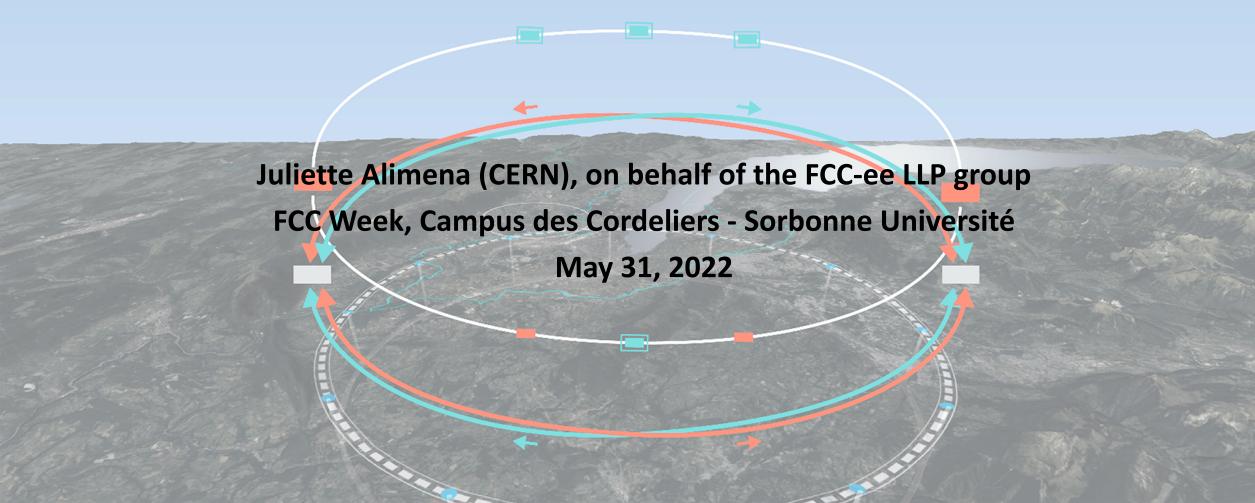




This project is supported from the European Union's Horizon 2020 research and innovation program under grant agreement No 951754.



## Long-Lived Particles at the FCC-ee

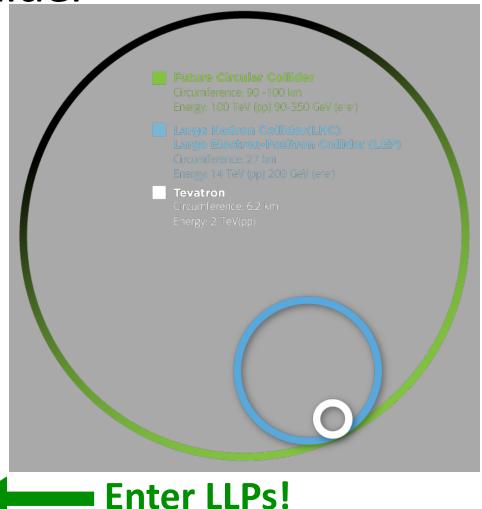


**Future Circular Collider** 

Future Circular Collider (FCC) will have one 100 km tunnel, two stages:

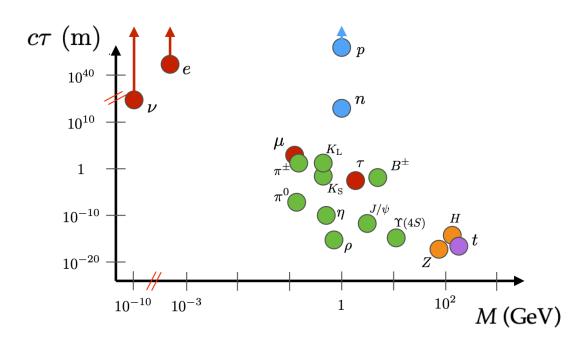
- Stage 1: FCC-ee (Z, W, H, tt) as Higgs EW and top factory at high luminosities
- Stage 2: FCC-hh (~100 TeV) as natural continuation at energy frontier, with ion and eh options

The FCC is is a frontier Higgs, top, electroweak, and flavor factory where we can directly discover new physics



