

Towards an all-orders calculation of the electroweak bubble wall velocity

Jessica Turner

Institute of Particle Physics Phenomenology, Durham University

Gravitational Probes of Physics Beyond the Standard Model



First order electroweak phase transition

Within the Standard Model the EWPT is a crossover

D'Onofrio & Rummukainen (2015)

Minimal new physics → first order PT

Anderson & Hall (1992)

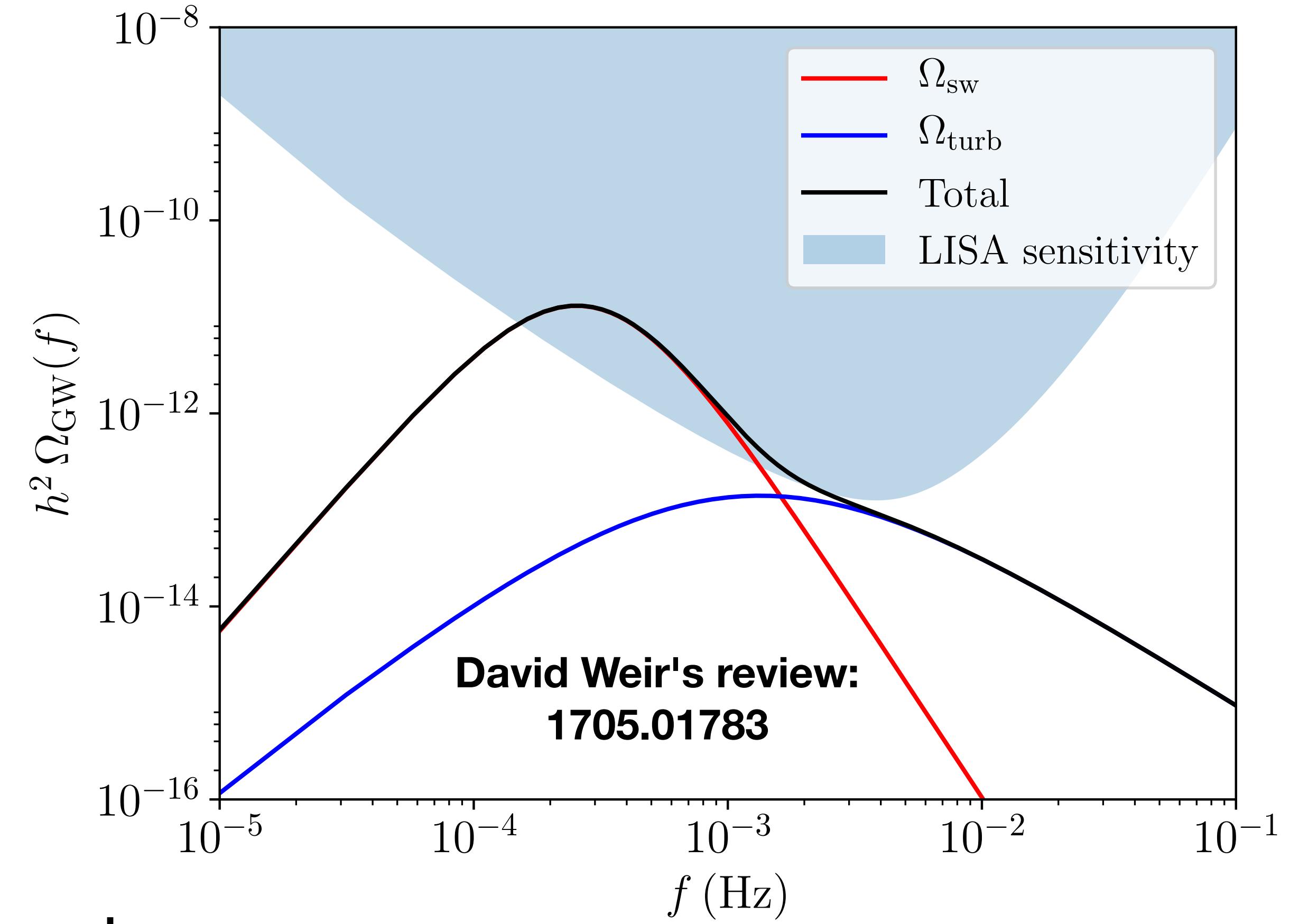
[See overview by Caprini](#)

- Matter-antimatter asymmetry Kuzmin, Rubakov & Shaposhnikov (1985)
- Topological defects Achucarro & Vachaspati (2000)
- Primordial magnetic fields Vachaspati (1991)
- Stochastic gravitational wave background Kamionkowski, Kosowsky & Turner (1993)

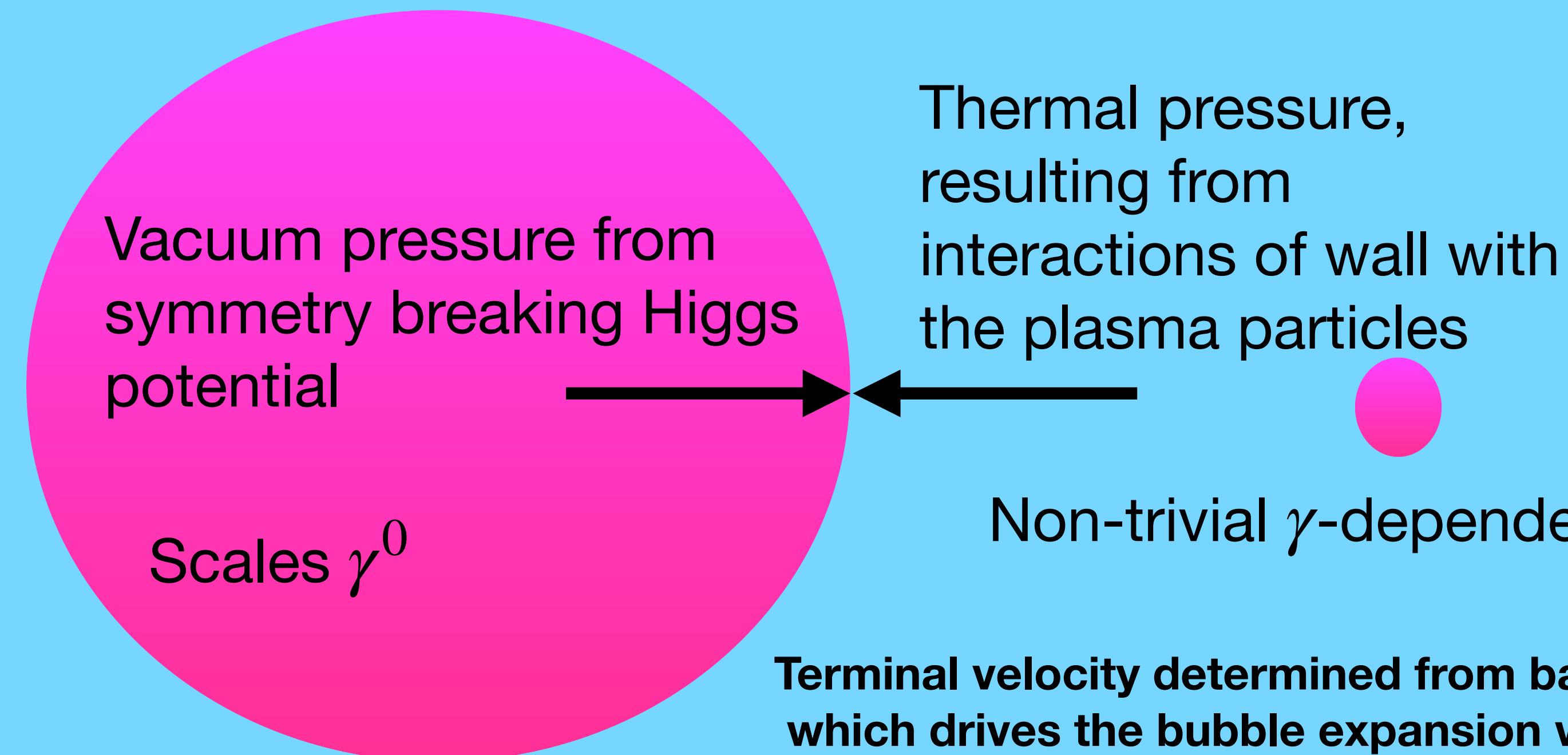
[See talks by Weir, Lewicki, de Vis & Pol](#)

First order electroweak phase transition

- Collision bubble walls: Ω_{env}
- Sound waves: Ω_{sw}
- Turbulence: Ω_{turb}
- Velocity affects gravitational wave spectrum
- Velocity affects efficiency of electroweak baryogenesis. But EWBG still viable with fast walls [Cline & Kainulainen \(2007.10935\)](#)
see also [Dorsch et al \(2106.06547\)](#)



I will discuss relativistic bubble walls so the following calculations would not be valid for subsonic bubble wall



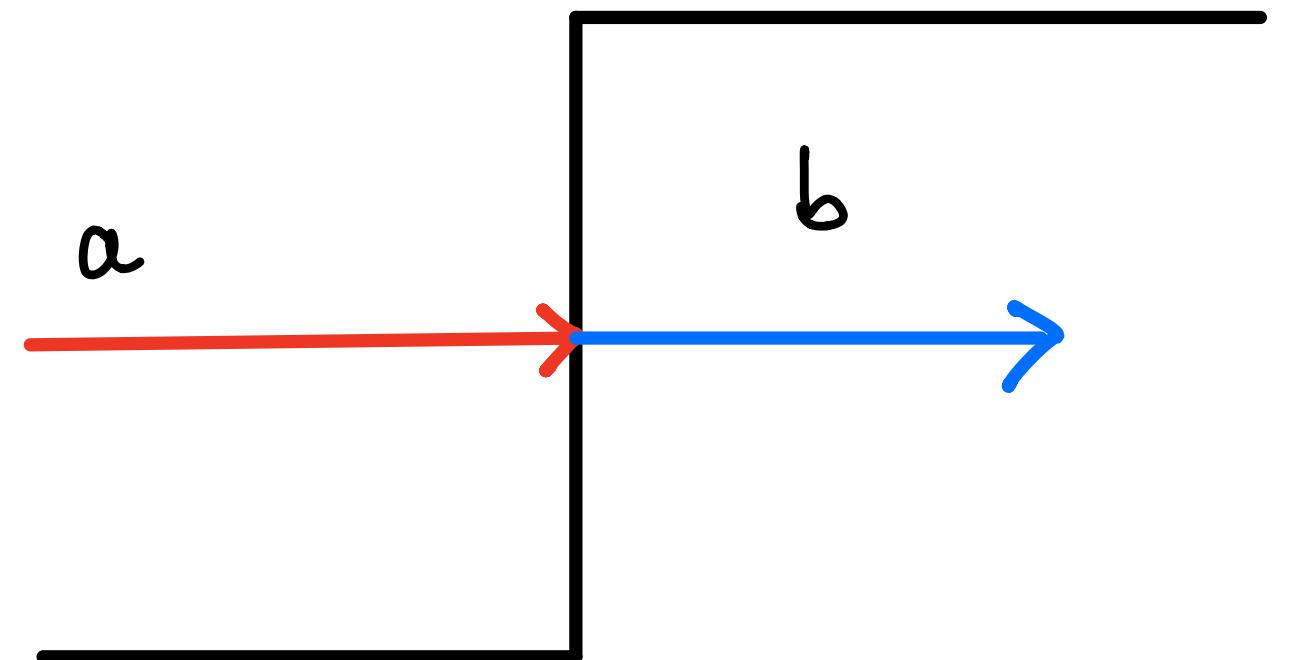
Kinematics and Assumptions

z-translation invariance broken by bubble wall

$$p_{a,z,s} = \sqrt{E_a^2 - |p_{a,\perp}|^2 - m_{a,s}^2}$$
$$p_{a,z,h} = \sqrt{E_a^2 - |p_{a,\perp}|^2 - m_{a,h}^2}$$

$s \equiv$ symmetric phase

$h \equiv$ broken phase



Throughout we assume
infinitely thin walls:
 $L_w \rightarrow 0$

Transverse momentum conserved:

$$p_{a,\perp} = p_{b,\perp}$$

Energy conserved:

$$E_a = E_b$$

1-to-1 pressure calculation for relativistic bubble walls

1-to-1 calculation argument

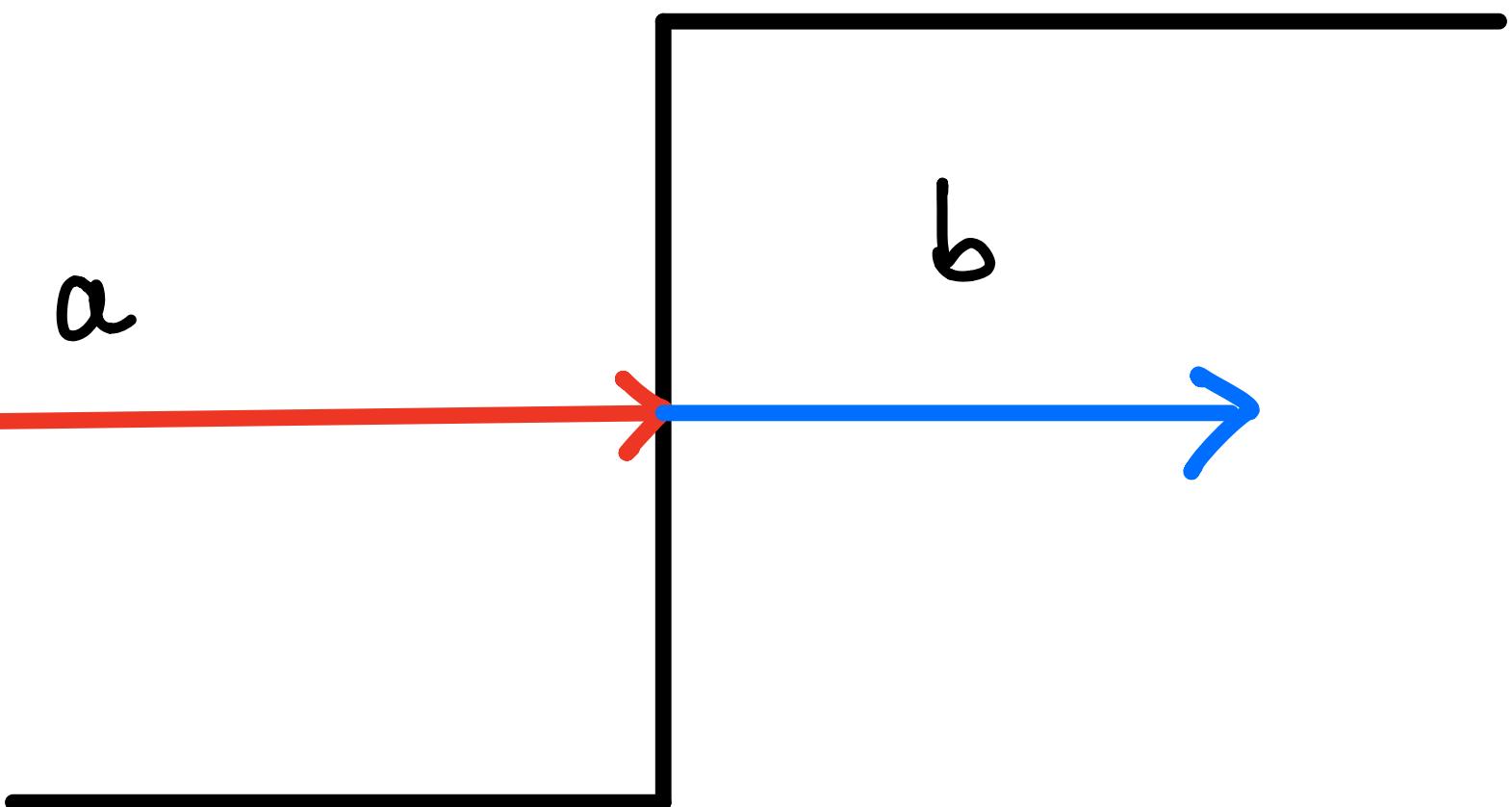
Arnold (1993)

