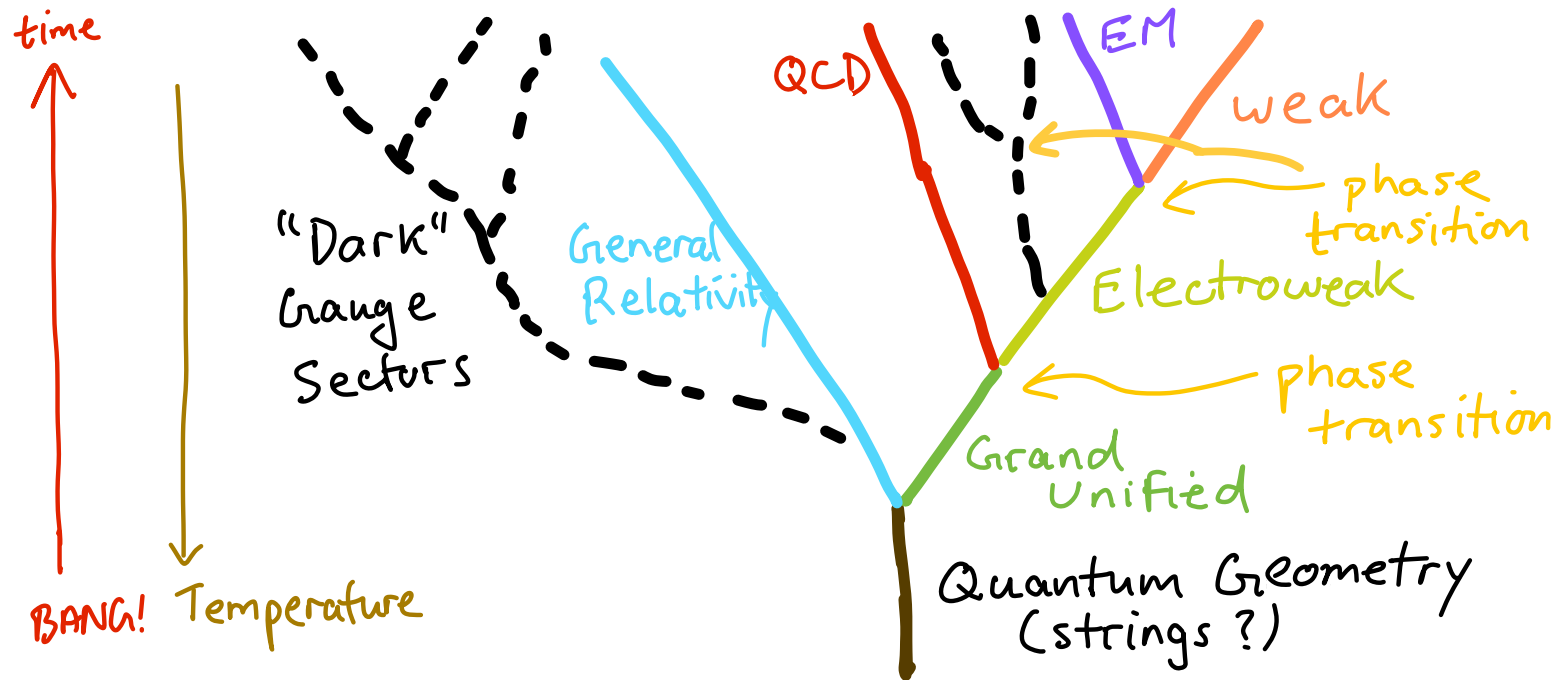


B S M FOR  
FUTURE  
COLLIDERS

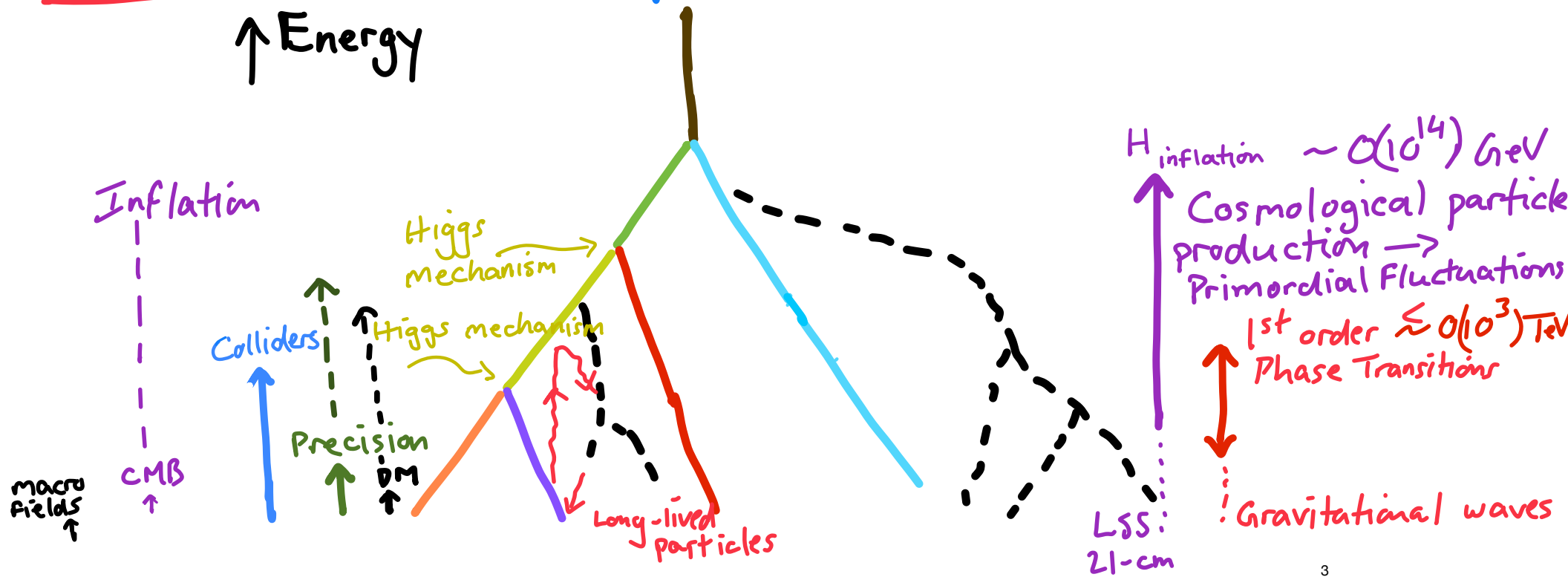
Raman Sundrum  
University of Maryland

# FAMILY TREE OF FUNDAMENTAL PHYSICS



# AN EXPLORER'S MAP

Future colliders would fit into a diverse & heroic push to map Nature's laws.



# OUTLINE

Higgs-Ploration

Composite Resonances

Gravitational Waves from Cosmological Phase Transitions

WIMP Dark Matter

Matter-Antimatter Asymmetry

Long-Lived Particles

Supersymmetry

The Little Hierarchy Problem & the Anthropic Principle

Ghost ("dark") Sectors

Higgs Portal

Hidden Naturalness

# The HIGGS ENIGMA

2012 Higgs discovery revealed a new QUANTUM  
FUNDAMENTAL FORCE, with a

SCALAR "EQUIVALENCE PRINCIPLE" (EP):

COUPLING  $\propto$  MASS  
to Higgs boson

from Higgs Mechanism  
+ QUANTUM CORRECTIONS

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COUPLING  $\propto$  MASS

to Higgs boson

from Higgs Mechanism

+ QUANTUM CORRECTIONS

$m_h = 125 \text{ GeV} \Rightarrow$  rich in accessible production/decay channels.

$\Rightarrow$  SHOULD SIGNIFICANTLY IMPROVE PRECISION TESTS OF quantum-corrected "EP": Yukawas<sup>6</sup>, di-Higgs, couplings to gauge bosons, EWPT

SM accommodates, but does not explain:

Overall mass scale  $v_{\text{weak}} \sim \langle H \rangle \lll M_{\text{Pl}}$  Hierarchy Problem

Relative mass scales, mixing angles exhibit mysterious patterns (Flavor Puzzle (gauge coupling unification))

And yet, attractive explanatory mechanisms may be within reach...

# HIGGS MAY BE LIGHT COMPOSITE OF NEW SECTOR

Model (in)dependently testing Higgs  
for compositeness presents classic goal  
for experimental exploration.