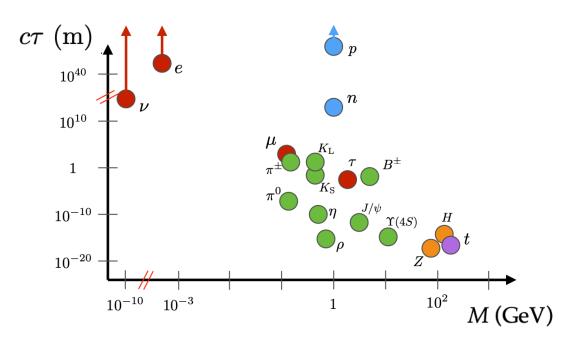
## Long-Lived Particles (LLPs)

# Standard model particles span a wide range of lifetimes ( $\tau$ )



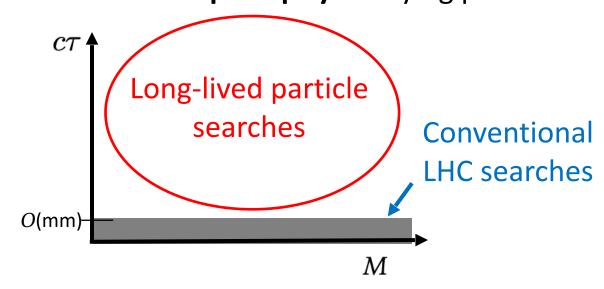
## Long-Lived Particles (LLPs)

Standard model particles span a wide range of lifetimes ( $\tau$ )



We expect **new phenomena** to have a wide range of lifetimes as well

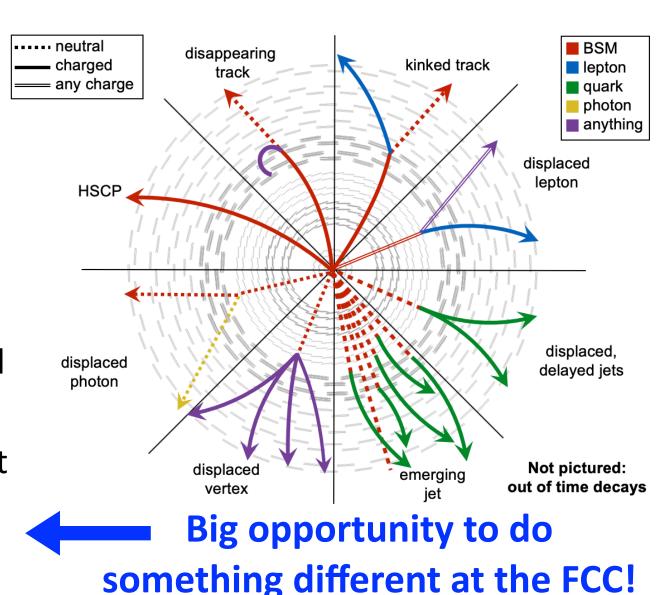
But **conventional searches** for new phenomena at the LHC are for **promptly** decaying particles



We also need to look for new particles with long lifetimes!

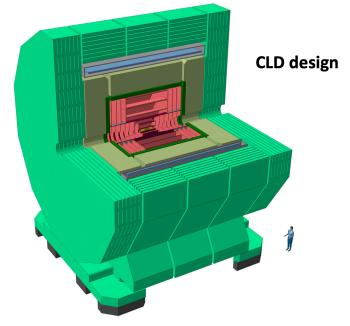
### Long-Lived Particle Searches

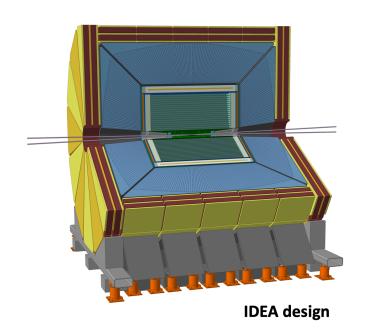
- Wide variety of:
  - Charges
  - Final states
  - Decay locations
  - Lifetimes
- Design **signature-driven** searches
- Often interpret results with a benchmark model, but can expand variety of scenarios
- Challenges of the LHC: main detect triggers, offline reconstruction not designed for displaced particles