Bodeker & Moore (2009)

Friction → scattering particles that couple to Higgs condensate

$$\gamma = \frac{1}{\sqrt{1 - v^2}}$$

Lorentz factor of the wall



$$\Delta p_{\text{wall}} \equiv p_{a,z,s} - p_{b,z,h} = \sqrt{E_a^2 - m_{a,s}^2 - p_{a,\perp}^2} - \sqrt{E_b^2 - m_{b,h}^2 - p_{b,\perp}^2}$$

$$E_a^2 \sim p_{a,z,s}^2 \sim \gamma^2 T^2 \gg m_{a,s}^2, m_{b,h}^2, p_{a,\perp}^2 \implies \Delta p_{\text{wall}} \approx \frac{m_{b,h}^2 - m_{a,s}^2}{2E_a} = \frac{m_{b,h}^2 - m_{a,s}^2}{2\gamma T}$$

$$\begin{aligned} \mathcal{P}_{1 \rightarrow 1} &= \frac{[\text{force}]}{[\text{area}]} = \frac{\Delta[\text{momentum}]}{[\text{area}] \times [\text{time}]} = [\text{flux}] \times \Delta[\text{momentum}] \\ &\sim \gamma T^3 \times \frac{\Delta m^2}{2\gamma T} = \frac{\gamma^0 T^2 \Delta m^2}{2} \end{aligned}$$

- 1-to-1: no flavour change. Assume no particles reflected

$$\mathcal{P}_{1 \rightarrow 1} = \sum_a \int d\mathcal{F}_a \sum_b \int d\mathbb{P}_{a \rightarrow b} \Delta p_z (1 \pm f_b)$$

Flux incoming particle
 Momentum transfer “observable”
 Differential probability
 Quantum Statistical factors

$$d\mathbb{P}_{a \rightarrow b} = \frac{d^3 p_b}{(2\pi)^3} \frac{1}{2E_b} \times (2\pi)^3 \delta^2(p_{a,\perp} - p_{b,\perp}) \delta(E_a - E_b) (2p_{b,z,h}) \delta_{ab}$$

$$\mathcal{P}_{1 \rightarrow 1} = \sum_a \nu_a \int \frac{d^3 p_a}{(2\pi)^3} \frac{d^3 p_b}{(2\pi)^3} \frac{p_{b,z,s}}{E_b} f_a (1 \pm f_b) (p_{a,z,s} - p_{b,z,h}) (2\pi)^3 \delta^2(p_{a,\perp} - p_{b,\perp}) \delta(E_a - E_b)$$

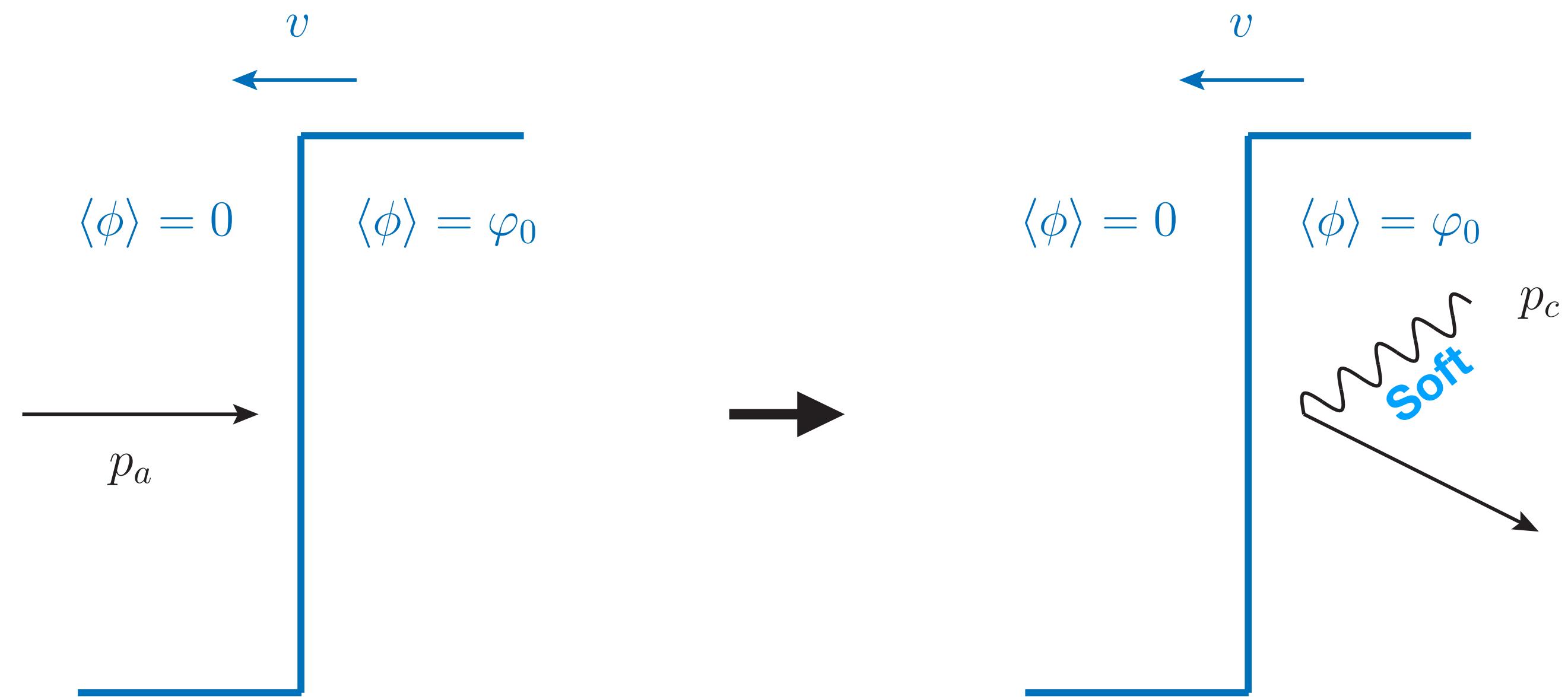
$$\frac{d^3 p_b}{(2\pi)^3 2E_b} = \frac{d^2 p_{b,\perp}}{(2\pi)^3} \frac{dE_b}{2E_b} \frac{E_b}{p_{b,z,s}}$$

$$\mathcal{P}_{1 \rightarrow 1} = \sum_a \nu_a \int \frac{d^3 p_a}{(2\pi)^3} f_a (1 \pm f_b) \Delta p_{\text{wall}} \approx \gamma^0 T^2 \Delta m^2$$

1-to-2 pressure calculation for relativistic bubble walls

1-to-2 calculation

Bodeker & Moore (2017)



Incident particle's energy \rightarrow mass second particle + transverse momentum

$\Delta p_{1 \rightarrow 1} < \Delta p_{1 \rightarrow 2}$ unless $p_\perp = 0$ and $m_c = 0$

1-to-2 calculation

Bodeker & Moore (2017)

$$\mathcal{P}_{1 \rightarrow 2} = \sum \nu_{a,b,c} \int [dp_a] [dp_b] [dp_c] f(p_a, p_b, p_c) \Delta p_z (2\pi)^3 \delta^2 (p_{a,\perp} - p_{c,\perp} - p_{b,\perp}) \delta(E_a - E_c - E_b) |\mathcal{M}|^2$$

Quantum
Statistical factors

Momentum
transfer
“observable”

**Combination of PS +
“observable”**

$$\mathcal{P}_{1 \rightarrow 2} = \sum_{abc} \nu_a \int \frac{d^3 p_a}{(2\pi)^3 2E_a} \int \frac{d^2 p_{c,\perp}}{(2\pi)^2} \frac{dE_c}{(2\pi) 2E_c} f(p_a, p_b, p_c) \Delta p_z \frac{1}{2p_{b,z,s}} \frac{E_c}{p_{c,z,s}} |\mathcal{M}|^2$$

B&M region of interest:

Ingoing hard

$$p_{a,\perp} \sim T \quad p_{a,z,s} \sim E_a \sim \gamma_w T$$

$$m_{a,s}, m_{a,h} \ll \gamma_w T$$

Outgoing hard

$$p_{b,\perp} \sim \max[T, m_c] \quad p_{b,z,s} \sim E_b \sim \gamma_w T$$

$$m_{b,s}, m_{b,h} \ll \gamma_w T$$

Outgoing soft

$$p_{c,\perp} \sim \max[T, m_c] \quad p_{c,z,s} \sim \max[T, m_c]$$

$$m_{c,s}, m_{c,h} \ll \gamma_w T$$

1-to-2 calculation

Bodeker & Moore (2017)

$$\mathcal{P}_{1 \rightarrow 2} = \sum_{abc} \nu_a \int \frac{d^3 p_a}{(2\pi)^3 2E_a} \int \frac{d^2 p_{c,\perp}}{(2\pi)^2} \frac{dE_c}{(2\pi) 2E_c} f(p_a, p_b, p_c) \Delta p_z \frac{1}{2p_{b,z,s}} \frac{E_c}{p_{c,z,s}} |\mathcal{M}|^2$$

B&M region of interest:

$$p_{c,z}(z) = E_c \sqrt{1 - \underbrace{\frac{m_{c,z}^2 + p_{c,\perp}^2}{E_c^2}}_{\epsilon}} \approx E_c \left(1 - \frac{\epsilon}{2} + \mathcal{O}(\epsilon^2) + \dots \right)$$

Collinearity: ϵ

$$p_{b,z} = \sqrt{E_b^2 - p_{\perp,b}^2 - m_b^2(z)} \approx E_a(1-x)$$

Softness: $x = \frac{E_c}{E_a}$

From PS

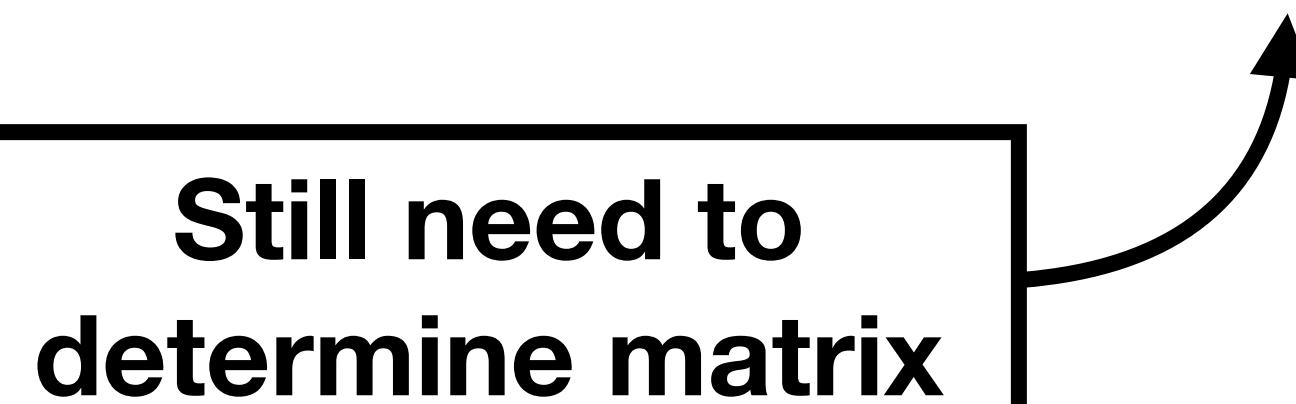
“Observable”

$$(p_{a,z,s} - p_{c,z,h} - p_{b,z,h}) \frac{1}{2p_{b,z,s}} \frac{E_c}{p_{c,z,s}} \approx \underbrace{\frac{1}{2E_a} \frac{m_c^2(z) + p_{c,\perp}^2}{2E_c}}_{\sim x\epsilon + O(\epsilon^2 x) + \dots}$$

B&M 1-to-2 master equation:

$$\mathcal{P}_{1 \rightarrow 2} = \sum \nu_a \int \frac{d^3 p_a}{(2\pi)^3 2E_a} \int \frac{d^2 p_{c,\perp}}{(2\pi)^2} \frac{dE_c}{(2\pi) 2E_c} f(p_a, p_b, p_c) \frac{m_{c,h}^2(z) + p_{c,\perp}^2}{2E_c} |\mathcal{M}|^2$$

**Still need to
determine matrix
element squared**



$$\langle p_c p_b | T | p_a \rangle = \int d^4x \langle p_c p_b | H_{\text{int}} | p_a \rangle = (2\pi)^3 \delta^2(p_{a,\perp} - p_{b,\perp} - p_{c,\perp}) \delta(E_a - E_b - E_c) \mathcal{M}$$

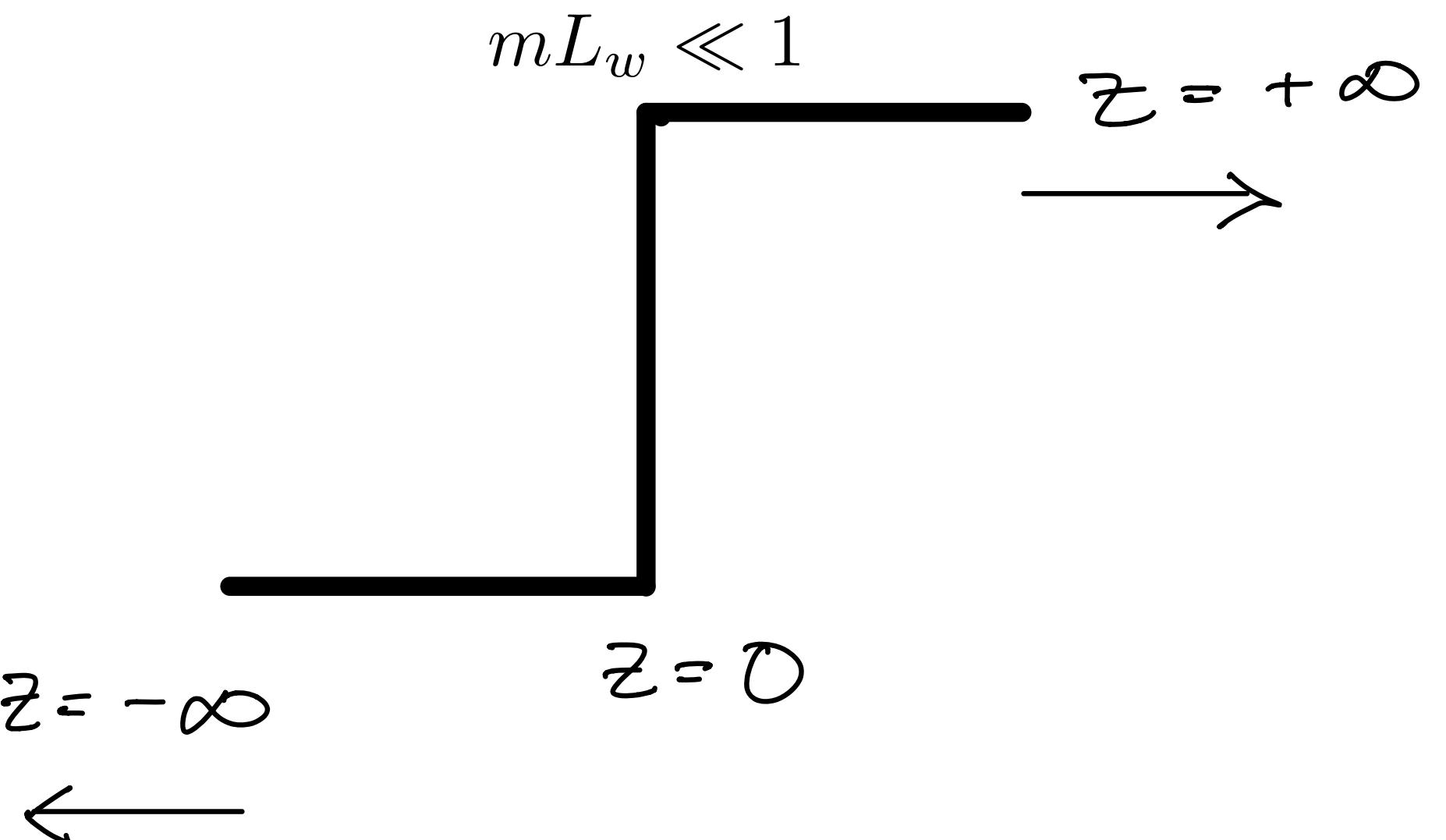
$$\mathcal{M} \equiv \int dz \chi_{p_c}^*(z) \chi_{p_b}^*(z) V(z) \chi_{p_a}(z)$$

Mode functions are treated using the WKB approximation:

$$\chi_{p_c}(z) \simeq \sqrt{\frac{p_{c,z,s}}{p_{c,z}(z)}} \exp\left(i \int_0^z p_{c,z}(z') dz'\right)$$

$$\mathcal{M} = \int dz \chi_b(p_{b,z,s}, z)^* \chi_c(p_{c,z,s}, z)^* \chi_a(p_{a,z,s}, z) V(z)$$

$$\mathcal{M} = V_s \int_{-\infty}^0 dz \exp \left[iz \frac{A_s}{2E_a} \right] + V_h \int_0^\infty dz \exp \left[iz \frac{A_h}{2E_a} \right] = 2iE_a \left(\frac{V_h}{A_h} - \frac{V_s}{A_s} \right)$$



$$A_s = E_a (p_{a,z,s} - p_{b,z,s} - p_{c,z,s})$$

$$A_h = E_a (p_{a,z,h} - p_{b,z,h} - p_{c,z,h})$$

$1/A$ terms resemble propagators, but they only propagate in the z -direction!

