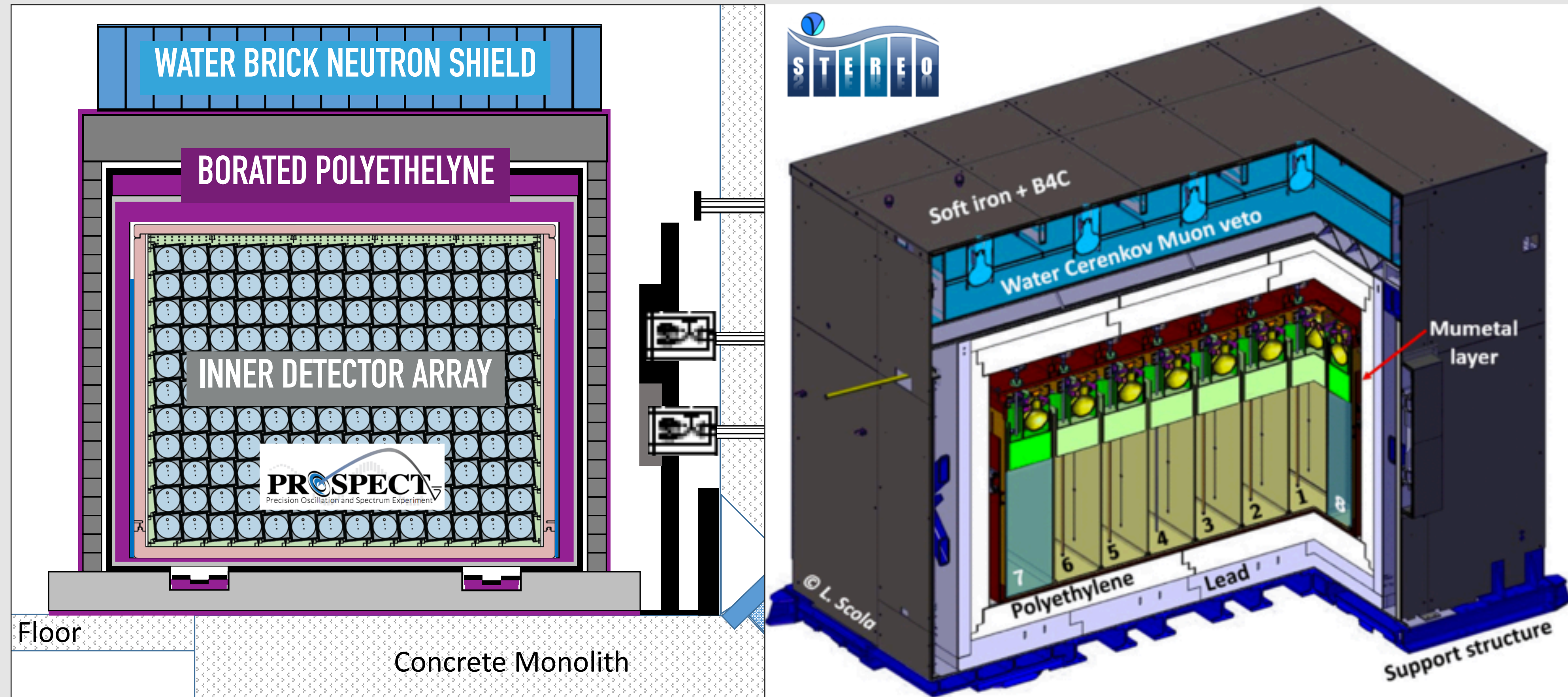
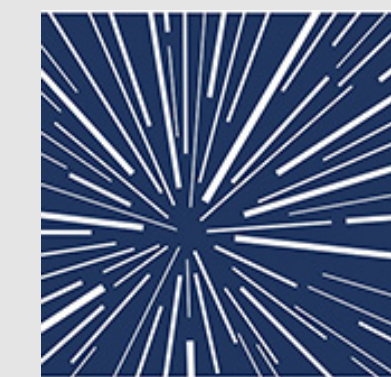


# JOINT MEASUREMENT OF THE $^{235}\text{U}$ ANTINEUTRINO ENERGY SPECTRUM BY PROSPECT AND STEREO



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ON BEHALF OF THE PROSPECT COLLABORATION

[arXiv:2107.03371](https://arxiv.org/abs/2107.03371)

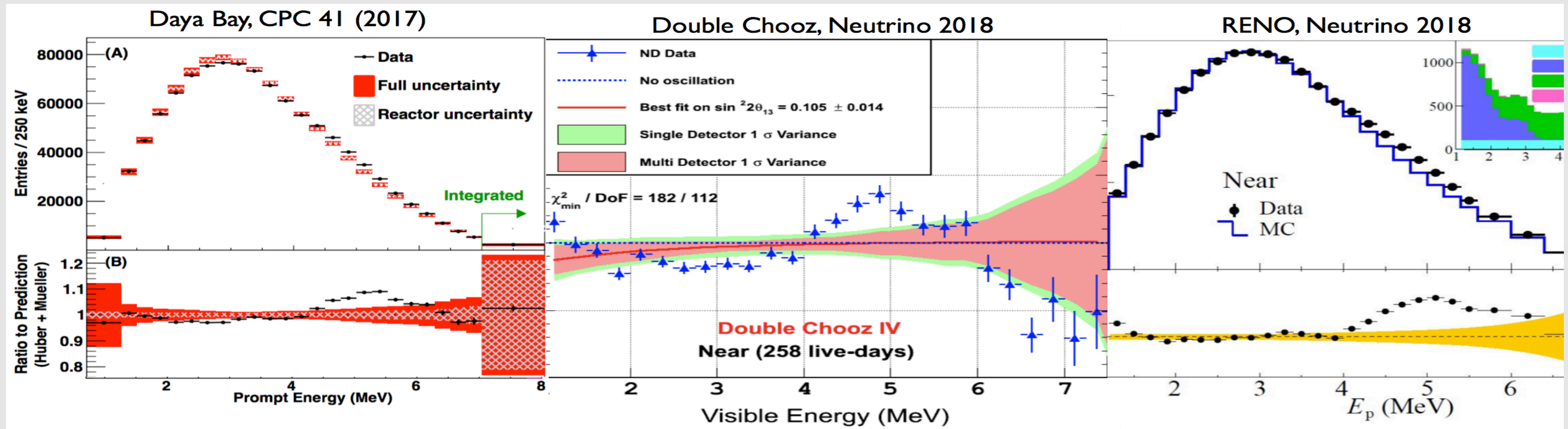


Wright  
Laboratory



# NEUTRINO SPECTRUM MEASUREMENTS FROM POWER REACTORS

- ▶ Spectrum models don't match experimental data in low enriched uranium (LEU) power reactors
- ▶ Poor fit overall to leading reactor models (Huber/Mueller).
- ▶ 'Bump' in 4-6 MeV (prompt energy) range
- ▶ Neutrino events come from a mixture of fissile isotopes:  $^{235}\text{U}$ ,  $^{238}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$
- ▶ Need new reactor data to clarify source of deviations



# WHY A JOINT MEASUREMENT

- ▶ Reactor models do not provide a sufficient prediction of the antineutrino spectrum
- ▶ PROSPECT and STEREO are the leading measurements of the pure  $^{235}\text{U}$  spectrum without significant contributions from other isotopes
- ▶ Both experiments' spectrum measurements are still statistics limited with relatively low systematic uncertainties
- ▶ By combining the measurements, we can increase the statistical power and produce a reference spectrum of  $^{235}\text{U}$  for use by the community

