

Status of the DUNE experiment

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XXVIII Cracow EPIPHANY Conference

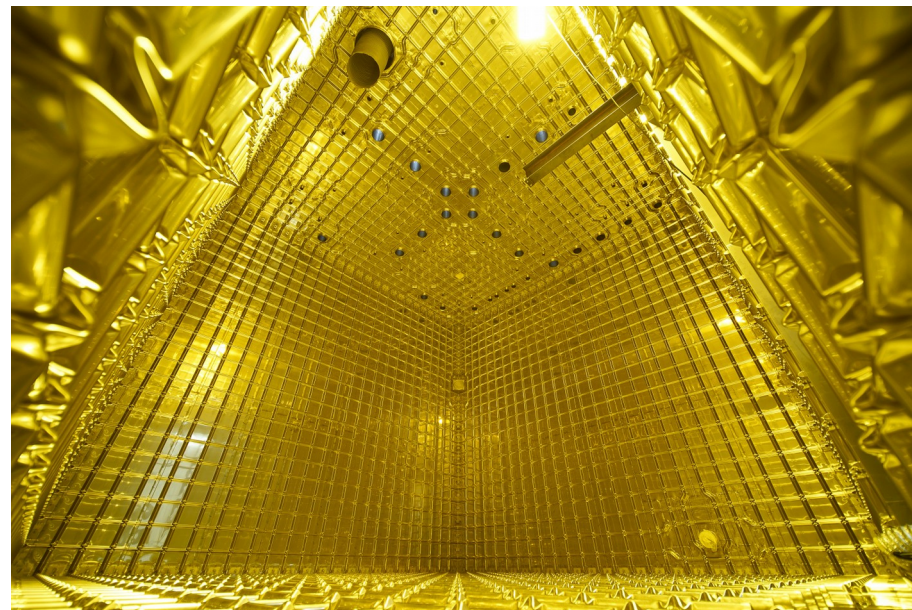
on Recent Advances in Astroparticle Physics

10-14 January 2022



Introduction

- A crash course in neutrino physics
- The DUNE experiment and its main elements
- DUNE physics goals
- DUNE prototypes and plans





Crash Course in Neutrino Physics

- Three neutrino flavors: ν_e , ν_μ and ν_τ

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

- Neutrinos have mass and can oscillate changing flavour.

$$P(\nu_x \rightarrow \nu_y) = \underbrace{\sin^2(2\theta)}_{\text{“amplitude”}} \underbrace{\sin^2\left(1.27 \Delta m^2 \frac{L(\text{km})}{E(\text{GeV})}\right)}_{\text{“frequency”}}$$

Two flavour approximation is good enough in most cases.

- Energy (E) and baseline (L) define which parameters we can probe.





The Current State of Neutrino Knowledge

23

13

12

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Atmospheric

Reactor/Interference

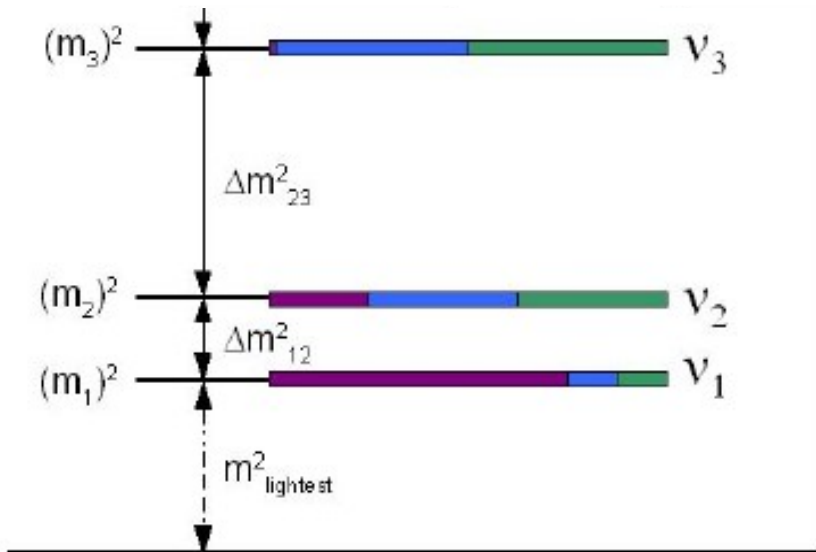
Solar

θ_{13} was first measured less than 10 years ago.

$\mu \Rightarrow \tau$

$\mu \Leftrightarrow e$

$e \Leftrightarrow \mu$



**“Known”
neutrino
physics**

$\Delta m_{21}^2 [10^{-5} \text{eV}^2]$ 7.56 ± 0.19

$|\Delta m_{31}^2| [10^{-3} \text{eV}^2]$ (NO) 2.55 ± 0.04

$|\Delta m_{31}^2| [10^{-3} \text{eV}^2]$ (IO) 2.49 ± 0.04

$\theta_{12}/^\circ$ $34.5_{-1.0}^{+1.1}$

$\theta_{13}/^\circ$ $8.44_{-0.15}^{+0.18}$

$\sin^2 \theta_{23}/10^{-1}$ (NO) $4.30_{-0.18}^{+0.20}$

$\theta_{23}/^\circ$ 41.0 ± 1.1

$\sin^2 \theta_{23}/10^{-1}$ (IO) $5.96_{-0.18}^{+0.17}$

$\theta_{23}/^\circ$ 50.5 ± 1.0

ν_e ν_μ ν_τ ν_s

Salas, Forero, Ternes, Tortola, Valle: 2017

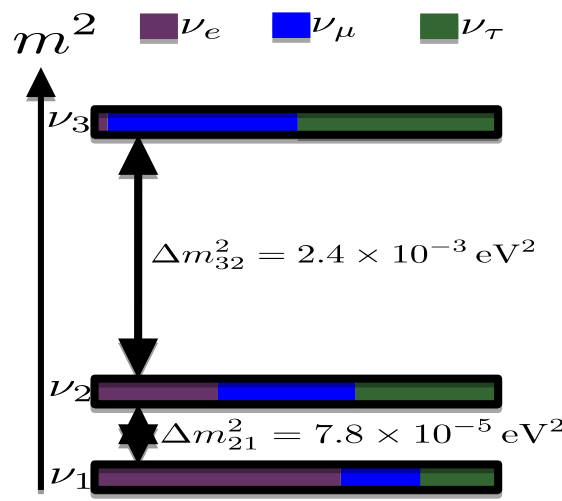




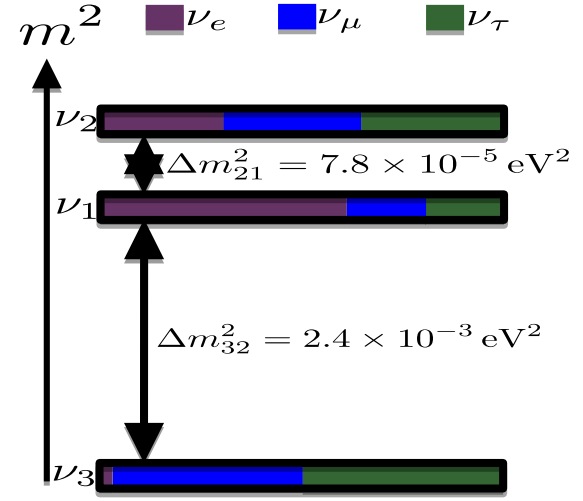
Open Questions in (oscillation) Neutrino Physics

- Neutrino Mass Hierarchy
- CP Violation – is the δ phase non-zero?
- Is θ_{23} maximal?
- *Are there more than 3 neutrino families (sterile neutrinos?)*

Probed with oscillations from muon-neutrino beams.
 Searching for electron-neutrinos (ν_e) appearing or muon-neutrinos (ν_μ) disappearing



Normal



Inverted



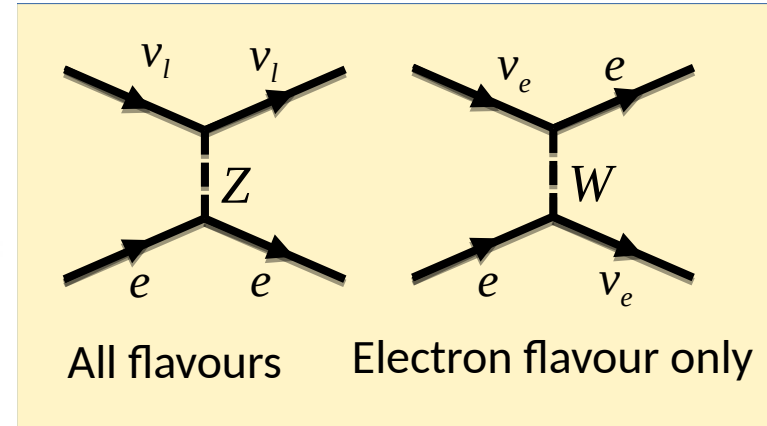


Long-Baseline Oscillations

$$\begin{aligned}
 P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) &\simeq \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2(\Delta_{31} - aL)}{(\Delta_{31} - aL)^2} \Delta_{31}^2 \\
 &+ \sin 2\theta_{23} \sin 2\theta_{13} \sin 2\theta_{12} \frac{\sin(\Delta_{31} - aL)}{(\Delta_{31} - aL)} \Delta_{31} \\
 &\times \frac{\sin(aL)}{(aL)} \Delta_{21} \cos(\Delta_{31} \pm \delta_{CP}) \\
 &+ \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(aL)}{(aL)^2} \Delta_{21}^2,
 \end{aligned}$$

$$a = \pm \frac{G_F N_e}{\sqrt{2}} \approx \pm \frac{1}{3500 \text{ km}} \left(\frac{\rho}{3.0 \text{ g/cm}^3} \right),$$

- Matter effects (mass hierarchy)
- CP violation
- θ_{23} octant



Matter Effects

Need to:

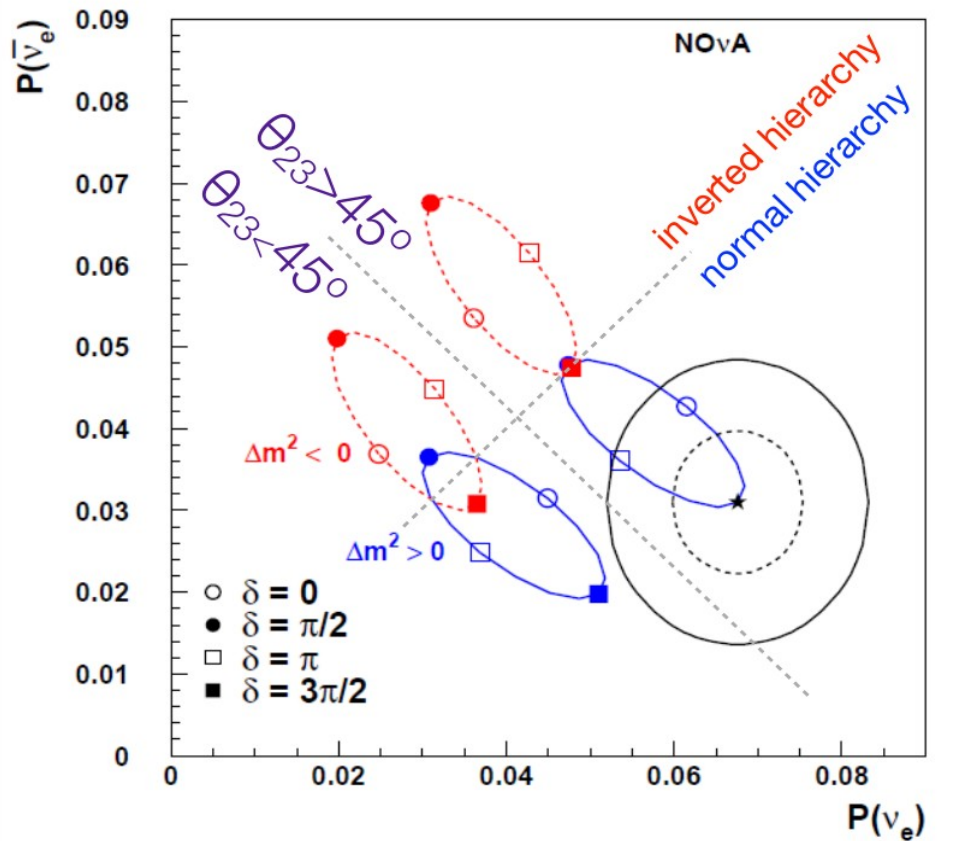
- measure ν_μ and ν_e after oscillation (and their energies)
- Repeat for $\bar{\nu}_\mu$ and $\bar{\nu}_e$
- Know the amount of ν_μ and ν_e before oscillation (near detector)
- Be able to differentiate effects of mass hierarchy, CPV and θ_{23}





