### **Di-Higgs Searches: Connections to Cosmology**

### M.J. Ramsey-Musolf

- T.D. Lee Institute/Shanghai Jiao Tong Univ.
- UMass Amherst
- Caltech

#### About MJRM:



Science



Family



Friends

My pronouns: he/him/his # MeToo Muon Collider Workshop, June 4 2021

## Key Ideas for this Talk

- Extensions of the Standard Model scalar sector can address key open questions in cosmology
- Di-Higgs searches provide one important window on the cosmological implications of extended scalar sectors
- This talk: focus on delineating the thermal history of EWSB and consequences for baryogenesis and gravitation wave searches
- There are exciting opportunities and synergies involving the LHC and prospective future colliders
   → how might a muon collider fit into this picture?

## **Outline**

- I. Cosmological Implications
- *II.* Was There an EW Phase Transition?
- III. Model Illustration: Real Singlet
- IV. Di-Higgs: Opportunities
- V. Outlook

## **Di-Higgs & Triscalar Interactions**

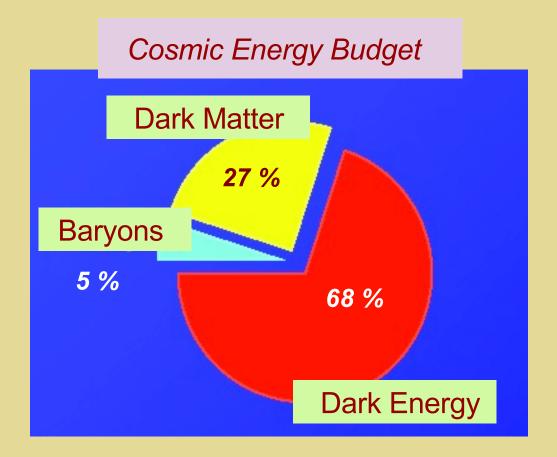
*h*<sub>1</sub> SM-like*h*<sub>2</sub> SM-like

$$\lambda_{ijk} h_i h_j h_k$$

$\lambda_{111}$	Non-resonant
$\lambda_{211}$	Resonant
λ <sub>122</sub>	Resonant – exotic decays & non-resonant

## I. Cosmological Implications

## The Origin of Matter



How can extended scalar sectors address this puzzle ?

### **Connections with Cosmology**

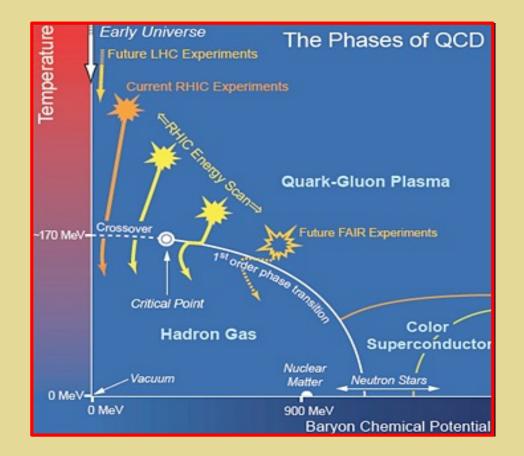
• Dark Matter: stable h<sub>2</sub>

• Thermal history of EWSB: Was there an electroweak phase transition ?

### **Electroweak Phase Transition**

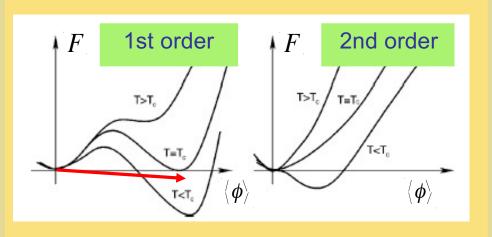
- Higgs discovery → What was the thermal history of EWSB ?
- Baryogenesis → Was the matter-antimatter asymmetry generated in conjunction with EWSB (EW baryogenesis) ?
- Gravitational waves → If a signal observed in next generation probes, could a cosmological phase transition be responsible ?

## **Thermal History of Symmetry Breaking**



QCD Phase Diagram  $\rightarrow$  EW Theory Analog?

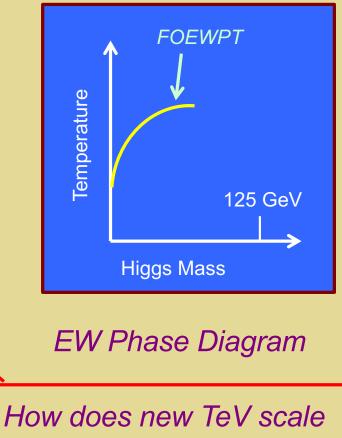
## **EWSB Transition: St'd Model**



Increasing m<sub>h</sub>

Lattice	Authors	$M_{\rm h}^C$ (GeV)
4D Isotropic	[76]	$80 \pm 7$
4D Anisotropic	[74]	$72.4\pm1.7$
3D Isotropic	[72]	$72.3 \pm 0.7$
3D Isotropic	[70]	$72.4\pm0.9$