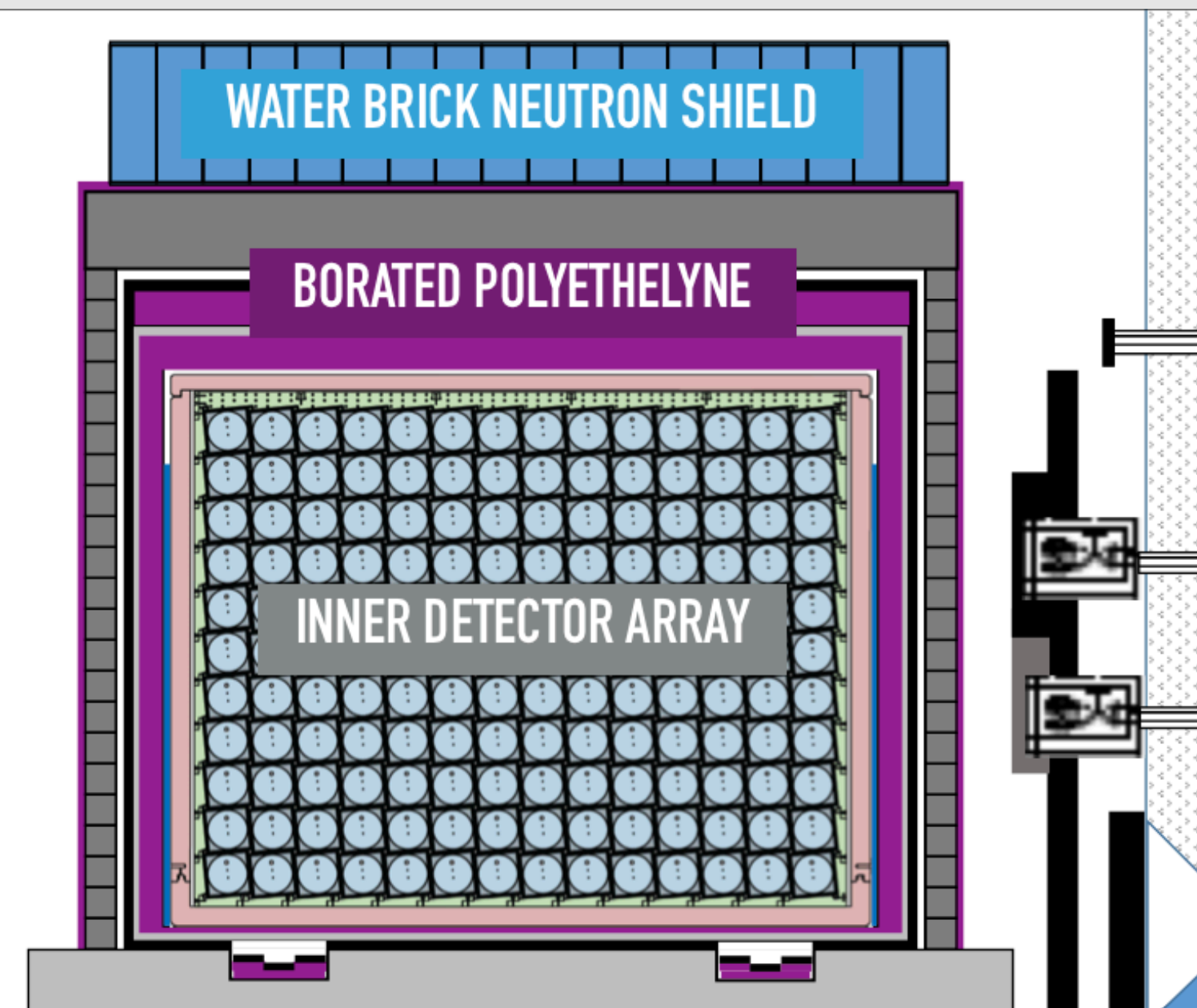
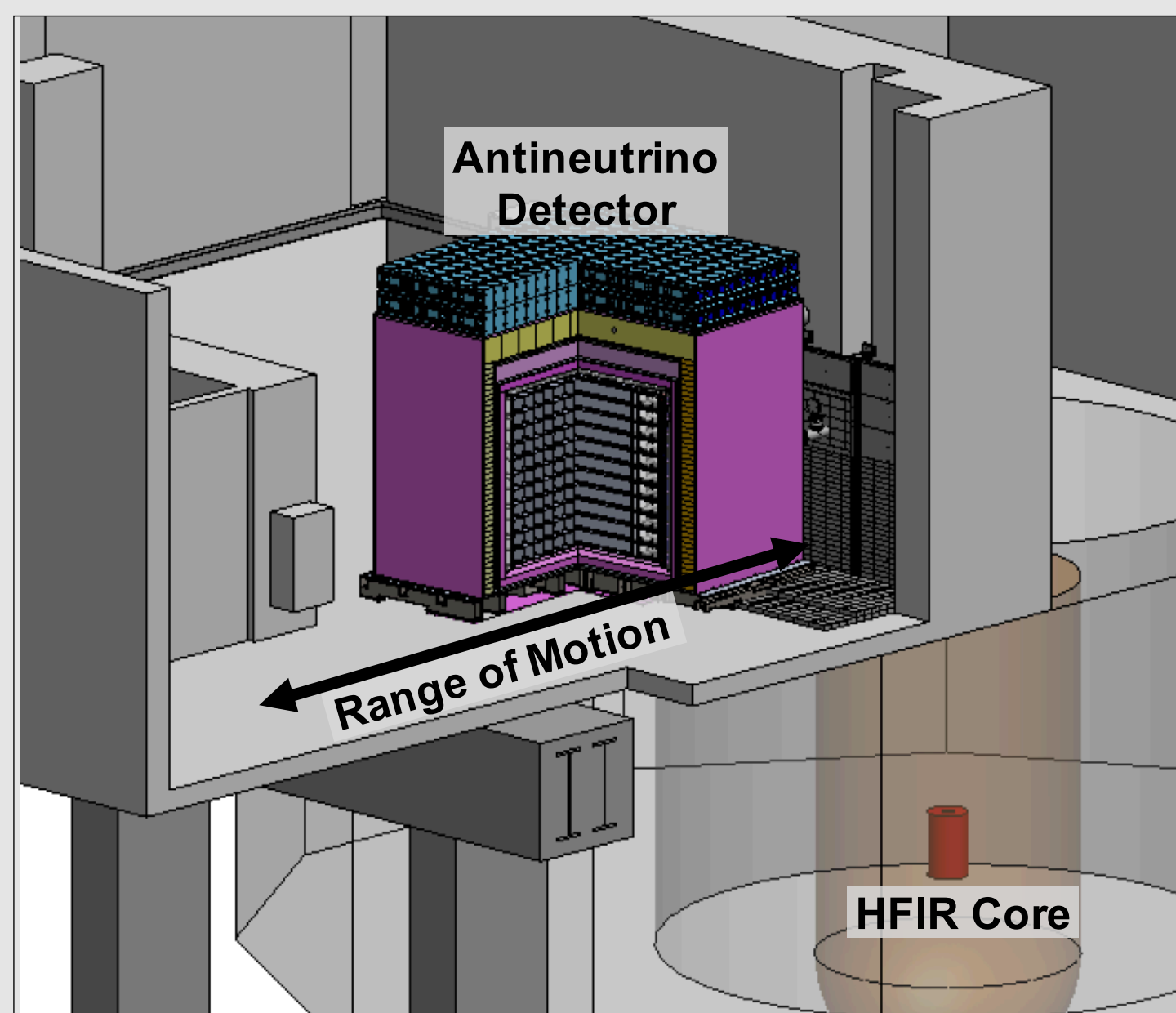
# THE PROSPECT EXPERIMENT

- ▶ Experimental Site (HFIR, ORNL):

- ▶ 85 MW HEU reactor core with 46% duty cycle
- ▶ >99% of  $\bar{\nu}_e$  flux from  $^{235}\text{U}$  fissions

- ▶ Detector Design

- ▶ Segmented design for calibration access
- ▶ Optimized for background suppression
- ▶ Particle identification with pulse shape discrimination



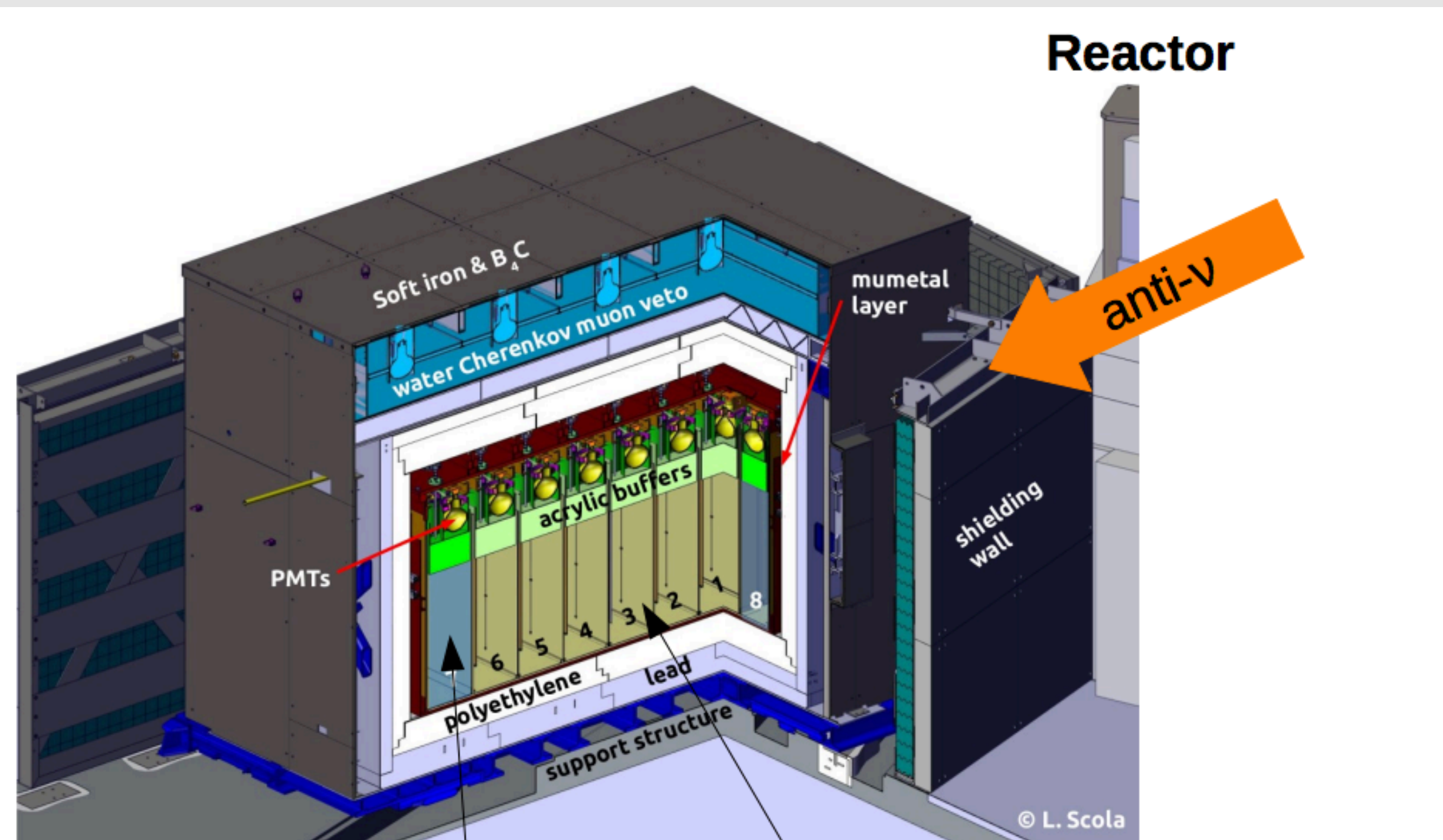
J. Ashenfelter et al., NIM A [2018.12.079](https://doi.org/10.1016/j.nima.2018.12.079)

