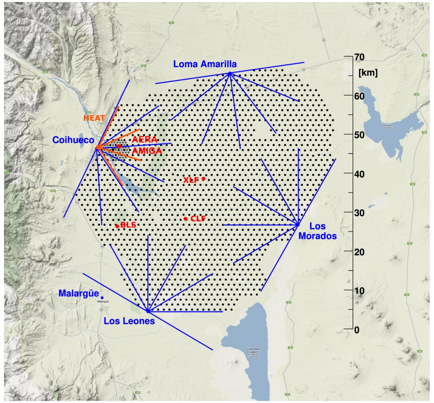
Simulation program

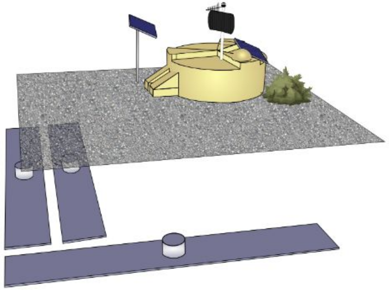
Detectors designs

Background: LAGO, AMIGA and MuTe

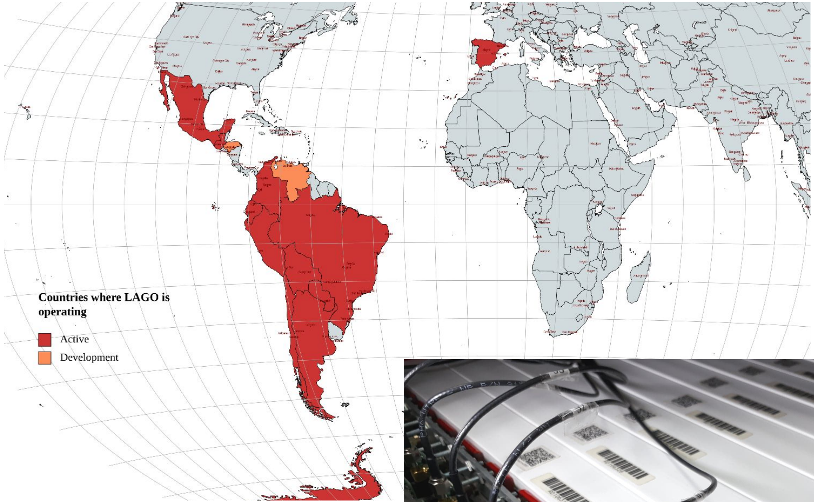
AMIGA is a buried muon counter designed to study the UHECR composition at the Pierre Auger Observatory by measuring the EAS muon distribution density at ground





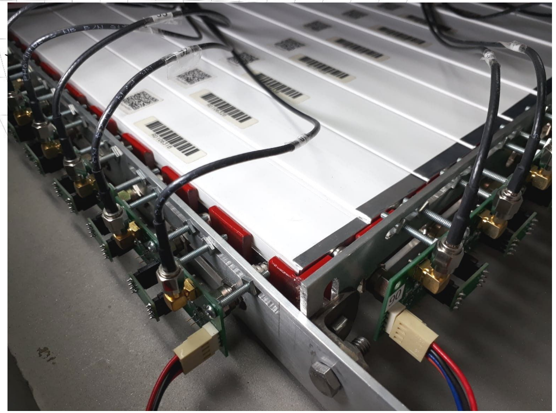


LAGO is a giant network of astroparticle WCD detectors, currently operating in 11 countries. The LAGO network measures the time-evolving flux of secondary particles produced by the modulated flux of GCR



MuTe is a two-panel, 900px and hybrid muographer, scintillators & WCD, designed, funded and built in Latin America

(L. Núñez et al, this workshop)



Muography developments within the MuAR project: advances in simulations and new detectors designs

Background

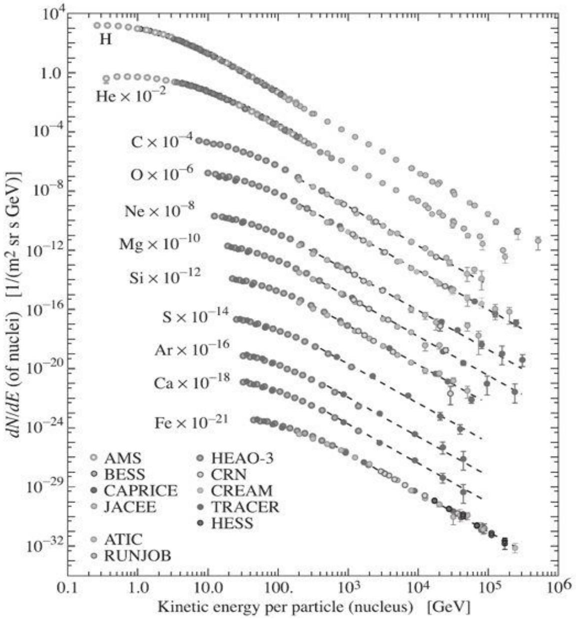
Simulation program

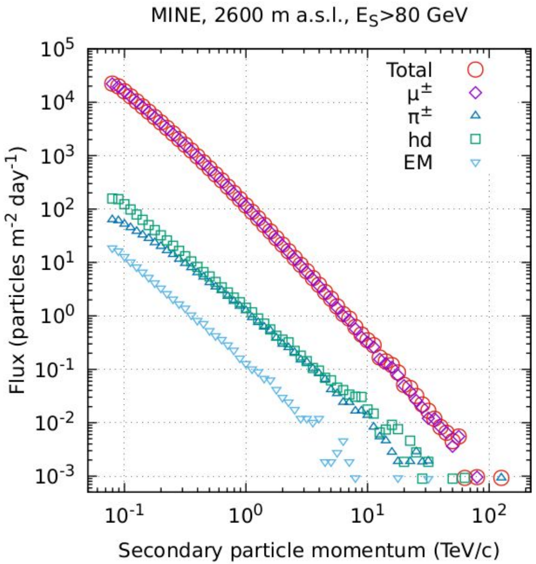
Detectors designs

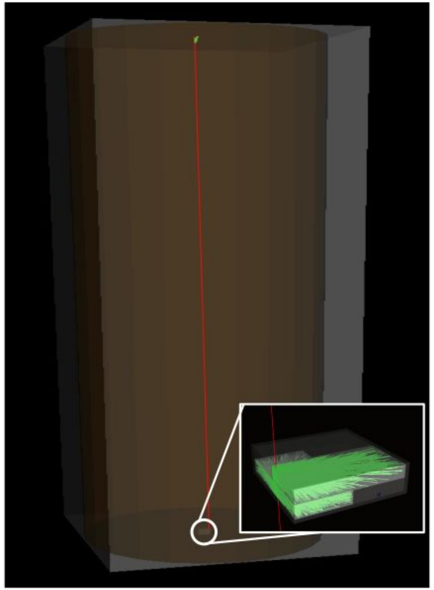
from primary flux to applications

$$\text{GCRFlux} \xrightarrow{\text{Heliosphere}} \text{Modulated Flux} \xrightarrow{\text{Magnetosphere}} \text{Primaries} \rightarrow \cdots$$

$$\cdots \rightarrow \text{Primaries} \xrightarrow{\text{Atmosphere}} \text{Secondaries} \xrightarrow{\text{Detector}} \text{Signals}$$

