

Dark matter at the GeV scale and searches at low-energy experiments

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Anomalies and Precision in the Belle II Era

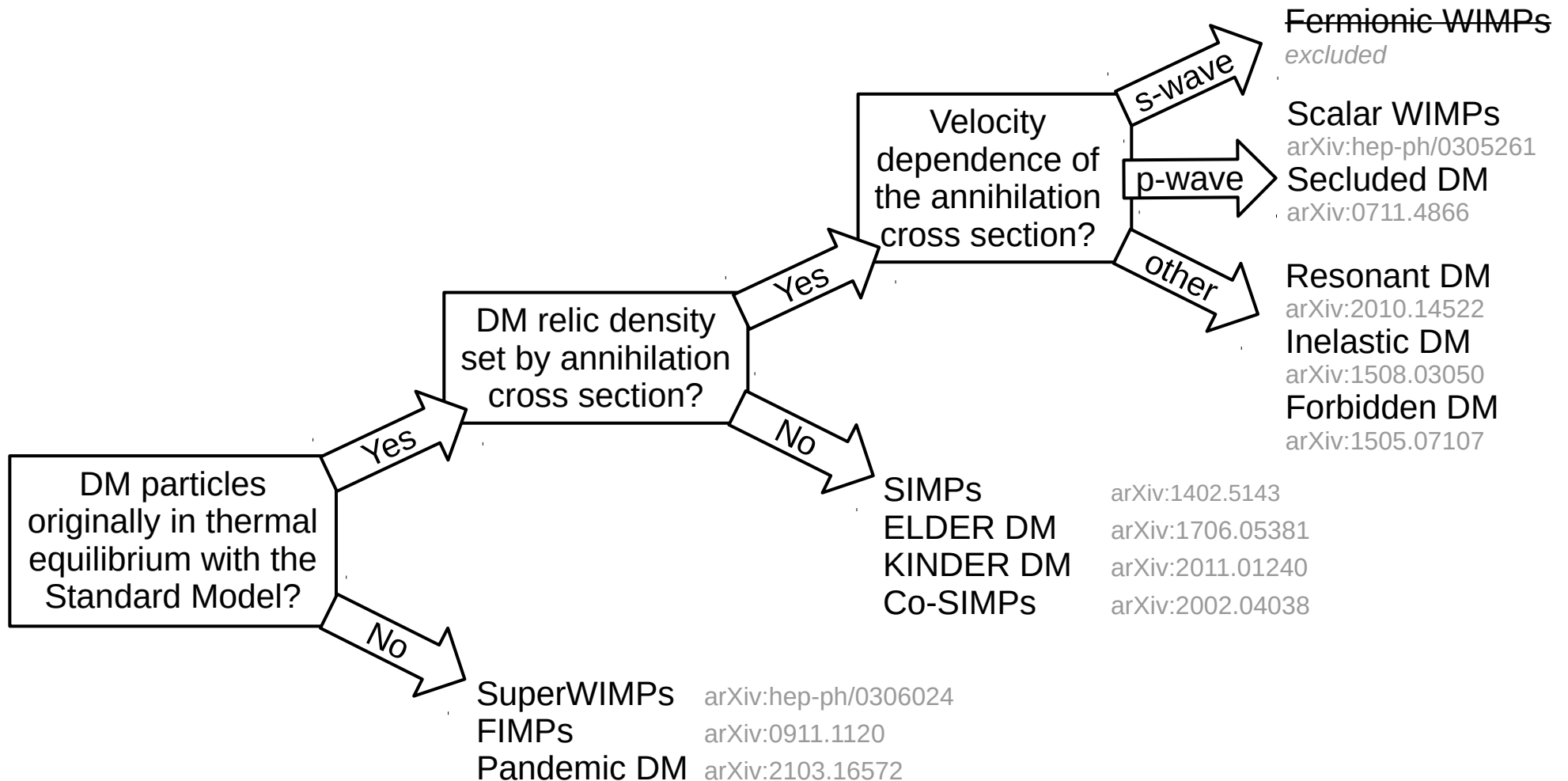
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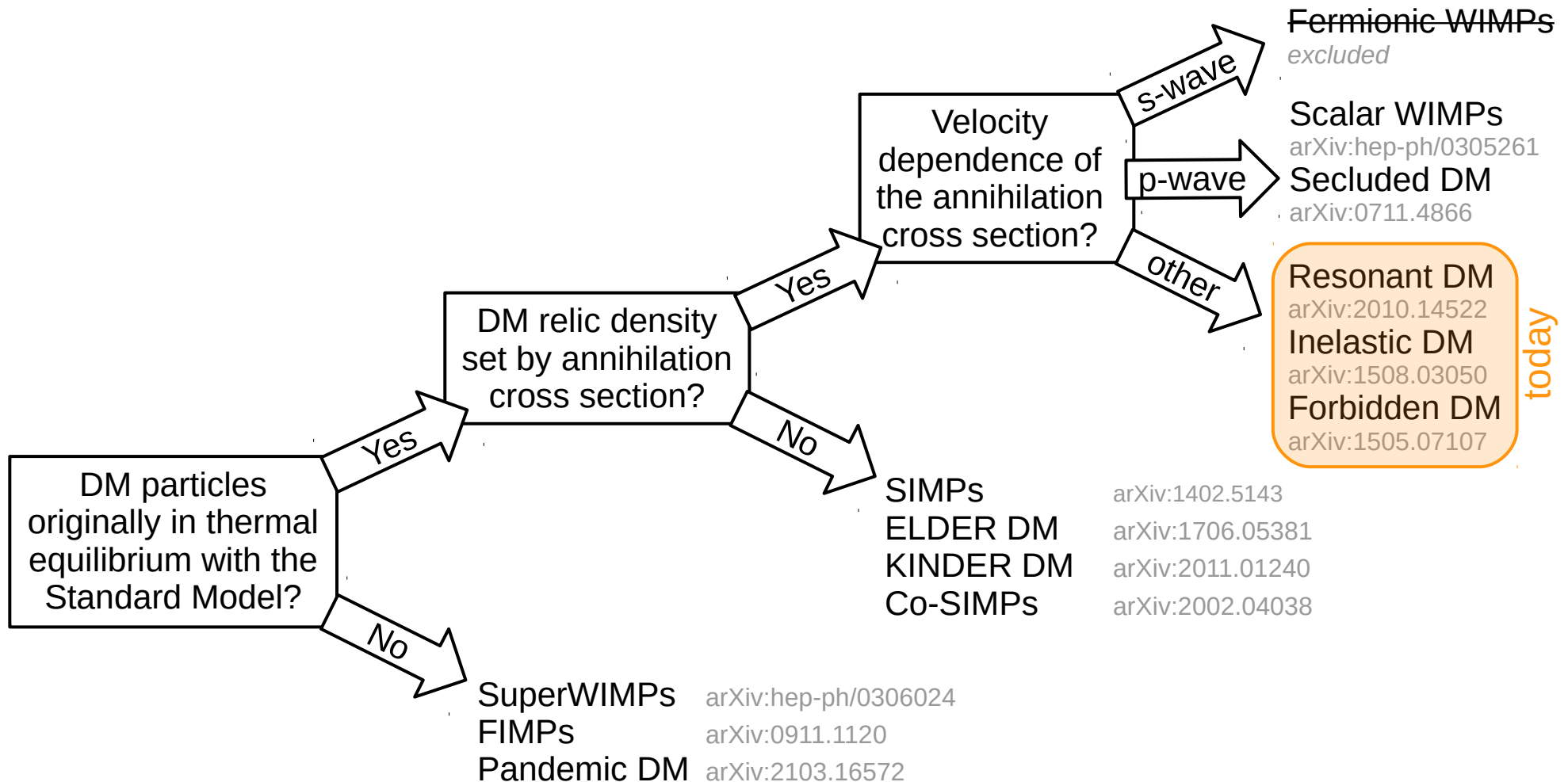
Including results from **arXiv:1907.04346**, **arXiv:1911.03176**, **arXiv:2010.14522**, **arXiv:2011.06604** and **ongoing work** in collaboration with Kai Böse, Juliana Carrasco Mejia, Elias Bernreuther, Michael Duerr, Torben Ferber, Chris Hearty, Saniya Heeba, Michael Krämer, Alessandro Morandini, Kai Schmidt-Hoberg and Patrick Tunney