Higgs Compositeness  $\Rightarrow$  small violations of "EP"  
 $\Rightarrow$  precision Higgs, EW tests

+ more indirectly, FCNCs & ~~CP~~  
in flavor tests, EDMs

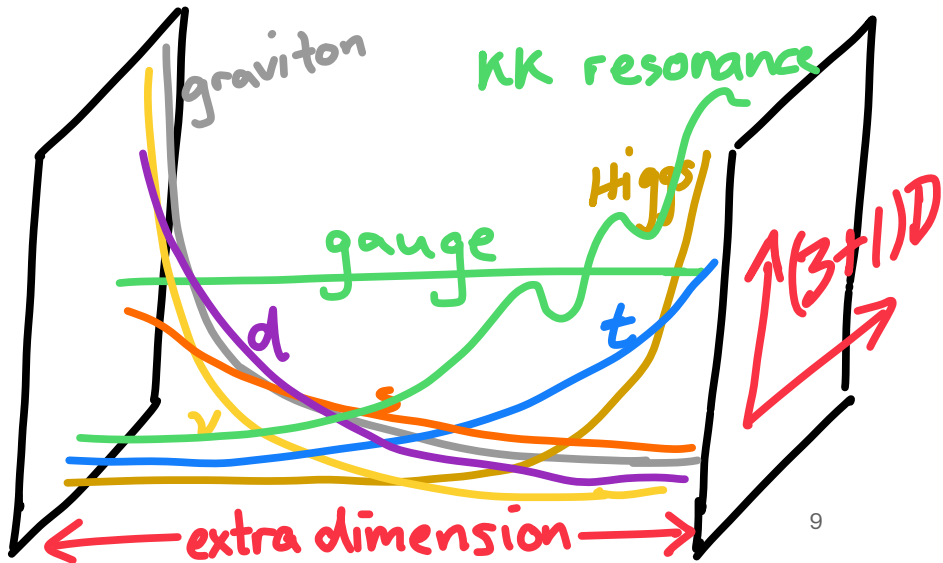


# HIGGS COMPOSITENESS

(& PARTIAL COMPOSITENESS)

Theoretically tough due to strong confining coupling.  
But can be "geometrized" via AdS/CFT duality:

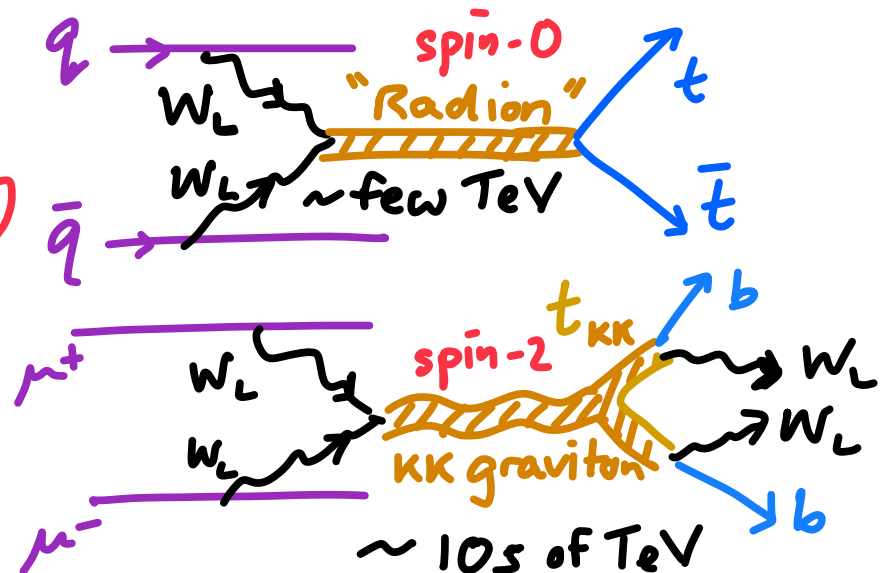
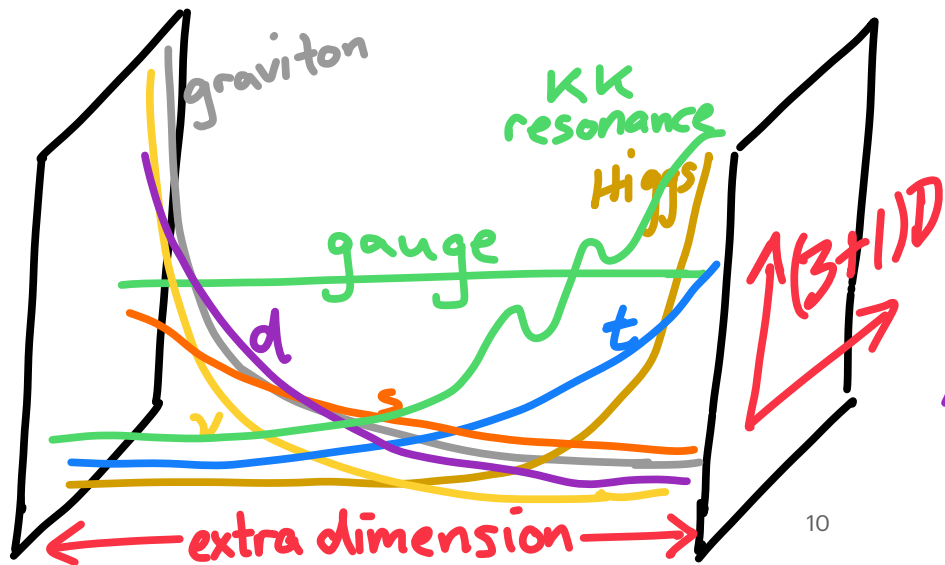
Strong Coupling  $\xrightarrow{\text{Quantum Magic}}$  Emergent "Warped" Extra Dimension



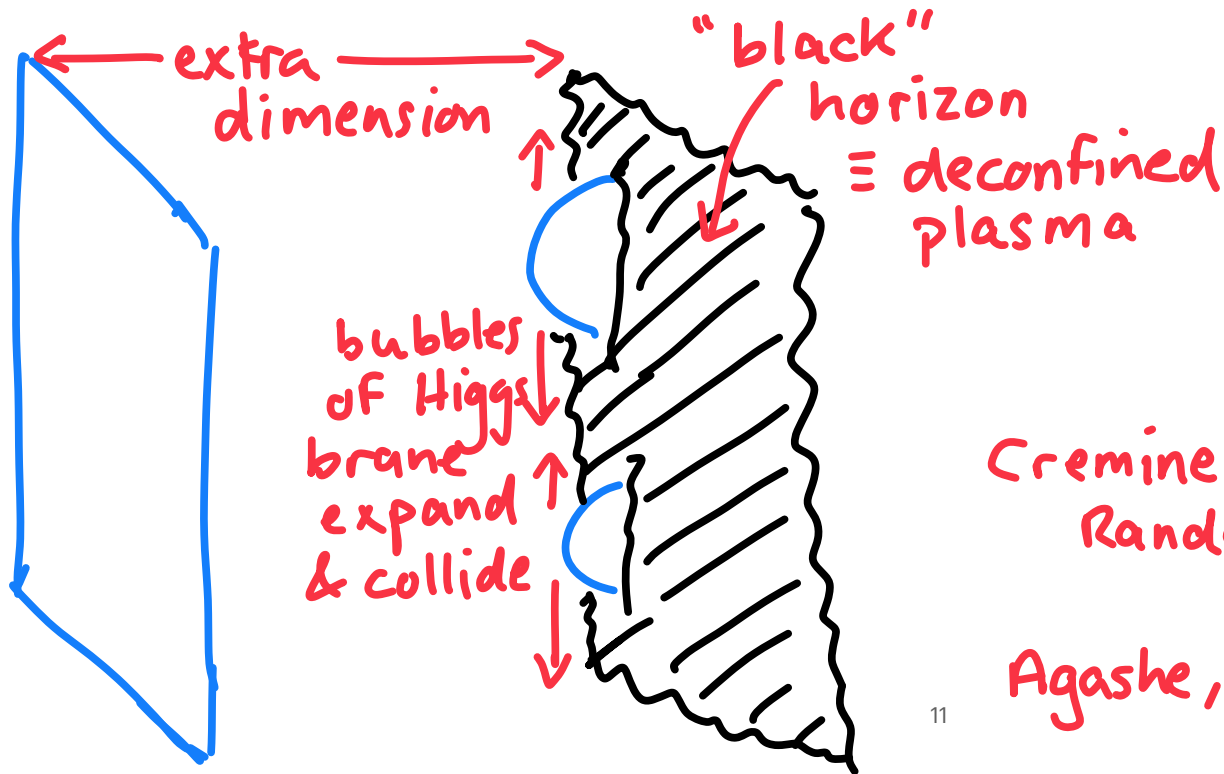
# WARPED EXTRA DIMENSION

Extra-dimensional wavefunction overlaps  $\rightarrow$  attractive resolutions of Hierarchy Problem + Flavor Puzzle.

Known models can have Kaluza-Klein resonances typically at  $\gtrsim 20$  TeV consistent with stringent flavor/CP tests. But some resonances could be significantly lighter:



# (DE)CONFINEMENT PHASE TRANSITION of Composite Higgs sector in v. Early Universe

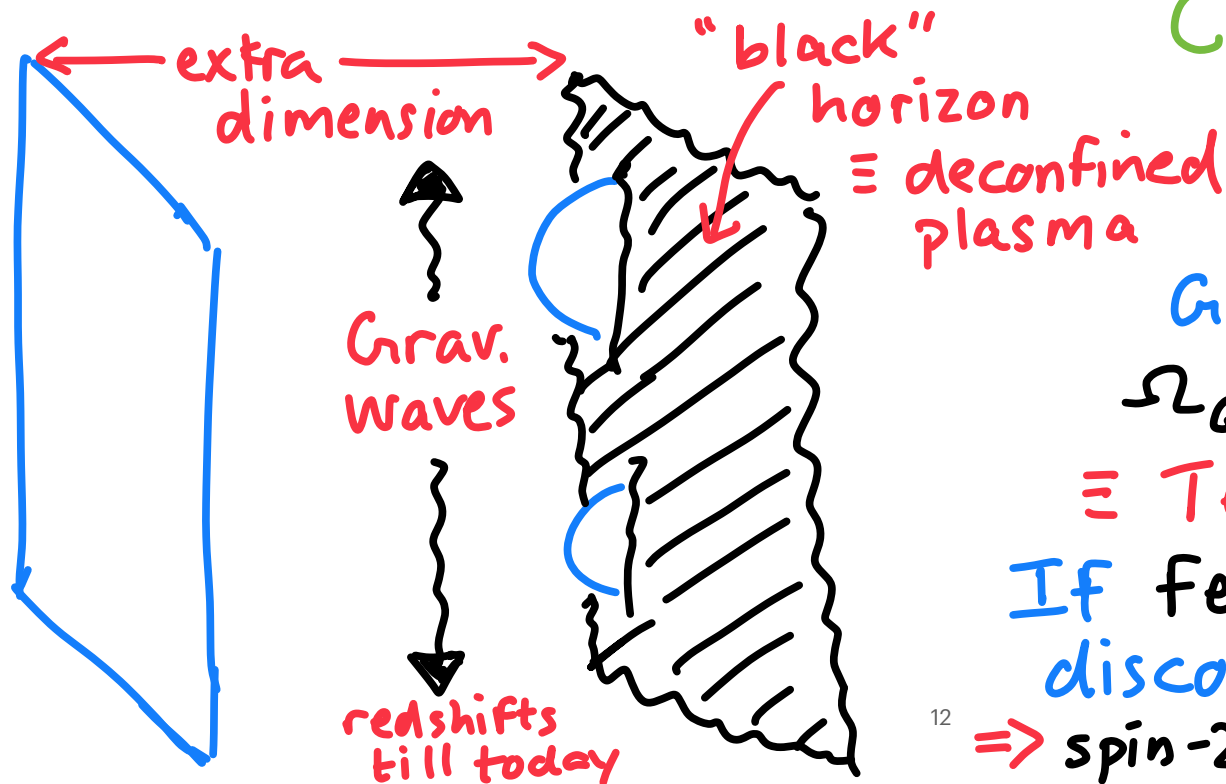


Creminelli, Nicolis, Rattazzi '02  
Randall, Servant '06

⋮  
Agashe, Du, Ekhterachian, Kumar,  
Sundrum '20

# GRAVITATIONAL WAVES

## ← (DE)CONFINEMENT PHASE TRANSITION



## COLLIDER COMPLEMENTARITY

Eg. Proposed LISA detector may see 1mHz GW stochastic background

$$\Omega_{GW} \sim \text{few} \cdot 10^{-10}$$

≡ TeV critical Temperature

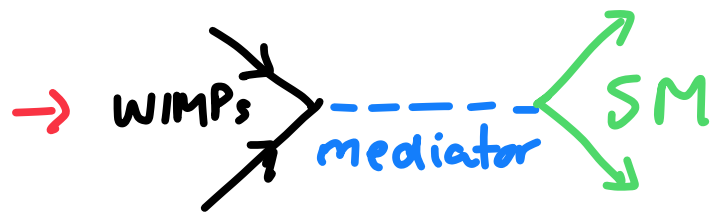
IF few TeV scalar radion discovered at collider

⇒ spin-2 KK graviton  $\approx 20$  TeV

# THE HEART OF THE MATTER

## Dark Matter WIMP "Miracle":

Early Universe equilibrium ...  
 $WIMP \leftrightarrow SM$



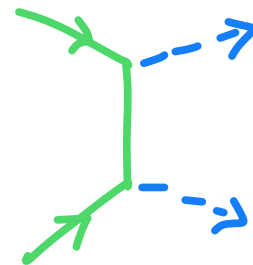
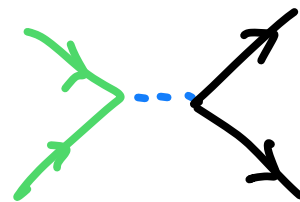
Freezes out  
 as universe expands

Energy fraction of universe today:

$$\Omega_{WIMP} \approx 0.1 \left( \frac{g_{weak}}{g_{WIMP}} \right)^4 \cdot \frac{m_{mediator}^4}{m_{WIMP}^2 \text{TeV}^2}$$

$\Rightarrow$  TeV WIMPs attractive

Broadest  
 Collider exploration  
 $\rightarrow$  with optimal  $E$  detection,  
 complementary to DM  
 detection experiments



# THE HEART OF THE MATTER S

## Dark Matter WIMP "Miracle":

Early Universe  
equilibrium ...  
WIMP  $\leftrightarrow$  SM



Freezes out  
as universe expands

Energy fraction of  
surviving WIMPs today

$$\Omega_{\text{WIMP}} \approx$$

$$0.1 \left( \frac{g_{\text{weak}}}{g_{\text{WIMP}}} \right)^4 \cdot \frac{m_{\text{mediator}}^4}{m_{\text{WIMP}}^2 \text{TeV}^2}$$

## BUT WHAT ABOUT THAT OTHER

## FORM OF MATTER - US?!

ie. Baryons

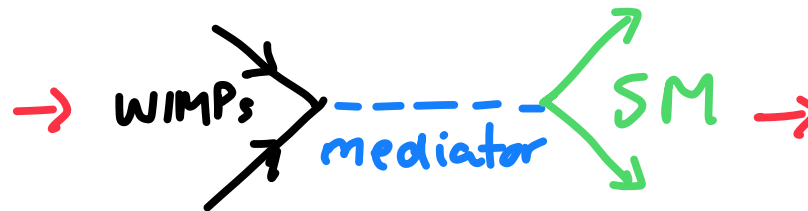
# BARYOGENESIS FOR WIMPS!

Cui, Sundrum '12

## Baryogenesis WIMP "Miracle":

Energy fraction of surviving WIMPs

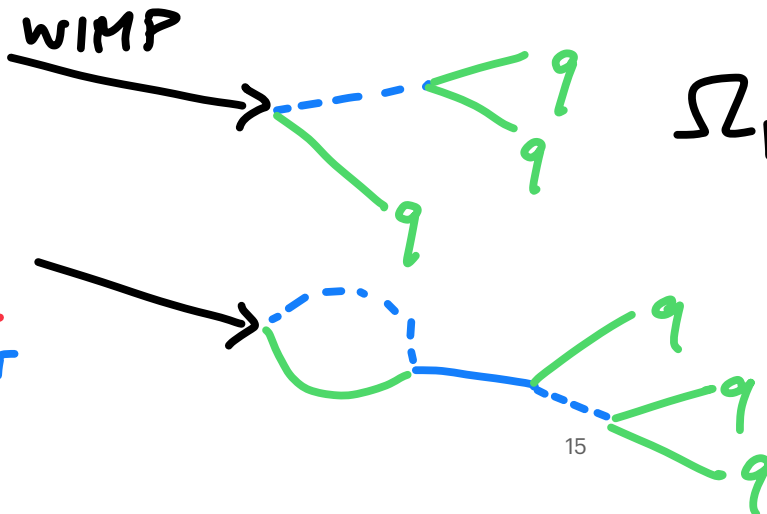
Early Universe equilibrium ...  
WIMP  $\leftrightarrow$  SM



Freezes out as universe expands

$$\Omega_{\text{WIMP}} \approx 0.1 \left( \frac{g_{\text{weak}}}{g_{\text{WIMP}}} \right)^4 \frac{m_{\text{mediator}}^4}{m_{\text{WIMP}}^2 \text{TeV}^2}$$

long-lived unstable WIMP  
O(1) ~~CP~~  
Baryon number ~~number~~



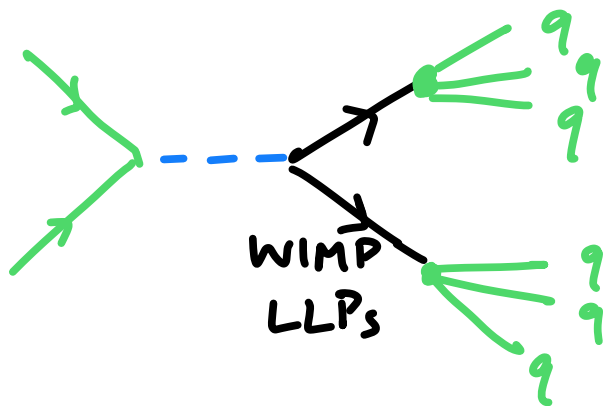
$$\Omega_{\text{baryon}} \approx \text{loop} \cdot \frac{m_{\text{proton}}}{m_{\text{WIMP}}} \Omega_{\text{WIMP}}$$

O(1) differences in stable/unstable WIMPs can accommodate  $\Omega_{\text{baryon}} \sim 1/5 \Omega_{\text{DM}}$

# BARYOGENESIS FOR WIMPS!

→ Broadest Collider Exploration  
with optimal sensitivity to v. weakly  
produced Long-Lived Particles (LLPs)

$\tau_{\text{WIMP}} \gtrsim \text{cm}$  for out-of-equilibrium WIMP decay  
after freezeout.



Even more broadly, other  
baryogenesis mechanisms,  
EW or other phase transition  
physics, may be within reach.