SM EW: Cross over transition

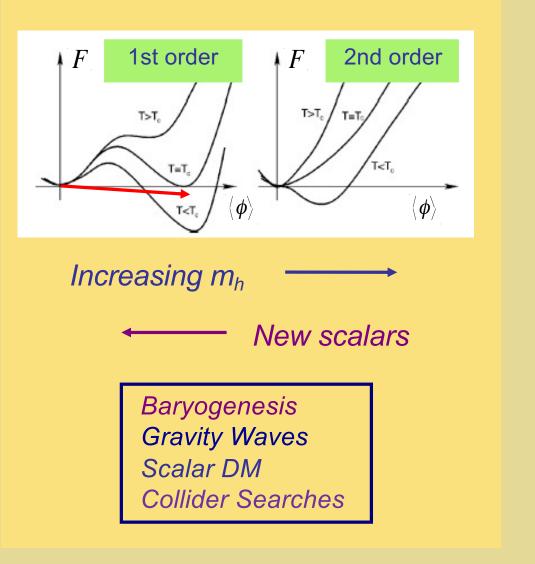


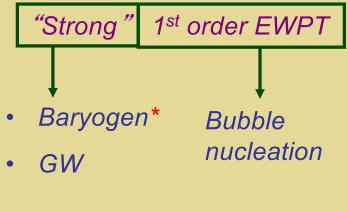
How does new TeV scale physics change this picture ? What is the phase diagram ? EWPT ? If so, what kind ?

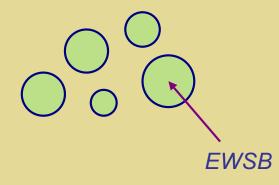
### **Electroweak Phase Transition**

- Higgs discovery → What was the thermal history of EWSB ?
- Baryogenesis → Was the matter-antimatter asymmetry generated in conjunction with EWSB (EW baryogenesis) ?
- Gravitational waves → If a signal observed in next generation probes, could a cosmological phase transition be responsible ?

## **EW Phase Transition: Baryogen & GW**



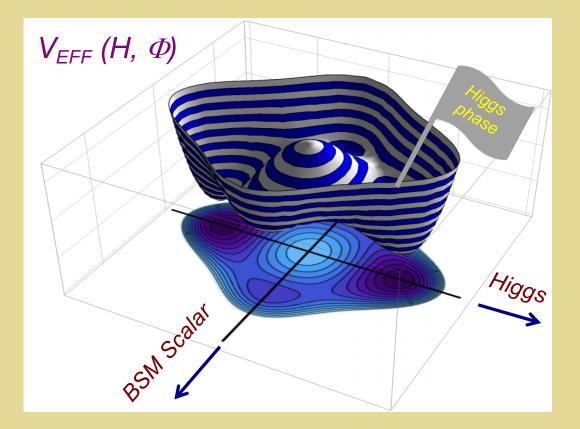




#### \* Need BSM CPV

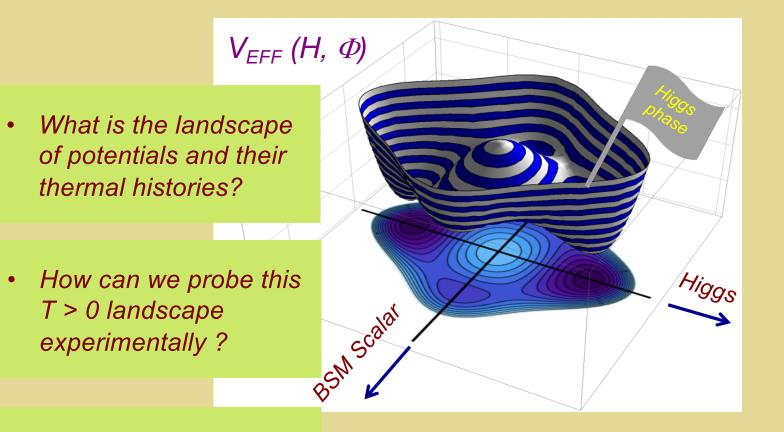
## **II. Was There an EW Phase Transition ?**

### **Thermal History of EWSB**



Extrema can evolve differently as T evolves → rich possibilities for symmetry breaking

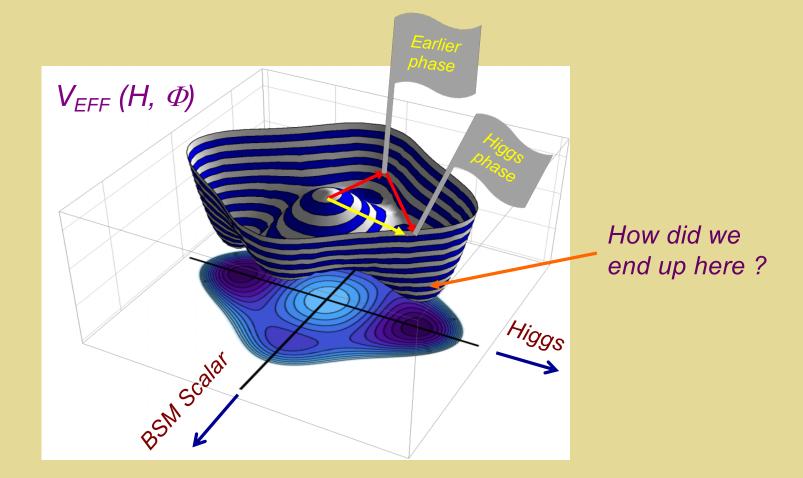
## **Thermal History of EWSB**



 How reliably can we compute the thermodynamics ?

