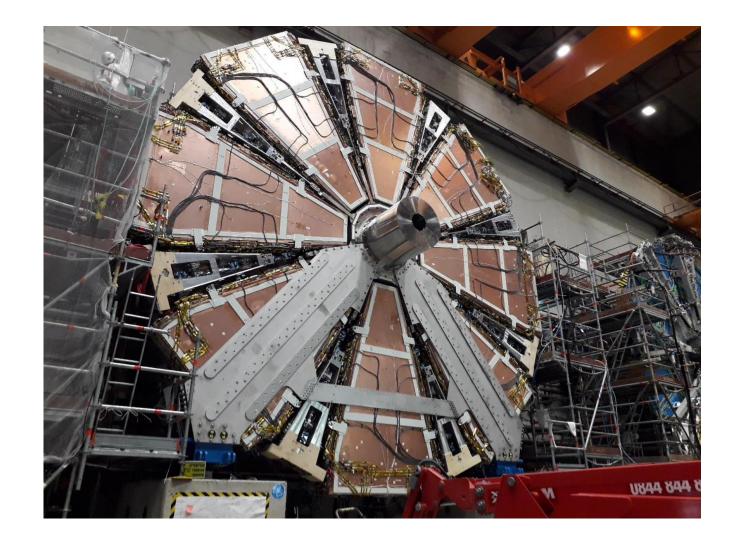
DE LA RECHERCHE À L'INDUSTRIE





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### LM1 production report – RD51 Meeting – 06/21

M Vandenbroucke on behalf of F. Jeanneau for the Saclay team



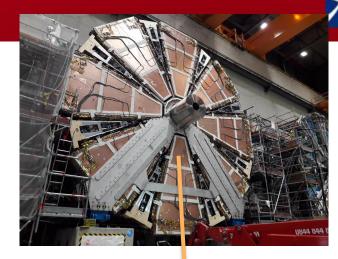
#### STATUS OF LM1 PRODUCTION TODAY

- 1 mechanical module (2018)
- 32 standard production modules already at Cern
- First spare module shipped last week

#### Still to be produced:

- Second spare module (additional PCB needed to build Eta panel)
- 1 stereo doublet for aging studies.

- 177 panels built and measured
- → 108 drift panels (36 central, 72 external)
- → 144 mesh stretched and transferred
- → 69 RO panels (35 Stereo and 34 Eta)



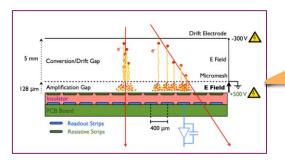




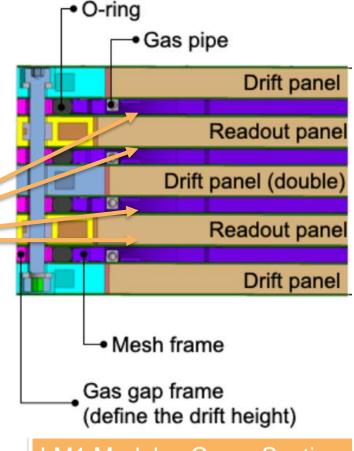
# STEPS FOR MODULE CONSTRUCTION AND VALIDATION



- Panel building and planarity measurements of panels
- Readout Panels Passivation
- Washing and Drying
- Drift Panels Preparation
- Module Assembly
- HV Tests
- Gas Tightness
- Alignment (Rasfork) Meas.
- Module Planarity
- Conditioning/Drying of Module
- Test in Cosmic rays



1 Module = 5 Pannels = 3 Drifts +2 Readout = 4 MM gaps



LM1 Modules Cross-Section

From: "The large inner Micromegas modules for the Atlas MuonSpectrometer Upgrade: construction, quality control andcharacterization" https://arxiv.org/pdf/2105.13709.pdf



### PANEL PLANARITY



Pannel building is glueing PCB on honey comb + bars

Planarity of each panel is measured in 3 steps:

