## Global Fits of Dark Matter Effective Field Theories with



http://gambit.hepforge.org

### Jonathan Cornell, on behalf of

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### GAMBIT: The Global And Modular BSM Inference Tool

gambit.hepforge.org

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arXiv: 1705.07908

- Extensive model database not just SUSY
- Extensive observable/data libraries
- Many statistical and scanning options (Bayesian & frequentist)
- Fast LHC likelihood calculator
- Massively parallel
- Fully open-source

#### Members of:

ATLAS, Belle-II, CLiC, CMS, CTA, *Fermi*-LAT, DARWIN, IceCube, LHCb, SHiP, XENON

#### Authors of:

DarkSUSY, DDCalc, Diver, FlexibleSUSY, gamlike, GM2Calc, IsaTools, nulike, PolyChord, Rivet, SoftSUSY, SuperISO, SUSY-AI, WIMPSim



- Fast definition of new datasets and theories
- Plug and play scanning, physics and likelihood packages



### **Recent collaborators:**

P Athron, C Balázs, A Beniwal, S Bloor, T Bringmann, A Buckley, J Eliel Camargo-Molina, C Chang, M Chrzaszcz, J Conrad, J Cornell, M Danninger, J Edsjö, B Farmer, A Fowlie, T Gonzalo, P Grace, W Handley, J Harz, S Hoof, F Kahlhoefer, N Avis Kozar, A Kvellestad, P Jackson, R Jardine, A Ladhu, N Mahmoudi, G Martinez, M Prim, F Rajec, A Raklev, J Renk, C Rogan, R Ruiz, I Sáez Casares, N Serra, A Scaffidi, P Scott, P Stöcker, W Su, J Van den Abeele, A Vincent, C Weniger, M White, Y Zhang

### 40+ participants in 11 experiments and 14 major theory codes

## (Some of the) Recent Developments in GAMBIT

- CosmoBit Constraints from large-scale structure, Type Ia supernovae, Big Bang Nucleosynthesis and the cosmic microwave background (arXiv:2009.03286)
- GAMBIT Universal Model machine (GUM) Tool for easy implementation of new particle physics models into GAMBIT (arXiv:2107.00030)
- New DarkBit Backends
  - DirectDM automatic calculation of running and matching of DM EFTs (Brod, et al., arXiv:1708.02678)
  - Capt'n General accurate calculation of capture rate of DM in the sun (Avis Kozar, et al., arXiv:2105.06810)



## Question

How viable is a generic WIMP that interacts with quarks and/or gluons in light of all available constraints?

# Model: 14 Operator DM EFT

$$\mathcal{L}_{int} = \sum_{a,d} \frac{\mathcal{C}_a^{(d)}}{\Lambda^{d-4}} \mathcal{Q}_a^{(d)}$$
Dimension 7

Dimension 6

 $\mathcal{Q}_{1,q}^{(6)} = (\overline{\chi}\gamma_{\mu}\chi)(\overline{q}\gamma^{\mu}q),$   $\mathcal{Q}_{2,q}^{(6)} = (\overline{\chi}\gamma_{\mu}\gamma_{5}\chi)(\overline{q}\gamma^{\mu}q),$   $\mathcal{Q}_{3,q}^{(6)} = (\overline{\chi}\gamma_{\mu}\chi)(\overline{q}\gamma^{\mu}\gamma_{5}q),$  $\mathcal{Q}_{4,q}^{(6)} = (\overline{\chi}\gamma_{\mu}\gamma_{5}\chi)(\overline{q}\gamma^{\mu}\gamma_{5}q)$ 

$$\begin{aligned} \mathcal{Q}_{1}^{(7)} &= \frac{\alpha_{s}}{12\pi} (\overline{\chi}\chi) G^{a\mu\nu} G^{a}_{\mu\nu} , \qquad \mathcal{Q}_{6,q}^{(7)} = m_{q} (\overline{\chi}i\gamma_{5}\chi) (\overline{q}q) , \\ \mathcal{Q}_{2}^{(7)} &= \frac{\alpha_{s}}{12\pi} (\overline{\chi}i\gamma_{5}\chi) G^{a\mu\nu} G^{a}_{\mu\nu} , \qquad \mathcal{Q}_{7,q}^{(7)} = m_{q} (\overline{\chi}\chi) (\overline{q}i\gamma_{5}q) , \\ \mathcal{Q}_{3}^{(7)} &= \frac{\alpha_{s}}{8\pi} (\overline{\chi}\chi) G^{a\mu\nu} \widetilde{G}^{a}_{\mu\nu} , \qquad \mathcal{Q}_{8,q}^{(7)} = m_{q} (\overline{\chi}i\gamma_{5}\chi) (\overline{q}i\gamma_{5}q) , \\ \mathcal{Q}_{4}^{(7)} &= \frac{\alpha_{s}}{8\pi} (\overline{\chi}i\gamma_{5}\chi) G^{a\mu\nu} \widetilde{G}^{a}_{\mu\nu} , \qquad \mathcal{Q}_{9,q}^{(7)} = m_{q} (\overline{\chi}\sigma^{\mu\nu}\chi) (\overline{q}\sigma_{\mu\nu}q) , \\ \mathcal{Q}_{5,q}^{(7)} &= m_{q} (\overline{\chi}\chi) (\overline{q}q) , \end{aligned}$$

24 total parameters to scan over: 14 Wilson coefficients, DM mass, new physics scale, and 8 nuisance parameters