Vertex Function

Bodeker & Moore (2017)

$$|\mathcal{M}|^2 \simeq 4E_a^2|V|^2 \frac{(A_h - A_s)^2}{A_h^2 A_s^2}$$

$a(p) \rightarrow b(k)c(p-k)$	$ V ^2$
$S \rightarrow V_T S$	
$F \rightarrow V_T F$	$4g^2 C_2[R] \frac{1}{x^2} k_\perp^2$
$V \rightarrow V_T V$	
$S \rightarrow V_L S$	
$F \rightarrow V_L F$	$4g^2 C_2[R] \frac{1}{x^2} m^2$
$V \rightarrow V_L V$	
$F \rightarrow F V_T$	$2g^2 C_2[R] \frac{1}{x} (k_\perp^2 + m_b^2)$
$V \rightarrow F F$	$2g^2 T[R] \frac{1}{x} (k_\perp^2 + m_b^2)$
$S \rightarrow S V_T$	$4g^2 C_2[R] k_\perp^2$
$F \rightarrow S F$	$y^2 (k_\perp^2 + 4m_a^2)$
$S \rightarrow S S$	$\lambda^2 \varphi^2$

$$k_\perp \equiv p_{c,\perp}$$

These are splitting functions up to the normalisation $P_{b \leftarrow a}(x) = |V|^2 x(1-x)/16\pi^2 k_\perp^2$

$$A_s \approx 2E_a \times \left(-\frac{|p_{a,\perp}|^2 + m_{a,s}^2}{2E_a} + \frac{|p_{b,\perp}|^2 + m_{b,s}^2}{2E_b} + \frac{|p_{c,\perp}|^2 + m_{c,s}^2}{2E_c} \right) \approx \frac{|p_{c,\perp}|^2 + m_{c,s}^2}{E_c/E_a}$$

$$A_h \approx 2E_a \times \left(-\frac{|p_{a,\perp}|^2 + m_{a,h}^2}{2E_a} + \frac{|p_{b,\perp}|^2 + m_{b,h}^2}{2E_b} + \frac{|p_{c,\perp}|^2 + m_{c,h}^2}{2E_c} \right) \approx \frac{|p_{c,\perp}|^2 + m_{c,h}^2}{E_c/E_a}$$

$$|\mathcal{M}|^2 = 4E_a^2 |V_s|^2 \frac{(A_s - A_h)^2}{A_s^2 A_h^2}$$

$$\approx 4E_a \underbrace{\frac{|V_s|^2}{(E_c/E_a)^{-2}}}_{x} \frac{\left(m_{c,h}^2 - m_{c,s}^2\right)^2}{\left(|p_{c,\perp}|^2 + m_{c,s}^2\right)^2 \left(|p_{c,\perp}|^2 + m_{c,h}^2\right)^2}$$

Lets keep track of what cancels where...

$$\begin{aligned}
 |\mathcal{M}|^2 &= 4E_a^2 |V_s|^2 x^2 \frac{m_{c,h}^4}{|p_{c,\perp}|^4 \left(|p_{c,\perp}|^2 + m_{c,h}^2 \right)^2} & m_{c,s} \ll m_{c,h} \\
 &= 16g^2 C_2[R] \frac{|p_{c,\perp}|^2}{\cancel{x^2}} x^2 \frac{m_{c,h}^4}{|p_{c,\perp}|^4 \left(|p_{c,\perp}|^2 + m_{c,h}^2 \right)^2} \\
 &= 16g^2 C_2[R] E_a^2 \frac{m_{c,h}^4}{|p_{c,\perp}|^2 \left(|p_{c,\perp}|^2 + m_{c,h}^2 \right)^2}
 \end{aligned}$$

Combine matrix element squared with momentum transfer observable:

$$\begin{aligned}
 |\mathcal{M}|^2 \times \Delta p_z &\approx 16E_a^2 g^2 C_2[R] \frac{m_{c,h}^4}{|p_{c,\perp}|^2 \left(|p_{c,\perp}|^2 + m_{c,h}^2 \right)^2} \times \frac{m_{c,h}^2 + p_{c,\perp}^2}{2E_c} \\
 &\approx \frac{8g^2 E_a}{x} C_2[R] \frac{m_{c,h}^4}{|p_{c,\perp}|^2 \left(|p_{c,\perp}|^2 + m_{c,h}^2 \right)}
 \end{aligned}$$

Two pieces left: the integration of PS of incoming “a” gives the flux. We also need to integrate over phase space of particle “c” our soft emission

$$\int_{g^2 T^2}^{m^2} \frac{dp_{c,\perp}^2}{(2\pi)^2 |p_{c,\perp}|^2 \left(|p_{c,\perp}|^2 + m_{c,h}^2\right)} \sim \frac{1}{24\pi m_{c,h}^2}$$

$$\int \frac{dE_c}{E_c^2} \sim \frac{1}{m_{c,h}}$$

Gamma dependence
enters from flux

$$\mathcal{P}_{1 \rightarrow 2} \sim \gamma T^3 \frac{1}{\underline{m^2}} \frac{1}{\underline{m}} \underline{m^4}$$

Transverse momentum
integration

Energy
integration

B&M assume $g^2 T^2 \ll m^2$

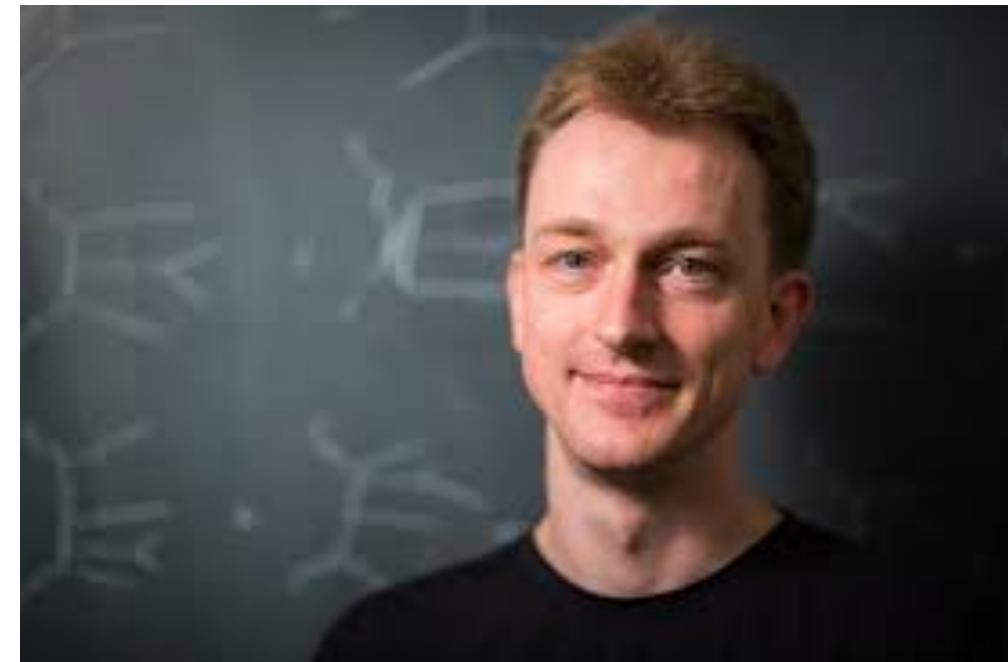
The thermal mass would cut off the integral and you'd get some (possibly large) $\log(m_{c,h}^2/m_{c,s}^2)$ see [Azatov et al \(2010.02590\)](#)

$$\mathcal{P}_{1 \rightarrow 2} \sim m\gamma T^3$$

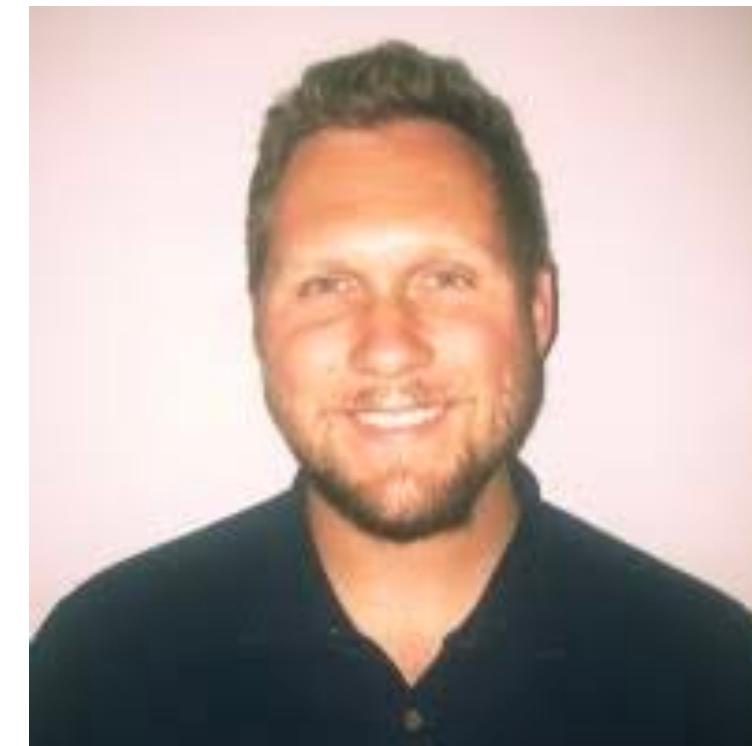
- $\mathcal{P}_{1 \rightarrow 2} \propto \gamma$ while $\mathcal{P}_{1 \rightarrow 1} \propto \gamma^0$. Since, vacuum pressure does not grow in λ , a terminal velocity will be reached.
- $\mathcal{P} \propto m^2 \implies$ no phase change pressure goes to zero. This comes from integrating over the mode functions
- \mathcal{M} : WKB and vertex. The vertex part is dominated in the soft regime.
- B&M cut off the k_\perp and k^0 integration by gauge boson mass.
- This interaction looks like a collider/scattering experiment where the collision occurs between the ingoing particle and the wall \implies the centre of mass energy will be large \implies many soft emission.

1-to-n pressure calculation for relativistic bubble walls

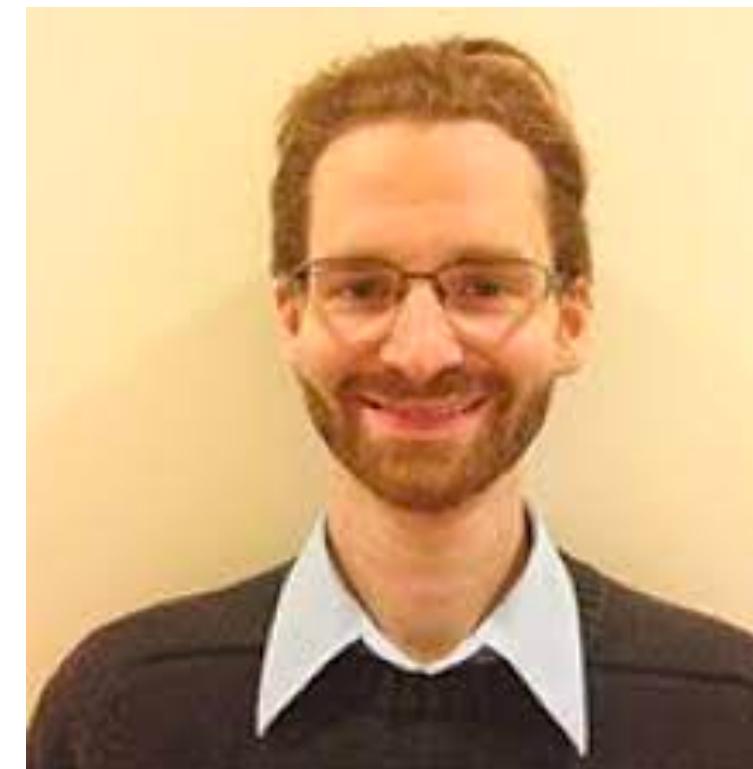
2007.10343 (JCAP 2103 (2021) 009)



