

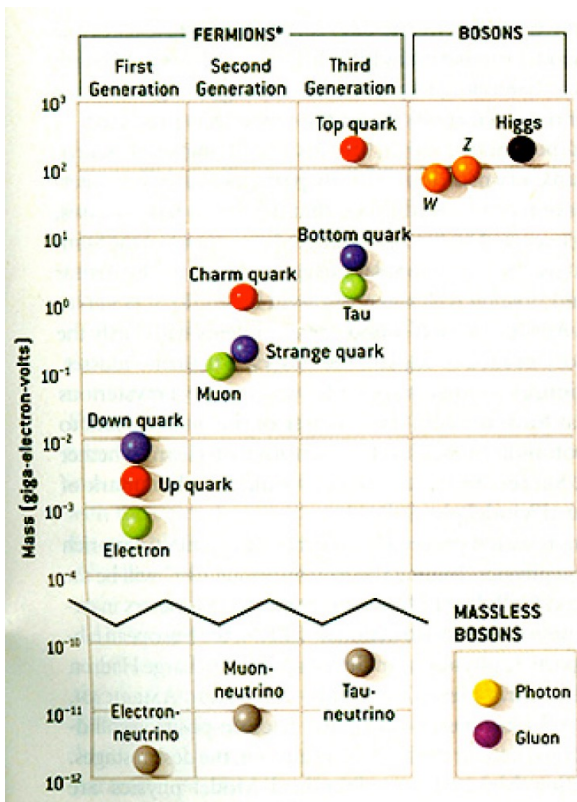
Shaping the future of High-Energy Physics (Experiment)

*Paris Sphicas
CERN & NKUA
JENAS 2022
Madrid, May 3-6, 2022*

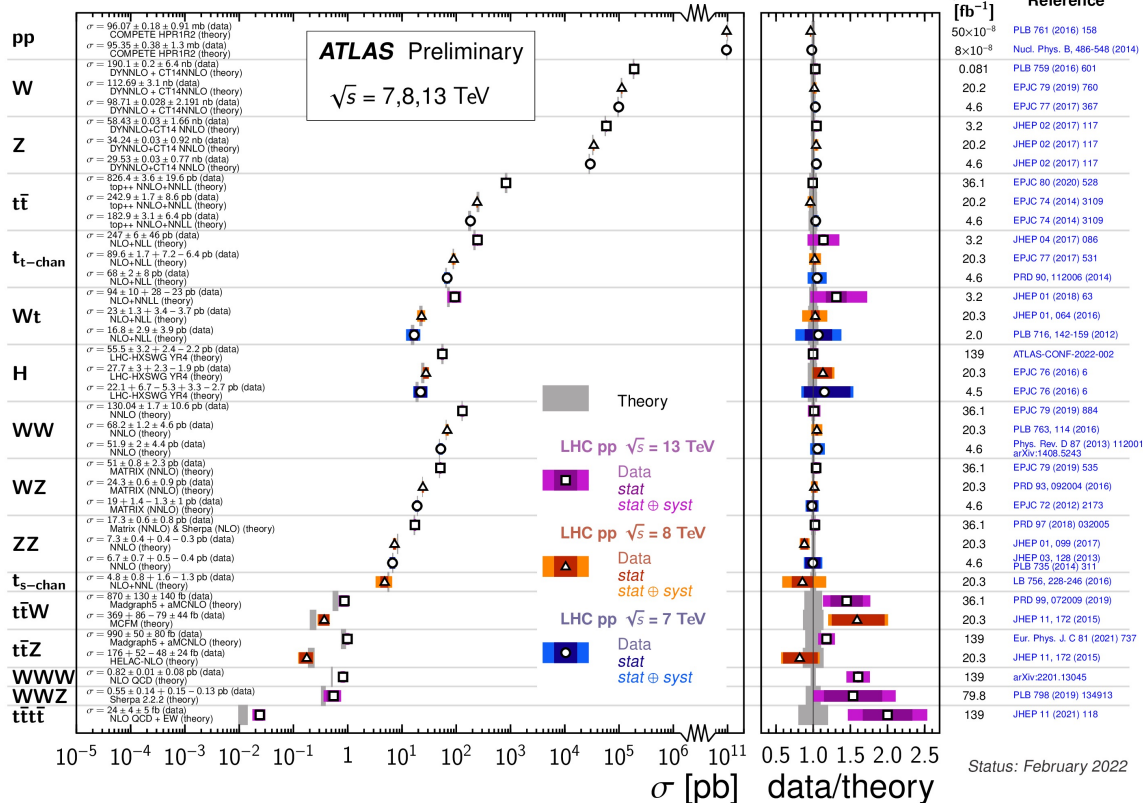
- **Introduction**
 - **Where we stand & the road to obtaining answers**
- **The (fundamental?) scalar sector of the SM and direct search for BSM physics**
 - **Couplings as per SM? Self-coupling? Fundamental or composite?**
 - **Direct searches for new interactions or new particles around or above the electroweak scale.**
- **Dark Matter**
 - **Interplay of Direct and Indirect (collider) searches**
- **Physics of Flavor**
 - **Quark sector: CP violation, rare decays of K and B mesons**
 - **Neutrinos: New source of CP violation(?) Mass ordering? Nature? New particles/new interactions?**
 - **Other high-precision measurements (e.g. $g-2$); tests of fundamental symmetries**
- **Pseudo-summary**

Where we stand

- Most successful Theory ever: Standard Model (Theory) of Particle Physics

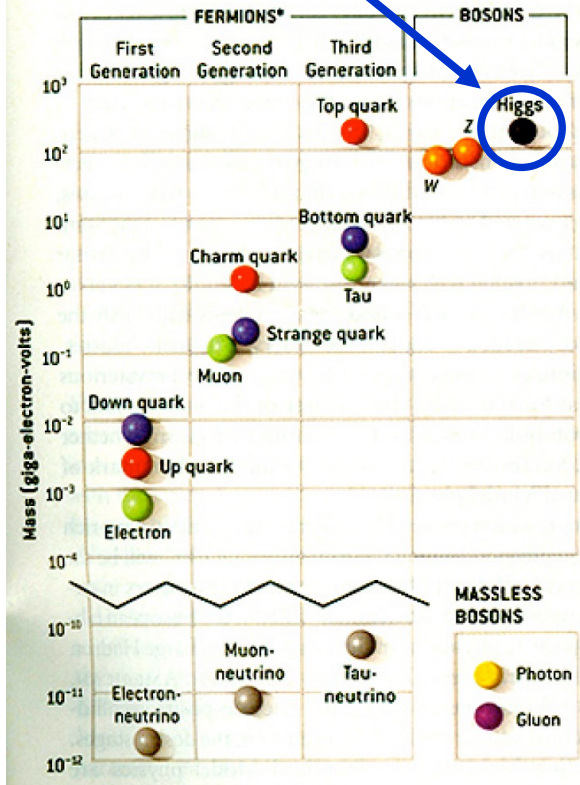


Standard Model Total Production Cross Section Measurements

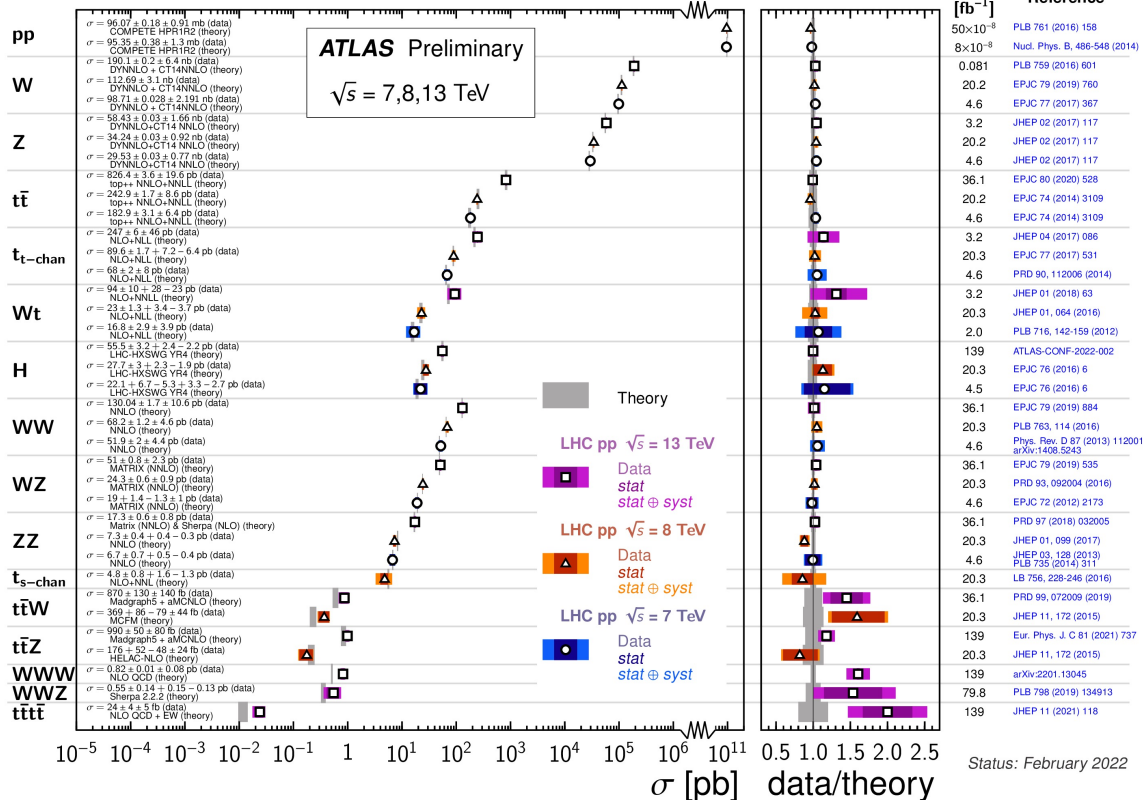


Where we stand

- Most successful Theory ever: Standard Model (Theory) of Particle Physics
 - Highest priority: *extend understanding of SM and its newly discovered scalar sector.*



Standard Model Total Production Cross Section Measurements

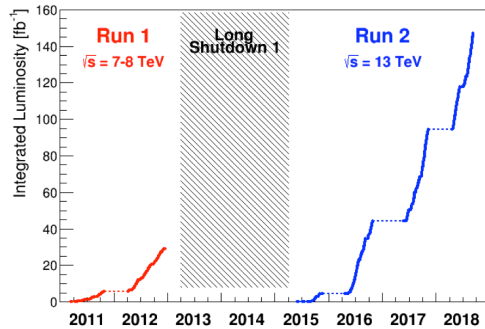


Status: February 2022

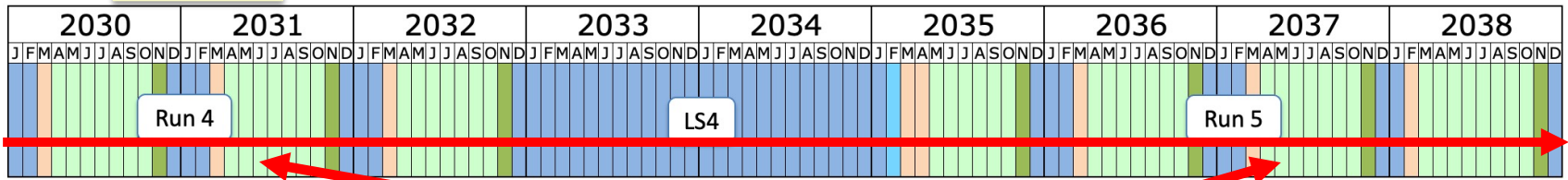
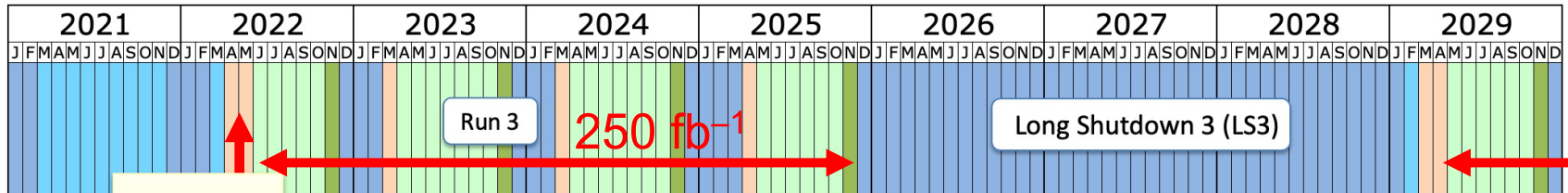
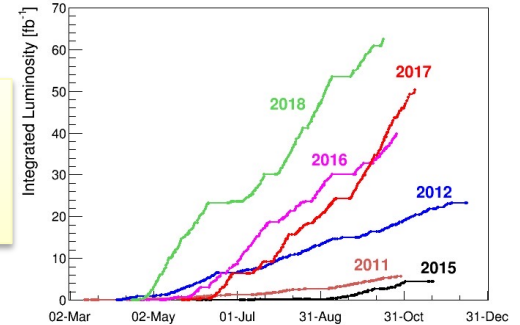
The road to obtaining answers: measurements...

- **Physics of fundamental scalars and direct searches for New Physics**
 - LHC, HL-LHC & future collider(s) (FCC-ee, FCC-hh, FCC-eh, CEPC, ILC, CLIC)
 - Understand the strangest of all elements of the SM.
 - Search for new particles and physics phenomena.
- **Dark Matter searches**
 - Very large number of experiments aiming at detecting DM
 - Direct detection, indirect detection (→ talk by S. DeJongh);
 - Matter-matter collisions (→ this talk)
- **Physics of Flavor:**
 - Neutrinos: Complementarity: beams [DUNE (US), HyperK (JP)], reactors (JUNO), atmosph (ORCA)
 - Is there CP violation in the lepton sector? PMNS matrix (à la CKM)?
 - Mass hierarchy of three ν generations (“normal” or “inverted”)? Dirac or Majorana?
 - Quark sector: Kaon and B hadron decays (LHC, HL-LHC, Fixed-Target (CERN, JP), Belle II (JP))
 - Measure CP sources as precisely as possible. Probe sacred (but also accidental...) laws (e.g. lepton number conservation). Universality? Probe rarest decay modes.
 - High-precision measurements; tests of Fundamental Symmetries

LHC & HL-LHC schedule



Run 1: 2010-2012; total $L \sim 30 \text{ fb}^{-1}$
 Run 2: 2015-2018 (150 fb^{-1})
 Peak lumi $\approx 2.1 \times 10^{34} \text{ cm}^{-2}\text{s}^{-1}$ (2018)



Last updated: January 2022

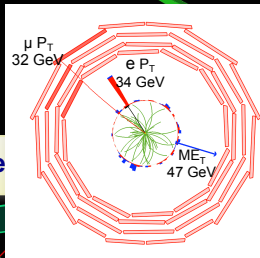
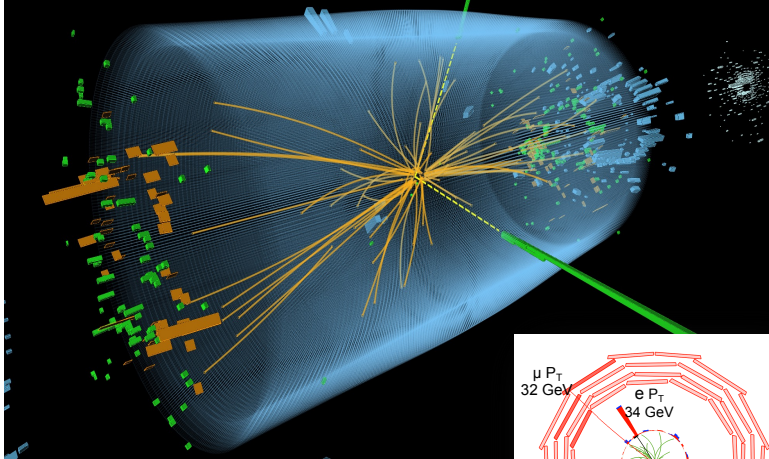
- Shutdown/Technical stop
- Protons physics
- Ions
- Commissioning with beam
- Hardware commissioning/magnet training

HL-LHC 3000 fb^{-1}

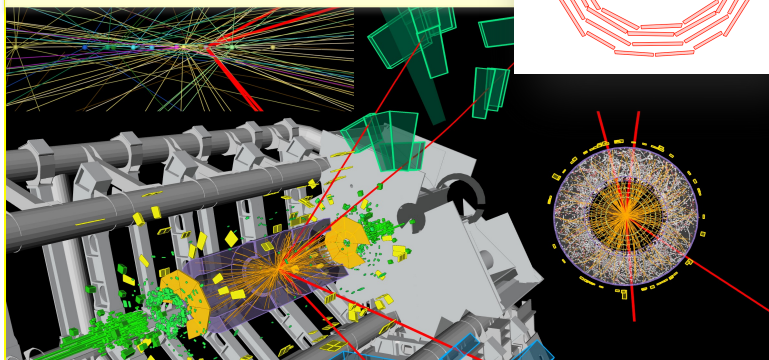


CMS Experiment at the LHC, CERN
 Data recorded: 2012-May-13 20:08:14.621490 GMT
 Run/Event: 134108 / 564224000

$H \rightarrow \gamma\gamma$

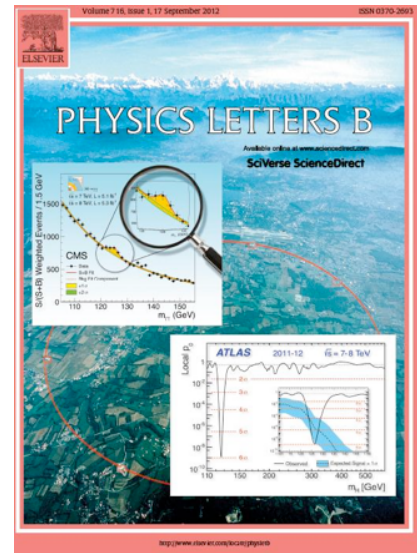
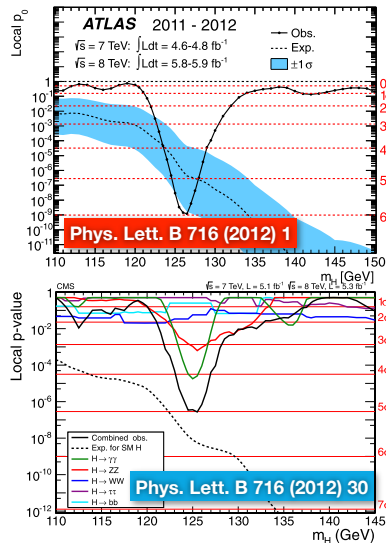
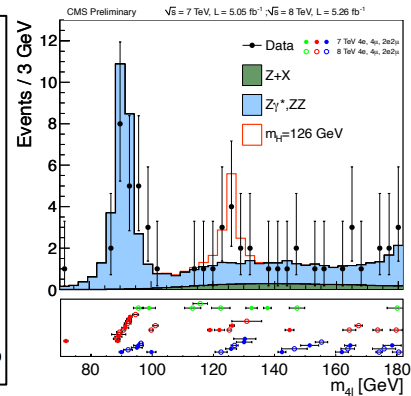
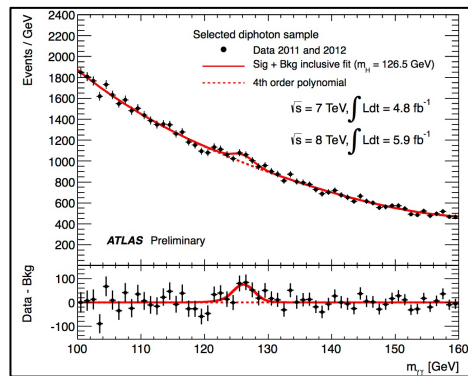


$p_T(\mu) = 36, 48, 26, 72 \text{ GeV}; m_{12} = 86.3 \text{ GeV}$



$H \rightarrow ZZ \rightarrow 4\mu$
 $m_{4\mu} = 125.1 \text{ GeV}$

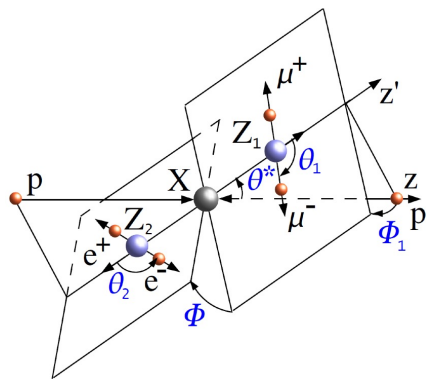
Recall... new particle @ $m \approx 125 \text{ GeV}$ $s=0$ or 2 boson (decays to $\gamma\gamma, ZZ$)



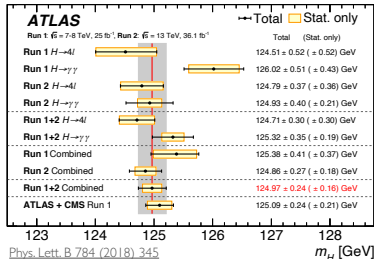
Higgs Physics: after the discovery

Measurement of Spin-Parity, mass: it IS the vacuum particle...