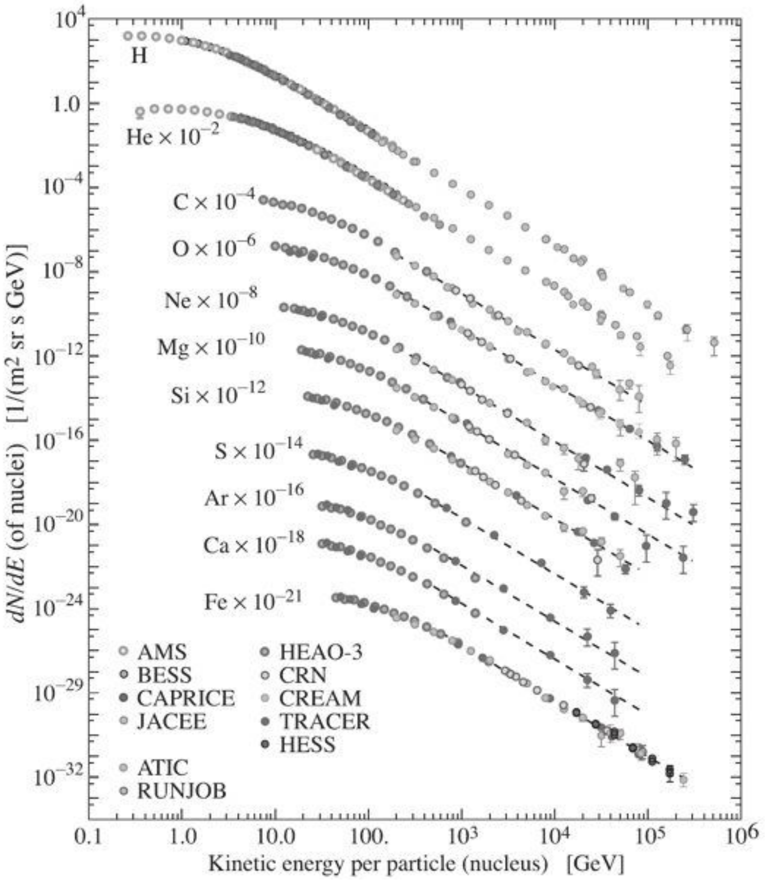






H. Asorey et al., COLAGE2021, Adv. Space Res., in preparation

primary flux integration



For each primary, we need to integrate its spectrum to get the expected (Poissonian) number of primaries at the top of the atmosphere

$$N_{t,S} = \int_t \int_S \int_\Omega \int_{E_p} j_0(E_p, Z_p)^{\alpha(E_p, Z_p)} \, dt \, dS \, d\Omega \, dE$$

We integrate:

full spectra, $1 < Z < 26$

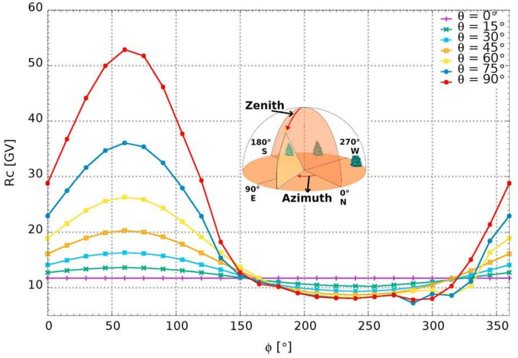
hemisphere, $0 \leq \theta \leq \pi/2, -\pi \leq \phi \leq \pi$

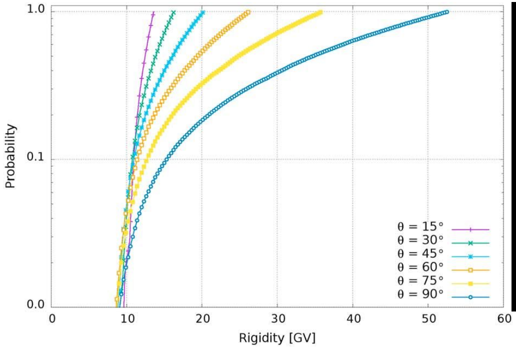
energy range, $(R_C \times Z_p) < E/GeV < E_{\max}$

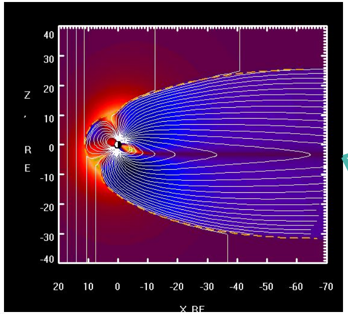
R_C is the local, time-dependent, geomagnetic rigidity cut-off
E_{max} depending on application

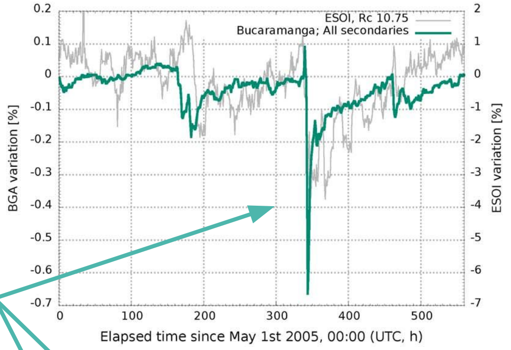
M. Suárez-Durán, H. Asorey and L.A. Núñez, Space Weather, (2018), [DOI:10.1002/2017SW001774]

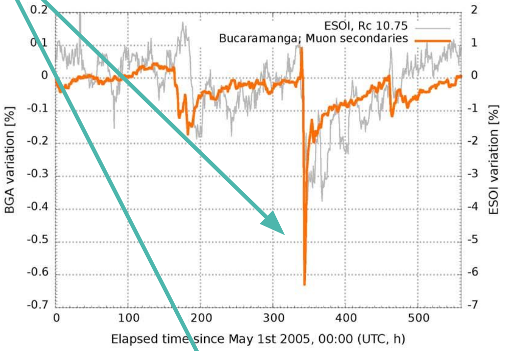
time-dependent local geomagnetic effects