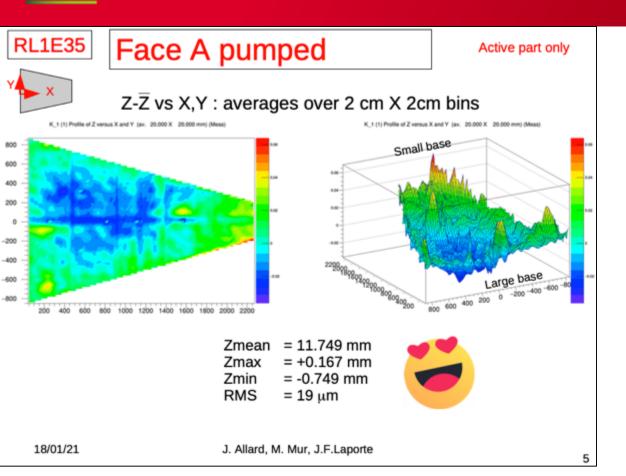
- Face A vacuum ON
- Face A on shimes
- Face B on shimes

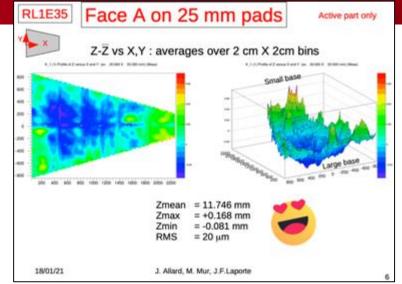


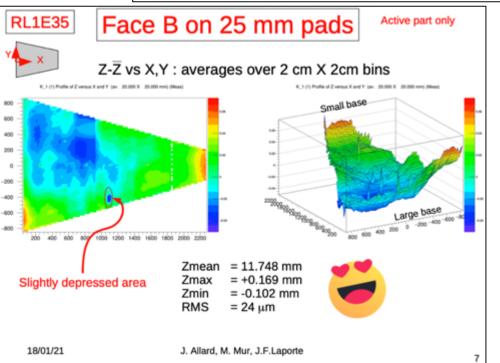


### PANEL PLANARITY: MEASUREMENT SET





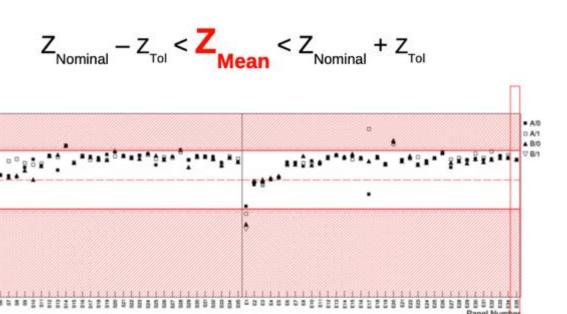




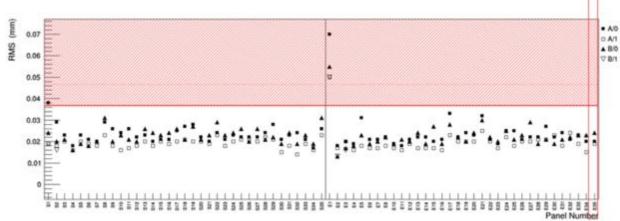


## **RO PANEL PLANARITY**





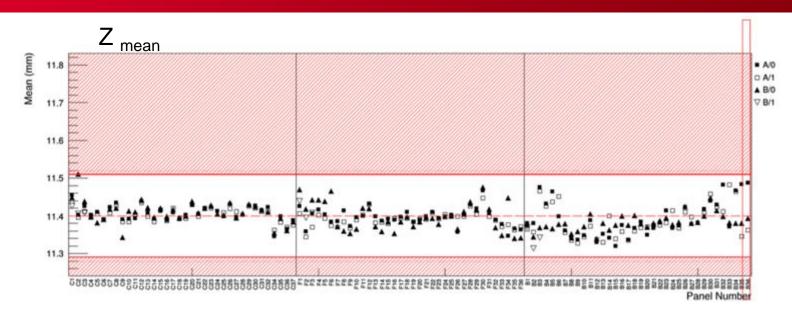


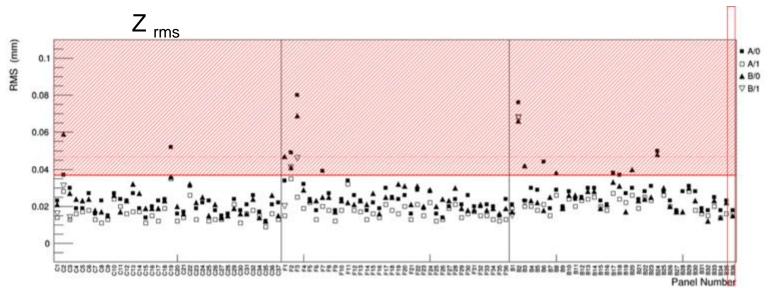




## **DRIFT PANEL PLANARITY**







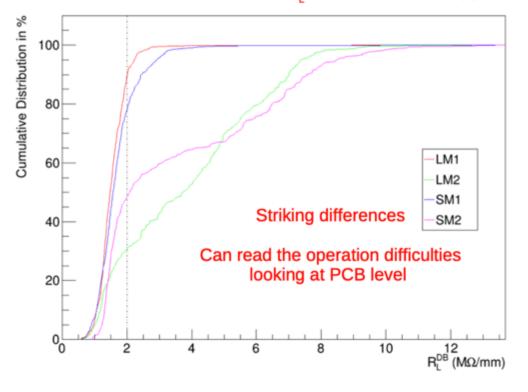


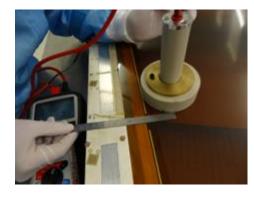
## RO PANEL PREPARATION - PASSIVATION

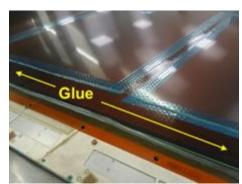


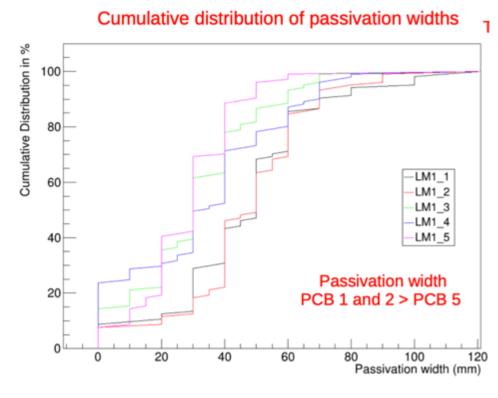
- Panel finalization → FEB pin and alignement pin gluing