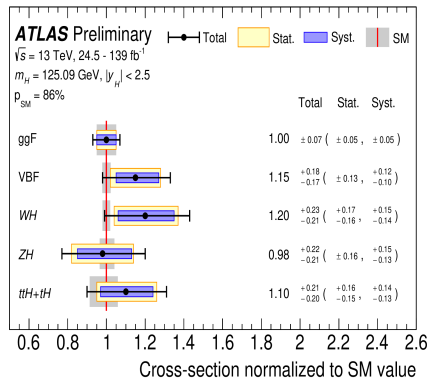
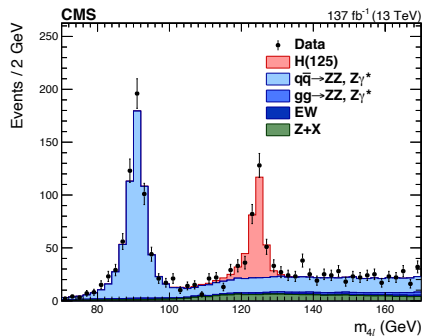
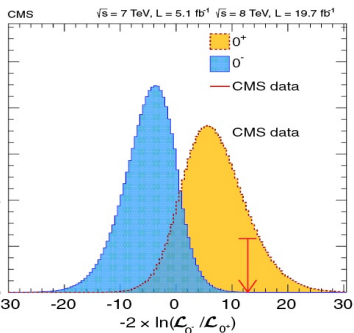
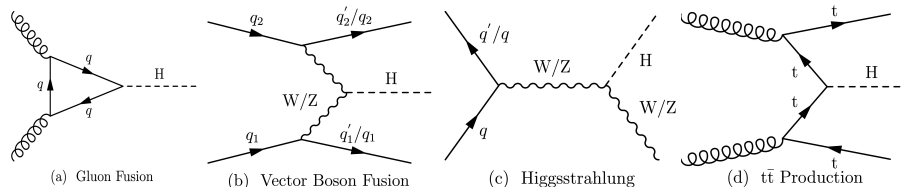
Production mechanisms, differential distributions



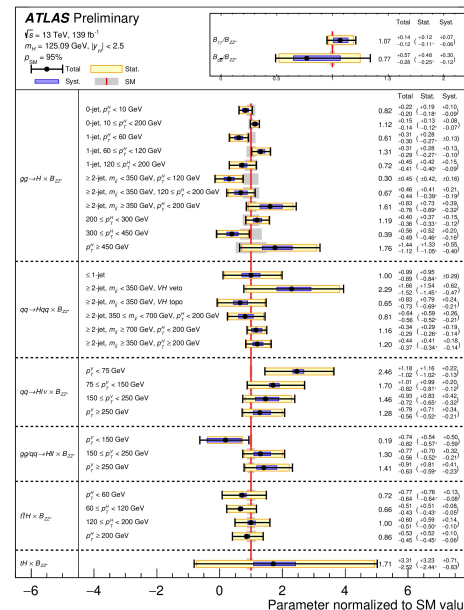
$$\delta M_H / M_H \sim 0.2\% (!)$$



$J^P = 0^+ (!!!)$



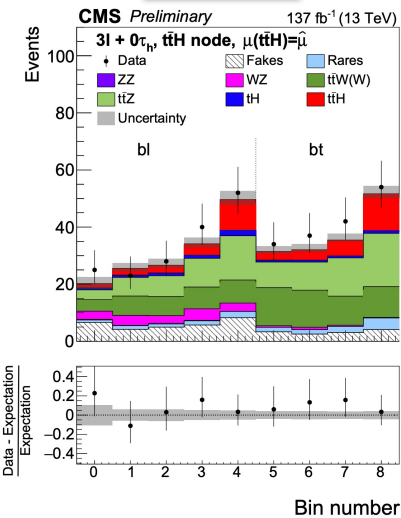
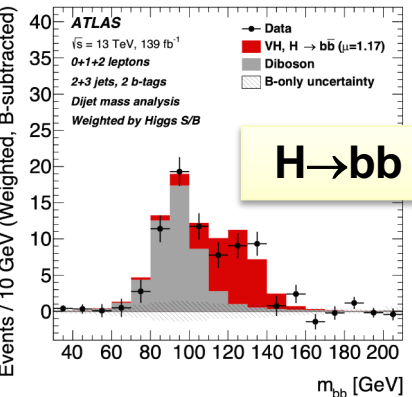
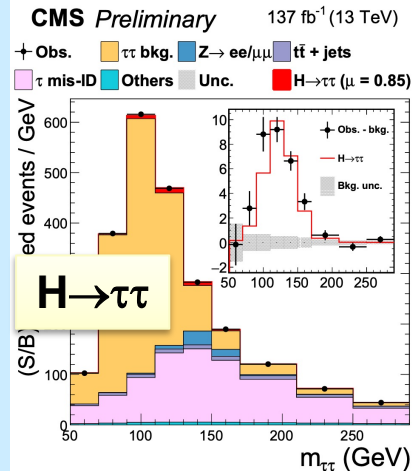
$$\mu = 1.06 \pm 0.07$$



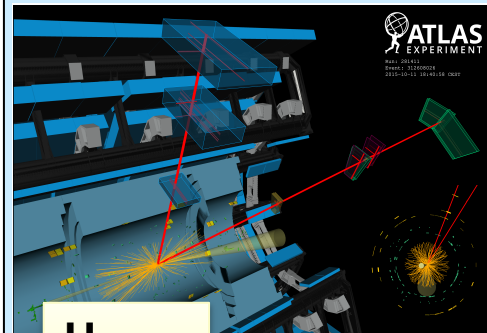
SM @ the highest E; EWSB (“Higgs” sector) (I)

Clear observation of couplings to 3rd-gen fermions

ttH



Evidence of coupling to 2nd-gen fermions



H $\rightarrow \mu\mu$

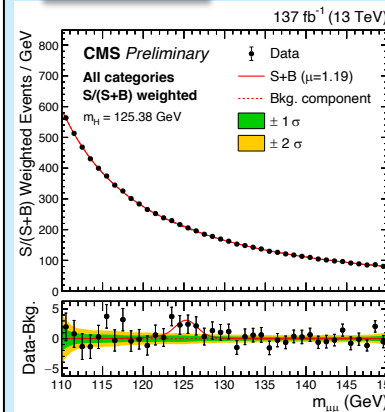
CMS:

obs: 3.0 σ (exp: 2.5 σ)

ATLAS:

obs: 2.0 σ (exp: 1.7 σ)

$$\mu = 1.2 \pm 0.6$$



H $\rightarrow cc$

$$\mu_{VZ(Z \rightarrow cc)} = 1.01^{+0.23}_{-0.21}$$

$$\mu_{VH(H \rightarrow cc)} < 14 \text{ (7.6 exp.)}$$

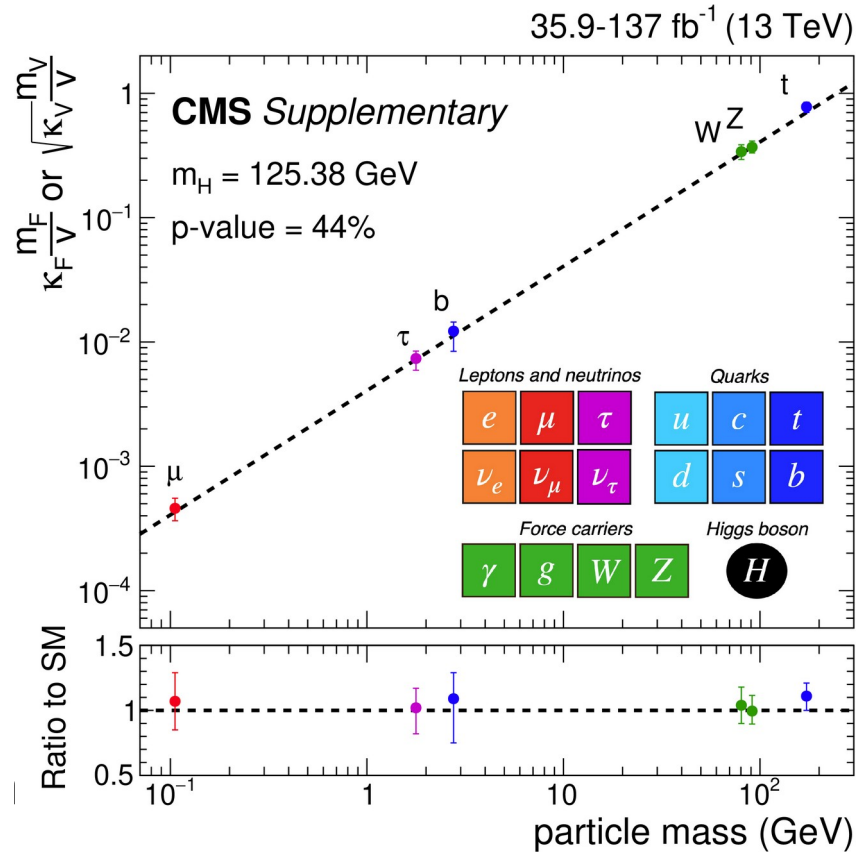
$$1.1 < |\kappa_c| < 5.5 \text{ (} |\kappa_c| < 3.4 \text{ exp.)}$$

The overall picture of the measured/seen Higgs couplings

$$\lambda_f = \kappa_f \left(\frac{m_f}{v} \right)$$

$$\left(\frac{g_V}{2v} \right)^{1/2} = \kappa_v^{1/2} \left(\frac{m_V}{v} \right)$$

A new kind of “force”,
with non-universal
coupling



Putting it all together: SM reigns supreme

Goodness of fit

$\chi^2_{\min} = 18.6 \rightarrow \text{Prob} = 23\%$

Fit result often more accurate than measurement

Small pulls for M_H , M_Z , $\Delta\alpha_{\text{had}}^{(5)}(M_Z^2)$, m_c , $m_b \rightarrow$ input accuracies exceed fit requirements

Knowledge of $m_H \rightarrow$ huge improvement in:

m_W (28 \rightarrow 11 MeV) arXiv:1407.3792

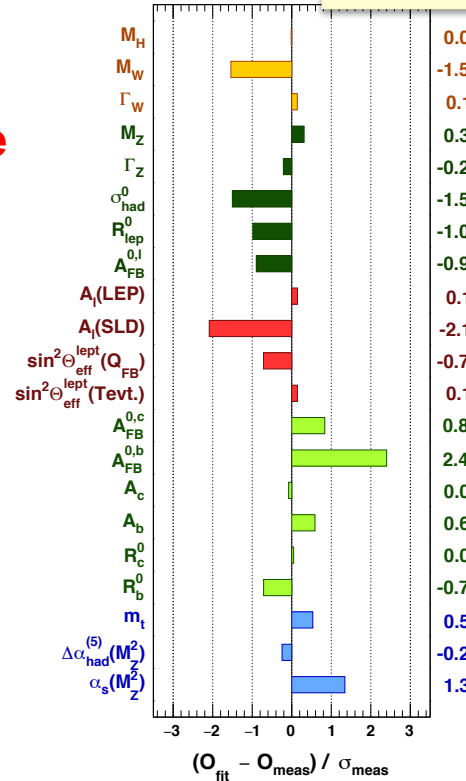
m_t (6.2 \rightarrow 2.5 GeV)

$\sin^2\theta_W$ (2.3 \rightarrow 1.0×10^{-3})

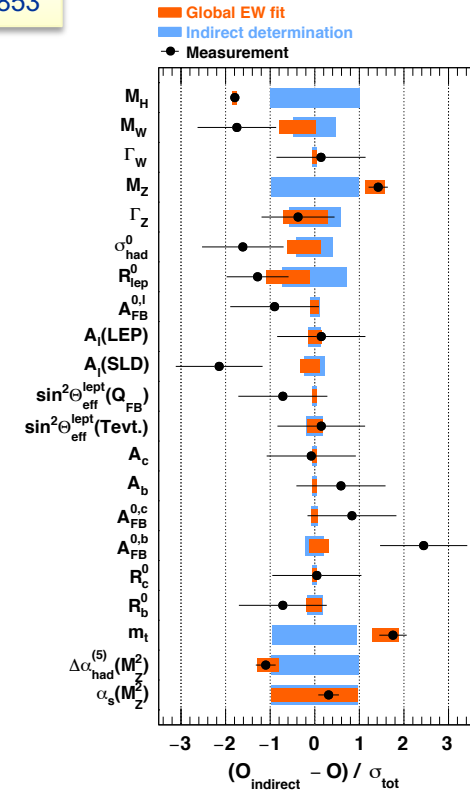
Largest discrepancy:

$A_{\text{FB}}(b)$: 2.5σ

arXiv:1803.01853

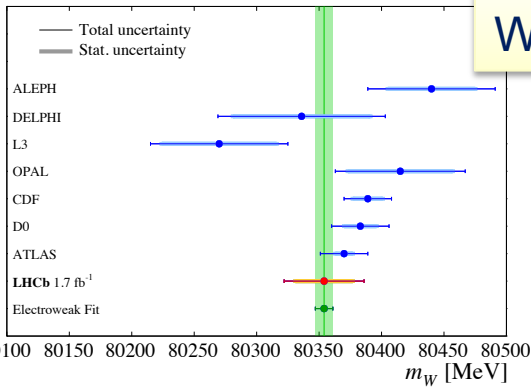


<https://project-gfitter.web.cern.ch/>

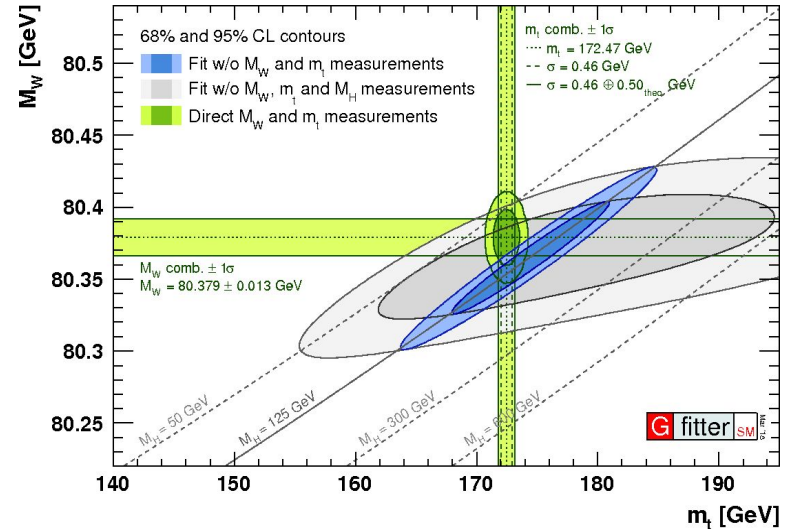
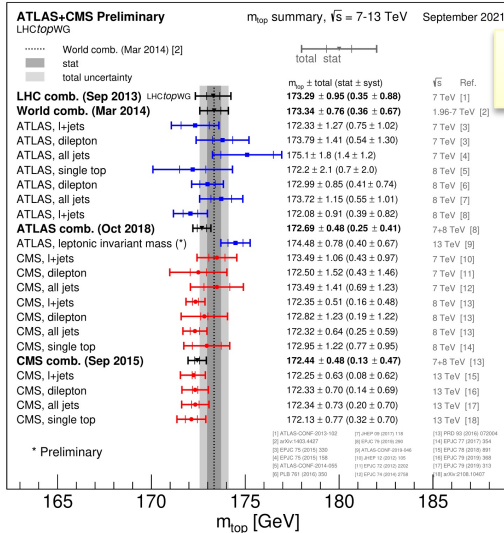
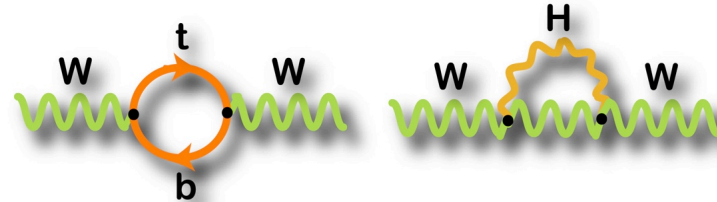


Light blue: fit excluding input from row

A very sensitive SM test: the m_W - m_t plane

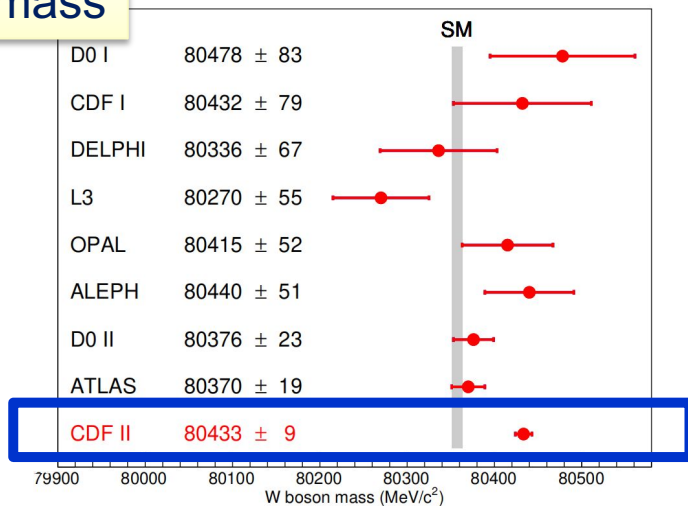


$$m_W^2 \left(1 - \frac{m_W^2}{m_Z^2} \right) = \frac{\pi\alpha}{\sqrt{2}G_\mu} (1 + \Delta r)$$



A very sensitive SM test: the m_W - m_t plane

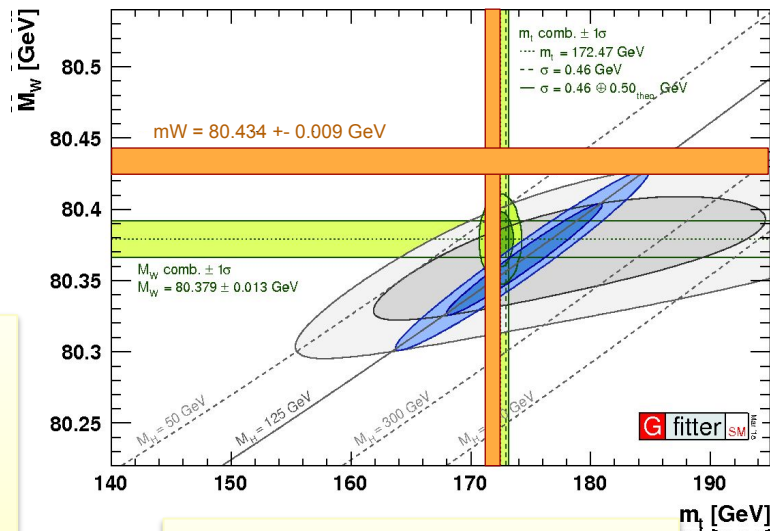
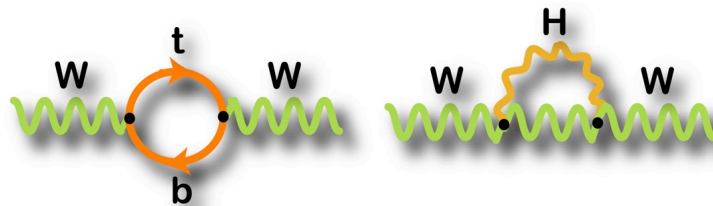
W mass



$m_W(\text{SM}) = 80375 \pm 4(\text{inp}) \pm 4(\text{th}) \text{ MeV}$
 $m_W(\text{CDF}) = 80433 \pm 9.4 \text{ MeV}$
 $\Delta m_W(\text{SM-CDF}) = 7\sigma \dots$

More precise than all previous m_W measurements combined
 ~ 3σ higher than most precise other measurements
 ~ 3σ higher than previous CDF result (1/4 of data, correlations taken into account)

$$m_W^2 \left(1 - \frac{m_W^2}{m_Z^2} \right) = \frac{\pi\alpha}{\sqrt{2}G_\mu} (1 + \Delta r)$$