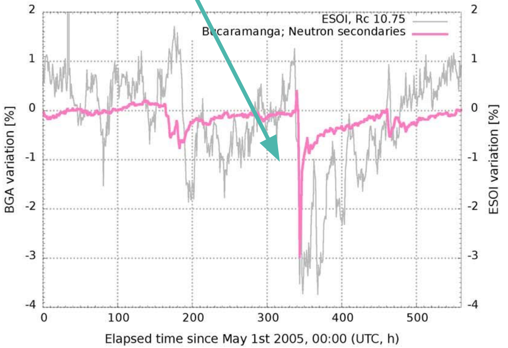


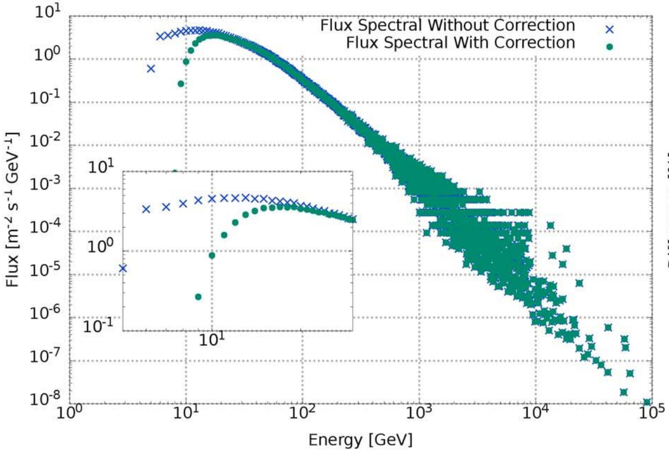


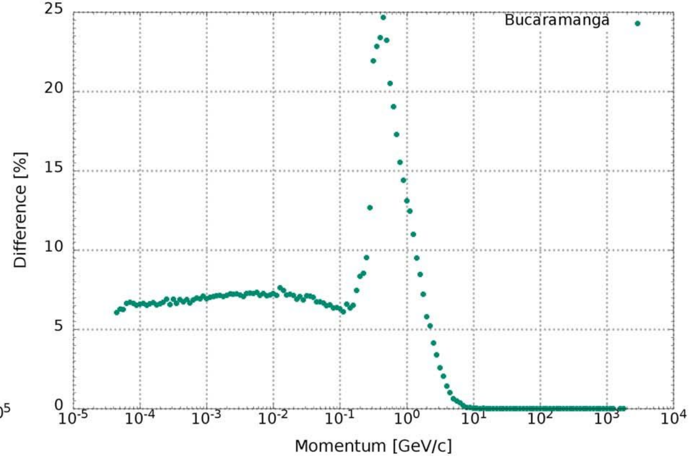








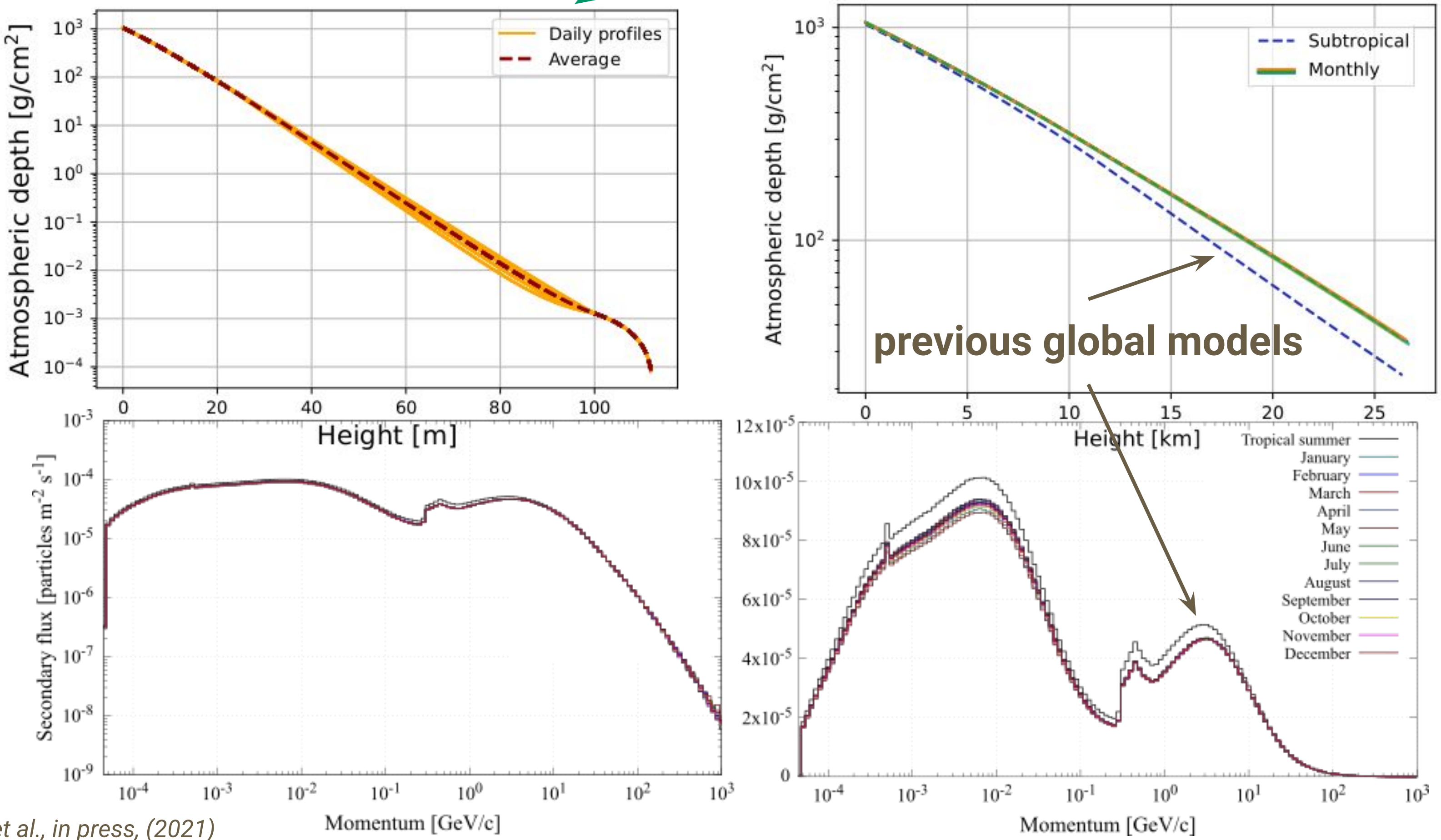




M. Suárez-Durán, H. Asorey and LA Núñez, Space Weather, (2018), [DOI:10.1002/2017SW001774](DOI:10.1002/2017SW001774)

local atmospheric effects

Monthly-averaged or instantaneous local atmospheric profiles from GDAS



J. Grisales et al., in press, (2021)

Our secondary flux simulation framework

Transient astrophysical phenomena

GRBs, Solar Activity, …

Inputs

Flux of CR

4.5×10^{8} day^{-1} m^{-2}

Primaries

Secondaries

4×10^{8} day^{-1}m^{-2} @4600 m asl

EM : MU : NE : HD= 0.875 : 0.065 : 0.050 : 0.010

Secondary propagation and detector response

type, geometry, materials...