<https://prospect.yale.edu/>



# THE STEREO EXPERIMENT

- ▶ Experimental site (RHF, ILL):
  - ▶ 58 MW HEU reactor
  - ▶ Compact core
  - ▶ >99% of flux from  $^{235}\text{U}$  fissions
- ▶ Detector Design:
  - ▶ 6 fiducial cells
  - ▶ Liq. Scintillator + Gd
  - ▶ Pulse shape discrimination



**Gamma-Catcher:** unloaded liquid scintillator    **Target:** Gd-loaded liquid scintillator

[arxiv:2010.01876](https://arxiv.org/abs/2010.01876)

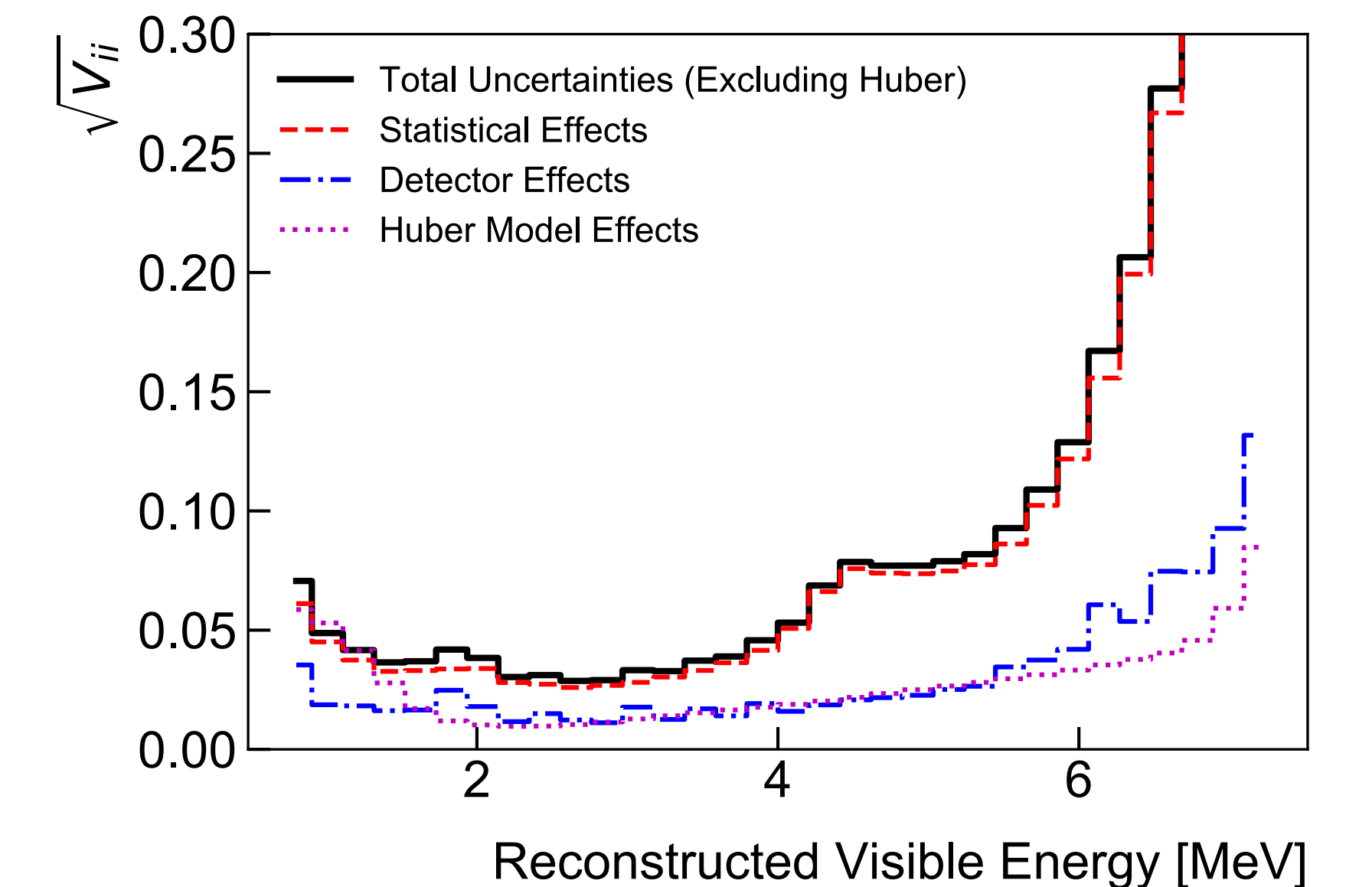
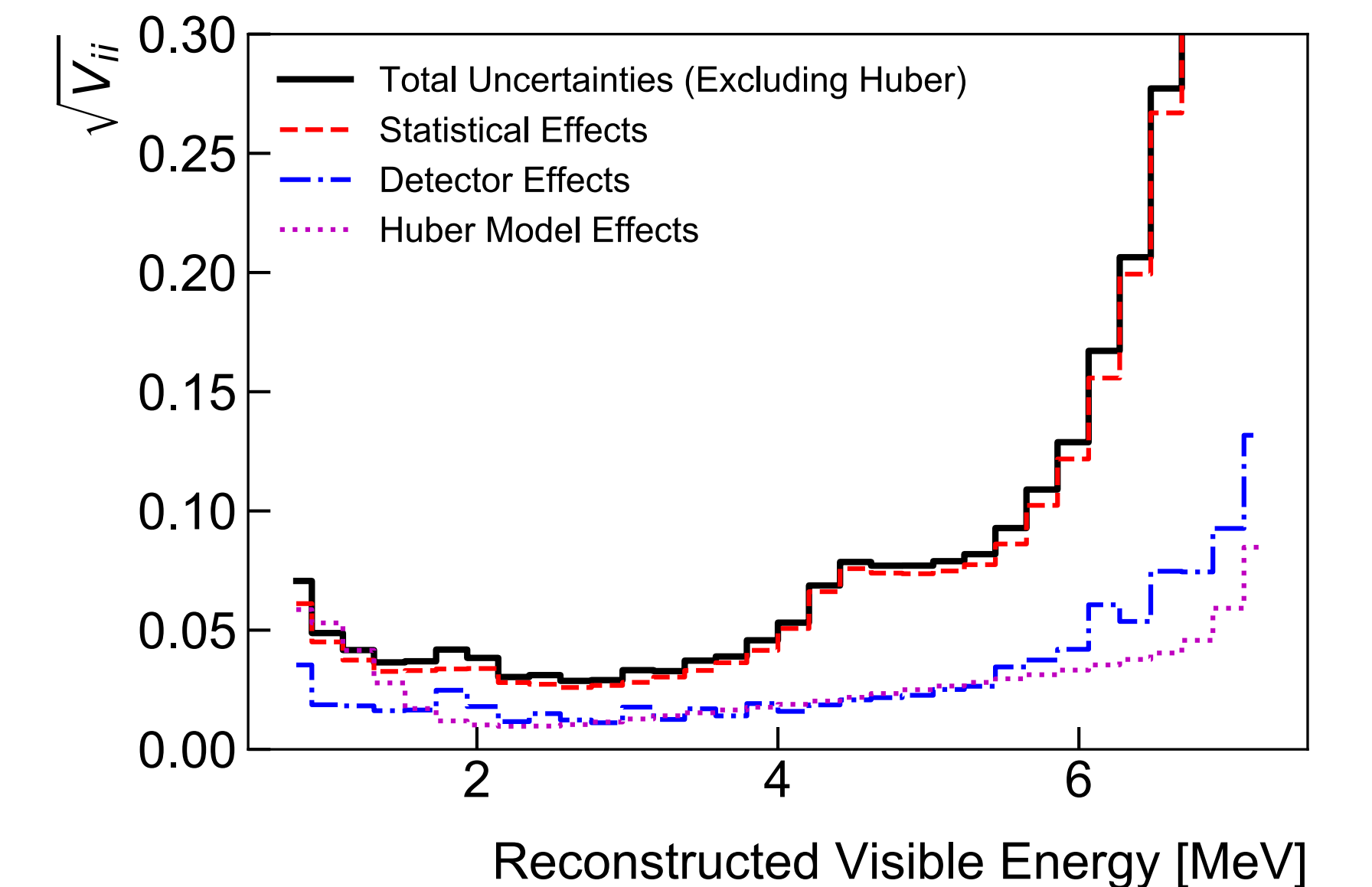
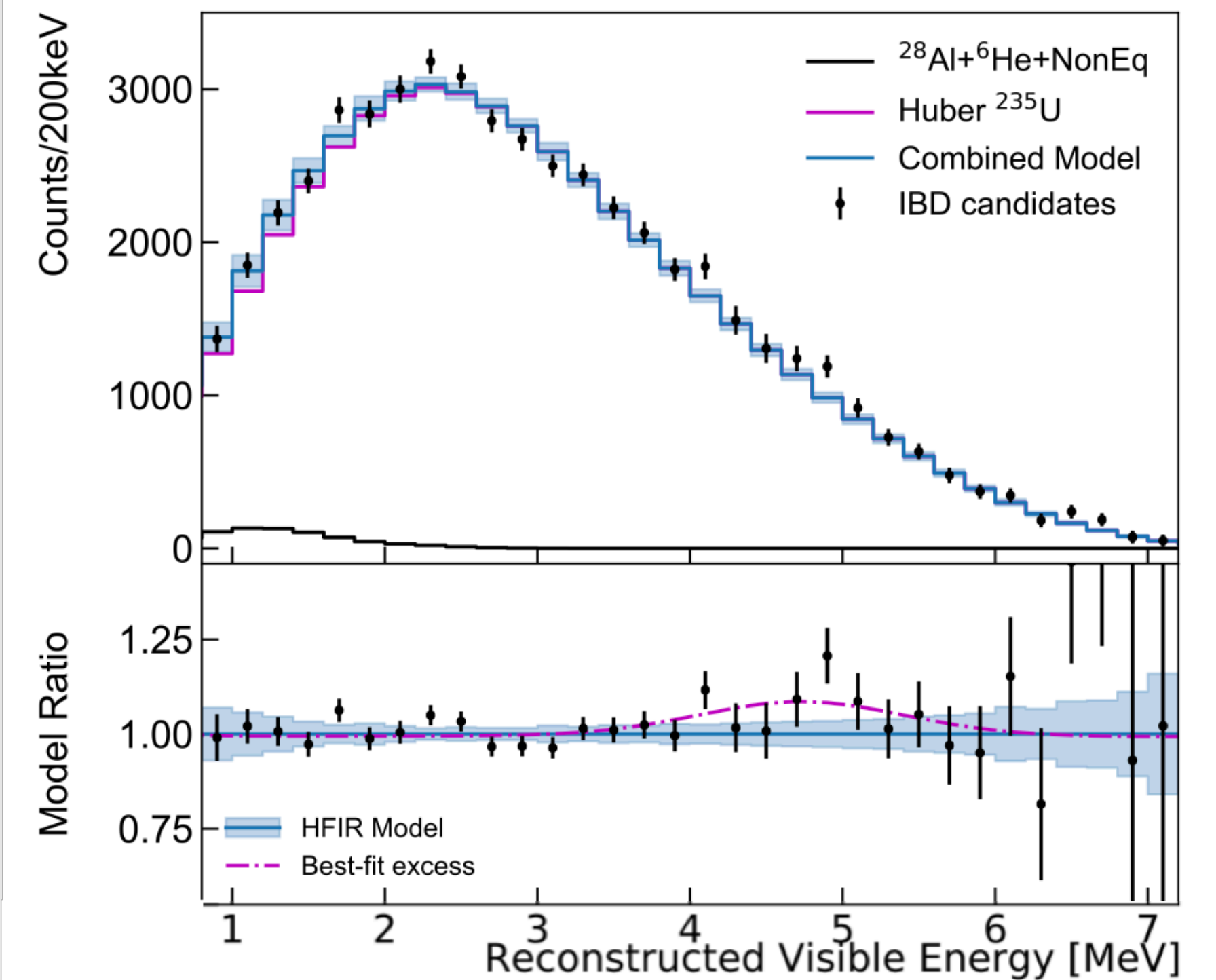
<https://www.stereo-experiment.org/>