### Detectors at the FCC

- **Two detector concepts** used for integration, performance, and cost estimates:
  - CLD design: adapted for the FCC-ee by the CERN Linear Collider Detector group
  - **IDEA design:** specifically designed for the FCC-ee (and CEPC)
- Now ready to take a broader look at the physics potential and optimize detector designs for a complete physics program
- Have the opportunity to design general-purpose detectors with LLPs in mind!
  - Can prioritize e.g. displaced tracking and precision timing information
  - Can also prioritize LLPs in the online filtering and offline reconstruction
- FCC-ee new baseline is consistent with having 2 or 4 detectors
  - Opportunities for new, creative designs!
  - E.g. HECATE dedicated to long lifetimes (<u>arXiv:2011.01005</u>)





### Past and Ongoing Work

- Snowmass LOI, white paper (arXiv:2203.05502)
- Several Masters student theses done or in progress:
  - <u>Sissel Bay Nielsen</u> (University of Copenhagen, 2017)
  - Rohini Sengupta (Uppsala University, 2021)
  - Lovisa Rygaard (Uppsala University, 2022)
  - Tanishq Sharma (University of Geneva, 2022)
- Will discuss 3 long-lived physics cases, focusing on what is new since Feb FCC Physics Workshop:
  - 1. Heavy Neutral Leptons (HNLs)
  - 2. Axion-like Particles (ALPs)
  - 3. Higgs bosons with exotic decays to LLPs

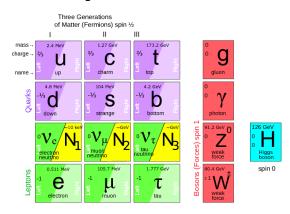
### **Analysis Overview**

#### Analysis goals

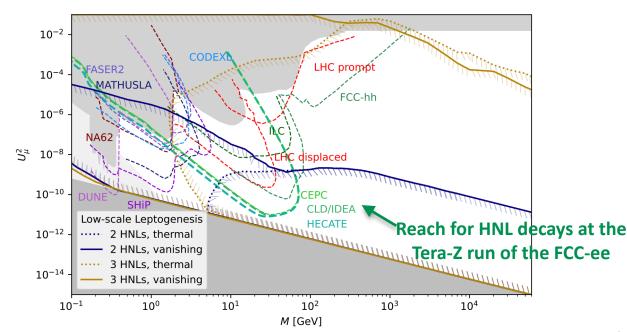
- Perform an FCC case study with the "official" analysis tools and framework available for the FCC
  - FCCAnalysis, Delphes, latest IDEA card, centrally-produced samples, etc.
- Try to be as realistic as possible
  - E.g. use high stats background samples
- Latest improvements (since FCC Physics Workshop)
  - Increased background MC sample size
    - Due to now able to run the analysis in batch and store processed samples in common eos area thanks!
  - We profit also from recent technical developments in FCCAnalyses
    - Much smoother analysis workflow

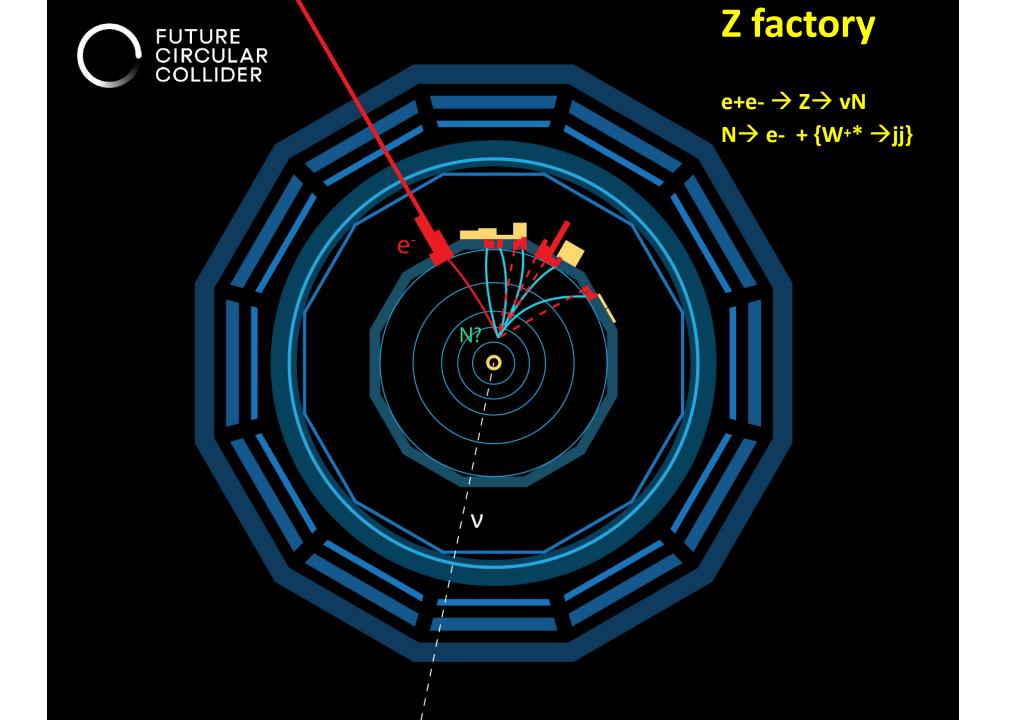
### 1st Physics Case: LL Heavy Neutral Leptons