GEANT4 detector models

Signals

Site characteristics

location, altitude, geomagnetic field, atmosphere, …

time-evolving conditions:

MAGNETOCOSMICS (IGRF13&TSY), GDAS and CORSIKA

C. Sarmiento-Cano et al. (2019), DOI:10.22323/1.358.0412; Hernán Asorey et al., Astropart. Phys., in preparation

Cloud-based implementation

Any HPC provider assigns resources to the cloud based implementation: cores, memory and local storage

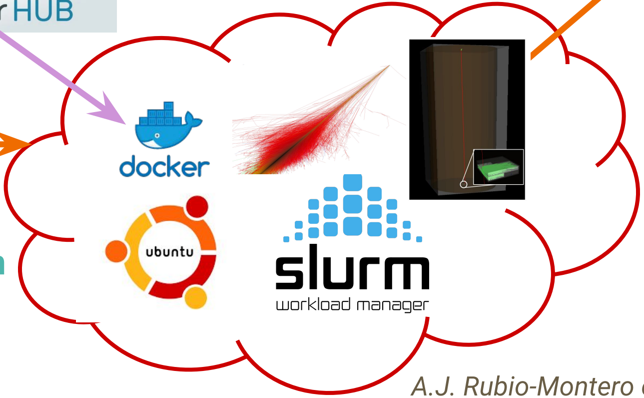
Infrastructure Manager

A web-based service for distributing all the available resources on user-configurable virtual clusters and install the OS (ubuntu 20.04 + slurm manager + docker)

docker HUB

docker containers with our codes are deployed from docker-HUB in the virtual cluster

virtual cluster



Google Cloud

Microsoft Azure

FAIR: produced data catalogs are findable, accessible, interoperable and reusable

B2FIND

ONEDATA

Results are stored at cloud-based storage services. Publicly accessible by requested personal tokens

B2HANDLE

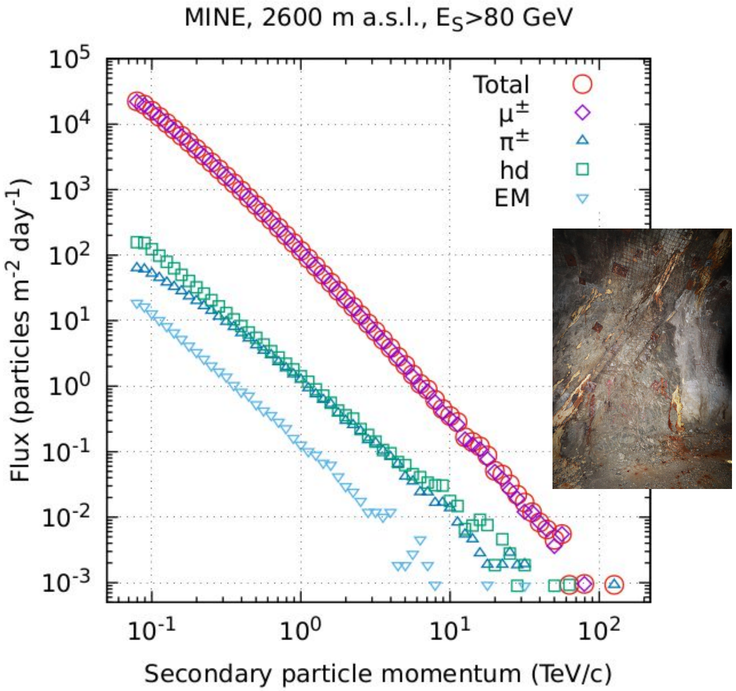
PID (Persistent IDentifiers) are assigned for each data catalog

A.J. Rubio-Montero et al (ICRC 2021) DOI:10.22323/1.395.0261

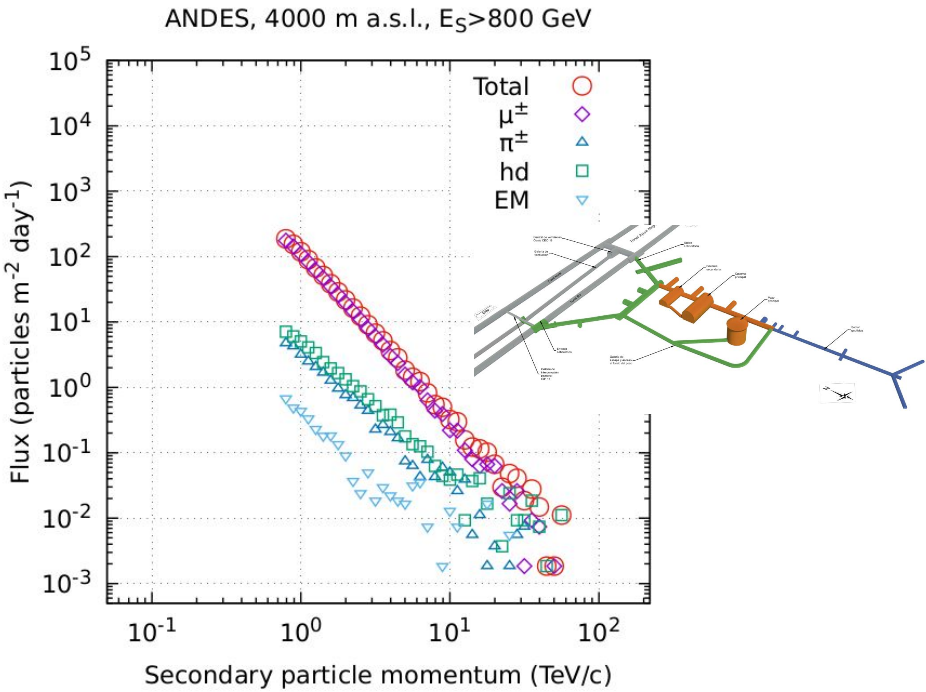
Muography and Underground LABs

One-year simulated flux of secondary particles at ground level (~1.5 kCPU·h/site)

MINE, 900 m.w.e., p_{CUT}=80 GeV/c



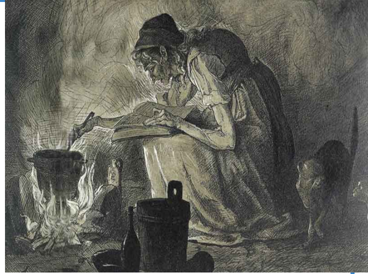
ANDES, 4600 m.w.e., p_{CUT}=800 GeV/c



High-momentum (p_s>p_{CUT}) secondary particle flux at different sites around the World

A.J. Rubio-Montero et al (ICRC 2021) DOI:10.22323/1.395.0261

Meiga, the sorceress (work in progress)



Transient astrophysical phenomena
GRBs, Solar Activity, ...
Inputs

Flux of CR

docker

Secondaries

Primaries

docker

Site characteristics
location, altitude, geomagnetic field, atmosphere, ...
time-evolving conditions:
MAGNETOCOSMICS (IGRF13&TSY), GDAS and CORSIKA

docker

Secondary propagation and detector response
type, geometry, materials...
GEANT4 detector models

docker

Signals

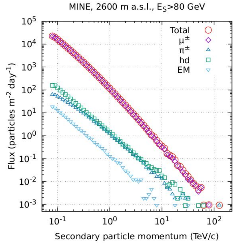
Meiga, a dedicated framework used for muography applications

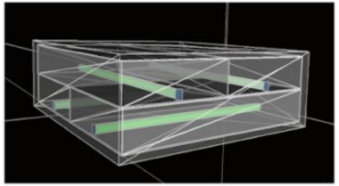
*A. Taboada and C. Sarmiento for the MuAr group - **Poster #26 - Nov 25, 2021, 15:30 CET - https://indi.to/N4Rqh***