# Likelihoods

- Direct Detection: Xenon1T, PandaX, DarkSide 50, CRESST-III, etc.
- Indirect Detection:
  - Fermi search for gamma rays from dwarf spheroidal galaxies
  - IceCube solar neutrinos
- Large Hadron Collider: ATLAS and CMS monojet analyses (capped and uncapped)
- Cosmology Planck:
  - DM density (as an overclosure limit)
  - Limits on energy injection from DM annihilation in the early universe
- Nuisance likelihoods: DM halo parameters, top quark mass, DD nuclear parameters

$$\mathcal{L}_{\text{total}} = \mathcal{L}_{\text{DD}} \mathcal{L}_{\text{ID}} \mathcal{L}_{\text{LHC}} \mathcal{L}_{\text{cosmo}} \mathcal{L}_{\text{nuisance}}$$



## **Dimension 6 and 7 Operators**



 Little difference between allowed parameter space for the two sets of operators

arXiv:2106.02056



• With just dimension 6 operators,  $\chi$  can only be a portion of the total DM of the universe when  $m_{\chi}$  < 100 GeV

# Large Hadron Collider

Removing the cap on the LHC likelihood (so that it can exceed the SM value), the best fit region is determined by small excesses in the ATLAS and CMS monojet searches.



The preferred region here is to some degree an artifact of how we implemented the EFT validity requirement. It should be interpreted with care!

# Summary

- GAMBIT is a publicly available tool (download it and try it!) for undertaking global fits of essentially any BSM model. It is rapidly growing in scope, with particularly substantial new additions in cosmology.
- We have used this code to do a global fit of a set of EFT operators that give rise to interactions between a WIMP, quarks, and gluons.
- We find that even just considering the dimension 6 operators, nearly all of the WIMP parameter space remains open, and with the addition of the dimension 7 operators, it is possible for this WIMP to make up 100% of the DM down to low masses.



## **Operator Properties**

	SI scattering	SD scattering	Annihilations
Dimension-6 operators			
$\mathcal{Q}_{1,q}^{(6)} = (\overline{\chi}\gamma_{\mu}\chi)(\overline{q}\gamma^{\mu}q)$	unsuppressed		s-wave
$\mathcal{Q}_{2,q}^{(6)} = (\overline{\chi}\gamma_{\mu}\gamma_{5}\chi)(\overline{q}\gamma^{\mu}q)$	suppressed		<i>p</i> -wave
$\mathcal{Q}_{3,q}^{(6)} = (\overline{\chi}\gamma_{\mu}\chi)(\overline{q}\gamma^{\mu}\gamma_{5}q)$		suppressed	s-wave
$\mathcal{Q}_{4,q}^{(6)} = (\overline{\chi}\gamma_{\mu}\gamma_{5}\chi)(\overline{q}\gamma^{\mu}\gamma_{5}q)$		unsuppressed	s-wave $\propto m_q^2/m_\chi^2$
Dimension-7 operators			
$\mathcal{Q}_1^{(7)} = \frac{\alpha_s}{12\pi} (\overline{\chi}\chi) G^{a\mu\nu} G^a_{\mu\nu}$	unsuppressed		<i>p</i> -wave
$\mathcal{Q}_2^{(7)} = \frac{\alpha_s}{12\pi} (\overline{\chi} i \gamma_5 \chi) G^{a\mu\nu} G^a_{\mu\nu}$	suppressed		s-wave
$\mathcal{Q}_{3}^{(7)} = \frac{\alpha_s}{8\pi} (\overline{\chi}\chi) G^{a\mu\nu} \widetilde{G}^a_{\mu\nu}$		suppressed	<i>p</i> -wave
$\mathcal{Q}_4^{(7)} = \frac{\alpha_s}{8\pi} (\overline{\chi} i \gamma_5 \chi) G^{a\mu\nu} \widetilde{G}^a_{\mu\nu}$		suppressed	s-wave
$\mathcal{Q}_{5,q}^{(7)} = m_q(\overline{\chi}\chi)(\overline{q}q)$	unsuppressed		$p ext{-wave} \propto m_q^2/m_\chi^2$
$\mathcal{Q}_{6,q}^{(7)} = m_q(\overline{\chi}i\gamma_5\chi)(\overline{q}q)$	suppressed		s-wave $\propto m_q^2/m_\chi^2$
$\mathcal{Q}_{7,q}^{(7)} = m_q(\overline{\chi}\chi)(\overline{q}i\gamma_5 q)$		suppressed	$p$ -wave $\propto m_q^2/m_\chi^2$
$\mathcal{Q}_{8,q}^{(7)} = m_q(\overline{\chi}i\gamma_5\chi)(\overline{q}i\gamma_5q)$		suppressed	s-wave $\propto m_q^2/m_\chi^2$
$\mathcal{Q}_{9,q}^{(7)} = m_q (\overline{\chi} \sigma^{\mu\nu} \chi) (\overline{q} \sigma_{\mu\nu} q)$	loop-induced	unsuppressed	s-wave $\propto m_q^2/m_\chi^2$
$\mathcal{Q}_{10,q}^{(7)} = m_q (\overline{\chi} i \sigma^{\mu\nu} \gamma_5 \chi) (\overline{q} \sigma_{\mu\nu} q)$	loop-induced	suppressed	s-wave $\propto m_q^2/m_\chi^2$

# Missing Energy Spectra