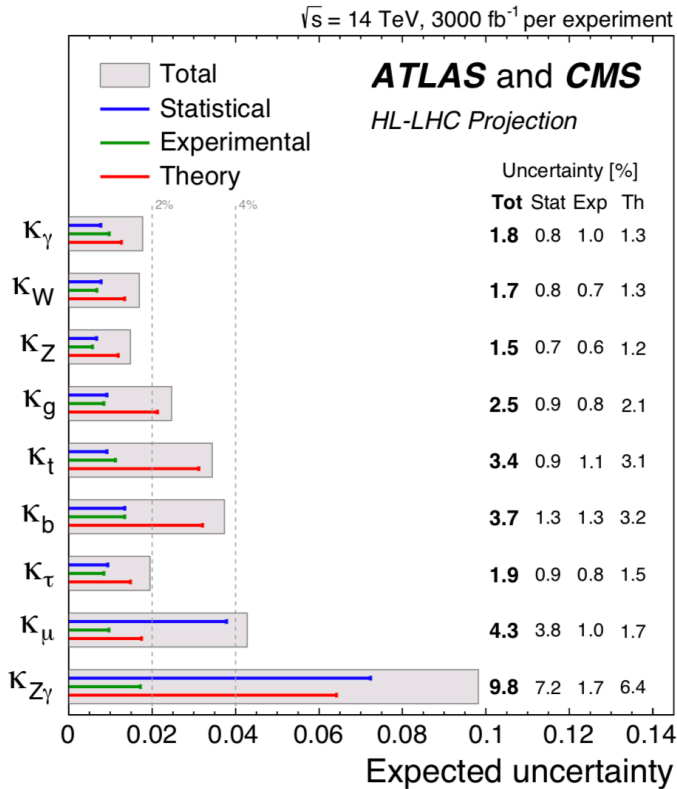
“in tension” with SM pred

The Standard Model

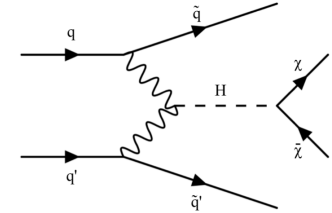
The Scalar Sector

Medium-term Higgs physics: the LHC/HL-LHC program

HL-LHC reach: Higgs

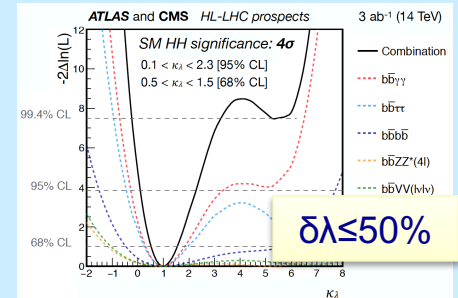
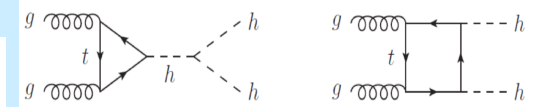
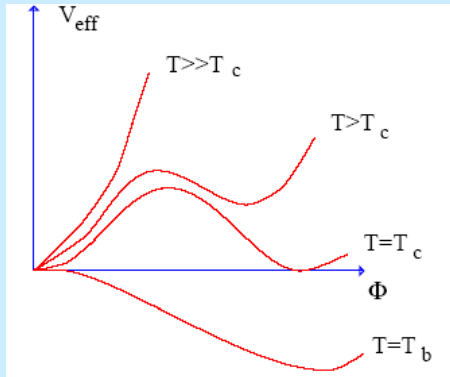


H width to invisible:
 $h(125) \rightarrow XX$.
Includes BSM decays
and rare SM decays: $\leq 4\%$

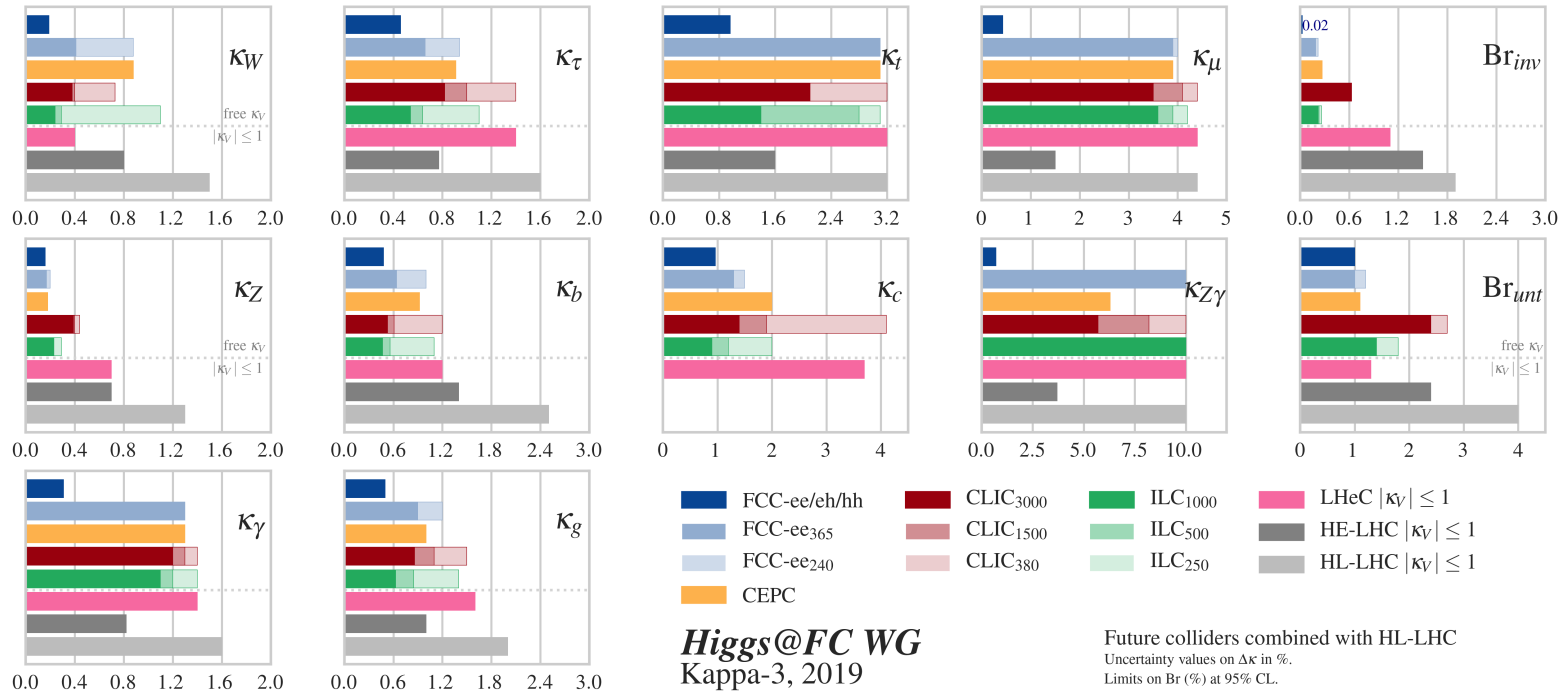


The ultimate frontier: Higgs self-coupling

$$V_h = \frac{m_h^2}{2} h^2 + (1 + \kappa_3) \lambda_{hhh}^{\text{SM}} v h^3 + \frac{1}{4} (1 + \kappa_4) \lambda_{hhhh}^{\text{SM}} h^4$$



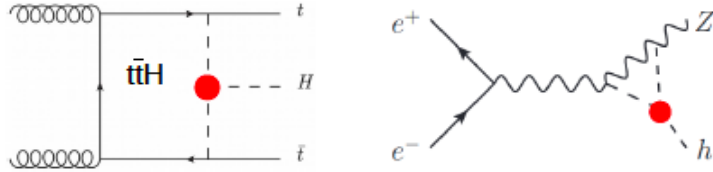
Grand summary: Higgs couplings (κ framework) (Future Collider+HL-LHC)



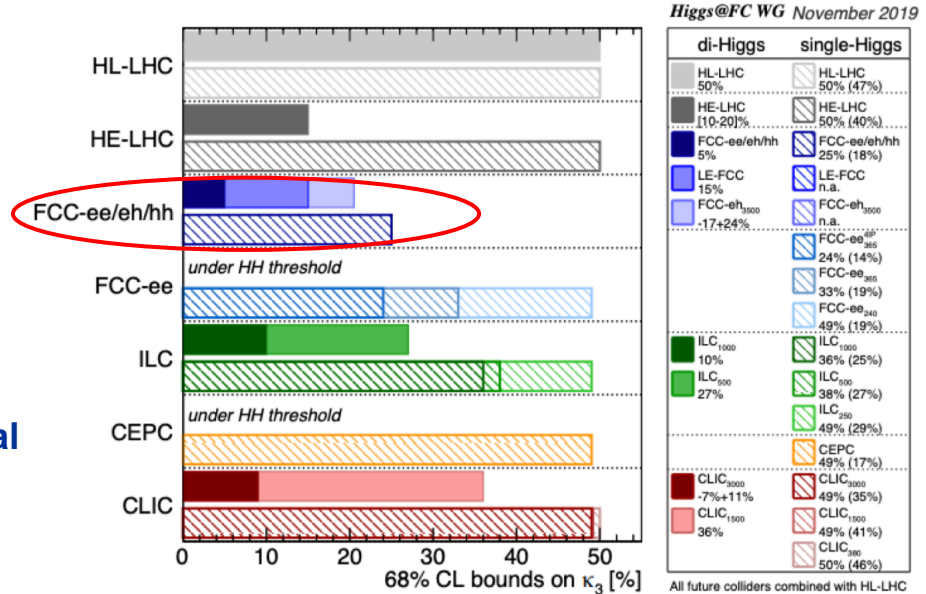
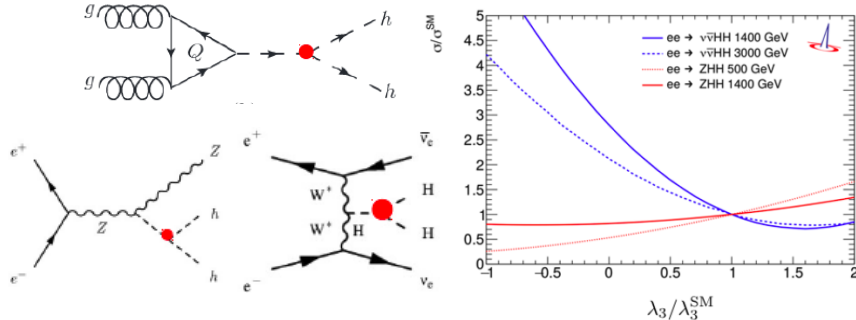
Plus of κ framework: it is simple; Minus: underestimates effects of polarization
 Can show deviation from SM, but no real further information on nature of source of deviation;
 Untagged and invisible BRs constrained by measurements.

Higgs self-coupling

- Single-H production
 - Sensitivity via loop diagrams



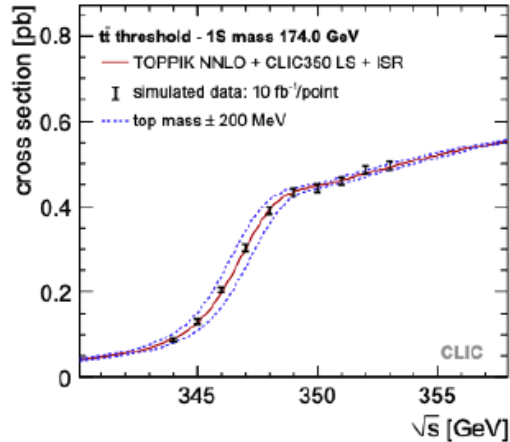
- For $k_\lambda=2$, hh: 3%; ee: 1%.
- HH production
 - hh: $\sigma(\text{HH}) \sim 0.01\sigma(\text{H})$; must use differential measurements;
 - ee: Complementarity of ZHH and VBF production



Single-H: FCC-ee or ILC ~ 35% (global analysis)
HH production:
 HL-LHC: ~50% (can probably do better?)
 HE-LHC: ~15%
 ILC500: ~27%; CLIC1500 ~36%
 CLIC3000: ~9%; FCC-hh ~5%

Precision EWK Observables

Top mass from threshold scan



Currently:

$\delta m_t \approx 400$ MeV

ILC/CLIC/FCC: ~ 25 MeV

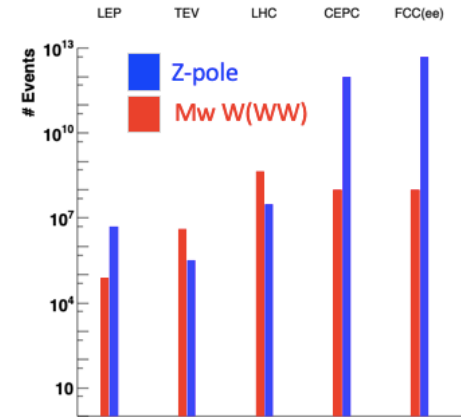
Note: $\delta m_t^{\text{TH}} \sim 40$ MeV (α_s), 40 MeV (HO)

Theory errors should decrease from Z-pole running; e.g. FCC-ee: 5×10^{12} Z (!!!)

EWPO: circular ee colliders

+ linear colliders for $\sin^2 \theta_W$.

Note: currently, discussion/plan for a large Z run for the linear colliders...



EWPO	Current	CEPC	FCC (ee)
M_Z [MeV]	2.1	0.5	0.1
Γ_Z [MeV]	2.1	0.5	0.1
N_ν [%]	1.7	0.05	0.03
M_W [MeV]	12	1	0.67
$A_{FB}^{0,b}$ [$\times 10^4$]	16	1	< 1
$\sin^2 \theta_W^{\text{eff}}$ [$\times 10^5$]	16	1	0.6
R_b^0 [$\times 10^5$]	66	4	2-6
R_μ^0 [$\times 10^5$]	2500	200	100

Beyond the Standard Model (I)

Composite Higgs?

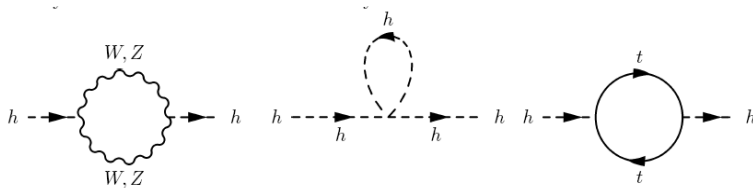
Sypersymmetry?

Contact Interactions?

New resonances?

The magic of the Higgs boson mass

- **Quantum Mechanics: ultimate destructor of small numbers (in nature) that are not protected by some symmetry (thus “law”)**
- **Higgs boson: the ultimate example**
 - **Quadratic divergence in the Higgs mass**

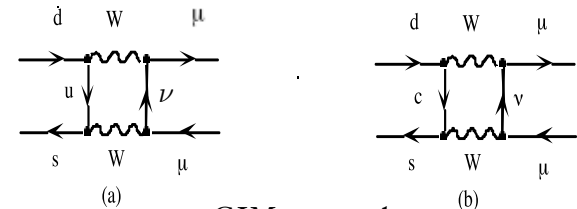


$$m^2(p^2) = m^2(\Lambda^2) + Cg^2 \int_{p^2}^{\Lambda^2} dk^2$$

- **H mass should be $\sim 10^{19}\text{GeV}$**
 - **Yet, it lies at 125 GeV...**
- **Put differently: if cut off at Λ_{Pl} , why $m_W \ll M_{\text{Pl}}$?**
 - **Or, why is gravity ($G \sim 1/M_{\text{Pl}}$) so weak?**

- **Reminder of just two past applications of naturalness argument:**
- **Weisskopf (1939): “the self-energy of charged particles obeying Bose statistics is found to be quadratically divergent...,”**
 - in theories of elementary bosons, new phenomena must enter at an energy scale of m/e
 - + positron, doubling of particles...

- **Rare Kaon decays: $K_L \rightarrow \mu^+ \mu^-$**



GIM proposal

$$\sim G \Lambda^2 = G(m_c^2 - m_u^2)$$

$$\rightarrow \Lambda^2 \sim 3-4 \text{ GeV}^2 \text{ !!!}$$

Solutions to the hierarchy problem

- **Solution #1: a composite Higgs**

- H: bound state (e.g. due to some new strong interaction)

- **Solution #2: supersymmetry**

- Partners for ALL SM particles, spin different by $\frac{1}{2}$
- Presumably broken symmetry (since partners unseen)

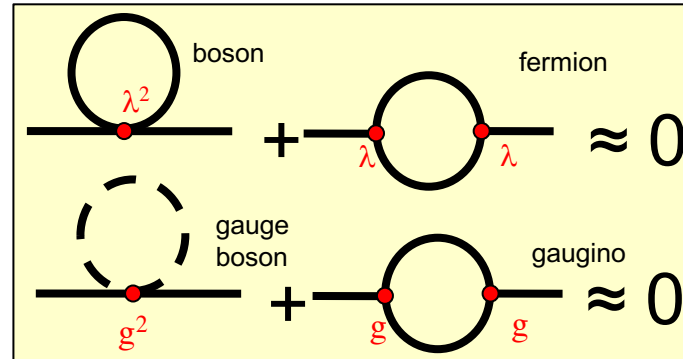
- **Solution #3: little Higgs**

- H: pseudo-Goldstone boson of Ultimate Theory; just another effective theory, e.g. valid to ~ 10 TeV. Loops cancel by particles of same spin (so need to introduce these particles)

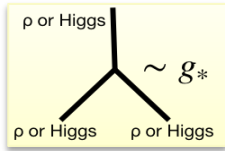
- **Solution #4: extra dimensions**

- N-dim space; gravity propagates in all dims, SM only in “our” 3 dims; e.g. warped extra dimension can explain weakness of G

- **All of these “solutions” introduce either deviations in tails or new particles, with masses whose “natural” values should be O(few TeV)**



Higgs Compositeness?



$$\text{gauge } \rho \sim \frac{g_W}{g_*}$$

H

$$\frac{c_\phi}{\Lambda^2} = \frac{g_*^2}{m_*^2}$$

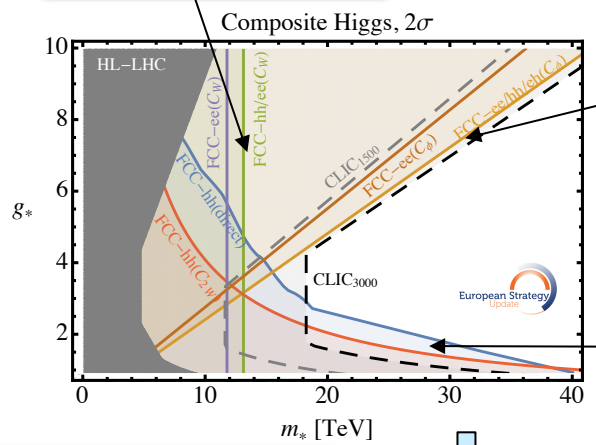
2f-2b

$$\frac{c_W}{\Lambda^2} = \frac{1}{m_*^2}$$

4f, W'/Z'

$$\frac{c_{2W}}{\Lambda^2} = \frac{1}{g_*^2 m_*^2}$$

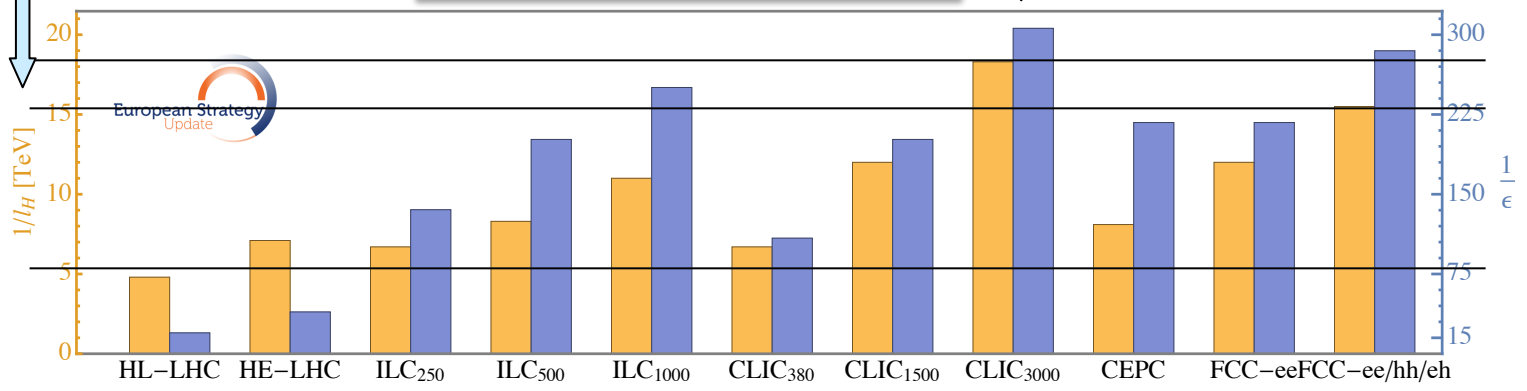
O_W (independent of g_*)



O_ϕ (Higgs measurements, primarily FCC-ee & CLIC)

Direct resonance searches (generic “ ρ ” EWK triplet; dileptons & dibosons)