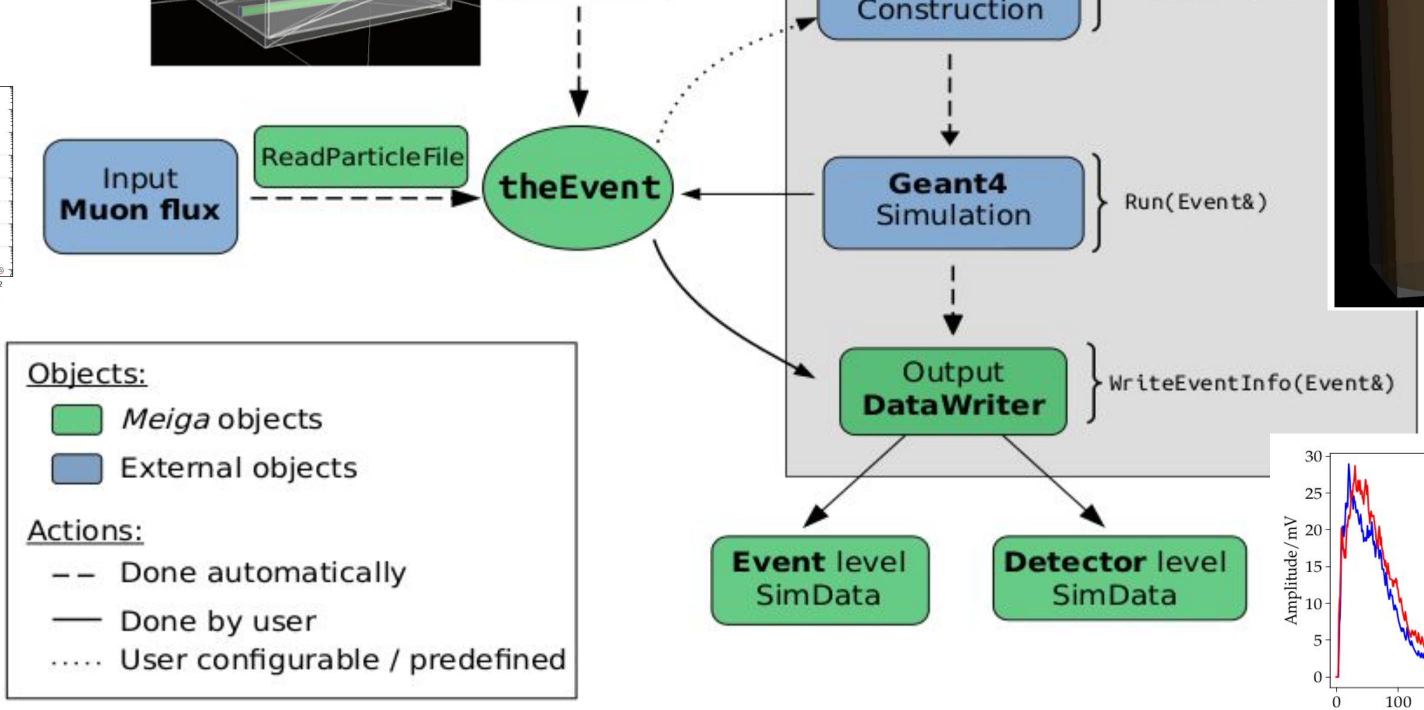
A. Taboada et al, this conference

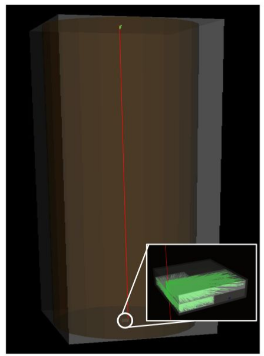
Meiga workflow

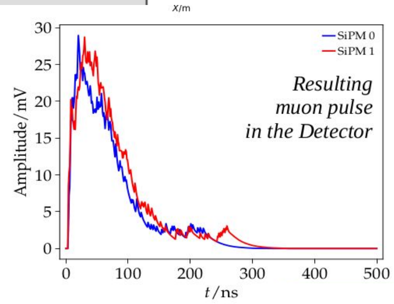
Poster #26 - Nov 25, 2021, 15:30 CET





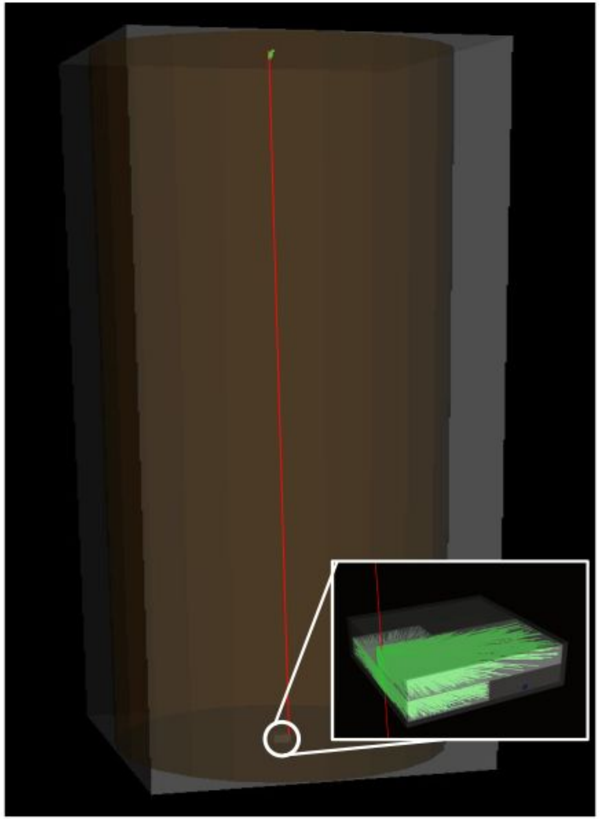




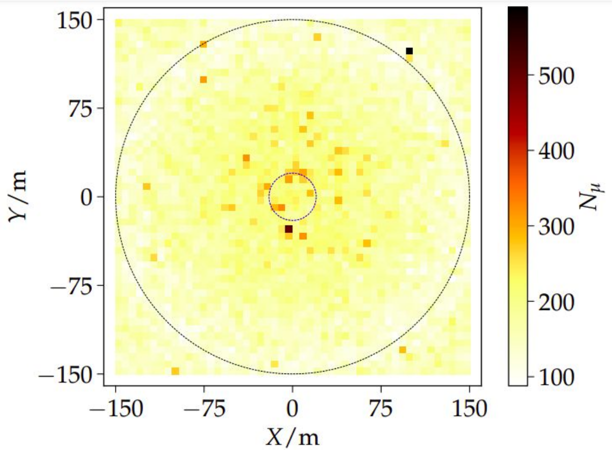


A. Taboada et al, this conference

Meiga example:
muon flux propagation through 500m of rock impinging on a single *Musaic'* tile