Stefan Höche



Jonathan Kozaczuk

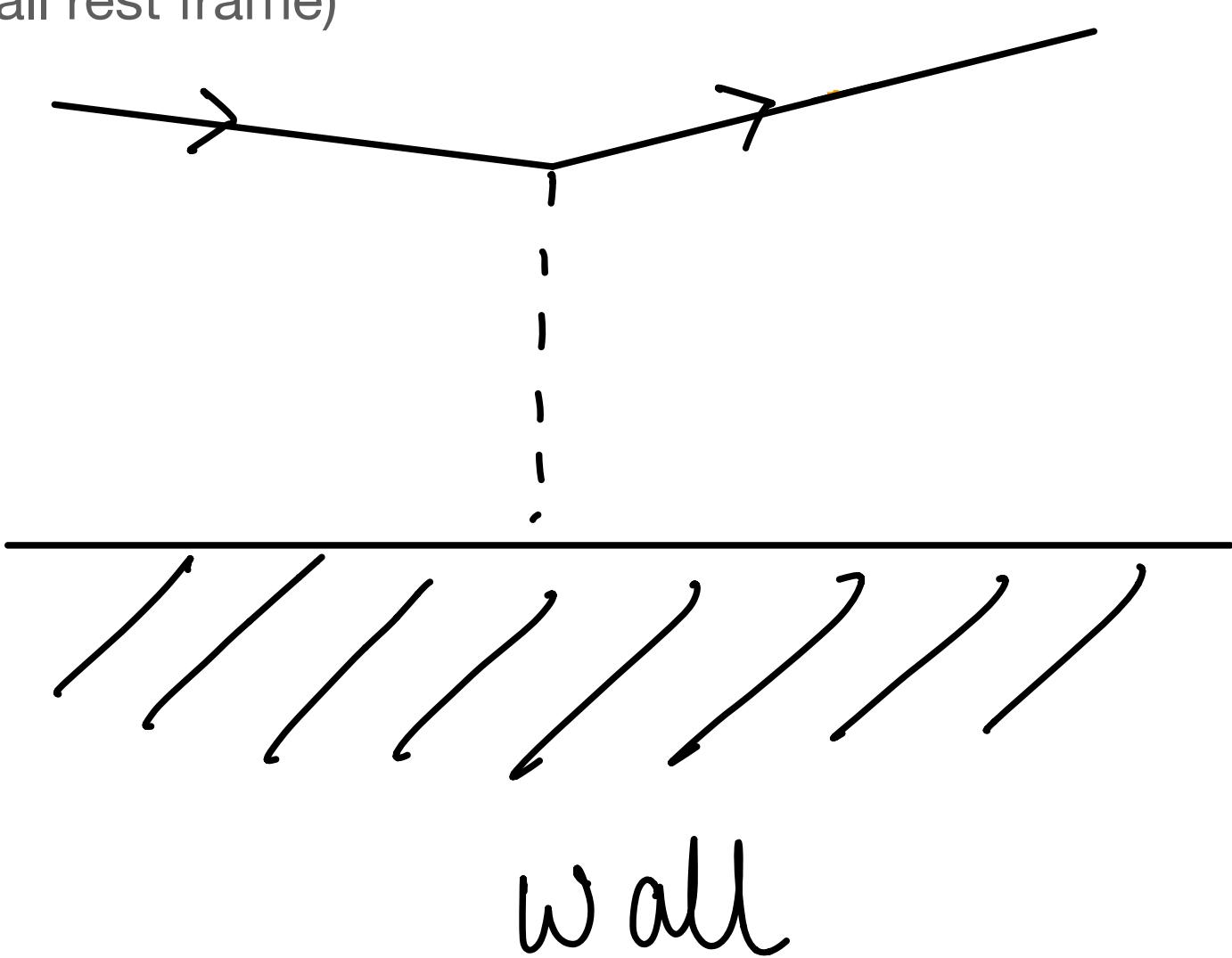


Andrew Long

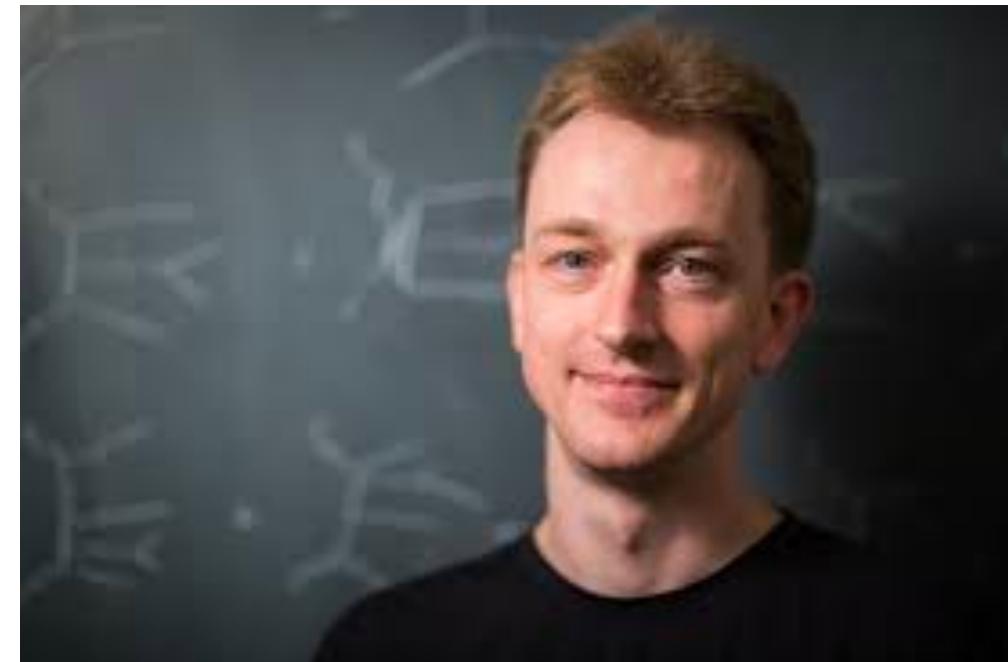


Yikun Wang

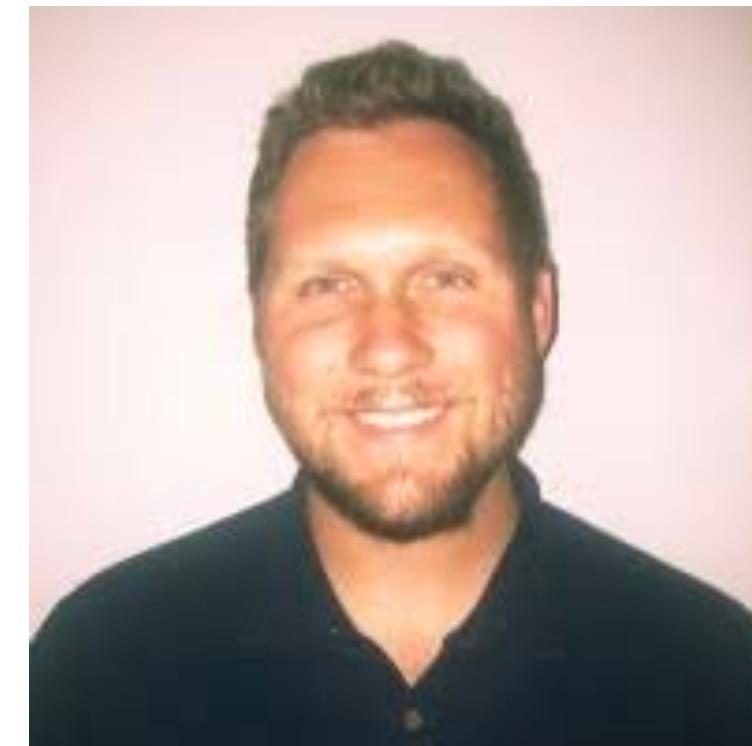
Incoming fast on-shell
particle
(in wall rest frame)



B&M investigated 1-to-1
which we will treat as leading order
process.



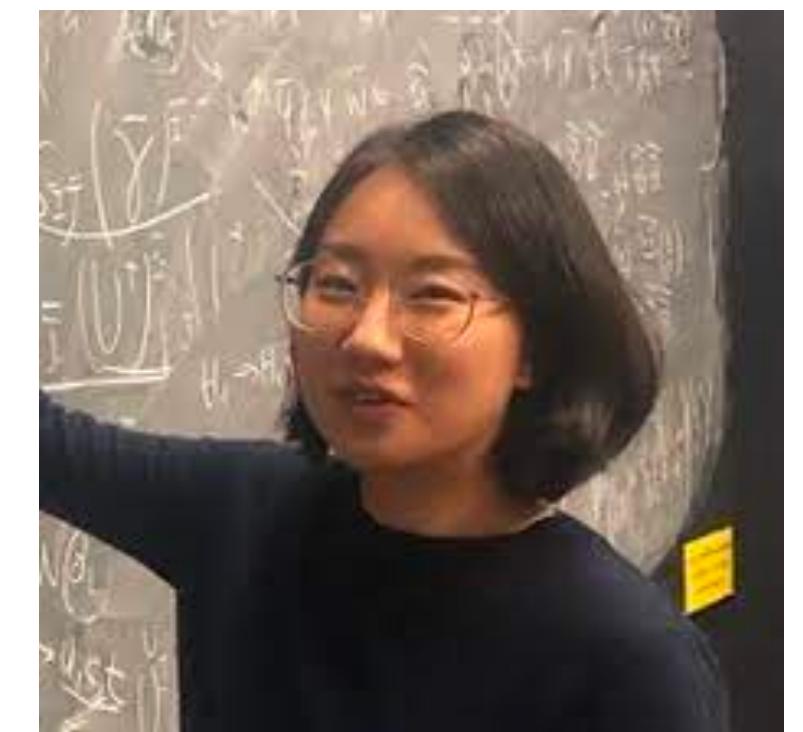
Stefan Höche



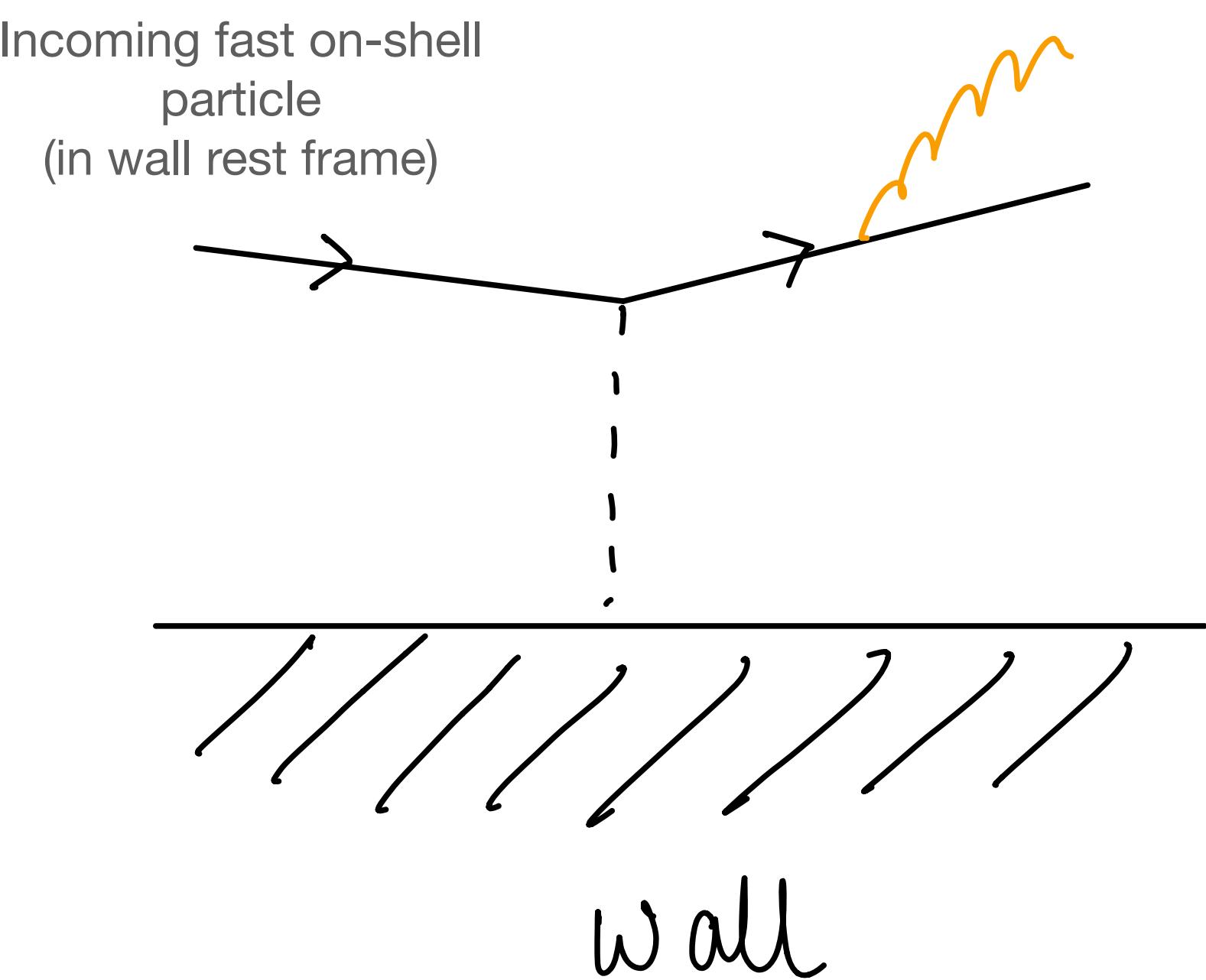
Jonathan Kozaczuk



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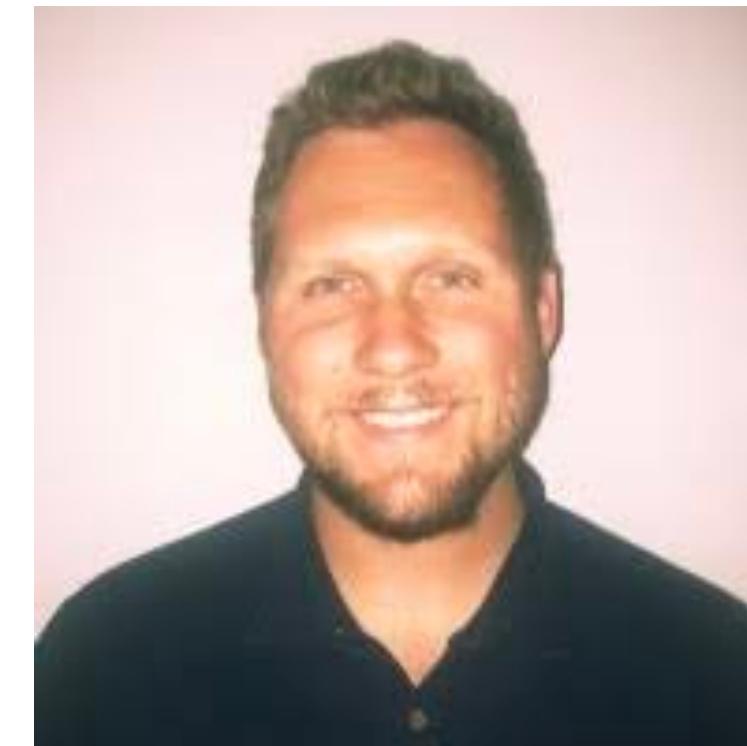
Yikun Wang



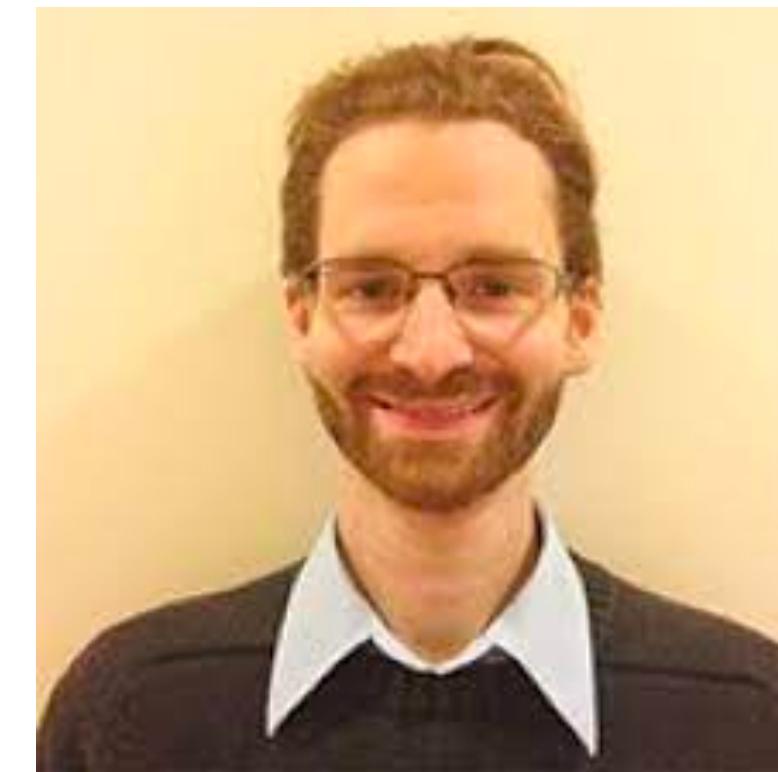
Treat 1-to-2 as quantum correction
to 1-to-1 process.