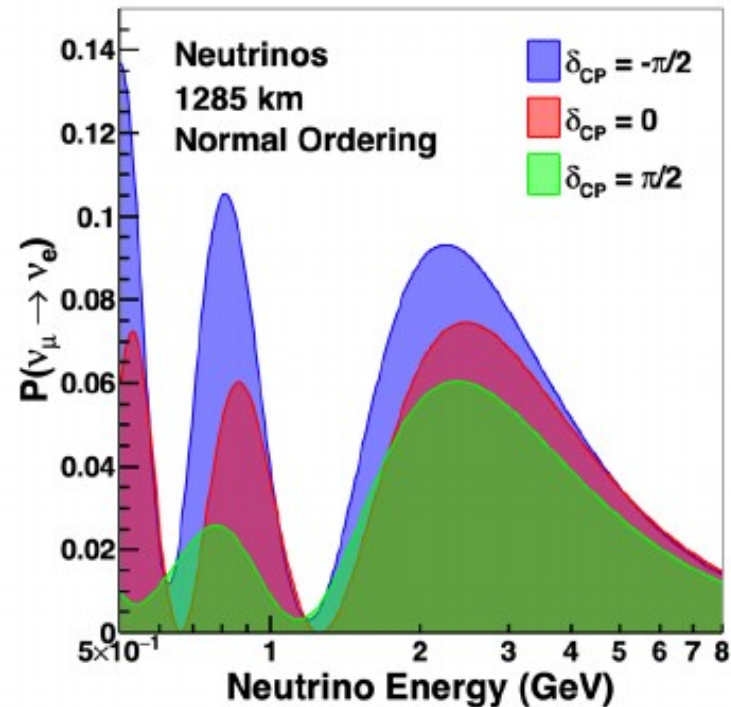
Long-Baseline Oscillations

1 and 2 σ Contours for Starred Point



M. Messier

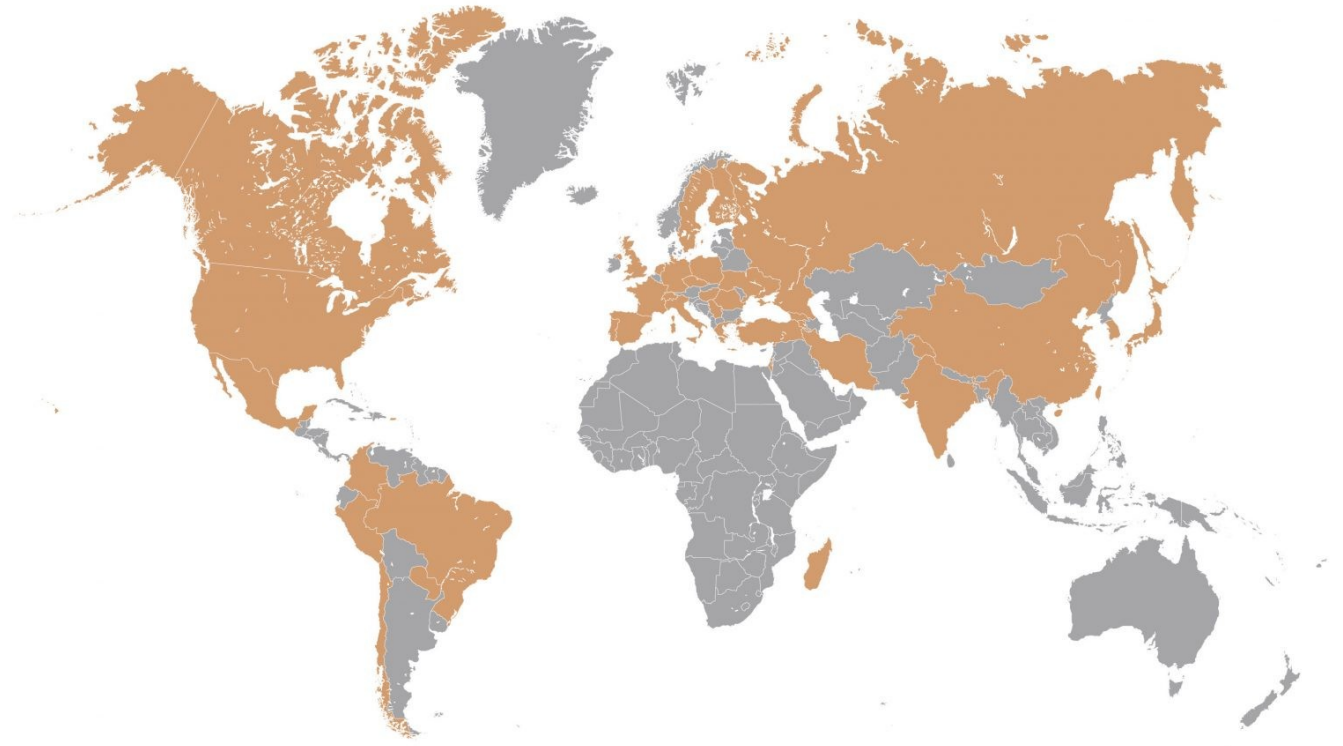
	$\nu_{\mu} \rightarrow \nu_e$	$\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$
Normal hierarchy	Enhanced	Decreased
Inverted hierarchy	Decreased	Enhanced





DUNE collaboration

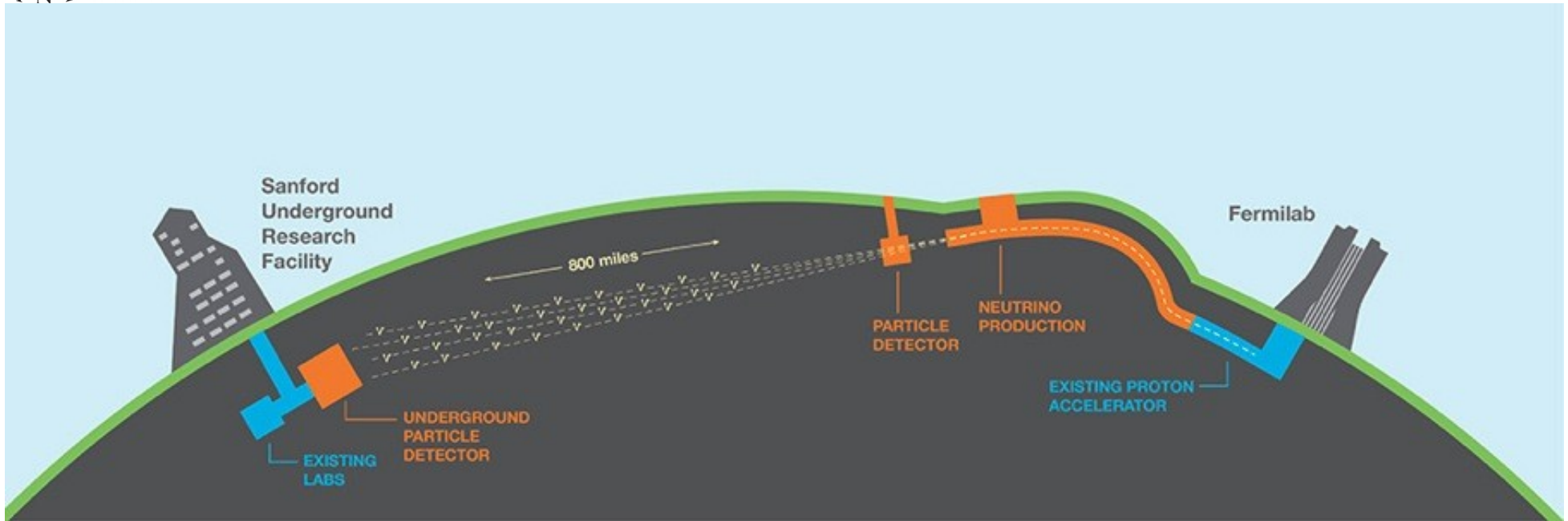
Over 1300
collaborators from
over 200 institutions
in over 30 countries
plus CERN



DUNE CM 2020
At CERN



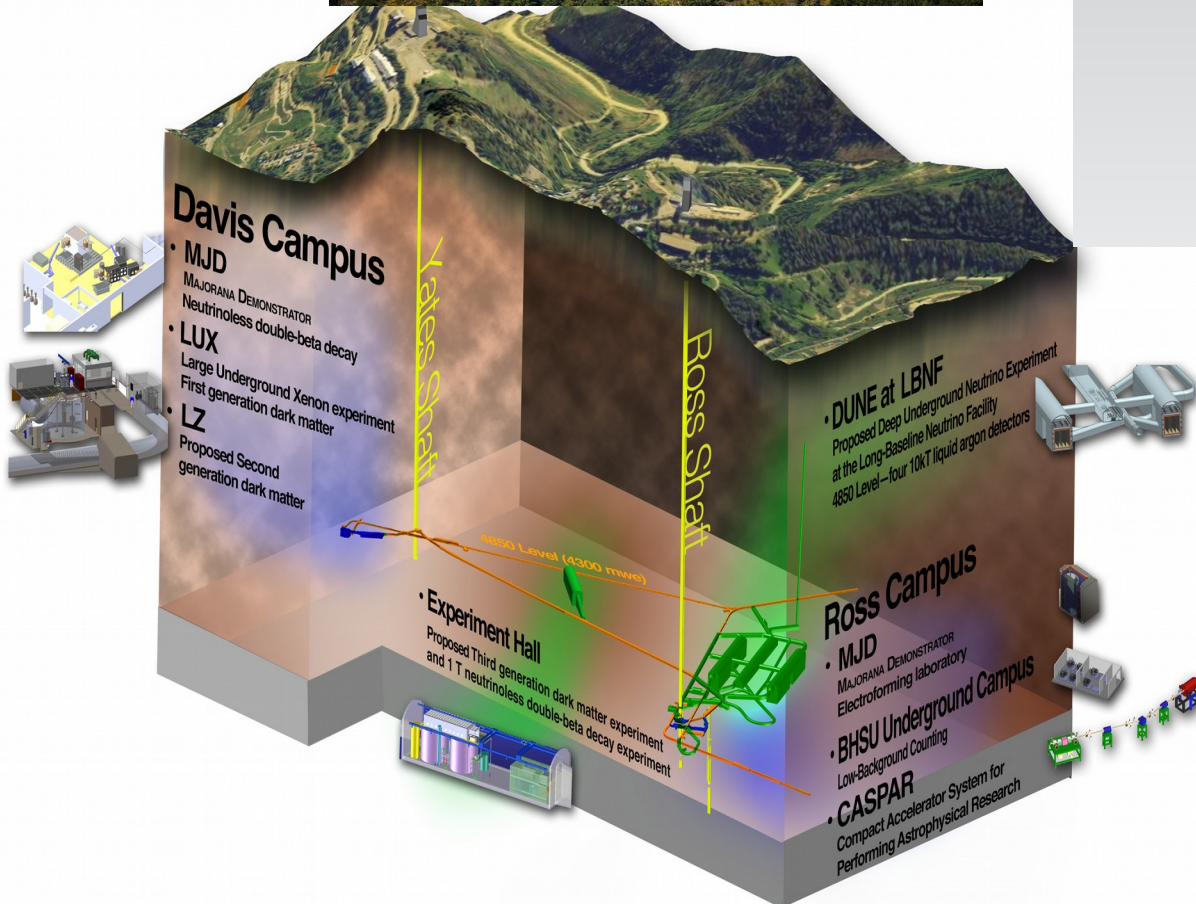
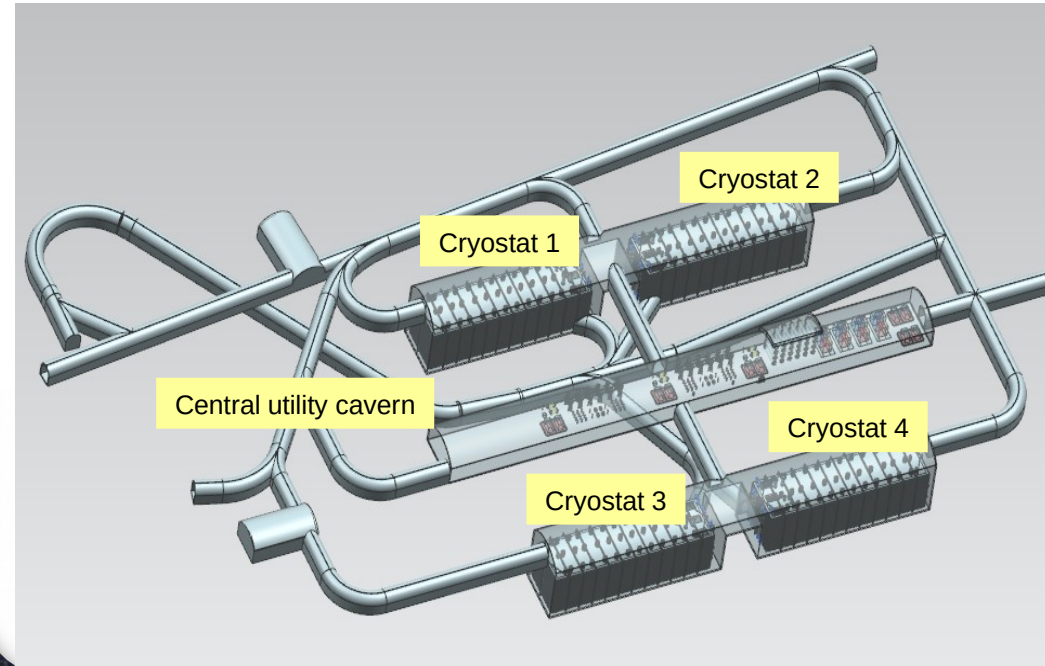
DUNE at a Glance



- High-intensity neutrino beam (1.2 MW – upgradeable to 2.4 MW)
- 70 kTons of liquid argon, 1.5 km underground.
- Detectors on-axis, observing a wide-band neutrino beam.
- 1300 km baseline, ample to provide matter effects.
- Multi-element near detector for excellent control of systematic effects.



The FD Complex

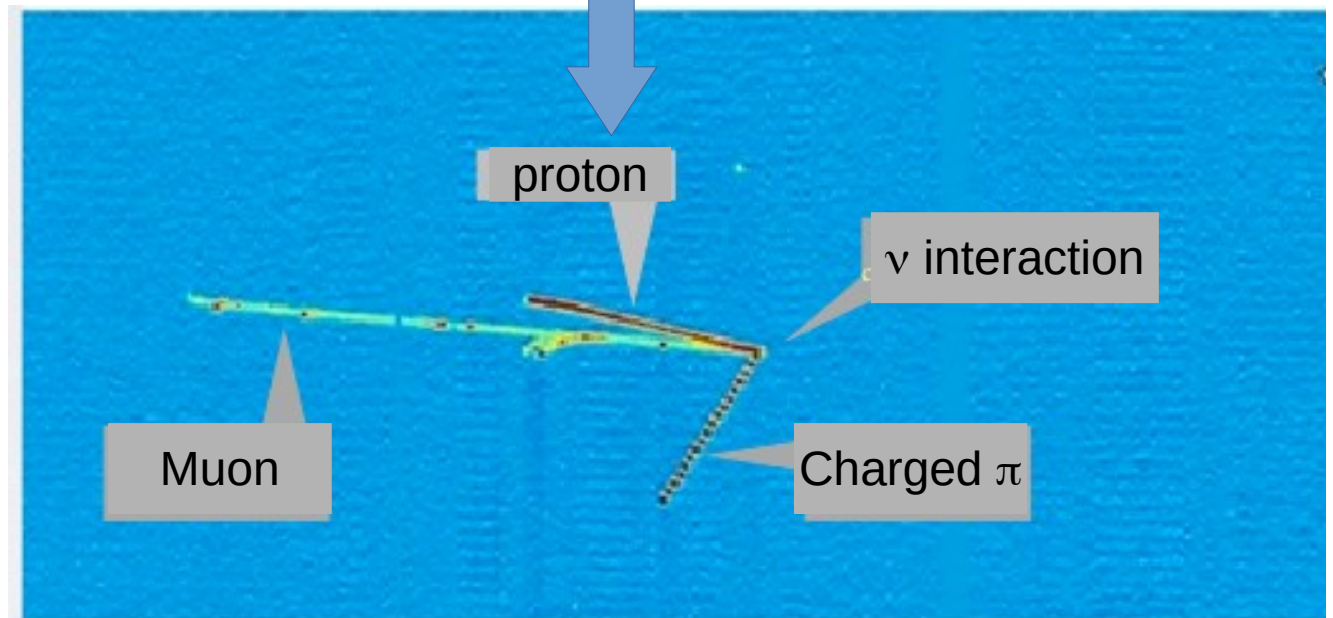
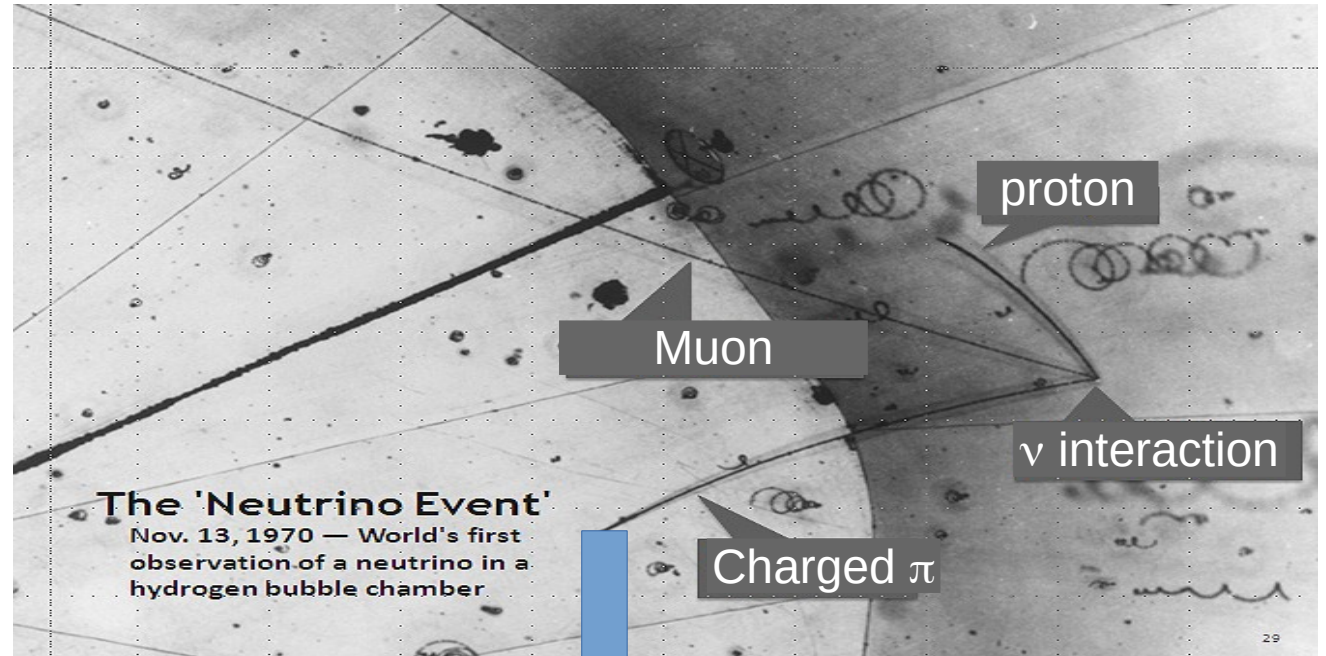