Models for GeV-scale dark matter: Overview

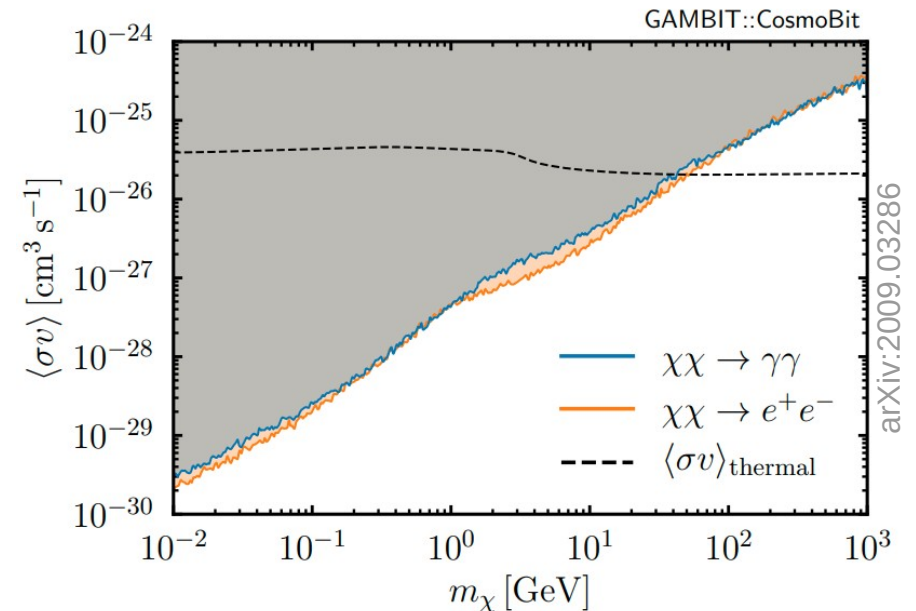


Models for GeV-scale dark matter: Overview



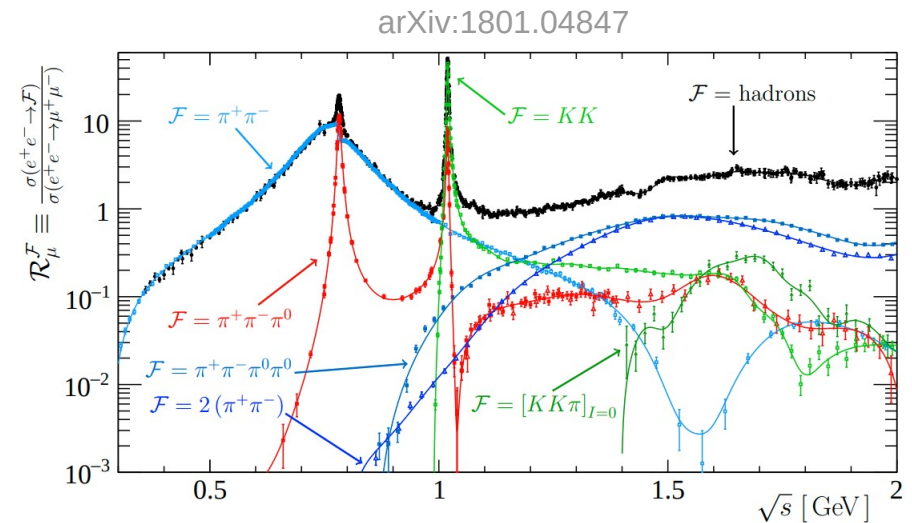
Velocity dependence of DM annihilations

- Thermal freeze-out requires annihilation cross section $\sigma v \sim 10^{-26} \text{ cm}^3/\text{s}$
- If σv is velocity independent, DM annihilations are still ongoing during recombination and in the present universe
 - Strong constraints on GeV-scale DM from CMB observations and indirect detection experiments
- To evade these constraints, it is necessary to suppress the annihilation cross section at small velocities
- Three main avenues:
 - Suppression of co-annihilation partners → Inelastic dark matter
 - Strongly energy-dependent matrix element → Resonant dark matter
 - Strongly energy-dependent phase space → Forbidden dark matter