Obtain limit on m_* , i.e. on $1/\text{length}(H)$



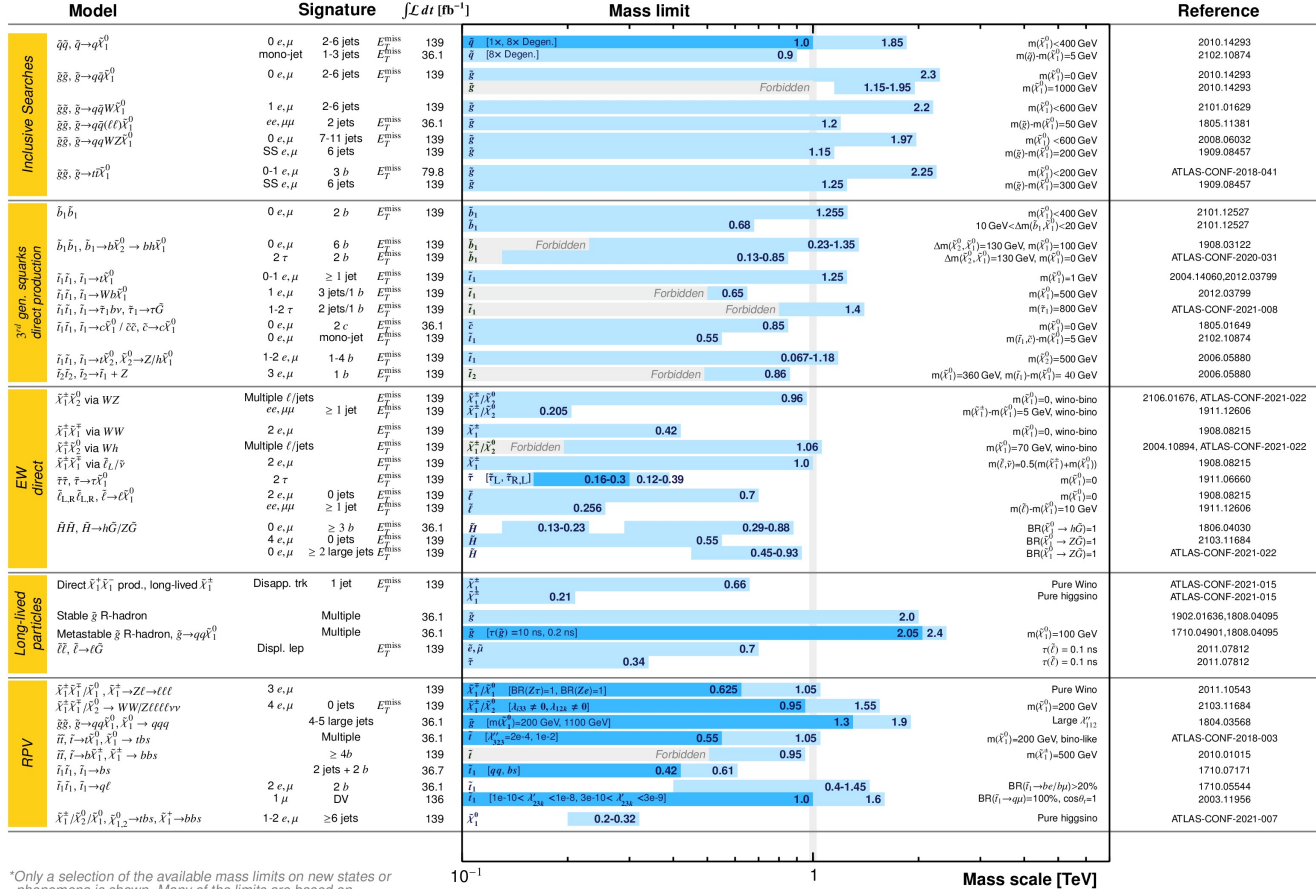
The never-ending search for SUSY... (and it's not like we didn't look for them)

ATLAS SUSY Searches* - 95% CL Lower Limits

June 2021

ATLAS Preliminary

$\sqrt{s} = 13$ TeV



*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10 $^{-1}$ 1 Mass scale [TeV]

Long-term future: SUSY at >1TeV.... <10 TeV?

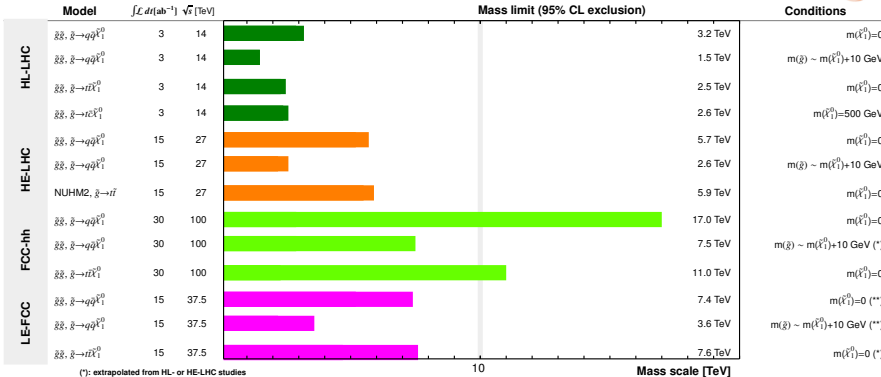
□ The questions:

- If {SUSY} which masses (and mass differences) of strongly- or weakly-coupled superpartners can we reach?
- Is nature fundamentally fine-tuned? If the solution is SUSY, how well can we test this?
- Is dark matter a thermal SUSY WIMP?

□ Strongly-interacting SUSY (gluinos and squarks): simply, the purview of hadron colliders

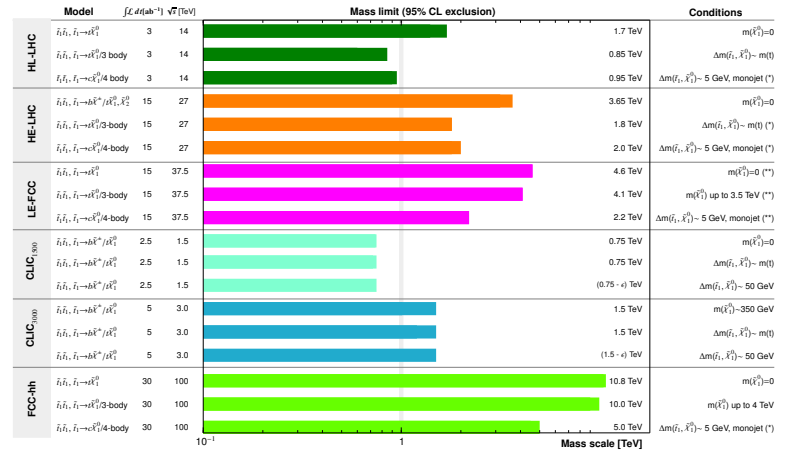
Hadron Colliders: gluino projections

(R-parity conserving SUSY, prompt searches)



HE-LHC: $\sim 2 \times M_{\text{HL-LHC}}$
FCC-hh: $\sim 5 \times M_{\text{HL-LHC}}$

All Colliders: Top squark projections (R-parity conserving SUSY, prompt searches)



ILC 500: discovery in all scenarios up to kinematic limit $\sqrt{s}/2$

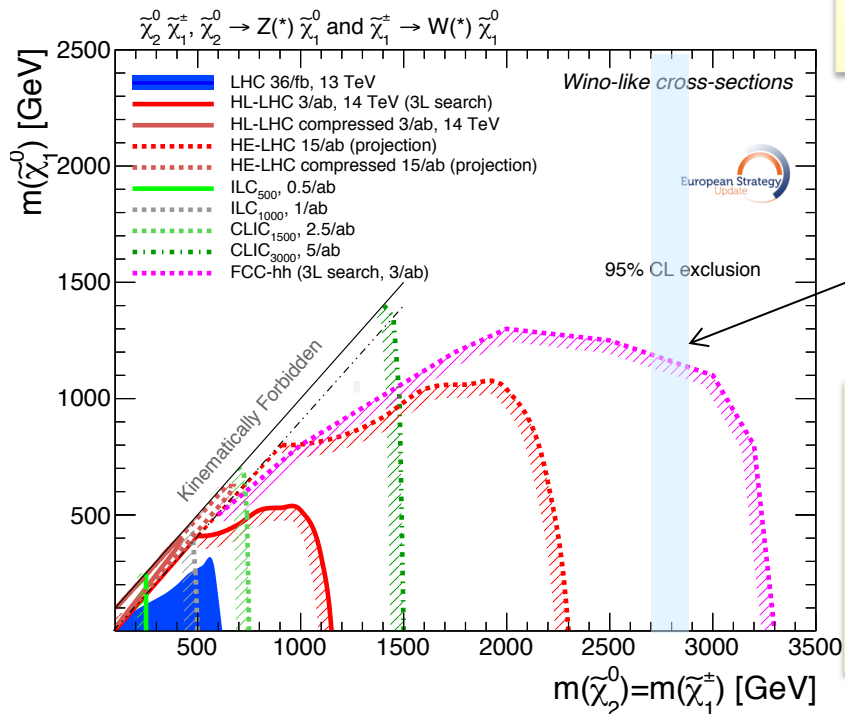
Long-term future: SUSY EWK sector

ee: mainly $\chi_{2,1}^0 \chi_{2,1}^0$ and $\chi_{1,1}^+ \chi_{1,1}^-$ ($\Delta m > 1 \text{ GeV}$)
 pp: mainly $\chi_{2,1}^0 \chi_{2,1}^\pm$ (gaps at low Δm)

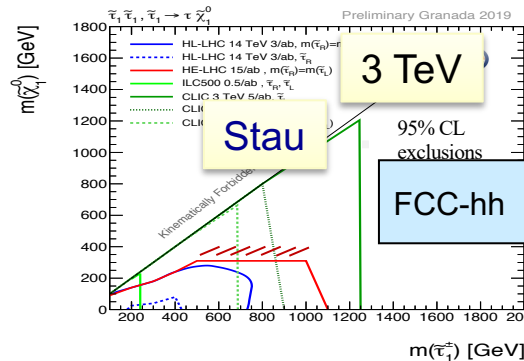
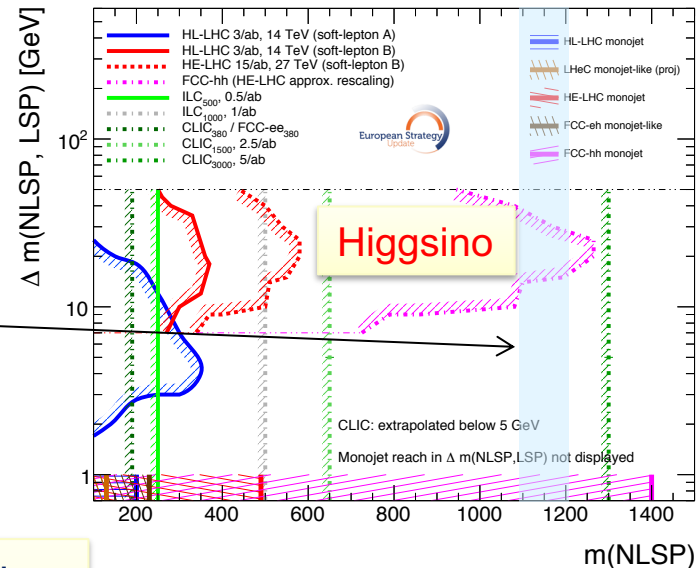
ee sensitivity
 ~independent
 of nature of
 LSP
 (but up to $\sqrt{s}/2$)

Thermal
 WIMP

pp sensitivity:
 depends on LSP;
 ISR jet for very low
 Δm (up to $\sim 5 \text{ GeV}$);
 Pure Higgsino:
 stopped track



Higgsino-like EWK processes

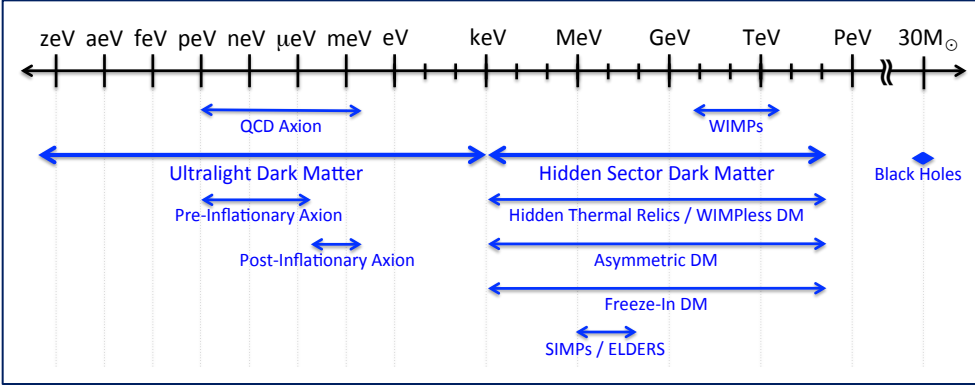


Beyond the Standard Model (II)

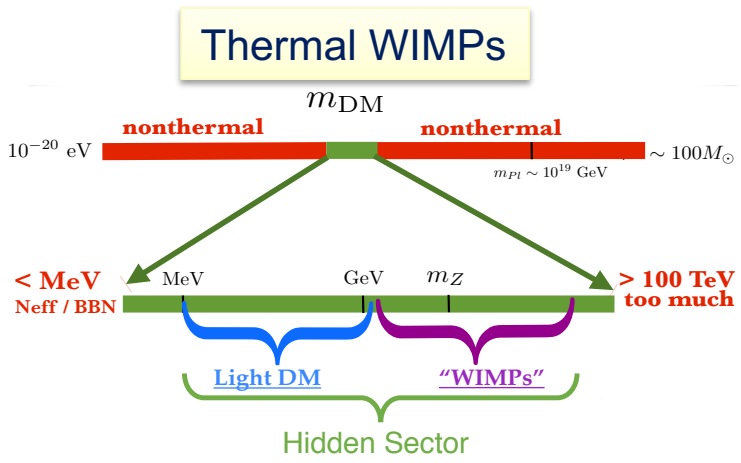
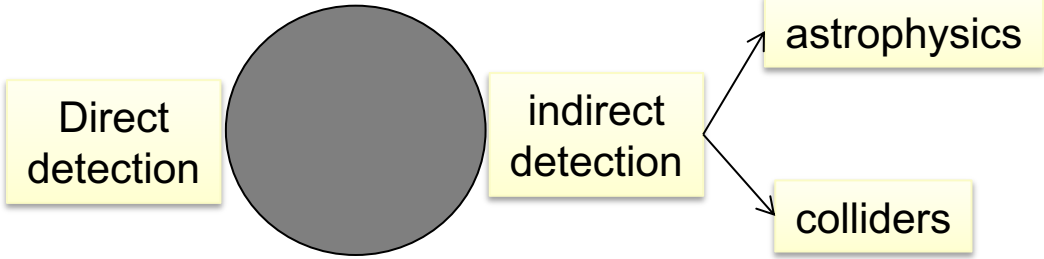
The Dark Sector
Feebly Interacting Particles

The Dark Sector

- **An experimental fact & yet, still a total mystery**
 - **And masses span over 80 orders of magnitude**
- **Nightmare scenario: totally dark**
 - **Only Gravity to play with...**



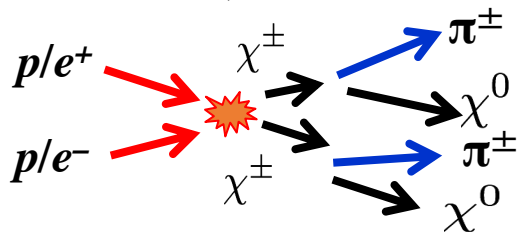
- **More promising: some shade of grey**



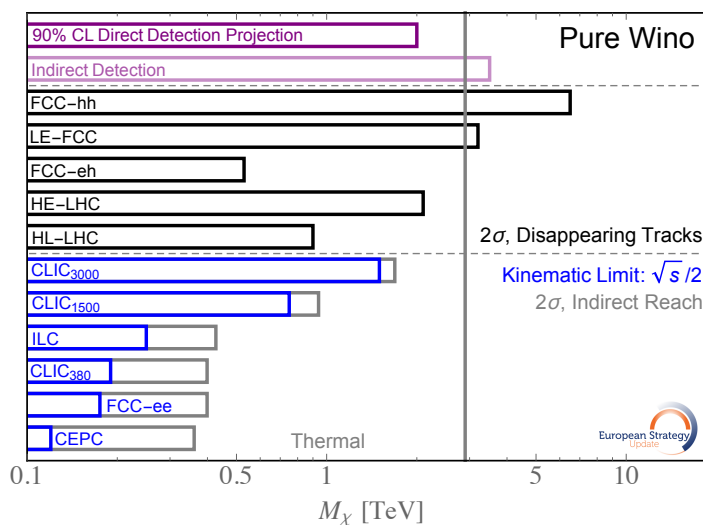
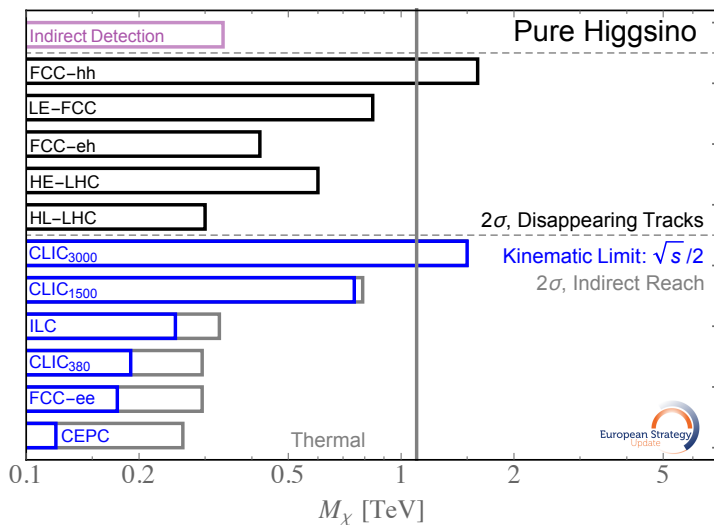
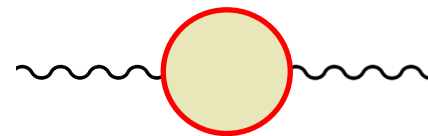
DM: Classic WIMPs

- Two (SUSY) “extremes”, pure Wino, pure Higgsino
 - Main “tools”: disappearing track, propagator modifications

For small Δm , soft π^\pm ...

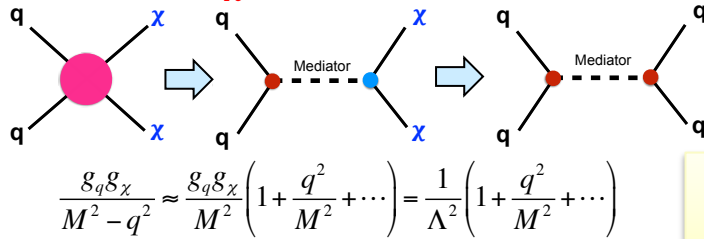


EWKinops in loop change prop
(W, Y parameters)



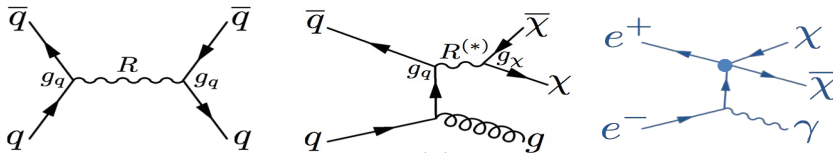
Simplified Models: (axial)vector&(pseudo)scalar mediator

Light DM, $m_\chi=1\text{GeV}$



V/A: $g_\chi=1, g_q=0.25$
P/S: $g_\chi=g_q=1$

Resonances + mono-object (photon, jet, top, W/Z...)



pp: assumes mediator couplings to quarks only.

750 GeV, HL-LHC

1.5 TeV, HE-LHC

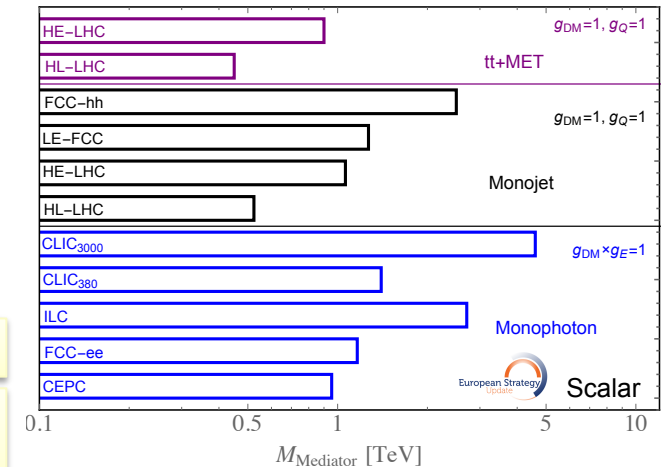
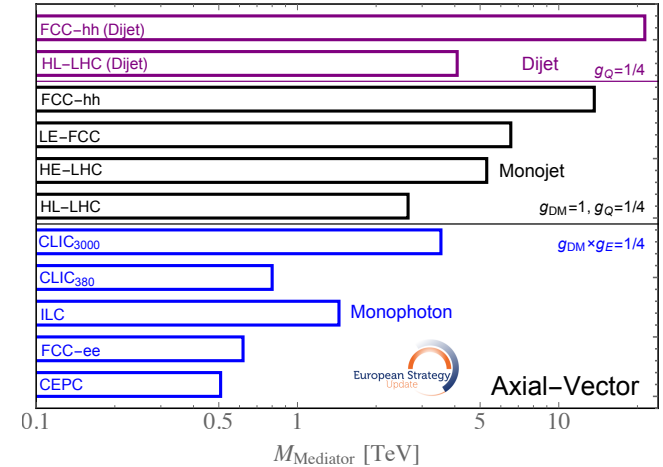
3.9 TeV for FCC-hh

Note: Dependence on couplings!

ee: assumes mediator couplings to leptons only. Also in EFT limit, so can be easily rescaled for modified couplings.

