

Parallelisation of (Track) Fits

HighRR Lecture Week

27. Sep. - 1. Oct .2022

DFG Deutsche
Forschungsgemeinschaft

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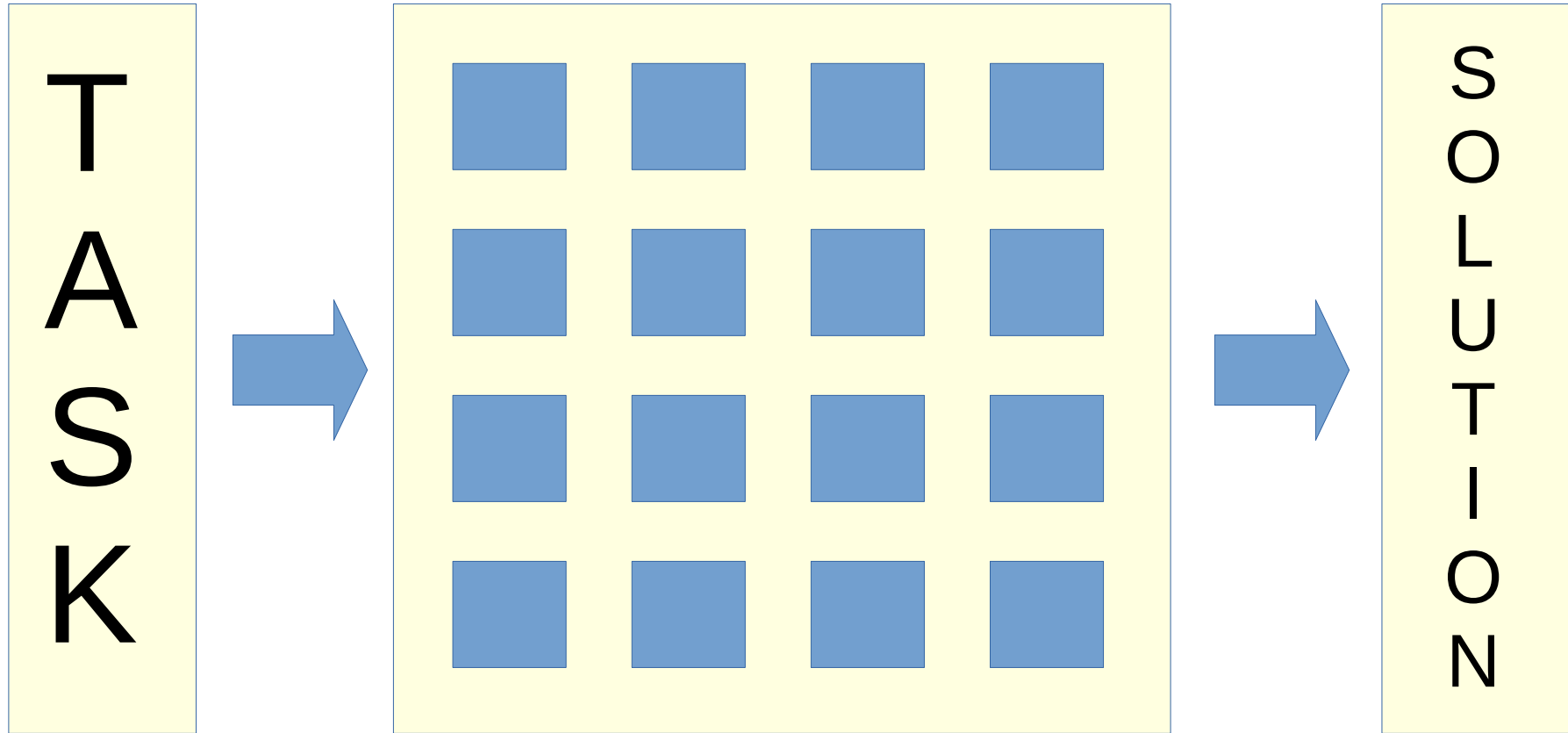


→ with a clear focus on algorithms



YOU

Parallelisability



Overview

- Motivation
- Intro to Track Fits (Overview)
- Fitting Tracks with Hit Uncertainties
- Linearisation
- **A New Hit Uncertainty Track Fit**
 - Triplet Representation (parallelisable)
 - Cholesky Decomposition (non-parallelisable)
- Summary

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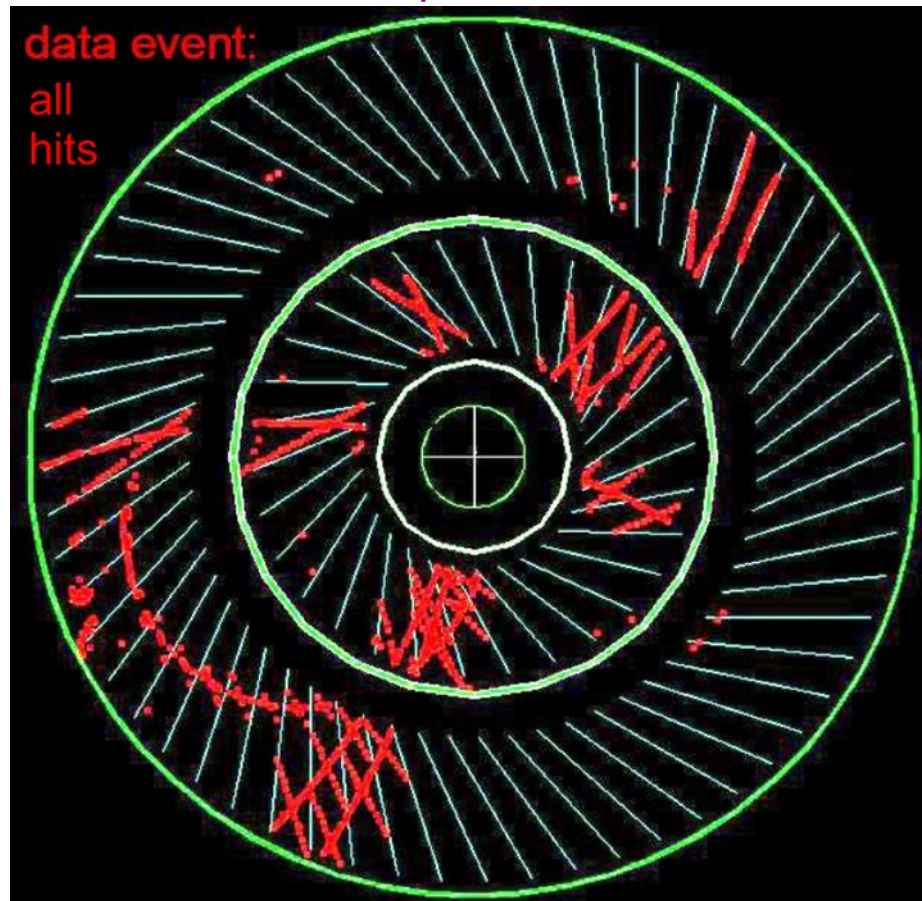


I will try to avoid formulas whenever possible

What do you see?

Drift chamber hits of the H1 experiment
(1991-2007)

- low track multiplicity
- many measurement points per track
- despite some ambiguities track finding is very easy

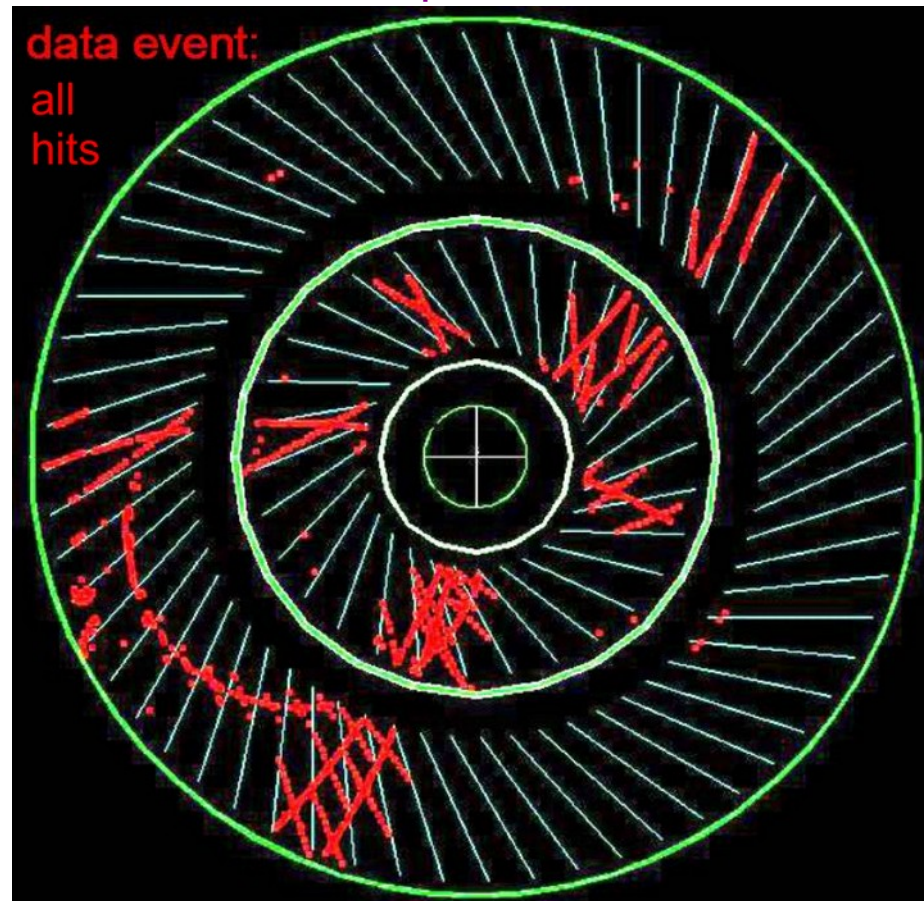


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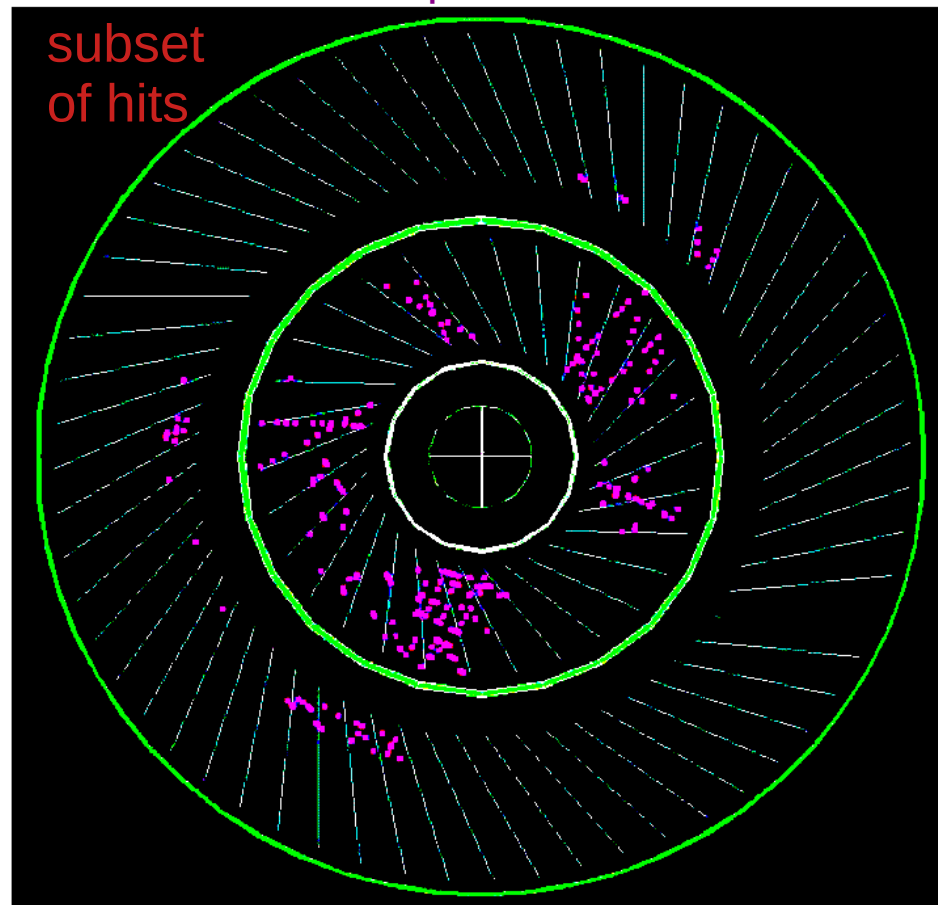
- no physics model is needed
- any amateur can find tracks!



What do you see?

Drift chamber hits of the H1 experiment
(1991-2007)

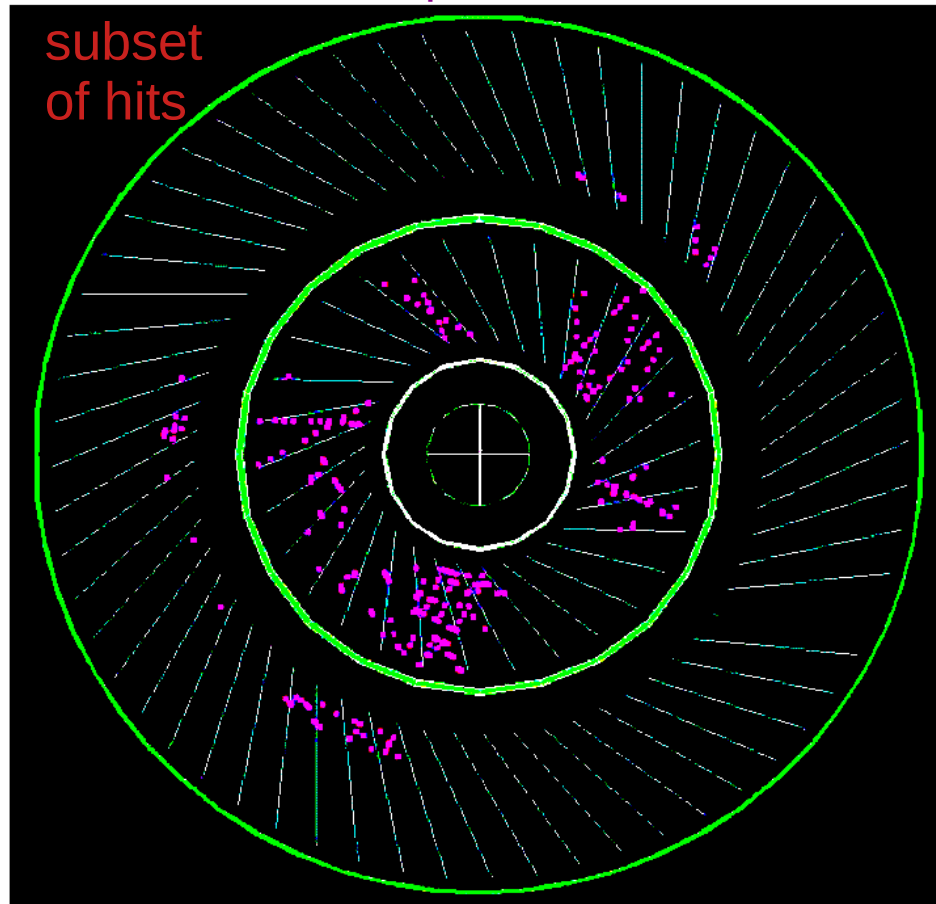
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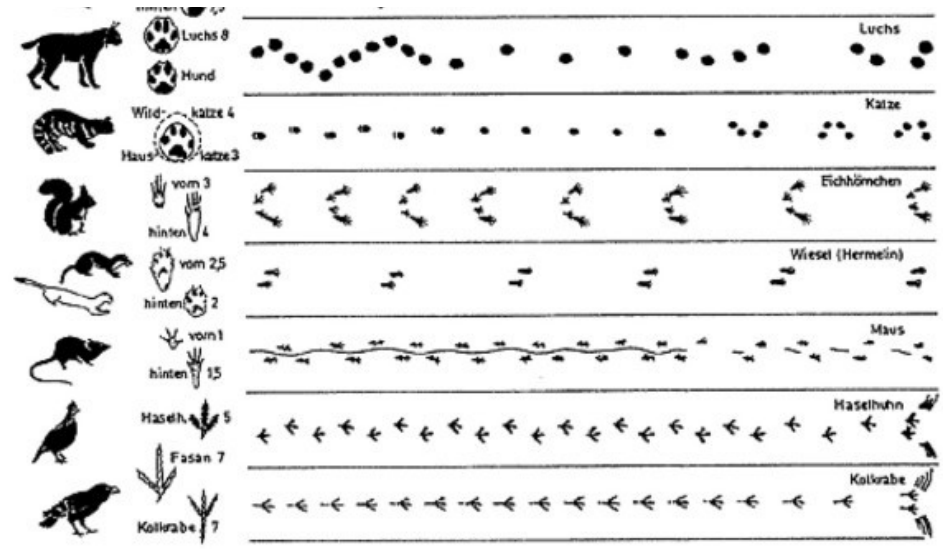
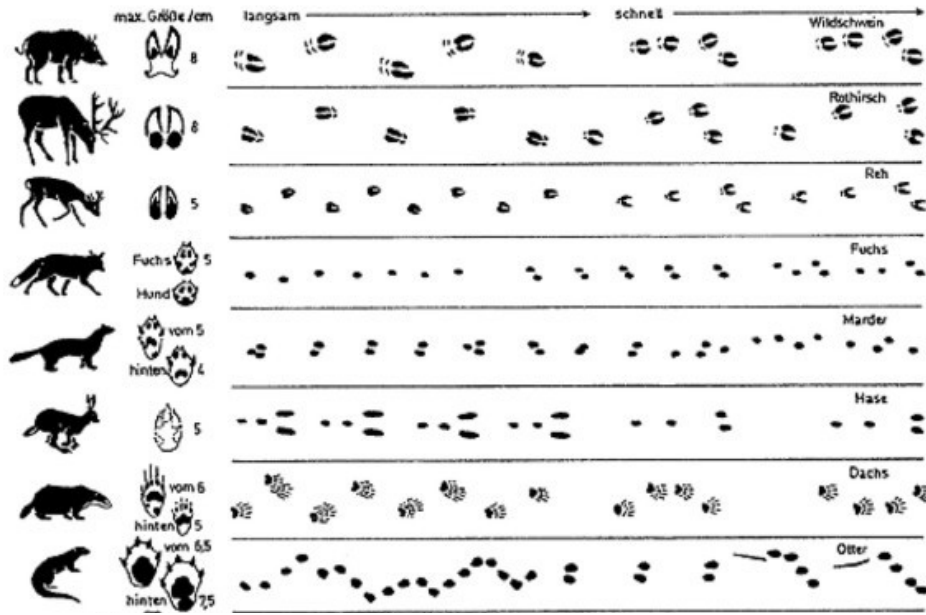
What do you see?