# PROSPECT PROMPT SPECTRUM

- ▶ 50560 +/- 406 IBD signal events
- ▶ Best fit bump size relative to Daya Bay: 84% +/- 39%
- ▶ Disfavor both 'No  $^{235}\text{U}$  Contribution to' and 'Only  $^{235}\text{U}$  Contributes to' LEU bump cases at  $>2\sigma$
- ▶ Still statistics limited

[M. Andriamirado et al., Phys Rev D 103, 032001](#)

<https://prospect.yale.edu/>



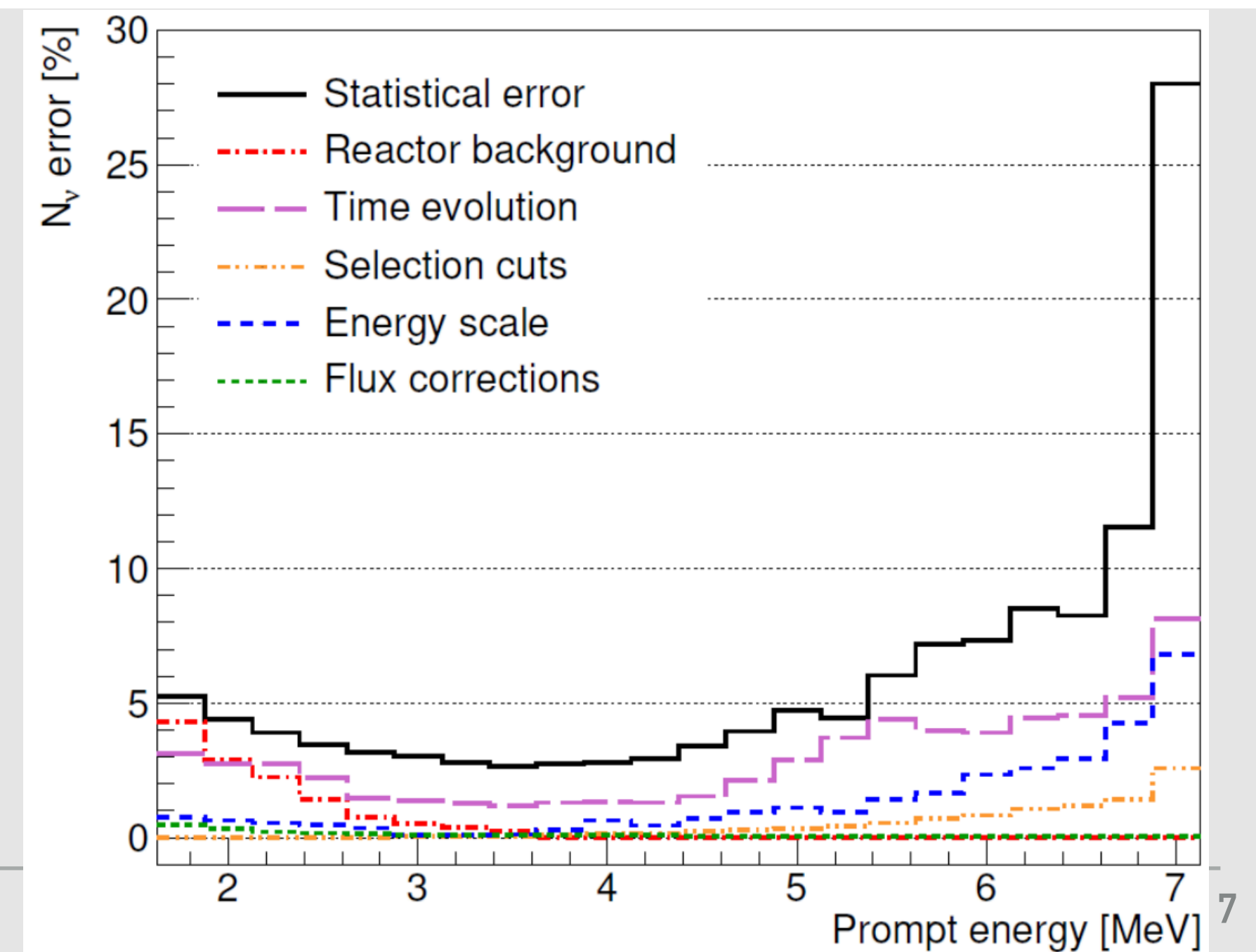
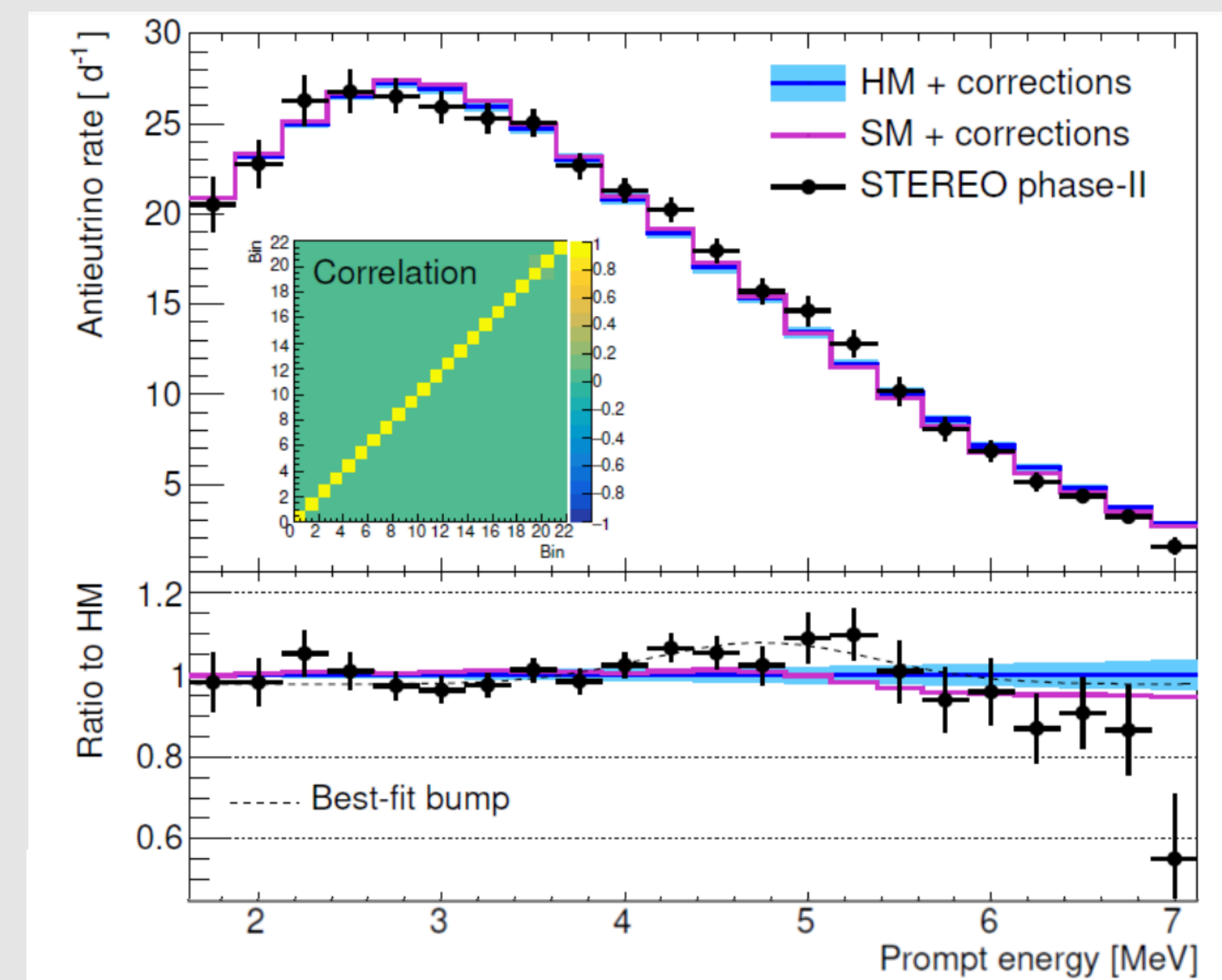