- Right-handed, sterile neutrinos
- Dirac or Majorana fermions with sterile neutrino quantum numbers
- Heavy enough to not disrupt the simplest BBN bounds and/or unstable on cosmological timescales
- Could shed light some open questions of the SM: Neutrino masses,
   Baryon asymmetry, Dark matter
- FCC will probe space not constrained by astrophysics or cosmology, complementary to accelerator and neutrino prospects
- HNLs at the FCC-ee are right in the parameter region that is good for baryogenesis! <u>arXiv:2106.16226</u>
- See Saw type I is simplest way we can probe HNLs at the FCC, but not only one
- Most studies only consider N1 to be light and longlived, but not necessarily true: <u>arXiv:1907.13034</u>

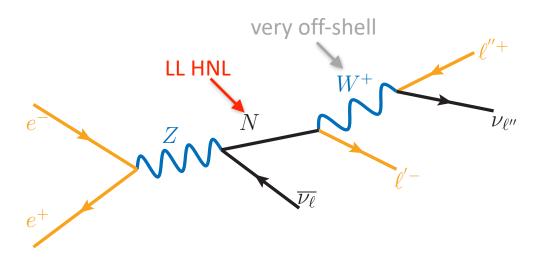


arXiv:2203.05502

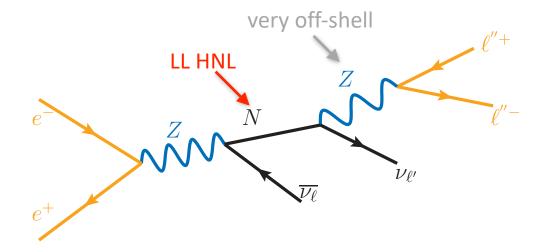




### Simulating HNLs



- Generated Majorana and Dirac HNLs with the SM\_HeavyN\_CKM\_AllMasses\_LO and SM\_HeavyN\_Dirac\_CKM\_Masses\_LO models (arXiv:1411.7305, arXiv:1602.06957)
- Started with the  $ee\nu$  final state (suggested as early as 1984(!) by <u>S. Petcov</u>)
- FCC-ee,  $\sqrt{s} = 91 \text{ GeV}$
- Generated in Madgraph5 v3.2.0 + Pythia8 + Delphes, with the latest IDEA card
- Experimental signature of LL HNLs: displaced vertex

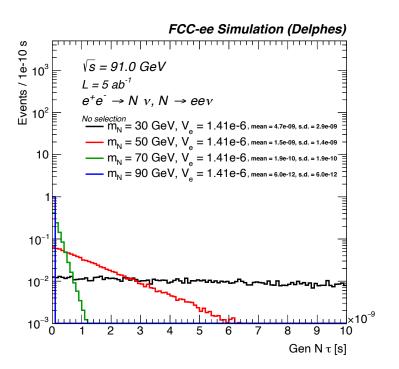


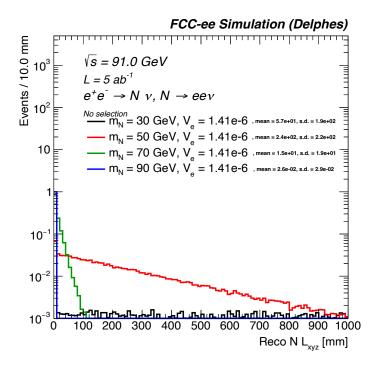
$$L \sim 0.025 \text{m} \left(\frac{10^{-6}}{V_l}\right)^2 \left(\frac{100 \text{ GeV}}{m_N}\right)^5$$