Drift chamber hits of the H1 experiment
(1991-2007)

- low track multiplicity
 - many measurement points per track
 - despite some ambiguities track finding is very easy
-
- **but in the situation of sparse information the task becomes much more difficult!**
 - **physics model** (B-field, lorentz force, momentum conservation) **is required to reconstruct the tracks**



Sparse Detector (Hit) Information



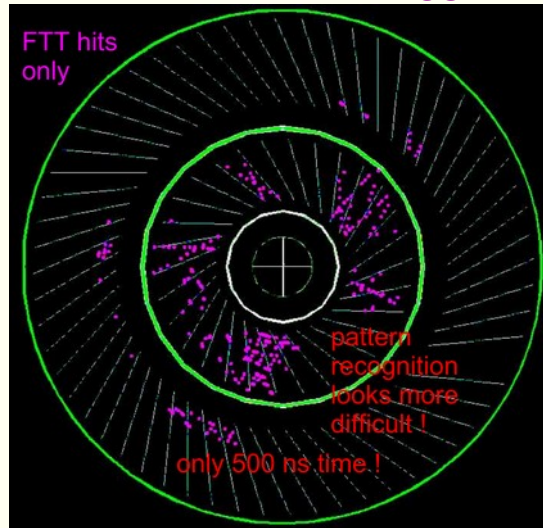
Knowledge helps to find (identify) tracks → **should make use of it!**

Examples for Sparse Hit Information

Trigger:

- limitations from bandwidth & processing power

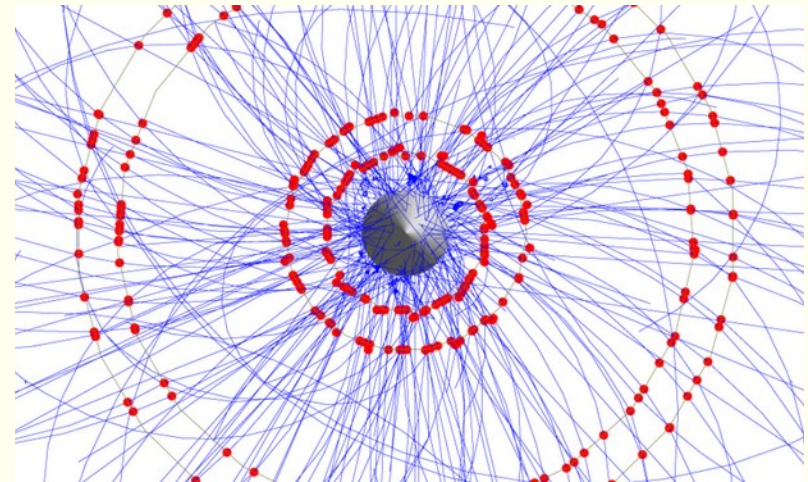
FTT = Fast Track Trigger



Semiconductor Trackers:

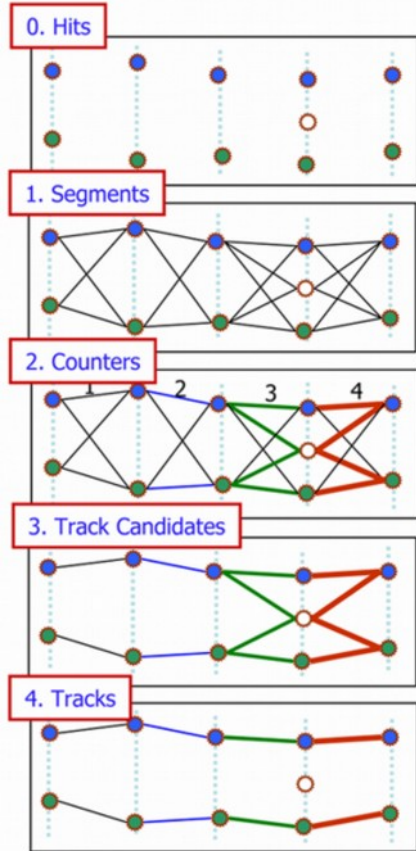
- limitations from multiple scattering (resolution) & powering & costs

Mu3e: only four (pixel) tracking layers



Track Reconstruction Example I

Cellular Automaton (e.g. CBM experiment, ALICE):



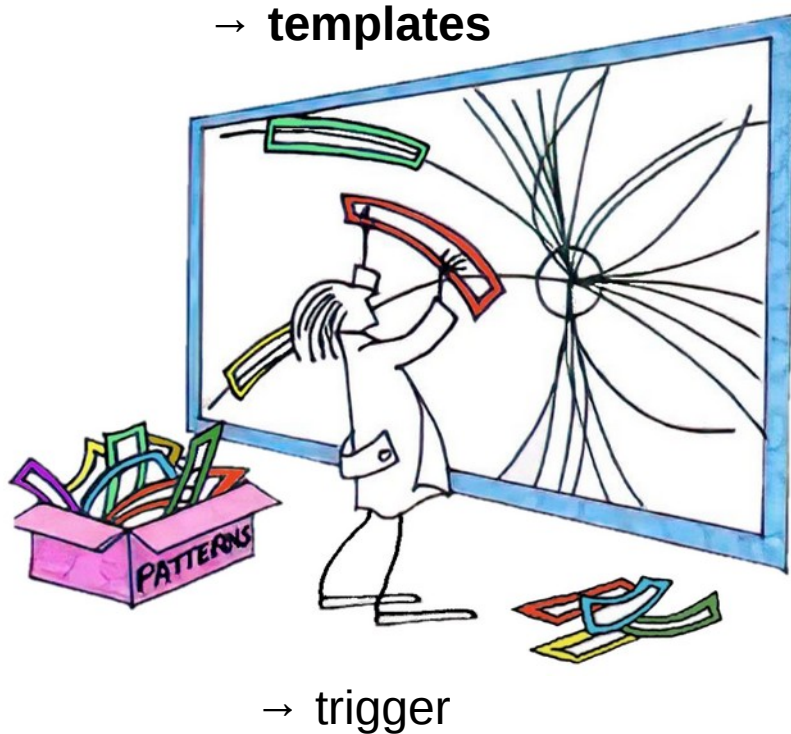
- **local** method based on segments
- uses mostly **topological** information
- parallelisable
- implemented e.g. on GPUs

works only if hit density is dense enough

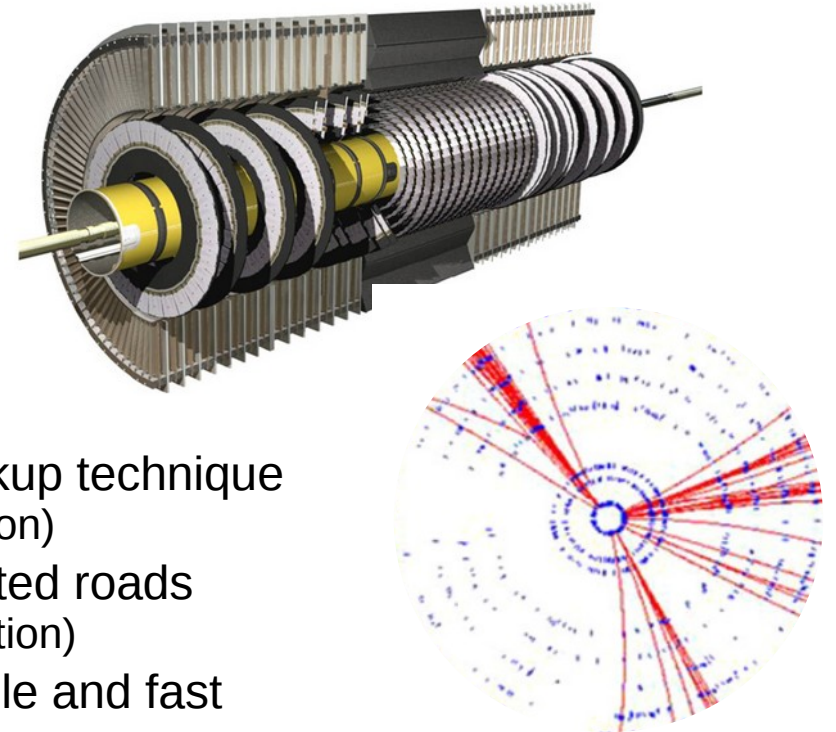
sketch from I.Kisel

Track Reconstruction Example II

ATLAS Fast Tracker Project (2019†)



ATLAS SCT



- pattern lookup technique (approximation)
- pre-calculated roads (from simulation)
- parallelisable and fast
- implemented in ASICs (AMchip)

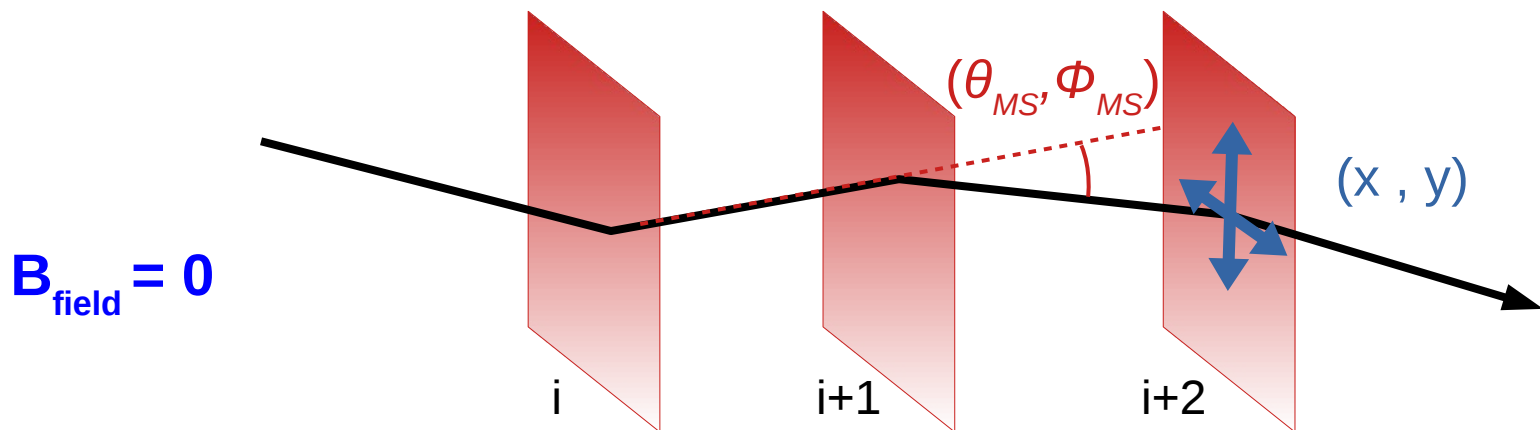
Track Reconstruction Example III

Full track fit:

$$\chi^2 = \sum_{\text{layer } i} \frac{\Theta_{\text{MS},i}^2}{\sigma_{\theta,i}^2} + \frac{\Phi_{\text{MS},i}^2}{\sigma_{\phi,i}^2} + \sum_{\text{hits } jk} (x_j - \xi_j) V_{jk}^{-1} (x_k - \xi_k)$$

includes **multiple scattering** and **hit uncertainties**

not best option
to fit all track
hypothesis



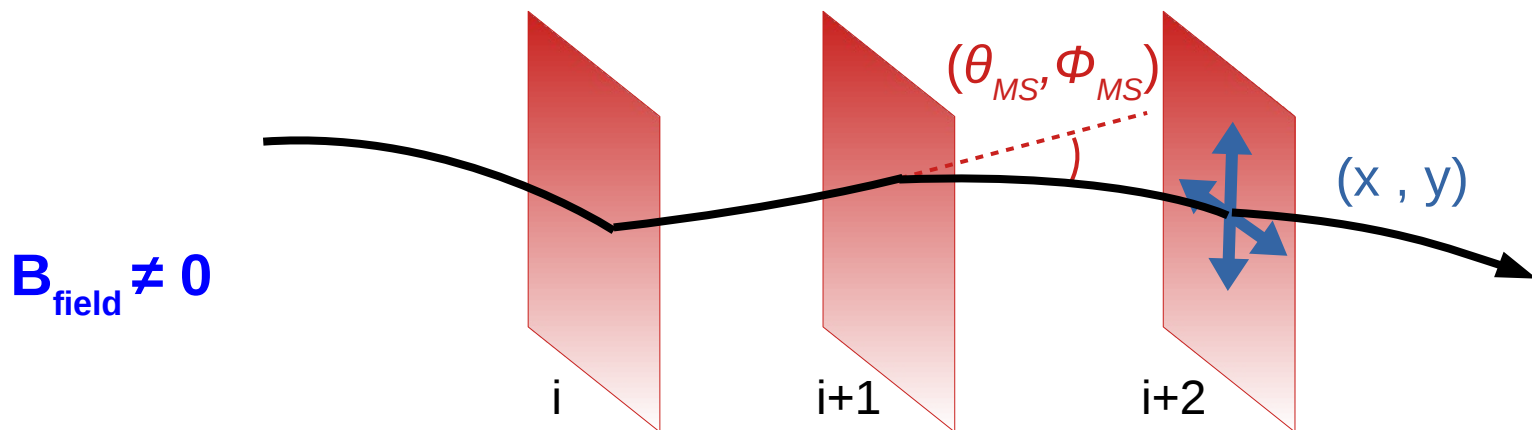
Track Reconstruction Example III

Full track fit:

$$\chi^2 = \sum_{\text{layer } i} \frac{\Theta_{\text{MS},i}^2}{\sigma_{\theta,i}^2} + \frac{\Phi_{\text{MS},i}^2}{\sigma_{\phi,i}^2} + \sum_{\text{hits } jk} (x_j - \xi_j) V_{jk}^{-1} (x_k - \xi_k)$$

includes **multiple scattering** and **hit uncertainties**

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hypothesis



Full glory fit (magnetic field) is computationally intensive!

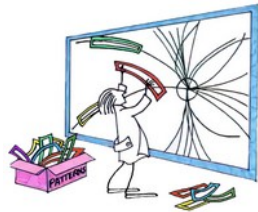
High Track Multiplicities (ATLAS)

Motivation for fast (& full) tracking:

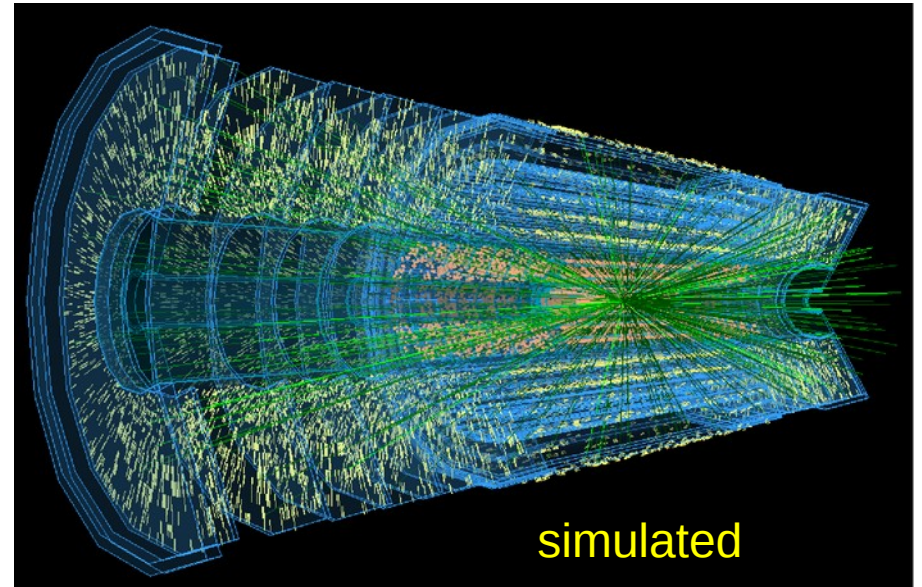
- Identify special or rate track based signatures (e.g. long lived particles)
- track-assisted object reconstruction for tracker (e.g. high energy particles)

ATLAS Approaches

- **FTK** Project (2019†)
- **Phase II** (High Lumi-LHC) Hardware Track Trigger Project (**HTT**, 2021†)
- **New**: fast tracking on **Event Filter** (EF)
 - option A: **CPU** only?
 - option B: **GPU-** or **FPGA-** accelerated?
 - provide highly parallel computing architectures!



Possible to reconstruct all tracks?



ATLAS High Luminosity Inner Tracker (ITK)
with **200 pileup** events at **40 MHz** collisions

Question to students:

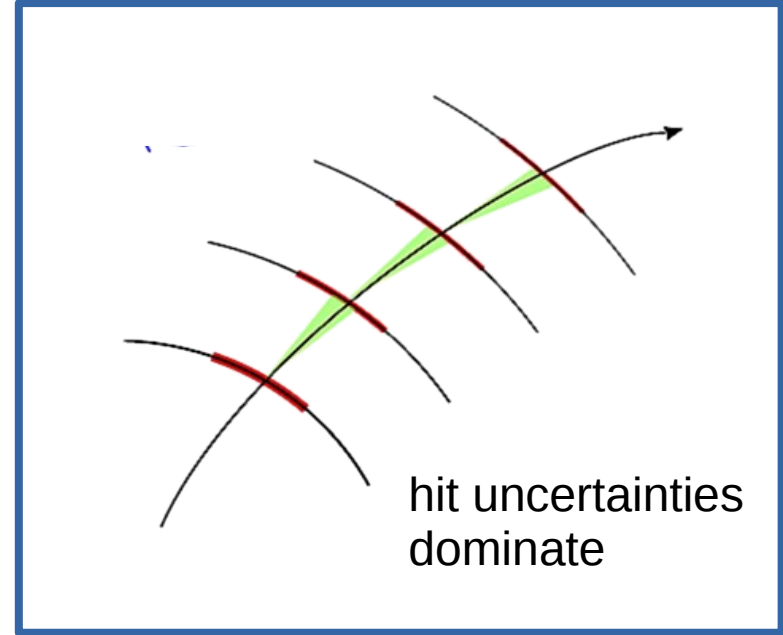
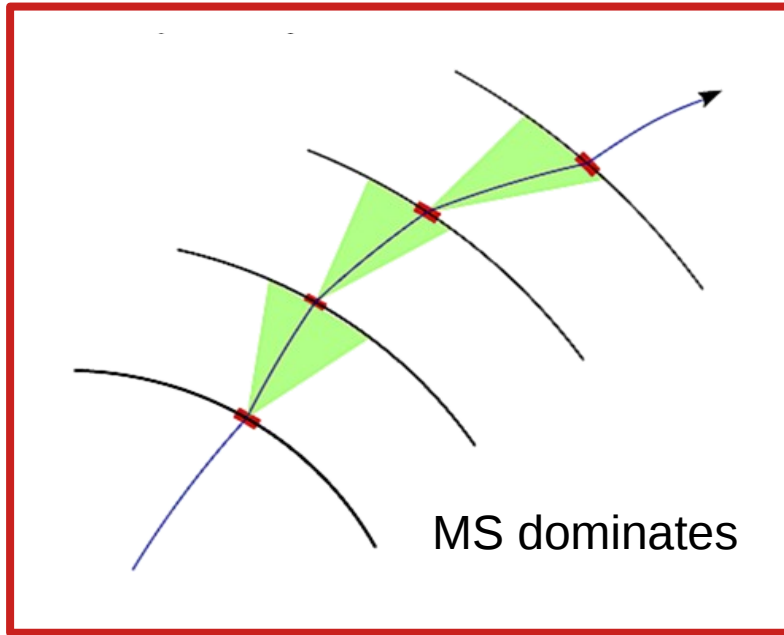
What is your favorite tracking
concept or algorithm?
And why?