Stefan Höche



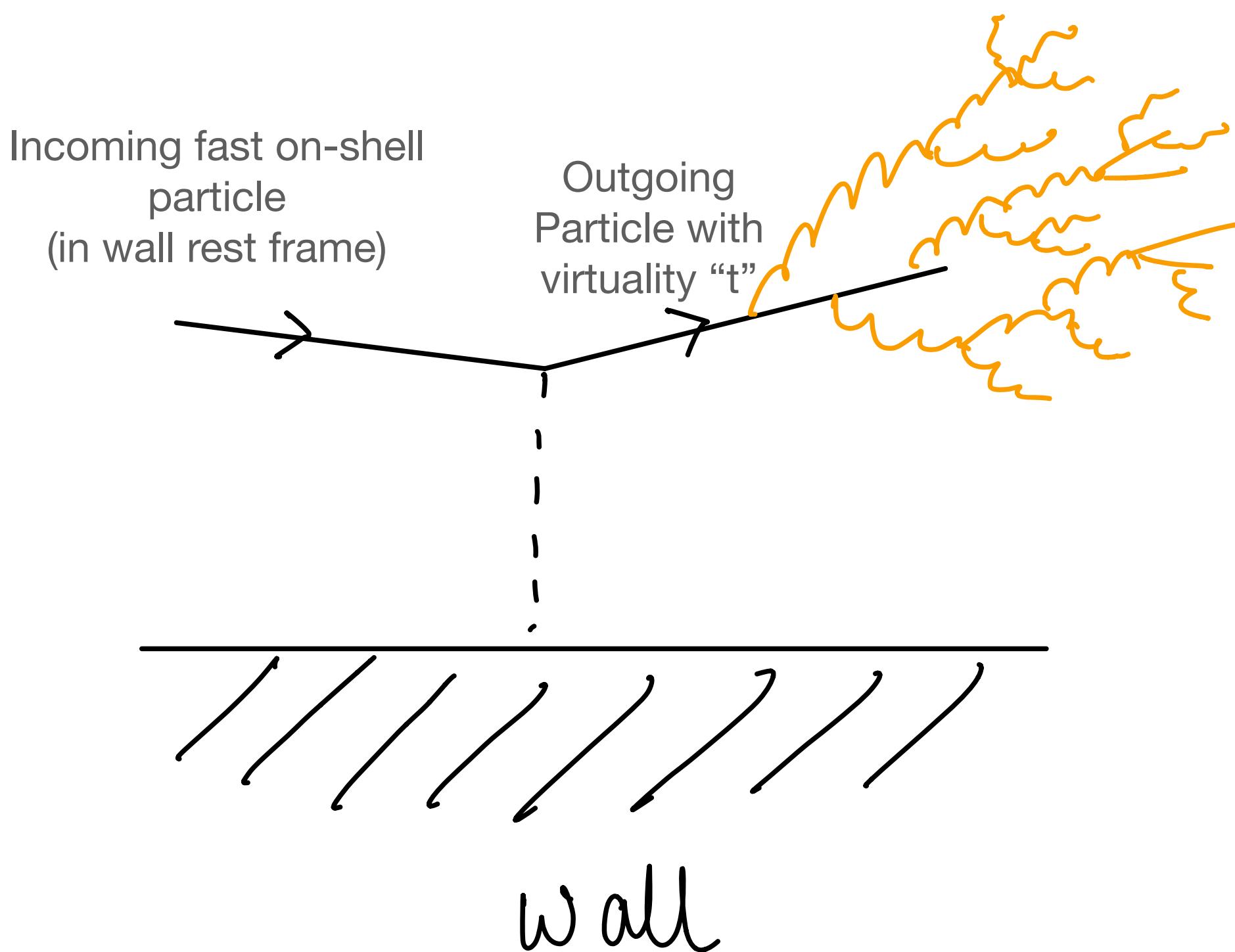
Jonathan Kozaczuk



Andrew Long



Yikun Wang



If 1-to-2 is important, will 1-to-n be even more important?

Reformulation of matrix element

B&M

$$\begin{aligned}
 A_s &= -2E_a (p_{a,z,s} - p_{b,z,s} - p_{c,z,s}) \\
 &\approx p_{a,\perp}^2 + m_{a,s}^2 - \frac{p_{b,\perp}^2 + m_{b,s}^2}{E_b/E_a} - \frac{k_\perp^2 + m_{c,s}^2}{E_c/E_a} \\
 &= -\frac{k_\perp^2 + m_{b,s}^2}{(1-x)} - \frac{k_\perp^2 + m_{c,s}^2}{x} \\
 &\approx -\frac{k_\perp^2 + m_{c,s}^2}{x} \quad (x \ll 1)
 \end{aligned}$$

Ultrarelativistic incoming particle
Thermal mass incoming particle negligible. Frame where incoming particle is "headon"
Soft limit

$$\begin{aligned}
 A_h &= -2E_a (p_{a,z,h} - p_{b,z,h} - p_{c,z,h}) \\
 &\approx p_{a,\perp}^2 + m_{a,h}^2 - \frac{p_{b,\perp}^2 + m_{b,h}^2}{E_b/E_a} - \frac{k_\perp^2 + m_{c,h}^2}{E_c/E_a} \\
 &= -\frac{k_\perp^2 + m_{b,h}^2}{(1-x)} - \frac{k_\perp^2 + m_{c,h}^2}{x} \\
 &\approx -\frac{k_\perp^2 + m_{c,h}^2}{x} \quad (x \ll 1)
 \end{aligned}$$

JT & co

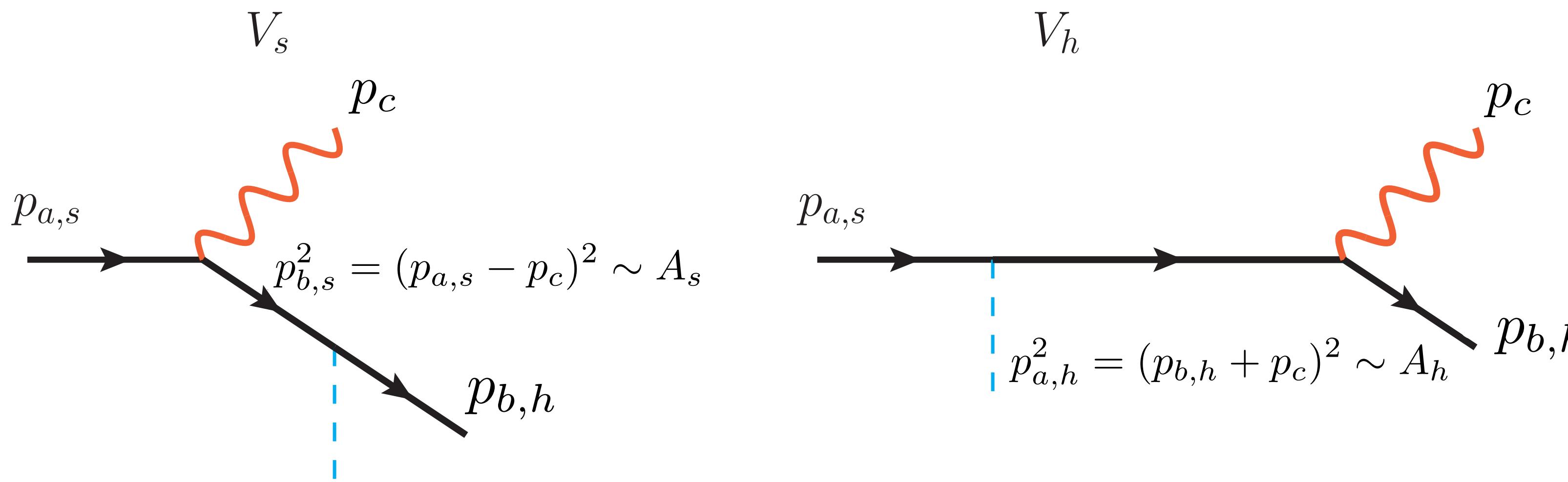
$$\begin{aligned}
 A_s &= -2p_{a,s} \cdot p_{c,s} \\
 &\approx -\frac{k_\perp^2 + m_{c,s}^2}{x(1-x)} - xm_{a,s}^2 \\
 &\approx -\frac{k_\perp^2 + m_{c,s}^2}{x} \quad x \ll 1
 \end{aligned}$$

$$\begin{aligned}
 A_h &= -2p_{b,h} \cdot p_{c,h} \\
 &\approx -\frac{k_\perp^2}{x(1-x)} - \frac{1-x}{x} m_{c,h}^2 - \frac{x}{1-x} m_{b,h}^2 \\
 &\approx -\frac{k_\perp^2}{x(1-x)} - \frac{1-x}{x} m_{c,h}^2 - \frac{x}{1-x} m_{b,h}^2 \\
 &\approx -\frac{k_\perp^2 + m_{c,h}^2}{x} \quad x \ll 1
 \end{aligned}$$

Two expressions same in the soft-limit but not same outside of soft limit.

Why the redefinition of A's?

Treat A's as covariant propagators. Wall provides spatial discontinuity off of which incoming particle scatters.



$$A_s = -2p_{a,s} \cdot p_{c,s} \approx 2p_{a,\perp} \cdot k_\perp - \frac{p_{a,\perp}^2 + m_{a,s}^2}{E_c/E_c} - \frac{k_\perp^2 + m_{c,s}^2}{E_c/E_c} \xrightarrow{x \ll 1} -\frac{k_\perp^2 + m_{c,s}^2}{x}$$
$$A_h = -2p_{b,h} \cdot p_{c,h} \approx 2p_{b,\perp} \cdot k_\perp - \frac{p_{b,\perp}^2 + m_{b,h}^2}{E_b/E_c} - \frac{k_\perp^2 + m_{c,h}^2}{E_c/E_b} \xrightarrow{x \ll 1} -\frac{k_\perp^2 + m_{c,h}^2}{x}$$

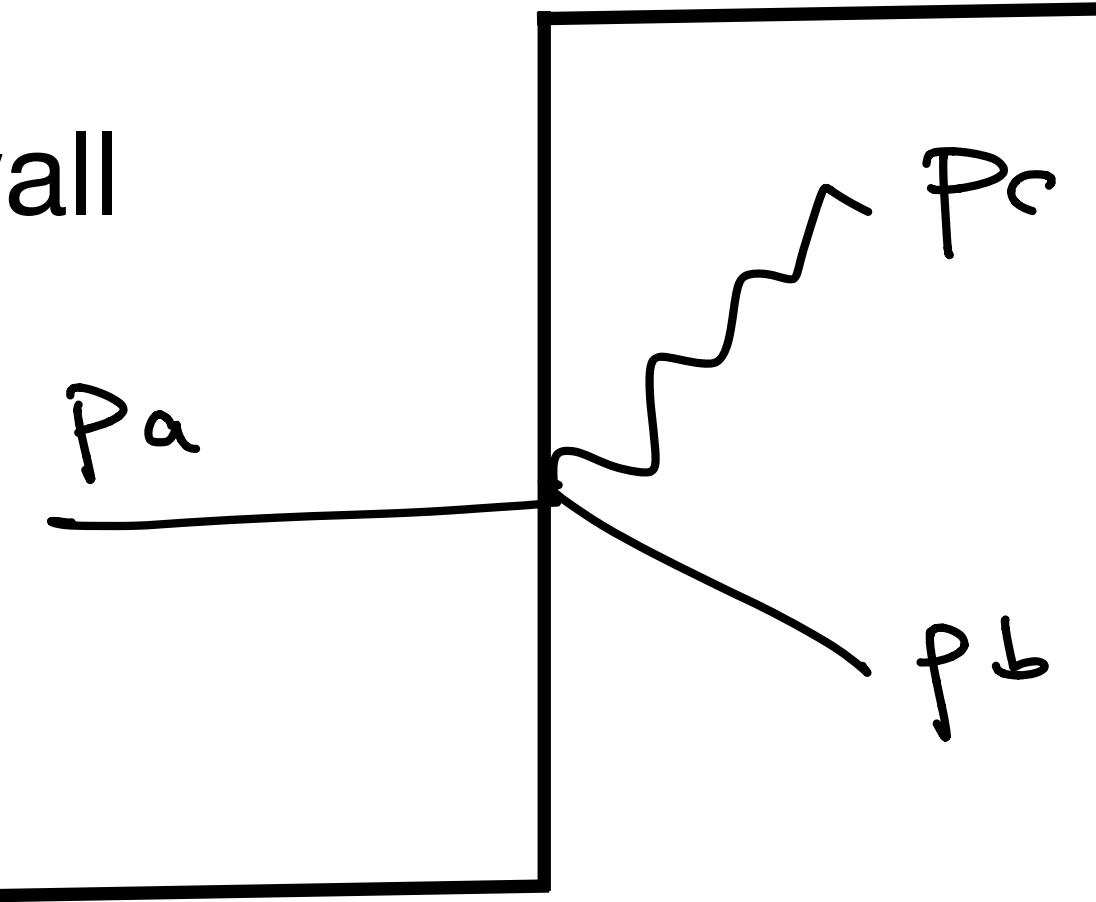
Reformulation of the Matrix Element

Our matrix element has the following form:

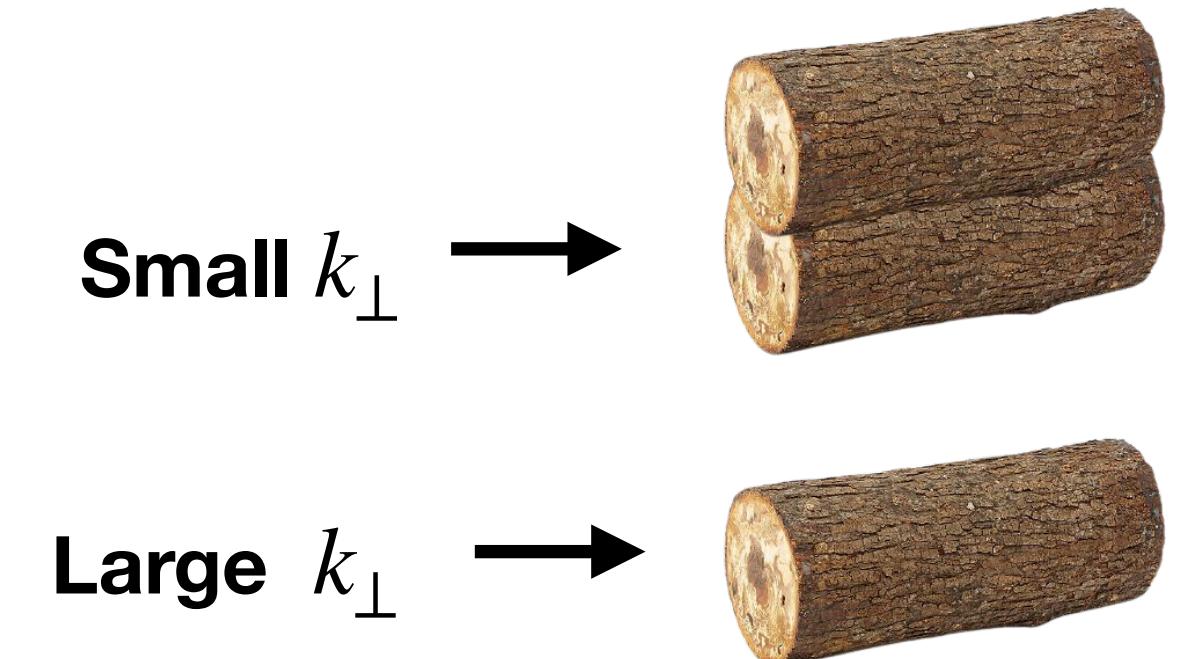
$$\left| M_{a \rightarrow bc}^{(0)} \right|^2 = 4E_a^2 |g|^2 \left(\frac{2p_{a,s}p_{b,h}}{p_{a,s}p_c p_{b,h} p_c} - \frac{m_{a,s}^2}{(p_{a,s}p_c)^2} - \frac{m_{b,h}^2}{(p_{b,h}p_c)^2} \right)$$

This is a gauge invariant squared matrix element that corresponds to decelerated charges emitting soft gauge field quanta.