# STEREO'S PROMPT SPECTRUM

- ▶ 43,000 Antineutrinos detected
- ▶ Significant bump observed in antineutrino energy:  $A = 12.1 \pm 3.4 \% (3.5\sigma)$  of spectrum at peak
- ▶ Findings consistent with case of  $^{235}\text{U}$  equally contributing to LEU bump
- ▶ Still statistics limited

[arxiv:2010.01876](https://arxiv.org/abs/2010.01876)

<https://www.stereo-experiment.org/>



# PROMPT COMPATIBILITY

- ▶ Prompt comparison avoids uncertainties of filtered unfolding!
- ▶ Move one experiment's data into the prompt space of the other with unfiltered unfolding, then refolding with the other's response

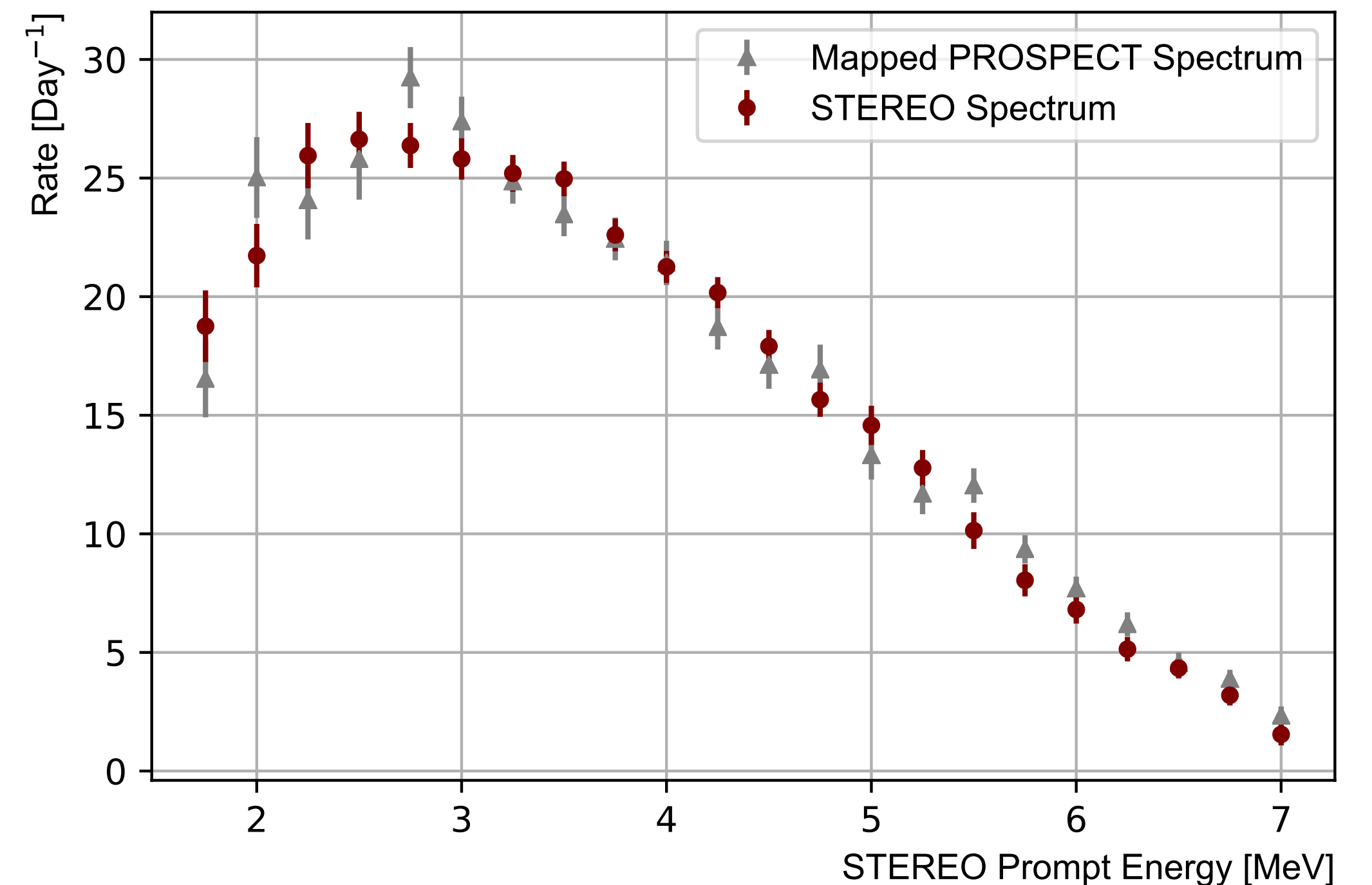
$$R_{map} = R_{STE} \cdot R_{PRO}^{-1}$$

$$M_{map} = R_{map} \cdot M_{PRO}$$

- ▶ Fit spectra with free floating normalization

$$\chi^2/ndf = 24.1/21$$

▶ Statistically Compatible Inputs





# ANALYSIS METHOD: DATA UNFOLDING

- ▶ To create a measurement independent of factors unique to each experiment, we must convert from the prompt space of each to true antineutrino energy space via 'unfolding'