- Panel passivation: due to a lack of resistivity of resistive anodes and flaws in the layout of resitive strips, the edge of the panels, where resistance values are lower, must be passivated by applying a thin layer of glue.

#### Cumulative distribution of R, DB for all modules types









From JF Laporte



## **WASHING AND DRYING**



All panels are washed (sometimes several times) and dried



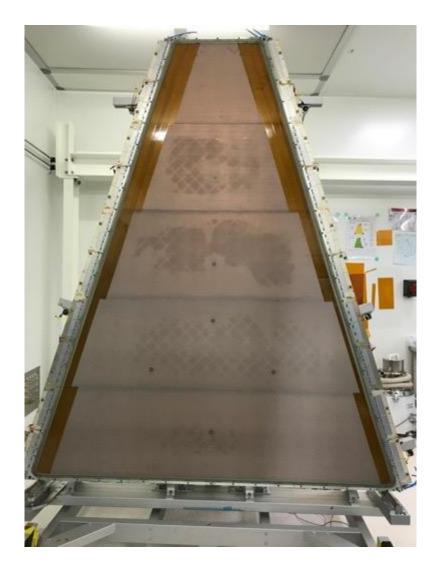




#### **DRIFT PANEL PREPARATION**



- Mesh tension is checked and must stay between 7 and 10 N.m
- Kapton tape is applied on the mesh to match the glue passivation on the corresponding RO panel.
- Electrical insulation between mesh and drift electrode is checked.

