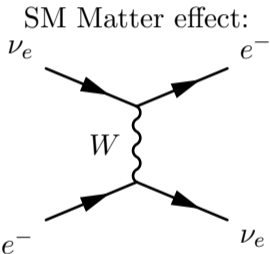


CP-Violating Neutrino Non-Standard Interactions in Long-Baseline-Accelerator Data

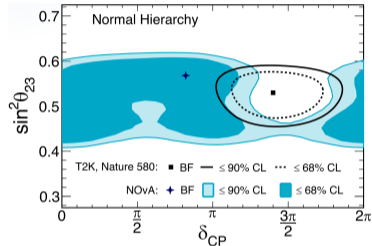
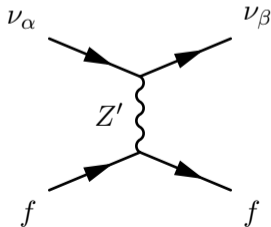
By **Peter B. Denton** with Julia Gehrlein and Rebekah Pestes: [2008.01110](#) (PRL)



L. Wolfenstein [PRD 17, 2369 \(1978\)](#)

- ▶ modifies oscillations
- ▶ sort of measured:
- ▶ compare solar and KamLAND

What if there is a new “matter effect” style interaction?



A. Himmel [10.5281/zenodo.3959581](#)

T2K & NOvA see different δ_{CP}
NSI effect $\propto \rho E$

	ρ [g/cc]	E [GeV]
T2K	2.6	0.6
NOvA	2.84	1.9

Does NSI fit?

CP-Violating Neutrino Non-Standard Interactions in Long-Baseline-Accelerator Data

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How big of NSI?

$$|\epsilon_{e\mu}| \approx \frac{s_{12}c_{12}c_{23}\pi\Delta m_{21}^2}{2s_{23}^2}$$

$$\times \left| \frac{\sin \delta_{\text{T2K}} - \sin \delta_{\text{NOvA}}}{a_{\text{NOvA}} - a_{\text{T2K}}} \right|$$

$$\approx 0.22$$

... and the phase?

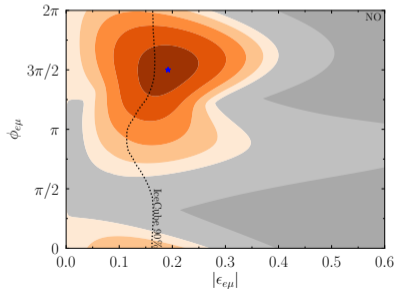
$$\cos(\delta_{\text{true}} + \phi_{e\beta}) \approx -1$$

$$\delta_{\text{true}} \approx \delta_{\text{T2K}} \approx 1.5\pi$$

$$\Rightarrow \phi_{e\beta} \approx 1.5\pi$$

Fit the data:

- ▶ NOvA & T2K app & dis
- ▶ Daya Bay
- ▶ KamLAND
- ▶ \Rightarrow reactors have no matter effect



Where do we go from here?

- ▶ NOvA & T2K will improve
- ▶ DUNE & HK will improve more
- ▶ IceCube & SK are comparably sensitive
- ▶ COHERENT too (depends on mediator)

Backup

$$\mathcal{L}_{\text{NSI}} = -2\sqrt{2}G_F \sum_{\alpha,\beta,f,P} \epsilon_{\alpha\beta}^{f,P} (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{f} \gamma_\mu P f)$$

Models with large NSIs consistent with CLFV:

Y. Farzan, I. Shoemaker [1512.09147](#) Y. Farzan, J. Heeck [1607.07616](#) D. Forero and W. Huang [1608.04719](#)
 K. Babu, A. Friedland, P. Machado, I. Mocioiu [1705.01822](#) [PBD](#), Y. Farzan, I. Shoemaker [1804.03660](#)
 U. Dey, N. Nath, S. Sadhukhan [1804.05808](#) K. Babu, et al. [1907.09498](#) Y. Farzan [1912.09408](#)

Affects oscillations via new matter effect

$$H = \frac{1}{2E} \left[UM^2U^\dagger + a \begin{pmatrix} 1 + \epsilon_{ee} & \epsilon_{e\mu} & \epsilon_{e\tau} \\ \epsilon_{e\mu}^* & \epsilon_{\mu\mu} & \epsilon_{\mu\tau} \\ \epsilon_{e\tau}^* & \epsilon_{\mu\tau}^* & \epsilon_{\tau\tau} \end{pmatrix} \right]$$

Matter potential: $a \propto G_F \rho E$

B. Dev, K. Babu, [PBD](#), P. Machado, et al. [1907.00991](#)