- Four modules (17kTons of argon each)
- Membrane Cryostats hold the detectors.
- Different module designs:
 - Horizontal Drift (HD)
 - Vertical Drift (VD)
 - Module of Opportunity - TBD



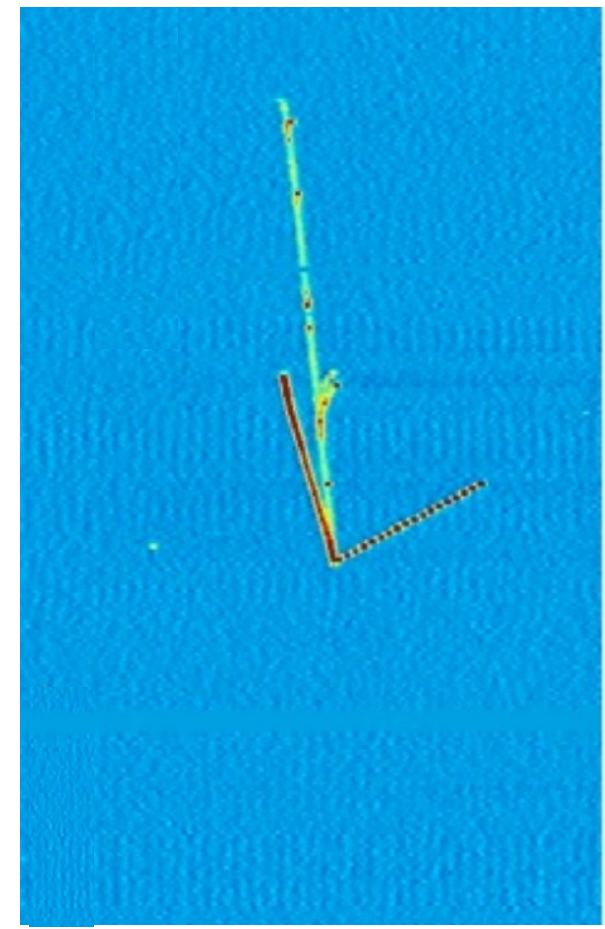
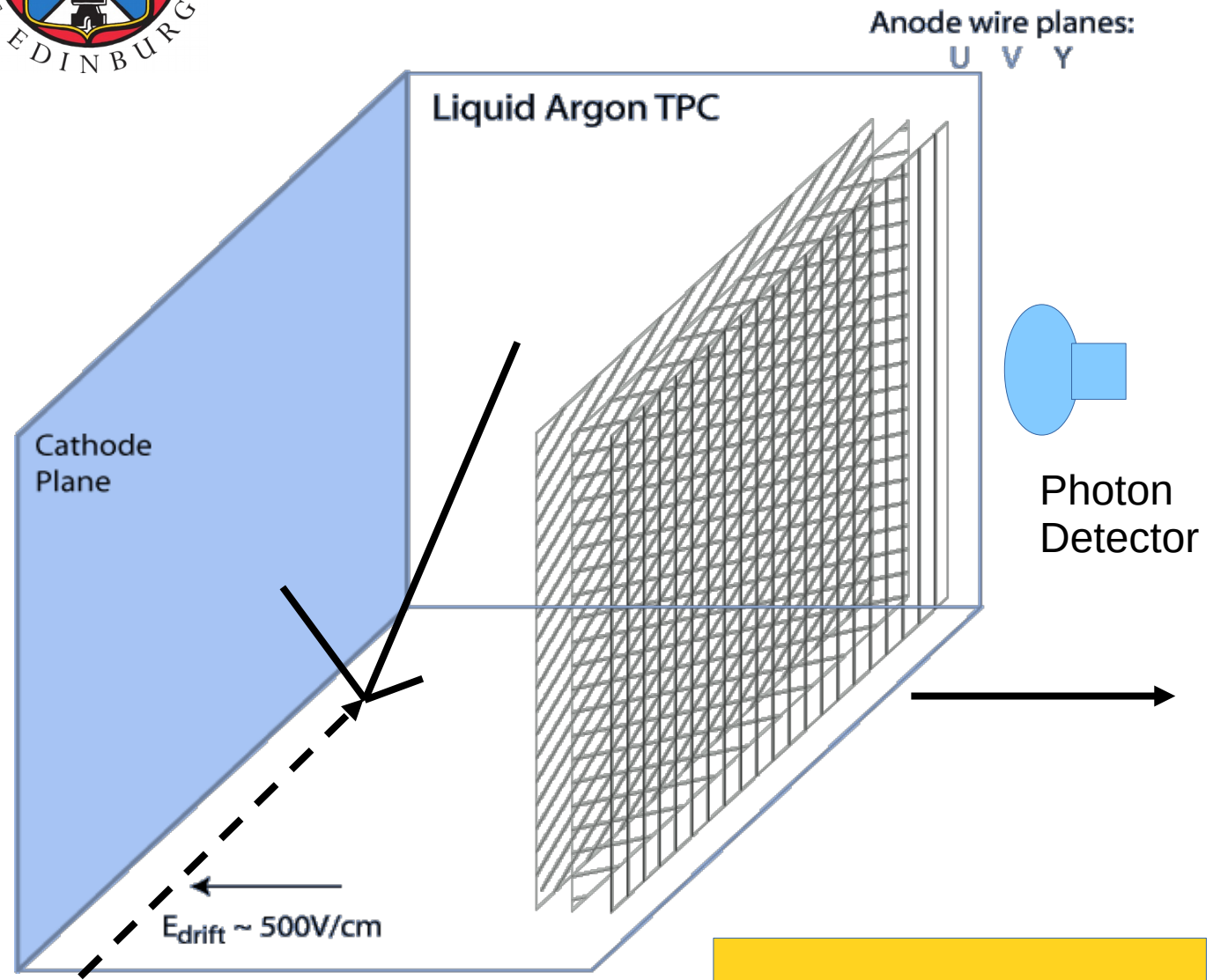
Why Liquid Argon?

Bubble chamber
quality of data with
added full
calorimetry.





LArTPCs



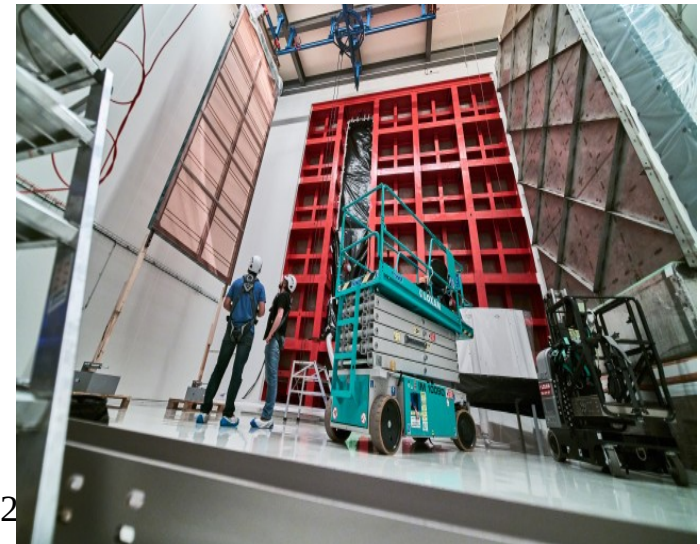
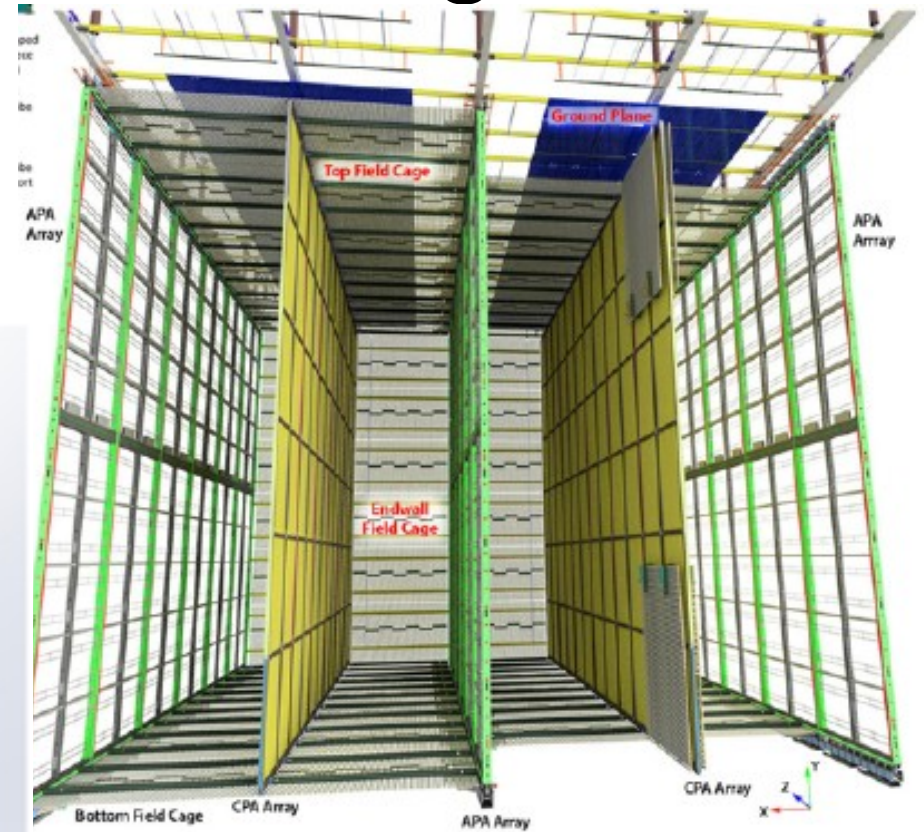
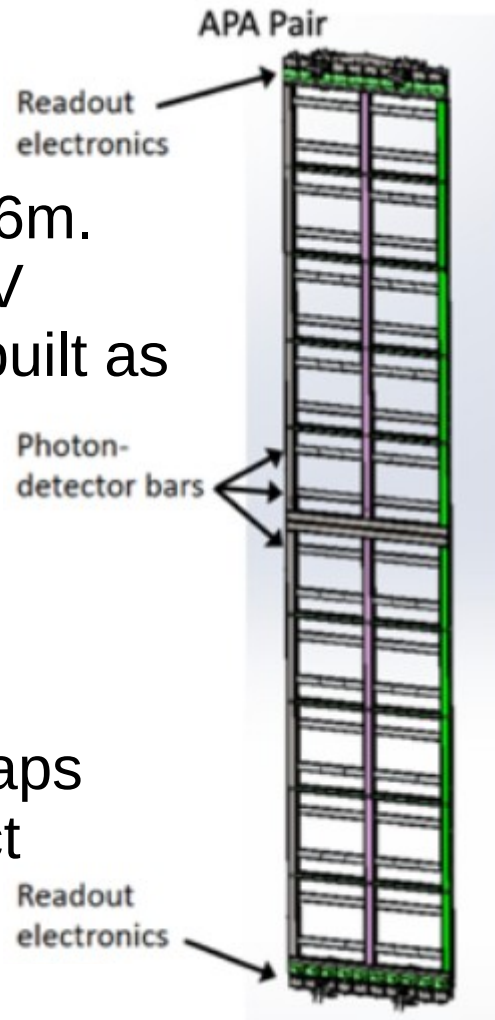
Can drift electrons for very large distances.