Here, consider massless radiation with radiator acquiring mass across wall



$$\frac{1}{\left| \mathcal{M}_{a \rightarrow b}^{(0)} \right|^2} \int \frac{d^3 p_c}{(2\pi)^3 2E_c} \left| M_{a \rightarrow b c}^{(0)} \right|^2 \approx \begin{cases} \frac{\alpha}{2\pi} \int \frac{dk_\perp^2}{k_\perp^2} C_{abc} \log \frac{m_{b,h}^2}{k_\perp^2} & \text{if } m_{a,s}^2 \ll k_\perp^2 \ll m_{b,h}^2 \\ \frac{\alpha}{2\pi} \int \frac{dk_\perp^2}{k_\perp^2} C_{abc} & \text{if } m_{a,s}^2, m_{b,h}^2 \ll k_\perp^2 \end{cases}$$



Analytic Resummation

Banfi, Salam & Zanderighi (2005)

$R(v)$ probability for decay $a \rightarrow bc$. For this splitting to produce momentum transfer of v we require that it did not produce a large momentum transfer before.

$$R(v) = \int [dk] |M^2(k)| \Theta[V(\{p\}, k) - v] \quad V(p_a, p_b, p_c) = \frac{\Delta p_z}{\gamma T} \approx \frac{k_\perp^2 / (2E_a^2)}{x(1-x)}$$

$$R_{abc}(V) = C_{abc} |g|^2 \int \frac{d^3 p_c}{(2\pi)^3 2E_c} \left(\frac{2p_{b,h} p_{a,s}}{p_{a,s} p_c p_{b,h} p_c} + O\left(\frac{m_{a,s}^2}{k_\perp^2}, \frac{m_{b,h}^2}{k_\perp^2}\right) \right) \Theta(V(p_a, p_b, p_c) - V) \Theta(p_{b,z,h}) \Theta(p_{c,z,h})$$

Rewrite phase space and matrix element squared in terms of observable V

$$R_{abc}(V) = C_{abc} \frac{\alpha}{2\pi} \int_V^1 \frac{dV'}{V'} \int_0^1 dx 2x \Theta\left(\frac{1}{1+V'} - x\right) \Theta\left(x - \frac{V'}{1+V'}\right)$$

$$R_{abc}(V) = \frac{\alpha}{2\pi} C_{abc} (L + 2 \log(1 + e^{-L})) \text{ where } L = \log \frac{1}{V}$$

Sudakov
form factor

Banfi, Salam & Zanderighi (2005)

$$\Delta_a(V) = \exp \left\{ - \sum_b R_{ab}(V) \right\}, \quad \text{where} \quad R_{ab}(V) = \sum_c R_{abc}(V)$$

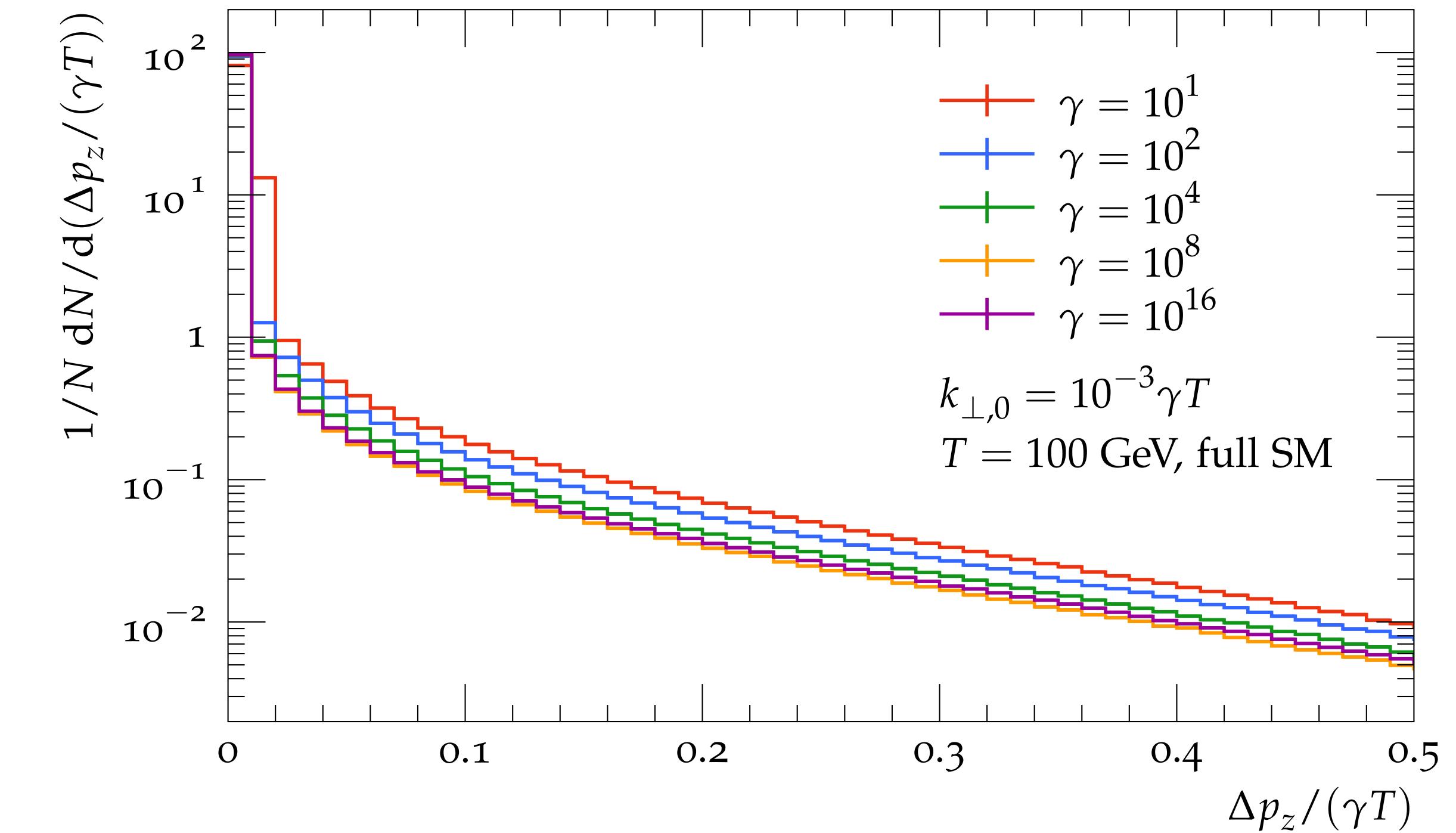
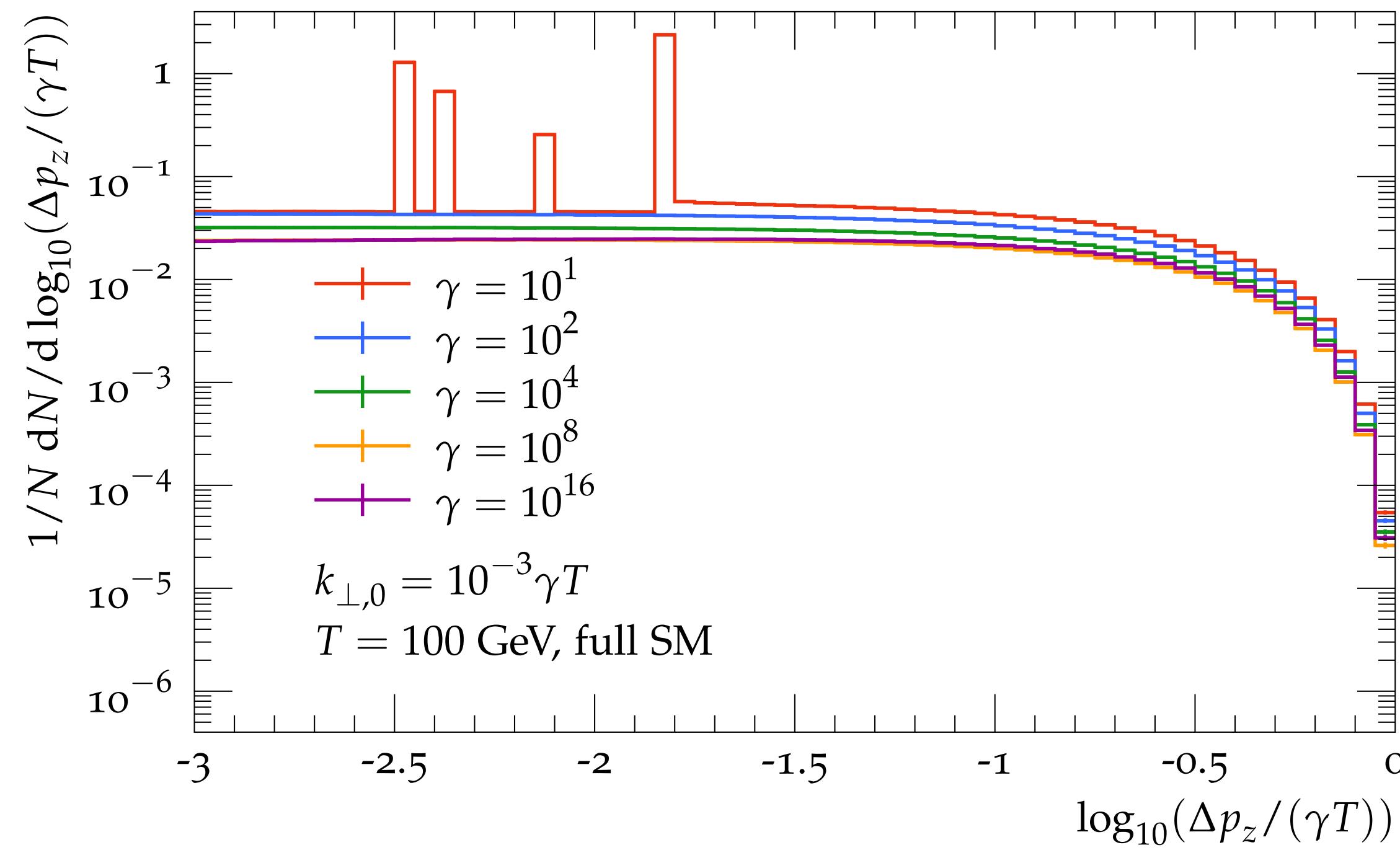
$$\left\langle \frac{\Delta p_z}{\gamma T} \right\rangle = \int_0^1 dV V \frac{d}{dV} \prod_{a \in \mathcal{S}} \Delta_a(V)$$

Average momentum
transfer from all
branchings

$$\left\langle \frac{\Delta p_z}{\gamma T} \right\rangle_{\text{FC}} = \int_0^\infty dL e^{-L} \frac{(\alpha C)_\Sigma}{2\pi} \frac{e^L - 1}{e^L + 1} \exp \left\{ - \frac{(\alpha C)_\Sigma}{2\pi} (L + 2 \log(1 + e^{-L})) \right\} \approx \zeta (\log 4 - 1) \quad \zeta = \frac{(\alpha C)_\Sigma}{2\pi}$$

$$\boxed{\langle \Delta p_z \rangle \sim \gamma T}$$

Numerical Resummation



$$\left\langle \frac{\Delta p_z}{\gamma T} \right\rangle = 0.89(17)\% - 0.14(3)\% \log_{10} \gamma \implies P \propto \gamma^2 T^4$$

Comparing analytic and numerics

Particle	n_f	ν_a	analytic ($\langle \Delta p_z / (\gamma T) \rangle_{\text{FC}}$)	numeric ($\langle \Delta p_z / (\gamma T) \rangle_{\text{FC}}$)
l^\pm	2×3	2	0.17%	0.17%
u	2×3	2×3	0.46%	0.44%
d	2×3	2×3	0.46%	0.44%
W^\pm	2	2	0.52%	0.52%
Z	1	2	0.41%	0.40%
h	1	1	0.22%	0.21%
G_{W^\pm}	2	1	0.22%	0.21%
G_Z	1	1	0.22%	0.21%

Good agreement between analytics and numerics. Largest pressure contribution
From QCD processes

Summary

- 1st order EWPT has many interesting physical consequences such as baryogenesis & GW production. Both quantitatively depend on the velocity of the bubble wall.
- Bubble wall velocity is a force balancing exercise: pressure from Higgs potential versus frictional pressure from plasma.
- We reformulated the calculation of the latter in a GI way and calculated the average pressure to all orders. This has a physical correspondence with a charged particle decelerating and emitting a soft radiation pattern.
- We find pressure $\propto \gamma^2$ also massless GB contribute the largest pressure of all SM. NB phase change is still required by fermion across wall. Numerical and analytic resumption agree to 10% level.
- Still **many** interesting avenues to explore.

A scenic view of Durham Cathedral, a large Gothic cathedral built on a rocky outcrop overlooking the River Wear. The cathedral's tall, light-colored stone towers and intricate stonework are prominent against a clear blue sky with some wispy clouds. In the foreground, the calm water of the river reflects the surrounding trees, which are displaying autumn colors. A row of traditional stone houses with red roofs sits along the riverbank at the base of the cathedral.

Thank you for your time

Mode functions

Mode function quick summary

Bodeker & Moore (2017)

scalar interacting with wall

$$\mathcal{L} = \sum_{f=a,b,c} \left[\frac{1}{2} (\partial_\mu \phi_f)^2 - \frac{1}{2} m_f^2(z) \phi_f^2 \right]$$

KG field equation

$$\square \phi_f + m_f^2(z) \phi_f = 0$$

Z-varying mass
parametrises spatial
inhomogeneity

solve with an **homogeneous** mass parameter,
solutions can be labeled by a 3-vector \vec{p}

$$\chi_f(\vec{p}, x) = e^{-iE_f(\vec{p})t} e^{i\vec{p}\cdot\vec{x}} \quad \text{with} \quad E_f(\vec{p}) \equiv \sqrt{|\vec{p}|^2 + m_f^2}$$

$$\phi_f(x) = \int \frac{d^3 p}{(2\pi)^3} \tilde{\phi}_f(\vec{p}) \chi_f(\vec{p}, x)$$

Mode function quick summary

Bodeker & Moore (2017)

But we have **inhomogeneous** mass term, make Ansatz for solution to KG equation

$$\phi_f(x) = \int \frac{d^2\vec{p}_\perp}{(2\pi)^2} \frac{dp_{z,s}}{(2\pi)} \tilde{\phi}_f(\vec{p}_\perp, p_{z,s}) \chi_f(p_{z,s}, z) e^{-iE_f(\vec{p}_\perp, p_z, s)t} e^{i\vec{p}_\perp \cdot \vec{x}_\perp}$$

Substitute into Klein Gordon \rightarrow WKB solution for a particle with
inhomogeneous mass

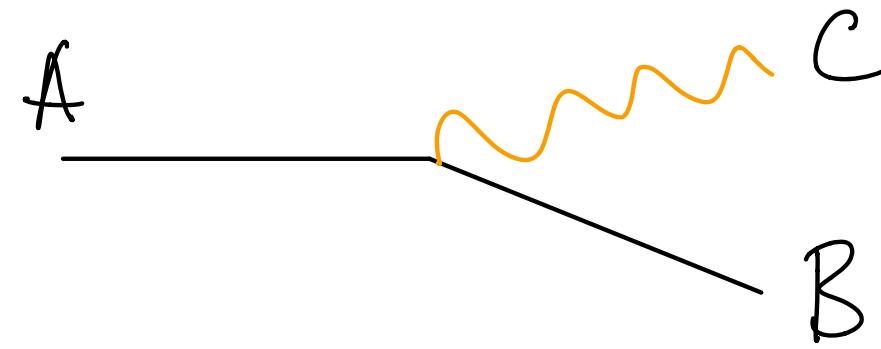
$$\chi_p(z) = \sqrt{\frac{p_{z,s}}{p_z(z)}} \exp\left(i \int_0^z p_z(z') dz'\right)$$

Analogous 1D-scattering off a potential well. Normally there would be a wave function with a negative phase (reflected) but here all particles transmitted

Splitting function

Quick Recap on splitting functions

Gribov & Lipatov (1972)
 Dokshitzer (1977)
 Altarelli & Parisi (1977)



P_{CA} is probability that A emits (collinear) C which carries $x = E_C/E_A$ energy fraction of parent particle

**Sudakov parametrisation
used in Altarelli & Parisi's paper**

$$P_{BA}(z) = \frac{1}{2} z(1-z) \sum_{\text{spins}} \frac{|V_{A \rightarrow B+C}|^2}{p_\perp^2}$$