Chapter 2

Introduction to Track Fits

The Master Equation

$$\chi^2 = \sum_{\text{layer } i} \frac{\Theta_{\text{MS},i}^2}{\sigma_{\theta,i}^2} + \frac{\Phi_{\text{MS},i}^2}{\sigma_{\phi,i}^2} + \sum_{\text{hits } jk} (x_j - \xi_j) V_{jk}^{-1} (x_k - \xi_k)$$



The Master Equation

$$\chi^2 = \sum_{\text{layer } i} \frac{\Theta_{\text{MS},i}^2}{\sigma_{\theta,i}^2} + \frac{\Phi_{\text{MS},i}^2}{\sigma_{\phi,i}^2} + \sum_{\text{hits } jk} (x_j - \xi_j) V_{jk}^{-1} (x_k - \xi_k)$$

Problems and difficulties:

- hit error matrix V^{-1} depends on the trajectory (result)
- scattering uncertainties and scattering angles depend on trajectory (result)
- fitted hit positions are all correlated → **no local processing possible**
- single outliers can spoil the fit → **iterative outlier rejection**

non-locality and iterations are the enemy of parallelisation

Kálmán Filter (KF)

The Kálmán Filter is an algorithm method which combines


- **track finding** (aka hit linking) and
- trajectory determination (**track parameter fitting**)


Note, the KF as such does not implement any physics!

Simple example:

Calculation of an average

mean value over n measurements


$$\mu_n = \frac{1}{n} \sum_{i=1}^n x_i$$


$$\mu_{n+1} = \frac{1}{n+1} \sum_{i=1}^{n+1} x_i = \frac{n}{n+1} \left(\frac{1}{n} \sum_{i=1}^n x_i + \frac{1}{n} x_{n+1} \right) = \frac{n}{n+1} \mu_n + \frac{1}{n+1} x_{n+1} = \mu_n + \frac{1}{n+1} (x_{n+1} - \mu_n)$$

mean value over $n+1$ measurements

previous estimation

new measurement

weight

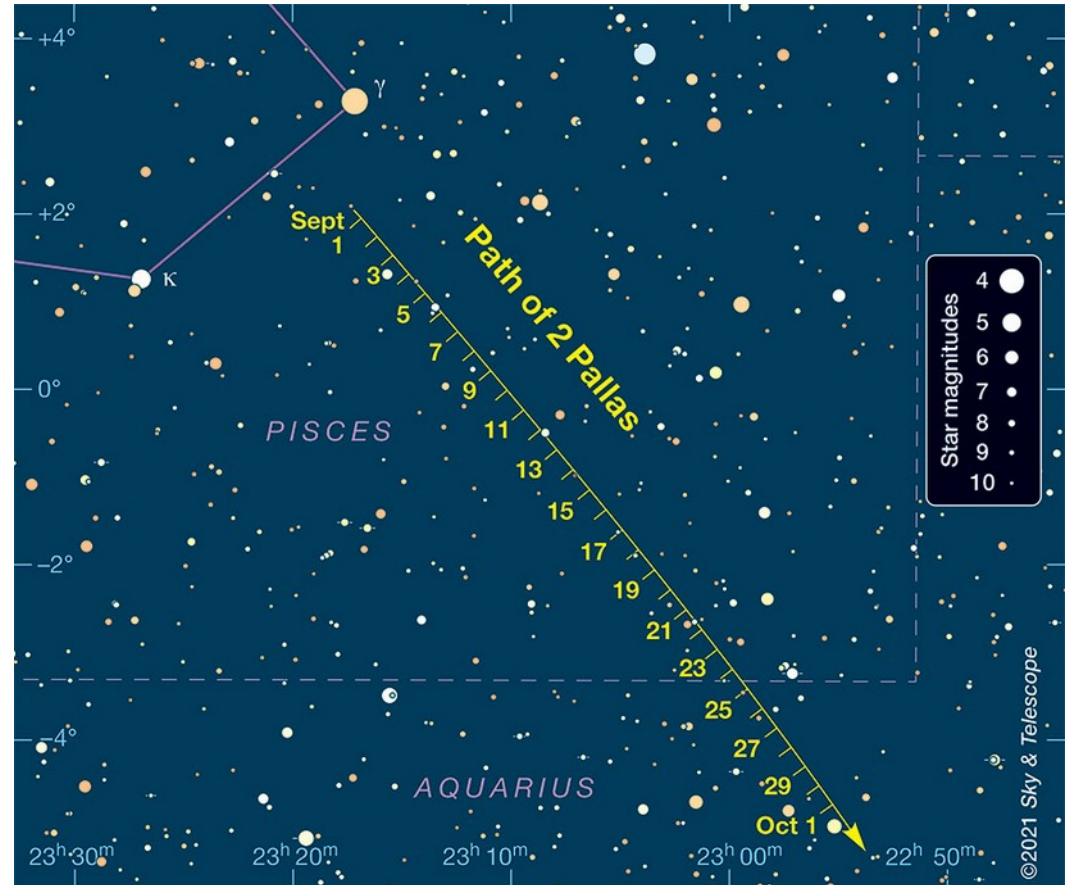
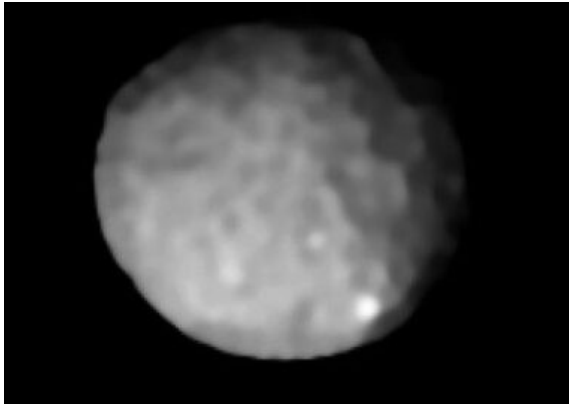
correction

from I.Kisel

Example: Asteroids

Asteroid Pallas

(first found by German astronomer Heinrich Olbers in 1807)



Tracking of Asteroids

Suppose we discover an asteroid!

State vector described by

$$\vec{r} = (x, y, z, v_x, v_y, v_z)$$

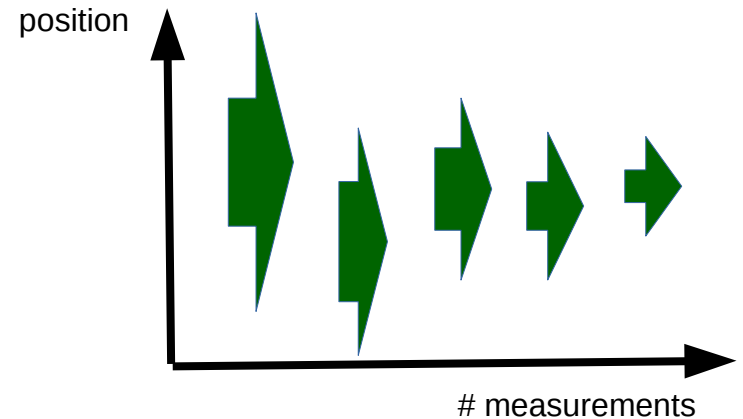
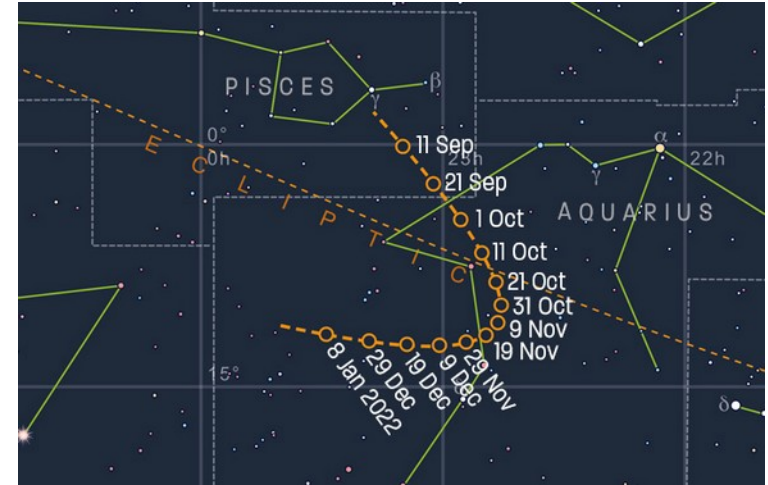
- 1st measurement (11.9): → **2d** information
- 2nd measurement (21.9): → 2d+2d=**4d** information
- 3rd measurement (1.10): → 4d+2d=**6d** information

→ able to reconstruct state vector \mathbf{r}
and first guess of error matrix

- 4th measurement (11.10): → **update** state and error matrix
- 5th measurement (21.10): → **update** state and error matrix

...

→ **the more measurements, the more precise!**

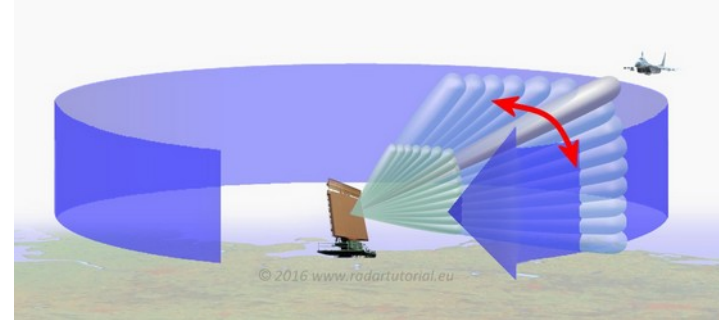


Rudolf Kálmán (1930-2016)

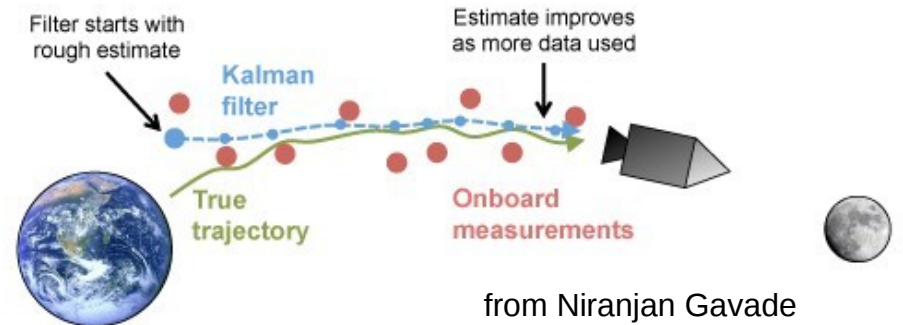


Kálmán (emeritus ETH professor) receiving the National Medal of Science from US president Obama in Oktober 2009.

Tracking of air-planes

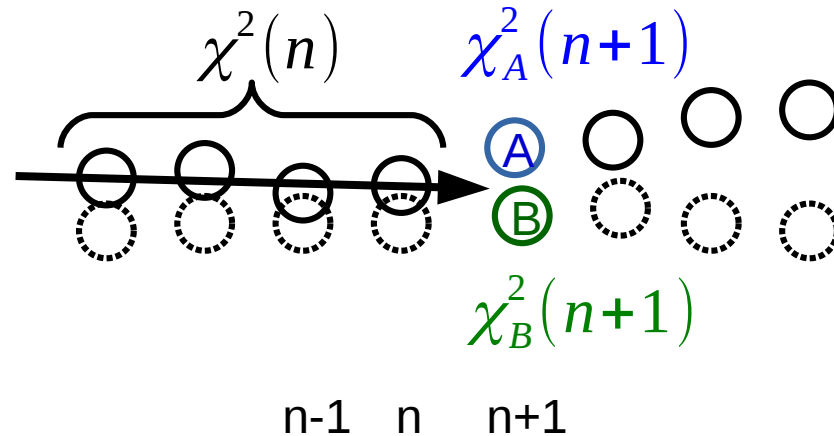


Tracking of space crafts (NASA Apollo mission)



Kálmán Filter Applied to Tracking

Hit linking example:



Properties:

- **flexible**
- relies on track **extrapolation** (works also in inhomogeneous magnetic fields)
- **iterative** algorithm (**not parallelisable**)
- results depend on the **order** and **direction** (e.g. **inside-out** versus **outside-in** tracking)

The Kálmán fitter is the gold standard in particle physics, nowadays

Is there any parallelisable track fit?

The Master Equation

$$\chi^2 = \sum_{\text{layer } i} \frac{\Theta_{\text{MS},i}^2}{\sigma_{\theta,i}^2} + \frac{\Phi_{\text{MS},i}^2}{\sigma_{\phi,i}^2} + \sum_{\text{hits } jk} (x_j - \xi_j) V_{jk}^{-1} (x_k - \xi_k)$$

Multiple scattering uncertainties dominate for:

- **low momentum** tracks (→ low energy physics)
- **high precision** trackers (→ instrumentation technology)

Highland formula (PDG):
(RMS of scattering angle)

$$\theta_0 = \theta_{\text{plane}}^{\text{rms}} = \frac{13.6 \text{ MeV}}{\beta c p} z \sqrt{\frac{x}{X_0}} \left[1 + 0.038 \ln\left(\frac{x z^2}{X_0 \beta^2}\right) \right]$$

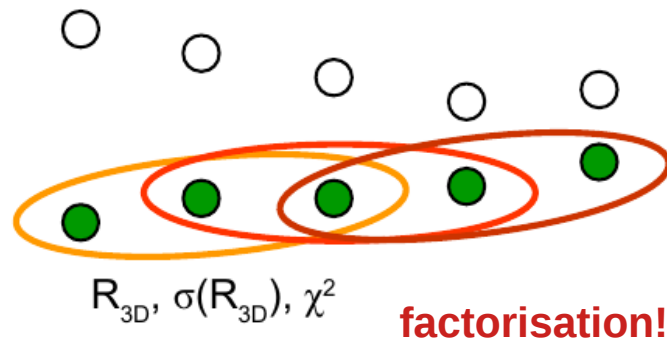
→ A multiple scattering fit can be **parallelised!**

The Multiple Scattering Fit

Literature:

- **A New Track Reconstruction Algorithm suitable for Parallel Processing based on Hit Triplets and Broken Lines**

[AS], EPJ Web Conf. 127 (2016) 00015

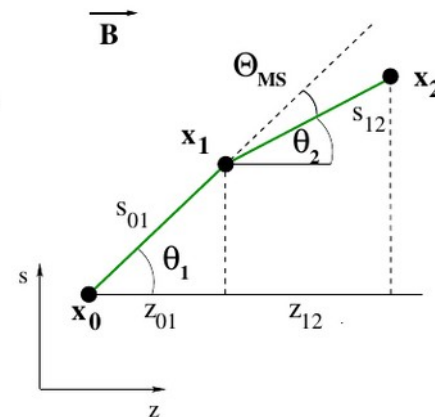
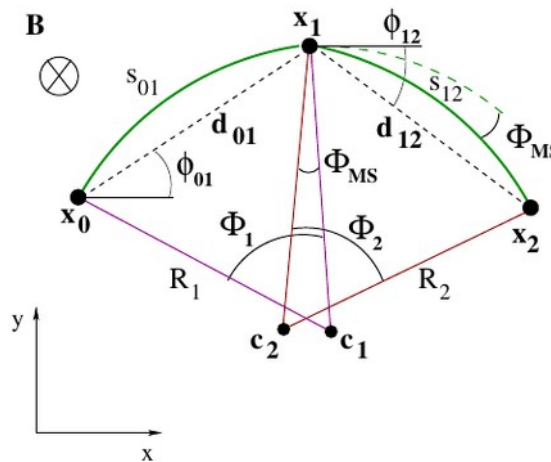


- **A New Three-Dimensional Track Fit with Multiple Scattering,**

N.Berger, M.Kiehn, A.Kozlinsky, [AS], NIMA 844C, 135 (2017)

fit of hit triplet:

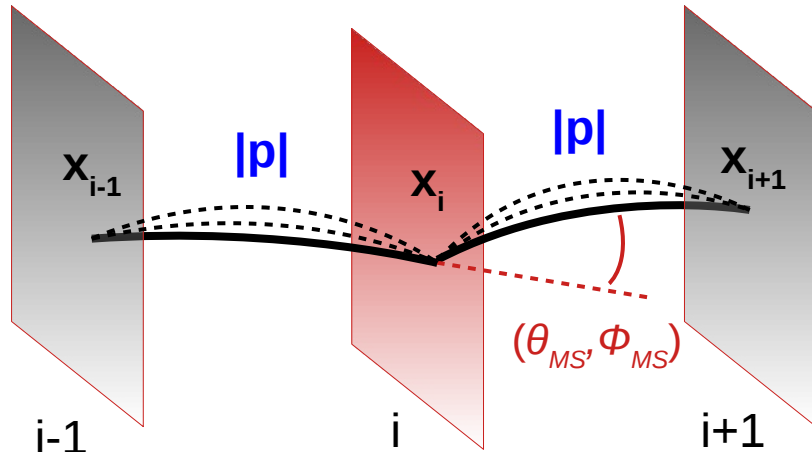
→ used by **Mu3e** and **Belle2** experiments



The Multiple Scattering Triplet Fit

Assumptions:

- All points \mathbf{x}_{i-1} , \mathbf{x}_i , \mathbf{x}_{i+1} (instrumentation layers) are given
- the modulus of the **momentum** \mathbf{p} of the particle is **conserved**
- the magnetic **field** \mathbf{B} is constant
- the **material** in **layer** i is known



(convenient to use cylinder coordinates)

- The **momentum** $|\mathbf{p}|$ is the only free (unknown) parameter of the particle
- The MS angles $(\theta_{MS}, \phi_{MS}) \rightarrow$ **minimised** depend on $|\mathbf{p}|$

$$\chi^2 = \frac{\Theta_{MS}^2}{\sigma_\theta^2} + \frac{\Phi_{MS}^2}{\sigma_\phi^2}$$

- All other parameters (e.g. direction) can be derived from triplet geometry if $|\mathbf{p}|$ is known!