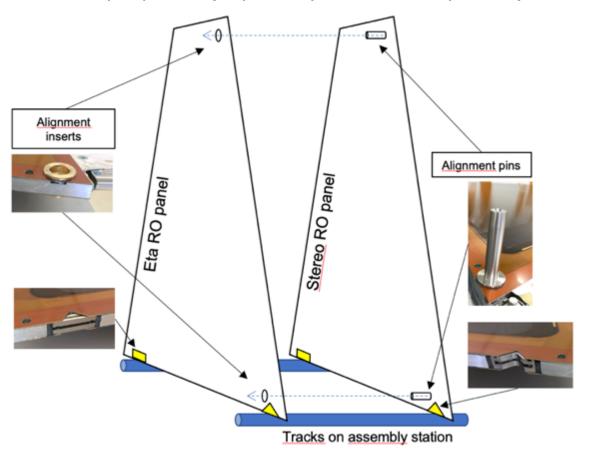


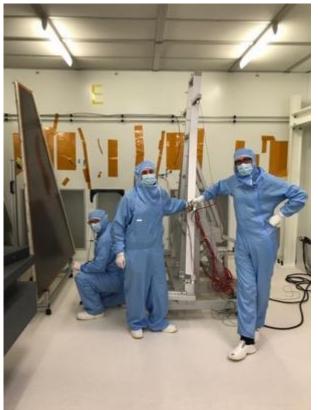


#### **MODULE ASSEMBLY**



- Modules are assembled by stacking 5 panels (2 RO and 3 drift) onto the assembly station.
- In our case the alignment is done by construction, using the mechanical inserts precisely glued during the panel construction on the marble table.
- Each panel is carefully checked and cleaned: vacuum-cleaned, antistatic roller and a lot of isopropanol (especially for the RO panels)





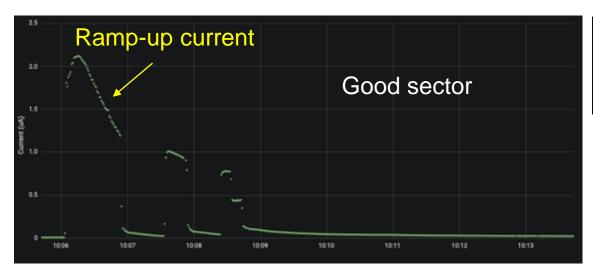


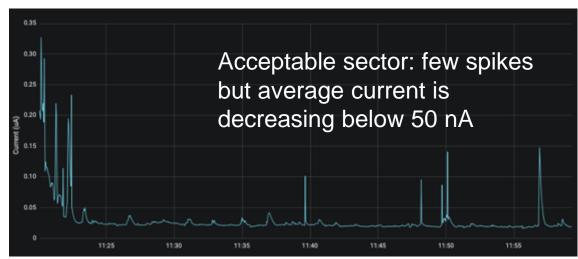


### **HV** TESTS



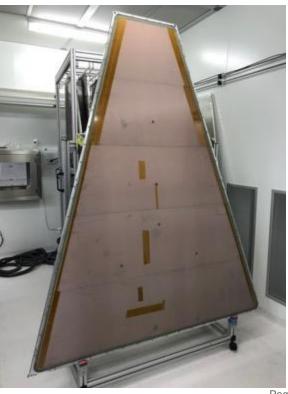
- Each module is assembled gap by gap
- Each gap is validated if it sustains 850V in air with low dark current (<50nA)





- If too much current → new dry cleaning
- If short → kapton method (use of large kapton foils to neutralize sector and localize the weak area by dichotomy)





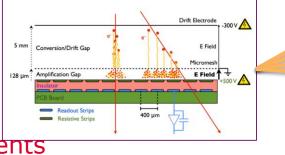
Page 1



# STEPS FOR MODULE CONSTRUCTION AND VALIDATION



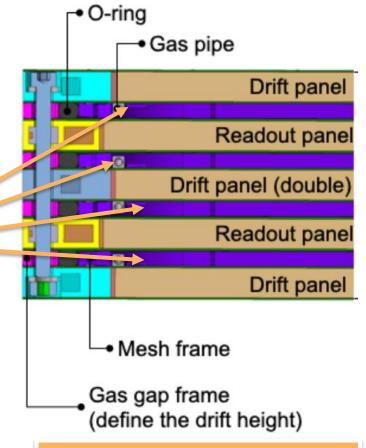
- Panel building and planarity measurements of panels
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- Alignment Rasfork Measurements
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1 Module = 5 Pannels