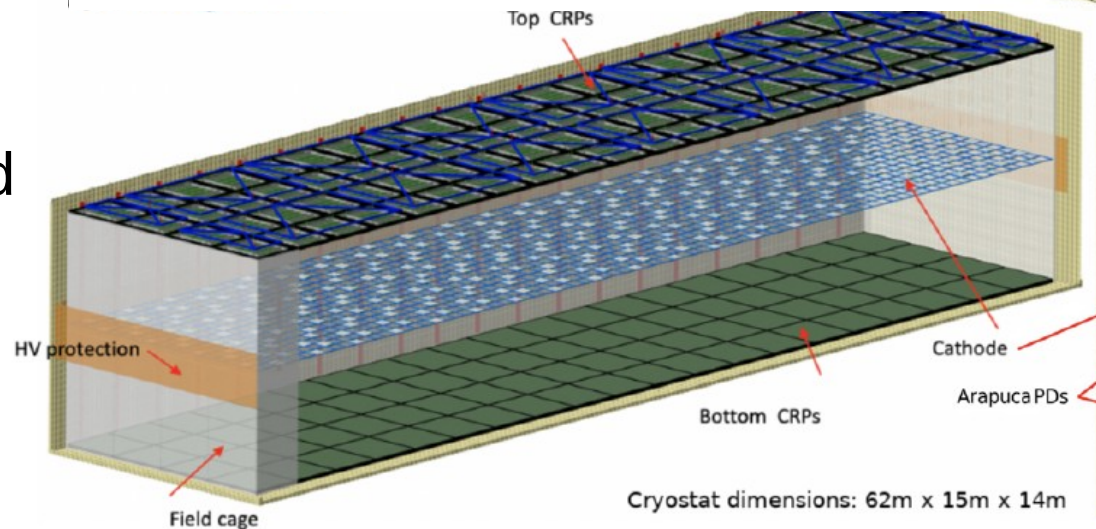
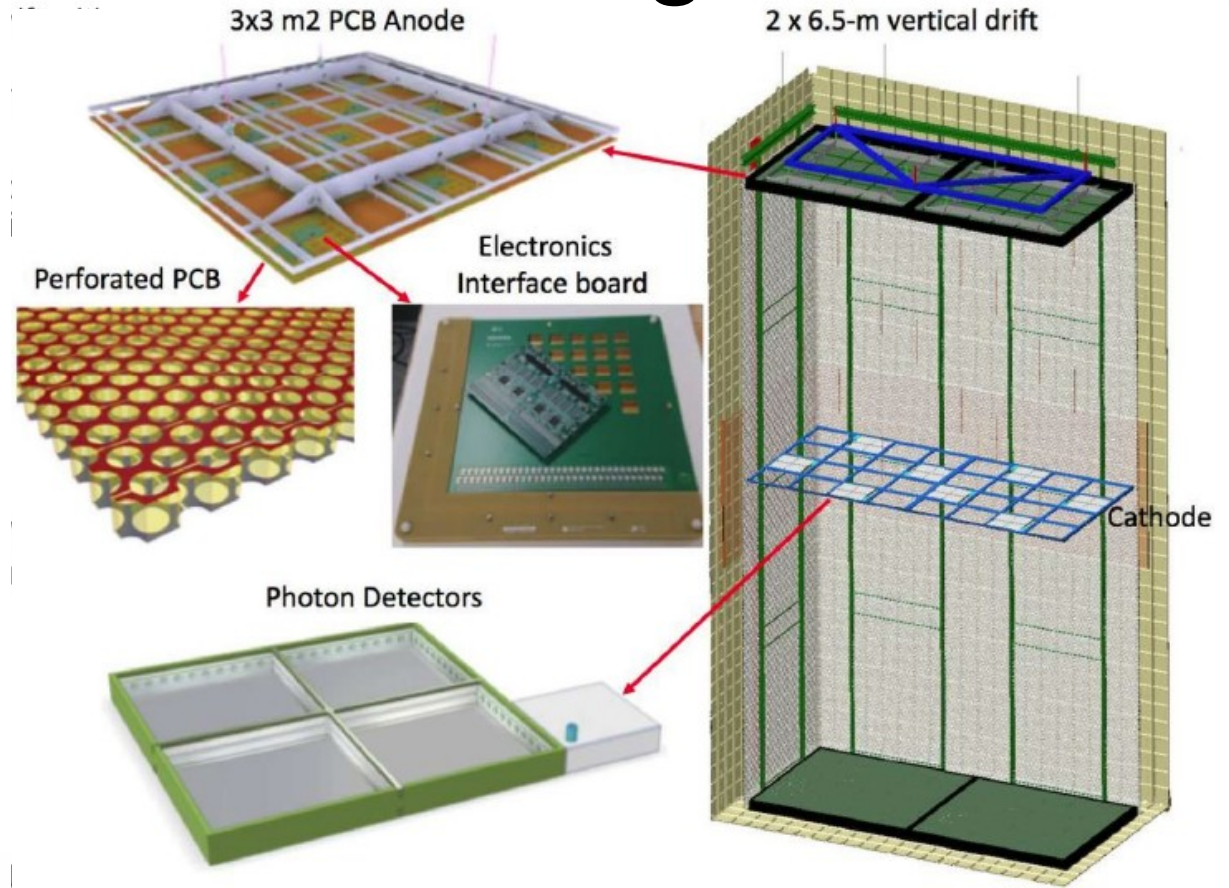
Horizontal Drift design

- First module built will be Horizontal Drift
- Alternating Anode and Cathode Plane assemblies – drift 3.6m.
- Requires HV=-180kV
- Anode = 150 APAs built as wire planes.
- Cathode highly resistive to prevent fast discharges.
- X-ARAPUCA light traps inside APAs to detect scintillation light



Vertical Drift design

- Second module built will be Vertical Drift.
- Evolution of the double phase concept.
- Charge drifted vertically (6.5m), cathode in the middle.
- Requires HV= -300kV
- Anode built out of strips etched on PCBs.
- Large X-ARAPUCA detectors on cathode and walls.



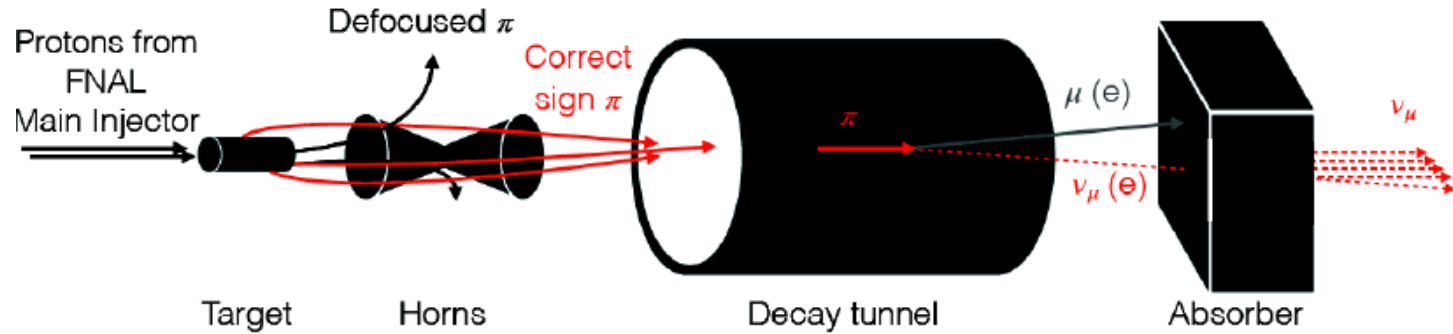


Excavation is ongoing!





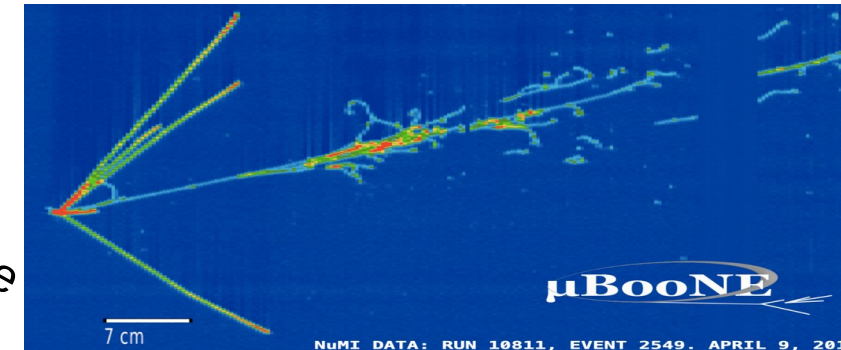
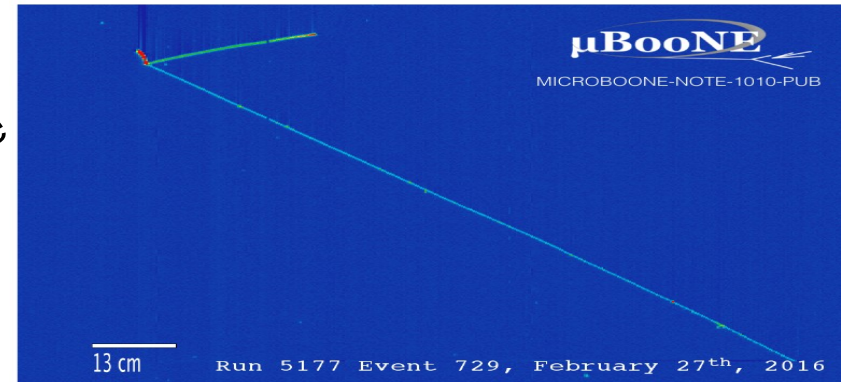
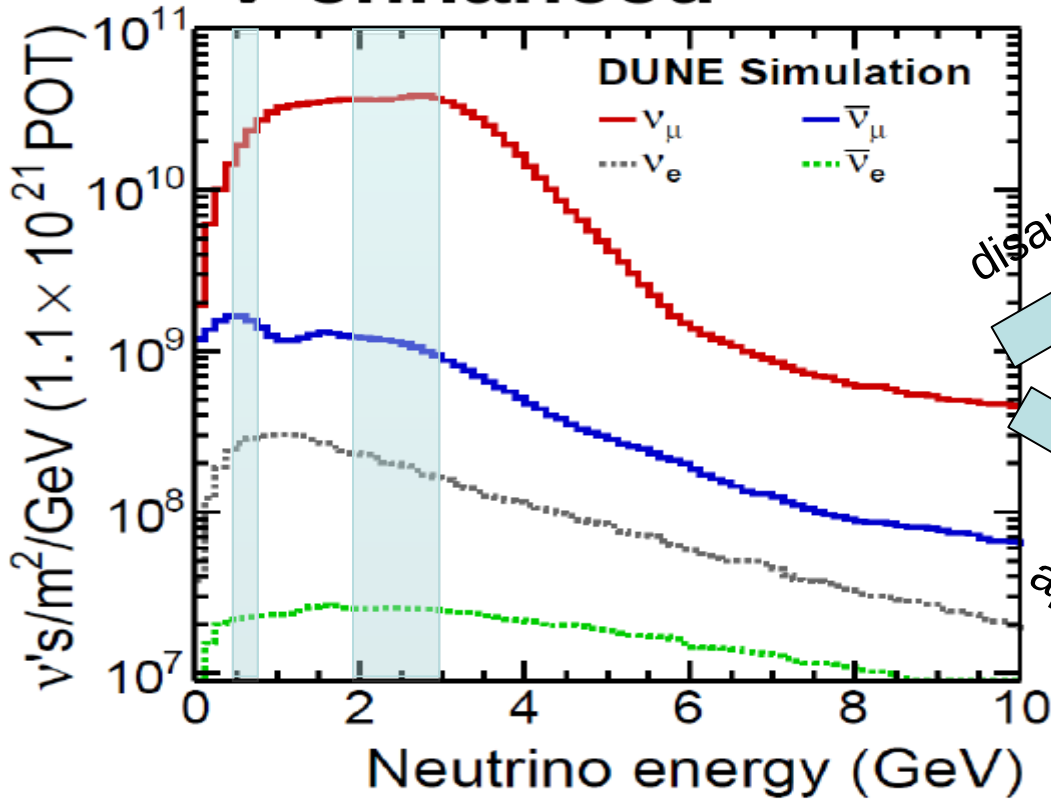
ν_μ beam



$$\pi^\pm \rightarrow \mu^\pm + (\bar{\nu})_\mu$$

Oscillation maxima

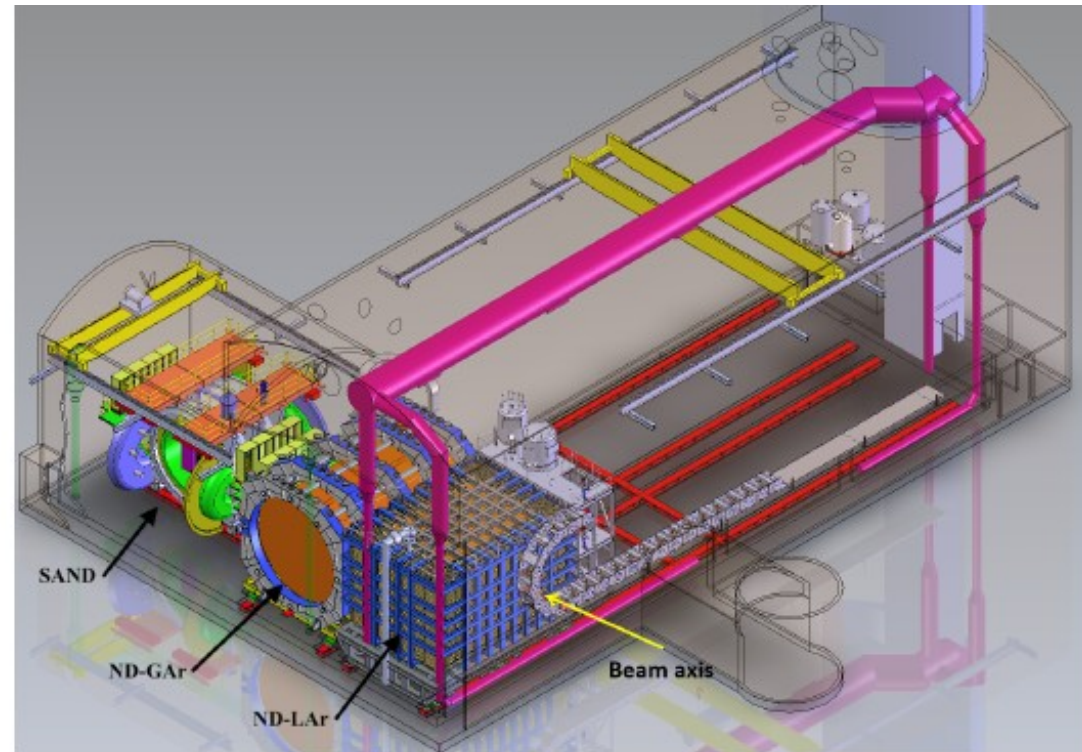
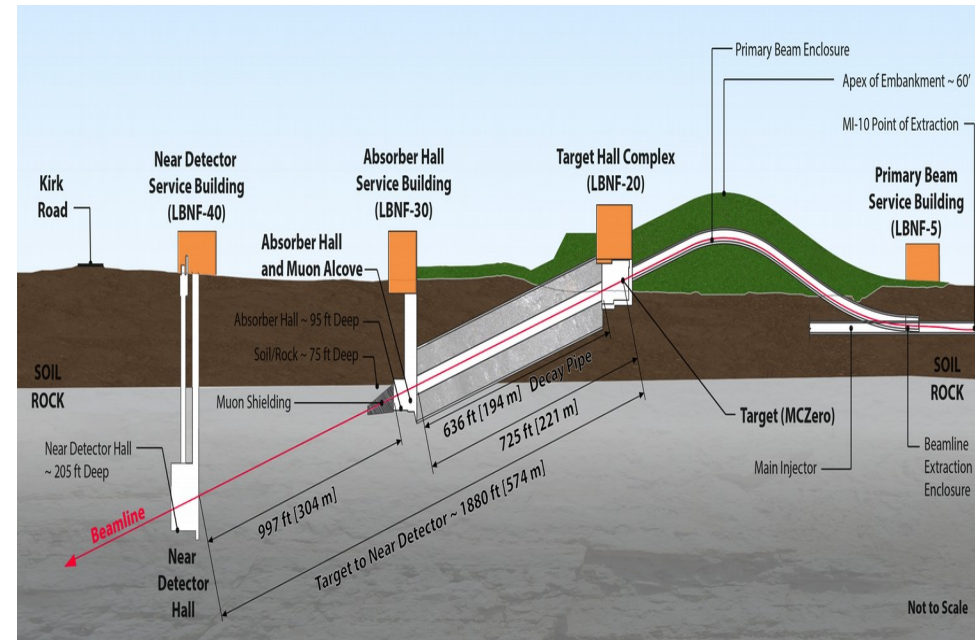
ν -enhanced