B-Field: The Helix

Parameterisation (Cartesian coordinates):

$$\mathbf{r}(t) = R [\cos(2\pi t)\mathbf{e}_x + \sin(2\pi t)\mathbf{e}_y] + ht\mathbf{e}_z$$

In transverse plane (2D) of a magnetic field

$$R = \frac{p_{\perp}}{qB}$$

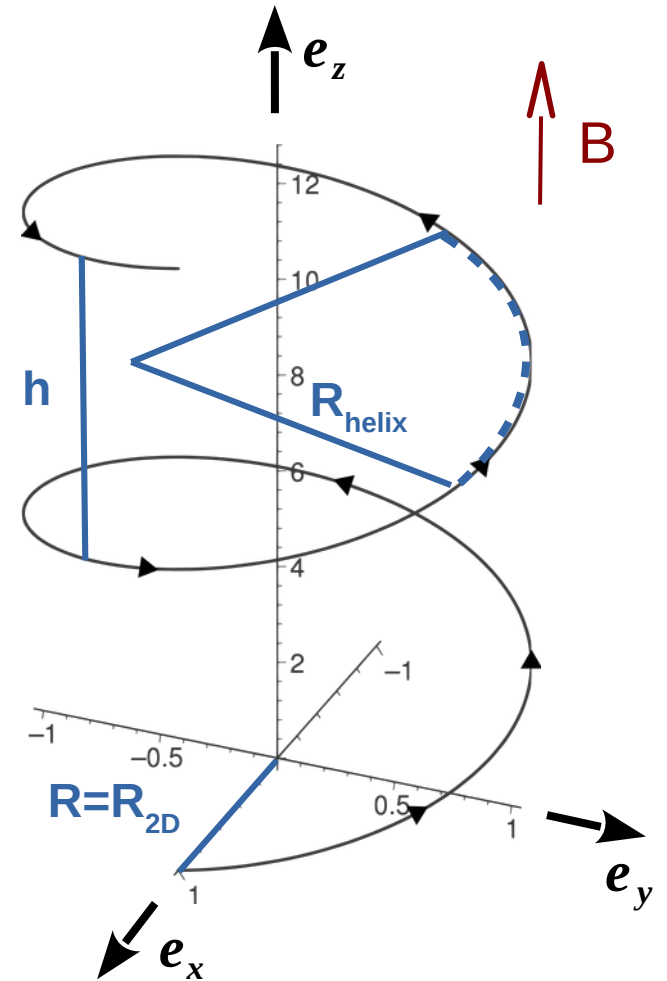
Define:

$$R_{3D} = \frac{R}{\sin(\theta)} = \frac{p}{qB}$$

invariant for MS!

Relation:

$$R_{3D}^2 = R_{2D} R_{\text{helix}} \quad (\text{geometric average})$$



Solution of MS Triplet Fit

Calculate:

$$\Phi_{MS} = \Phi_{MS}(R_{3D})$$

$$\Theta_{MS} = \Theta_{MS}(R_{3D})$$

Solution given by

$$\sin^2 \frac{\Phi_1}{2} = \frac{d_{01}^2}{4R_{3D}^2} + \frac{z_{01}^2 \sin^2(\Phi_1/2)}{R_{3D}^2 \Phi_1^2}$$

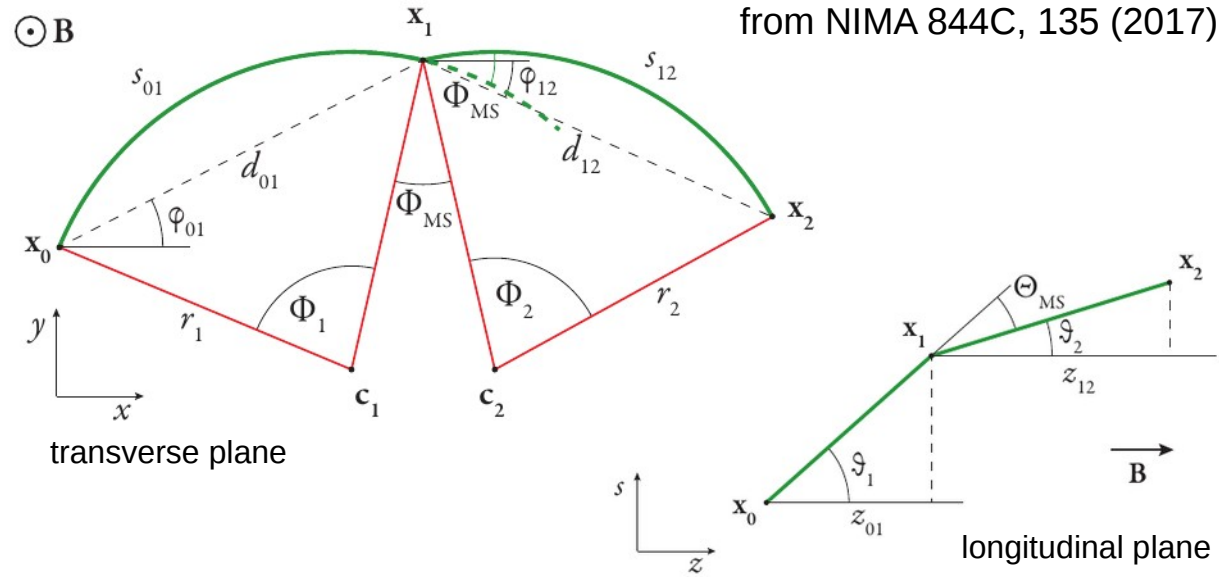
$$\sin^2 \frac{\Phi_2}{2} = \frac{d_{12}^2}{4R_{3D}^2} + \frac{z_{12}^2 \sin^2(\Phi_2/2)}{R_{3D}^2 \Phi_2^2}$$

and

$$\sin \theta_1 = \frac{d_{01}}{2R_{3D}} \operatorname{cosec} \left(\frac{z_{01}}{2R_{3D} \cos \theta_1} \right)$$

$$\sin \theta_2 = \frac{d_{12}}{2R_{3D}} \operatorname{cosec} \left(\frac{z_{12}}{2R_{3D} \cos \theta_2} \right)$$

→ transcendent equations



Important geometric relations for solution:

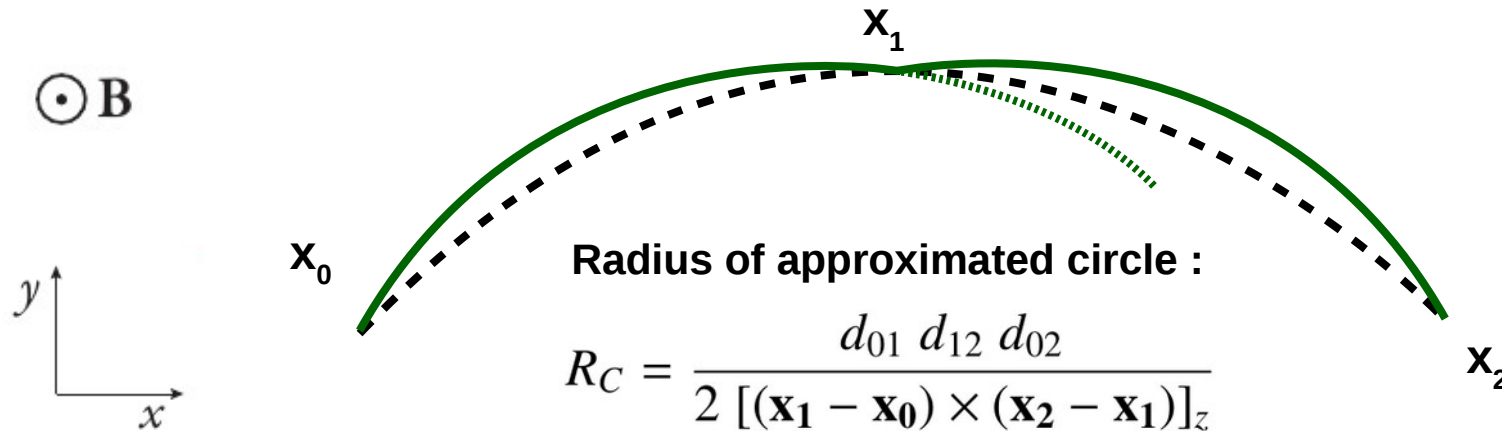
$$R_{3D}^2 = R_1^2 + \frac{z_{01}^2}{\Phi_1^2} = R_2^2 + \frac{z_{12}^2}{\Phi_2^2}$$

$$\Phi_{MS}(R_{3D}) = (\phi_{12} - \phi_{01}) - \frac{\Phi_1(R_{3D}) + \Phi_2(R_{3D})}{2}$$

Linearisation of MS Triplet Fit

- Minimisation of the χ^2 -function requires the **derivative** of **transcendent** equations (\rightarrow **OK**)
- But the derivatives are again transcendent equations; **no algebraic solution** (\rightarrow **NOK**)
- However, the functions are **analytical** \rightarrow **linearisation ansatz**

Trick: assume that the scattering **angles are small!** (\rightarrow good assumption)



transverse plane

\rightarrow **treat multiple scattering as small perturbation!**

Single MS Triplet Fit

3D Radius (momentum):

$$R_{3D}^{min} = -\frac{\eta \tilde{\Phi} \sin^2 \vartheta + \beta \tilde{\Theta}}{\eta^2 \sin^2 \vartheta + \beta^2}$$

independent
of MS uncertainty!

Fit quality:

$$\chi_{min}^2 = \frac{1}{\sigma_{MS}^2} \frac{(\beta \tilde{\Phi} - \eta \tilde{\Theta})^2}{\eta^2 + \beta^2 / \sin^2 \vartheta}$$

3D Radius uncertainty:

$$\sigma(R_{3D}) = \sigma_{MS} \sqrt{\frac{1}{\eta^2 \sin^2 \vartheta + \beta^2}}$$

Note that σ_{MS} is calculated from MS-formula using above momentum result

Geometry parameters are based on circle solution:

$$\tilde{\Phi} = -\frac{1}{2}(\Phi_{1C}\alpha_1 + \Phi_{2C}\alpha_2),$$

$$\eta = \frac{d\Phi_{MS}}{dR_{3D}} = \frac{\Phi_{1C} \alpha_1}{2R_{3D,1C}} + \frac{\Phi_{2C} \alpha_2}{2R_{3D,2C}}$$

$$\tilde{\Theta} = \vartheta_{2C} - \vartheta_{1C} - \left((1 - \alpha_2) \cot \vartheta_{2C} - (1 - \alpha_1) \cot \vartheta_{1C} \right)$$

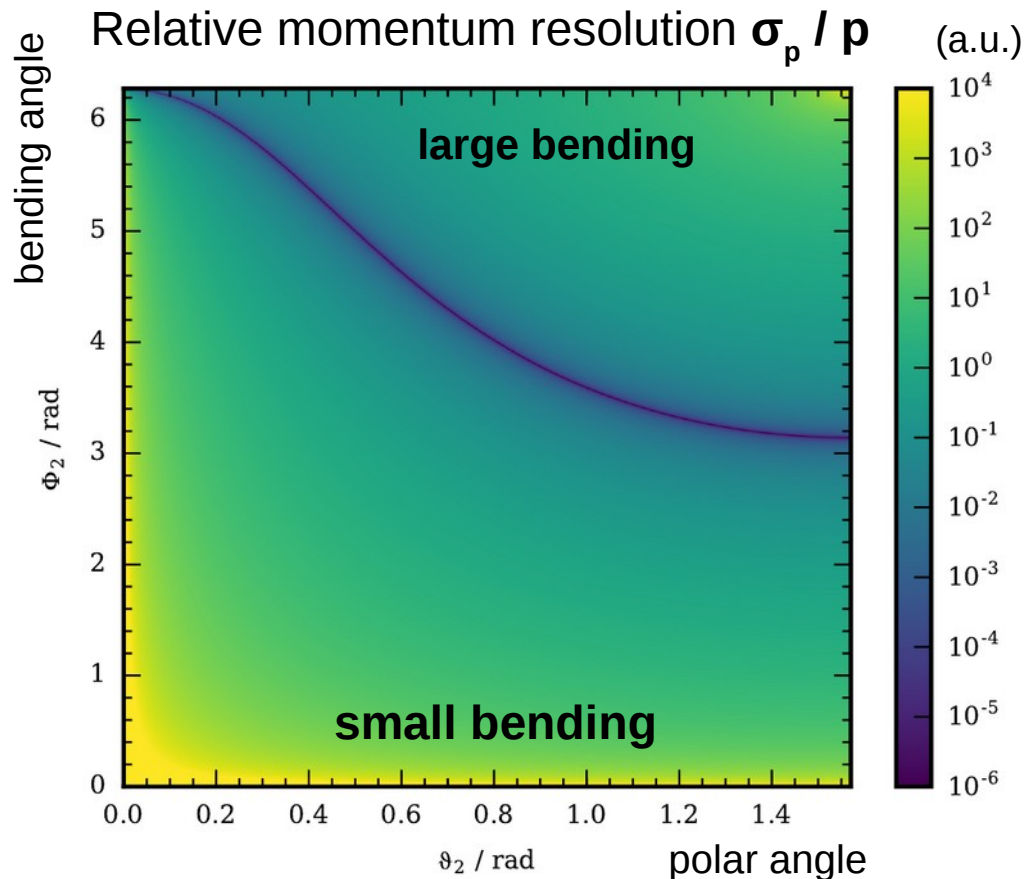
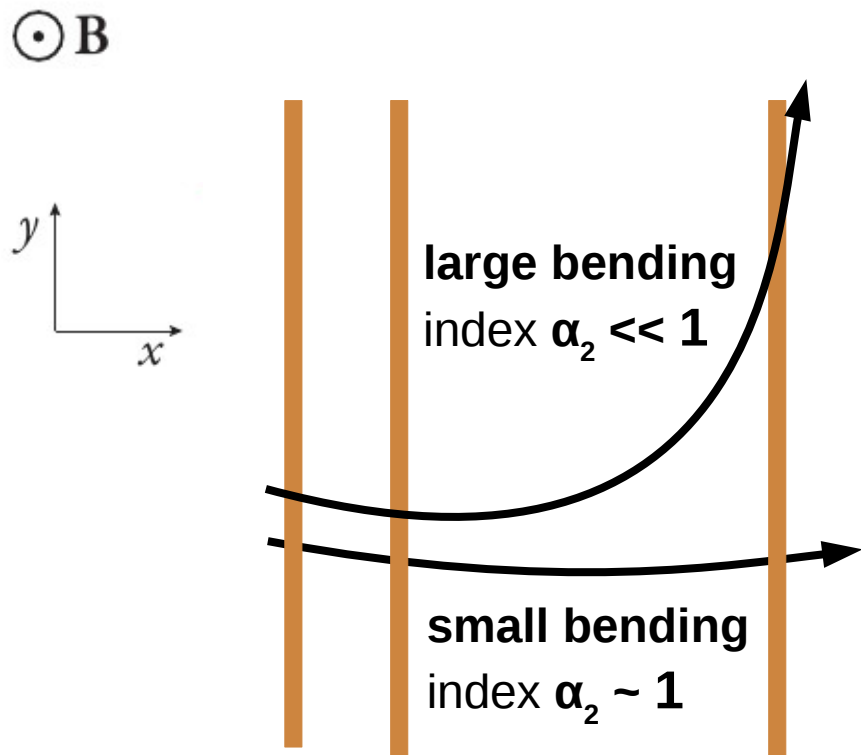
$$\beta = \frac{d\Theta_{MS}}{dR_{3D}} = \frac{(1 - \alpha_2) \cot \vartheta_{2C}}{R_{3D,2C}} - \frac{(1 - \alpha_1) \cot \vartheta_{1C}}{R_{3D,1C}}.$$

with index
parameters:

$$\alpha_1 = \frac{R_C^2 \Phi_{1C}^2 + z_{01}^2}{\frac{1}{2} R_C^2 \Phi_{1C}^3 \cot \frac{\Phi_{1C}}{2} + z_{01}^2}$$

$$\alpha_2 = \frac{R_C^2 \Phi_{2C}^2 + z_{12}^2}{\frac{1}{2} R_C^2 \Phi_{2C}^3 \cot \frac{\Phi_{2C}}{2} + z_{12}^2}$$

Example Spectrometer



Combination of Triplets

Each triplet fit provides:

$$R_{3D,i}, \sigma(R_{3D})_i, \chi_i^2$$

averaging:

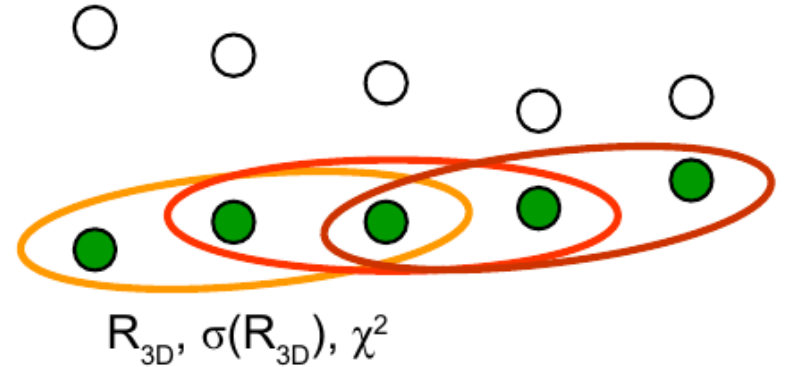
$$\overline{R_{3D}} = \frac{\sum_i^{n_{hit}-2} \frac{R_{3D,i}}{\sigma_i(R_{3D})^2}}{\sum_i^{n_{hit}-2} \frac{1}{\sigma_i(R_{3D})^2}}$$

combination:

$$\chi_{comb}^2 = \sum_{i=triplet} \chi_i^2 + \frac{(R_{3D,i} - \overline{R_{3D}})^2}{\sigma_i(R_{3D})^2}$$

Comments:

- every triplet is independent (hit positions are given)
- thus, all momentum measurements are independent!
- errors are uncorrelated



number of hits: N_{hit}

number of triplets: $N_{triplet} = N_{hit} - 2$

Remark: track building is simple:

- connecting triplets share two hits
- connecting triplets should have compatible momenta

→ **graph theory**

Question to students:

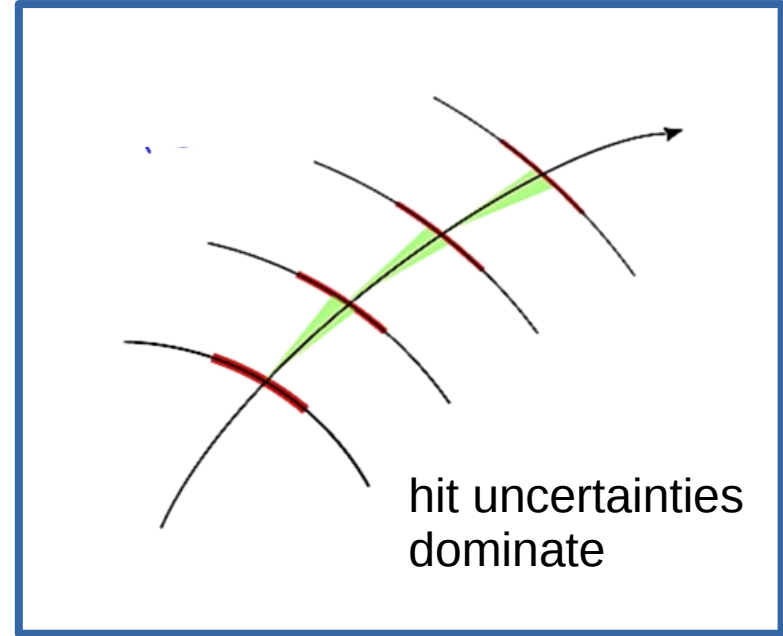
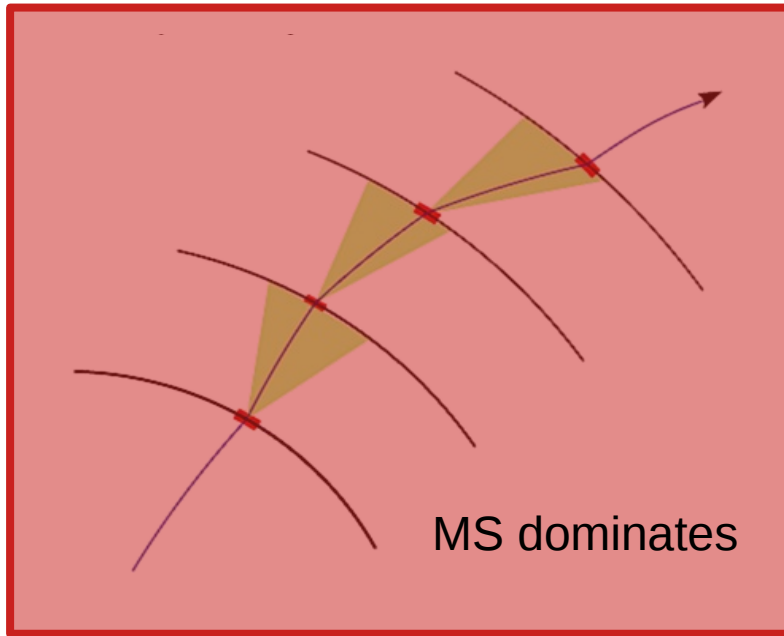
What are the advantages and achievements of the MS Triplet fit?