n evolve differently as T evolves → ilities for symmetry breaking

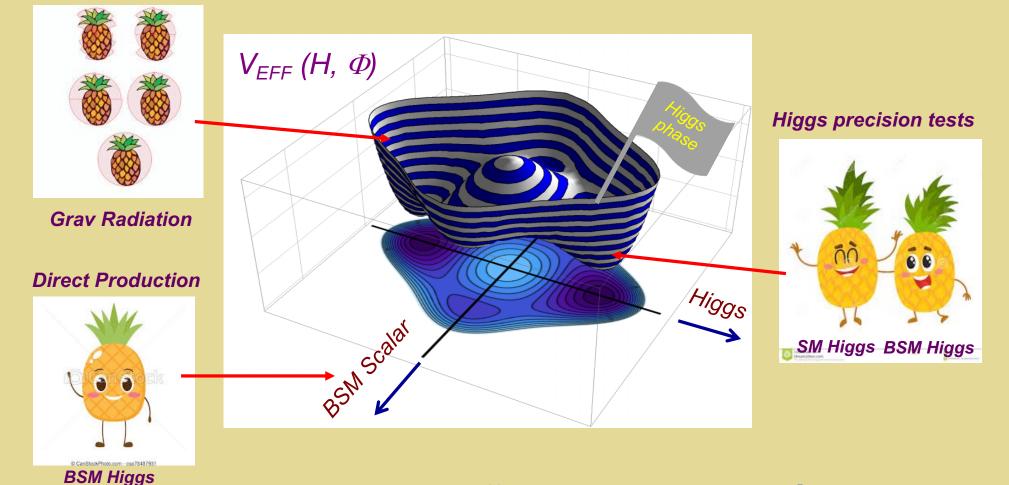
### **Patterns of Symmetry Breaking**



Extrema can evolve differently as T evolves → rich possibilities for symmetry breaking

## **Experimental Probes**

#### **Bubble Collisions**



Extrema can evolve differently as T evolves → rich possibilities for symmetry breaking

# **T<sub>EW</sub> Sets a Scale for Colliders**

### **High-T SM Effective Potential**

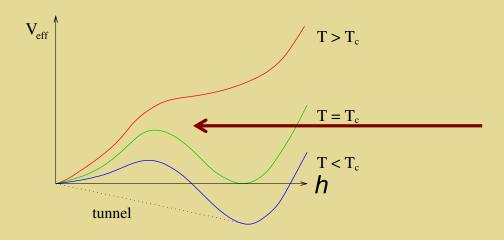
$$V(h,T)_{\rm SM} = D(T^2 - T_0^2) \, h^2 + \lambda \, h^4 \quad {\rm +} \ \cdots$$

$$T_0^2 = (8\lambda + \text{ loops}) \left( 4\lambda + \frac{3}{2}g^2 + \frac{1}{2}g'^2 + 2y_t^2 + \cdots \right)^{-1} v^2$$

$$T_0 \sim 140 \; \text{GeV} \equiv T_{EW}$$

18

# First Order EWPT from BSM Physics

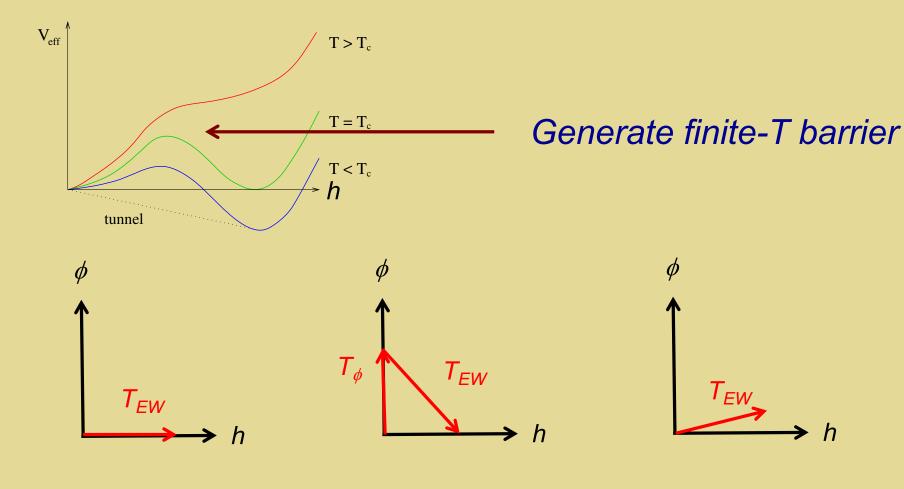


Generate finite-T barrier

Introduce new scalar  $\phi$ interaction with h via the Higgs Portal



# First Order EWPT from BSM Physics



 $a_2 H^2 \phi^2$  : T > 0loop effect

 $a_2 H^2 \phi^2$  : T = 0tree-level effect

 $a_1 H^2 \phi$  : T = 0tree-level effect

## **III. Model Illustrations**



Simple Higgs portal models:

- Real gauge singlet (SM + 1)
- Real EW triplet (SM + 3)

# **Real Singlet**

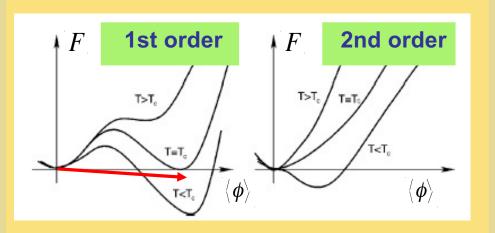
#### **Potential & conventions**

$$V = -\mu^{2} |H|^{2} + \lambda |H|^{4} + \frac{1}{2}a_{1} |H|^{2} S + \frac{1}{2}a_{2} |H|^{2} S^{2} + b_{1}S + \frac{1}{2}b_{2}S^{2} + \frac{1}{3}b_{3}S^{3} + \frac{1}{4}b_{4}S^{4},$$

$$h_{1} = h\cos\theta + s\sin\theta$$

$$h_{2} = -h\sin\theta + s\cos\theta,$$
• Profumo, RM, Shaugnessy:  $h_{1} = SM$ -like  
• Kozaczuk, RM, Shelton:  $h_{1} = lightest$ 