Near Detector

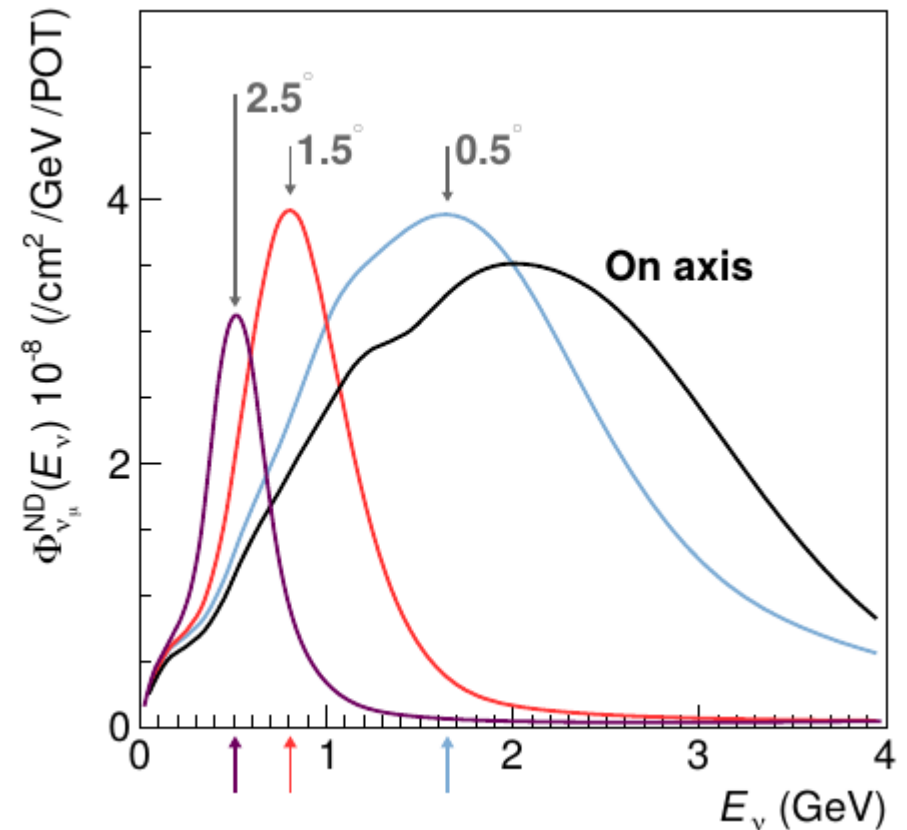
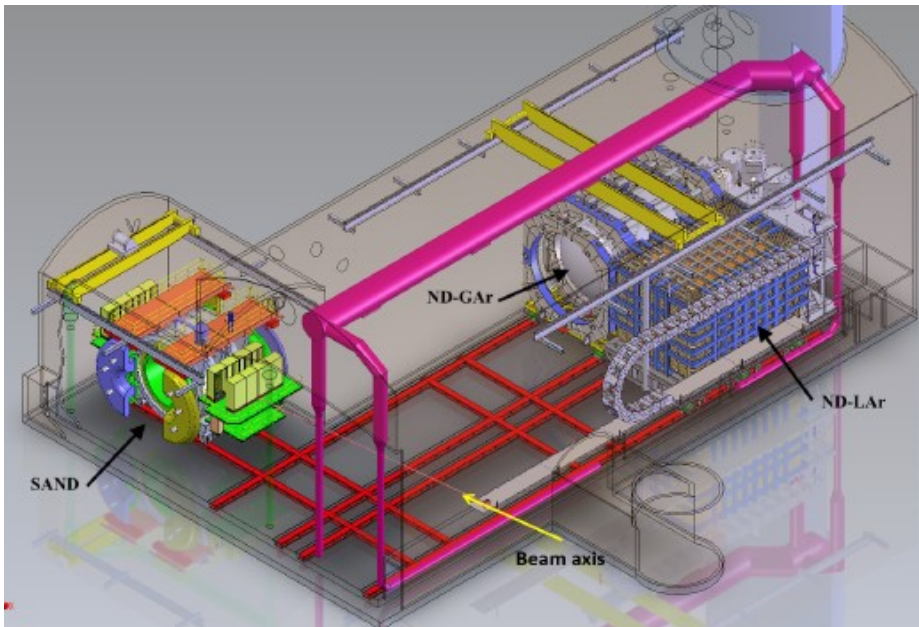
- Multi-component near detector will ensure excellent control of systematics.
- **ND-LAr**: Modular LAr TPC - pixel readout.
- **ND-GAr**: HP GAr TPC + magnet.
- **SAND**: tracker detector, surrounded by an ECAL and magnet: on-axis beam monitor.





DUNE - Prism

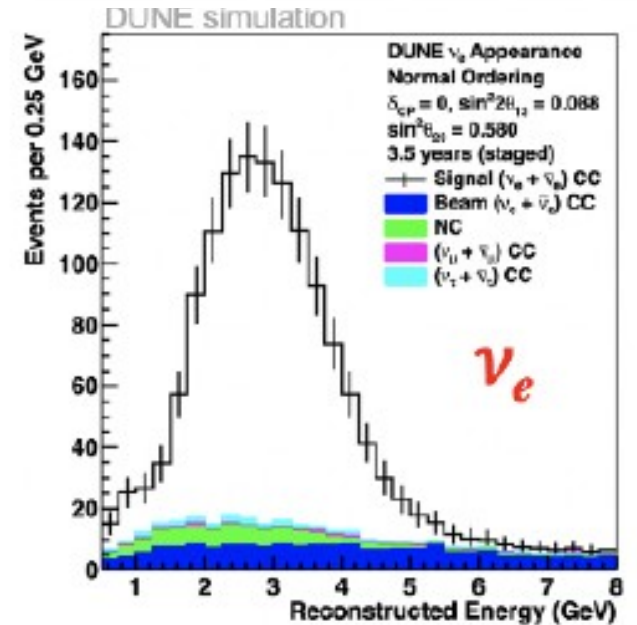
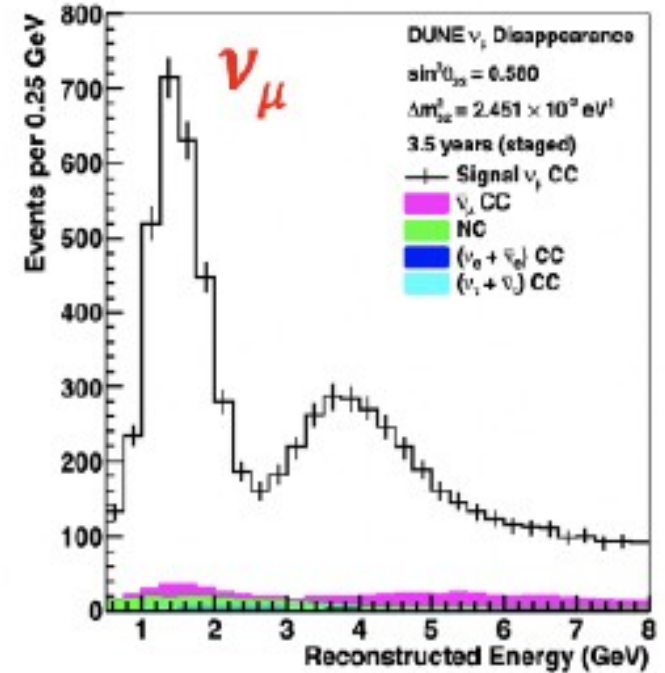
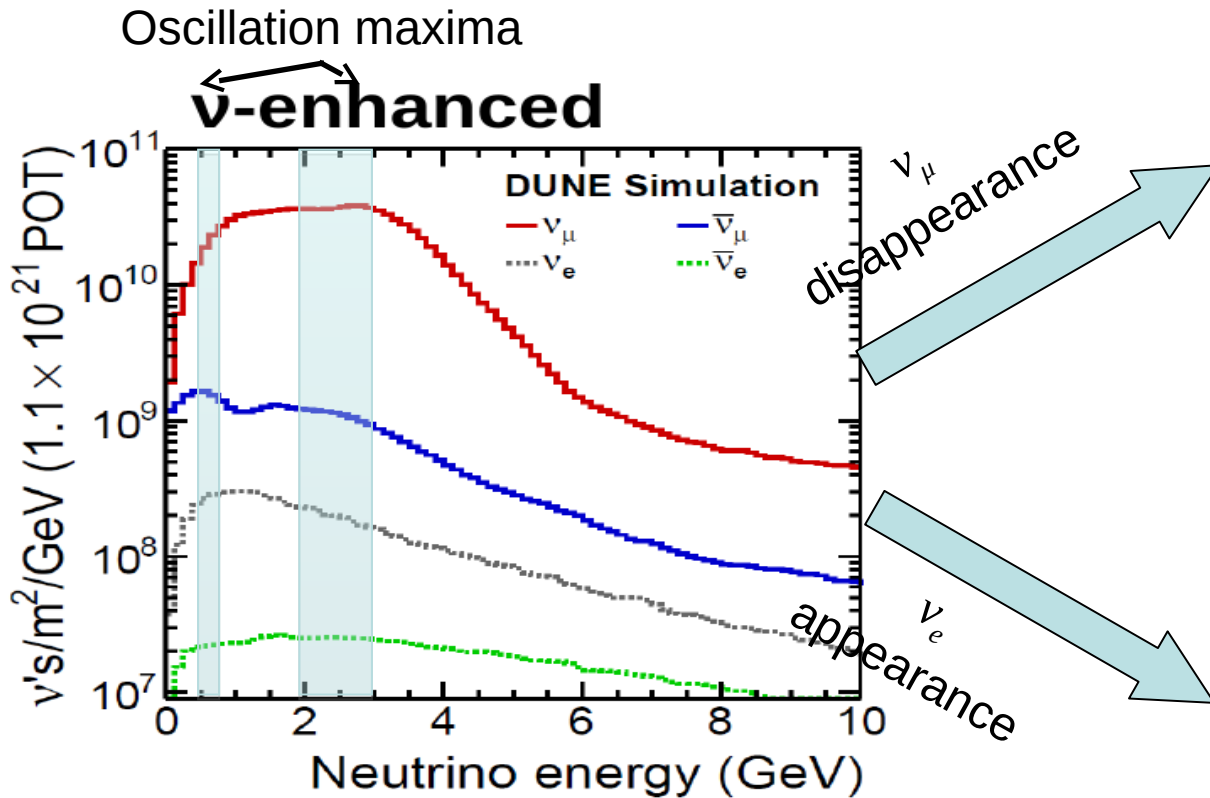
DUNE-PRISM: **ND-LAr** and **ND-GAr** can move sideways to constrain flux and interaction uncertainties.





Physics Sensitivities

~10k events in 7 years
(surviving)

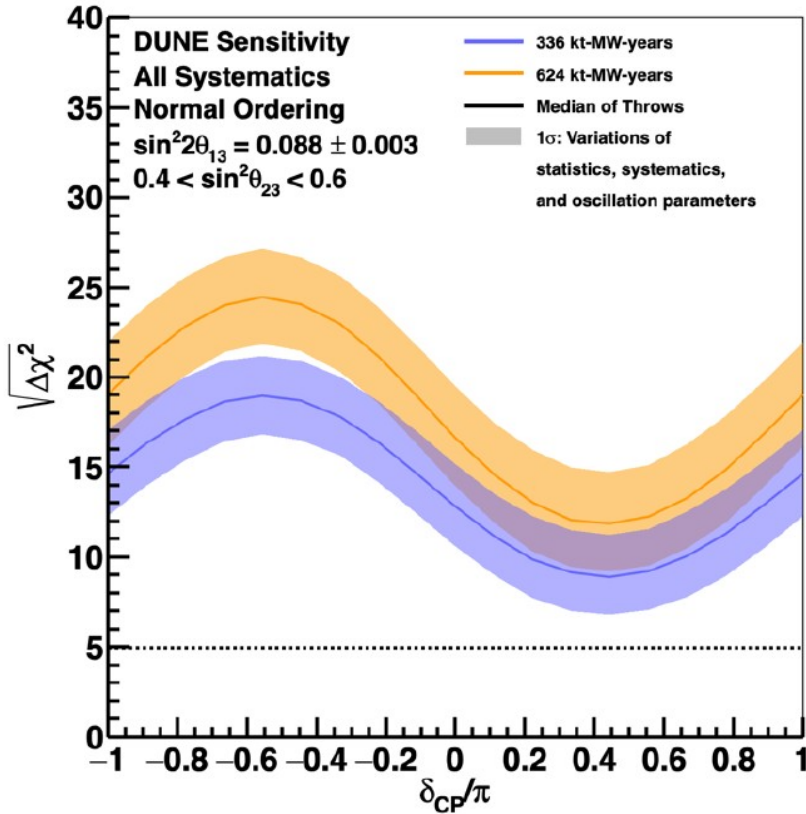


~1k events in 7 years
(appearing)

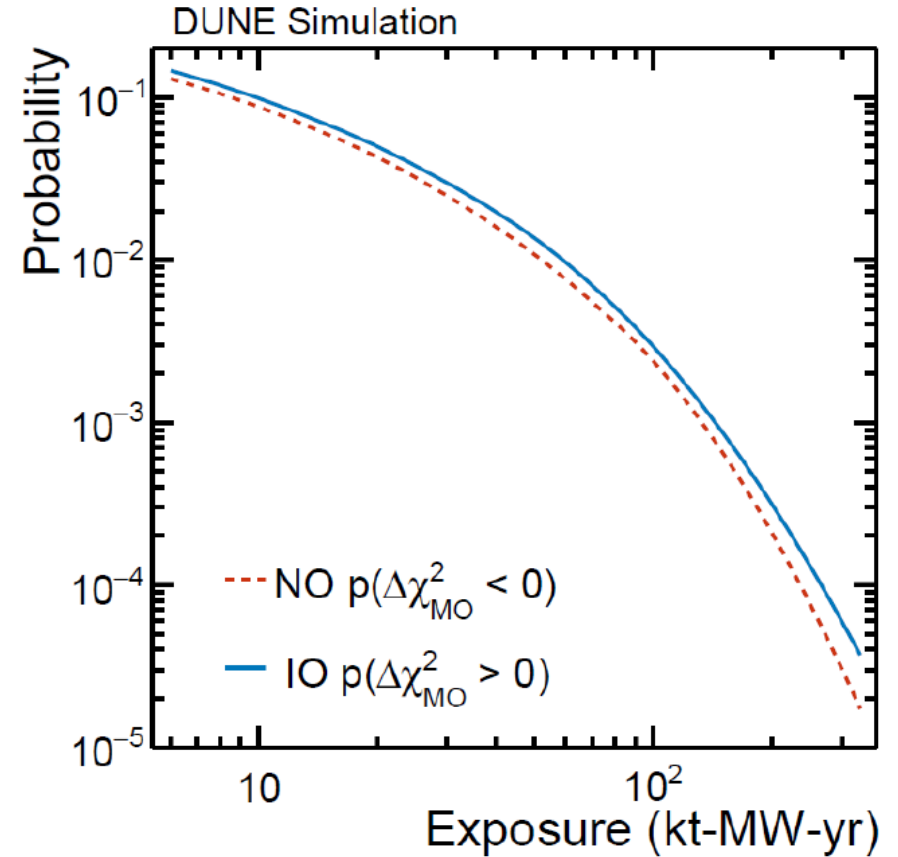




Mass Ordering



EPJC 80 (2020) 978



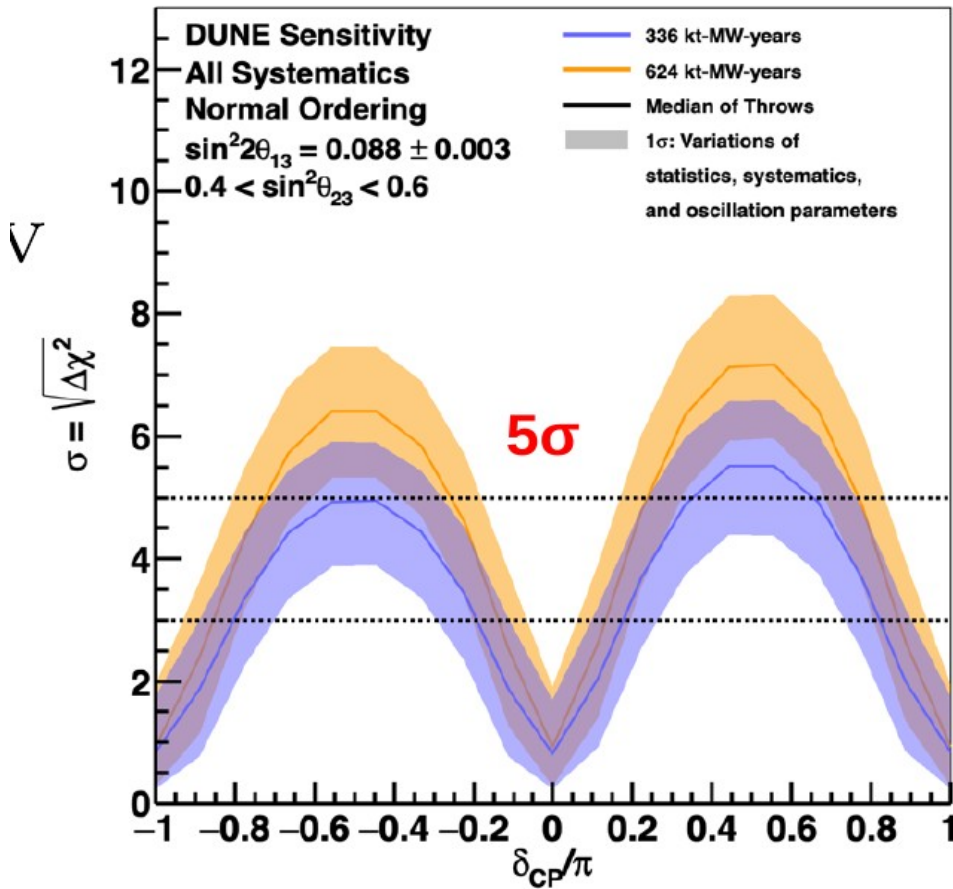
arXiv:2109.01304

- DUNE will determine the MO soon in its run.
- By exposure of 66 kt-MW-yr, probability to extract wrong ordering < 0.01 .