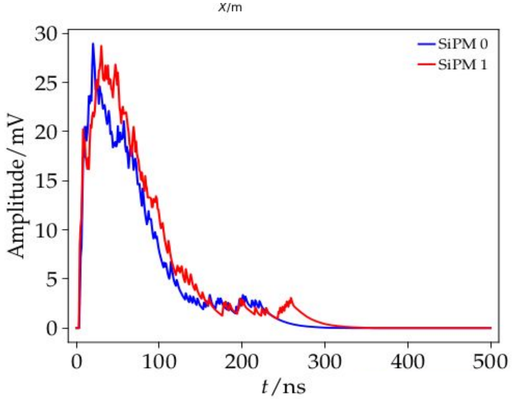


standard rock and
detector models



Muon density footprint
after propagation through
500m of standard rock

Expected pulse at each SiPM



A. Taboada et al, this conference

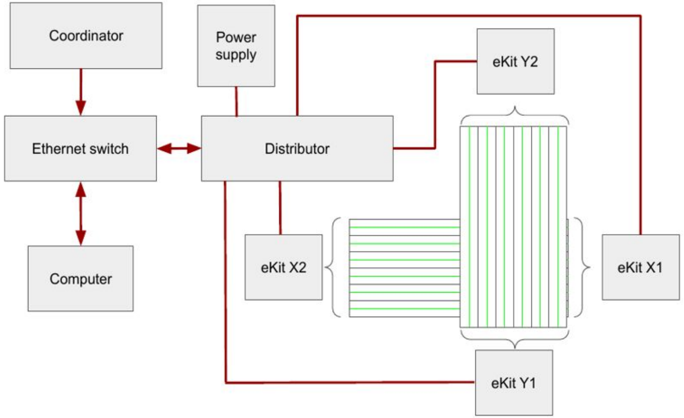
Muography developments within the MuAR project: advances in simulations and new detectors designs

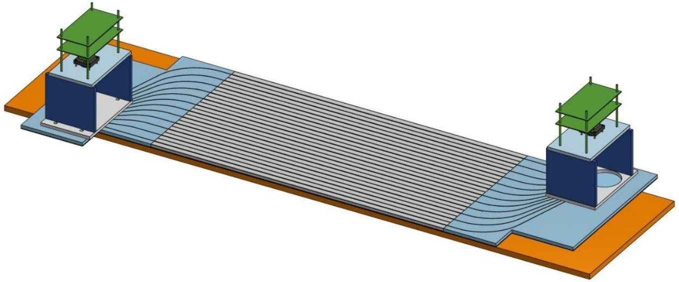
Background

Simulation program

Detectors designs

Modulus, our modular design

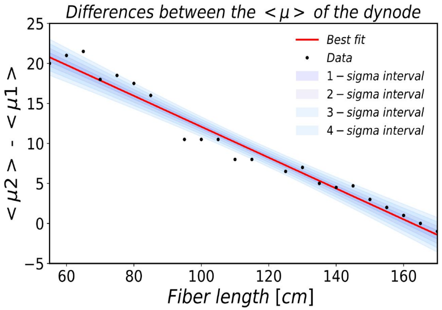


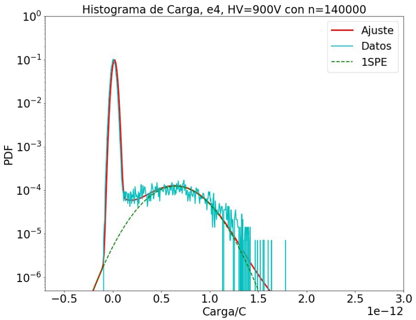


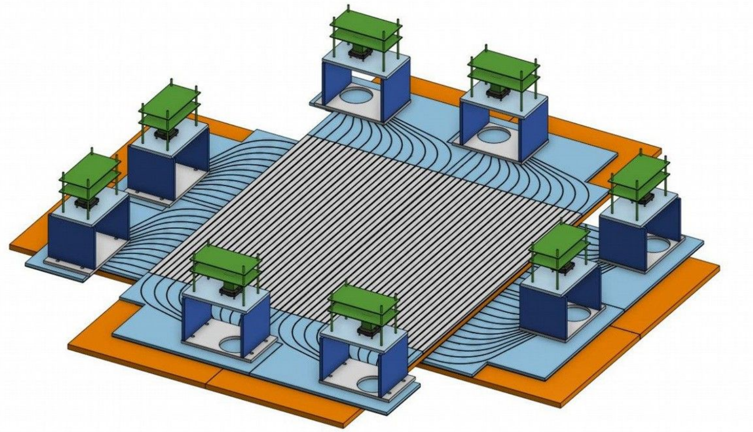
Development of Mudulus, a muography detector based on double-synchronized electronics for geophysical applications

R. Calderón-Ardila et al for the MuAr group

Poster #26 - Nov 25, 2021, 15:30 CET - https://indi.to/HQBZK





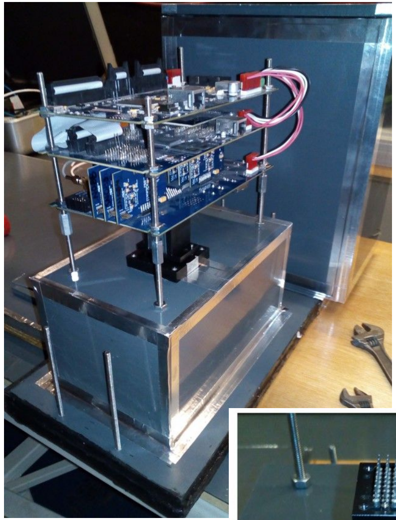


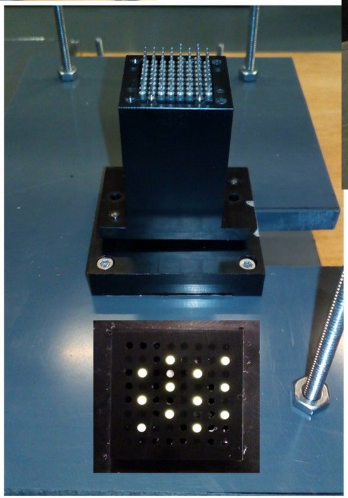
$$S_{real}(x) = S_{ideal}(x) \otimes N(x)$$

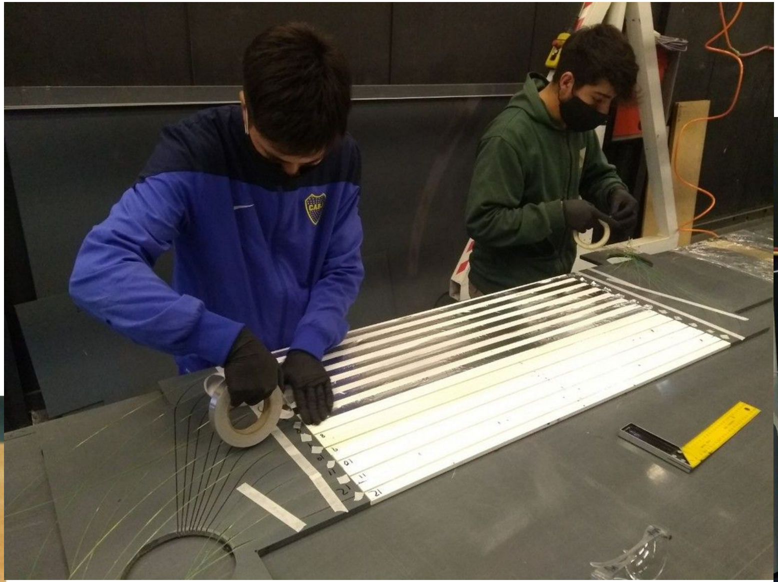
$$S_{ideal}(x) = \sum_{n=0}^{\infty} \frac{u^n \cdot e^{-u}}{n!} \cdot \frac{1}{\sigma\sqrt{2\pi}} \cdot e^{-\frac{(x-n\cdot q)^2}{2n\sigma^2}}$$

R. Calderón-Ardila et al., JINST, submitted, R. Calderón-Ardila et al, this conference

First functional prototype (2021)