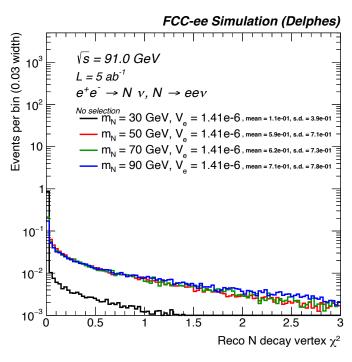
[Valid when  $m_N \lesssim 100~{
m GeV}$ , arXiv:1905.11889]

Get long-lived HNLs when coupling and mass are small

### HNL Lifetime and Decay Vertex





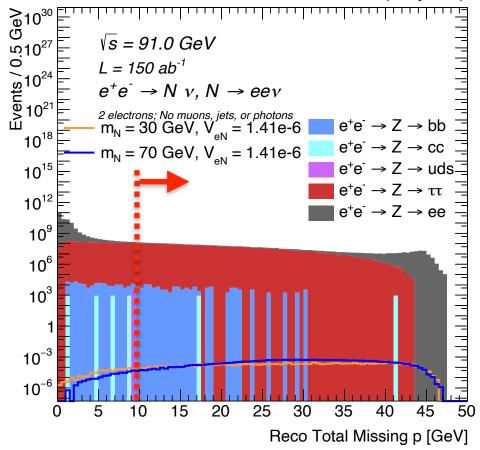


- Confirmed HNL signal kinematics behave as expected, at gen and reco level
- For example, for m = 50 GeV, Ve = 1.41e-6, the mean of the generated lifetime is 1.5E-9 s —> 45 cm, which is what we expected
  - On the other hand, m = 90 GeV, Ve = 1.41e-6 is pretty prompt
- Reco L<sub>xyz</sub> (3D decay length) and vertex  $\chi^2$  distributions are also understood
  - m = 30 GeV, Ve = 1.41e-6 is fairly displaced, so less isolated electrons are reconstructed

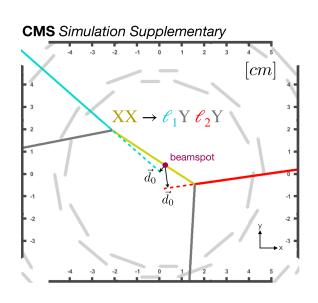
### S vs B: Missing Energy

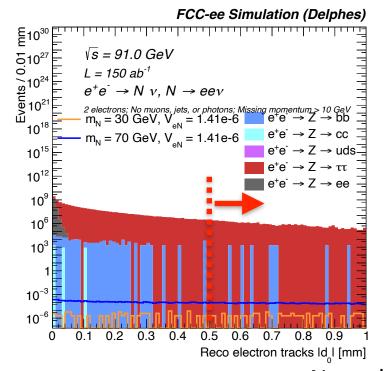
- Then added centrally-produced "spring2021" background samples with the IDEA detector, at 91 GeV CME
- Recent update: more background statistics
- Can look at the total missing energy at an e+ecollider!
- Require missing energy > 10 GeV

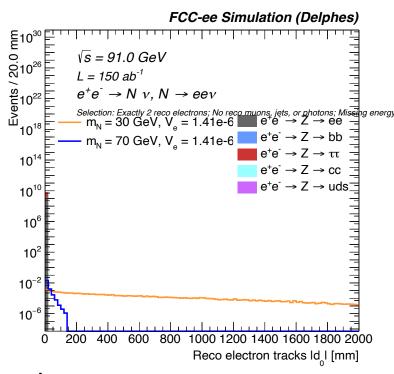
### FCC-ee Simulation (Delphes)