# SUPERSYMMETRY

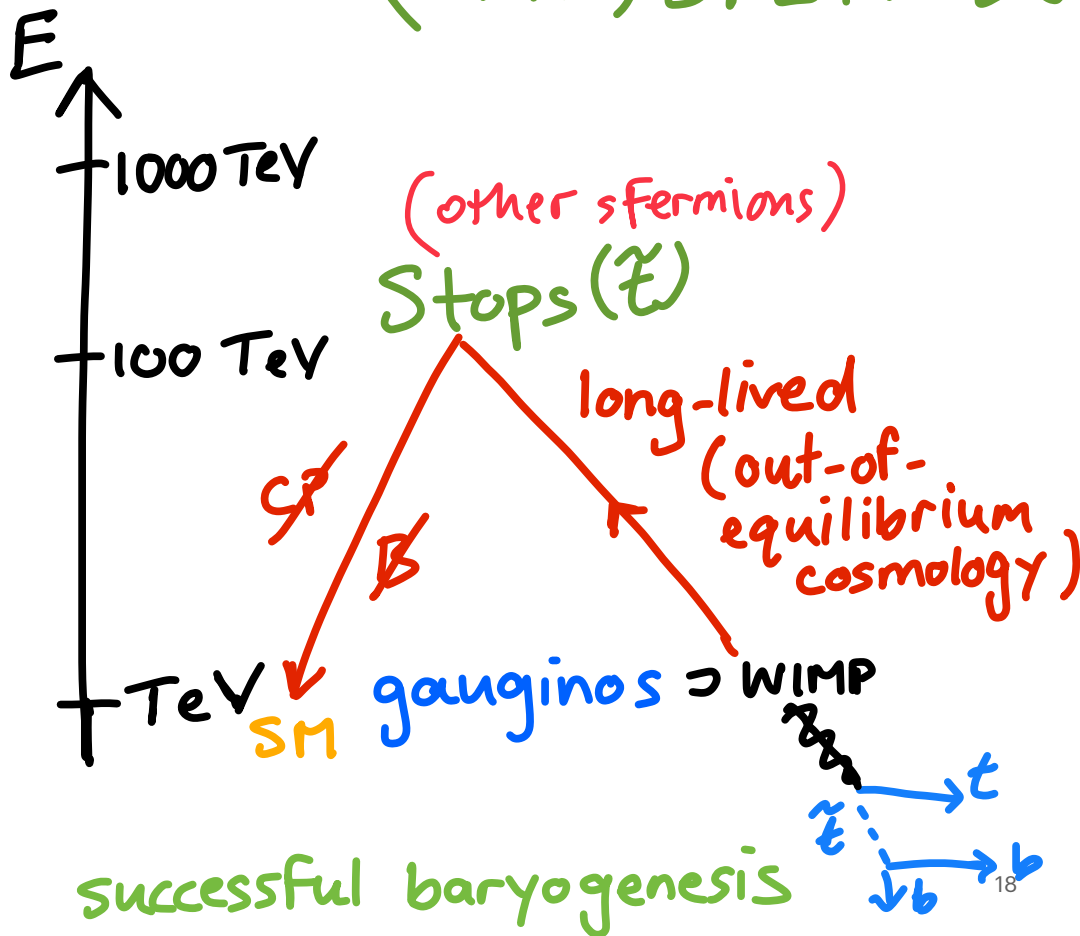
remains a strongly motivated paradigm for resolving Hierarchy Problem, & scaffolding for other mechanisms (eg. DM, Baryogenesis, ...)

$m_h = 125 \text{ GeV}$  (+ absence of superpartners @ LHC) suggests  $M_{\text{stops}} > 10 \text{ TeV}$

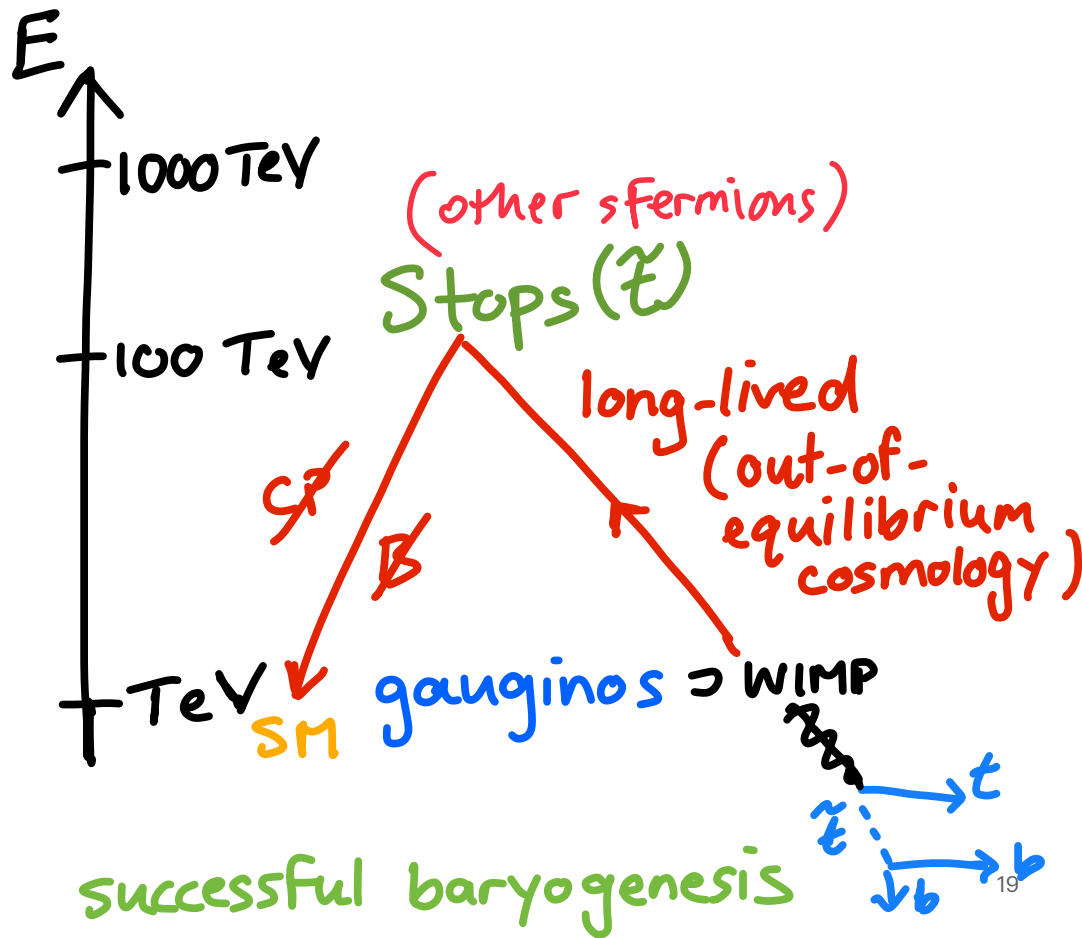
# BARYOGENESIS FOR WIMPS in (mini-)SPLIT SUSY

Cui, Sundrum '12  
Cui '13

R-Parity violating  
SUSY



# BARYOGENESIS vs. SUSY ?

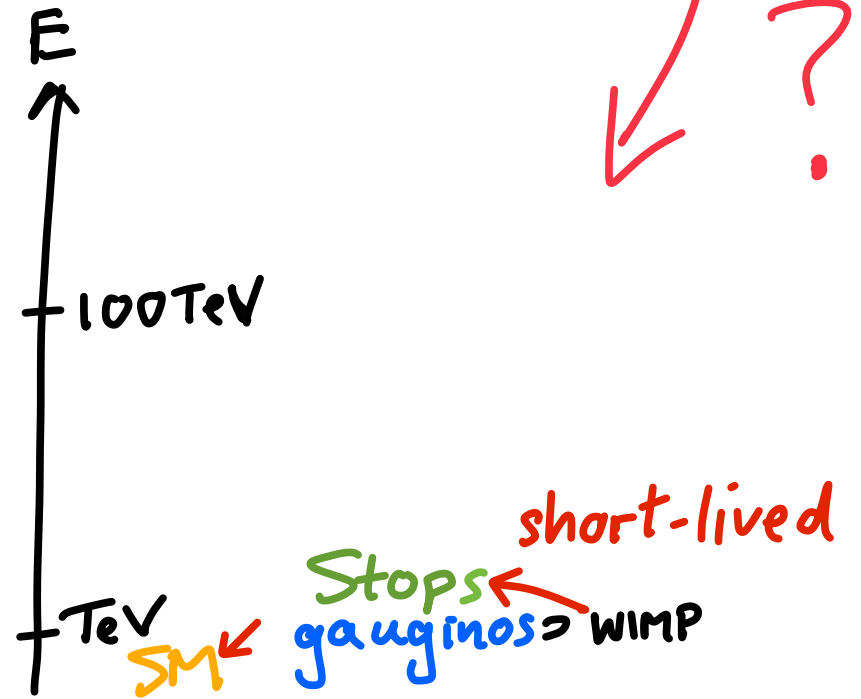
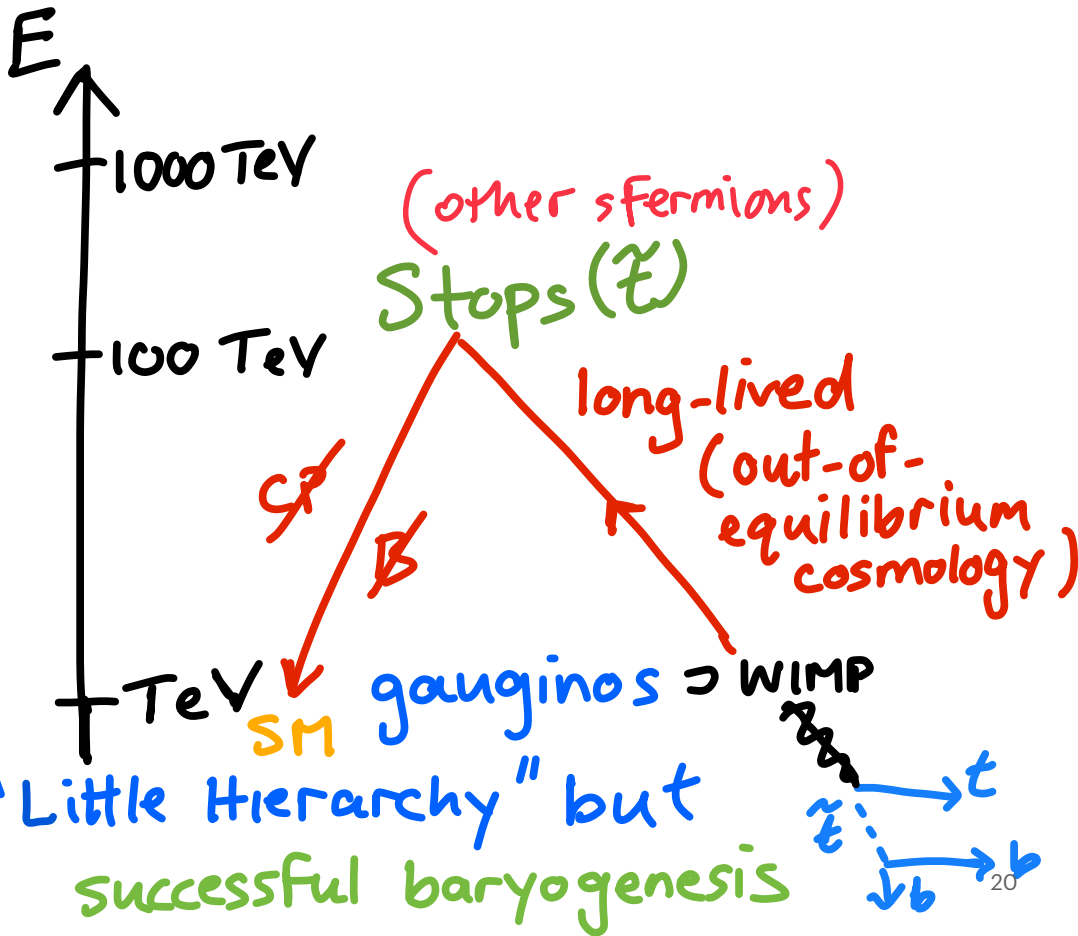