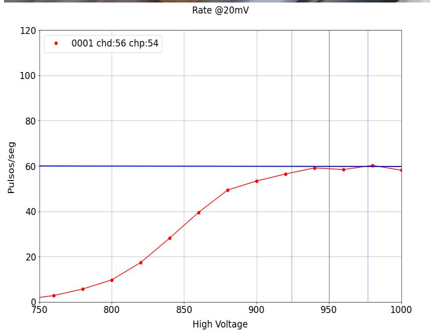


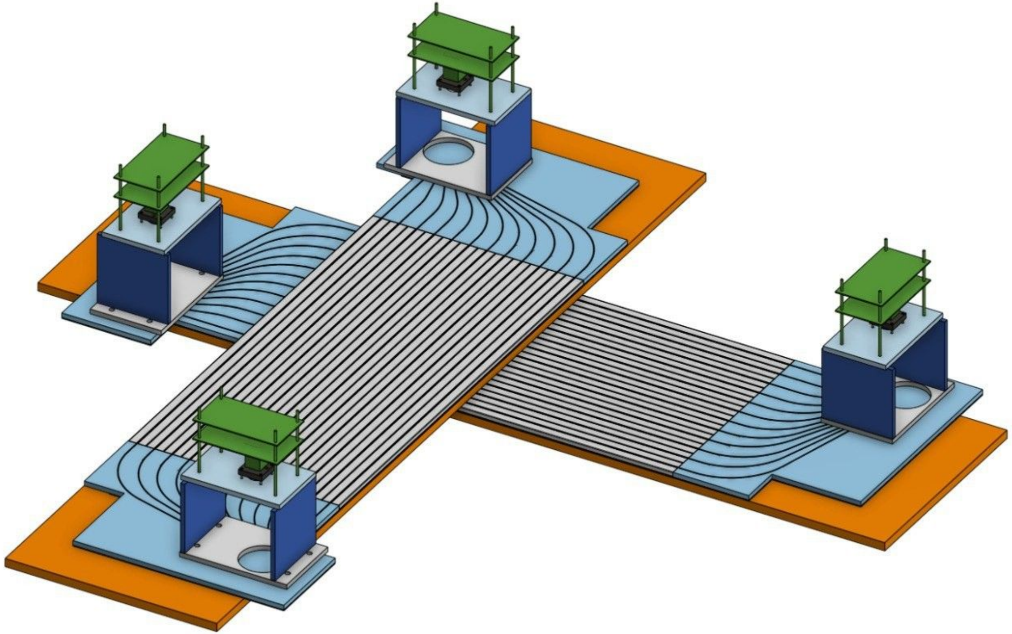




Technology transference from astroparticle detection to muography

Modular assembly, calibration, testing and coincidence detection





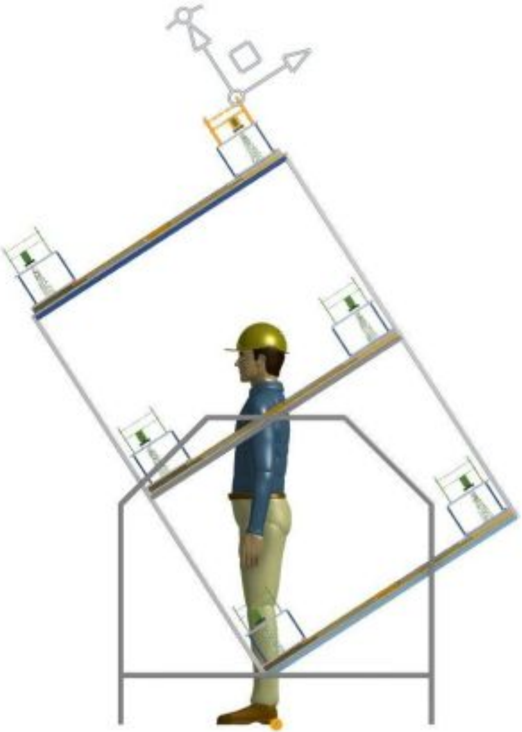
Single or double-head configurable signal acquisition electronics, depending on possible targets characteristics

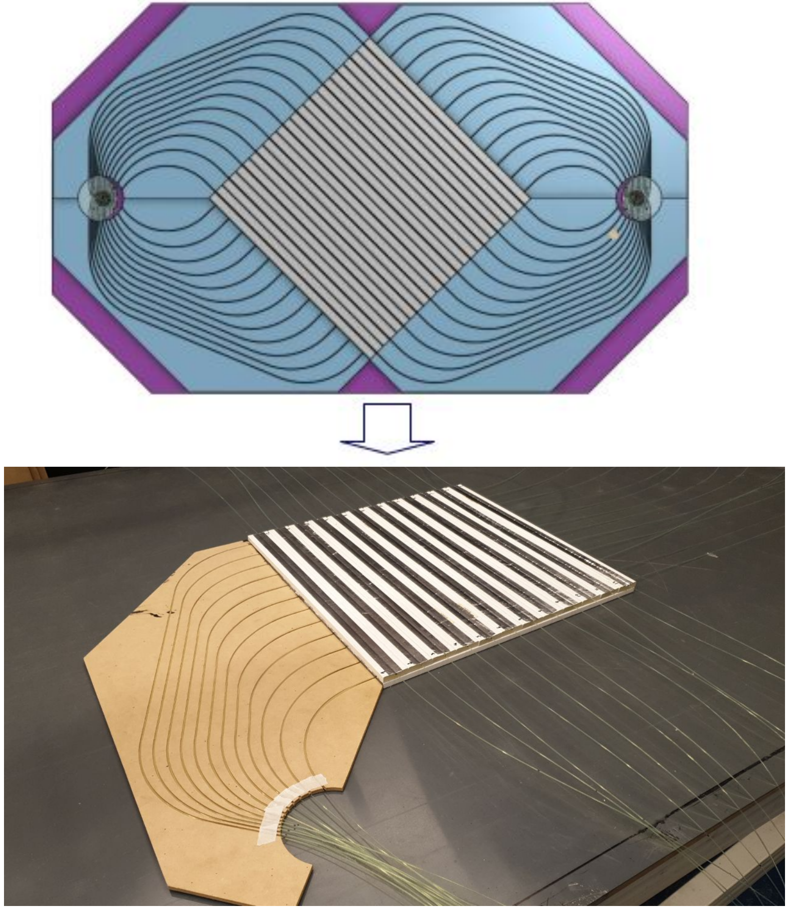
Mining prospecting applications

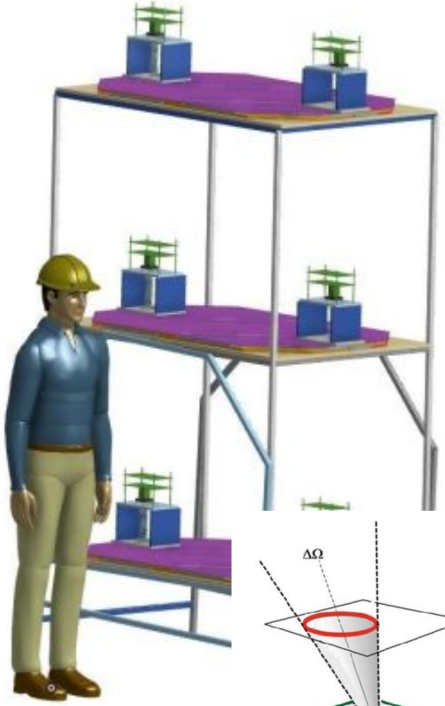
Detector deployment at selected site planned for 2022Q1