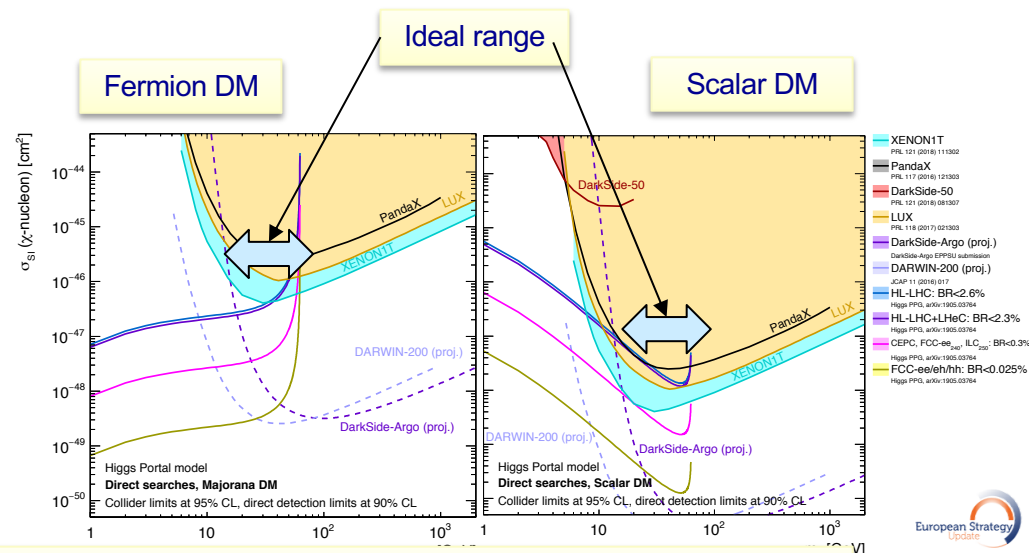
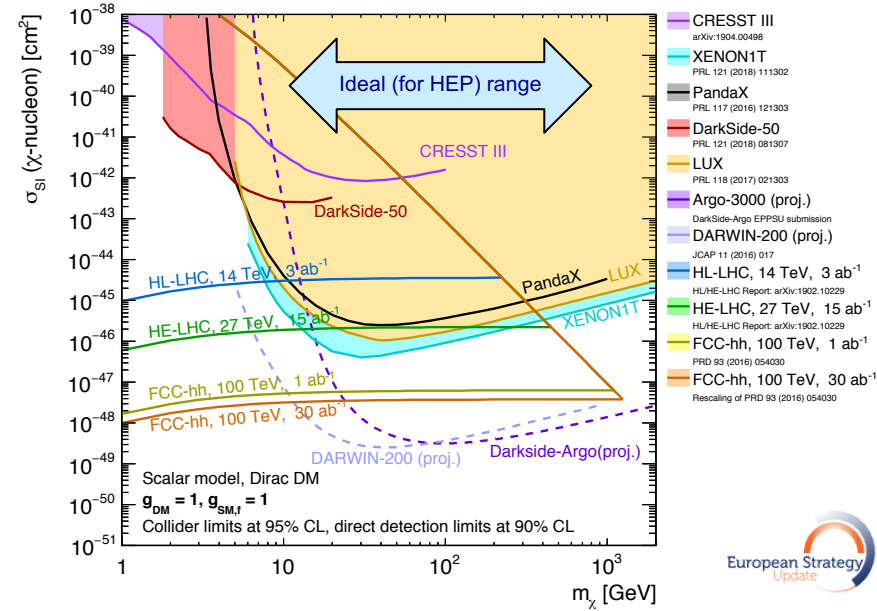
Note: taking EFT scale as free parameter, M_{DM} reach \sim kinematic reach of collider.

Significant model dependence. UV models may have comparable quark and lepton couplings. If both present, can also use dilepton resonances.



Scalar mediator: Higgs portal and BSM scalar

A collider discovery will need confirmation from DD/ID for cosmological origin
 A DD/ID discovery will need confirmation from colliders to understand the nature of the interaction



A future collider program that optimizes sensitivity to invisible particles coherently with DD/ID serves us well. Need maximum overlap with DD/ID!

Feebly Interacting Particles (FIPs)

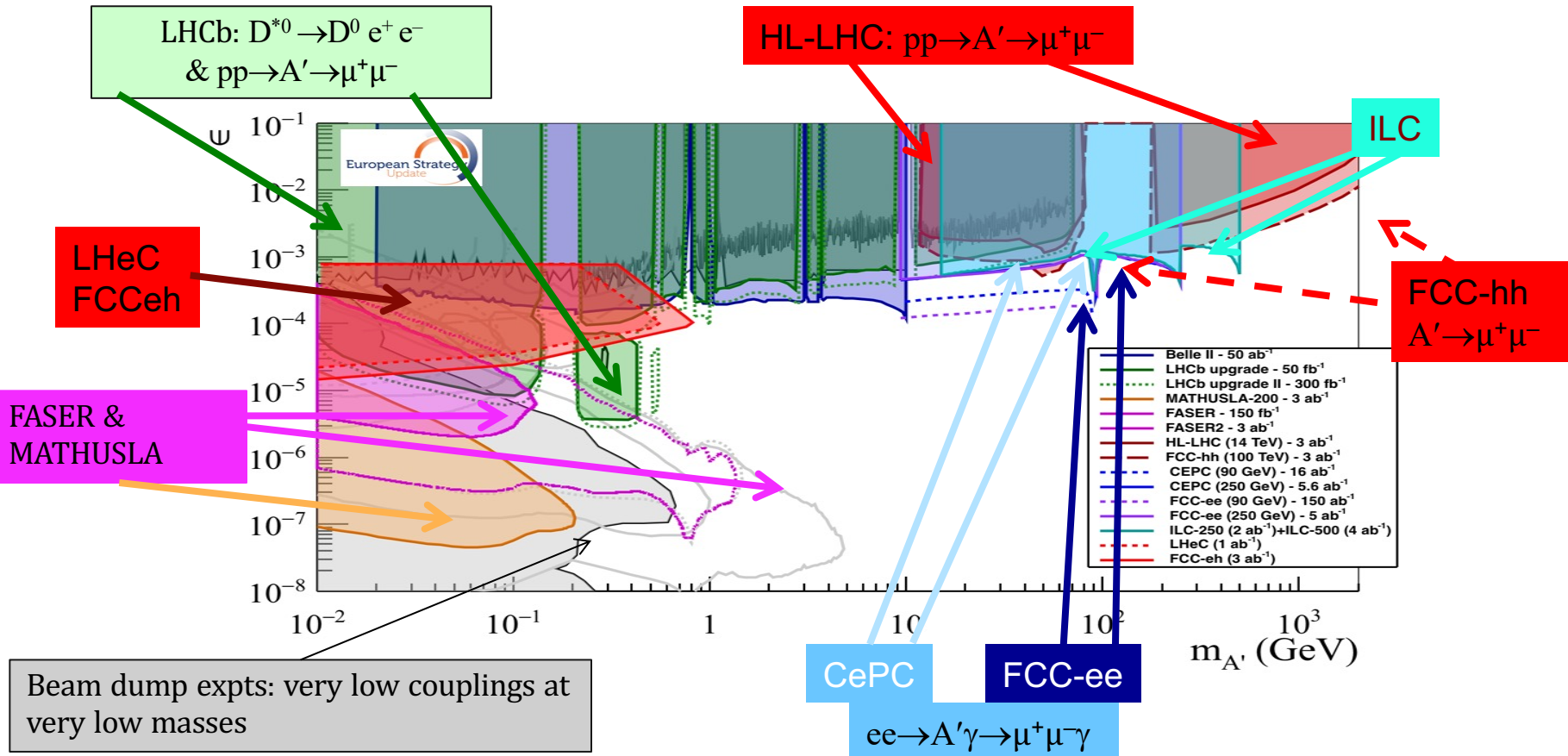
- **FIP mass(es) can span several orders of magnitude**
 - Though there are preferred regions for motivated models (Dark photon for thermal dark matter, relaxion in its natural region, right-handed neutrinos below EW scale down to the see-saw limit) that are within reach for accelerator-based experiments
- **Very (very...) wide range of possibilities .AND. Models**
 - How to search for such broad class of models? → Simplified models
 - How to compare frontiers & Experiments? → Use benchmarks.
 - Simplified models: four “portals”

PBC report, arXiv:1901.09966

Portal	Coupling	
Dark Photon, A_μ	$-\frac{\epsilon}{2 \cos \theta_W} F'_{\mu\nu} B^{\mu\nu}$	
Dark Higgs, S	$(\mu S + \lambda S^2) H^\dagger H$	(Relaxion toy model, mixes \w Higgs)
Axion, a	$\frac{a}{f_a} F_{\mu\nu} \tilde{F}^{\mu\nu}, \frac{a}{f_a} G_{i,\mu\nu} \tilde{G}_i^{\mu\nu}, \frac{\delta_{\mu\alpha}}{f_a} \bar{\psi} \gamma^\mu \gamma^5 \psi$	
Sterile Neutrino, N	$y_N L H N$	

- **Beam dump and collider experiments: complementary in reach**
 - Very significant reach in several places. Not exhaustive – but this is only the beginning.
 - **Physics Beyond Colliders effort...**
 - From portals: identify benchmark cases to evaluate experimental sensitivities. Common ground to compare machines/experiments and put them in worldwide context

FIPs: Example from Vector Portal (Dark Photon)



FIPs represent a new paradigm that requires systematic exploration on multiple fronts

Physics of Flavor

Neutrinos

Quarks

Charged Leptons

Super-brief Intro: from hopeless to lucky strike(s)

From invisible particle ($\sigma \sim 10^{-44} \text{cm}^2$ @ $E_\nu \sim 2 \text{ MeV}$) to major source of physics:
 From Pontecorvo's few/day/ton with $\sim 10^{11} \nu/\text{cm}^2/\text{s}$ to \rightarrow PMNS matrix & CP violation(?!?!)

$$P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2 2\theta \sin^2 \left(1.27 \frac{\Delta m^2 (\text{eV}^2) L(\text{km})}{E_\nu (\text{GeV})} \right) \quad P(\nu_\alpha \rightarrow \nu_\alpha) = 1 - P(\nu_\alpha \rightarrow \nu_\beta)$$

SuperK, SNO: large mixing; good Δm
 $\Delta m_{\text{sol}}^2 \approx 7 - 8 \times 10^{-5} \text{eV}^2$

$\Delta m_{\text{sol}}^2 \sim E_\nu (\text{MeV}) / L (100 \text{ km})$
 \rightarrow Reactors! \rightarrow KamLAND

IMB, Kamiokande, Soudan(2)... SuperK:
 large mixing; super Δm : $\Delta m_{\text{atm}}^2 \approx 2.5 \times 10^{-3} \text{eV}^2$

$\Delta m_{\text{atm}}^2 \sim E_\nu (\text{MeV}) / L (100 \text{ m}) = (\text{GeV}) / (100 \text{ km})$
 \rightarrow Accelerators! \rightarrow K2K, MINOS(+), ..., T2K, NOVA
 + Short baselines: Daya Bay, RENO, ..., xBOONE

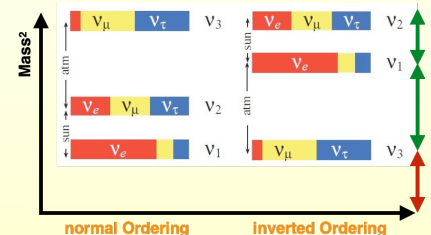
ν physics: PMNS; 3x3 or 4x4? nature (Majorana or Dirac); mass ordering

$$U_{3 \times 3} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times (U_{\text{Maj}})$$

atmospheric +
 accelerator disapp

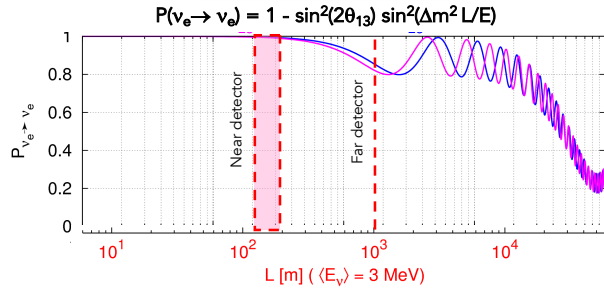
SBL reactor +
 accelerator app

solar +
 KamLAND

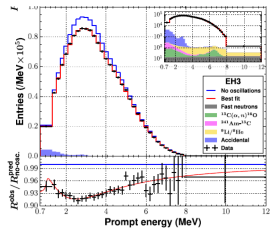
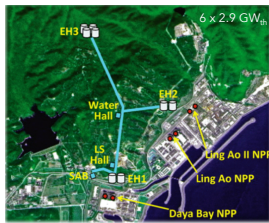


Neutrinos: Reactor expts, Short Baselines

Principle: θ_{13}



The prized measurement



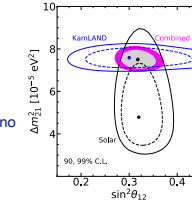
Third lucky strike (in ν phys)

$$J_{\text{CP}}^{\text{lep}} = \text{Im}(U_{\mu 3} U_{e 3}^* U_{e 2} U_{\mu 2}^*) = (1/8) \times \\ \times \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13} \sin \delta_{\text{CP}} \\ \approx 0.033 \sin \delta_{\text{CP}} \text{ BUT : } J_{\text{CP}}^{\text{q}} \approx 3 \times 10^{-5}$$

Long series of measurements have brought us to a “precision era”

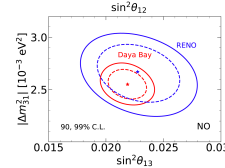
solar sector

Cl, Ga, SK
SNO, Borexino
KamLAND



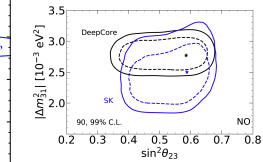
SBL reactors

Daya Bay
RENO



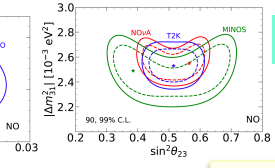
atmospheric results

Super-K
IC-DeepCore



LBL experiments

MINOS
T2K
NOvA



M. Tortola ICHEP2020

de Salas et al, arXiv:2006.11237

parameter	best fit $\pm 1\sigma$	3σ range
Δm_{21}^2 [10^{-5}eV^2]	$7.50^{+0.22}_{-0.20}$	6.94–8.14
$ \Delta m_{31}^2 $ [10^{-3}eV^2] (NO)	$2.56^{+0.03}_{-0.04}$	2.46–2.65
$ \Delta m_{31}^2 $ [10^{-3}eV^2] (IO)	2.46 ± 0.03	2.37–2.55
$\sin^2 \theta_{12} / 10^{-1}$	3.18 ± 0.16	2.71–3.70
$\sin^2 \theta_{23} / 10^{-1}$ (NO)	$5.66^{+0.16}_{-0.22}$	4.41–6.09
$\sin^2 \theta_{23} / 10^{-1}$ (IO)	$5.66^{+0.18}_{-0.23}$	4.46–6.09
$\sin^2 \theta_{13} / 10^{-2}$ (NO)	$2.225^{+0.055}_{-0.078}$	2.015–2.417
$\sin^2 \theta_{13} / 10^{-2}$ (IO)	$2.250^{+0.056}_{-0.076}$	2.039–2.441
δ/π (NO)	$1.20^{+0.23}_{-0.14}$	0.80–2.00
δ/π (IO)	1.54 ± 0.13	1.14–1.90

2.7%

1.2%

5.2%

4.9%

4.8%

3.0%

17%

8%

relative 1σ uncertainty

3-5%

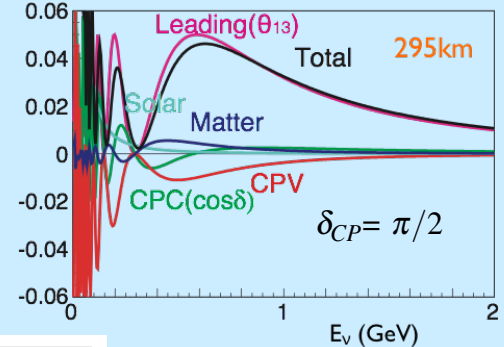
Neutrinos: oscillations & CPV

Currently (NOvA and T2K) a bit confused

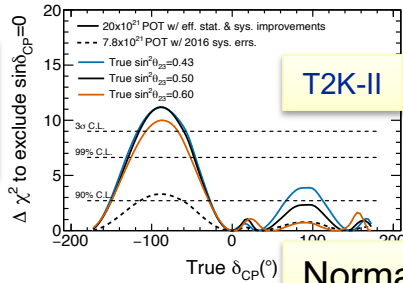
$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) = & \sin^2 \theta_{23} \frac{\sin^2 2\theta_{13}}{(A-1)^2} \sin^2[(A-1)\Delta_{31}] \\
 & + \alpha^2 \cos^2 \theta_{23} \frac{\sin^2 2\theta_{12}}{A^2} \sin^2(A\Delta_{31}) \\
 & - \alpha \frac{\sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23} \cos \theta_{13} \sin \delta_{CP}}{A(1-A)} \sin \Delta_{31} \sin(A\Delta_{31}) \sin[(1-A)\Delta_{31}] \\
 & + \alpha \frac{\sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23} \cos \theta_{13} \cos \delta_{CP}}{A(1-A)} \cos \Delta_{31} \sin(A\Delta_{31}) \sin[(1-A)\Delta_{31}]
 \end{aligned}$$

Favored (slightly):
Normal Ordering &
 $\delta_{CP} \sim 3\pi/2$

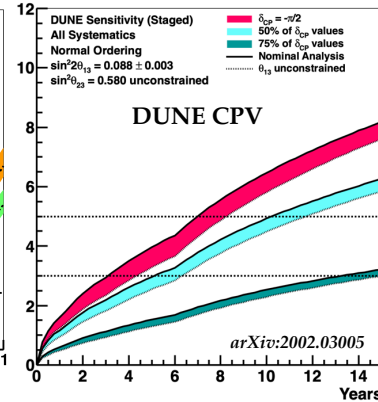
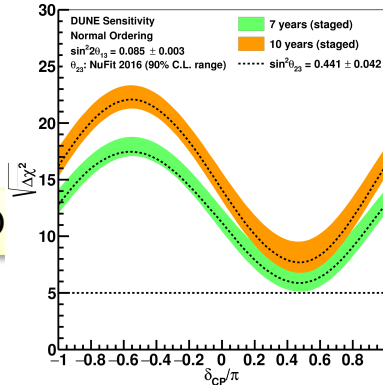
But... fits are complicated



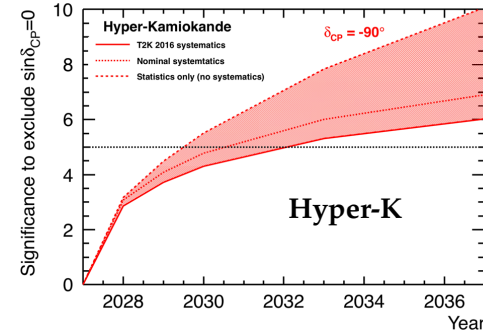
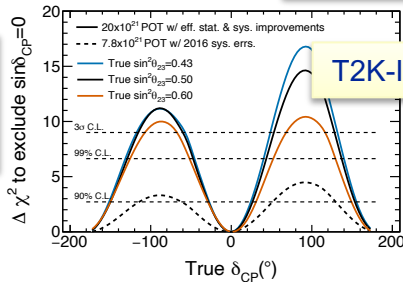
Unknown mass order



NormalO



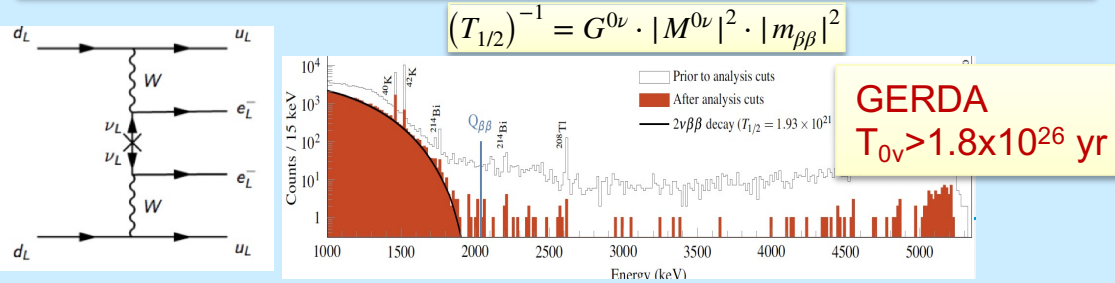
Known mass order



Very useful/nice: complementarity of NOvA/T2K and DUNE/HyperK: baseline (1300 vs 200 km) and detection technique

Neutrinos: Majorana/Dirac ($0\nu\beta\beta$) & mass measurement

As fundamental as a question can get & very hard to answer



Mass limits (measurements...)