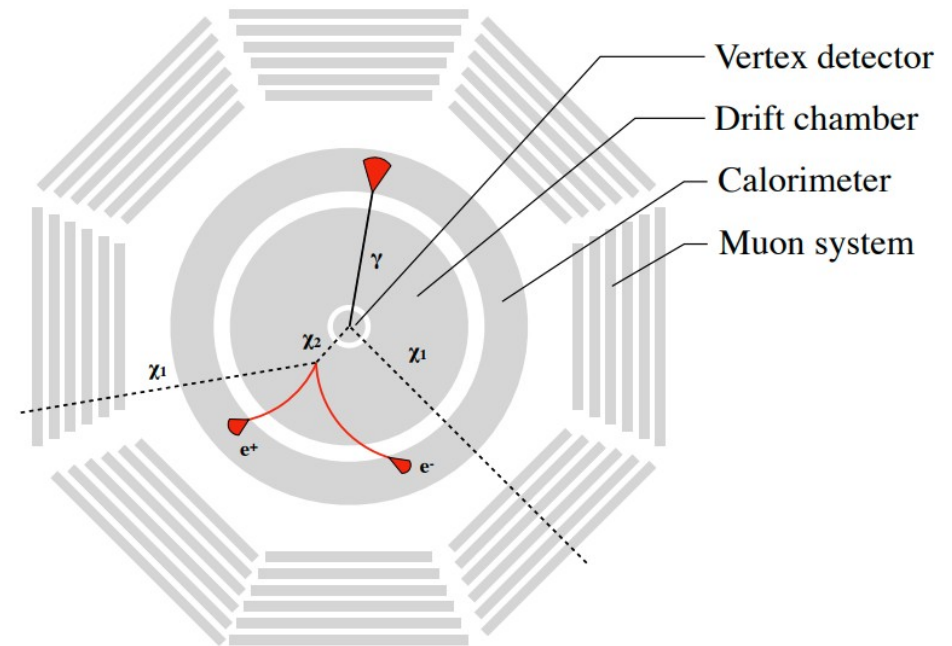
Quick interlude: Portal interactions

- Mechanisms discussed in this talk largely independent of how the dark sector couples to the Standard Model
- For simplicity, consider a dark photon mediator with kinetic mixing
 - Couplings proportional to electric charge
- Many other interesting options:
 - Gauged baryon-minus-lepton number (B-L)
 - Gauged baryon number (B)
 - Axial or chiral couplings
 - (Pseudo-)scalar mediators
- Determines experimental signatures
 - Final-state leptons?
 - Missing energy (neutrinos)?
 - Photonic decays?
- Ideally all of these possibilities should be explored!



Inelastic dark matter

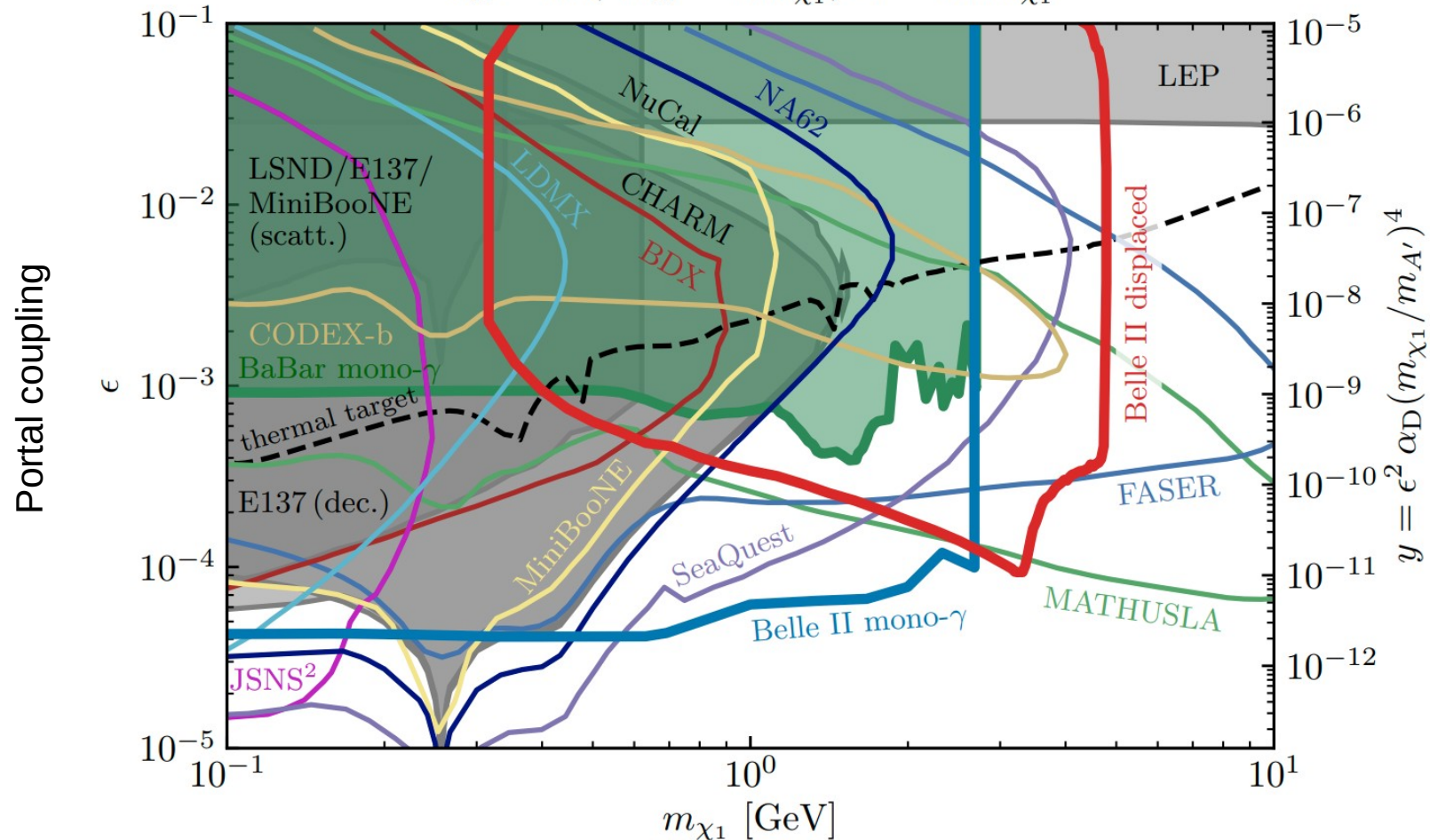
- Mass splitting Δ between ground state χ_1 and excited state χ_2
- Relative abundance of excited state scales $\sim \exp(-\Delta/T)$ in the early universe
- All interactions involve one ground state and one excited state
 - Annihilation rate becomes suppressed for $T < \Delta$
- Also: Strong suppression of scattering in direct detection experiments
- Key prediction: Long-lived excited state
 - Decay length may be in mm-cm range
 - Interesting decay mode: $\chi_2 \rightarrow \chi_1 e^+ e^-$
- Possible search at Belle II:
 - ISR photon allows for triggering (three isolated clusters)
 - Even better prospects with dedicated displaced vertex trigger
 - Complementary constraints from single-photon search