Chapter 3

Fitting Tracks with Hit Uncertainties

The Master Equation

$$\chi^2 = \sum_{\text{layer } i} \frac{\Theta_{\text{MS},i}^2}{\sigma_{\theta,i}^2} + \frac{\Phi_{\text{MS},i}^2}{\sigma_{\phi,i}^2} + \sum_{\text{hits } jk} (x_j - \xi_j) V_{jk}^{-1} (x_k - \xi_k)$$

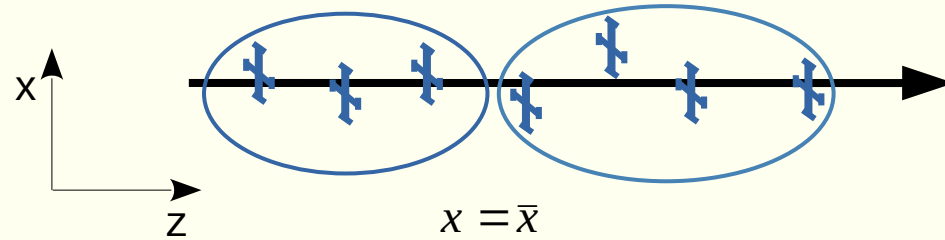


Fitting Tracks with Hit Uncertainties

(no multiple scattering)

Case A

- no B-field
- no slope
- **averaging**



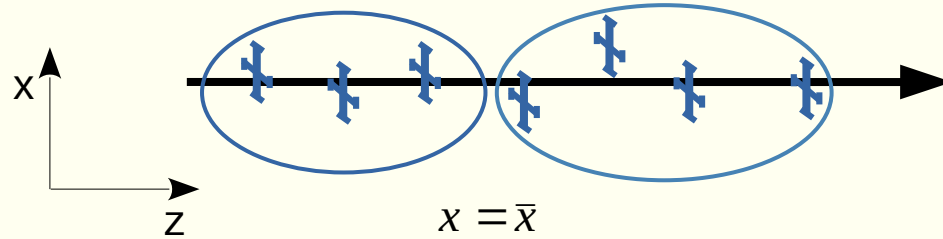
$$\bar{x} = \sum_{\text{hits } i} x_i w_i$$

Fitting Tracks with Hit Uncertainties

(no multiple scattering)

Case A

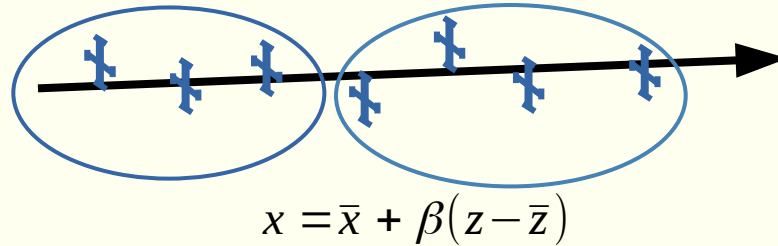
- no B-field
- no slope
- **averaging**



$$\bar{x} = \sum_{\text{hits } i} x_i w_i$$

Case B

- no B-field
- slope unknown
- **straight line fit**



$$\bar{x} = \sum_{\text{hits } i} x_i w_i$$

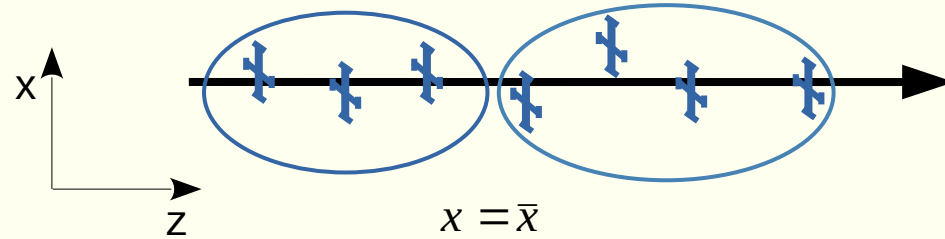
$$\beta = \frac{\text{Cov}(Z, X)}{\text{Var}(Z)}$$

Fitting Tracks with Hit Uncertainties

(no multiple scattering)

Case A

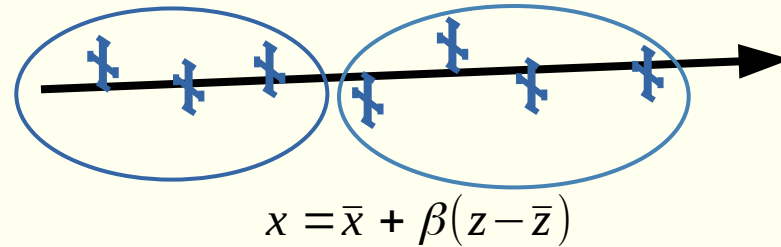
- no B-field
- no slope
- **averaging**



$$\bar{x} = \sum_{\text{hits } i} x_i w_i$$

Case B

- no B-field
- slope unknown
- **straight line fit**

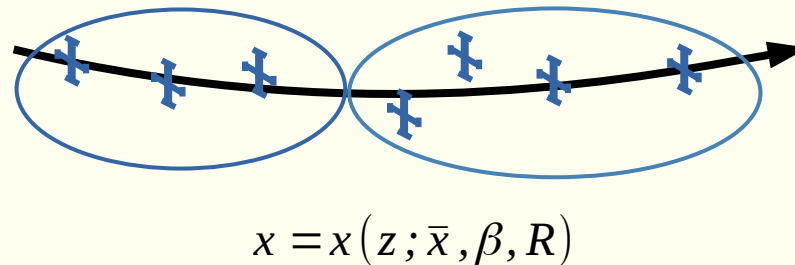


$$\bar{x} = \sum_{\text{hits } i} x_i w_i$$

$$\beta = \frac{\text{Cov}(Z, X)}{\text{Var}(Z)}$$

Case C

- B-field > 0
- slope unknown
- momentum unknown
- **helix fit**



$$\bar{x} = \sum_{\text{hits } i} x_i w_i$$

$$\beta = \dots$$

$$R = \dots$$

Question to students:

Which of the three cases
are parallelisable?

And if yes, how?

Fitting a Helix to Hits with Errors

Described by a non-linear equation:

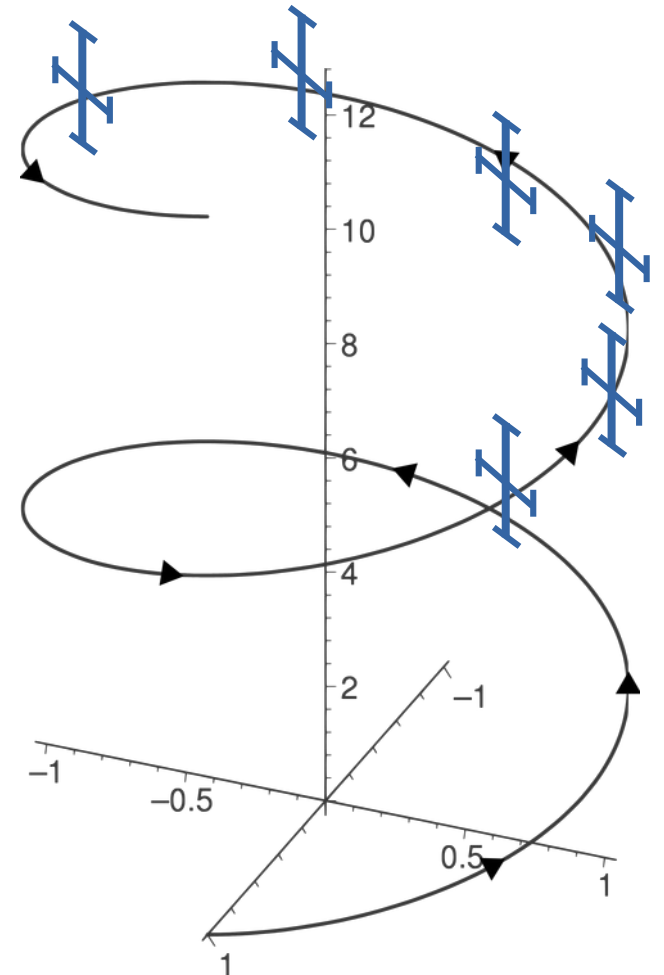
$$\mathbf{r}(t) = R [\cos(2\pi t)\mathbf{e}_x + \sin(2\pi t)\mathbf{e}_y] + ht\mathbf{e}_z$$

In general, difficult to solve:

- hit errors need to be **projected** on trajectory
- minimisation problem is **non-linear**

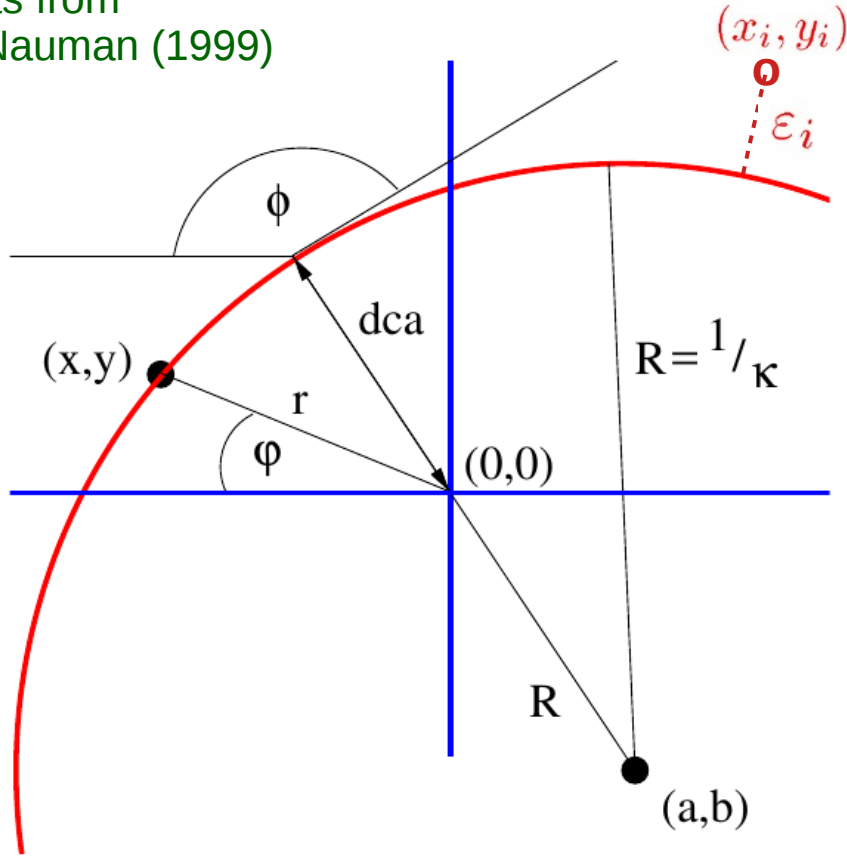
However, for the case of simple hit weights an **algebraic solution** exists for **circle** fit:

- **circle fit from Karimaki (1991)**



Karimaki Circle Fit

plots from
J. Nauman (1999)



New parameters defined:

dca = distance of closest approach to origin
(aka d_0)

Φ = initial angle at dca

$\kappa = 1/R$ = curvature of radius

Closest distance between hit and circle:

$$\epsilon_i = \pm \left[\sqrt{(x_i - a)^2 + (y_i - b)^2} - R \right]$$

Task: minimise χ^2 with respect to $\rightarrow \kappa, dca, \Phi$:

$$\chi^2 = \sum_i w_i \epsilon_i^2 \quad (w_i \text{ are weights})$$

Parameter Transformations

If hits are positioned close to circle:

$$\varepsilon_i \stackrel{|\varepsilon_i| \ll R}{\approx} \pm R^{-1} \left[(x_i - a)^2 + (y_i - b)^2 - R^2 \right]$$

$$\chi^2 = \sum_i w_i \varepsilon_i^2$$

Problems:

- parameters a , b , R can become very large (high momentum tracks) → numerical unstable
- uncertainties (for example on R) are not Gaussian distributed!

1. Switch to polar coordinates and use new parameters:

$$\varepsilon_i = \frac{1}{2} \kappa r_i^2 - (1 + \kappa d_{ca}) r_i \sin(\phi - \varphi_i) + \frac{1}{2} \kappa d_{ca}^2 + d_{ca}$$

2. simplify expression further by transforming χ^2 function:

$$\chi^2 = (1 + \kappa d_{ca}) \hat{\chi}^2$$

$$\varepsilon_i = (1 + \kappa d_{ca}) \eta_i$$

and minimisation of

$$\hat{\chi}^2 = \sum_i w_i \eta_i^2$$

result depends now weakly on position of origin!

Results

Parameters:

$$\kappa = \frac{2\beta}{\sqrt{1 - 4\delta\beta}}$$

$$d_{ca} = \frac{2\delta}{1 + \sqrt{1 - 4\delta\beta}}$$

$$\phi = \frac{1}{2} \arctan \left(\frac{2q_1}{q_2} \right)$$

Geometry parameters:

$$q_1 = C_{r^2 r^2} C_{xy} - C_{xr^2} C_{yr^2}$$

$$q_2 = C_{r^2 r^2} (C_{xx} - C_{yy}) - C_{xr^2}^2 + C_{yr^2}^2$$

$$\phi = 1/2 \arctan(2q_1/q_2)$$

$$\beta = (\sin\phi C_{xr^2} - \cos\phi C_{yr^2}) / C_{r^2 r^2}$$

$$\delta = -\beta \langle r^2 \rangle + \sin\phi \langle x \rangle - \cos\phi \langle y \rangle$$

C_{pq} are the covariance of samples p and q

Fit quality (only approximate):

$$\chi^2 = S_w (1 + \kappa d_{ca})^2 \cdot (\sin^2 \phi C_{xx} - 2\sin\phi \cos\phi C_{xy} + \cos^2 \phi C_{yy} - \kappa^2 C_{r^2 r^2})$$

- non-iterative track fit
- provides error matrix (not shown)
- complexity of calculation a bit higher than for MS fit (but different regime)

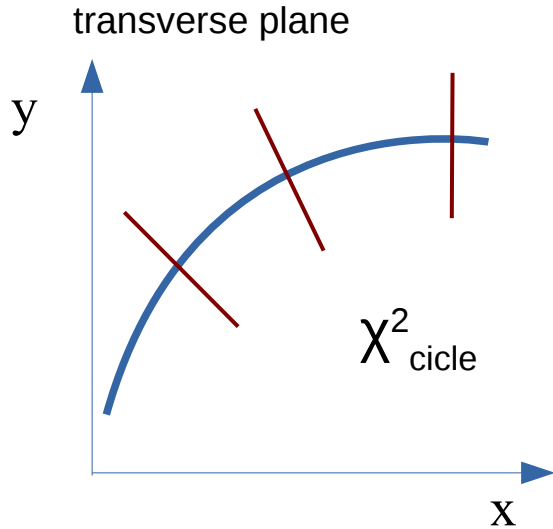
Question for students:

What are the advantages
of the circle fit?

What are the difficulties for
parallelisation?