= 3 Drifts +2 Readout

= 4 MM gaps



LM1 Modules Cross-Section

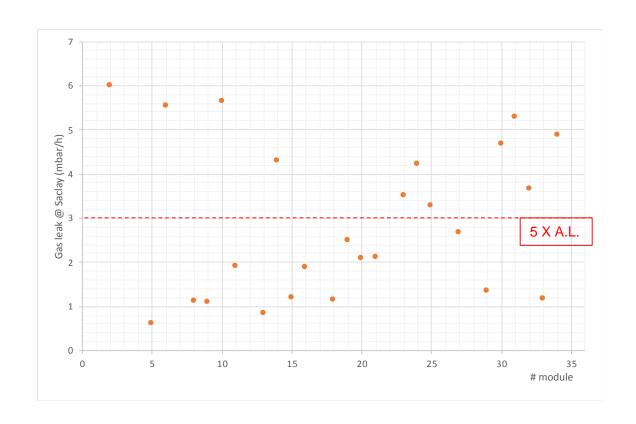
From: "The large inner Micromegas modules for the Atlas MuonSpectrometer Upgrade: construction, quality control andcharacterization"

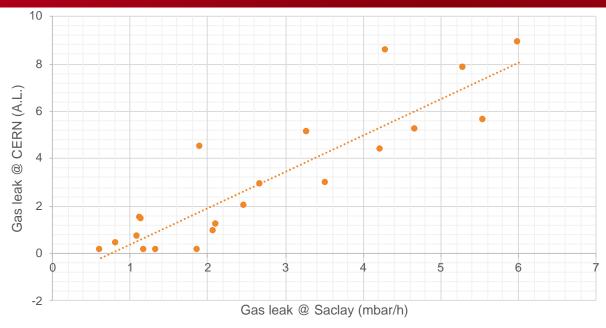
https://arxiv.org/pdf/2105.13709.pdf

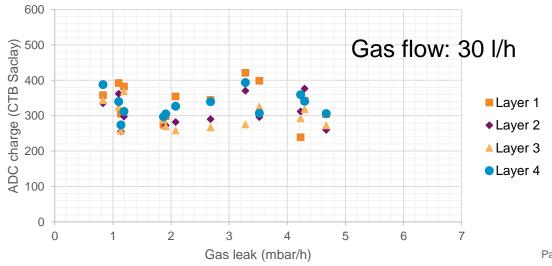


## **GAS TIGHTNESS**







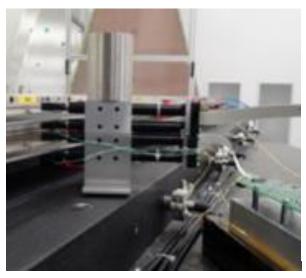


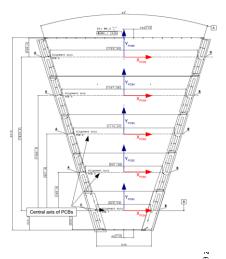


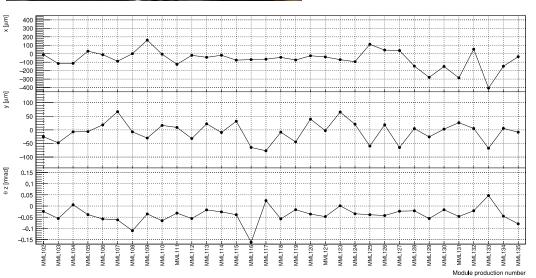
## ALIGNEMENT (RASFORK) MEASUREMENTS



- Rasfork has been developed at Saclay for doublet and quadruplet alignment checks.







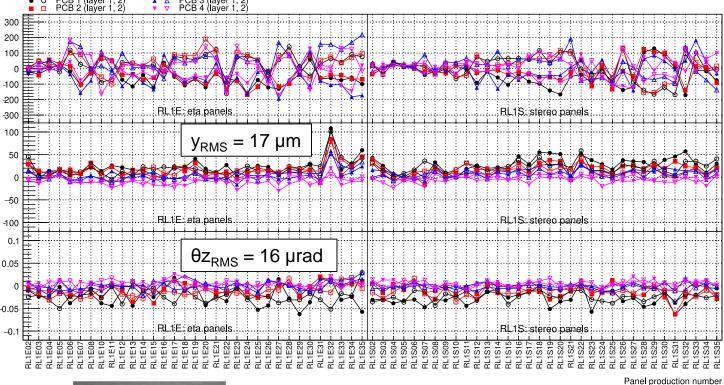




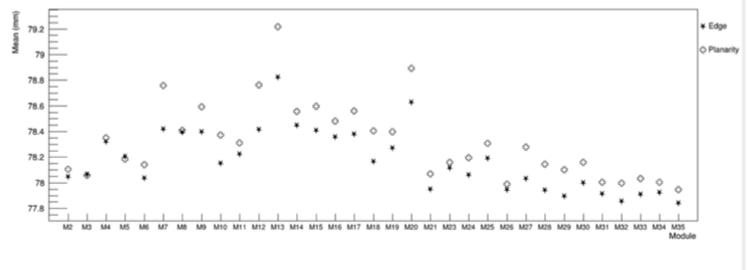
Figure 23: A Rasnik mask etched on PCB.