Inelastic dark matter: Results

Additional model parameters

$$\alpha_D = 0.1, m_{A'} = 3m_{\chi_1}, \Delta = 0.1m_{\chi_1}$$



arXiv:1911.03176

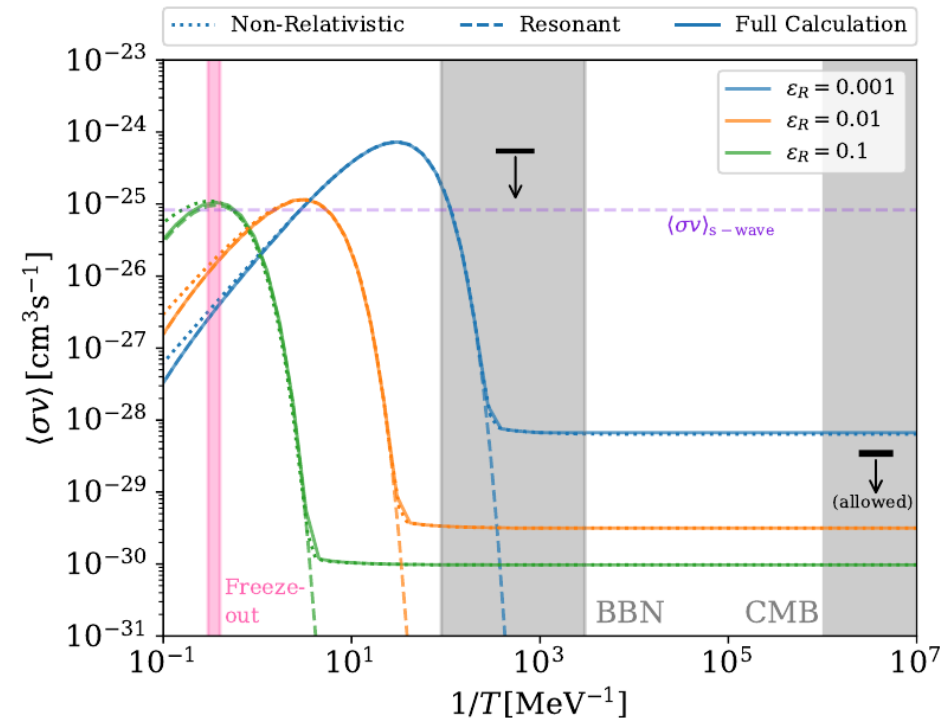
Dark matter mass

Resonant dark matter

- If the DM mass is close to twice the mediator mass, annihilations receive a resonant enhancement

$$\sigma v_{\text{lab}} = F(\epsilon) \frac{m_{A'} \Gamma_{A'}}{(s - m_{A'}^2)^2 + m_{A'}^2 \Gamma_{A'}^2}$$

- No strong constraints from CMB and indirect detection
- Relic density requirement can be satisfied even for tiny couplings