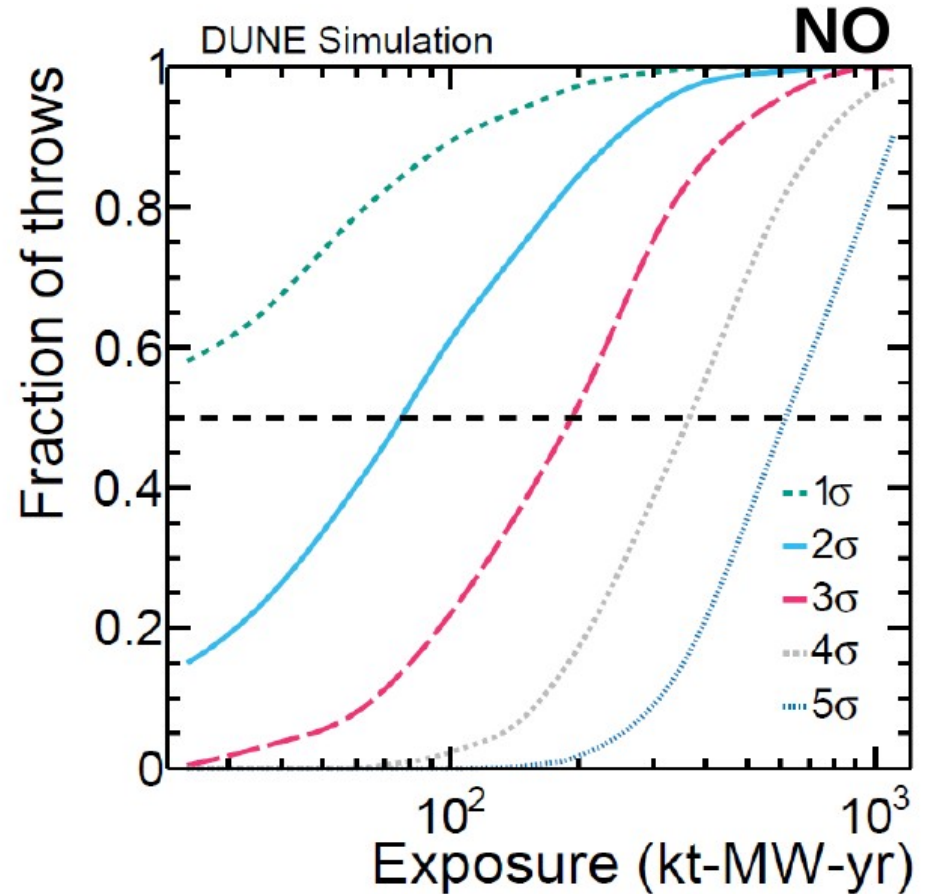




CP Violation



EPJ C 80 (2020) 978



arXiv:2109.01304

- Sensitivity to over 50% of δ_{CP} values at 5σ , with full exposure.
- Median sensitivity for 50% δ_{CP} above 3σ after 197 kt-MW-yr

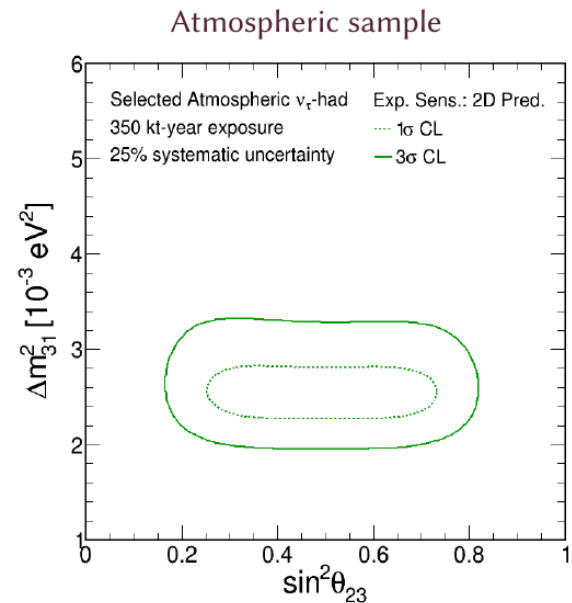
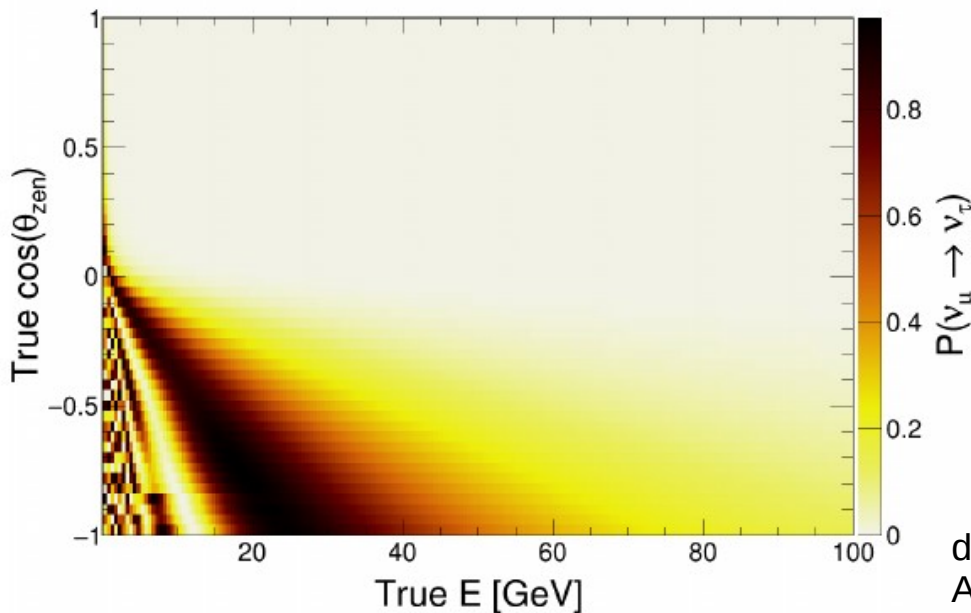




Atmospheric Neutrinos

- DUNE will detect atmospheric neutrinos:

Sample	Event rate per year
fully contained electron-like	1600
fully contained muon-like	2400
partly contained muon-like	790



ν_τ appearance is possible.
 Expect 1 ν_τ /kton-year
 Possible tests of non-unitarity.

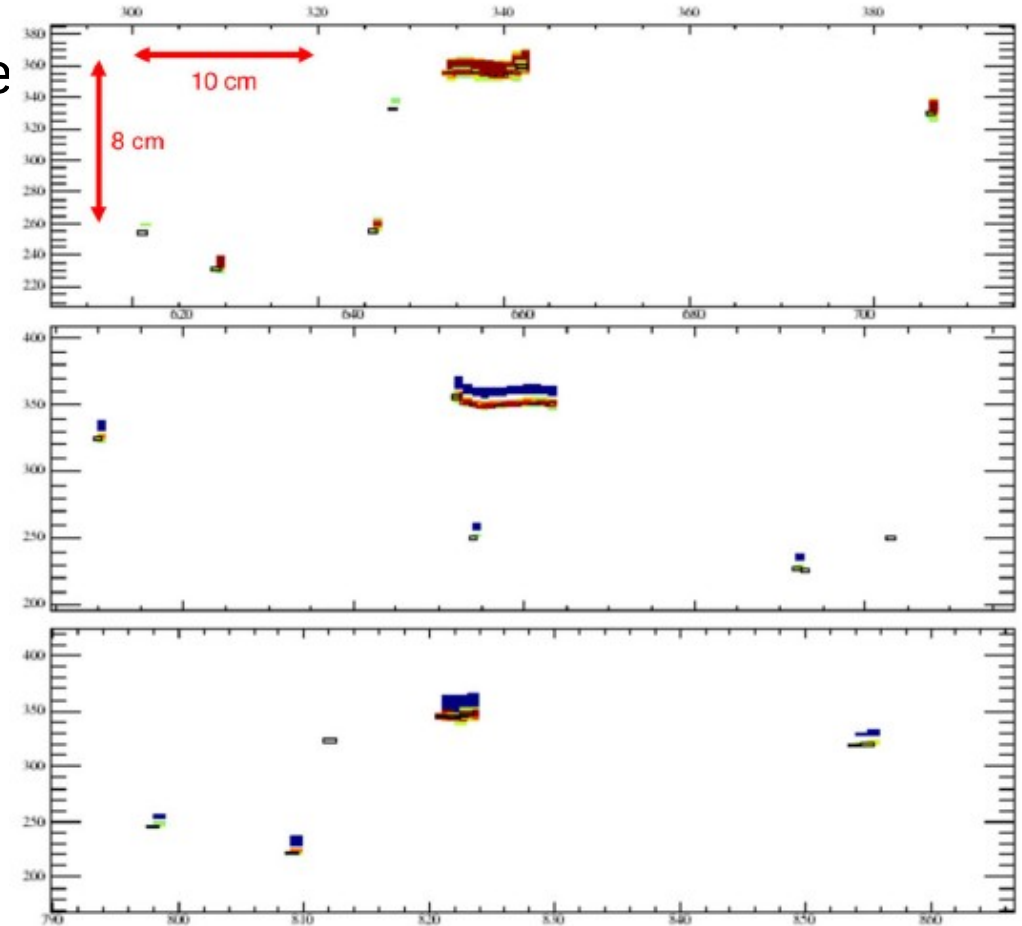
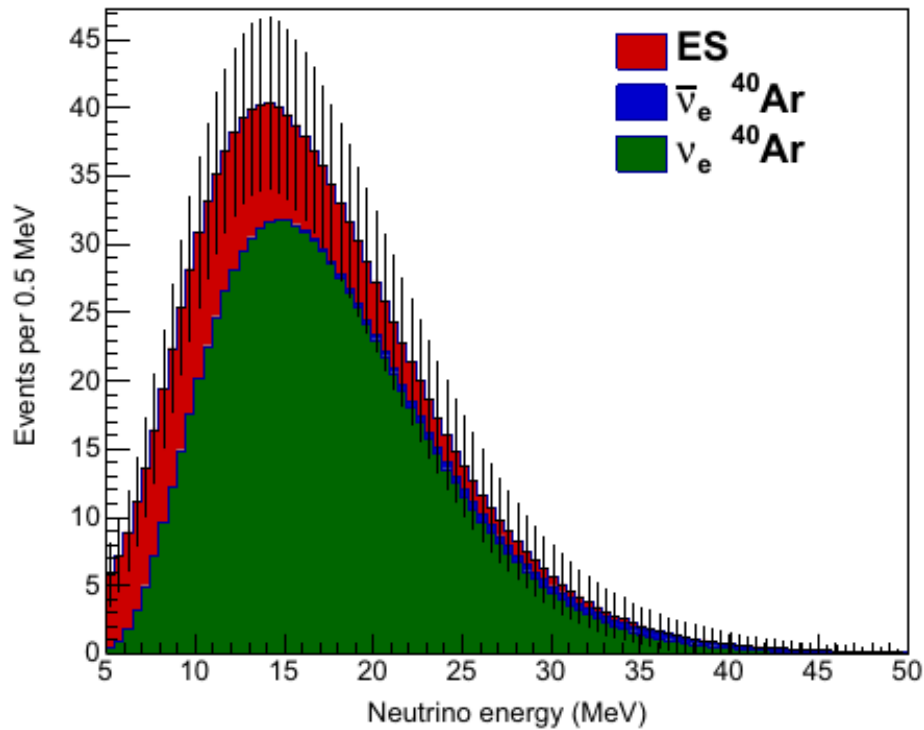
de Gouvea, Kelly, Stenico, Pasquini, PRD 100, 016004 (2019)
 And. A. Aurisano talk at NuTau 2021





Supernova neutrinos

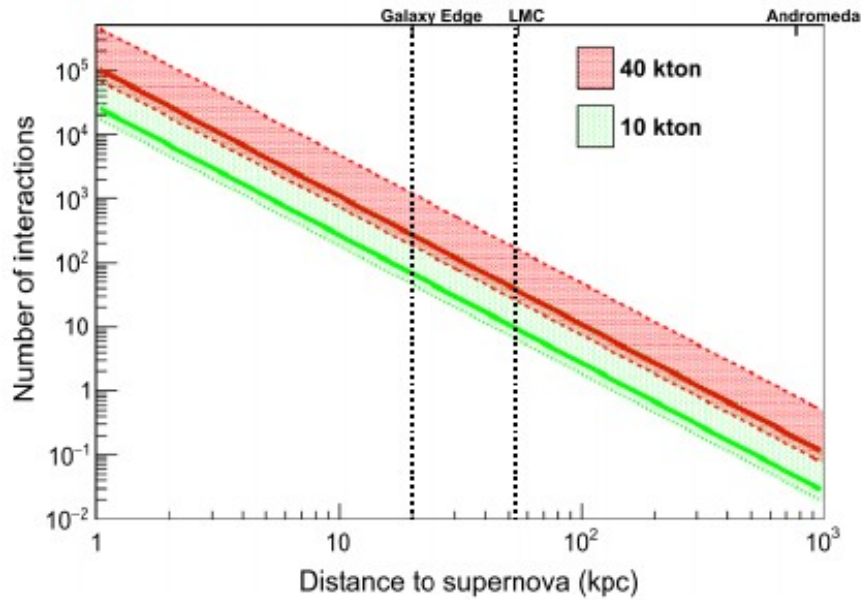
- LAr detectors are mainly sensitive to ν_e via: $\nu_e + {}^{40}\text{Ar} \rightarrow e^- + {}^{40}\text{K}^*$