### S vs B: Impact Parameter







Note the x-axis ranges

- Another good discriminating variable is the impact parameter
- Recent update: more background statistics
- Started by looking at transverse impact parameter ( $d_0$  or  $d_{xy}$ ), but will move to the 3D impact parameter ( $d_{xyz}$ )
- $|d_0| > 0.5$  mm removes the vast majority of the SM background

### Selection

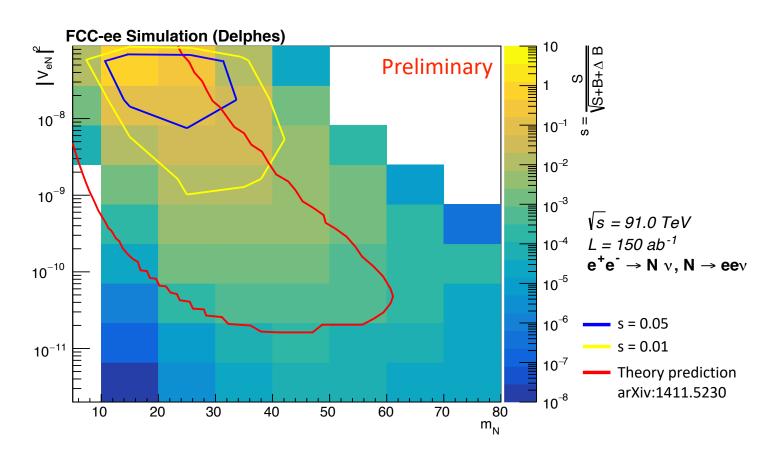
- Preliminary event selection
- Tables show the expected number of events at 150 ab-1, cumulative after each cut (on reco variables)
- Here used  $10^7 10^9$  (5\*104) raw/unscaled events for background (signal)
  - Will need to generate larger samples for some background
- Most discriminating variables explored so far: missing energy and  $|d_0|$

	Before selection	Exactly 2 reco e	Veto photons, muons, and jets		$ d_0  > 0.5 \; \mathrm{mm}$
$Z \rightarrow ee$	$2.19 \times 10^{11} \pm 6.94 \times 10^{7}$	$1.75 \times 10^{11} \pm 6.19 \times 10^7$	$1.53 \times 10^{11} \pm 5.80 \times 10^{7}$	$7.07 \times 10^8 \pm 3.94 \times 10^6$	$\leq 3.94 \times 10^6$
$Z \rightarrow bb$	$9.97 \times 10^{11} \pm 4.14 \times 10^7$	$5.64 \times 10^8 \pm 9.85 \times 10^5$	$3.25 \times 10^5 \pm 2.36 \times 10^4$	$1.22 \times 10^5 \pm 1.45 \times 10^4$	$1.72 \times 10^3 \pm 1.72 \times 10^3$
Z  o  au au	$2.21 \times 10^{11} \pm 7.00 \times 10^7$	$5.49 \times 10^9 \pm 1.10 \times 10^7$	$5.10 \times 10^9 \pm 1.06 \times 10^7$	$2.52 \times 10^9 \pm 7.47 \times 10^6$	$6.64 \times 10^4 \pm 3.84 \times 10^4$
$Z \to cc$	$7.82 \times 10^{11} \pm 2.61 \times 10^7$	$1.69 \times 10^7 \pm 1.21 \times 10^5$	$5.22 \times 10^3 \pm 2.13 \times 10^3$	$1.74 \times 10^3 \pm 1.23 \times 10^3$	$\leq 1.23 \times 10^3$
$Z \to uds$	$2.79 \times 10^{12} \pm 8.83 \times 10^{7}$	$2.30 \times 10^7 \pm 2.54 \times 10^5$	$2.79 \times 10^3 \pm 2.79 \times 10^3$	$\leq 2.79 \times 10^3$	$\leq 2.79 \times 10^3$