Cui, Sundrum '12  
Cui '13

R-Parity violating  
SUSY

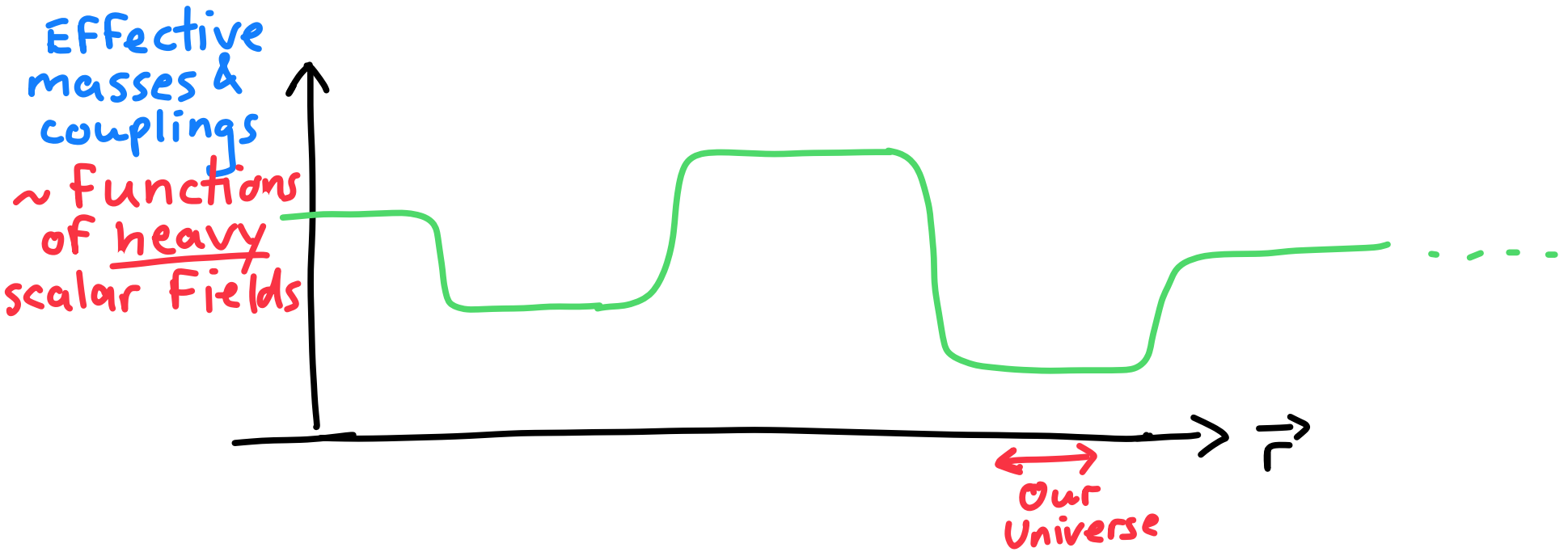
BUT  
Why is SUSY  
not more natural?  
Less fine-tuned in  
Solving Hierarchy  
Problem?

WHY NOT

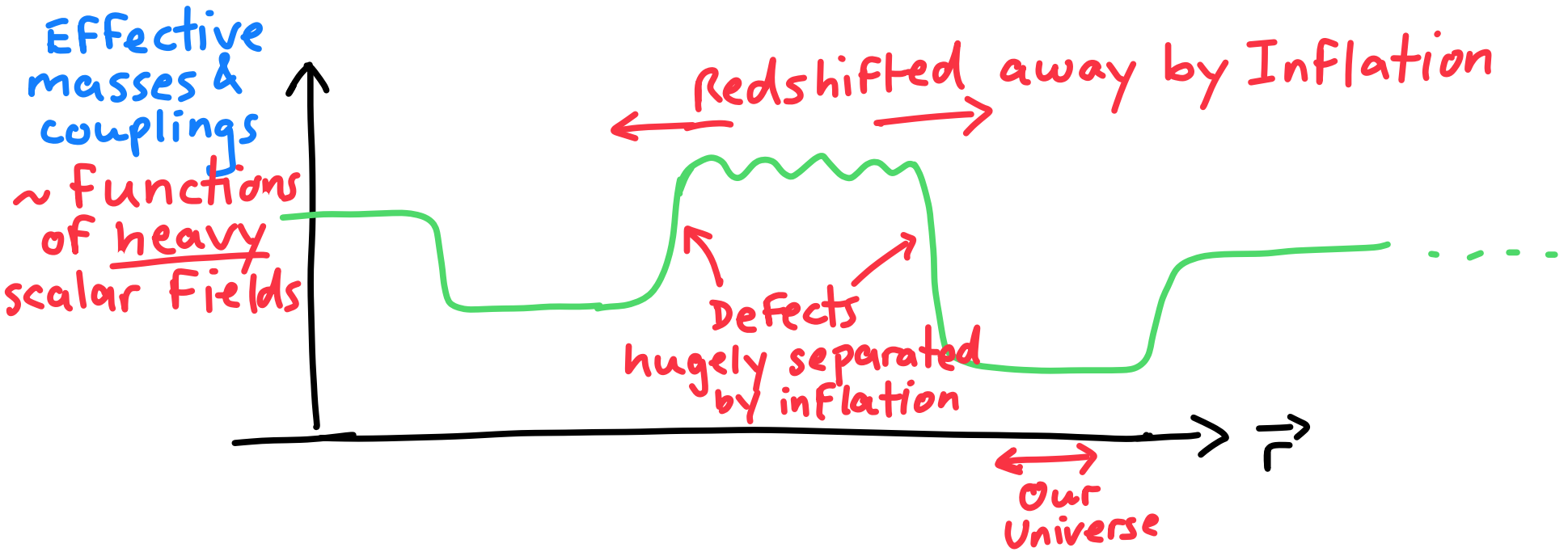


All "NATURAL" BUT NO BARYONS!