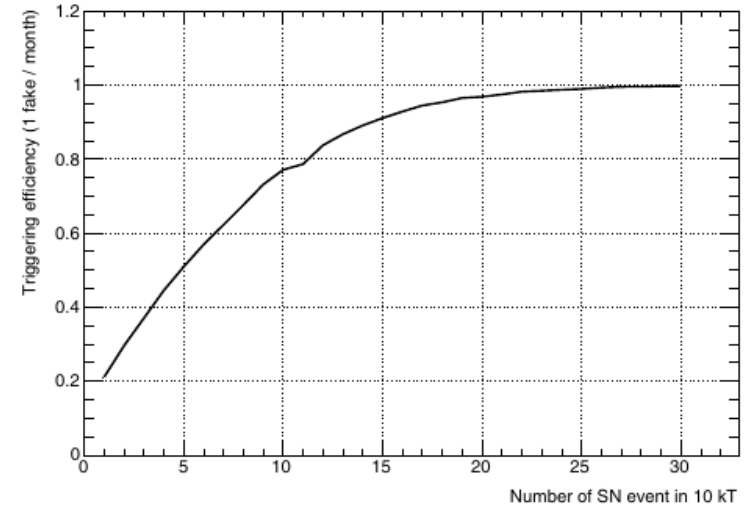
Eur. Phys. J. C 81 (2021) 5, 423]



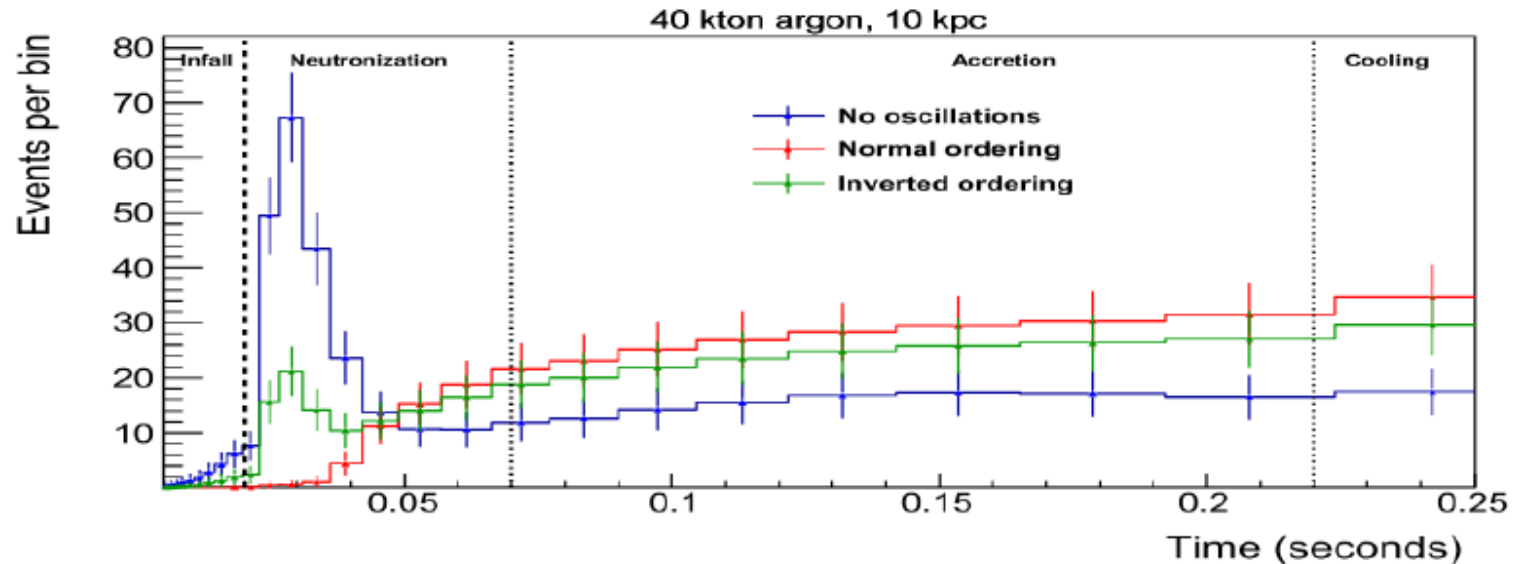
Supernova neutrinos



SN burst trigger efficiency

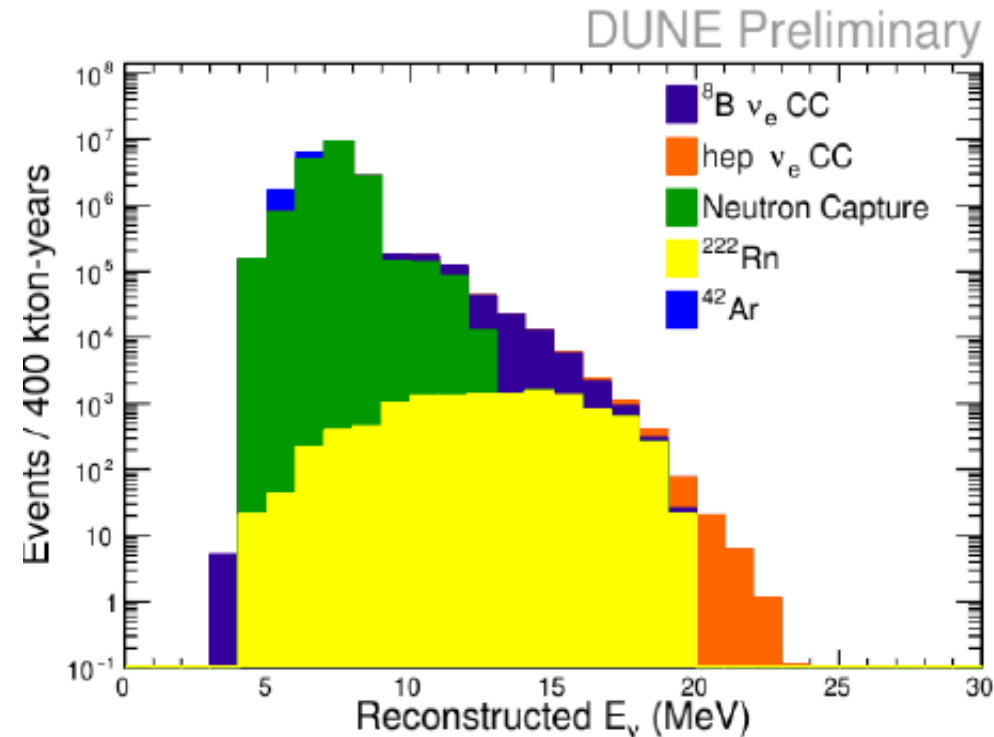
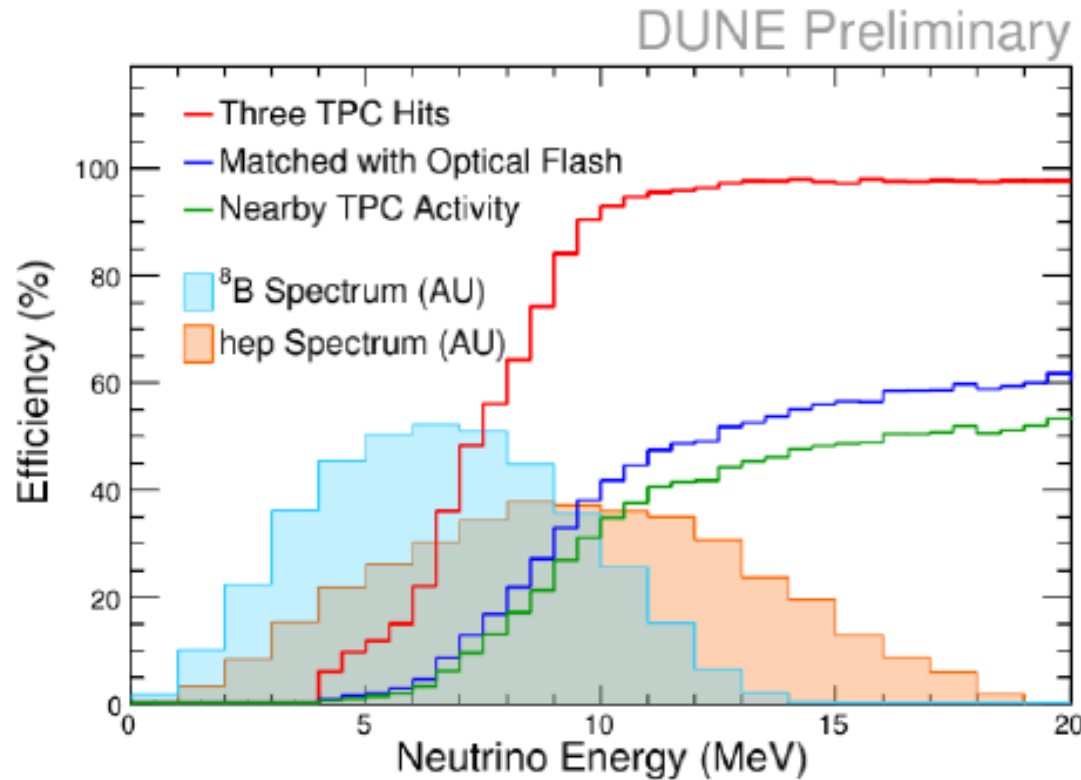


- Sensitivity to neutronization burst
- Sensitivity to mass hierarchy





Solar Neutrinos



- DUNE should be sensitive to solar neutrinos, with a potential to observe hep neutrinos.
- Main backgrounds: neutrons and Rn-induced alpha-gamma interactions.

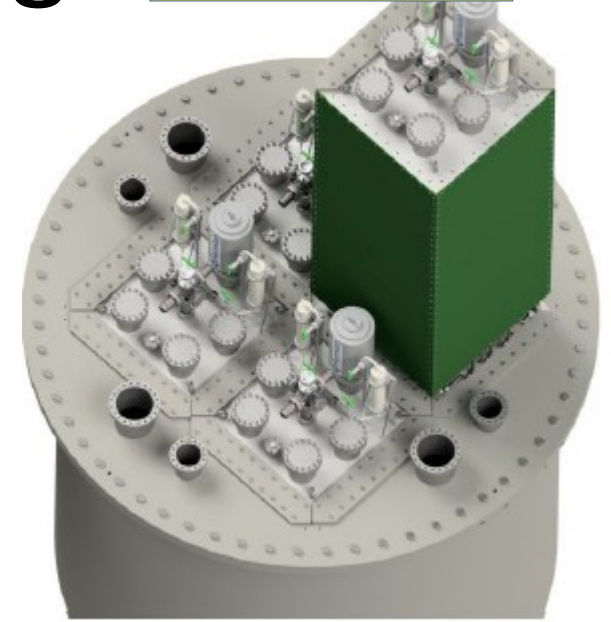
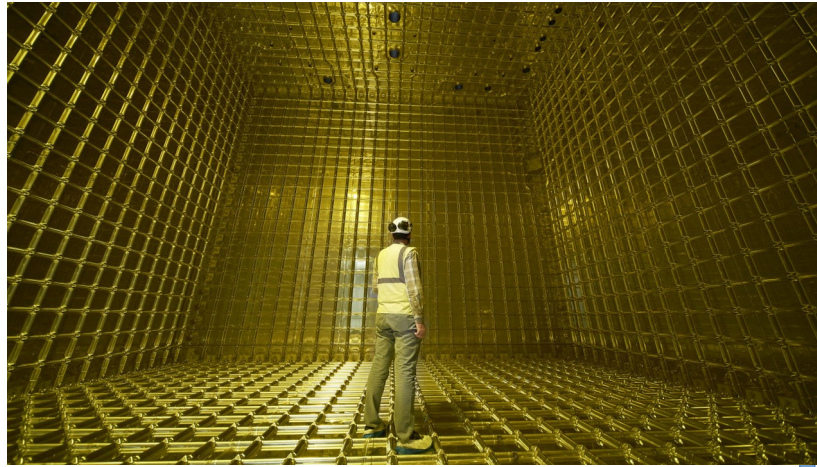




DUNE Prototypes

Near Detector

Far Detector



ArgonCube 2x2 ND-LAr Demonstrator.

4 independent modules, test
In NuMI beam scheduled in 2022

- ProtoDUNEs at CERN
- 2x1kTon cryostats used to validate FD components at full scale.
- Successful run from 2018-2020.
- Next phase starting in summer 2022.





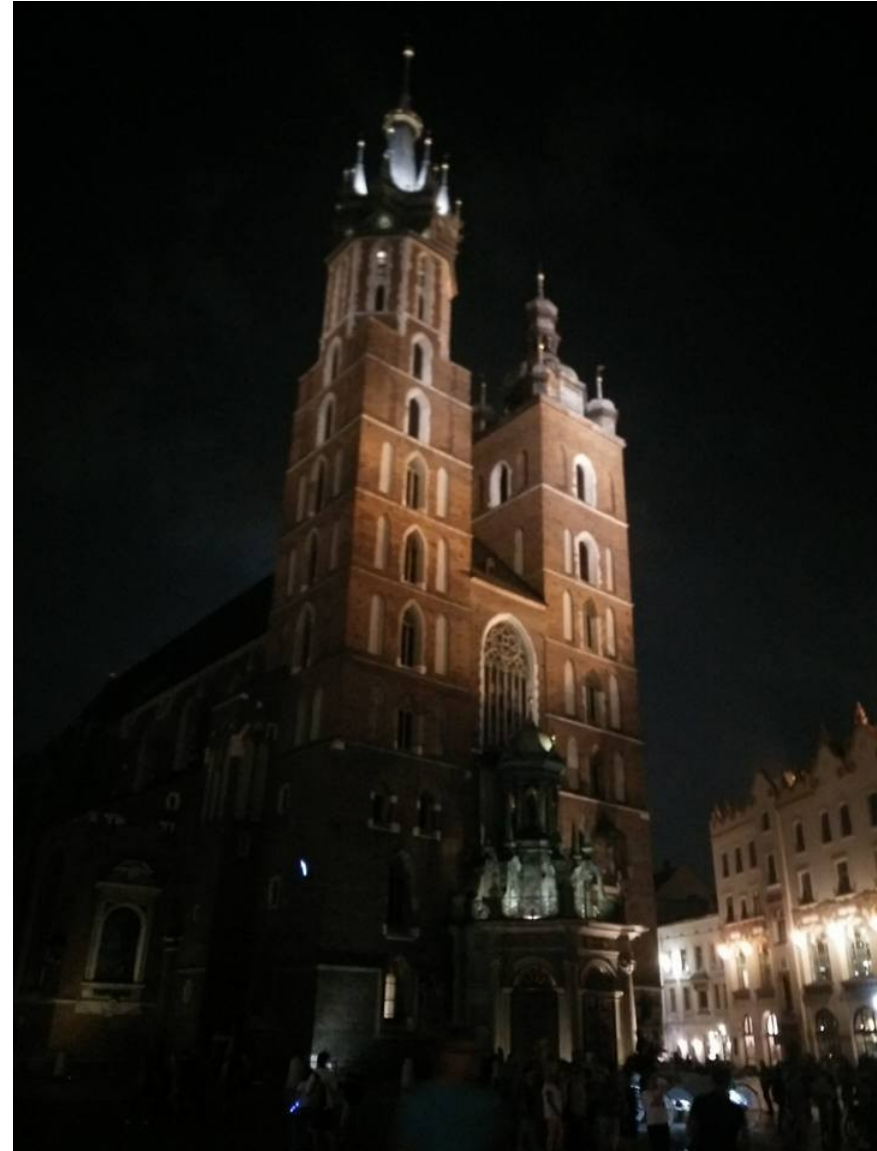
Conclusions

- DUNE is a next-generation long-baseline oscillation experiment set to discover CP violation in the lepton sector and determine the neutrino mass ordering.
- A rich physics programme includes astrophysical and atmospheric neutrinos and BSM physics (proton decay).
- Excavation of the far site is underway, and technologies for first two modules have been selected.
- Extensive prototyping efforts at CERN and Fermilab ongoing, with test runs planned in 2022-2023.
- First DUNE far-detector installation in mid-2020s, with neutrino beam data in late 2020s.