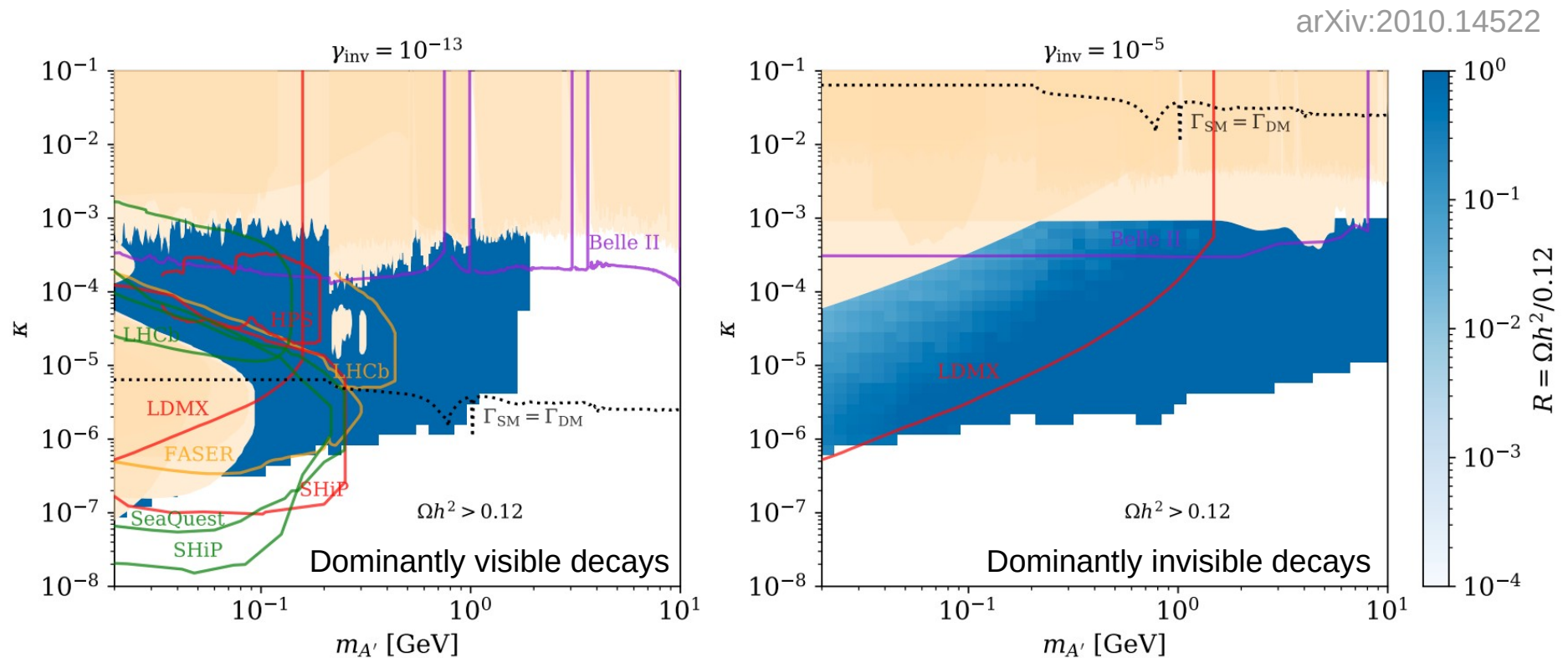
arXiv:2010.14522

- Dark sector characterised by two quantities

- Resonance parameter:
$$\epsilon_R = \frac{m_{A'}^2 - 4m_\chi^2}{4m_\chi^2}$$
- Reduced invisible width:
$$\gamma_{\text{inv}} \equiv \frac{\Gamma_{\text{DM}}}{m'_{A'}} = \frac{g_\chi^2}{12\pi} \left(1 - \frac{1}{1 + \epsilon_R}\right)^{1/2} \left(1 + \frac{1}{2(1 + \epsilon_R)}\right)$$

Resonant dark matter: Results

- Blue shading: Viable parameter space (light blue: viable only for DM sub-component)
- Orange shading: Existing constraints

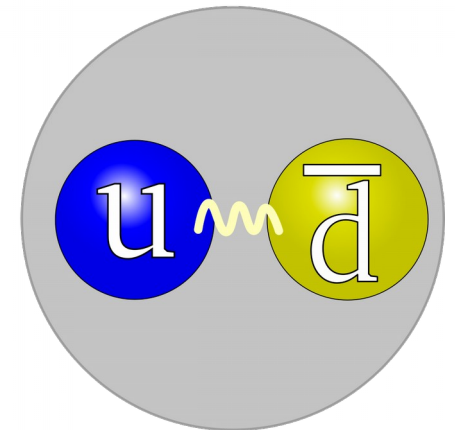


- Comprehensive exploration requires combination of searches for visible and invisible final states

Forbidden dark matter

- Dark matter freeze-out proceeds via annihilation into dark sector states: $x x \rightarrow A' A'$
 - Assumption: A' unstable with $m_{A'} < 2 m_x$
 - Any A' produced decays into SM states
 - Additional assumption: $m_{A'} > m_x$
 - Annihilations are kinematically allowed only at finite temperature/energy
 - Exponential suppression of annihilations at late times
- Original idea: A' is a dark photon that couples to SM particles via kinetic mixing
 - Similarity of $m_{A'}$ and m_x largely accidental
- Alternative perspective: x and A' both arise as bound states from a strongly-interacting dark sector that resembles QCD
 - Dark matter particles \leftrightarrow pseudoscalar mesons (dark pions)
 - Annihilation partners \leftrightarrow vector mesons (dark rho mesons)
 - Underlying motivation for $m_x < m_{A'} < 2 m_x$