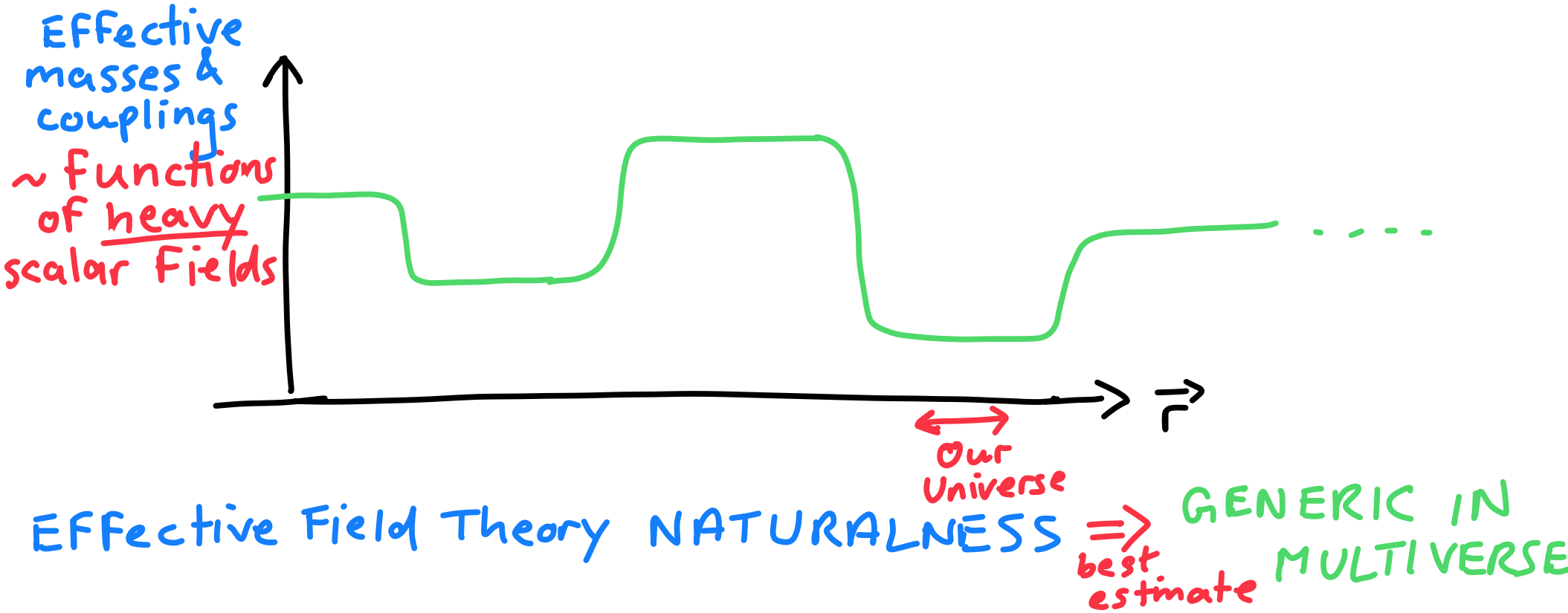
# SATELLITE VIEW OF THE MULTIVERSE



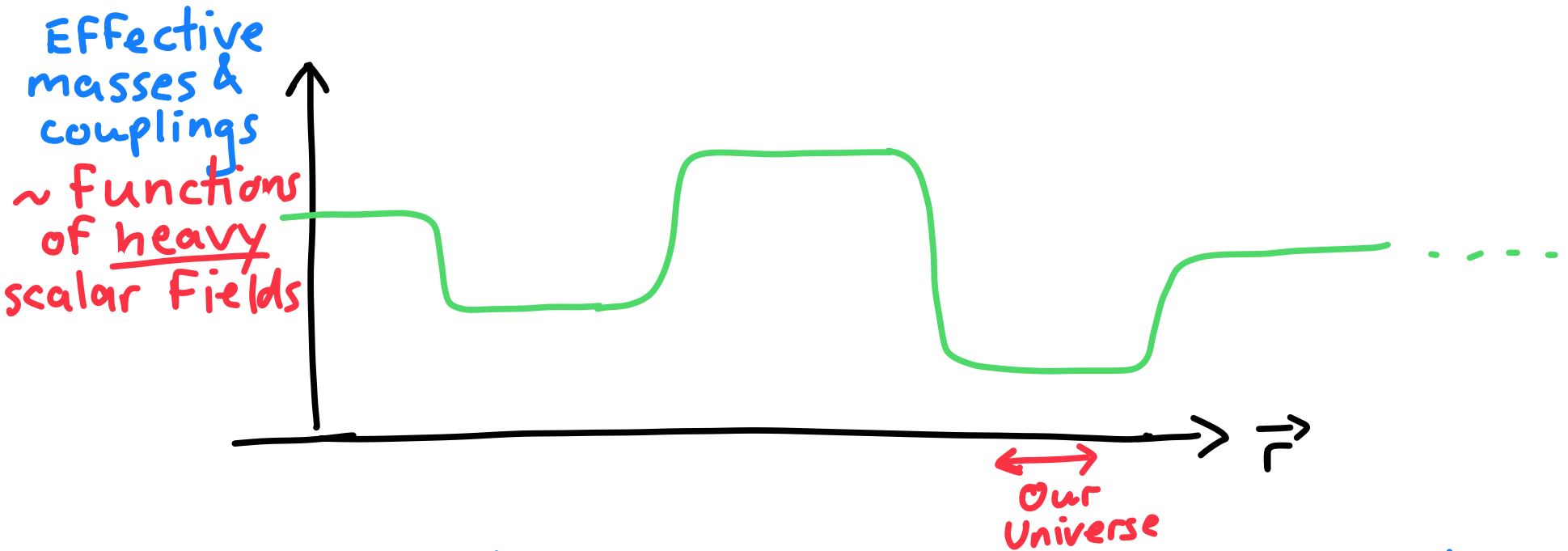
# SATELLITE VIEW OF THE MULTIVERSE



# SATELLITE VIEW OF THE MULTIVERSE



# FRUSTRATED NATURALNESS



There is a (top-sided) competition between Naturalness (genericity) & ANTHROPIC requirements (eg. baryons)  $\rightarrow$  LITTLE HIERARCHY<sup>24</sup> "PROBLEM"



# ARE WE ALONE ?

Gauge Field Theory naturally divides into "socially-distanced pods" (sectors) of gauge fields & charged matter.

SM is one such, but are there other "ghostly" gauge sectors awaiting discovery?

In Supergravity models, EWSB radiatively triggered by ~~SUSY~~  $\sim m_{\text{gravitino}}$ . Same mechanism can naturally lead to weak-scale "ghost" sectors

# SPECIES versus POPULATION

Massless and/or stable "ghost" populations are constrained:

$$\Delta N_{\text{eff}} < 0.4$$

cosmological constraints on new relativistic particles

$$\Omega_{\text{dark}} < 0.2$$

massive stable particles must fit within DM.

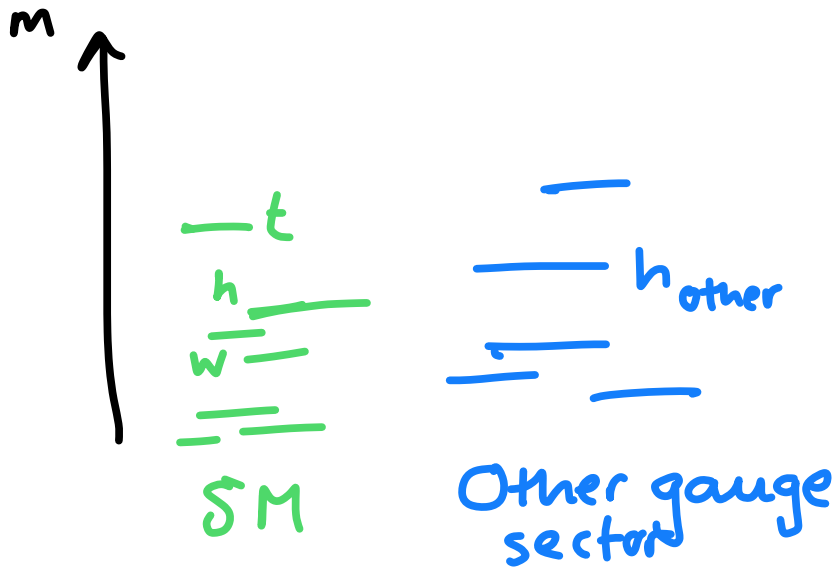
Even 1000's of new gauge sectors allowed phenomenologically/string-theoretically (!) if poorly "reheated" post-inflation, or unstable to SM decay.

# THE HIGGS PORTAL

to other gauge sectors

$$\mathcal{L} \supset \lambda H_{SM}^\dagger H_{SM} H_{other}^\dagger H_{other}$$

is special,  
renormalizable  
(ie. efficient)  
window of opportunity



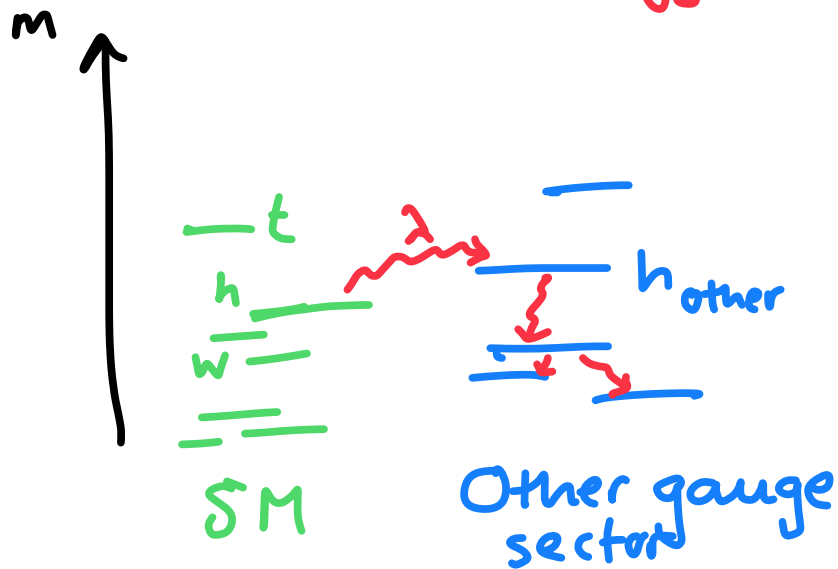
# THE HIGGS PORTAL

to other gauge sectors

$$\mathcal{L} \supset \lambda \langle H_{SM}^\dagger \rangle H_{SM} H_{other}^\dagger \langle H_{other} \rangle$$

Higgs mixing

is special,  
renormalizable  
(ie. efficient)  
window of opportunity



$\Rightarrow$  Exotic Higgs decays,  
important to confidently  
detect modest ~~E~~ at  
modest rates