Thank you





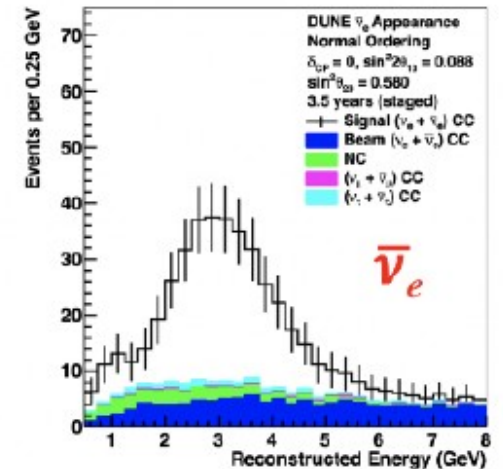
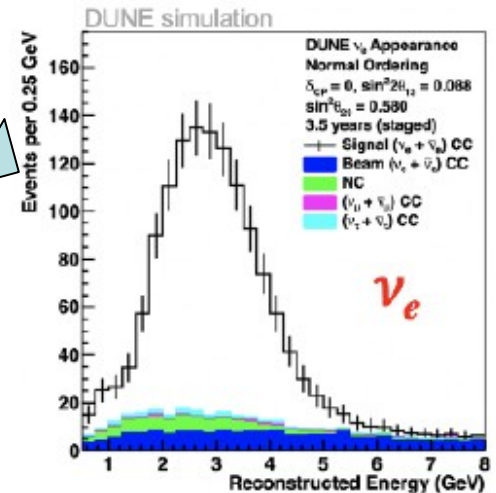
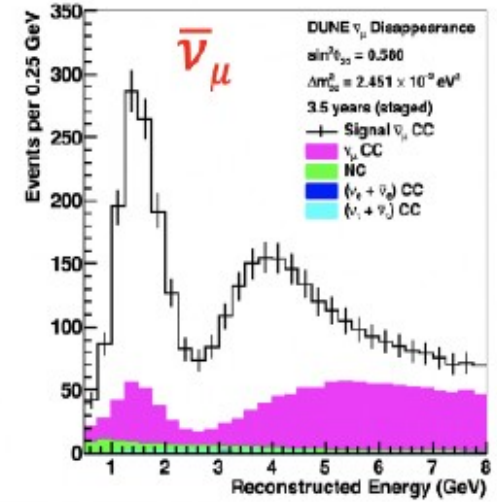
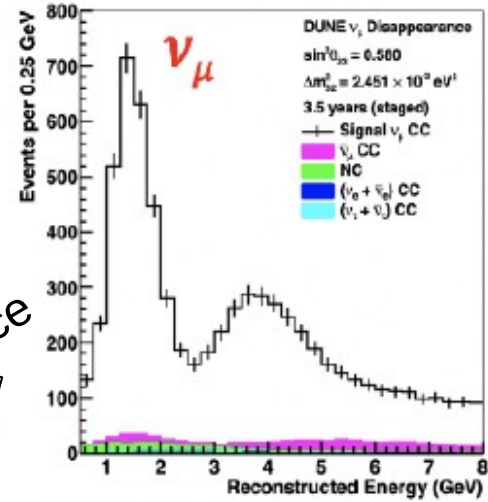
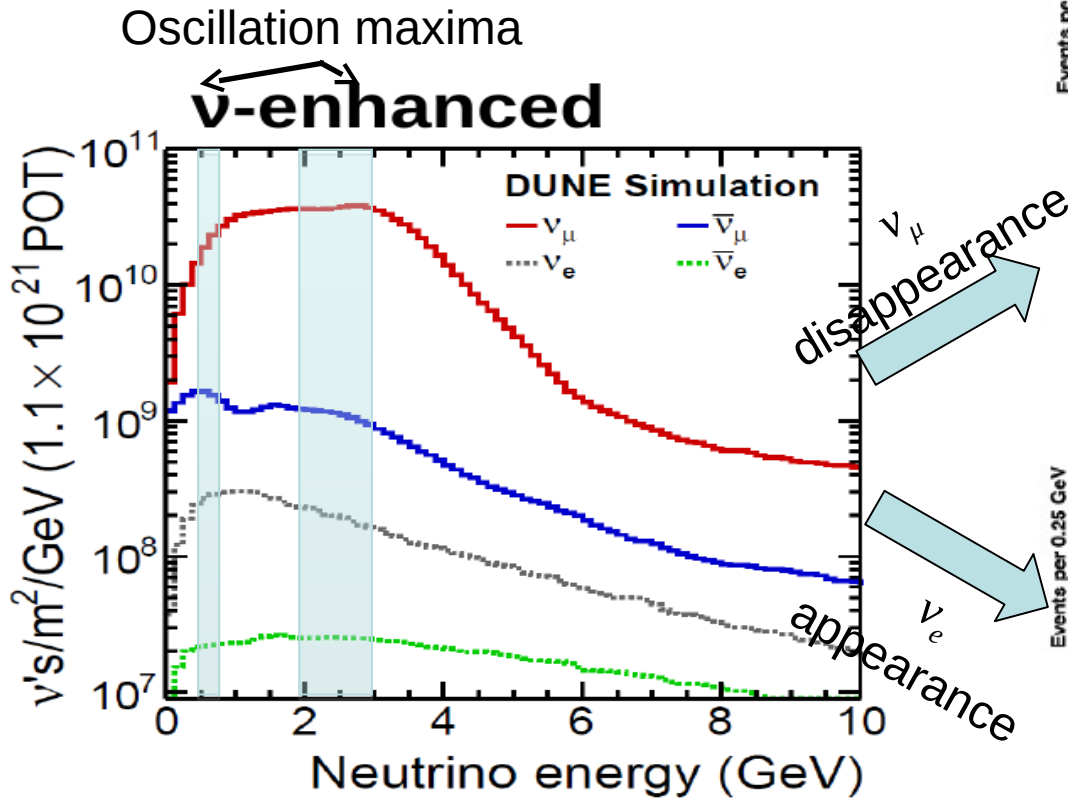
Backup Slides





Physics Sensitivities

~10k events in 7 years

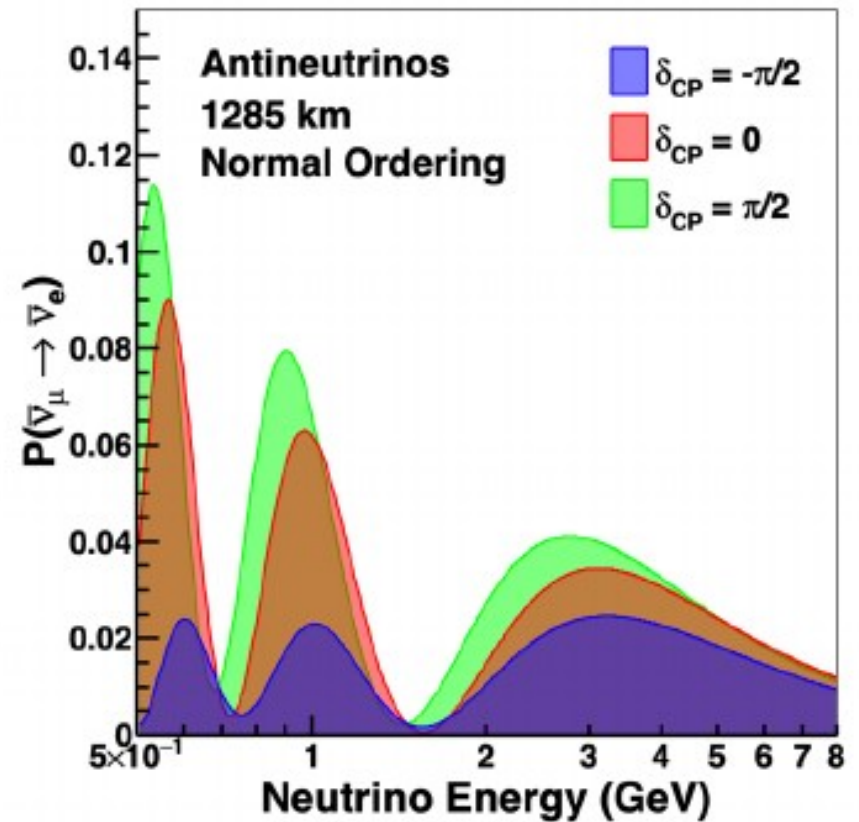
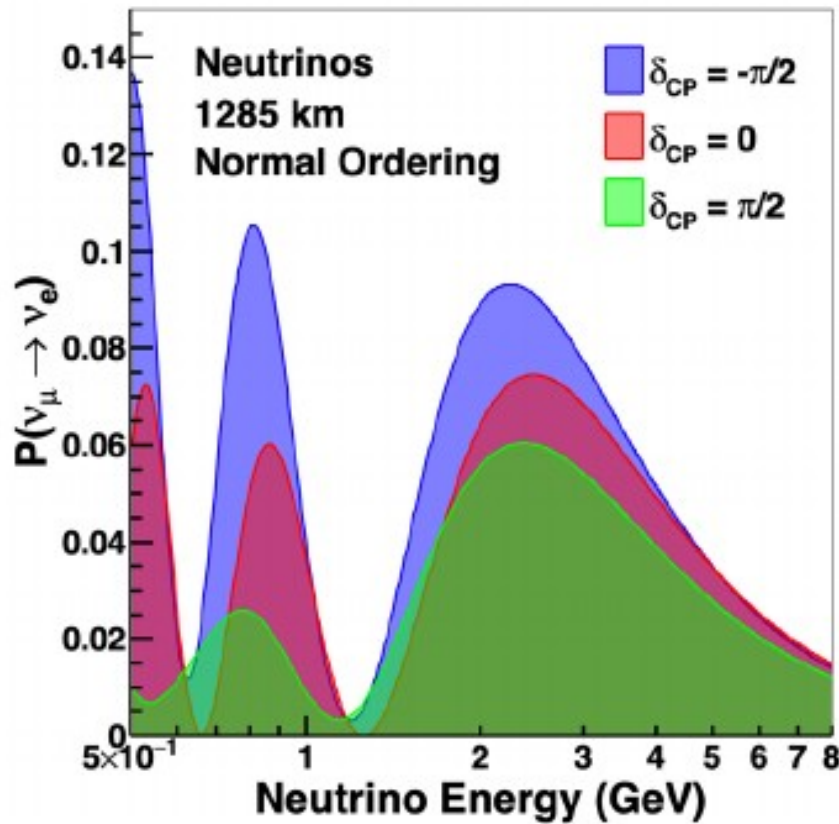


~1k events in 7 years





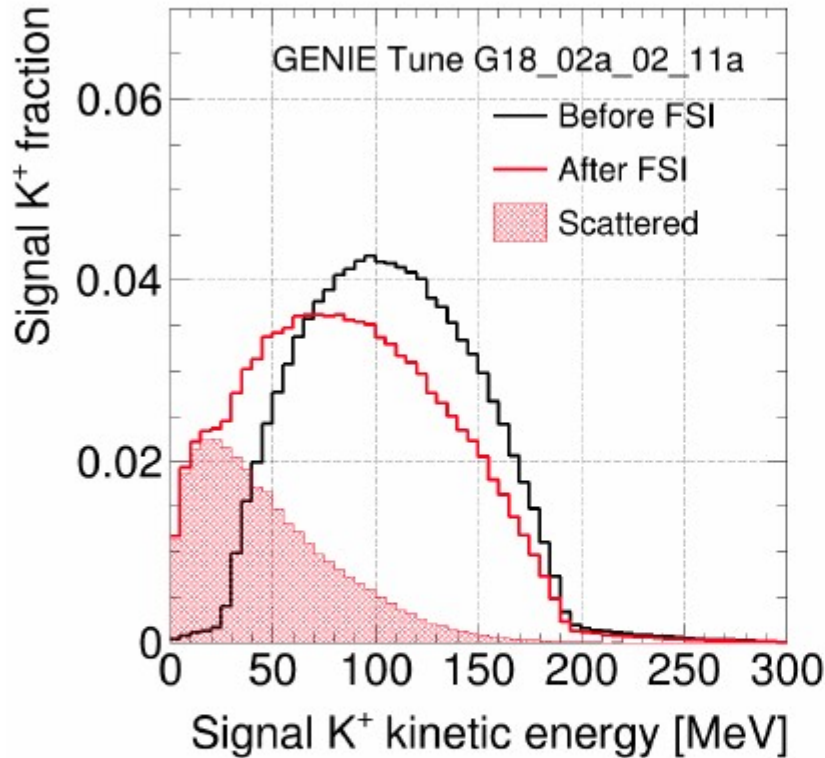
CP phase in ν_e appearance



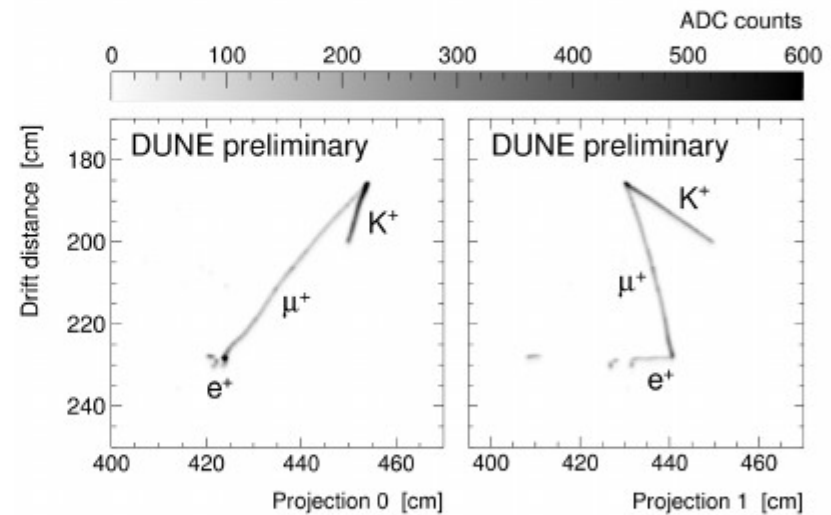


Nucleon Decay

<https://doi.org/10.22323/1.390.0226>



Backgrounds:
Atmospheric neutrinos



DUNE will be competitive for favoured SUSY decay channel:
 $\tau(p \rightarrow \bar{\nu}K) > 1.3 \times 10^{34}$ yrs @ 90 % CL after 400 kt · yrs