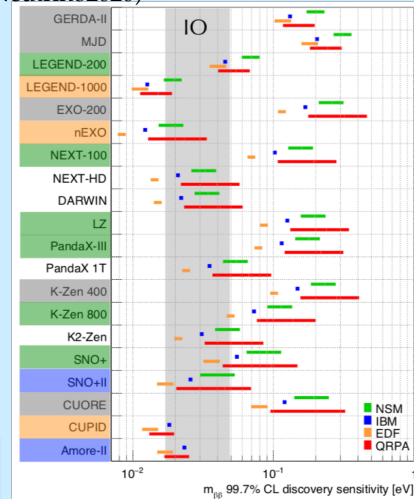
Today (BAO +struct. form):
 $\Sigma m_\nu < 0.12$ (95% CL)
Ultimate (DESI+BAO): $\Sigma m_\nu < 0.02$

Direct m_ν measurement(s)
KATRIN: $m_\beta < 1.1$ eV (90% CL)
Asymptotic: $m_\beta < 0.2$ eV
Ultimate (Project-8): 40 meV

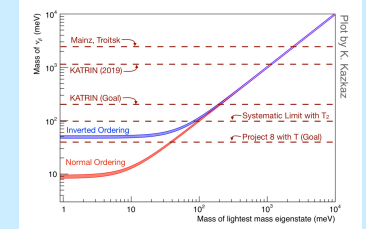
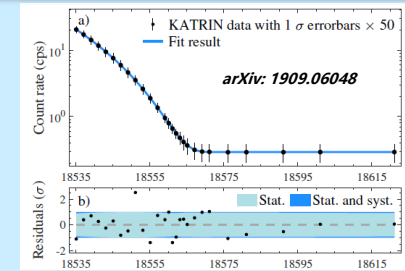
Extremely active field

(J. Detwiler @ Neutrino2020)

- Bolometers (CUPID, AMoRE, CANDLES IV)
 - Measure E ($\sigma \sim 0.1$ - 0.3%) from phonons; granularity gives position info
 - Instrumenting with photon detectors for background rejection
- External trackers (SuperNEMO)
 - Trackers + calorimeters, measure E ($\sigma \sim 3$ - 10%) + tracks / positions + PID
- Scintillators (KamLAND2-Zen, SNO+, Theia, ZICOS)
 - Measure E ($\sigma \sim 3$ - 10%) + position from scintillation light; some PID
- Semiconductors (LEGEND, SELENA)
 - Measure E ($\sigma \sim 0.05$ - 0.3%) from ionization; some tracking / position sensitivity
- TPCs (nEXO, NEXT, PandaX, AXEL, NvDex, DARWIN, LZ)
 - Collect scintillation + ionization: measure E ($\sigma \sim 0.4$ - 3%) + tracks / position + PID



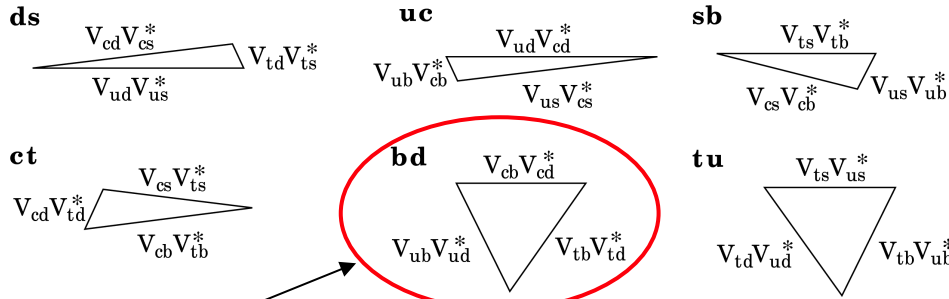
(But) one example: EXO-200 (0.2 t) \rightarrow nEXO (5 t)
Xe; $T_{1/2} < 3.5 \times 10^{25}$ yr \rightarrow $T_{1/2} < 10^{28}$ yr



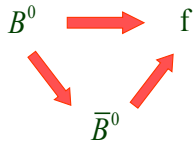
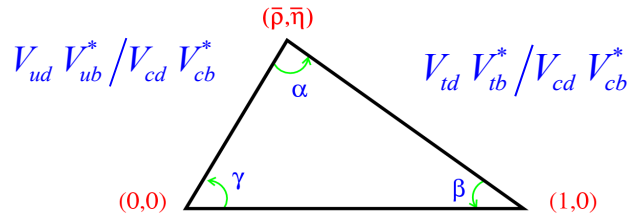
Flavor physics: quark sector (I)

$$V_{\text{CKM}} = \begin{bmatrix} 1 - \lambda^2/2 & \lambda & A\lambda^3(\rho - i\eta) \\ -\lambda & 1 - \lambda^2/2 & A\lambda^2 \\ A\lambda^3(1 - \rho - i\eta) & -A\lambda^2 & 1 \end{bmatrix}$$

$$V_{ui} V_{uj}^* + V_{ci} V_{cj}^* + V_{ti} V_{tj}^* = 0$$

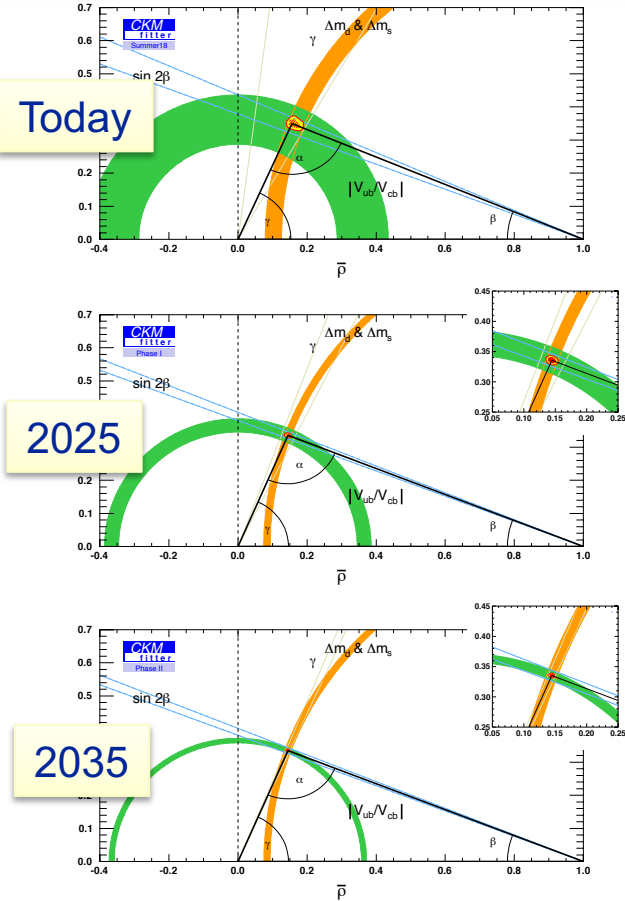


CP $\sim \sin 2\phi$



$$\frac{\Gamma(\bar{B}^0 \rightarrow \bar{f}) - \Gamma(B^0 \rightarrow f)}{\Gamma(\bar{B}^0 \rightarrow \bar{f}) + \Gamma(B^0 \rightarrow f)} = -\eta_f \sin(2\phi) \sin(\Delta M t)$$

CKM triangle: extensive precision program en route



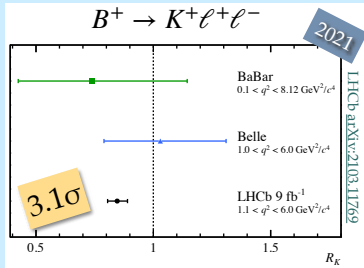
Flavor physics: quark sector (II)

Lots of attention on Lepton Universality.

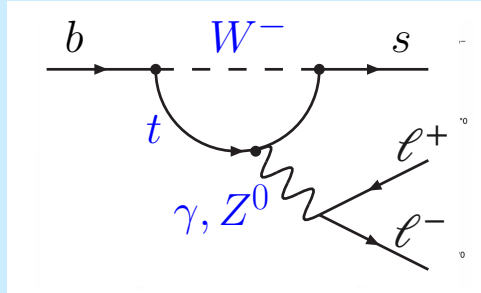
A. Greljo @ FAW21*

$$b \rightarrow s \ell^+ \ell^-$$

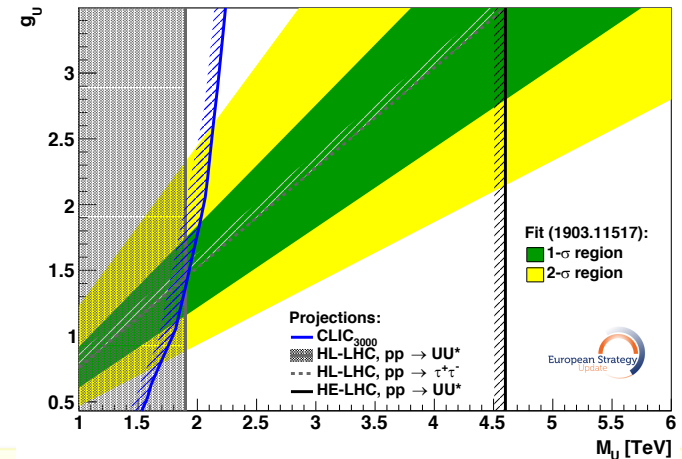
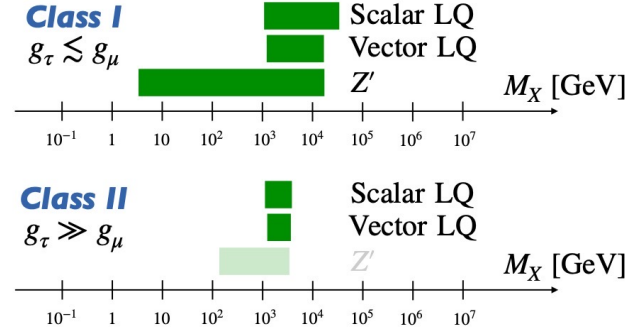
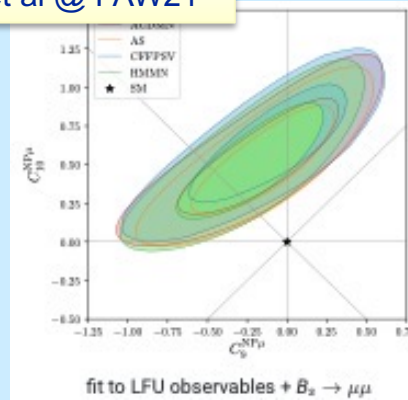
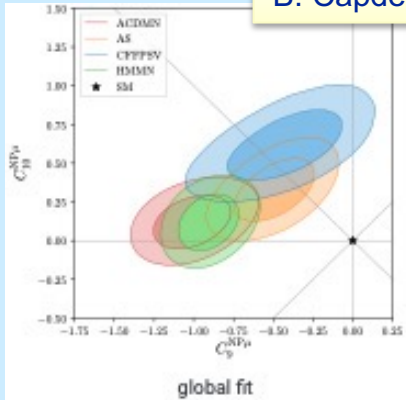
$$R_K = \mu\mu K / eeK \quad R_{K^*} = \mu\mu K^* / eeK^*$$



+semileptonic + ang. distrib.



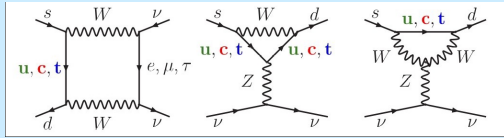
B. Capdevila et al @ FAW21*



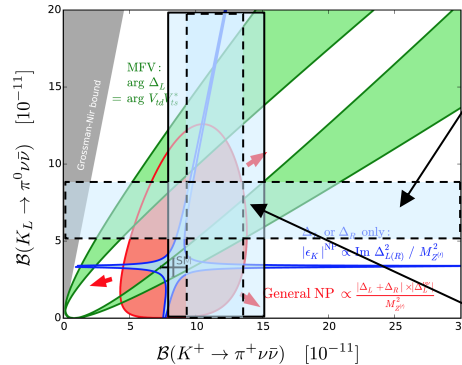
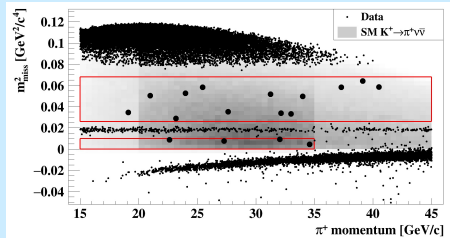
*Unpaid ad for FAW21: [Flavour Anomaly Workshop 2021](#)

Flavor physics: quark sector (III)

Today: NA62



$$(K^+ \rightarrow \pi^+ \nu \bar{\nu}) = (11.0^{+4.0}_{-3.5}) \times 10^{-11}$$



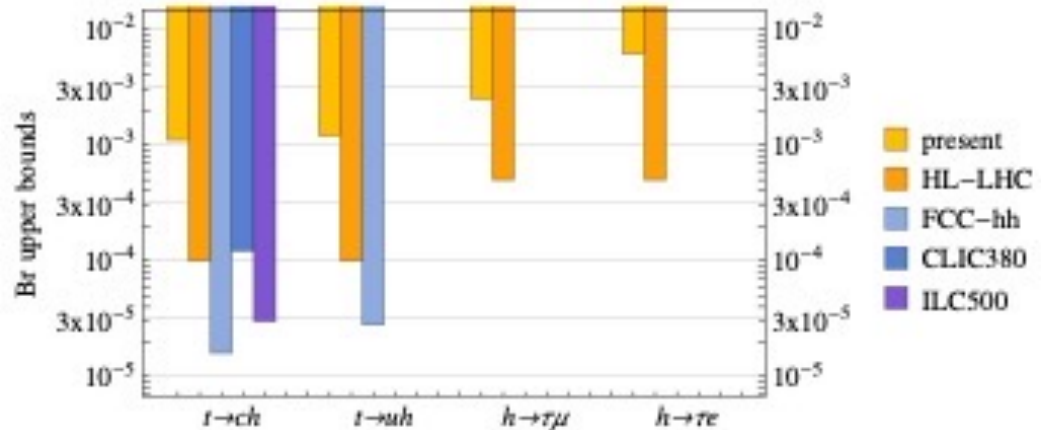
KOTO 2
KLEVER

NA62
(today+guess)

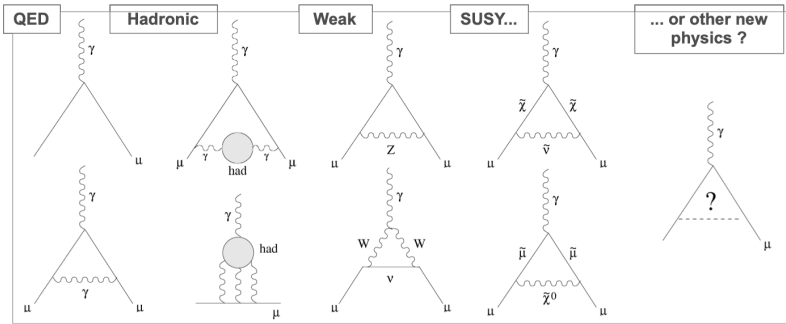
FCNC in top decays

$BR \times 10^5$	HL-LHC	HE-LHC	ILC ₅₀₀	CLIC ₃₈₀	LHeC	FCC-ee	FCC-hh	FCC-eh
$t \rightarrow Zq$	2.4 - 5.8				4	2.4	≈ 0.1	0.6
$t \rightarrow \gamma c$	7.4		≈ 1	2.6			0.024	
$t \rightarrow \gamma u$	0.86						0.018	
$t \rightarrow \gamma q$					1	1.7		0.085
$t \rightarrow gc$	3.2	0.19						
$t \rightarrow gu$	0.38	0.056						

FCNC with higgs decays



Muon $a_\mu = g_\mu - 2$

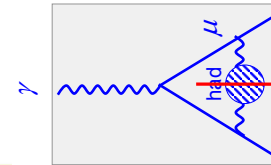


$$a_\mu^{\text{QED}} = \frac{\alpha}{2\pi} + A_2 \left(\frac{\alpha}{\pi}\right)^2 + A_3 \left(\frac{\alpha}{\pi}\right)^3 + A_4 \left(\frac{\alpha}{\pi}\right)^4 + A_5 \left(\frac{\alpha}{\pi}\right)^5$$

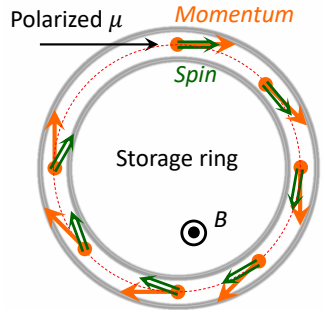
$$a_\mu^{\text{EW}} = 153.6 (1.0) \times 10^{-11}$$

$$a_\mu^{\text{HLbL}} = 94 (19) \times 10^{-11}$$

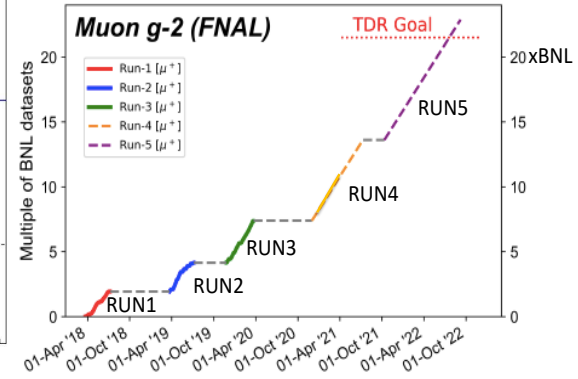
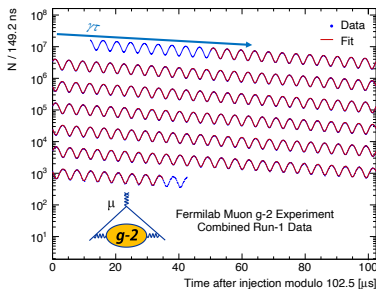
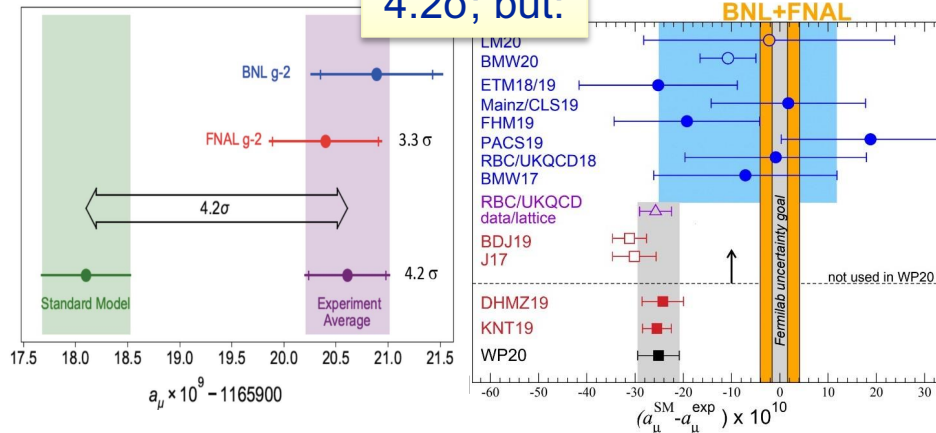
Main uncertainty: Hadronic Vacuum Polarization



Low-energy e^+e^- or Lattice QCD

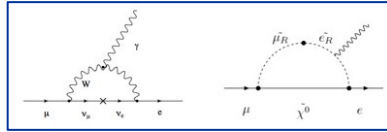


4.2 σ ; but:



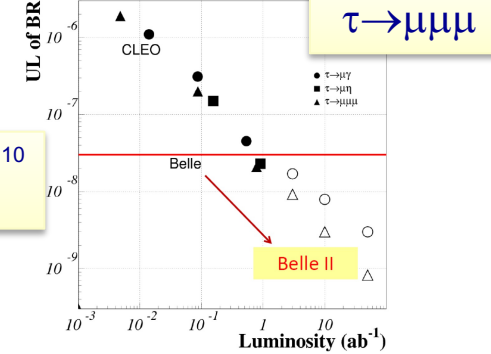
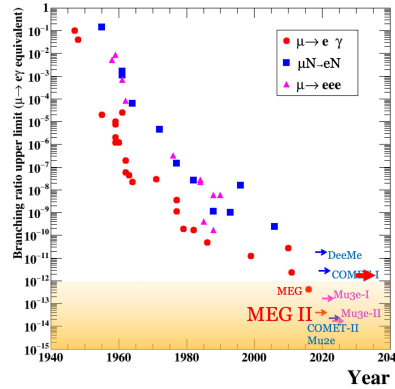
Flavor physics: charged lepton sector & summary

Final MEG upper limit
 $B(\mu \rightarrow e\gamma) < 4.2 \cdot 10^{-13}$
 @90% CL: there
 since 2016...