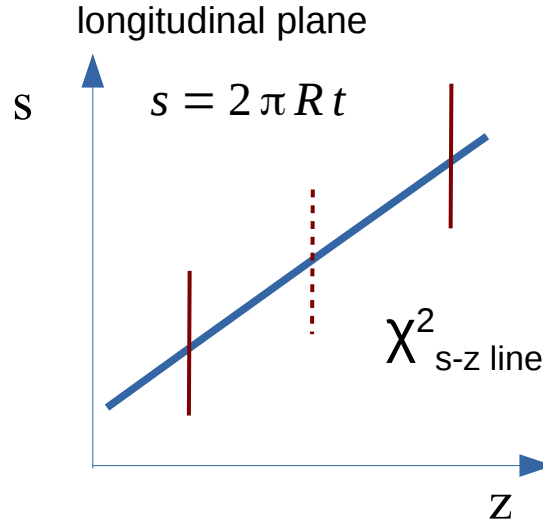
Helix Fit with Karimaki

“2.5D tracking”: fit transverse and longitudinal plane separately:



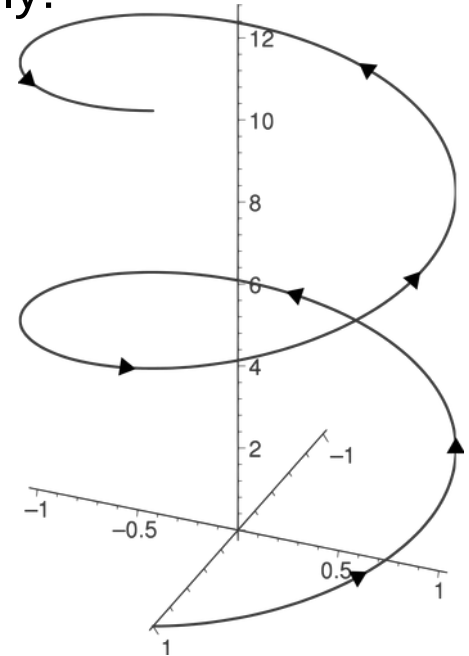
x-y fit provides:

- R = radius
- Φ = azimuth at dca
- dca = distance of closest approach



s-z fit provides:

- abscissa z_0
- $\tan \Theta = \Delta s / \Delta z$

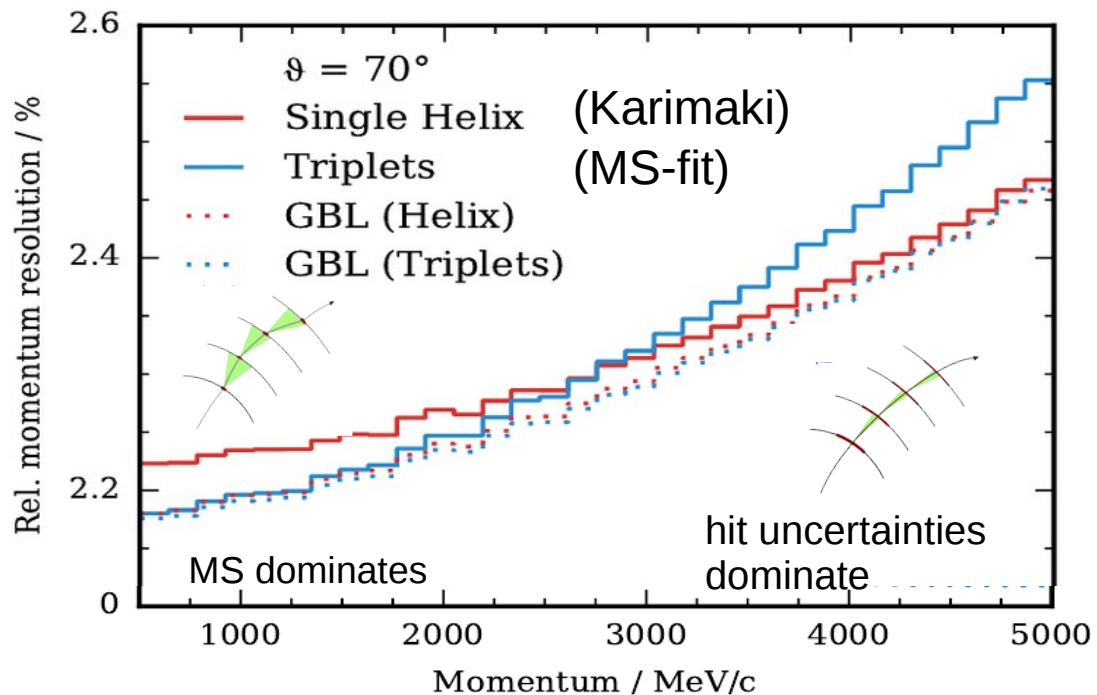
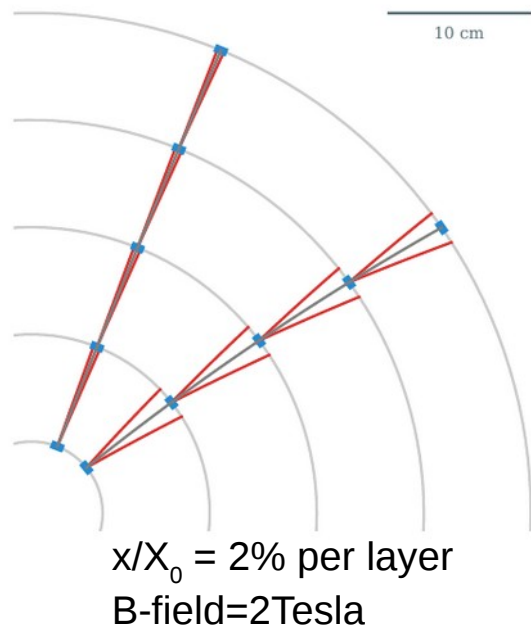


$$\mathbf{r}(t) = R [\cos(2\pi t)\mathbf{e}_x + \sin(2\pi t)\mathbf{e}_y] + ht\mathbf{e}_z$$

→ hit correlations between transverse and longitudinal plane are not considered!

Comparing Results

plot from NIMA 844C, 135 (2017)
(aka MS-fit)



GBL=General Broken Line

V. Blobel, NIMA, 566 (2006) 14.

- MS-fit is 2-5 times faster than Karimäki
- GBL is about $O(100)$ slower than the others

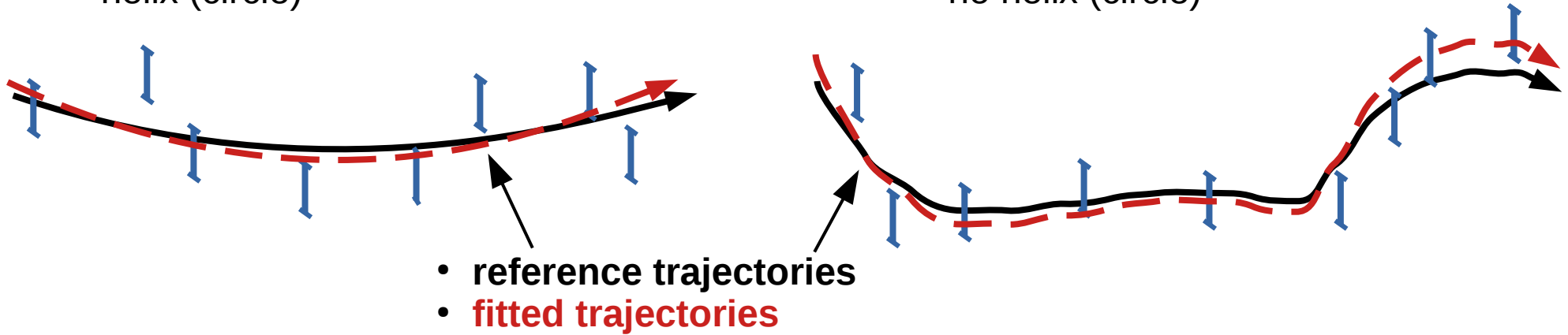
Chapter 4

Linearisation

Linearisation & Linear Fit

A) homogeneous magnetic field
→ helix (circle)

B) non-homogeneous magnetic field
→ no helix (circle)



If a reference trajectory close to the final resolution is given, the problem can be **linearised** by treating the hit displacements as **small corrections**

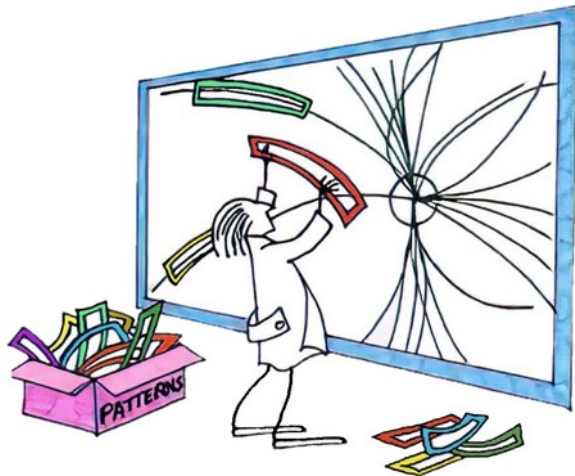
1. calculate hit positions and pulls with respect to reference trajectory
2. update **position \bar{x}** , **slope β** , **curvature radius R** (momentum)
3. can be **repeated** (iterated) for high **precision**

Applications

Linearised track fit is a good approach if

- the **track parameters** are **roughly known** by a previous reconstruction step (e.g. pattern match, other track finding techniques)
- Tracks are known to be roughly **straight lines** (no B-field, high momentum tracks)

Example: **ATLAS FTK & HTT** track trigger projects (similar project in CMS)



- The different **roads** describe/contain bundles of **similar trajectories**
- The roads provide an **initial guess** of the **track parameters**

Linearisation of a Circle/Helix Fit

Given a list of p track parameters
(e.g. R , Φ , dca , θ , z_0):

$$p_i^{\text{true}}, i = 1, \dots, p$$

and N hit displacements with respect to
reference orbit:

$$\delta x_j = x_j - \bar{x}_j$$

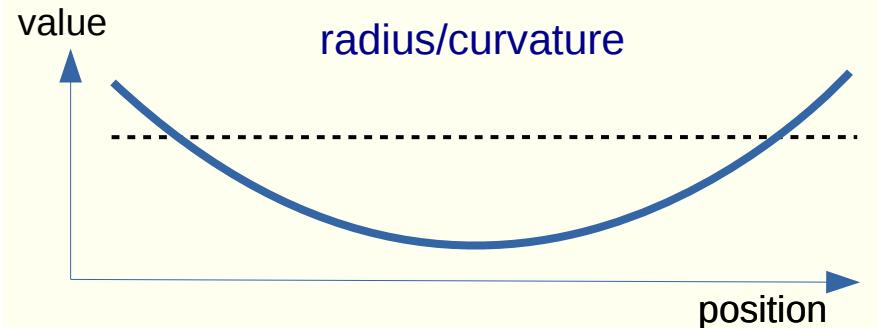
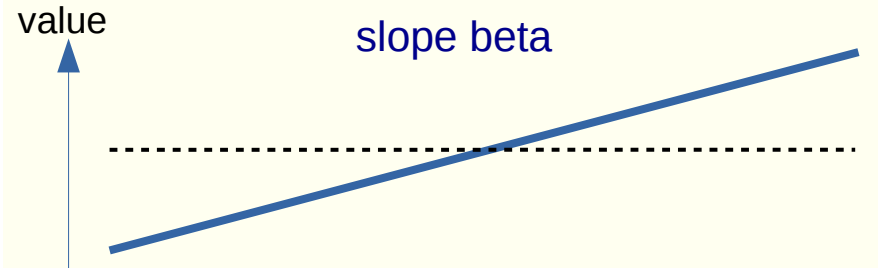
Then the track is linearised using:

$$p_i = \sum_{j=1}^N A_{ij} \delta x_j + \bar{p}_i$$

with coefficients \mathbf{A}_{ij} (matrix of $N \times p$ coefficients)

Example for coefficients (weights):

hit positions:



Linearisation of Fit Quality

For track finding (good/bad) or a track trigger the fit quality is crucial!

The calculation of the chi2 function can also be linearised using a principal component analysis:

$$\chi_i = \sum_{j=1}^N B_{ij} \delta x_j$$

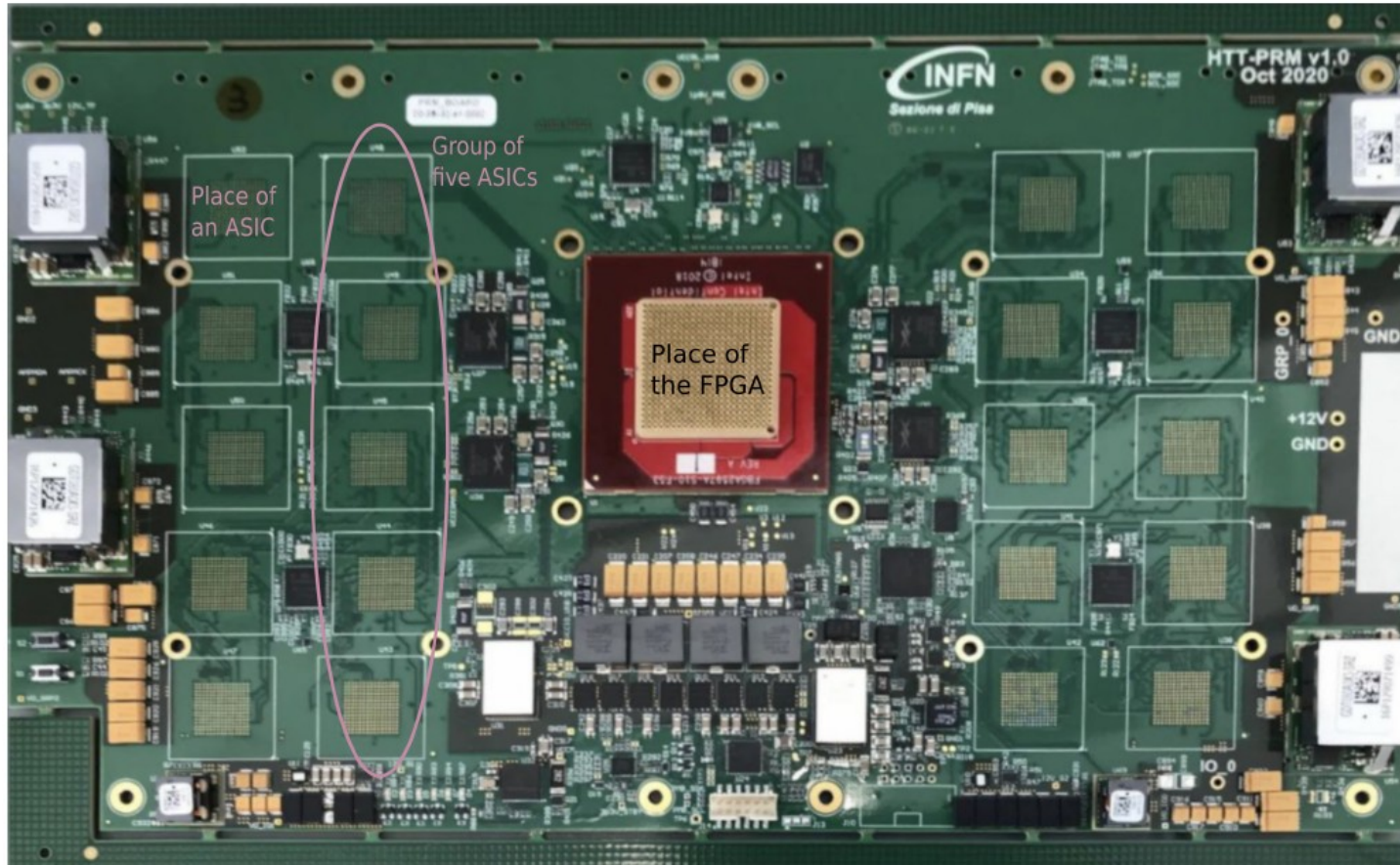
$$\chi^2 = \sum_{i=1}^{N-p} \chi_i^2$$

The coefficients B_{ij} can be represented by a **N x (N-p)** matrix.

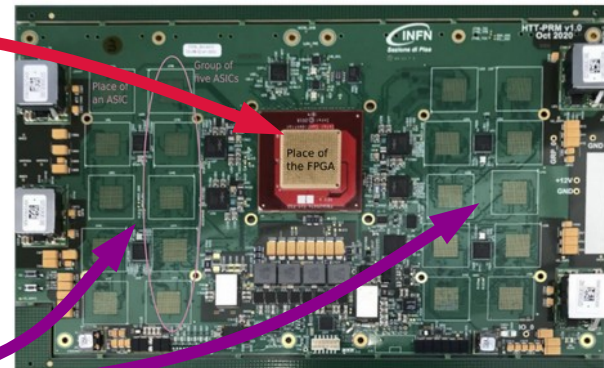
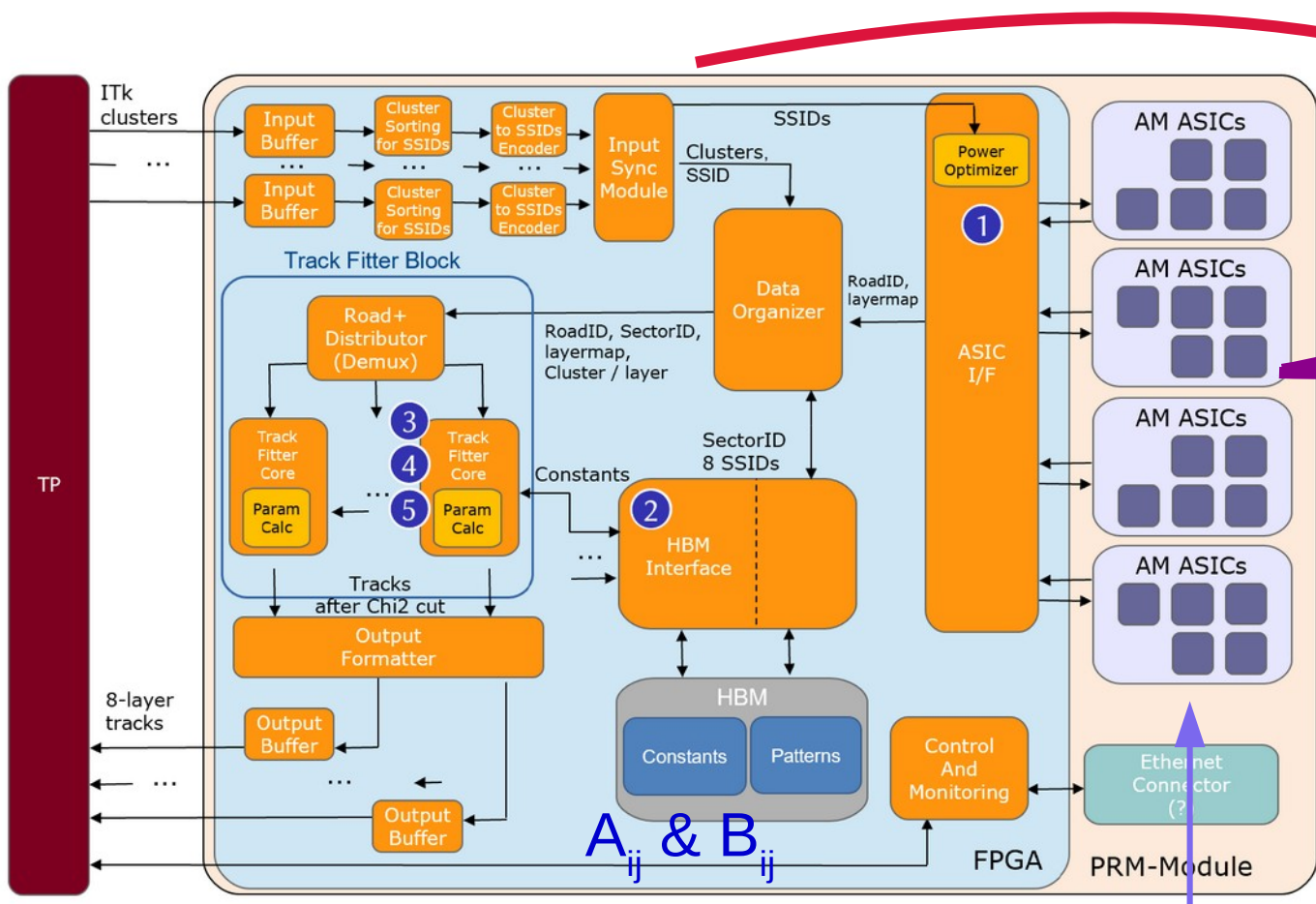
- Example: $p=5$ parameters, $N=12$ hits \rightarrow 84 parameters
- not all coefficients are significant
- clever choice of parameters can reduce the complexity (\rightarrow extra slide)

Example: ATLAS HTT

Pattern Recognition Mezzanine



Example: ATLAS HTT



- ① pattern match → ref. track
- ② request constants
- ③ linearised χ^2 calculation
- ④ apply χ^2 cut
- ⑤ linearised parameter fit

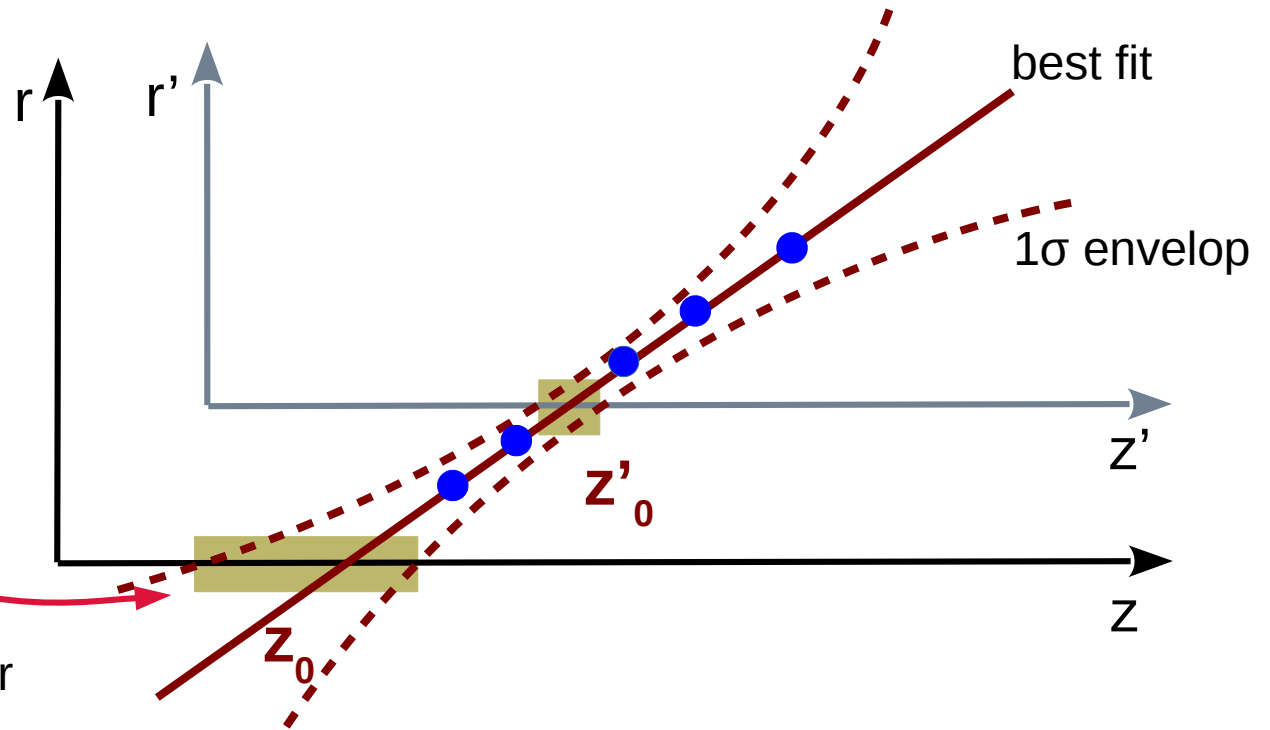
highly parallel

Question to students:

What do you consider is most challenging and technologically ambitious in this design?