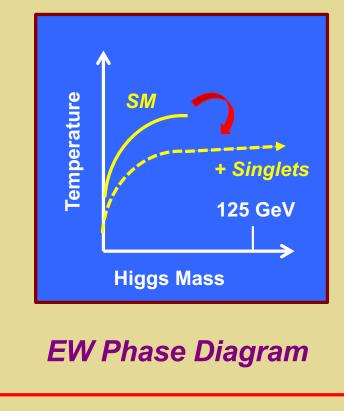
## **EW Phase Transition: Singlet Scalars**



Increasing m<sub>h</sub>

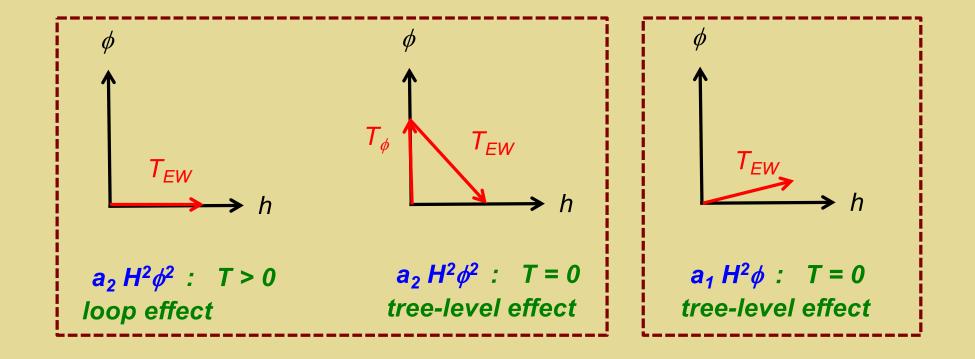
Lattice	Authors	$M_{\rm h}^C$ (GeV)
4D Isotropic	[76]	$80\pm7$
4D Anisotropic	[74]	$72.4 \pm 1.7$
3D Isotropic	[72]	$72.3 \pm 0.7$
3D Isotropic	[70]	$72.4 \pm 0.9$

SM EW: Cross over transition



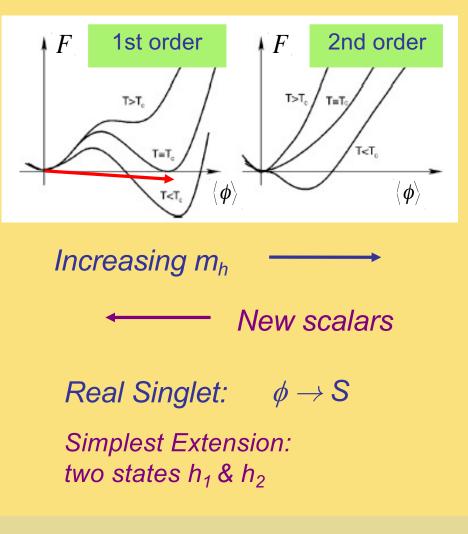
How does this picture change in presence of new TeV scale physics ? What is the phase diagram ?

## **Real Singlet**

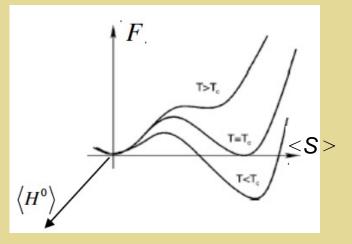


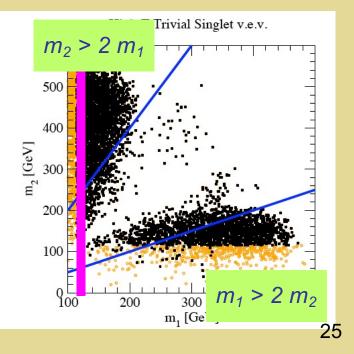
No Z<sub>2</sub> breaking at T = 0 required Z<sub>2</sub> breaking at T = 0 (explicit or spontaneous)

## **EW Phase Transition: New Scalars**

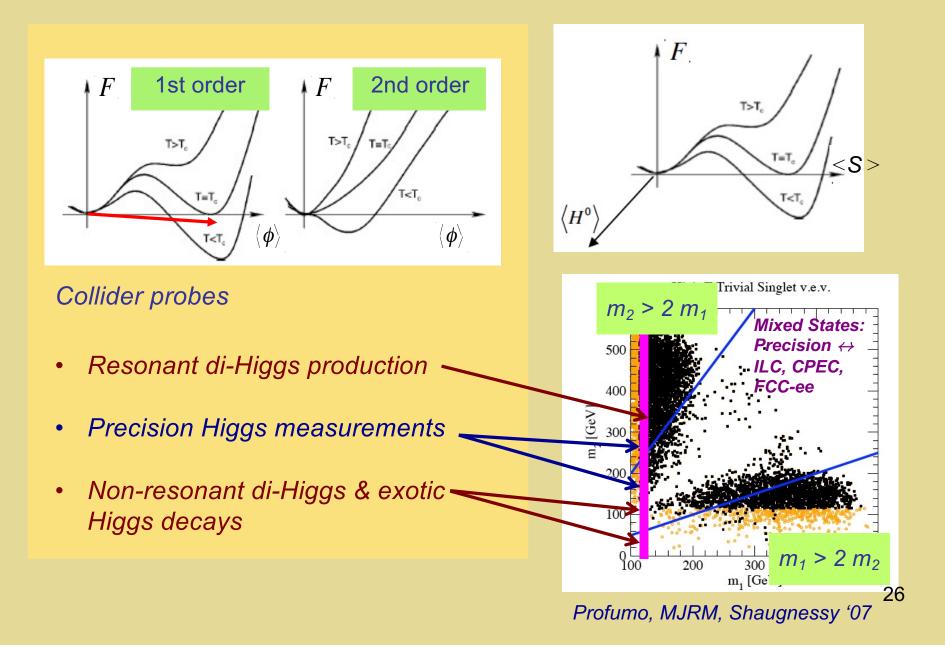


Profumo, R-M, Shaugnessy JHEP 0708 (2007) 010 Espinosa, Konstandin, Riva NPB 854 (2012) 592