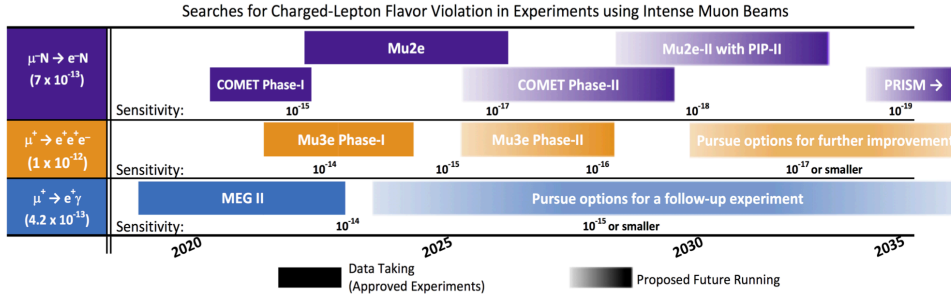
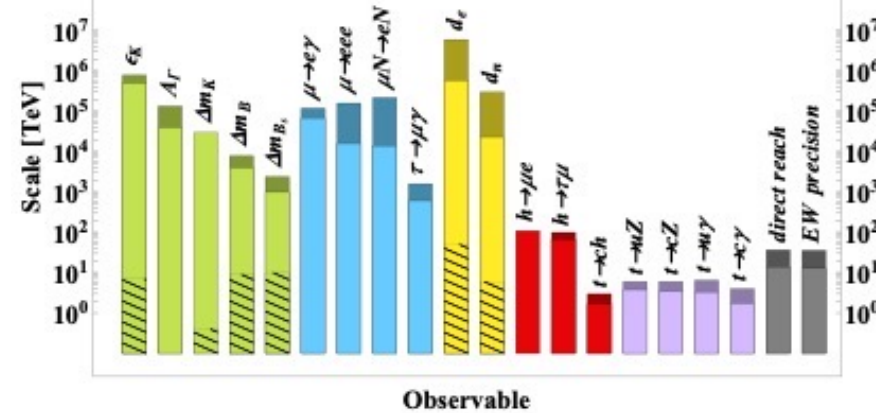
Muon LFV quiet since 2016...
 Exciting times ahead with
 MEGII, Mu2e, COMET, Mu3e...

$\mu \rightarrow e\gamma$
 $\mu \rightarrow eee$
 $\mu N \rightarrow eN$



Expect 5×10^{10}
 $\tau\tau$ evts...

Qualitative summary



Pseudo-summary

And, of course, the “other” issues

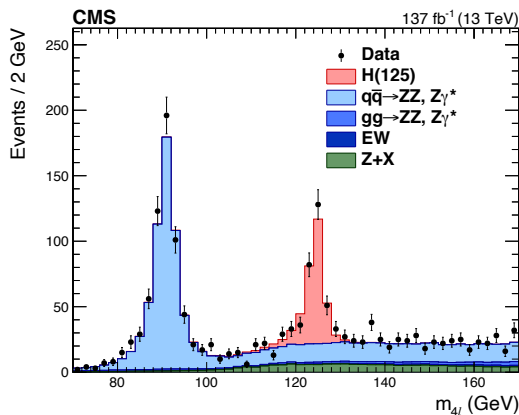
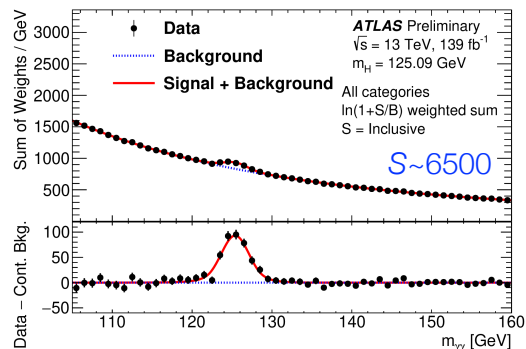
Pseudosummary/Outlook

- **Extremely rich physics program ahead to understand the scalar sector**
 - The LHC and HL-LHC will get us to 3-5% couplings;
 - All options for a Higgs “factory” bring in $\sim O(10^{-2}-10^{-3})$ understanding of couplings.
 - Important aspects: EWPO (needs next-gen Z factory) and top threshold.
 - **Actually, the linear and circular options are quite complementary...**
 - Fundamental scalar? Probe to 15-18 TeV. HE ee and pp colliders offer supplementary information.
 - FCC-ee/hh combination has the largest direct reach to new particles/phenomena. From new particles to Higgs self-coupling to Dark Matter searches...
- **Dark Matter:**
 - Complementarity with indirect searches at colliders (and astroparticle expts).
 - Next-gen colliders can cover the thermal WIMP scenario.
- **Flavor Physics: neutrinos have all one could ask for (albeit it with small σ_{int}). But quark sector may hold the first genuine surprises(?)**
 - In the next 5-10 years: could get very pleasant news on CP front. Definitive statement of mass hierarchy and CP from the full program (DUNE and HyperK).
 - **Again: complementarity also important here.**
 - Quark sector: the current situation will be resolved in the next five years. If current small deviations confirmed: we will have the first experimental evidence for New Physics from the current collider HEP program.
 - **Kaon program complementary; Very important ongoing experiments on charged lepton flavor studies.**
- **It is the best of times; as long as there is (soon) an international move towards a new collider.**

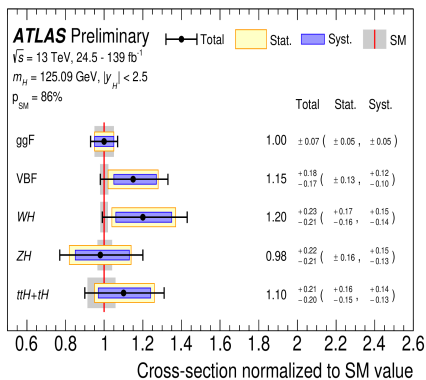
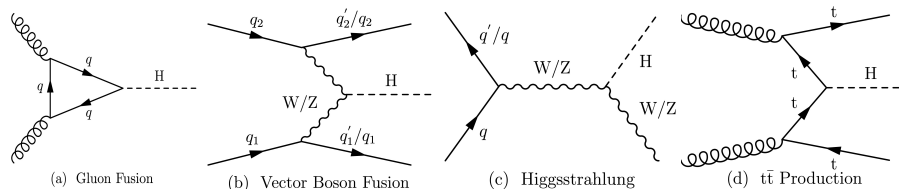
Backups

Higgs physics: probing its couplings/properties

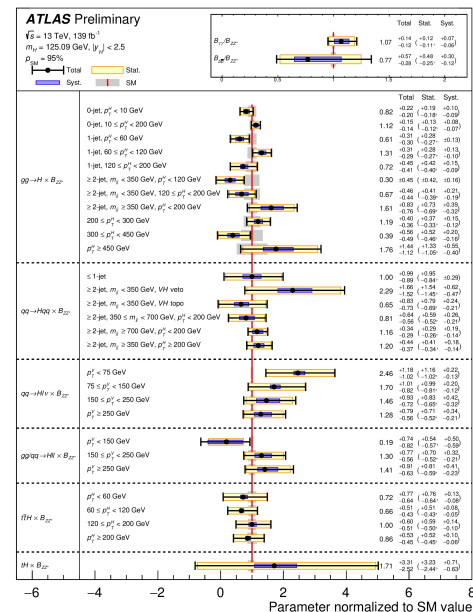
Current LHC:
increased stats...



Production mechanisms, differential distributions



$$\mu = 1.06 \pm 0.07$$



Grand summary

kappa-3 scenario	HL-LHC+								
	ILC ₂₅₀	ILC ₅₀₀	CLIC ₃₈₀	CLIC ₁₅₀₀	CLIC ₃₀₀₀	CEPC	FCC-ee ₂₄₀	FCC-ee ₃₆₅	FCC-ee/eh/hh
κ_W (%)	1.1	0.29	0.75	0.4	0.38	0.95	0.95	0.41	0.2
κ_Z (%)	0.29	0.23	0.44	0.39	0.39	0.18	0.19	0.17	0.17
κ_g (%)	1.4	0.84	1.5	1.1	0.86	1.1	1.2	0.89	0.53
κ_γ (%)	1.3	1.2	1.5*	1.3	1.1	1.2	1.3	1.2	0.36
$\kappa_{Z\gamma}$ (%)	11.*	11.*	11.*	8.4	5.7	6.3	11.*	10.	0.7
κ_c (%)	2.	1.2	4.1	1.9	1.4	2.	1.6	1.3	0.97
κ_t (%)	2.7	2.4	2.7	1.9	1.9	2.6	2.6	2.6	0.95
κ_b (%)	1.2	0.57	1.2	0.61	0.53	0.92	1.	0.64	0.48
κ_μ (%)	4.2	3.9	4.4*	4.1	3.5	3.9	4.	3.9	0.44
κ_τ (%)	1.1	0.64	1.4	0.99	0.82	0.96	0.98	0.66	0.49
BR _{inv} (<%, 95% CL)	0.26	0.22	0.63	0.62	0.61	0.27	0.22	0.19	0.024
BR _{unt} (<%, 95% CL)	1.8	1.4	2.7	2.4	2.4	1.1	1.2	1.	1.

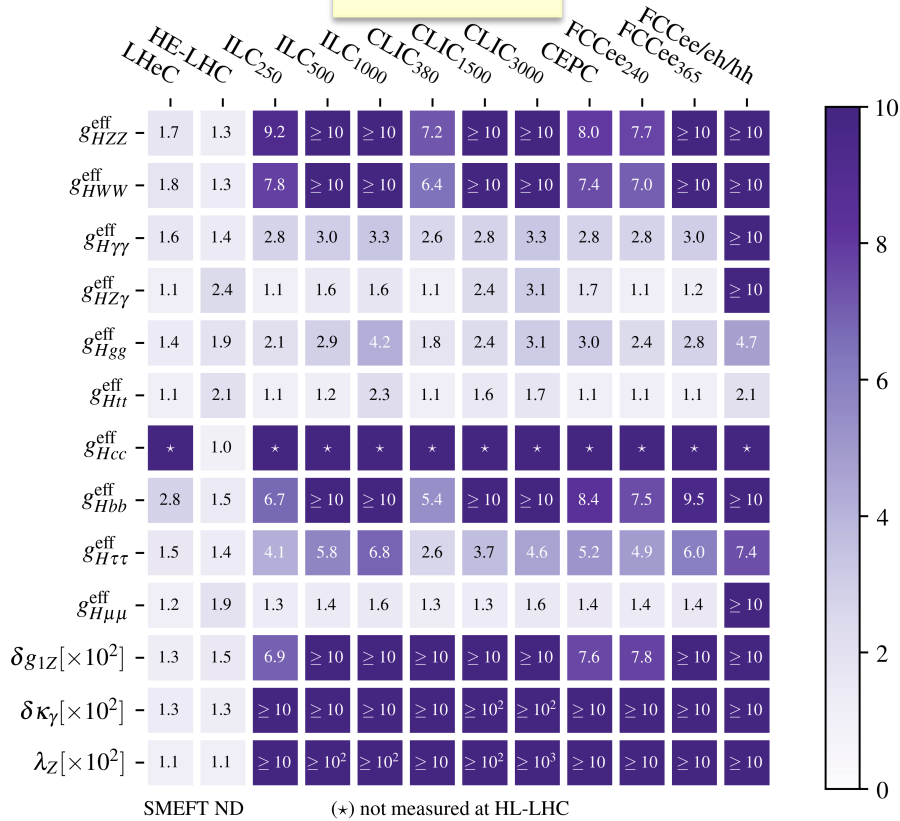
All ee colliders achieve major (and comparable) improvements in their first stage already in probing Higgs sector compared to HL-LHC: at least half of the couplings get improved by factor 5 or more W/Z effective couplings and BR(H → invisible) probed to $\sim 3 \times 10^{-3}$

Model-independent total cross section measurement → access to width, untagged BR

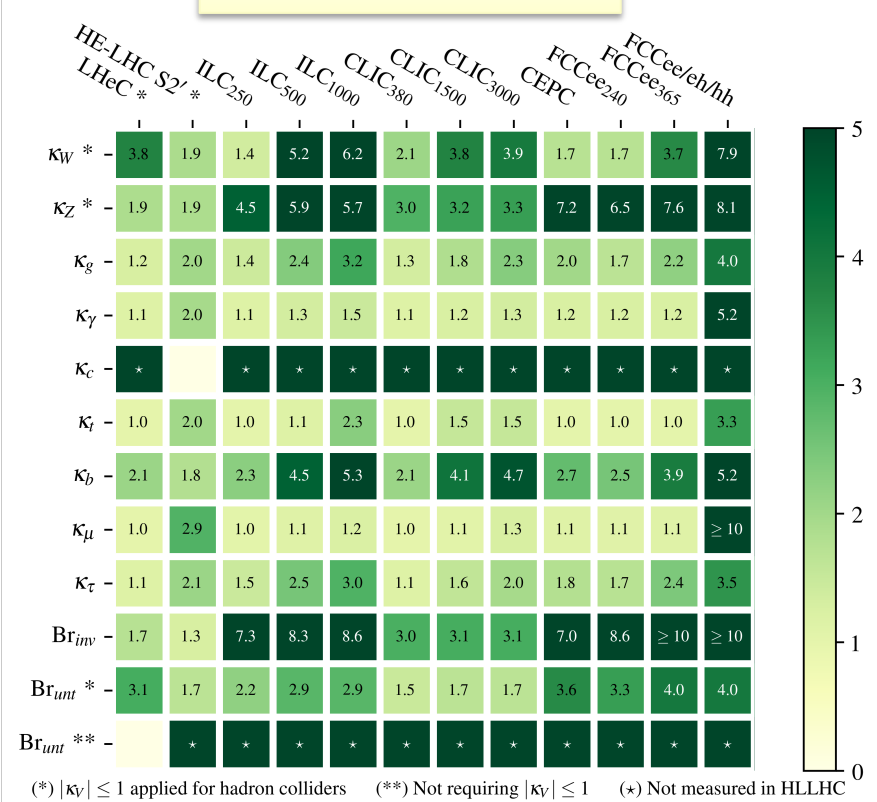
Clean environment to study H if/when anomalies are seen to understand underlying physics

Higgs couplings: SMEFT

Absolute



Gain wrt HL-LHC



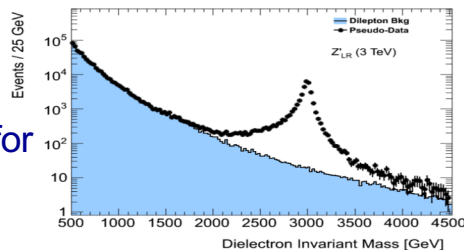
The scalar sector of the SM and direct search for BSM physics

- **Challenge #1: what is the nature of the Higgs boson?**
 - Is it really a fundamental scalar??? Not a composite structure?
 - Are its couplings to fermions and gauge bosons in full accord in with the SM? What about its coupling to itself?
 - How do precision electroweak observables inform us about the Higgs boson properties and/or BSM physics?
- **Challenge #2: are there new interactions or new particles around or above the electroweak scale?**
- **Challenge #3: can particle physics experiments probe the full range of masses/couplings relevant to thermal relic WIMPs**
- **Challenge #3: maximize extent to which experiments probe feebly interacting sectors**

New resonances/particles/forces?

Seeing the peak. Reach:

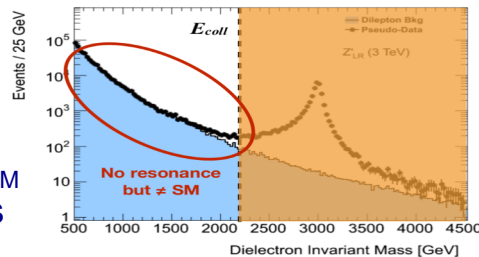
- $M < \sqrt{s}$ for lepton colliders
- $M \lesssim 0.3-0.5 \sqrt{s}$ in hadron colliders for couplings \sim weak couplings



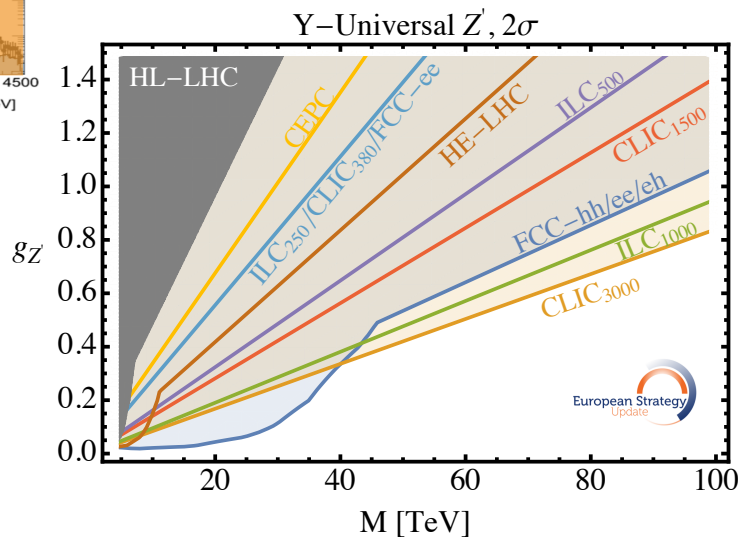
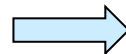
Courtesy:
J. De Blas

Deviations in high-M tails:

- Better suited for lepton colliders; sensitive to $[\text{mass/coupling}] \gg \sqrt{s}$
- Hadron colliders relevant for $g_{Z'} > g_{\text{SM}}$ couplings: $[\text{mass/coupling}] \gg 0.5\sqrt{s}$



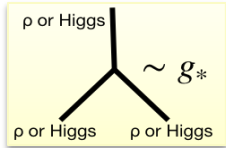
Use very simple model as example. Universal Z' . Clearly, many models with flavor dependence etc.



Higgs Compositeness?

□ Using fits from EWK/Higgs group ([arXiv:1905.03764](https://arxiv.org/abs/1905.03764))

□ Connection between notations:



$$\frac{c_\phi}{\Lambda^2} = \frac{g_*^2}{m_*^2}$$

H

$$\frac{c_W}{\Lambda^2} = \frac{1}{m_*^2}$$

2f-2b

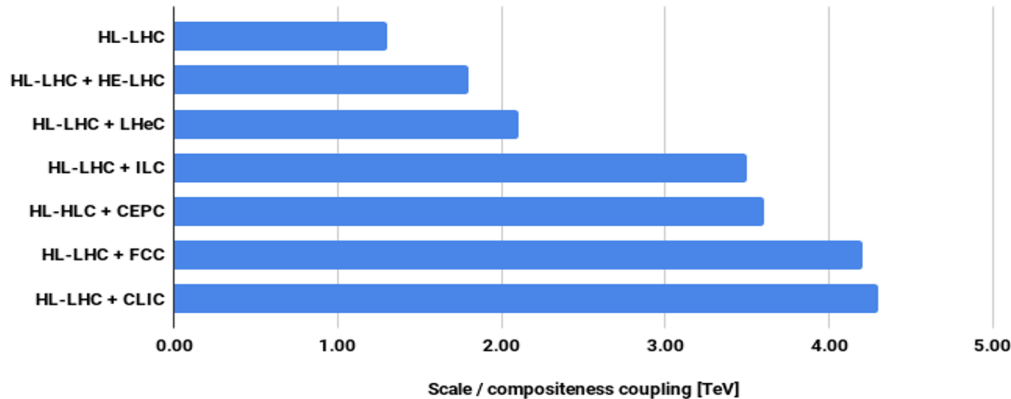
$$\frac{c_{2W}}{\Lambda^2} = \frac{1}{g_*^2 m_*^2}$$

4f, W'/Z'

$$\text{gauge } \rho \sim \frac{g_W}{g_*}$$

□ For mass/coupling ~ 2 TeV \rightarrow deviations $\sim 1\%$ in Higgs couplings

95% CL limits on compositeness scale (O_H operator)



Maximum sensitivities:
@CLIC and
FCC(ee+eh+hh)

Higgs Compositeness → fine-tuning?