From PF Giraud



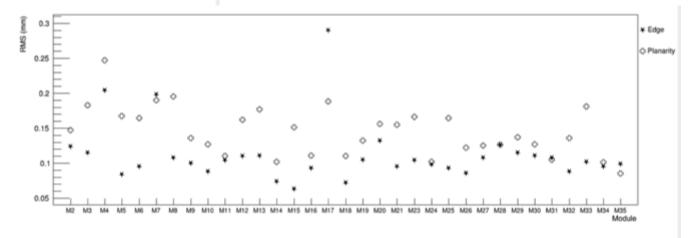
## **MODULE PLANARITY**





Edge: from 78.173 - 0.333 Core: from 78.318 - 0.373

78.173 + 0.649 78.318 + 0.899



Edge:  $\leq 0.290$ Core:  $\leq 0.247$ 

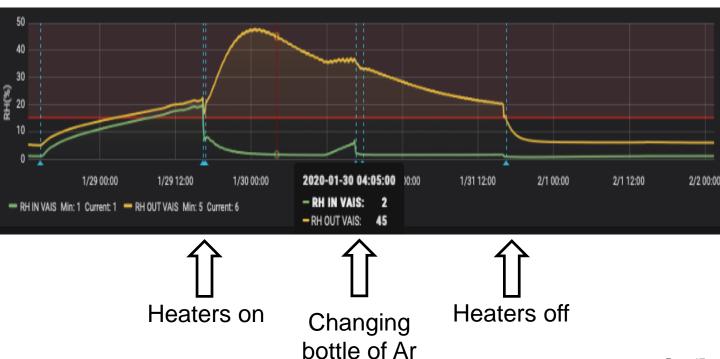


### CONDITIONING



- After module closure, conditioning is needed before applying HV in Ar CO2 mixture.
- Few days at 45°C under gas flow to remove humidity trapped in the material (essentially FR4 and kapton)
- Transfer on the cosmic bench then HV ramping-up for one day (controlled by software)





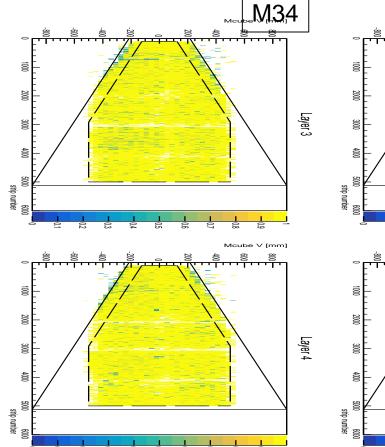


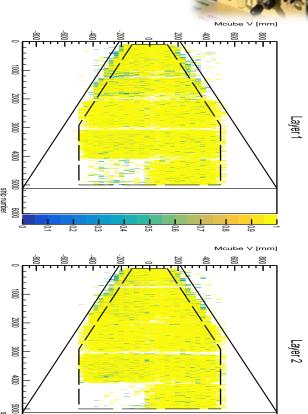
## **TESTS IN COSMICS**

# Plofu

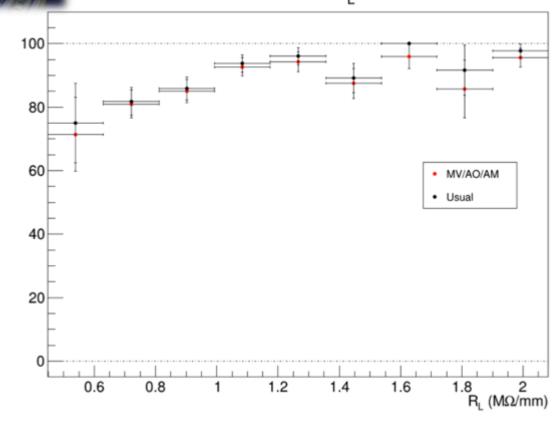
#### M34:

- Trip of L1R5 and L2R5 during test
- All sectors good at BB5





Fraction of good sectors - R Probe - Both Sides



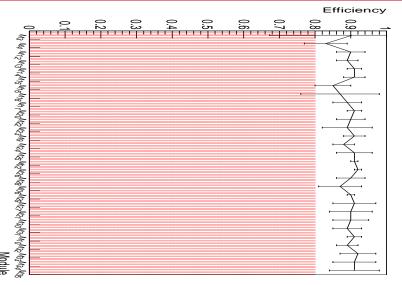


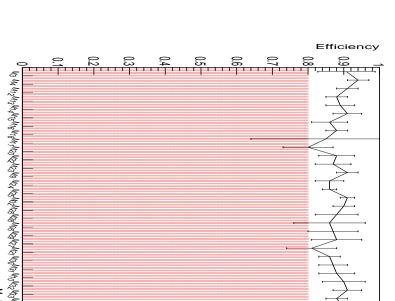
## **MODULE EFFICIENCIES**

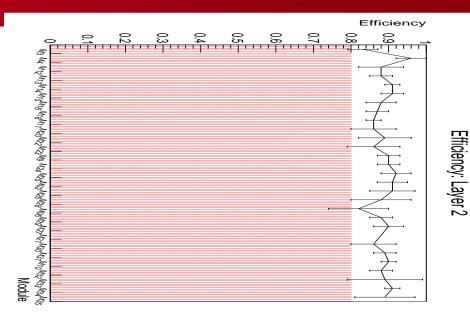
Efficiency: Layer 1

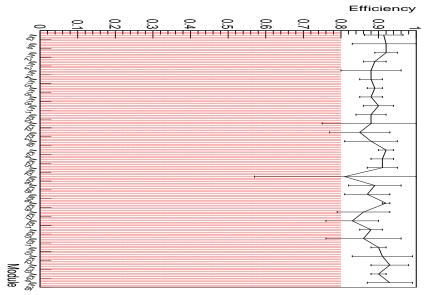
Efficiency: Layer 3











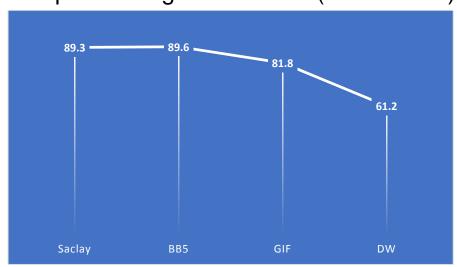
Efficiency: Layer 4



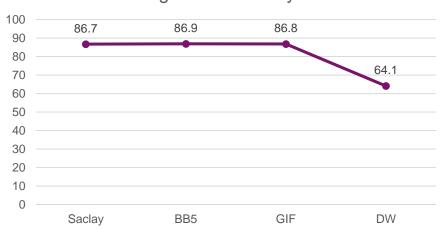
### **IMPORTANT DECONDITIONING WITH TIME**



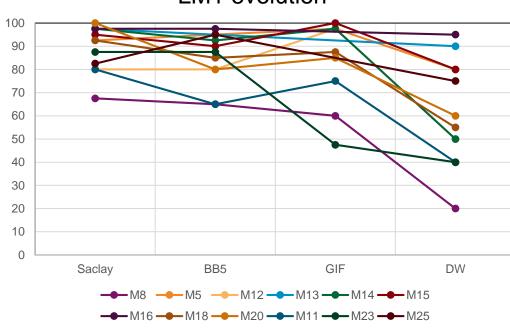
Proportion of good sections (HV > 550 V)



Average HV Efficiency LM1







- Main difference between Saclay and Cern is the time used for conditioning (drying and HV ramp-up)
- Gas flow without heating is not enough to remove humidity
- No storage under gas flow at Cern (humidity!)
- Continuous gas flow on the wheel → margin for improvement



#### CONCLUSION



- 32 LM1 modules + 1 spare have been produced and delivered at Cern
- Still one spare to produce and one Stereo doublet for studies
- This is (almost) the end of a very complicated production
- Despite the technical problems (mainly due to low resistivity) this technology is well adapted for large area. For future, a stronger prototype policy should be considered and avoid success-oriented planning.
- The work of the Isobutane group carries a large hope to operate these detectors in a more secure mode
- Special thanks to our Frascati colleagues who came several times at Saclay to help us during production



Thank you!