## **EW Phase Transition: Singlet Scalars**



$$\lambda_{211} = \frac{1}{4} [(a_1 + 2a_2x_0)\cos^3\theta + 4v_0(a_2 - 3\lambda)\cos^2\theta\sin\theta + (a_1 + 2a_2x_0 - 2b_3 - 6b_4x_0)\cos\theta\sin^2\theta - 2a_2v_0\sin^3\theta]$$
(12)

 $h_2 \rightarrow h_1 h_1$ 

 $g_{111} = \lambda v_0 \cos^3 \theta + \frac{1}{4} (a_1 + 2a_2 x_0) \cos^2 \theta \sin \theta + \frac{1}{2} a_2 v_0 \cos \theta \sin^2 \theta + \frac{b_3}{3} \sin^3 \theta + b_4 x_0 \sin^3 \theta \quad .$ 

 $h_1 h_1 h_1$ 

**Exotic decays**  $h_1 \rightarrow h_2 h_2 _{27}$ 

$$g_{122} = \frac{v_0 c_\theta}{2} \left( a_2 (c_\theta^2 - 2s_\theta^2) + 6\lambda s_\theta^2 \right) + \frac{a_1 + 2a_2 x_0}{2} \left( s_\theta^3 - c_\theta^2 s_\theta \right) + (b_3 + 3b_4 x_0) c_\theta^2 s_\theta, \qquad [s_\theta \equiv \sin \theta, \ c_\theta \equiv \cos \theta].$$

Insensitive to  $\theta$  and  $x_0$ 

$$\lambda_{211} = \frac{1}{4} [(a_1 + 2a_2x_0)\cos^3\theta + 4v_0(a_2 - 3\lambda)\cos^2\theta\sin\theta + (a_1 + 2a_2x_0 - 2b_3 - 6b_4x_0)\cos\theta\sin^2\theta - 2a_2v_0\sin^3\theta]$$
(12)

 $h_2 \rightarrow h_1 h_1$ 

$$g_{111} = \lambda v_0 \cos^3 \theta + \frac{1}{4} (a_1 + 2a_2 x_0) \cos^2 \theta \sin \theta + \frac{1}{2} a_2 v_0 \cos \theta \sin^2 \theta + \frac{b_3}{3} \sin^3 \theta + b_4 x_0 \sin^3 \theta .$$

 $h_1 h_1 h_1$ 

**Exotic decays**  $h_1 \rightarrow h_2 h_2 _{28}$ 

$$g_{122} = \frac{v_0 c_{\theta}}{2} \left( a_2 (c_{\theta}^2 - 2s_{\theta}^2) + 6\lambda s_{\theta}^2 \right) + \frac{a_1 + 2a_2 x_0}{2} \left( s_{\theta}^3 - c_{\theta}^2 s_{\theta} \right) + (b_3 + 3b_4 x_0) c_{\theta}^2 s_{\theta}, \qquad [s_{\theta} \equiv \sin \theta, \ c_{\theta} \equiv \cos \theta].$$

Same combination

 $\theta$  suppression

 $\lambda_{211} = \frac{1}{4} [(a_1 + 2a_2x_0)\cos^3\theta + 4v_0(a_2 - 3\lambda)\cos^2\theta\sin\theta + (a_1 + 2a_2x_0 - 2b_3 - 6b_4x_0)\cos\theta\sin^2\theta - 2a_2v_0\sin^3\theta]$ (12)

**Resonant di-Higgs** 

 $h_2 \rightarrow h_1 h_1$ 

 $g_{111} = \lambda v_0 \cos^3 \theta + \frac{1}{4} (a_1 + 2a_2 x_0) \cos^2 \theta \sin \theta + \frac{1}{2} a_2 v_0 \cos \theta \sin^2 \theta + \frac{b_3}{3} \sin^3 \theta + b_4 x_0 \sin^3 \theta \quad .$ 

Higgs self-coupling

 $h_1 h_1 h_1$ 

**Exotic decays & non-res di-Higgs**  $h_1 \rightarrow h_2 h_2$  29

$$g_{122} = \frac{v_0 \ c_\theta}{2} \left( a_2 (c_\theta^2 - 2s_\theta^2) + 6\lambda s_\theta^2 \right) + \frac{a_1 + 2a_2 x_0}{2} \left( s_\theta^3 - c_\theta^2 s_\theta \right) + (b_3 + 3b_4 x_0) \ c_\theta^2 s_\theta, \qquad [s_\theta \equiv \sin \theta, \ c_\theta \equiv \cos \theta].$$

Portal coupling sensitivity without  $\theta$  suppression

$$\lambda_{211} = \frac{1}{4} [(a_1 + 2a_2x_0)\cos^3\theta + 4v_0(a_2 - 3\lambda)\cos^2\theta\sin\theta + (a_1 + 2a_2x_0 - 2b_3 - 6b_4x_0)\cos\theta\sin^2\theta - 2a_2v_0\sin^3\theta]$$
(12)

 $h_2 \rightarrow h_1 h_1$ 

 $g_{111}=\lambda v_0\cos^3\theta+\frac{1}{4}(a_1+2a_2x_0)\cos^2\theta\sin\theta$  $+\frac{1}{2}a_2v_0\cos\theta\sin^2\theta + \frac{b_3}{3}\sin^3\theta + b_4x_0\sin^3\theta$ .