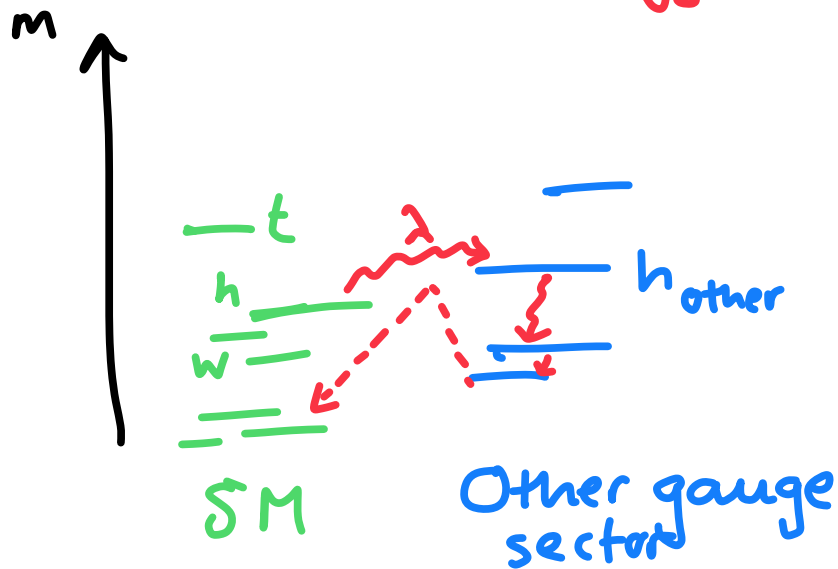
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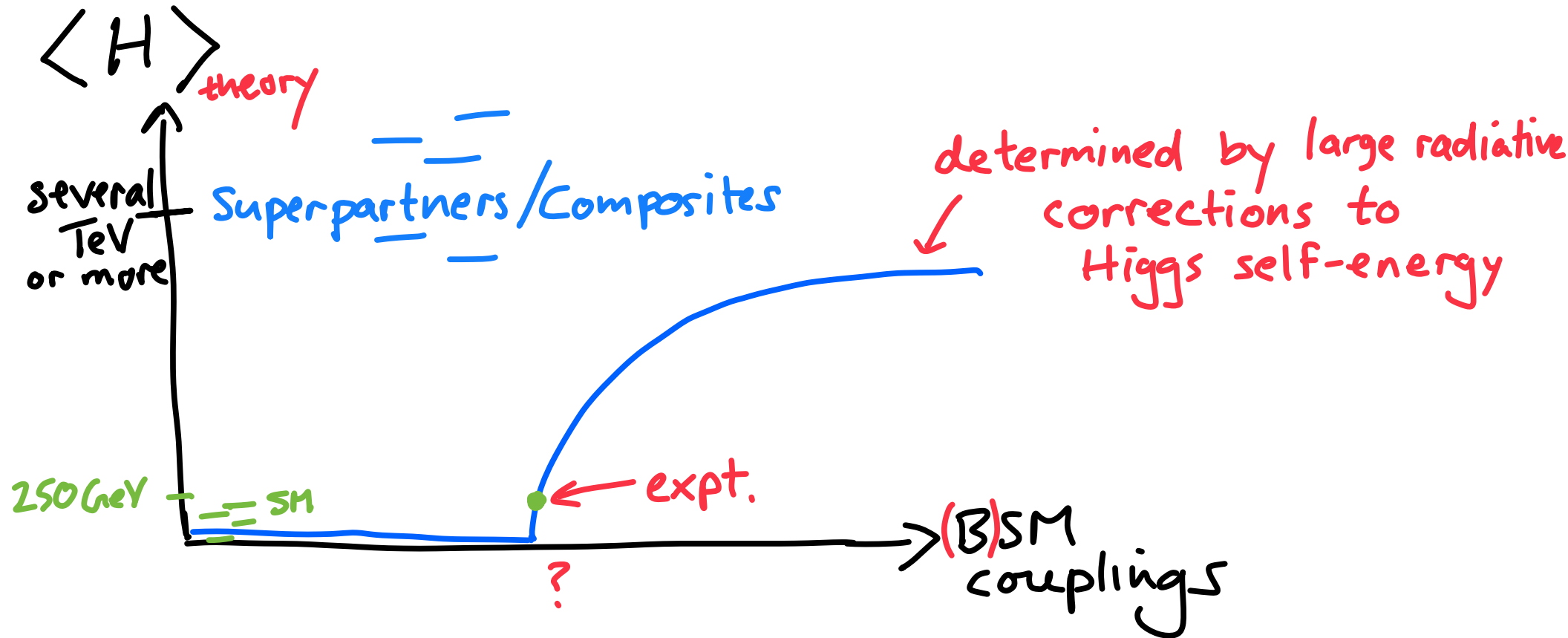
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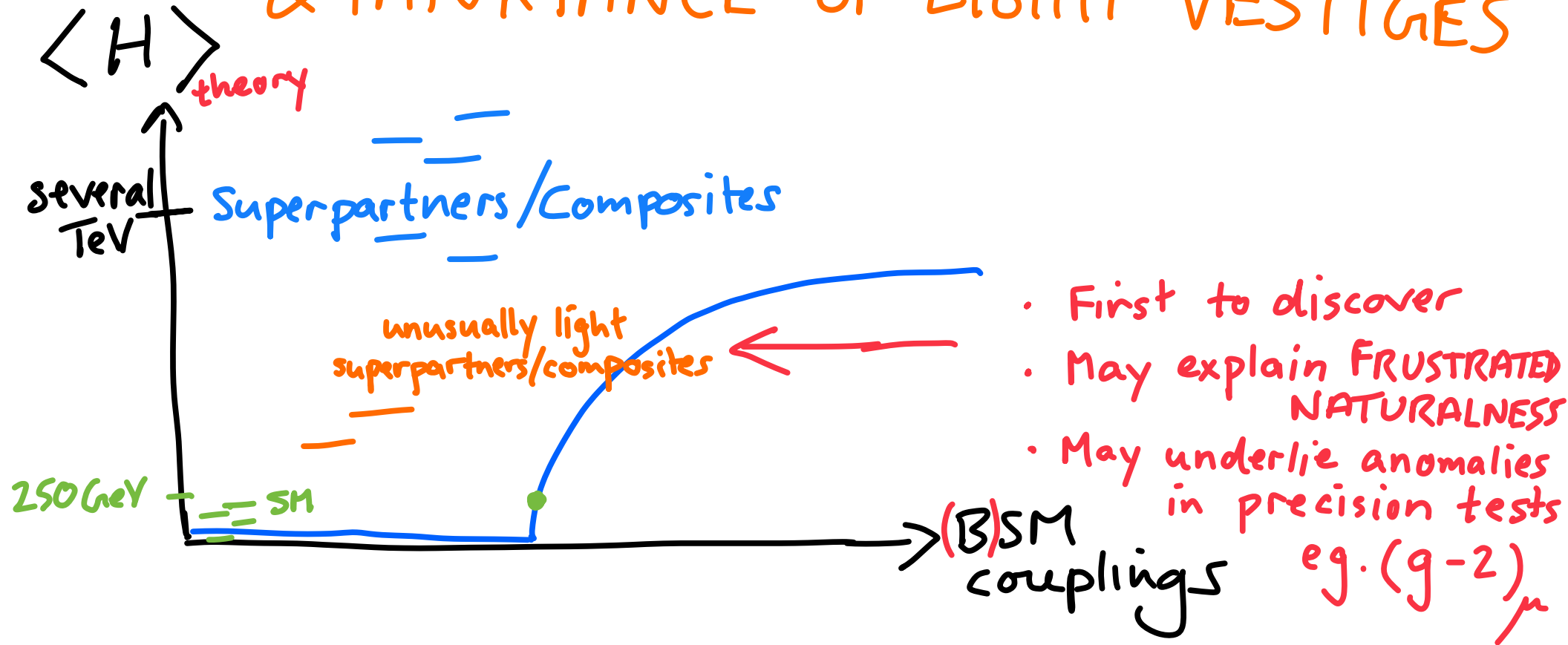


⇒ Important to detect long-lived decays from other sectors back to SM.

# THE (LITTLE) HIERARCHY



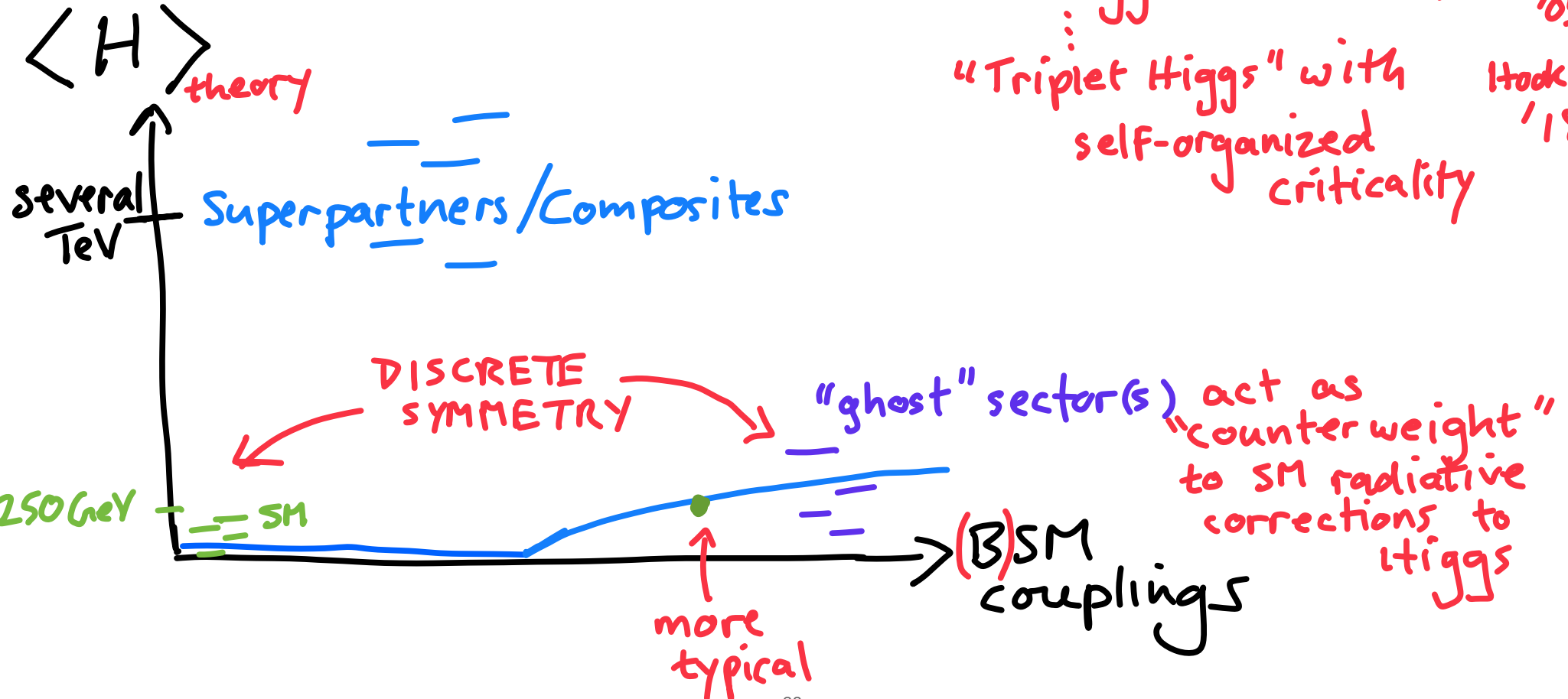
# THE (LITTLE) HIERARCHY & IMPORTANCE OF LIGHT VESTIGES