Best Fitting Parameters?

- The fitted value and its uncertainty and also the correlations depend on the choice of the **coordinate system!**
- a wrong coordinate system choice can lead to large **non-linearities**
- it is also possible to **redefine parameters** which behave better in the fit



e.g.
$$z' = z - \cot(\theta)(R - R') - \frac{\cot(\theta)R^3}{6(2\rho)^2}$$

from [arXiv:1809.01467]

Chapter 5

A New Hit Uncertainty Track Fit

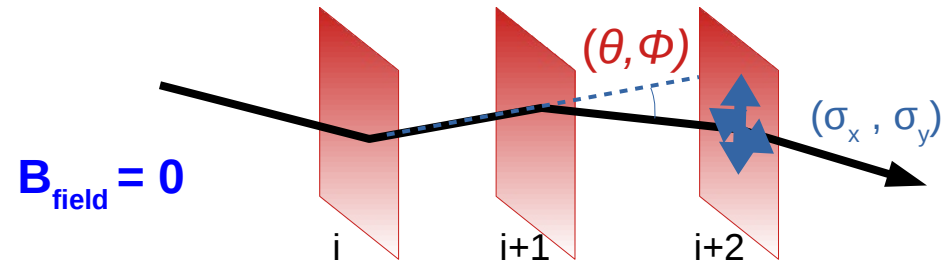
Recap

Full track fit:

$$\chi^2 = \sum_{\text{layer } i} \frac{\Theta_{\text{MS},i}^2}{\sigma_{\theta,i}^2} + \frac{\Phi_{\text{MS},i}^2}{\sigma_{\phi,i}^2} + \sum_{\text{hits } jk} (x_j - \xi_j) V_{jk}^{-1} (x_k - \xi_k)$$

includes **multiple scattering** and **hit uncertainties**

- **MS fit** alone can be **parallelised**
- This parallelisation is based in **hit triplets**
- Hit uncertainty fit can be **linearised**
→ good for parallelisation
- But linearisation needs a **reference** trajectory



Question to students:

What is the next logical step?

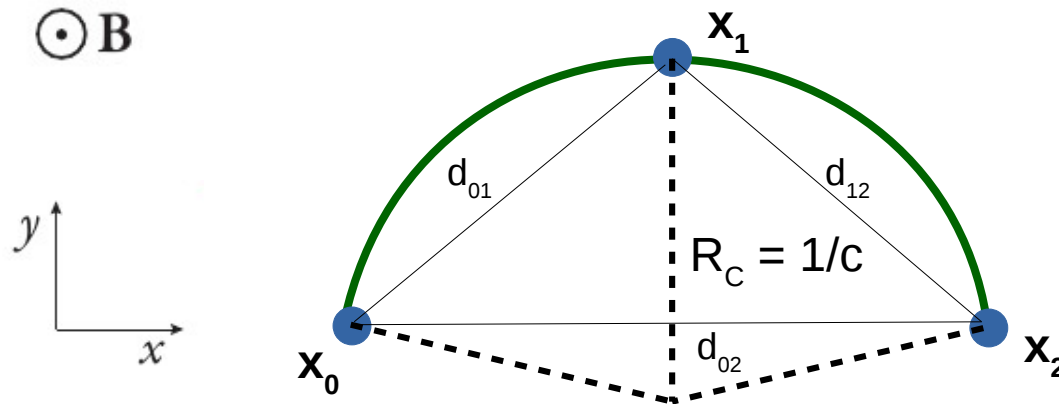
Hit Triplets

- A hit **triplet** is the smallest tracking **element** which contains all track parameter information
- The **precision** of the triplet track parameters depend on the **lever arm** (size)
- For track finding, triplets are often used as **seeds** (combinatorics is small):

seed finding = triplet finding!

Easy reconstruction in homogeneous magnetic field

- **three points** can always be connected by a **circle**
- all track parameters can be calculated → **reference track**



Radius of circle:

$$R_c = \frac{d_{01} d_{12} d_{02}}{2 [(\mathbf{x}_1 - \mathbf{x}_0) \times (\mathbf{x}_2 - \mathbf{x}_1)]_z}$$

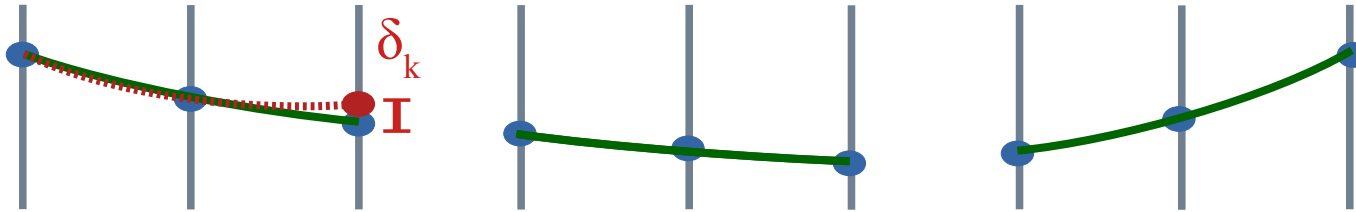
$$c = \frac{2 \sin(\phi_{12} - \phi_{01})}{d_{02}}$$

(and permutations)

Track Fit with N Hits

uncertainties are not shown

Possible **configuration** where all hits lie on their own reference trajectory



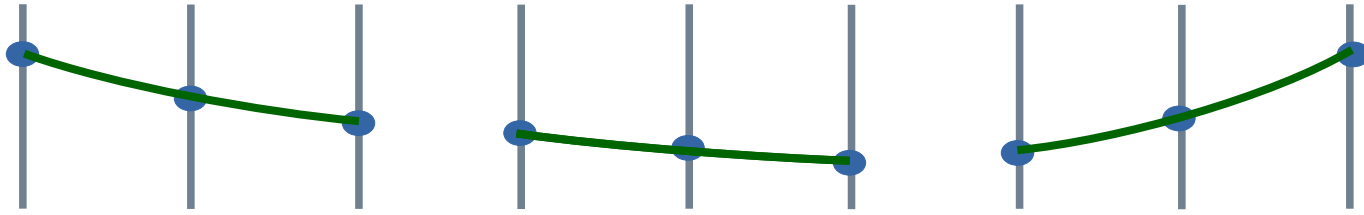
We can calculate how the track parameters change if we displace one point

$$\bar{x}_j, \beta_j, R_j \rightarrow \bar{x}(\delta_k)_j, \beta(\delta_k)_j, R(\delta_k)_j \quad (j=1, \dots, N_{\text{triplet}})$$

Now we can also calculate a **weight** or fit quality for small displacements δ_k

$$\chi_j^2 = \sum_{k=0}^2 \frac{\delta_k^2}{\sigma_k^2} \quad \leftarrow \text{hit pull}$$

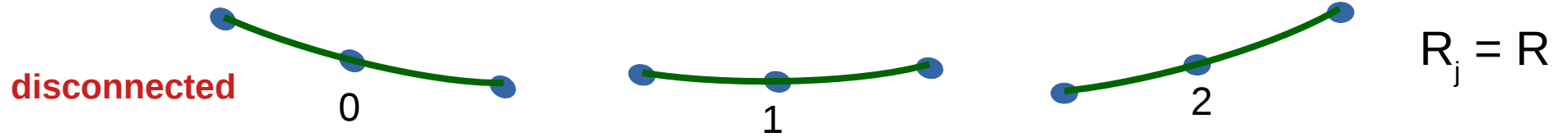
Idea of common fit \rightarrow combine all hit triplets with the constraint: $R_j = R$



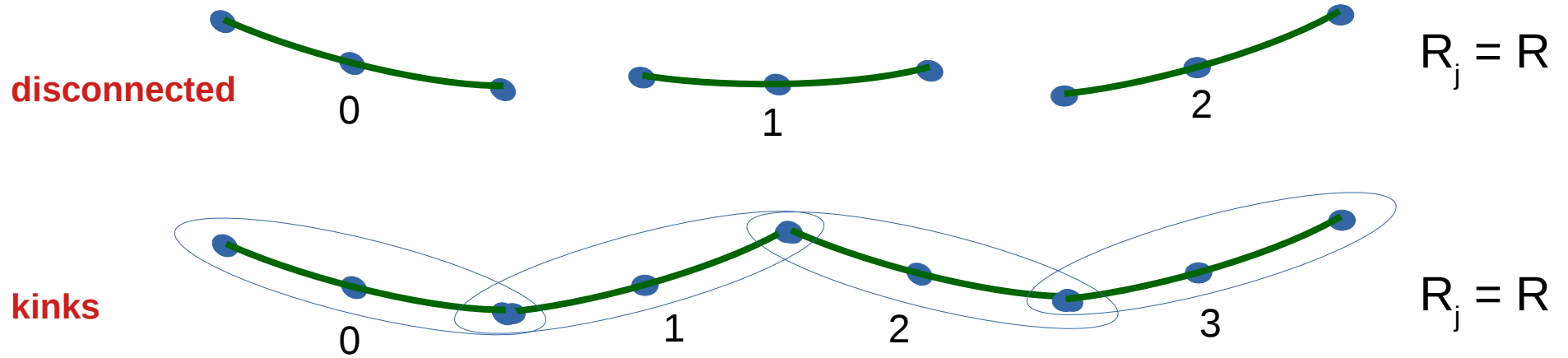
Question to students:

Will such a fit work?

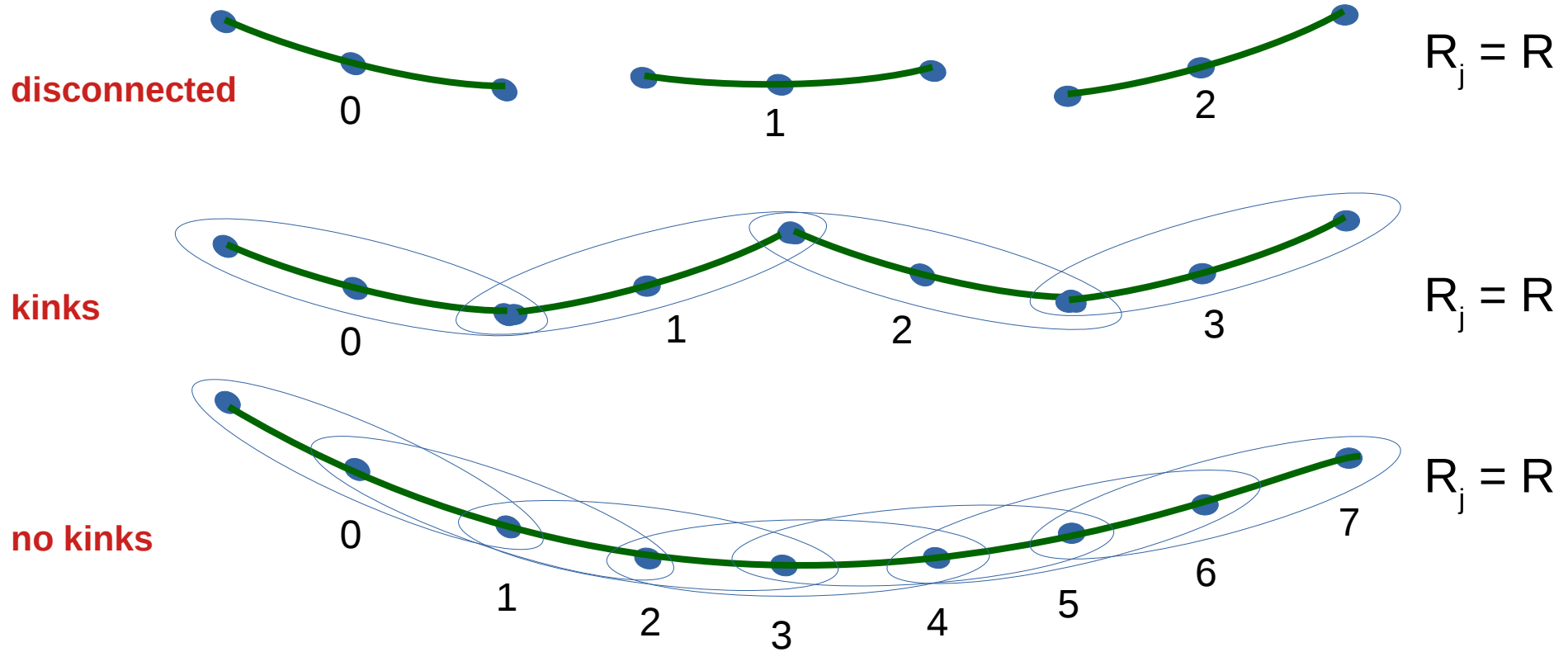
Correlations and triplet topologies



Correlations and triplet topologies



Correlations and triplet topologies



correlations are automatically taken into account if **consecutive** triplets are considered

Question to students:

How to make a constraint fit?

Method of Lagrange Multipliers

Lagrange function

$$\mathcal{L}(x, \lambda) = f(x) - \lambda g(x)$$

to be minimised

constraint $g(x) \rightarrow 0$

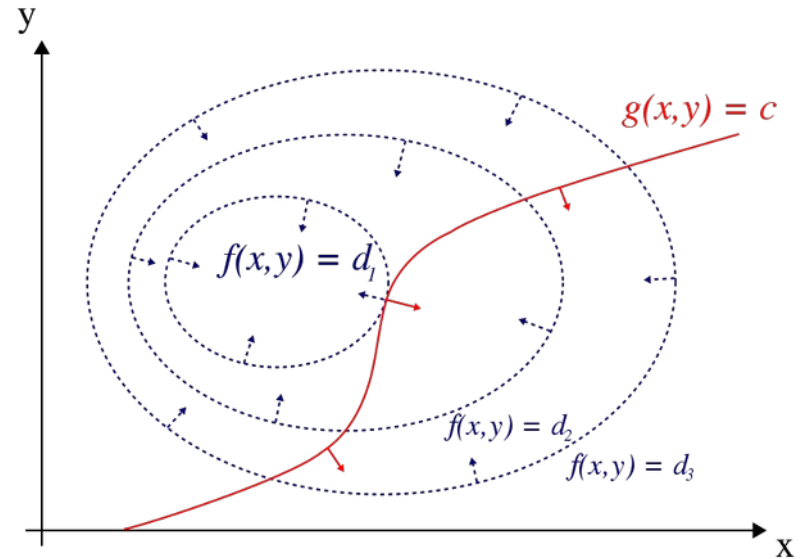
Minimisation of this Lagrangian results in:

$$Df(x^*) = \lambda^{*T} Dg(x^*)$$

with the partial derivatives: $D := \frac{\partial}{\partial x_k}$

The minimisation results in a **system of equations** yielding the new fitted hit positions x^*

Since the displacements x^* are small, the system can be linearised and solved!



The Lagrange parameters λ_k is the rate of change of the quantity being optimized as a function of the constraint parameter (from Wikipedia)

In other words, λ_k described how well the radius/curvature is measured!