arXiv:1505.07107



arXiv:1907.04346

Strongly-interacting dark sectors: Example

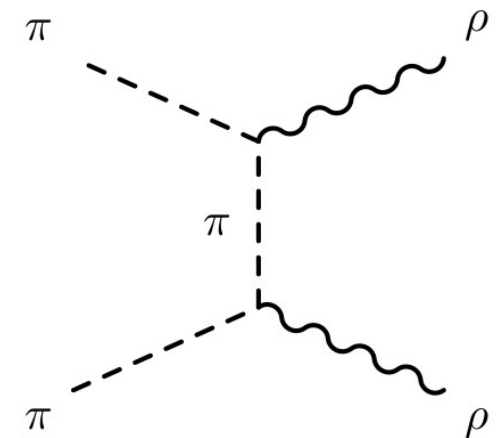
- Consider $SU(N)$ gauge group with $N = 3$ (like QCD)
- Focus on the case of two light quarks with equal mass
- Assume that the two quarks couple to a dark photon with opposite charge
 - Confinement gives rise to three light pions (π^+ , π^- , π^0)
 - All three pions may be stable and suitable dark matter candidates

arXiv:1801.05805

- The neutral vector meson (ρ^0) mixes with both the dark and the SM photon
 - Long-lived particle that decays into SM final states

- No $3 \rightarrow 2$ processes (as required for the SIMP mechanism)
- Relic density set by the forbidden annihilations $\pi \pi \rightarrow \rho^0 \rho^0$

arXiv:1907.04346



Strongly-interacting dark sectors at Belle II

- No preferred energy scale for the confinement of the dark sector
- Confinement scales < 50 MeV are in conflict with bounds on DM self-interactions
 - Interesting to think about dark sectors in the 100 MeV – 1 GeV range

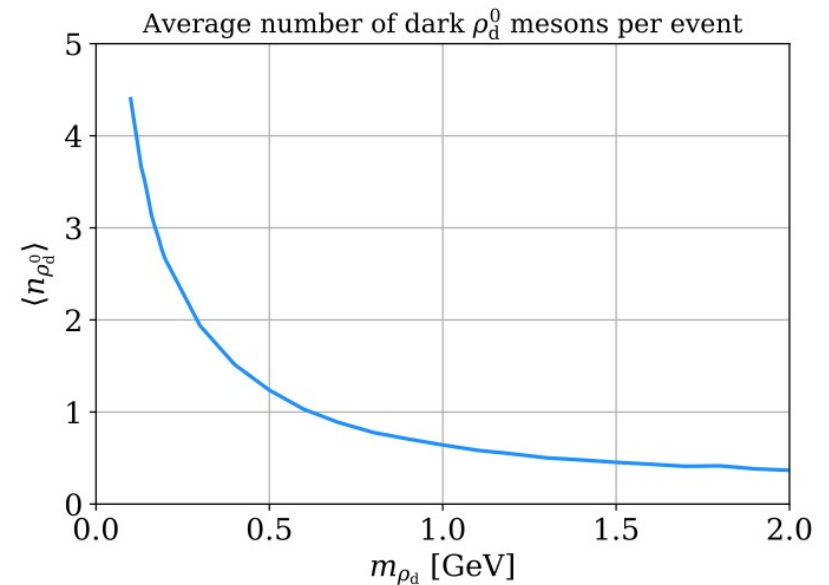
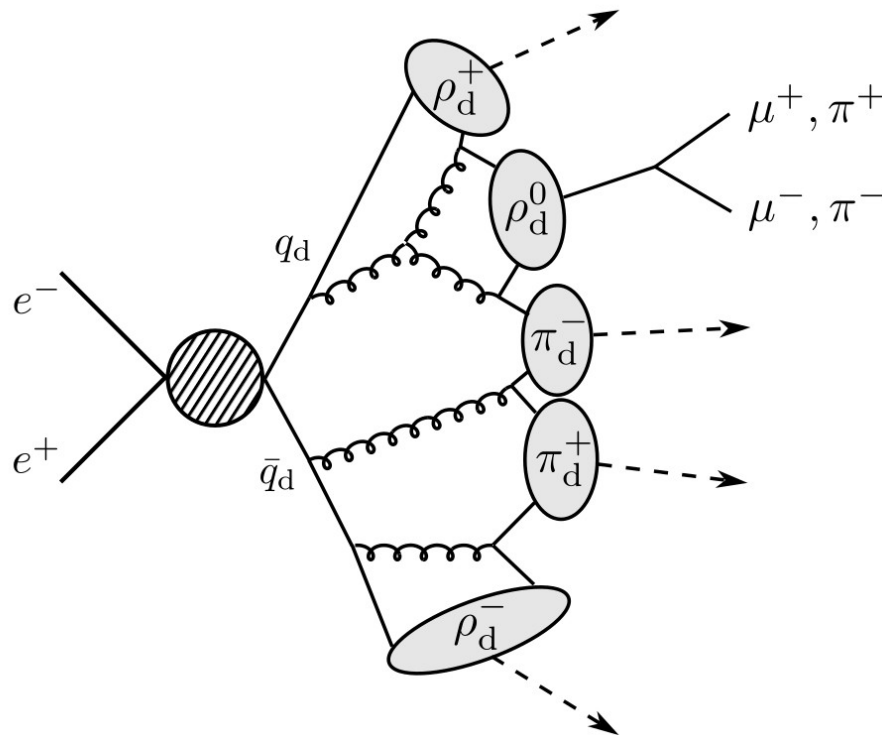
- Dark photon could be significantly heavier
 - Interactions between dark quarks and SM described by effective operator