□ Corollary question: is it “natural”?

$$(m_H^2)_{Phys.} = \int_0^\infty F_{true}(E; g_{true})$$

$$= \int_0^{\Lambda_{UV}} (\dots) + \int_{\Lambda_{UV}}^\infty (\dots)$$

SM Contribution

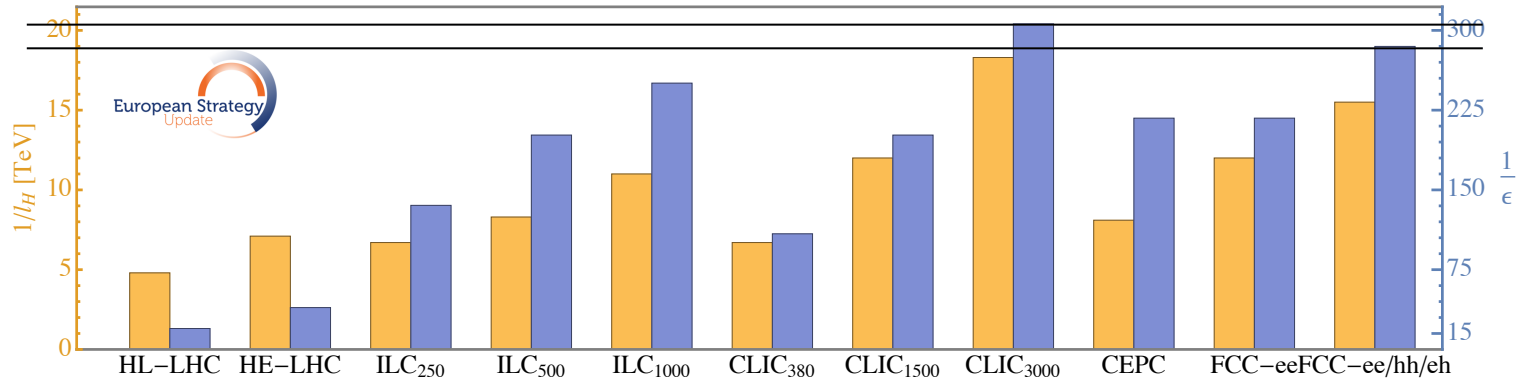
$$\delta m_H^2 = \frac{3y_t^2}{8\pi^2} \Lambda_{UV}^2$$

$$\Delta \geq \frac{\delta m_H^2}{m_H^2} \simeq \left(\frac{126 \text{ GeV}}{m_H} \right)^2 \left(\frac{\Lambda_{UV}}{500 \text{ GeV}} \right)^2$$

HL-LHC: gets to $\Lambda \sim 1.5 \text{ TeV} \rightarrow \Delta > 10$
 (FCC-hh: $\Lambda \sim 5 \text{ TeV} \rightarrow \Delta > 88$)

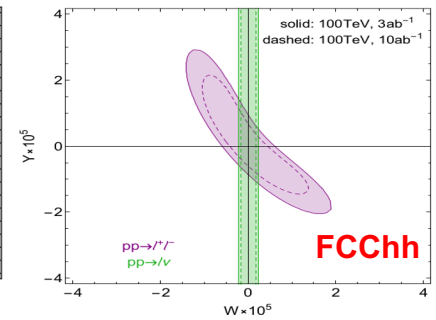
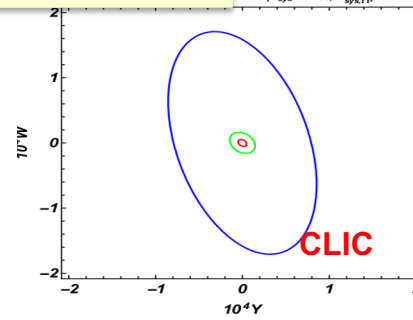
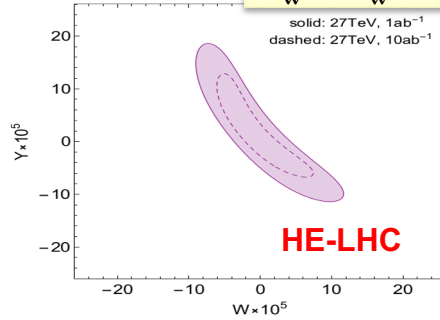
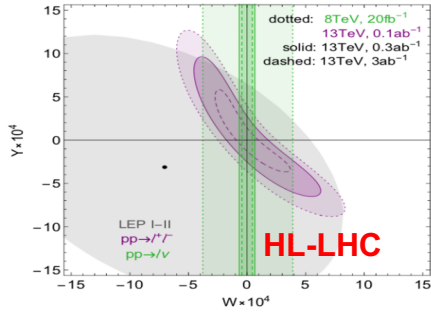
$$\frac{1}{\epsilon} > \left(\frac{m_T}{500 \text{ GeV}} \right)^2 > \left(\frac{m_*}{g_* v} \right)^2$$

$$\frac{c_\phi}{\Lambda^2} = \frac{g_*^2}{m_*^2} = \frac{\epsilon}{v^2} \rightarrow \frac{1}{\epsilon} = \frac{1}{v^2} \left(\frac{c_\phi}{\Lambda^2} \right)^{-1}$$



Contact Interactions

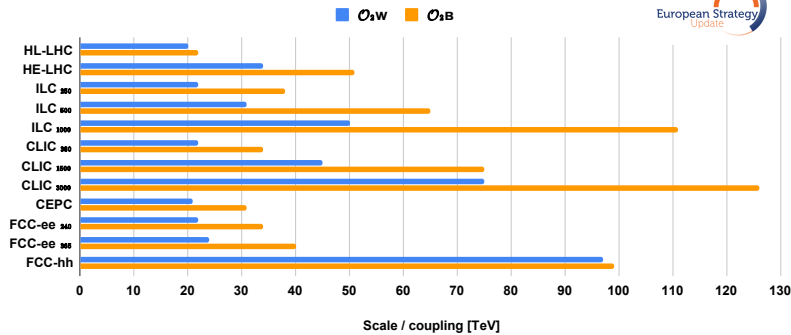
$$\frac{g^2 W}{m_W^2}, \frac{g'^2 Y}{m_W^2} \rightarrow \Delta \left(\left[\frac{\text{coupling}}{\text{Scale}} \right]^2 \right)$$



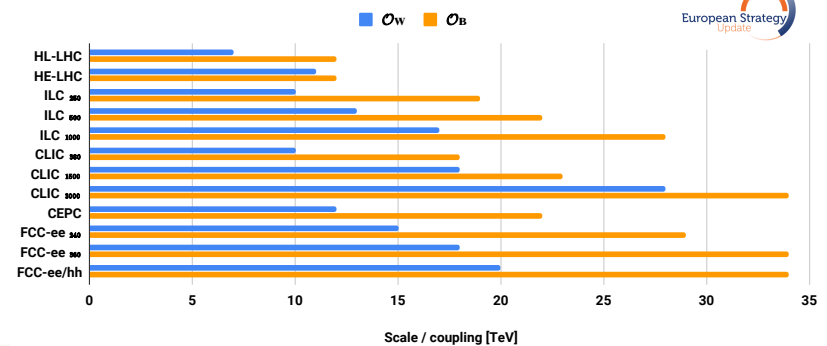
ee: all difermion final states, angular analysis (+polarization)
pp: Drell-Yan Z .AND. W

CLIC: ZH, angular analysis
FCC: $p_T(Z)$ in WZ production

95% CL scale limits on 4-fermion contact interactions



95% CL scale limits on 2-fermion 2-boson contact interactions

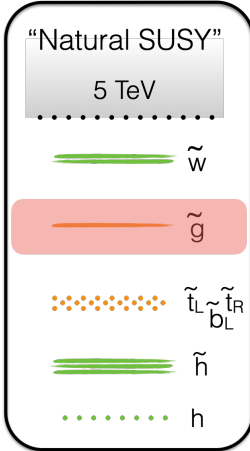
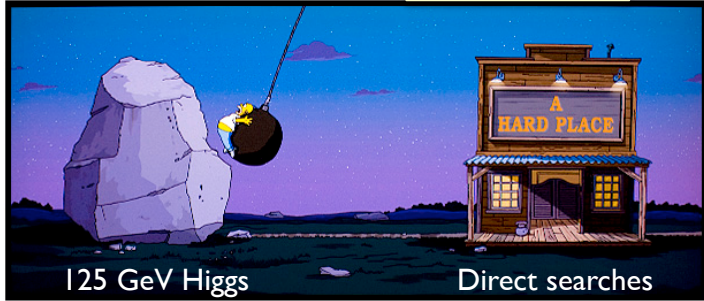


Sensitivity for ee colliders enhanced for couplings $\gtrsim 1$
(weak couplings \rightarrow direct searches become more sensitive)
Searches for W' & charged fermion currents more effective at hadron colliders

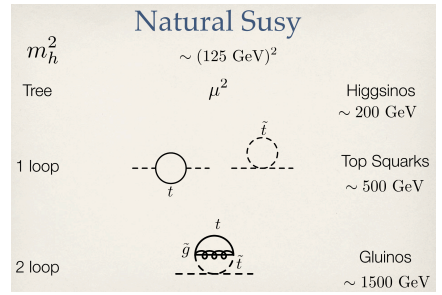
SUSY: what does it mean?

Corollary question: is SUSY “natural”?

Credit: D. Shih

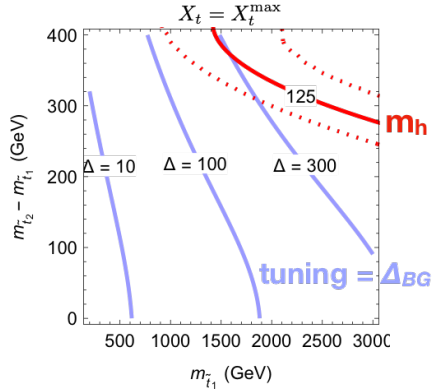


Fine tuning parameter: $\frac{1}{\epsilon} = \frac{\Delta m_h^2}{m_h^2}$



$$\frac{1}{\epsilon_{\tilde{t}}} = \frac{3y_t^2 m_{\tilde{t}}^2}{2\pi^2 m_h^2} \ln(\Lambda/m_{\tilde{t}})$$

$$\frac{1}{\epsilon_{\tilde{g}}} = \frac{4y_t^2 \alpha_s m_{\tilde{g}}^2}{\pi^3 m_h^2} \ln^2(\Lambda/m_{\tilde{g}})$$



ϵ	$\ln(\Lambda/m_{\tilde{t},\tilde{g}}) \approx 30$ High-scale mediation	$\ln(\Lambda/m_{\tilde{t},\tilde{g}}) \approx 1$ Low-scale mediation
stop	$5 \times 10^{-5} \left(\frac{10 \text{ TeV}}{m_{\tilde{t}}}\right)^2$	$2 \times 10^{-3} \left(\frac{10 \text{ TeV}}{m_{\tilde{t}}}\right)^2$
gluino	$7 \times 10^{-6} \left(\frac{17 \text{ TeV}}{m_{\tilde{g}}}\right)^2$	$6 \times 10^{-3} \left(\frac{17 \text{ TeV}}{m_{\tilde{g}}}\right)^2$

MSSM: already unnatural

Learn also (though less) from indirect information

$$r_g - 1 \approx \frac{1}{4} \frac{m_{\tilde{t}}^2}{m_{\tilde{t}_1}^2} \approx 0.7\% \left(\frac{1 \text{ TeV}}{m_{\tilde{t}_1}^2}\right)^2$$

$$\frac{\delta \mathcal{O}_{\text{SUSY}}}{\mathcal{O}_{\text{SM}}} \sim \frac{m_{\text{SM}}^2}{m_{\text{SUSY}}^2}$$

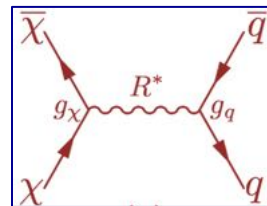
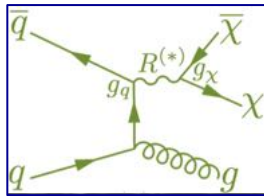
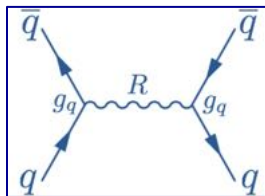
Thermal relic WIMPs (?)

- **Motivation for direct, indirect and collider searchers:**

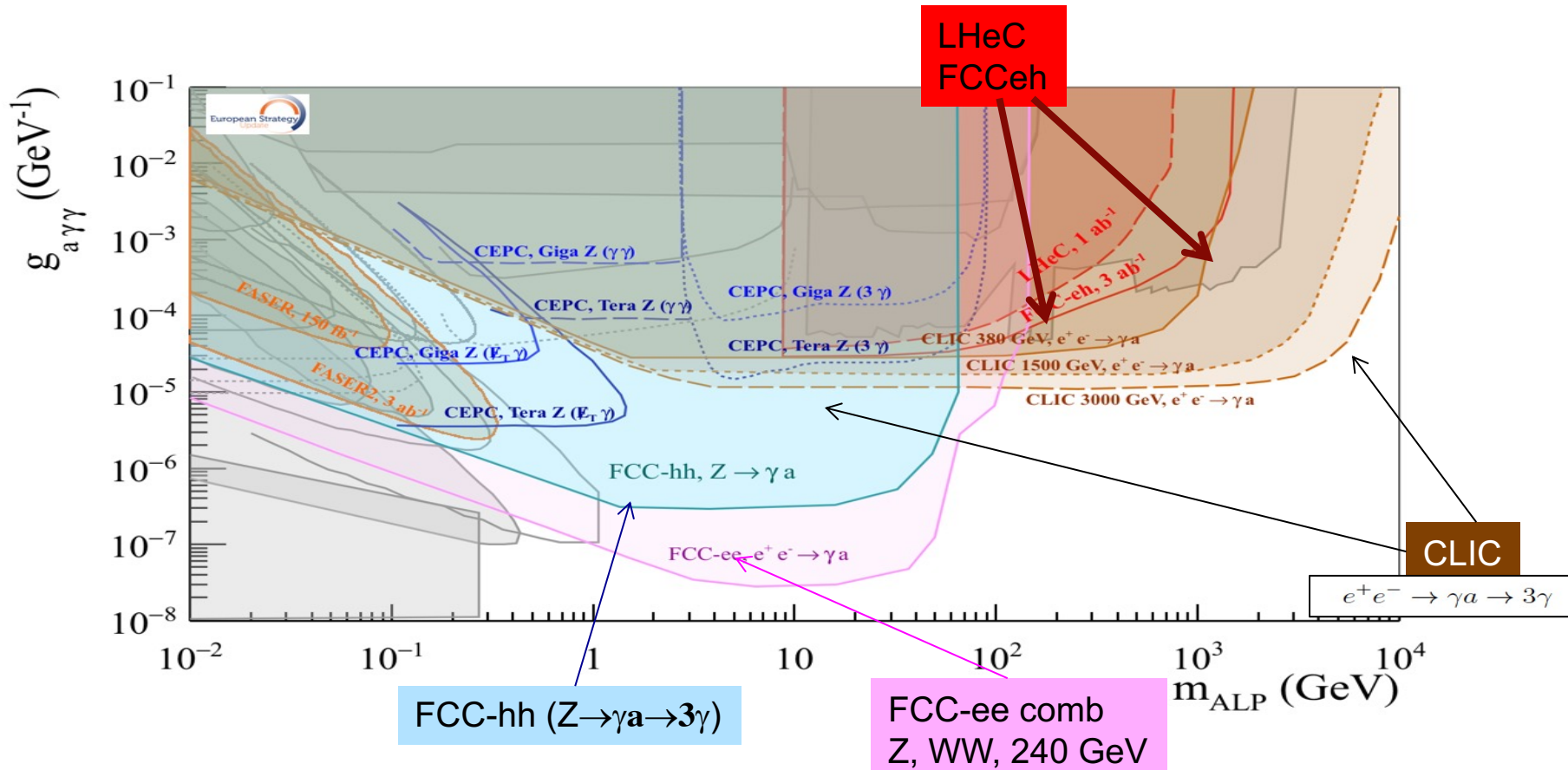
$$\Omega_{\text{DM}} h^2 \sim 0.12 \times \left(\frac{M_{\text{DM}}}{2 \text{ TeV}} \right)^2 \left(\frac{0.3}{g_{\text{eff}}} \right)^4 \implies M_{\text{DM}} \sim \mathcal{O}(\text{few GeV}) \rightarrow \mathcal{O}(10\text{'s TeV})$$

- **WIMP miracle has been moving (moved) upwards – to ~TeV.**
- **Focus of these searches: GeV–TeV region; two classes**
 - **Classic electroweak WIMP candidates (SUSY inspired)**
 - **Winos and Higgsinos (and linear combinations...)**
 - **Simplified models with mediator particles**
 - **Axial-vector simplified models**
 - **Scalar simplified models**

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\text{DM}} + \mathcal{L}_{\text{Int}}$$



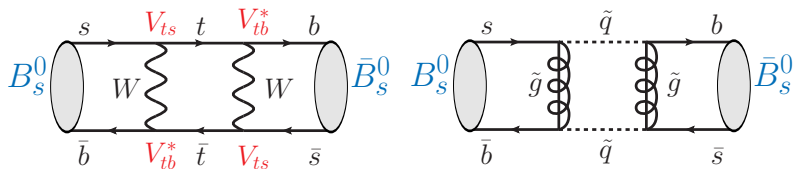
FIPs: Pseudo-Scalar Portal (Axions, ALPs)



FIPs represent a new paradigm that requires systematic exploration on multiple fronts

Flavor physics: quark sector (II)

Very rich B/D program (Belle II & LHCb, some ATLAS-CMS) in the next ten years.



Observables: γ_{CKM} ;
 φ_S ; φ_D ;
 $B_d \rightarrow \mu\mu / B_s \rightarrow \mu\mu \dots$

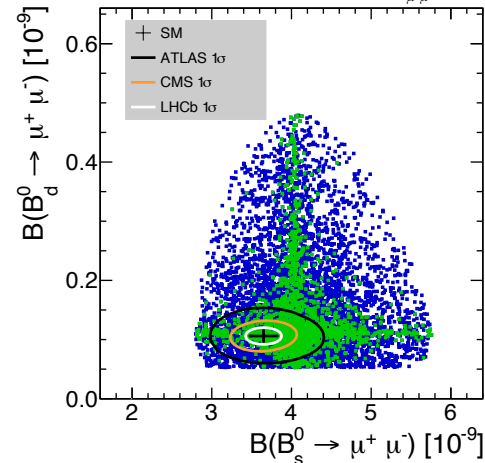
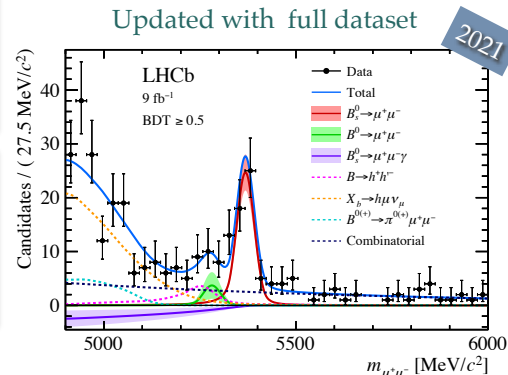
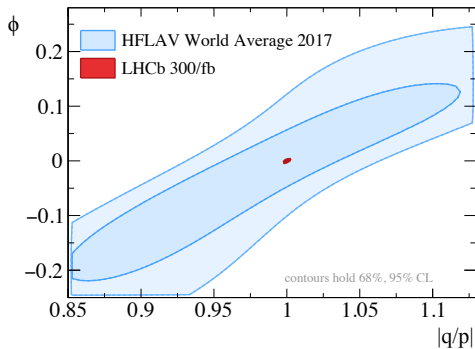
Next frontier in CP precision : D

$$A_{raw} = \frac{N(D^0 \rightarrow f) - N(\bar{D}^0 \rightarrow f)}{N(D^0 \rightarrow f) + N(\bar{D}^0 \rightarrow f)} = A_{CP} + A_D(\pi_s^+/\mu) + A_P(D^{*+}/D_{from B}^0)$$

FCC-ee:
 5×10^{12} Z;
 7.5×10^{11} bb

$D^0 \rightarrow KK, \pi\pi$; observation (5.3σ) $\Delta A_{CP} = -(0.185 \pm 0.029)\%$ arXiv:1903.08726

$\pm 80.0 \times 10^{-5}$	$\pm 96.0 \times 10^{-6}$	$\pm 14.0 \times 10^{-5}$	LHCb
Current			
$\pm 46.0 \times 10^{-5}$		$\pm 12.0 \times 10^{-5}$	Belle II
$\pm 32.0 \times 10^{-5}$	$\pm 40.0 \times 10^{-6}$	$\pm 6.2 \times 10^{-5}$	LHCb
2025			
$\pm 8.0 \times 10^{-5}$	$\pm 8.0 \times 10^{-6}$	$\pm 1.4 \times 10^{-5}$	HL-LHC
$D^0 \rightarrow K^\pm \pi^\mp$	$D^0 \rightarrow K^\mp \pi^\pm \pi^+ \pi^-$	$D^0 \rightarrow K_s \pi^+ \pi^-$	



Where we stand

- **Most successful Theory ever: Standard Model of Particle Physics**
 - **Highest priority:** *extend understanding of SM and its newly discovered scalar sector.*
- **Evidence that SM is either incomplete or an Effective Theory of some Ultimate Theory (or another step in a series of Effective Theories...)**
 - **Experimental evidence**
 - **Dark Matter (DM):** The “thing” we know the least about: Unknown nature, unknown number(s) of DM particles, unknown mass range(s) – 10^{80} .
 - **Neutrino masses:** SM gauge group allows for Majorana masses, “explaining” their tiny values; unknown (putative) Majorana scale.
 - **Matter-antimatter asymmetry of the Universe:** nowhere near what we measure in CP violation experiments. What about CP violation in the lepton sector?
 - **Theory issues**
 - **Electroweak (EW) hierarchy problem.** Why is the Higgs so light?
 - **The flavour puzzle.** Why three generations of quarks and leptons? With very different masses and mixings? Size of CP violation? (explain matter universe???)
 - **And lots more...** e.g. Strong CP problem. Another vacuum issue (this time QCD). Why is its θ parameter experimentally constrained to be extremely small? For a priori no good reason.... → Axions?