- ▶ Ideal Case: 
$$M = R \times S \Rightarrow S = R^{-1} \times M$$

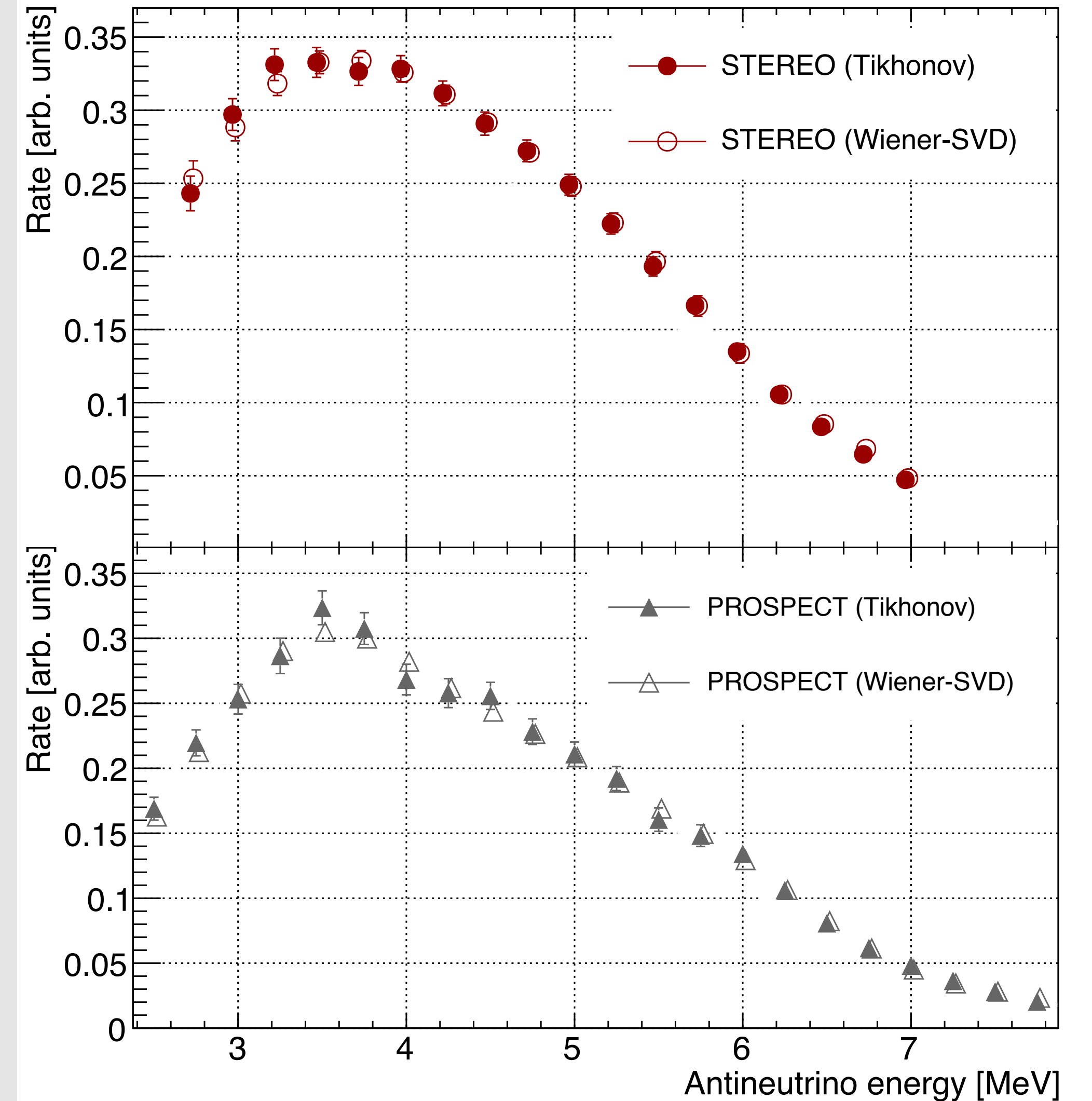
- ▶  $S$  = true signal in neutrino energy
  - ▶  $R$  = response matrix
  - ▶  $M$  = measured signal in prompt energy
- 
- ▶ Realistically:
    - ▶  $R$  not necessarily invertible
    - ▶  $M$  has non-signal noise elements which are blown out of proportions by  $R^{-1}$

# COMPARISON OF FRAMEWORKS

## ► Framework Validation:

1. STEREO's Tikhonov regularization
2. PROSPECT's WienerSVD unfolding method

► Cross-checked and Consistent Results

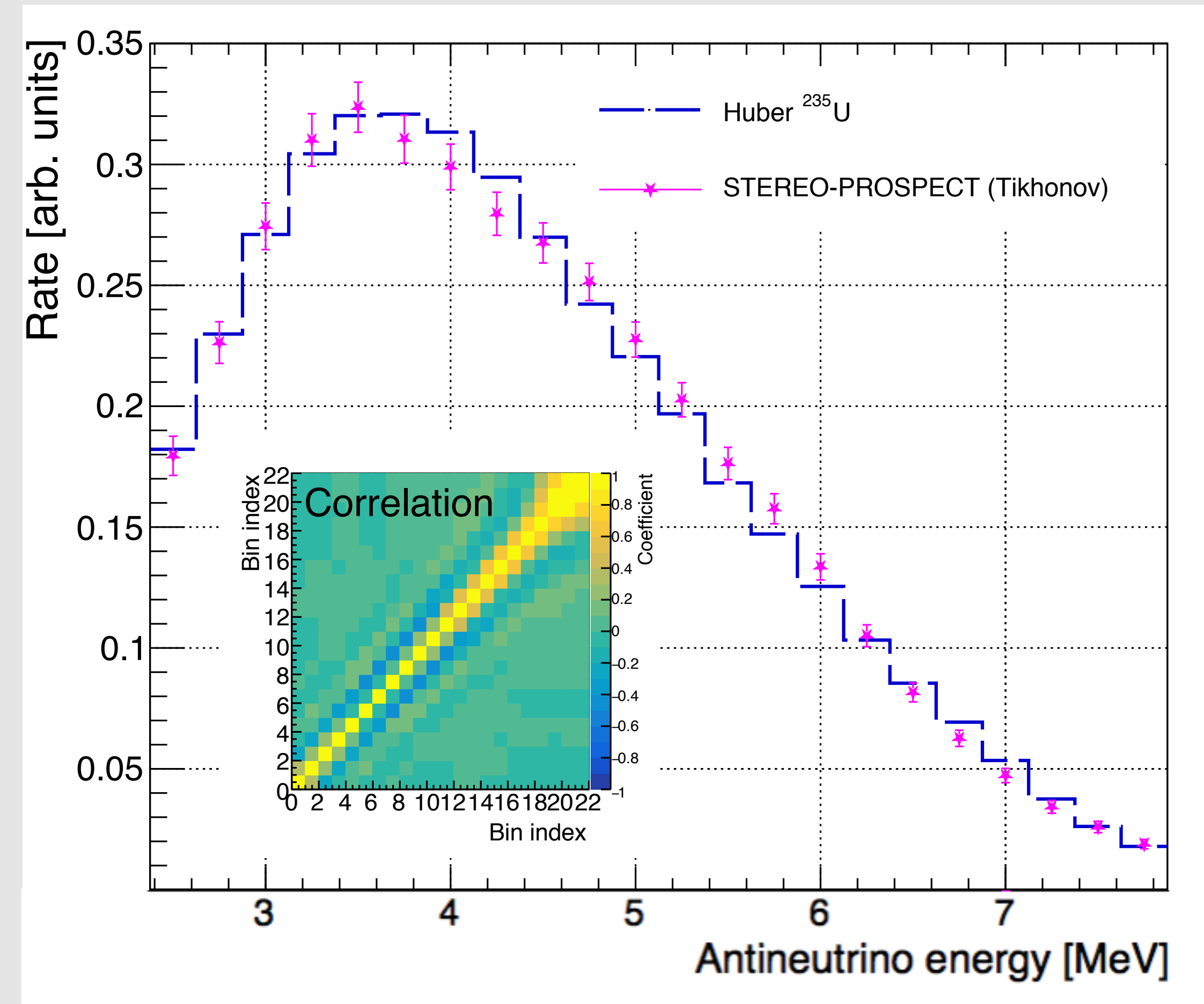




# UNFOLDED SPECTRUM

- ▶ Use the Tikhonov method to present result
- ▶ Using a free floating normalization, best fit to Huber model gives  $\chi^2/ndf = 30.8/21$