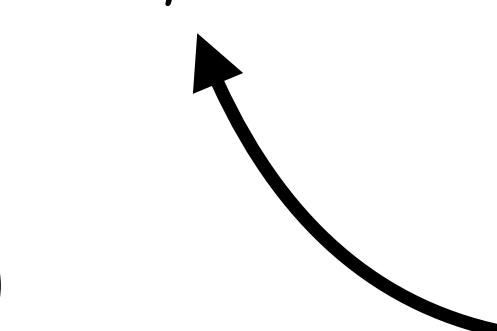
$$\sum_{\text{spin}} |V_{A \rightarrow B+C}|^2 = \frac{1}{2} C_2(R) \text{Tr} (k_B^\mu \gamma_\mu k_A^\nu \gamma_\nu) \sum_{\text{pol}} \epsilon^{*\mu} \epsilon_\mu$$

$$\sum_{\text{spin}} |V_{A \rightarrow B+C}|^2 = \frac{2p_\perp^2}{x(1-x)} \frac{1+(1-x)^2}{x} C_2(R)$$

$$k_A = (P, P, 0)$$

$$k_B = \left((1-x)P + \frac{p_\perp^2}{2(1-x)P}, (1-x)P, -p_\perp \right)$$

$$k_C = \left(xP + \frac{p_\perp^2}{2xP}, xP, p_\perp \right)$$



Calculated in light-like Axial gauge
in Altarelli Parisi (AP) paper

Matrix element reformulation

Reformulation of the Matrix Element

$$V_s = (-ig)\bar{u}_s(p_{b,s})\gamma^\mu \epsilon_\mu^*(p_{c,s}) u_s(p_{a,s})$$

$$V_h = (-ig)\bar{u}_h(p_{b,h})\gamma^\mu \epsilon_\mu^*(p_{c,h}) u_h(p_{a,h})$$

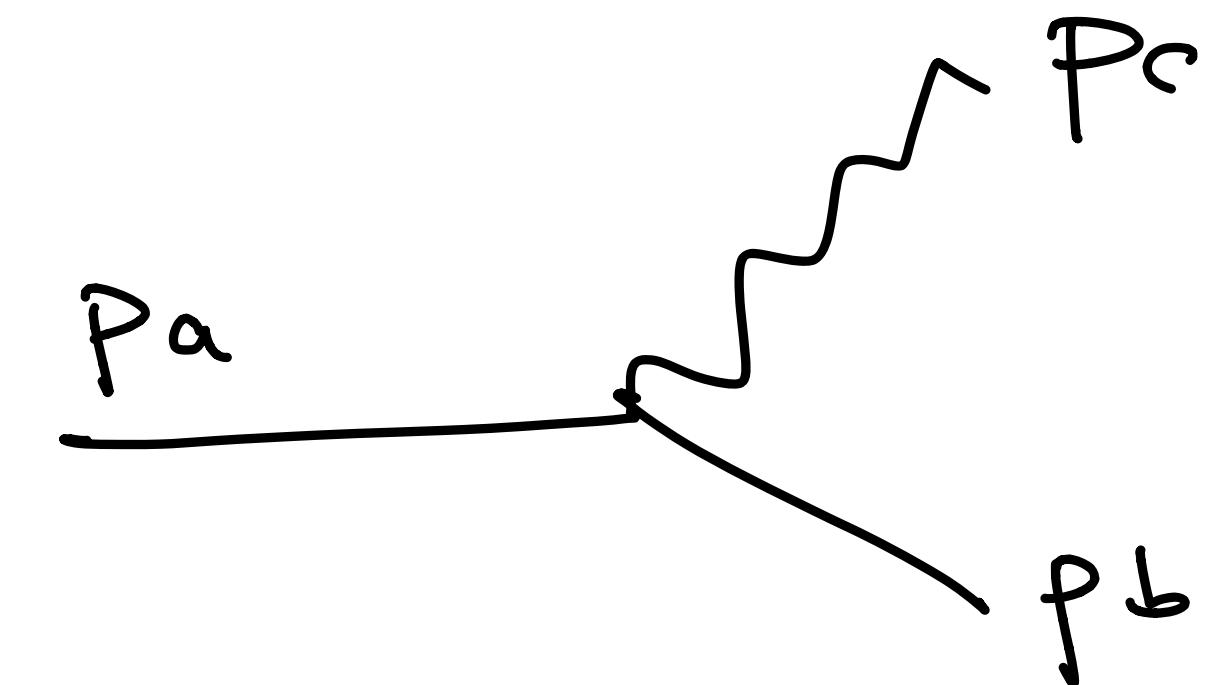
$$\sum_\lambda \epsilon \epsilon^* = -g^{\mu\nu} + \zeta \left(-\frac{n^2 p_c^\mu p_c^\nu}{(n \cdot p_c)^2} + \frac{n^\nu p_c^\mu + n^\mu p_c^\nu}{n \cdot p_c} \right)$$

$$V_s^* V_s = 2g^2 \left(\left(1 + \frac{2\zeta}{x^2}\right) k_\perp^2 - 2(1-\zeta)m_{b,s}^2 + 2\zeta m_c^2 \frac{1-x}{x^2} \right)$$

$$V_s^* V_h = 2g^2 \left(\left(1 + \frac{2\zeta}{x^2}\right) k_\perp^2 - (1-\zeta)(m_{b,s}^2 + m_{b,h}^2) + 2\zeta m_c^2 \frac{1-x}{x^2} \right)$$

$$V_h^* V_s = 2g^2 \left(\left(1 + \frac{2\zeta}{x^2}\right) k_\perp^2 - (1-\zeta)(m_{b,s}^2 + m_{b,h}^2) + 2\zeta m_c^2 \frac{1-x}{x^2} \right)$$

$$V_h^* V_h = 2g^2 \left(\left(1 + \frac{2\zeta}{x^2}\right) k_\perp^2 - 2(1-\zeta)m_{b,h}^2 + 2\zeta m_c^2 \frac{1-x}{x^2} \right)$$



$$\zeta = 1 \implies |V|^2 \propto \frac{k_\perp^2}{x^2}$$

from A&P and same as
B&M

Reformulation of the Matrix Element

Feynman gauge: $\zeta = 0$

$$V_s^* V_s = 2g^2 k_\perp^2$$

$$V_s^* V_h = 2g^2 (k_\perp^2 - m_{b,h}^2)$$

$$V_h^* V_s = 2g^2 (k_\perp^2 - m_{b,h}^2)$$

$$V_h^* V_h = 2g^2 (k_\perp^2 - 2m_{b,h}^2)$$

Our A's
→
Soft and
collinear limit
 $k_\perp \rightarrow \lambda k_\perp$,
 $x \rightarrow \lambda x$

$$\begin{aligned}\frac{V_s^* V_s}{A_s^2} &= O(1) \\ \frac{V_s^* V_h}{A_s A_h} &= -2g^2 \frac{1}{\lambda^2 k_\perp^2} \frac{x^2}{x^2 + k_\perp^2/m_{b,h}^2} + O(1/\lambda) \\ \frac{V_h^* V_s}{A_h A_s} &= -2g^2 \frac{1}{\lambda^2 k_\perp^2} \frac{x^2}{x^2 + k_\perp^2/m_{b,h}^2} + O(1/\lambda) \\ \frac{V_h^* V_h}{A_h^2} &= -2g^2 \frac{2}{\lambda^2 x^2 m_{b,h}^2} \left(\frac{x^2}{x^2 + k_\perp^2/m_{b,h}^2} \right)^2 + O(1/\lambda)\end{aligned}$$

$$|\mathcal{M}|^2 = 4g^2 \frac{1}{\lambda^2 k_\perp^2} \left(\frac{x^2}{x^2 + k_\perp^2/m_{b,h}^2} \right)^2 + O(1/\lambda)$$

Reformulation of the Matrix Element

Axial gauge: $\zeta = 1$ and choose auxiliary vector $n = p_{a,s}$

$$V_s = 0 \quad \implies \quad |V_h|^2 = \text{Tr} [(\not{p}_{b,h} - m_{b,h}) \gamma^\mu (\not{p}_{b,h} - m_{b,h}) \gamma^\nu] \left(-g^{\mu\nu} + \frac{p_c^\mu p_{a,s}^\nu + p_c^\nu p_{a,s}^\mu}{p_{a,s} p_c} \right)$$

$$= 8 p_{b,h}^\mu p_{b,h}^\nu \left(-g^{\mu\nu} + \frac{p_c^\mu p_{a,s}^\nu + p_c^\nu p_{a,s}^\mu}{p_{a,s} p_c} \right)$$

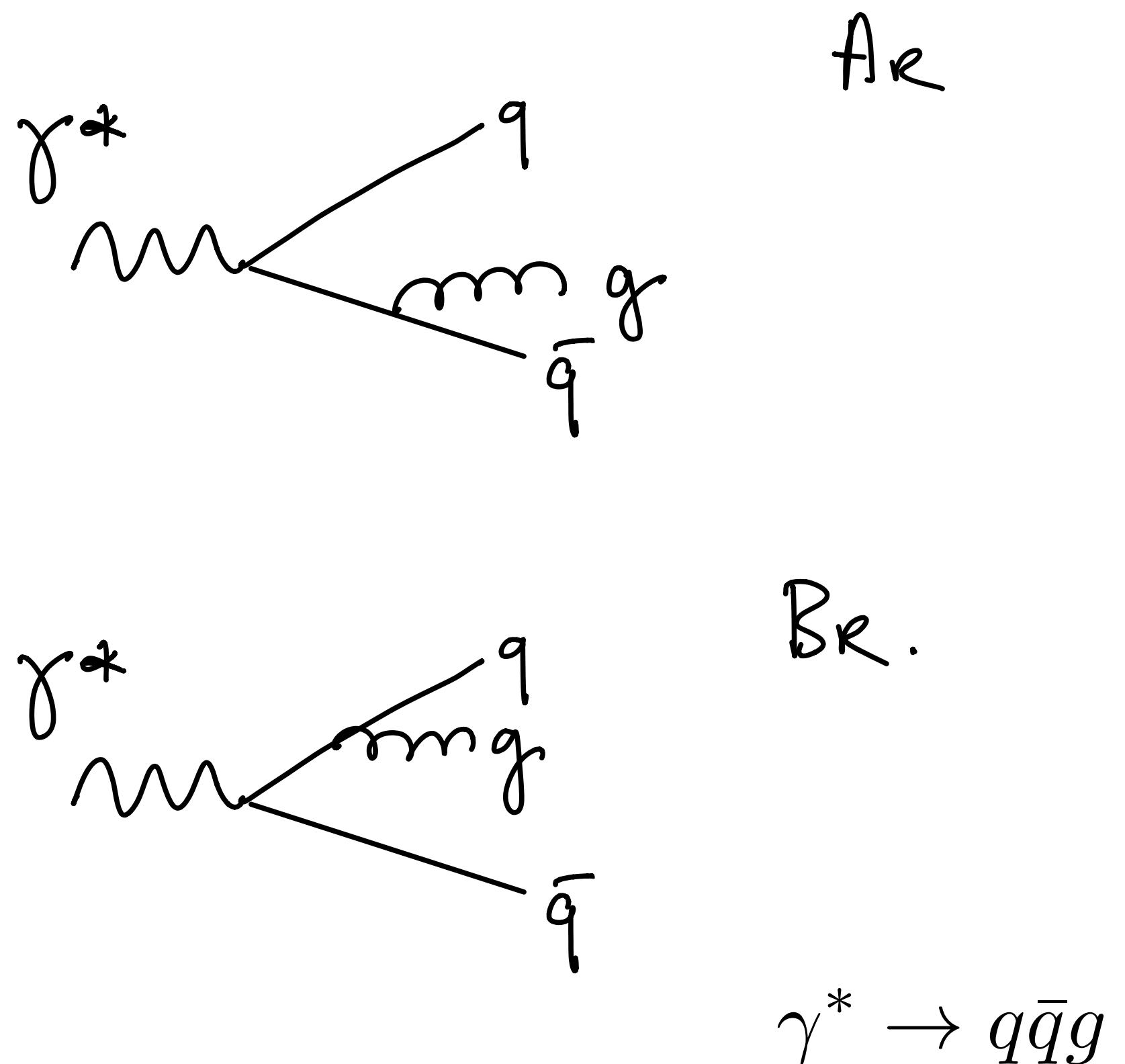
$$= 8 \left(2 p_{b,h} p_c \frac{p_{a,s} p_{b,h}}{p_{a,s} p_c} - m_{b,h}^2 \right)$$

$$\frac{\sum |V_h|^2}{A_h^2} = \frac{2 p_{a,s} p_{b,h}}{p_{a,s} p_c p_{b,h} p_c} - \frac{m_{b,h}^2}{(p_{b,h} p_c)^2} \rightarrow 4 g^2 \frac{1}{\lambda^2 k_\perp^2} \left(\frac{x^2}{x^2 + k_\perp^2 / m_{b,h}^2} \right)^2 + O(1/\lambda)$$

Same as Feynman gauge

Reformulation of the Matrix Element

See R Field's "Application of Perturbative QCD" Chp 3



Feynman gauge

$$S_{11} = |B_R|^2 \propto \frac{(1-x)}{t}$$
$$S_{22} = |A_R|^2 = 0$$
$$S_{12} = 2A_R^* B_R \propto \frac{x}{t}$$

Axial gauge

$$S_{11} = |B_R|^2 \propto \frac{1}{t}$$
$$S_{12} = S_{22} = 0$$

Total matrix element gauge invariant
But certain gauge choices can make subamplitudes differ between gauges.

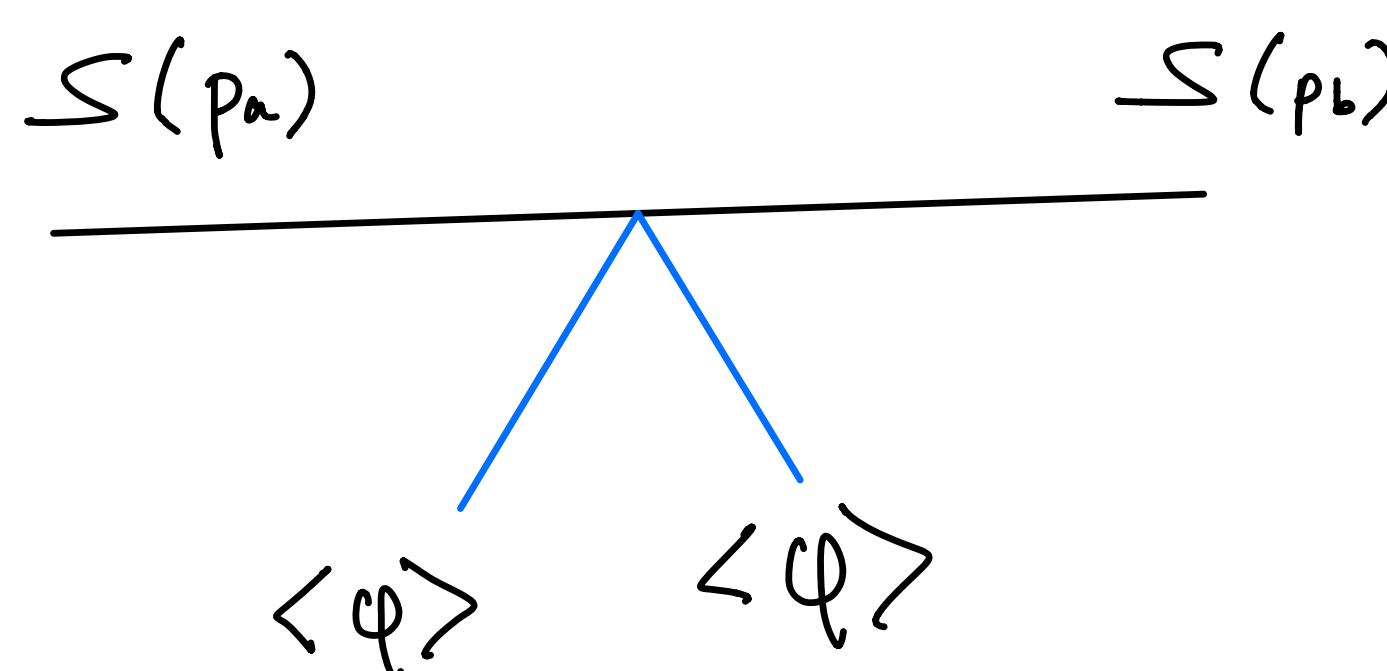
Phase change dependence

Points for discussion and future work

Where does the dependence in the phase change enter?