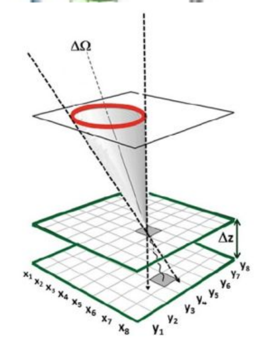
Detector site is selected at 900 mwe deepth

New detector geometry optimized for underground measurement





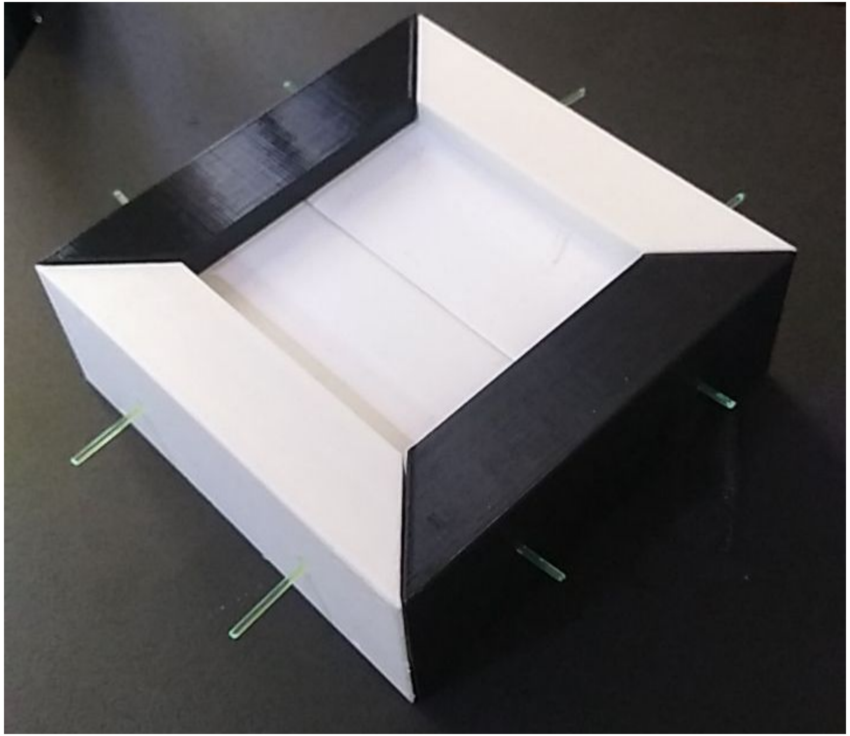




C. Ardila et al, ICES-16 (2021)

Next phase: Musaic

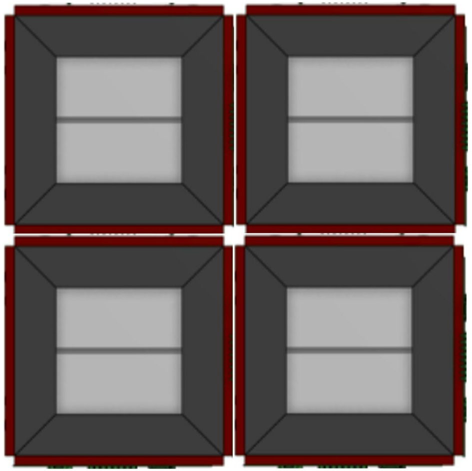
Portable, fully autonomous and interconnectable (2x2) to (6x6) pixel tiles

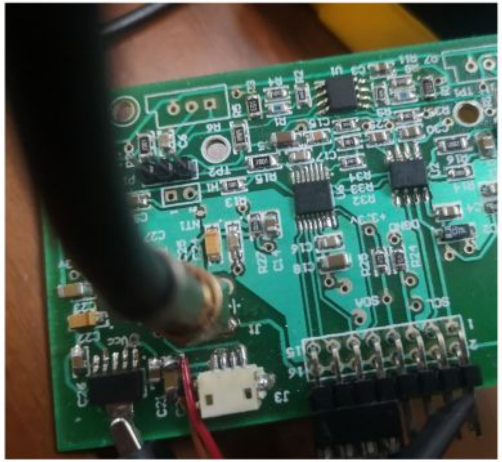


n=2-6, n^2 pixel fully autonomous tiles

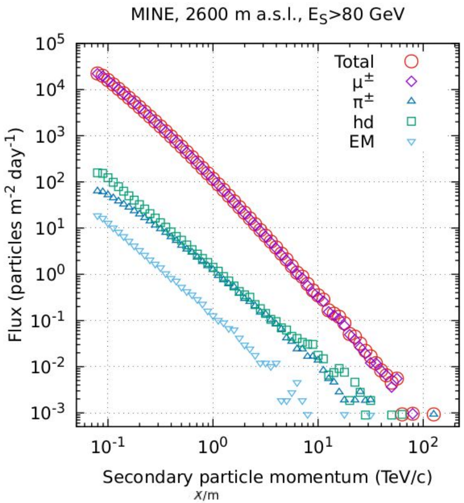
2n SiPM-based tiles

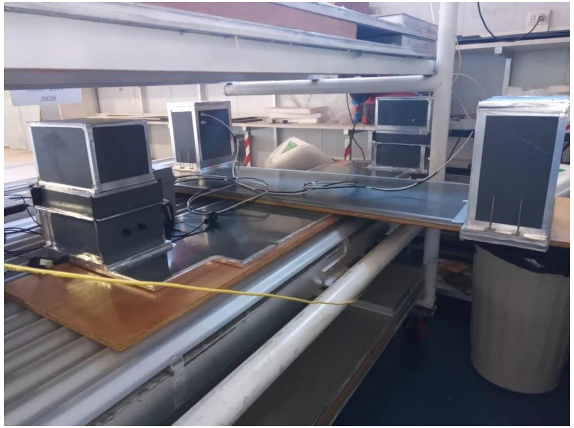
wireless synchronization

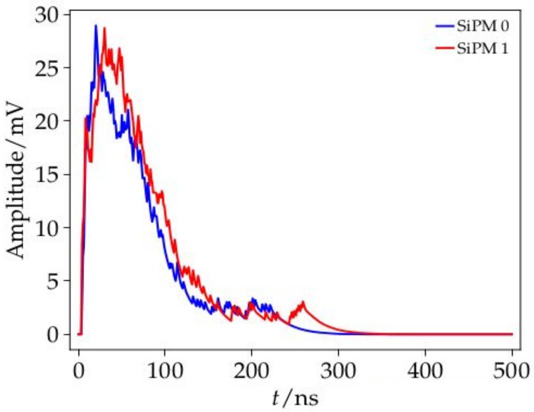




Conclusions







Fully integrated simulation sequence

Configurable and autonomous detector designs: portables, modulars and/or teseleables detectors

Starting field acquisition in 2022Q1

Muography developments within the MuAR project: advances in simulations and new detectors designs

Background

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¡Muchas gracias!