Lagrangian for Hit Uncertainty Fit

$$\mathcal{L}(c; \delta_k, \lambda_j) = \sum_{\text{hit } k}^{n_{\text{hit}}} \frac{\delta_k^2}{\sigma_k^2} + \sum_{\text{triplet } j}^{n_{\text{hit}}-2} \lambda_j [c_j(\delta_j, \delta_{j+1}, \delta_{j+2}) - c]$$

hit pulls
L.M.
reference trajectory
to be fitted curvature

note c=1/R

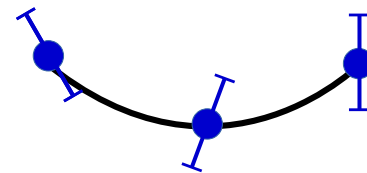
here use c=1/R (curvature),
which is numerically more stable

Linearisation:

$$c_j = \tilde{c}_j + \sum_{k=0}^3 \frac{dc_j}{d\delta_{j+k}} \delta_{j+k}$$

$$= \tilde{c}_j + \Xi_j \left(\frac{(\tau_0)_j}{(d_{01})_j} \delta_j + \frac{(\tau_3)_j}{(d_{12})_j} \delta_{j+2} - \left(\frac{(\tau_1)_j}{(d_{01})_j} + \frac{(\tau_2)_j}{(d_{12})_j} \right) \delta_{j+1} \right)$$

projection parameters



with: $\Xi_j = \sqrt{\frac{4}{(d_{02})_j^2} - \tilde{c}_j^2}$

Lagrangian for Hit Uncertainty Fit

Minimisation $\frac{\partial \mathcal{L}}{\partial \delta_k} = 0$ (hits) and $\frac{\partial \mathcal{L}}{\partial \lambda_j} = 0$ (constraints) yields:

$$\begin{pmatrix}
 2/\sigma_0^2 & 0 & 0 & 0 & 0 & (\xi_0)_0 & 0 & 0 \\
 0 & 2/\sigma_1^2 & 0 & 0 & 0 & -(\xi_{12})_0 & (\xi_0)_1 & 0 \\
 0 & 0 & 2/\sigma_2^2 & 0 & 0 & (\xi_3)_0 & -(\xi_{12})_1 & (\xi_0)_2 \\
 0 & 0 & 0 & 2/\sigma_3^2 & 0 & 0 & (\xi_3)_1 & -(\xi_{12})_2 \\
 0 & 0 & 0 & 0 & 2/\sigma_4^2 & 0 & 0 & (\xi_3)_2 \\
 (\xi_0)_0 & -(\xi_{12})_0 & (\xi_3)_0 & 0 & 0 & 0 & 0 & 0 \\
 0 & (\xi_0)_1 & -(\xi_{12})_1 & (\xi_3)_1 & 0 & 0 & 0 & 0 \\
 0 & 0 & (\xi_0)_2 & -(\xi_{12})_2 & (\xi_3)_2 & 0 & 0 & 0
 \end{pmatrix}
 \begin{pmatrix}
 \delta_0 \\
 \delta_1 \\
 \delta_2 \\
 \delta_3 \\
 \delta_4 \\
 \lambda_0 \\
 \lambda_1 \\
 \lambda_2
 \end{pmatrix}
 =
 \begin{pmatrix}
 0 \\
 0 \\
 0 \\
 0 \\
 0 \\
 c - \tilde{c}_0 \\
 c - \tilde{c}_1 \\
 c - \tilde{c}_2
 \end{pmatrix}$$

$\left. \begin{matrix} \text{---} \\ \text{---} \\ \text{---} \\ \text{---} \\ \text{---} \end{matrix} \right\} N_{\text{hits}}$
 $\left. \begin{matrix} \text{---} \\ \text{---} \\ \text{---} \end{matrix} \right\} N_{\text{triplets}}$

with

$$\begin{aligned}
 (\xi_0)_j &= \frac{\Xi_j(\tau_0)_j}{(d_{01})_j} \\
 (\xi_{12})_j &= \frac{\Xi_j(\tau_1)_j}{(d_{01})_j} + \frac{\Xi_j(\tau_2)_j}{(d_{12})_j} \\
 (\xi_3)_j &= \frac{\Xi_j(\tau_3)_j}{(d_{12})_j}
 \end{aligned}$$

hits and Lagrange multipliers

fitted curvature

Note the symmetry of the matrix!
How to solve the system of equations?

Lagrangian for Hit Uncertainty Fit

Minimisation $\frac{\partial \mathcal{L}}{\partial \delta_k} = 0$ (hits) and $\frac{\partial \mathcal{L}}{\partial \lambda_j} = 0$ (constraints) yields:

$$\begin{pmatrix}
 2/\sigma_0^2 & 0 & 0 & 0 & 0 & (\xi_0)_0 & 0 & 0 \\
 0 & 2/\sigma_1^2 & 0 & 0 & 0 & -(\xi_{12})_0 & (\xi_0)_1 & 0 \\
 0 & 0 & 2/\sigma_2^2 & 0 & 0 & (\xi_3)_0 & -(\xi_{12})_1 & (\xi_0)_2 \\
 0 & 0 & 0 & 2/\sigma_3^2 & 0 & 0 & (\xi_3)_1 & -(\xi_{12})_2 \\
 0 & 0 & 0 & 0 & 2/\sigma_4^2 & 0 & 0 & (\xi_3)_2 \\
 \hline
 (\xi_0)_0 & -(\xi_{12})_0 & (\xi_3)_0 & 0 & 0 & 0 & 0 & 0 \\
 0 & (\xi_0)_1 & -(\xi_{12})_1 & (\xi_3)_1 & 0 & 0 & 0 & 0 \\
 0 & 0 & (\xi_0)_2 & -(\xi_{12})_2 & (\xi_3)_2 & 0 & 0 & 0
 \end{pmatrix}
 \begin{pmatrix}
 \delta_0 \\
 \delta_1 \\
 \delta_2 \\
 \delta_3 \\
 \delta_4 \\
 \lambda_0 \\
 \lambda_1 \\
 \lambda_2
 \end{pmatrix}
 =
 \begin{pmatrix}
 0 \\
 0 \\
 0 \\
 0 \\
 0 \\
 c - \tilde{c}_0 \\
 c - \tilde{c}_1 \\
 c - \tilde{c}_2
 \end{pmatrix}$$

$\left. \begin{matrix} \text{---} \\ \text{---} \\ \text{---} \\ \text{---} \\ \text{---} \end{matrix} \right\} N_{\text{hits}}$
 $\left. \begin{matrix} \text{---} \\ \text{---} \\ \text{---} \end{matrix} \right\} N_{\text{triplets}}$

with

$$\begin{aligned}
 (\xi_0)_j &= \frac{\Xi_j(\tau_0)_j}{(d_{01})_j} \\
 (\xi_{12})_j &= \frac{\Xi_j(\tau_1)_j}{(d_{01})_j} + \frac{\Xi_j(\tau_2)_j}{(d_{12})_j} \\
 (\xi_3)_j &= \frac{\Xi_j(\tau_3)_j}{(d_{12})_j}
 \end{aligned}$$

hits and Lagrange
multipliers

fitted curvature

Note the symmetry of the matrix!

How to solve the system of equations?

Structure of Equations

system of equations

$$\mathbf{M} \cdot \begin{pmatrix} \vec{\delta} \\ \vec{\lambda} \end{pmatrix} = \begin{pmatrix} \vec{0} \\ c\vec{1} - \vec{\hat{c}} \end{pmatrix}$$

$$\vec{\xi}_j = (\xi_0, \xi_{12}, \xi_3)_j^T$$

$$\vec{\delta} = (\delta_0, \delta_1, \dots, \delta_{n_{\text{hit}}-1})^T$$

$$\vec{\lambda} = (\lambda_0, \lambda_1, \dots, \lambda_{n_{\text{triplet}}-1})^T$$

$$\vec{\hat{c}} = (\hat{c}_0, \hat{c}_1, \dots, \hat{c}_{n_{\text{triplet}}-1})^T$$

with

$$\mathbf{M} = \begin{pmatrix} \mathbf{D} & \mathbf{E} \\ \mathbf{E}^T & \mathbf{0} \end{pmatrix}$$

$$\mathbf{D} = \text{diag}\left(\frac{2}{\delta_0^2}, \frac{2}{\delta_1^2}, \dots, \frac{2}{\delta_{k-1}^2}\right)$$

$$\mathbf{E} = \underbrace{\begin{pmatrix} \vec{\xi}_0 & 0 & \dots & 0 \\ 0 & \vec{\xi}_1 & \dots & 0 \\ \dots & \dots & \dots & 0 \\ 0 & 0 & \dots & \vec{\xi}_{n_{\text{triplet}}-1} \end{pmatrix}}_{\mathbf{N}_{\text{triplet}}} \left. \vphantom{\begin{pmatrix} \vec{\xi}_0 \\ \vec{\xi}_1 \\ \dots \\ \vec{\xi}_{n_{\text{triplet}}-1} \end{pmatrix}} \right\} n_{\text{hit}}$$

matrix size is $\mathbf{N}_{\text{triplet}} \times \mathbf{N}_{\text{hit}}$

Solution

Ansatz for inverted matrix

$$\mathbf{M}^{-1} = \begin{pmatrix} \mathbf{A} & \mathbf{B} \\ \mathbf{B}^T & \mathbf{C} \end{pmatrix} \quad \text{with} \quad \begin{aligned} \mathbf{C} &= -(\mathbf{E}^T \mathbf{D}^{-1} \mathbf{E})^{-1} = \mathbf{C}^T \\ \mathbf{B} &= -\mathbf{D}^{-1} \mathbf{E} \mathbf{C} \end{aligned}$$

Task is the inversion of the matrix: $\mathbf{P} = \mathbf{E}^T \mathbf{D}^{-1} \mathbf{E}$ (rank N_{triplet})

$$\mathbf{P} = \begin{pmatrix} p_{00} & p_{01} & p_{02} & 0 & 0 & \dots & 0 \\ p_{10} & p_{11} & p_{12} & p_{13} & 0 & \dots & 0 \\ p_{20} & p_{21} & p_{22} & p_{23} & p_{24} & \dots & 0 \\ 0 & p_{31} & p_{32} & p_{33} & p_{34} & \dots & 0 \\ 0 & 0 & p_{42} & p_{43} & p_{44} & \dots & 0 \\ \dots & \dots & \dots & \dots & \dots & \dots & \dots \\ 0 & 0 & 0 & 0 & 0 & \dots & p_{n-1,n-1} \end{pmatrix}$$

\mathbf{P} is a symmetric and sparse matrix. It can be inverted by the method of **Cholesky Decomposition**

usually a small matrix!

similar problem appears in the (related) broken line fit

Result

The **fitted curvature** is given by a weighted sum of the triplet curvatures:

$$c_0 = \frac{\sum_j w_j \hat{c}_j}{\sum_j w_j} \quad \text{with} \quad w_j = -\frac{1}{2} \sum_i \mathbf{C}_{ij}$$

$$\sigma_c^{-2} = \frac{1}{2} \left(\frac{d^2 \chi^2(c)}{dc^2} \right)_{(c=c_0)} = \frac{1}{2} \sum_{i,j} (\mathbf{B}^T \mathbf{D} \mathbf{B})_{ij} = -\frac{1}{2} \sum_{i,j} \mathbf{C}_{ij}$$

The **hit positions** are obtained from:

$$\delta_k = \sum_j B_{kj} (c - \hat{c}_j)$$

$$\sigma_{\text{hit},k} = \sigma_c \sum_j B_{kj}$$

The **correlation** between hits can be calculated as well:

$$\text{cov}_{\text{hit},kl} = \sigma_c^2 \left(\sum_j B_{kj} \right) \left(\sum_i B_{li} \right)$$

→ **the fit provides all information!**

Cholesky Decomposition

Want to solve: $A x = b$

Matrix A is symmetric, so we can write:

$$A = L L^T$$

with L being a left-sided matrix.

(alternatively one can also use: $A = L D L^T$)

If L is known (\rightarrow next page)

the matrix inversion is done by a recursive

A) forward and

B) backward substitution:

$$\text{A) } L y = b \qquad \text{B) } L^T x = y$$

$$\begin{pmatrix} A_{11} & A_{21}^* & A_{31}^* & A_{41}^* \\ A_{21} & A_{22} & A_{32}^* & A_{42}^* \\ A_{31} & A_{32} & A_{33} & A_{43}^* \\ A_{41} & A_{42} & A_{43} & A_{44} \end{pmatrix} * \begin{pmatrix} x_1 \\ x_2 \\ x_3 \\ x_4 \end{pmatrix} = \begin{pmatrix} b_1 \\ b_2 \\ b_3 \\ b_4 \end{pmatrix}$$



$$\begin{pmatrix} A_{11} & A_{21}^* & A_{31}^* & A_{41}^* \\ A_{21} & A_{22} & A_{32}^* & A_{42}^* \\ A_{31} & A_{32} & A_{33} & A_{43}^* \\ A_{41} & A_{42} & A_{43} & A_{44} \end{pmatrix} = \begin{pmatrix} L_{11} & 0 & 0 & 0 \\ L_{21} & L_{22} & 0 & 0 \\ L_{31} & L_{32} & L_{33} & 0 \\ L_{41} & L_{42} & L_{43} & L_{44} \end{pmatrix} * \begin{pmatrix} L_{11} & L_{21}^* & L_{31}^* & L_{41}^* \\ 0 & L_{22} & L_{32}^* & L_{42}^* \\ 0 & 0 & L_{33} & L_{43}^* \\ 0 & 0 & 0 & L_{44} \end{pmatrix}$$

Note, this a sequential algorithm – not parallelisable!

Calculation of Matrix L

from Wikipedia

Cholesky–Banachiewicz algorithm

$$L_{j,j} = (\pm) \sqrt{A_{j,j} - \sum_{k=1}^{j-1} L_{j,k}^2},$$
$$L_{i,j} = \frac{1}{L_{j,j}} \left(A_{i,j} - \sum_{k=1}^{j-1} L_{i,k} L_{j,k} \right) \quad \text{for } i > j.$$

```
for (i = 0; i < dimensionSize; i++) {
  for (j = 0; j <= i; j++) {
    float sum = 0;
    for (k = 0; k < j; k++)
      sum += L[i][k] * L[j][k];
    if (i == j)
      L[i][j] = sqrt(A[i][i] - sum);
    else
      L[i][j] = (1.0 / L[j][j] * (A[i][j] - sum));
  }
}
```

- Note, that the algorithm needs to be executed in a predefined **sequential** order
- In general **NxN steps** are required, costs scale as **N³** (sums)
- However, for a **pentadiagonal matrix** the algorithm scales as **N²**
- (Other algorithms for a tri-diagonal matrix scale as **N logN**)

Example: Calculation of L and y and x

Cholesky–Banachiewicz algorithm

For 3x3 matrix:

$$1 \quad \mathbf{L} = \begin{pmatrix} \sqrt{A_{11}} & 0 & 0 \\ A_{21}/L_{11} & \sqrt{A_{22} - L_{21}^2} & 0 \\ A_{31}/L_{11} & (A_{32} - L_{31}L_{21})/L_{22} & \sqrt{A_{33} - L_{31}^2 - L_{32}^2} \end{pmatrix}$$

Example:

$$A = \begin{pmatrix} 1 & 1 & 1 \\ 1 & 2 & 5 \\ 1 & 5 & 21 \end{pmatrix} \quad \rightarrow \quad L = \begin{pmatrix} 1 & 0 & 0 \\ 1 & 1 & 0 \\ 1 & 4 & 2 \end{pmatrix}$$

$$L^T = \begin{pmatrix} 1 & 1 & 1 \\ 0 & 1 & 4 \\ 0 & 0 & 2 \end{pmatrix}$$

$$2 \quad L y = b : \quad \begin{matrix} \downarrow \\ \begin{pmatrix} 1 & y_0 & 0 & 0 \\ 1 & y_0 & 1 & y_1 & 0 \\ 1 & y_0 & 4 & y_1 & 2 & y_2 \end{pmatrix} = \begin{pmatrix} b_0 \\ b_1 \\ b_2 \end{pmatrix} \end{matrix}$$

$$\rightarrow y = \begin{pmatrix} b_0 \\ b_0 - b_1 \\ 1/2 b_2 + 3/2 b_1 - 2b_0 \end{pmatrix}$$

$$3 \quad L^T x = y : \quad \begin{pmatrix} 1 & x_0 & 1 & x_1 & 1 & x_2 \\ 0 & 1 & x_1 & 4 & x_2 \\ 0 & 0 & 2 & x_2 \end{pmatrix} \begin{matrix} \uparrow \\ \\ \end{matrix} = \begin{pmatrix} y_0 \\ y_1 \\ y_2 \end{pmatrix}$$

$$\rightarrow x = \begin{pmatrix} y_0 - y_1 + 3 y_2 \\ y_1 - 4 y_2 \\ y_2 \end{pmatrix}$$

Performance of Choleski Decomposition

Sparse Cholesky Factorization on FPGA Using Parameterized Model

(Yichun Sun et al., <https://doi.org/10.1155/2017/3021591>)

	CPU	CPU-GPU	FPGA (200MHz)	
Matrix	HSL_MA87 times (s)	CHOLMOD times (s)	Ours	
different big matrices	nd3k	2.02	2.92	1.96 ($m = 2, k = 256$)
	nd24k	28.56	22.17	10.08 ($m = 8, k = 32$)
	Trefethen_20000b	12.63	8.49	3.58 ($m = 4, k = 64$)

Conclusions

- FPGA implementation is about 2 times faster for large matrices
- The FPGA runs at O(10) times lower speed and consumes 50% of the power

For hit finding with O(10) hits, the gain using FPGAs is probably small

Summary

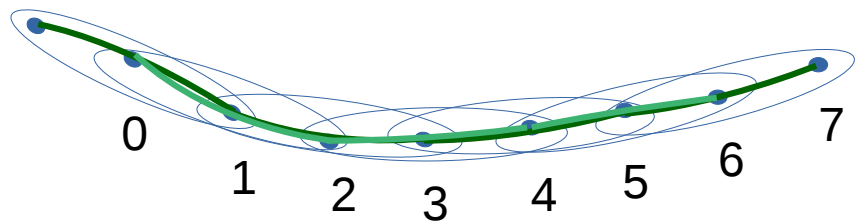
- **Linearisation** and **parallelisation** are very powerful methods to accelerate computations
- A new **parallelisable track fit** with hit uncertainties is presented (for 2D)
 - An extension to 3D is straight forward (not shown)
 - An extension to include multiple scattering can also easily be done (not shown)
→ **General Triplet Based Track Fit** (paper in preparation)
 - **Inverting** the sparse **matrix A** (e.g. Cholesky decomposition) is not parallelisable; but FPGA and GPU could maybe used for acceleration
- The proposed track fit is based on hit-triplets which are **ideal seeds for track finding** (→ graph theory)

Backup

Comparison: Triplet Fit versus GBL

Triplet Model:

fit MS angles and hit positions with respect to **reference triplets**



- **Single triplets** can be fitted including **MS and hit uncertainty algebraically** (not shown)
- triplets can be all fitted in **parallel** including **data preparation**
- looks like an **ideal** algorithm for track finding?
- But speed of algo not measured yet

GBL Model:

fit MS angles and hit positions with respect to a given **reference track**

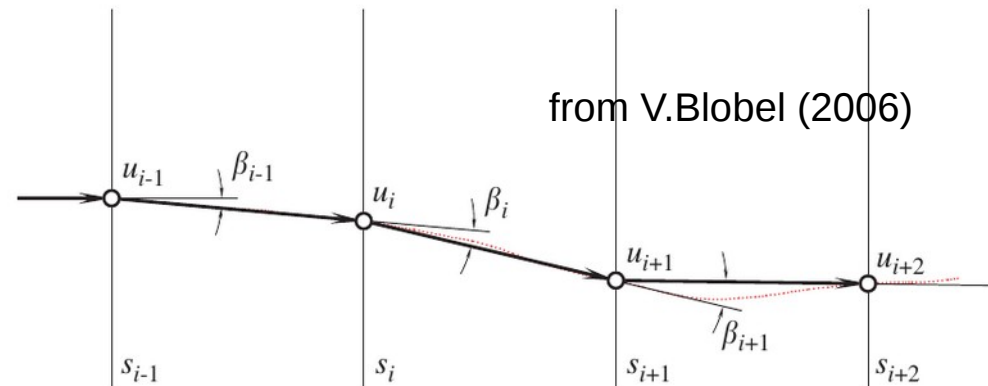


Fig. 3. Particle trajectory with fitted residuals u_i and kink angles β_i .

- **GBL** is faster than **Kalman** fitter
- GBL requires **reference trajectory** as input
- therefore not suitable for track finding