# NEUTRAL (HIDDEN) NATURALNESS

"Twin Higgs" Chacko, Groh, Harnik '05

"Triplet Higgs" with self-organized criticality Hodak '18



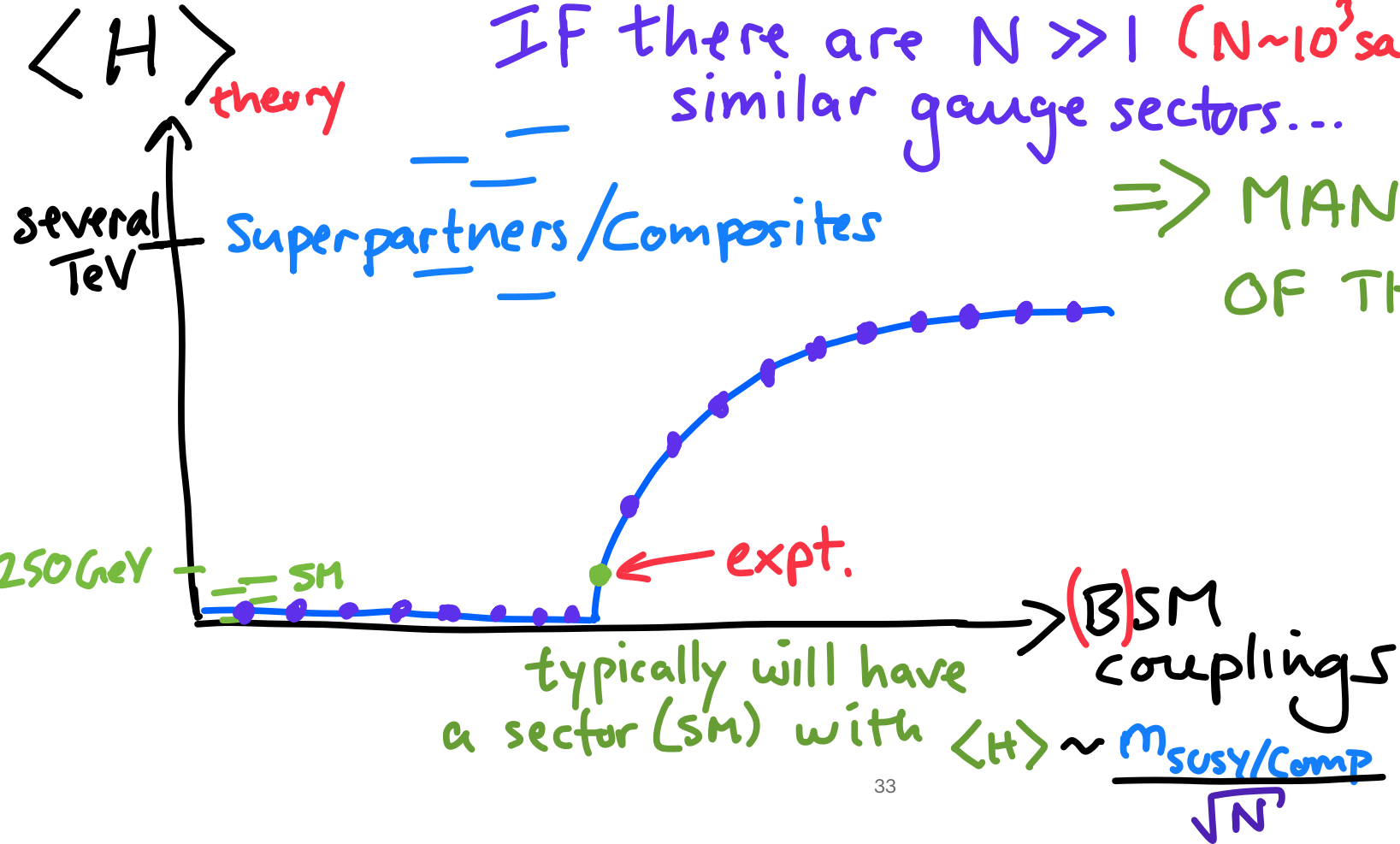


# Alternately, $N$ NATURALNESS

Arkani-Hamed, Cohen,  
D'Agnolo, Hook, Kim,  
Pinner '16

IF there are  $N \gg 1$  ( $N \sim 10^3$  say)  
similar gauge sectors...

$\Rightarrow$  MANY THROWS  
OF THE DICE!

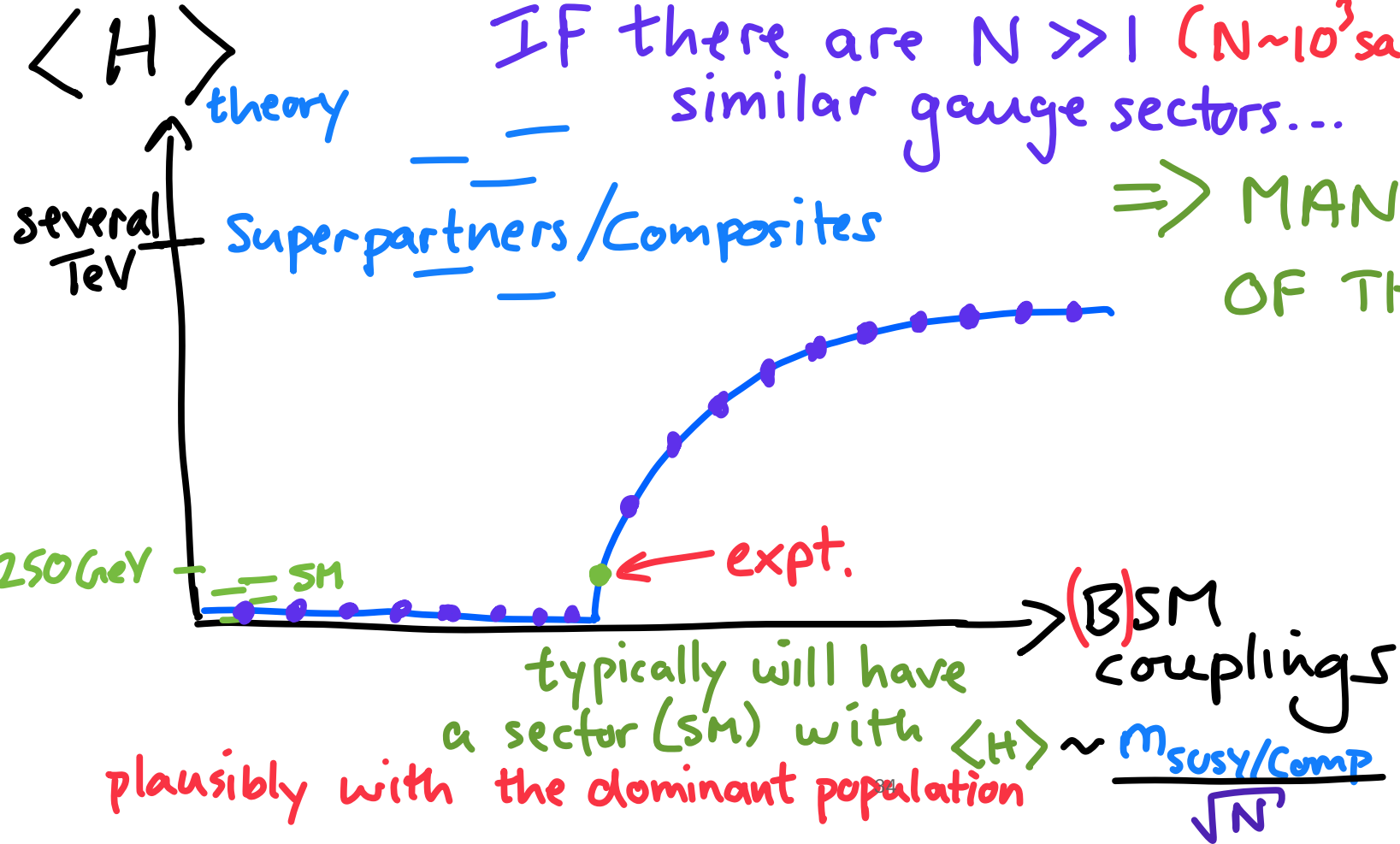


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# CONCLUSIONS

There are BROAD EXCITING PHYSICS THEMES to pursue at future colliders:

Dark Matter, Baryogenesis, SUSY, Compositeness, flavor origins, parallel gauge sectors, long-lived particles, precision Higgs structure

Need a collider at highest energies, clean enough & with sensitive enough detectors, to pursue both high mass &/or weakly coupled BSM at high precision & to excite & challenge next generation of experimentalists.

If new physics (dimly) seen in DM, flavor, EDM, precision, gravitational wave, cosmological expts., we need collider with reach/precision <sup>35</sup> to complement, corroborate, clarify