Higgs self-coupling

**Exotic decays &** 

non-res di-Higgs

30

 $h_1 \rightarrow h_2 h_2$ 

 $h_1 h_1 h_1$ 

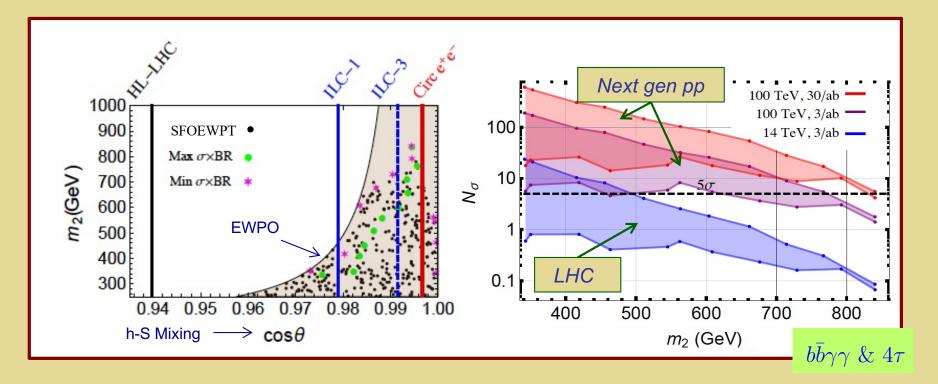
$$g_{122} = \frac{v_0 \ c_\theta}{2} \left( a_2 (c_\theta^2 - 2s_\theta^2) + 6\lambda s_\theta^2 \right) + \frac{a_1 + 2a_2 x_0}{2} \left( s_\theta^3 - c_\theta^2 s_\theta \right) + (b_3 + 3b_4 x_0) \ c_\theta^2 s_\theta, \qquad [s_\theta \equiv \sin \theta, \ c_\theta \equiv \cos \theta].$$

## **IV. Di-Higgs: Opportunities**

Apologies to all whose work I cannot cover here !

## Singlets: Precision & Res Di-Higgs Prod

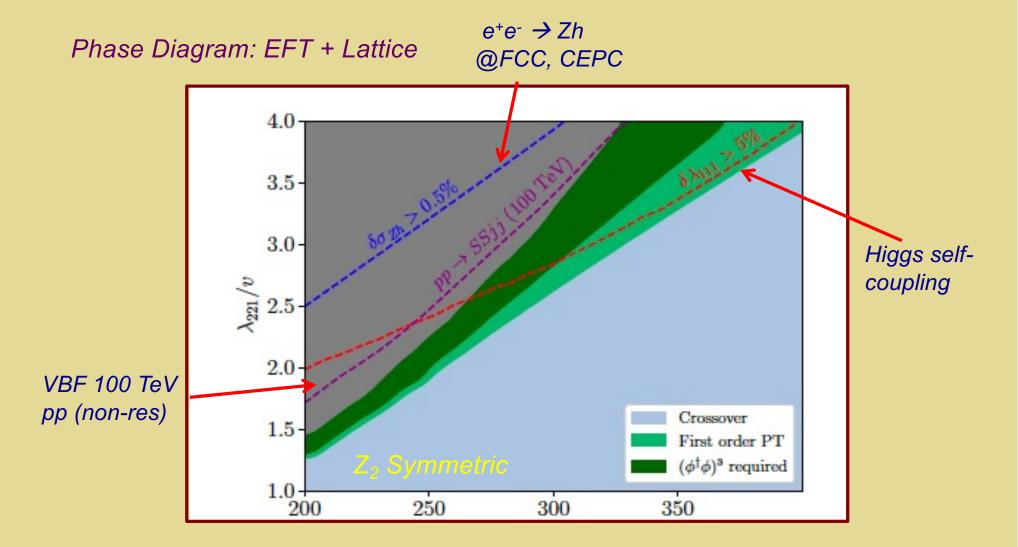
SFOEWPT Benchmarks: Resonant di-Higgs & precision Higgs studies



Kotwal, No, R-M, Winslow 1605.06123

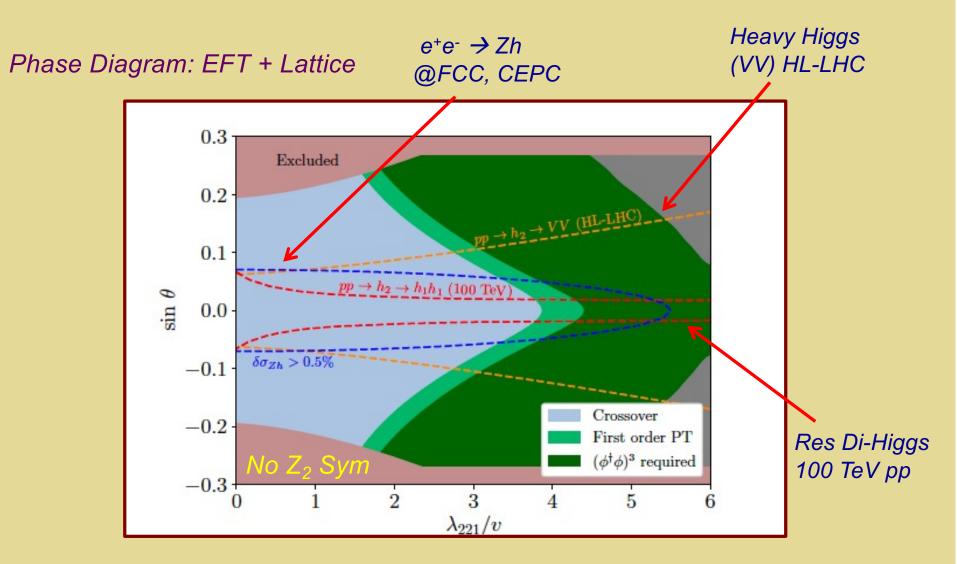
See also: Huang et al, 1701.04442; Li et al, 1906.05289