$$\mathcal{L}_{\text{eff}} \supset \frac{1}{\Lambda^2} \sum_f q_f \bar{f} \gamma^\mu f \bar{q}_d \gamma_\mu q_d$$

- For $\Lambda \sim \text{TeV}$ the dark rho meson has detector-size decay length
- Highly interesting scenario for electron-positron colliders!

Dark showers at Belle II

- Dark quarks produced in e^+e^- collisions will hadronise and create a dark shower
- Multiplicity (and boost) of long-lived dark rho mesons depends on mass scale



- Possible strategy: Search for events with a muon pair from a displaced vertex

Existing exclusion limits and projections

- Existing BaBar limit from model-independent search for LLPs

arXiv:1502.02580

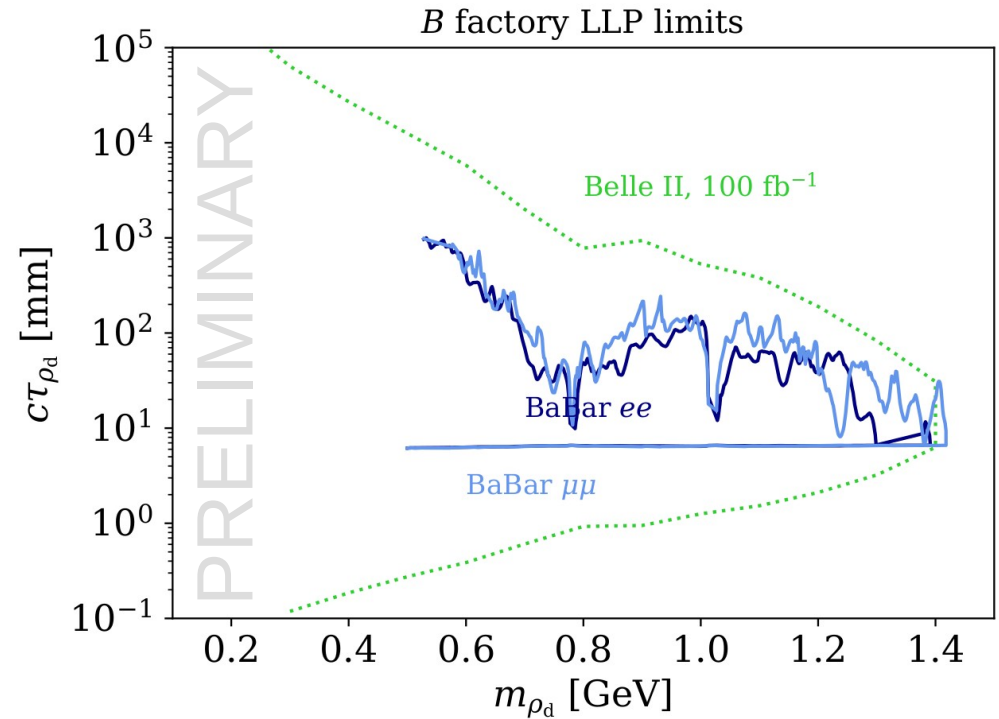
- Belle II projection based on similar assumptions as for inelastic DM

arXiv:2012.08595

- Interesting parameter regions compatible with other constraints (EWPT, Z boson invisible width, ...)

- Not shown: Additional exclusion limits from searches for displaced di-muon resonances in LHCb

→ Comparison cannot be done in EFT approach (dark photon produced on-shell)



Conclusions

- Huge variety of models for GeV-scale dark sectors
- Dark matter particles may reproduce observed relic abundance via freeze-out
- Constraints require annihilation rate with non-standard velocity dependence
- **Inelastic dark matter**
 - Excited state has three-body decay with macroscopic decay length
- **Resonant dark matter**
 - Tiny couplings viable for both visible and invisible decays
- **Forbidden dark matter**
 - Well-motivated annihilation partners from strongly-interacting dark sectors
 - Prediction: Dark showers at e^+e^- collisions
 - Evaluation of existing constraints and projected sensitivities ongoing