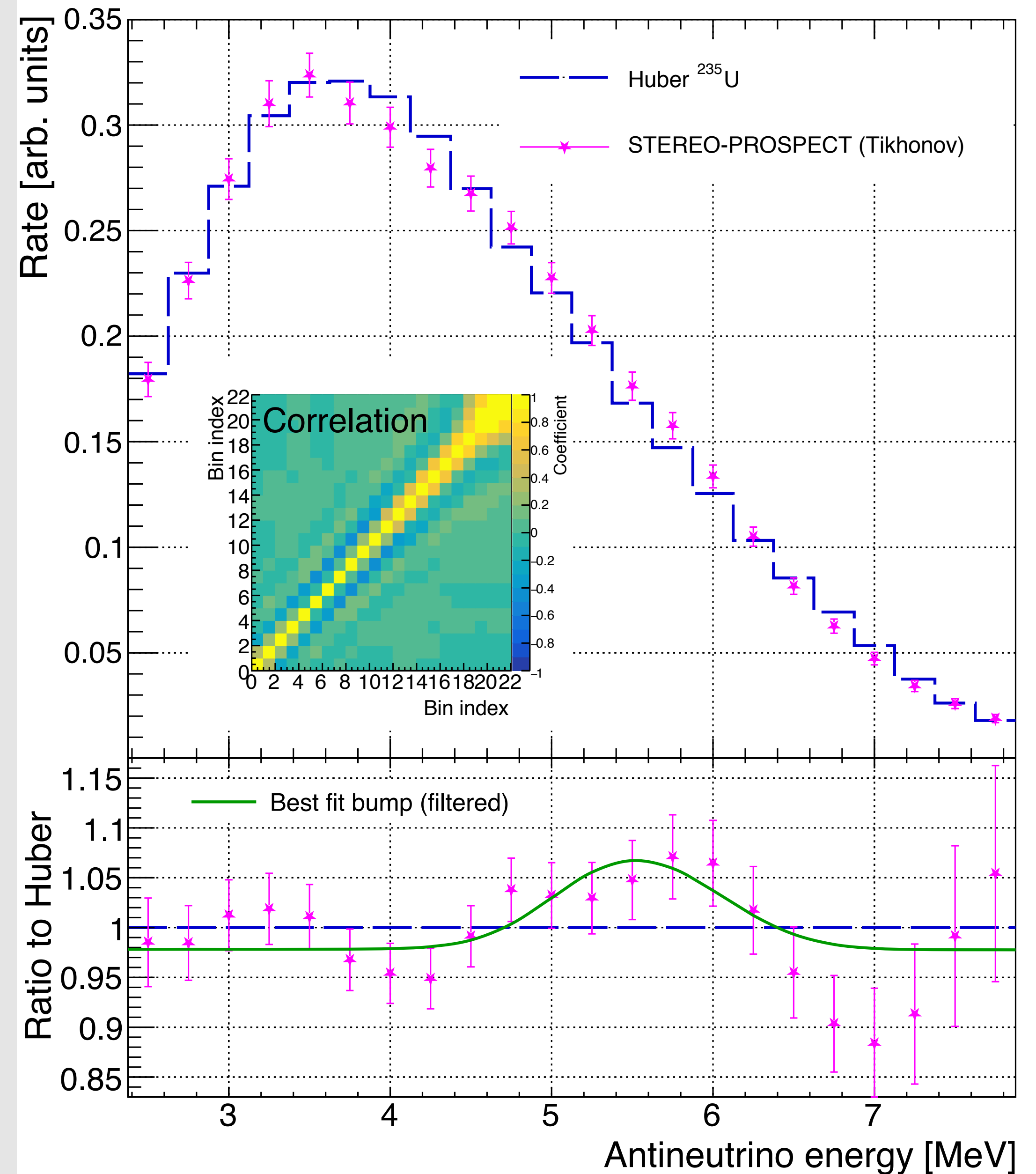
- ▶ Results available ([arXiv:2107.03371](https://arxiv.org/abs/2107.03371))
- ▶ Thorough supplemental materials, including filter matrix
- ▶ Can be directly compared to  $^{235}\text{U}$  model predictions



# BUMP SEARCH

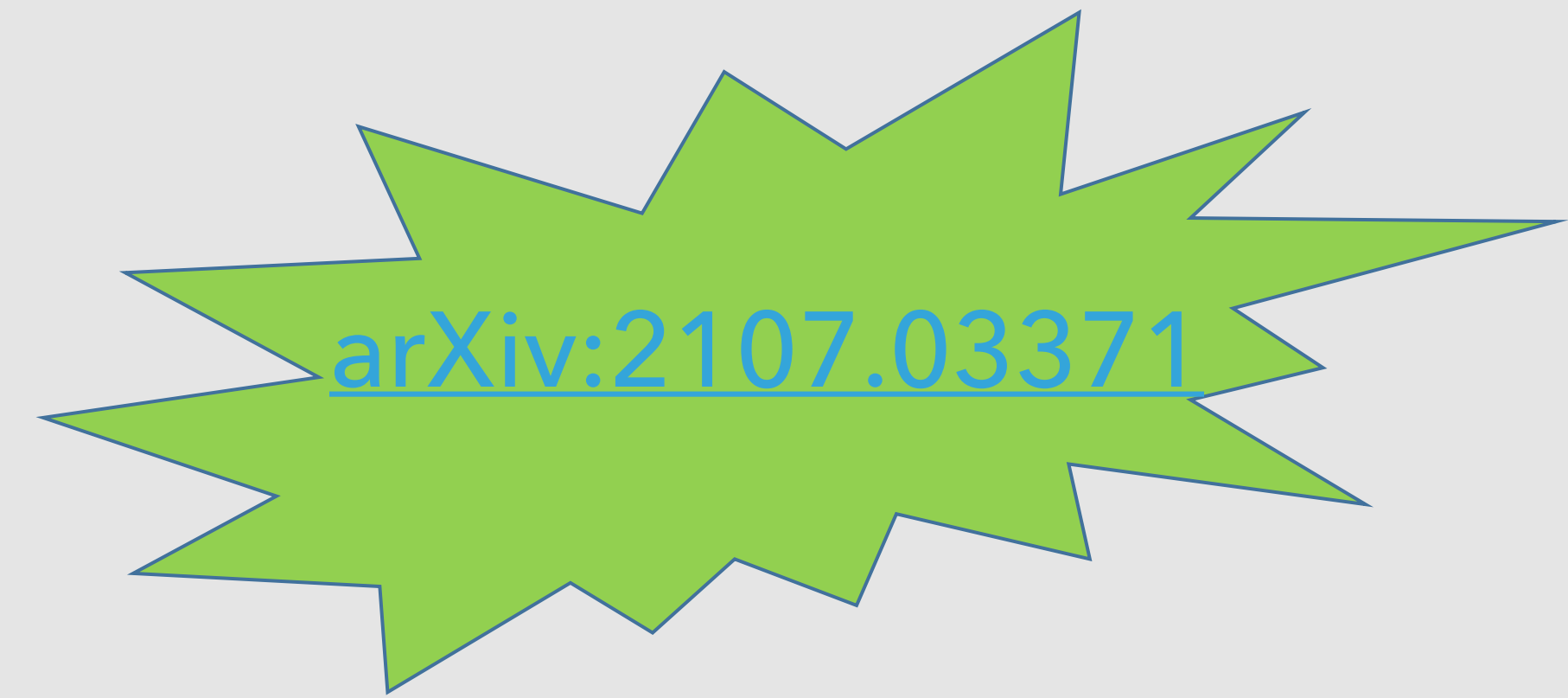
- ▶ Find an excess in the 5-6 MeV range
- ▶ Fit a Gaussian with free amplitude, mu, and sigma values to the excess
- ▶ The addition of the best-fit Gaussian to the Huber model gives  $\Delta\chi^2/\Delta ndf = 12.0/3$  ( $2.4\sigma$  significance)
- ▶ Consistent with the Daya Bay  $^{235}\text{U}$  spectrum in shape-only comparison ( $\chi^2/ndf = 21.0/21$ )

- ▶ Find an excess with significance  $2.4\sigma$
- ▶ Consistent with  $^{235}\text{U}$  equally contributing to LEU excess



# CLOSING STATEMENTS

- ▶ New results posted to arxiv just last week!



- ▶ PROSPECT and STEREO datasets found to be statistically compatible
- ▶ PROSPECT and STEREO have successfully combined their separately measured high precision  $^{235}\text{U}$  spectra
- ▶ The publication of the jointly unfolded result includes filter matrix for comparing to  $^{235}\text{U}$  antineutrino models, can be used as a reference spectrum by community
- ▶ Find an excess with  $2.4\sigma$  significance in the 5-6 MeV energy range



# PROSPECT

[prospect.yale.edu](http://prospect.yale.edu)



Funding provided by:



U.S. DEPARTMENT OF ENERGY



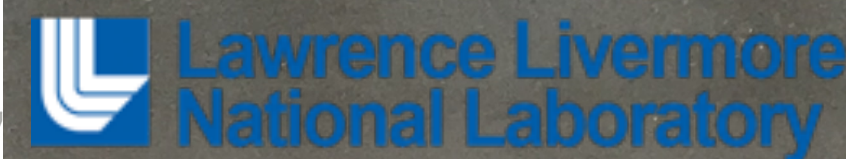
15 Institutions, 70 collaborators



NIST



W&M



Yale



# BACKUP SLIDES

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# COMPARISON TO DAYA BAY 235U SPECTRUM

- ▶ Use Daya Bay's published pure 235U spectrum (PRL 123, 111801 (2019))
- ▶ Interpolate Daya Bay's spectrum to match the joint binning
- ▶ Perform chi2 comparison, with a free-floating normalization parameter
- ▶  $\chi^2/ndf = 21.0/21$