## Singlets: Non-Resonant Di-Higgs Prod



O. Gould, J. Kozaczuk, L. Niemi, MJRM, TVI Tenkanen, D.J. Weir, 1903.11604

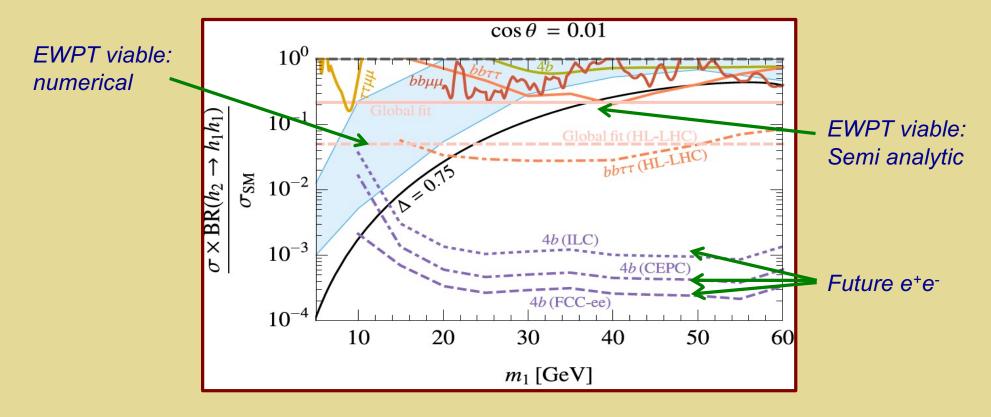
## Singlets: Non-Resonant Di-Higgs Prod



O. Gould, J. Kozaczuk, L. Niemi, MJRM, TVI Tenkanen, D.J. Weir, 1903.11604

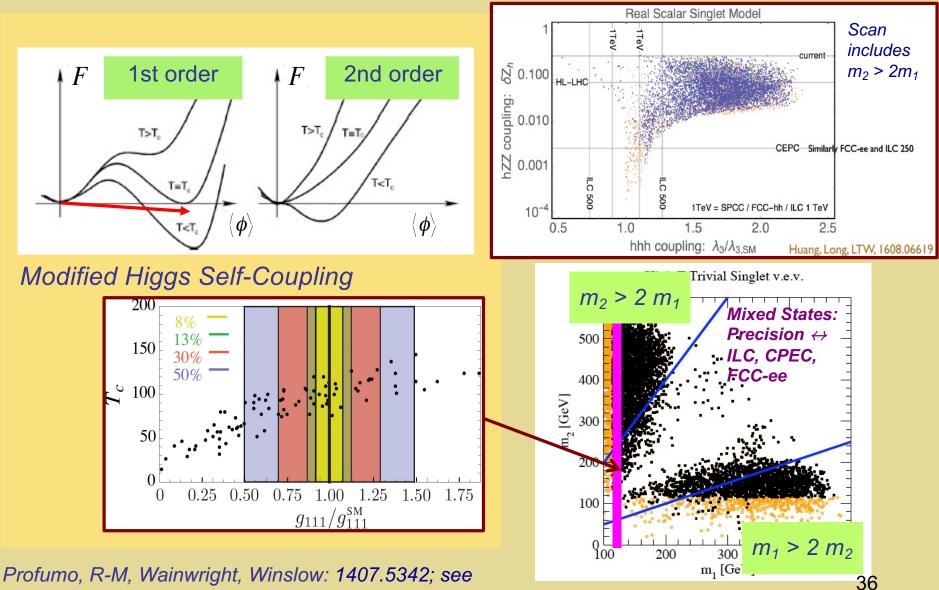
## Light Singlets: Exotic Decays

#### $h_2 \rightarrow h_1 h_1 \rightarrow 4b$



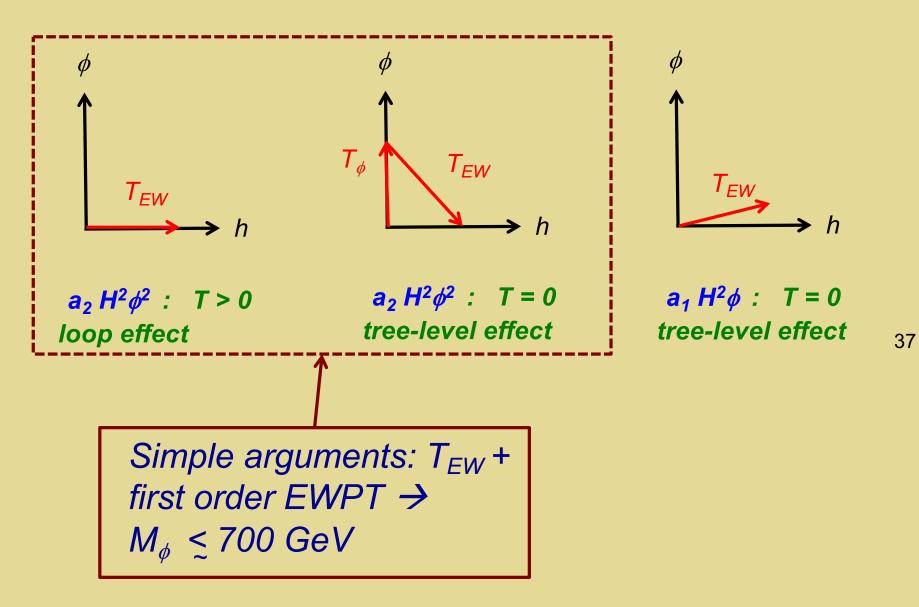
J. Kozaczuk, MR-M, J. Shelton 1911.10210 See also: Carena et al 1911.10206