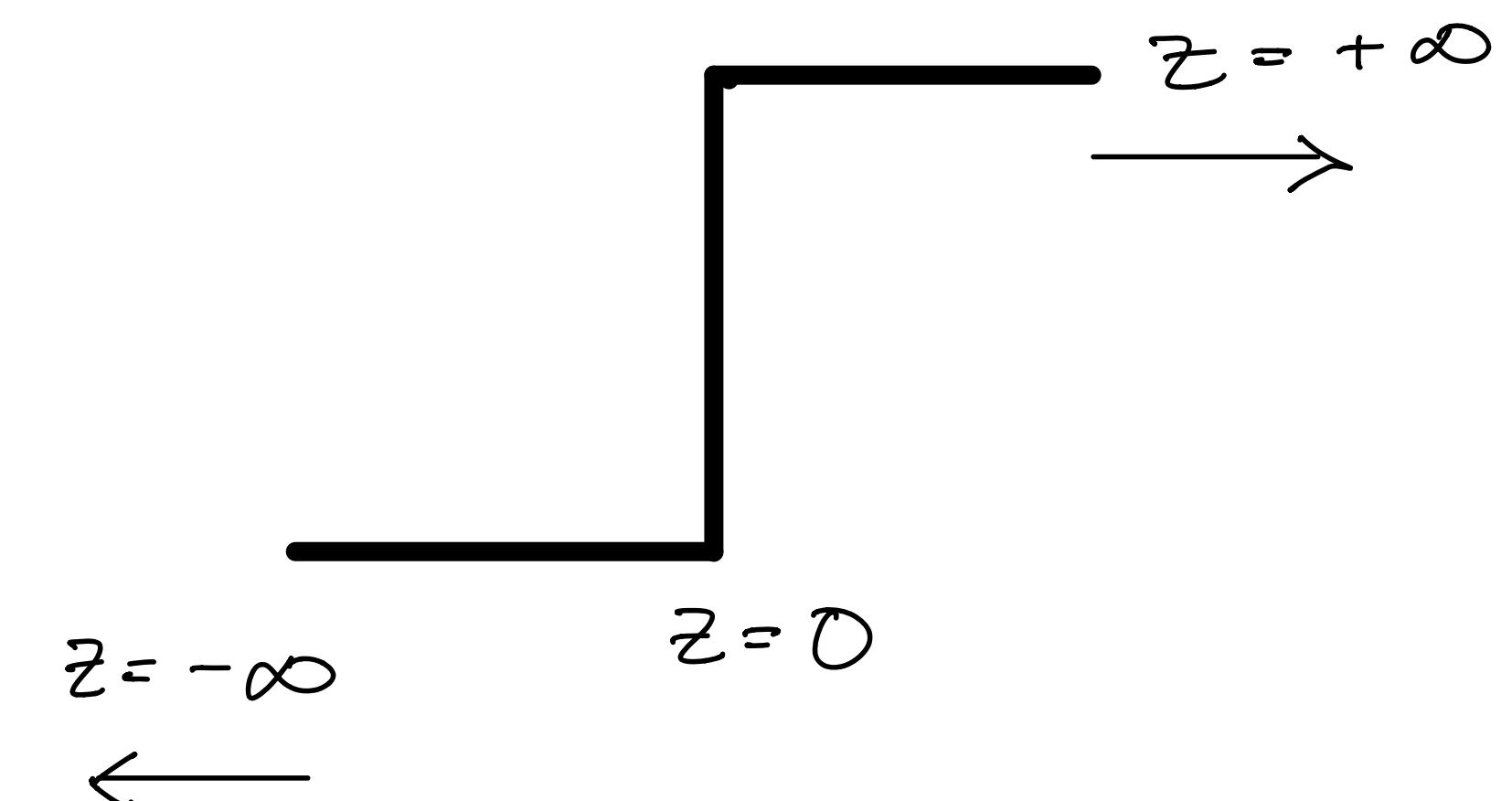
To perform a resummation or even define what the quantum correction is, we need to define what the leading order process is (aka Born-level process/hard function)

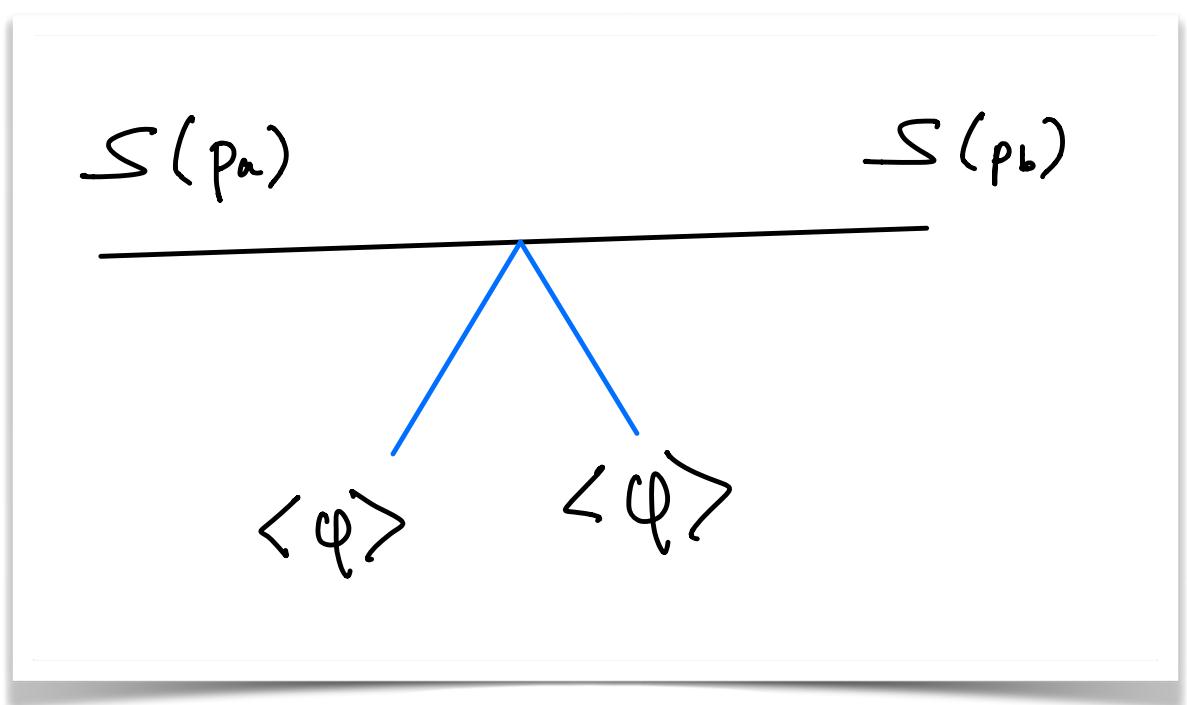
Lets consider our hard function to be a scalar interacting with the wall (fermions follow analogously). Leading order process is calculated from VEV-insertion:



A yellow box containing the following equations:

$$L_{\text{int}} \supset -\frac{\lambda}{4} S^2 \Phi^2$$
$$\Phi = \varphi(x, t) + \hat{\varphi}$$
$$\varphi(z) = \varphi_0 \Theta(z)$$





$$\mathcal{M}_{a \rightarrow b}^{(0)} = i \frac{\lambda}{4} \left\langle p_b \left| T \int d^4x S(x, t)^2 \varphi(x, t)^2 \right| p_a \right\rangle$$

**Fourier transform
step function profile of the wall**

$$\varphi(x, t)^2 = \varphi_0^2 \int \frac{dk_z}{2\pi} \int \frac{dk'_z}{2\pi} e^{i(k_z + k'_z)z} \left(\frac{\delta(k_z)}{2} - \frac{i}{2\pi} PV \frac{1}{k_z} \right) \left(\frac{\delta(k'_z)}{2} - \frac{i}{2\pi} PV \frac{1}{k'_z} \right)$$

Perform integrations → $\frac{1}{4\pi^3} \frac{1}{\Delta p_z}$

Scalar quantised usual way → $\langle p_b | S(x, t)^2 | p_a \rangle = \langle p_b | S(x, t) | 0 \rangle \langle 0 | S(x, t) | p_a \rangle = e^{i(p_a - p_b) \cdot x}$

$$S(x, t) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} (a_p e^{-ip \cdot x} + a_p^\dagger e^{ip \cdot x})$$

Put it all together

$$P_{a \rightarrow b} = \frac{\lambda \varphi_0^2}{8\pi^3} \delta(E_a - E_b) \delta^2(p_{a,\perp} - p_{b,\perp}) \frac{1}{\Delta p_z}$$

$E_a^2 \gg m^2$ →

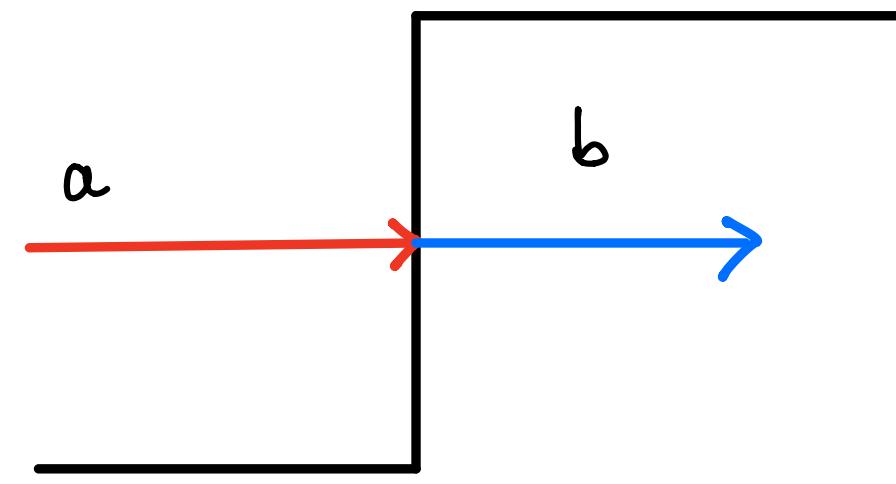
$$\Delta p_z = \frac{m_{b,h}^2 - m_{a,s}^2}{2E_a} \approx \frac{m_{b,h}^2}{2E_a} \quad (m_{a,s} \ll m_{b,h})$$

$$P_{a \rightarrow b} = \frac{\cancel{\lambda} \varphi_0^2}{8\pi^3} \delta(E_a - E_b) \delta^2(p_{a,\perp} - p_{b,\perp}) \frac{1}{\cancel{\Delta p_z}}$$

$m_{b,h}^2 \gg m_{a,s}$ →

$$P_{a \rightarrow b} = \frac{m_{b,h}^2}{8\pi^3} \delta(E_a - E_b) \delta^2(p_{a,\perp} - p_{b,\perp}) \frac{E_a}{m_{b,h}^2}$$

Recall the B&M 1-to-1 calculation



$$dP_{a \rightarrow b} = \frac{d^3 p_b}{(2\pi)^3} \frac{1}{2E_b} \times (2\pi)^3 \delta^2 (p_{a,\perp} - p_{b,\perp}) \delta(E_a - E_b) (2p_{b,z,h}) \delta_{ab}$$

$$\mathcal{P}_{1 \rightarrow 1} = \sum_a \nu_a \int \frac{d^3 p_a}{(2\pi)^3} f_a (1 \pm f_b) \Delta p_{\text{wall}} \approx \gamma^0 T^2 \Delta m^2$$

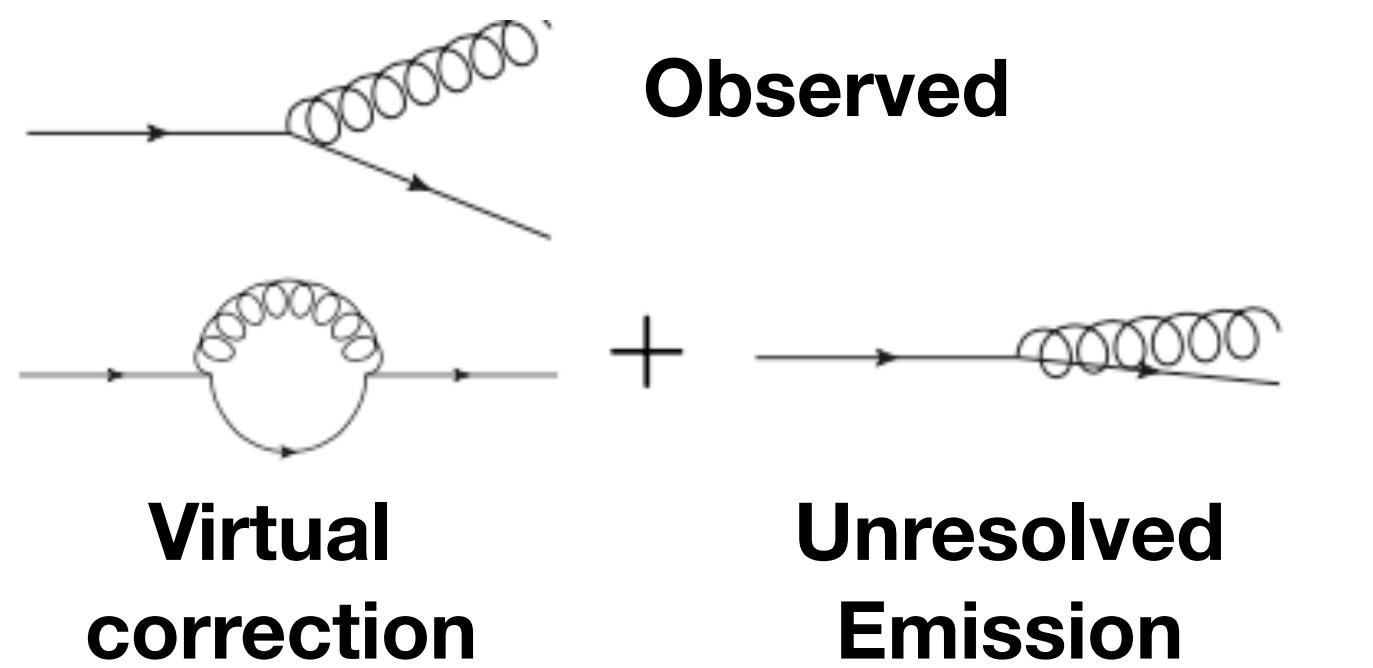
Mass dependence comes from observable not from matrix element

Summary: treat 1-to-1 as leading order process using infinitely thin wall approximation. As long as the ingoing particles gains a mass it will interact with a probability of 1 following from VEV-insertion method. This hard function/LO process normalises our NLO process. This procedure of defining LO process that normalises NLO process is tried and tested (> 30 years of collider data) and follows from the universal factorisation properties of gauge theories in the soft limit.

Resummation

Radiative corrections as branching processes

Marchesini & Webber (1983)
Sjöstrand (1985)



$$P_{\text{no em}} + P_{em} = 1$$

Unobserved processes captured by Sudakov factor

Expression for change in population:

N = population

$$dN = -\lambda N dt$$

λ = decay constant

Survival probability at time “t”: $e^{-\lambda t}$ where $t \leftrightarrow$ energy scale ($\log(1/v)$)

Change in population analogous to boson emission probability

$$\lambda N dt = \int [dk] M^2(k) \Theta(v - V(\{\tilde{p}\}, k)) \quad \xrightarrow{\hspace{1cm}} \quad [\text{virt.} + \text{unres.}] = e^{-\int [dk] M^2(k) \Theta(V(\{\tilde{p}\}, k) - v)}$$

Probability of not emitting bosons above “v” (Sudakov)

Kinematics

$$p_a^\mu = \left(E_a, \vec{0}, \sqrt{E_a^2 - m_a^2} \right) \approx \left(E_a, \vec{0}, E_a \left(1 - \frac{m_a^2}{2E_a^2} \right) \right)$$
$$p_b^\mu = \left((1-x)E_a, -\vec{k}_\perp, \sqrt{(1-x)^2 E_a^2 - \vec{k}_\perp^2 - m_b^2} \right) \approx \left((1-x)E_a, -\vec{k}_\perp, (1-x)E_a \left(1 - \frac{\vec{k}_\perp^2 + m_b^2}{2(1-x)^2 E_a^2} \right) \right)$$
$$p_c^\mu = \left(xE_a, \vec{k}_\perp, \sqrt{x^2 E_a^2 - \vec{k}_\perp^2 - m_c^2} \right) \approx \left(xE_a, \vec{k}_\perp, xE_a \left(1 - \frac{\vec{k}_\perp^2 + m_c^2}{2x^2 E_a^2} \right) \right)$$