# SMEARING MATRIX

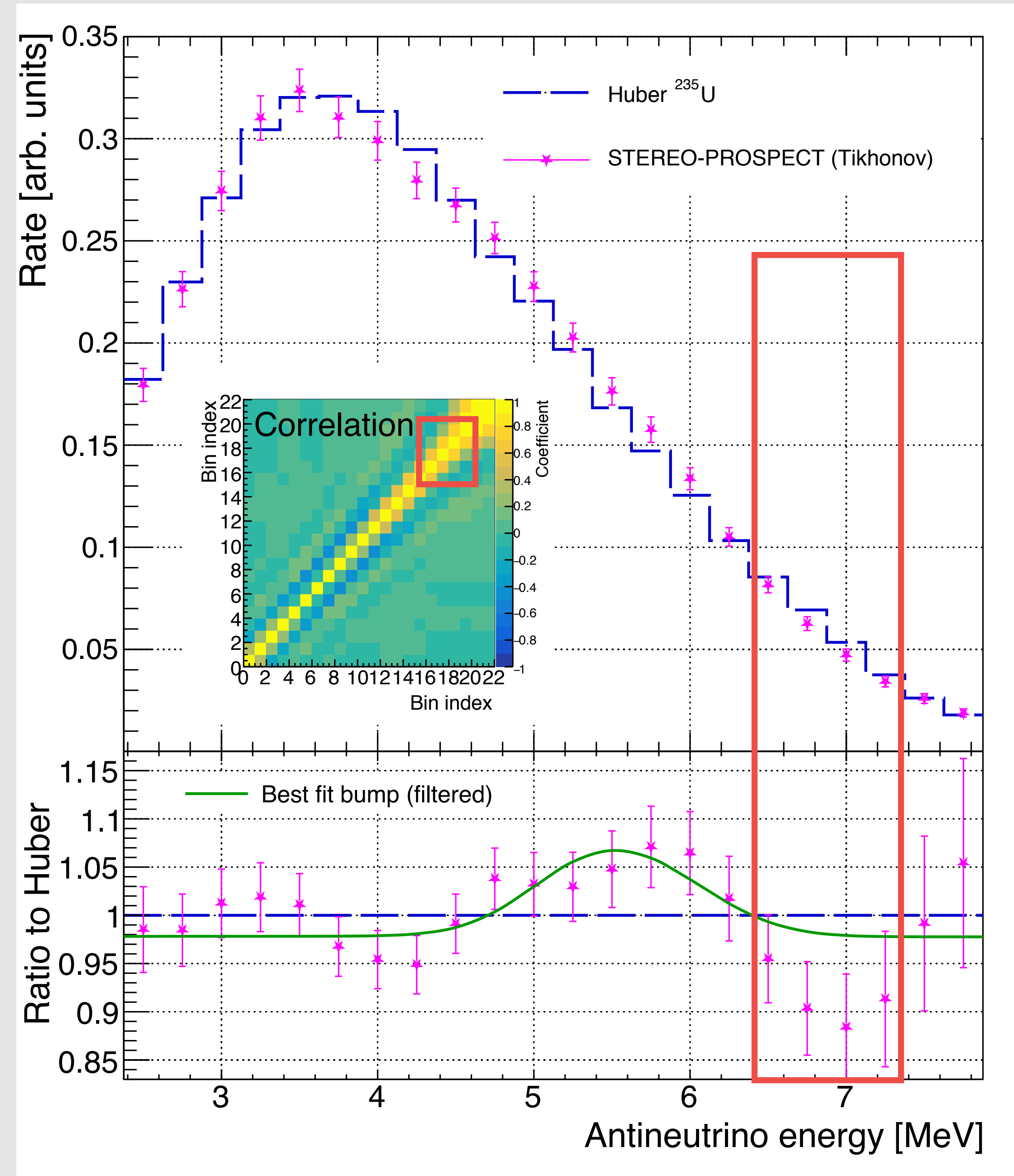
- ▶ The  $A_C$ , or smearing, matrix is used to apply the bias and smoothing effect from unfolding to a given spectrum in antineutrino energy
- ▶ This can and should be applied to models in true antineutrino energy to allow for an accurate comparison with the unfolded data
- ▶ The  $A_C$  matrix for the WienerSVD is:

$$D_{jnt} = \overset{A_{Cjnt}}{\boxed{C^{-1}V_C W_C V_C^T C (R^T (R R^T)^{-1} M_{jnt})}}$$

- ▶ For more information, please refer to [arxiv:1705.03568](https://arxiv.org/abs/1705.03568)

# 7 MEV DEFICIT SIGNIFICANCE

- ▶ Visually apparent deficit at 7 MeV (in the LEU bump region)
- ▶ Error bars imply more significant of a distortion than reality
- ▶ 7+ MeV region has much higher correlations than rest of spectrum
- ▶ Fitting a negative amplitude Gaussian shows significance of only  $1.3\sigma$



# PROSPECT METHOD: WIENER-SVD UNFOLDING

- ▶ Singular Value Decomposition (SVD) on  $R$

$$R = UDV^T$$

- ▶ Where  $D$  is a diagonal matrix of singular value in decreasing order ( $U$  and  $V$  are orthogonal matrices)
- ▶ If the elements  $d_i$  in  $D$  get too small, then  $D^{-1}$  will cause noise in  $M$  to blow up when calculating  $S$ .
- ▶ Use a Wiener filter  $W$  to keep  $d_i$  from getting too small

$$S = R^{-1} \times M = V(WD)^{-1}U^T \times M$$

- ▶ See arXiv 1705.03568



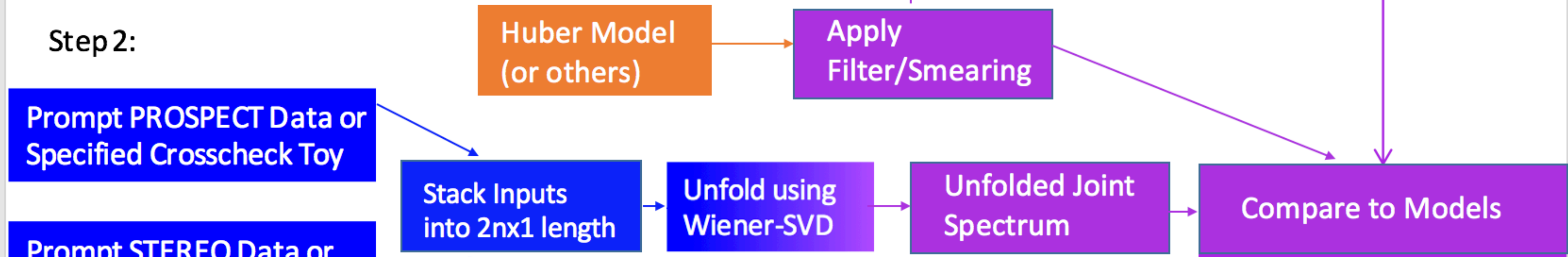
# JOINT WIENERSVD UNFOLDING


Uses 10,000 Huber-based toys produced by with Huber Model, folded into prompt energy, and given real, data-driven fluctuations



Step 1:



Step 2:



 = Prompt Energy

 = Smeared Energy  
 = Neutrino Energy

\*Incorporates a 3% uncertainty on Huber Model for unfolding

# STEREO METHOD: TIKHONOV REGULARIZATION

❖ Writing of the regularized  $\chi^2$  to be minimized:

$$\chi_{joint}^2(\Phi, \alpha) = (D_{ST} - R_{ST}\Phi)^T V_{ST}^{-1} (D_{ST} - R_{ST}\Phi) + (D_{PR} - \alpha * R_{PR}\Phi)^T V_{PR}^{-1} (D_{PR} - \alpha * R_{PR}\Phi) + \lambda * R(\Phi)$$

- $\alpha$  is a normalization parameter to account for PROSPECT's spectrum normalization.
- $R(\Phi)$  is a regularization term that can be written with the first or second discrete derivative of the deviation to the HM prior shape:

$$R(\Phi) = R_1(\Phi) = \sum_i \left( \frac{\Phi_{i+1}}{\Phi_{i+1}^{HM}} - \frac{\Phi_i}{\Phi_i^{HM}} \right)^2 := \Phi^T M_1 \Phi$$

$$R(\Phi) = R_2(\Phi) = \sum_i \left( \frac{\Phi_{i+1}}{\Phi_{i+1}^{HM}} - 2 \frac{\Phi_i}{\Phi_i^{HM}} + \frac{\Phi_{i-1}}{\Phi_{i-1}^{HM}} \right)^2 := \Phi^T M_2 \Phi$$

- $\lambda$  is the regularization strength trading-off the agreement to the data and the smoothness of the unfolding. It is tuned following the GCV prescription<sup>1</sup>.

❖ Unfolded spectrum reads:

$$\hat{\Phi} = \underbrace{\left( R_{ST}^T V_{ST}^{-1} R_{ST} + \alpha^2 R_{PR}^T V_{PR}^{-1} R_{PR} \right)}_{\text{Classical unfolding part (unregularized)}} \underbrace{+ \lambda M_{1,2}}_{\text{Regularization contribution}} \underbrace{\left( R_{ST}^T V_{ST}^{-1} D_{ST} + \alpha R_{PR}^T V_{PR}^{-1} D_{PR} \right)}_{\text{Data Spectra}}$$

14/8/21

11

<sup>1</sup>Grace Wahba, Gene H. Golub, Michael Heath, "Generalized Cross-Validation as a Method for Choosing a Good Ridge Parameter, Technometrics", Vol. 21, NO. 2, May 1979.