## Singlets: Triple Self-Coupling

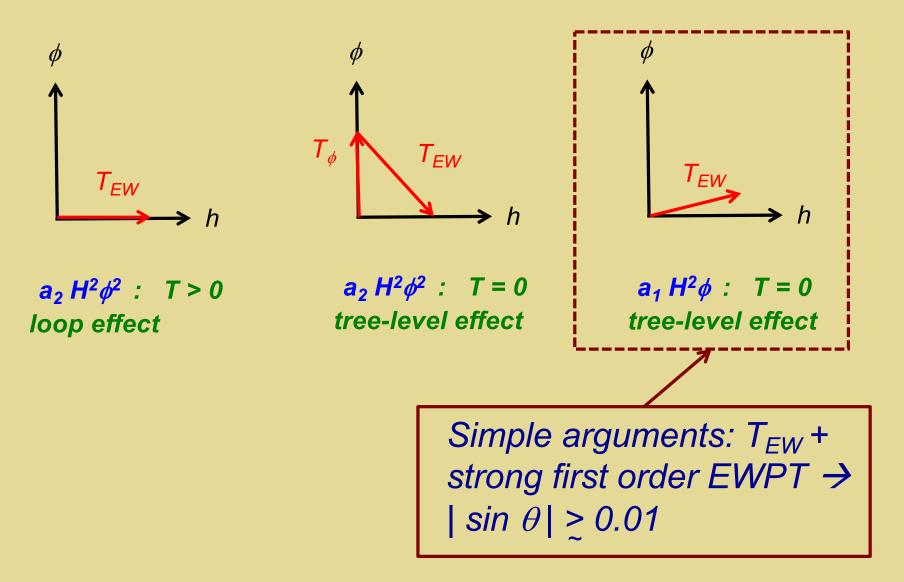


also Noble & Perelstein 0711.3018

### **Muon Collider: Comments**



### **Muon Collider: Comments**



## V. Outlook

- Extensions of the Standard Model scalar sector can address key open questions in cosmology
- Di-Higgs searches provide one important window on the cosmological implications of extended scalar sectors
- This talk: focus on delineating the thermal history of EWSB and consequences for baryogenesis and gravitation wave searches
- There are exciting opportunities and synergies involving the LHC and prospective future colliders
   → how might a muon collider fit into this picture?

# **Back Up Slides**

## References

- EWPT & Colliders General: MJRM 1912.07189
- EWPT & Di-Higgs:
  - Profumo, MJRM, Shaugnessy 0705.2425
  - No & MJRM 1310.6035
  - Kotwal, No, MJRM, Winslow 1605.06123
  - Huang, Pernie, MJRM, Safanov, Spannowsky, Winslow 1701.04442
  - *Li, MJRM, Willocq* 1906.05289
  - Papaefstathiou and White, 2010.00597
  - Ren et al 1706.05980

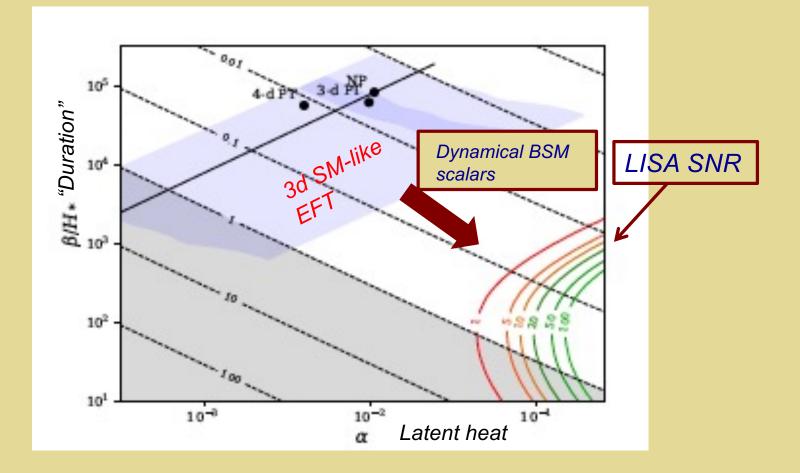
### Heavy Real Singlet: EWPT & GW

Non-dynamical heavy BSM scalars 4-d PT 3-d PT 105 B/H . "Duration" 3d SM-like LISA SNR 104 103  $10^{2}$ 10<sup>1</sup> 10-10-0  $10^{-2}$ Latent heat α

Gould, Kozaczuk, Niemi, R-M, Tenkanen, Weir 1903.11604

One-stepNon-perturbative

### Heavy Real Singlet: EWPT & GW



Gould, Kozaczuk, Niemi, R-M, Tenkanen, Weir 1903.11604

One-step

• Non-perturbative