	Before selection	Exactly 2 reco e	Vetoes	p > 10  GeV	$ d_0  > 0.5 \text{ mm}$
$m_N = 10 \text{ GeV},  V_{eN}  = 2 \times 10^{-4}$	$2534 \pm 11$	$1006 \pm 7$	$996 \pm 7$	$951 \pm 7$	$907 \pm 7$
$ m_N = 20 \text{ GeV},  V_{eN}  = 9 \times 10^{-5}$	$458 \pm 2$	$313 \pm 2$	$308 \pm 2$	$293\pm2$	$230 \pm 1$
$m_N = 20 \text{ GeV},  V_{eN}  = 3 \times 10^{-5}$	$51.0 \pm 0.2$	$34.7\pm0.2$	$34.2\pm0.2$	$32.6 \pm 0.2$	$31.2 \pm 0.2$
$m_N = 30 \text{ GeV},  V_{eN}  = 1 \times 10^{-5}$	$5.01 \pm 0.02$	$3.85\pm0.02$	$3.76 \pm 0.02$	$3.54 \pm 0.02$	$3.39 \pm 0.02$
$m_N = 50 \text{ GeV},  V_{eN}  = 6 \times 10^{-6}$	$1.23 \pm 0.01$	$0.99 \pm 0.01$	$0.96\pm0.01$	$0.92 \pm 0.01$	$0.729 \pm 0.004$

### Sensitivity

#### New!



• FOM = 
$$\frac{S}{\sqrt{S + B + \Delta B}}$$

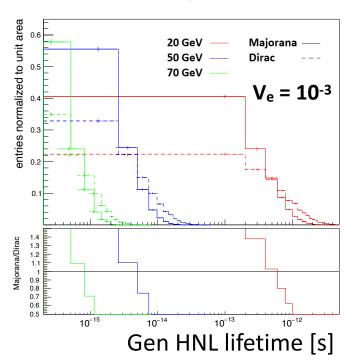
- This analysis:  $N \rightarrow ee\nu$ 
  - Contours show where FOM = 0.01 and 0.05
- Theory prediction from arXiv:1411.5230
  - Includes all HNL decay modes, not only electrons
- Next step: add more decay modes, particularly  $N \rightarrow ejj$

### Bonus: Dirac vs Majorana HNLs

Dirac (LNC) and Majorana (LNC+LNV) HNLs produce different kinematic distributions: <a href="mailto:arXiv:2105.06576">arXiv:2105.06576</a>
Variables that can distinguish between Majorana and Dirac HNLs:

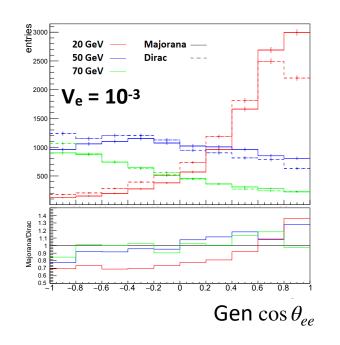
#### **HNL** Lifetime

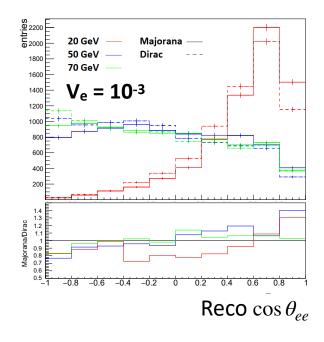
(model-dependent)



 $\cos \theta_{ee}$ 

(opening angle between final state electron/positron)

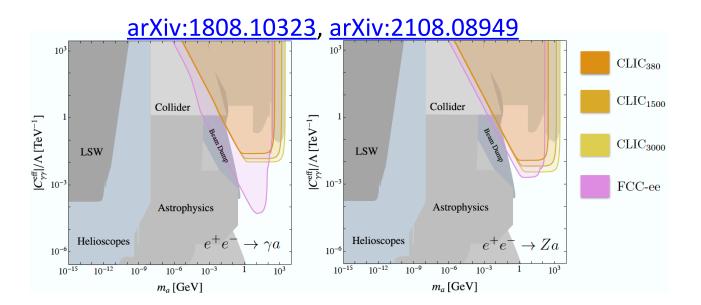


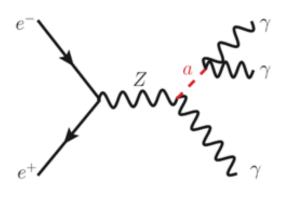


Next: improve reconstruction, find more discriminating variables

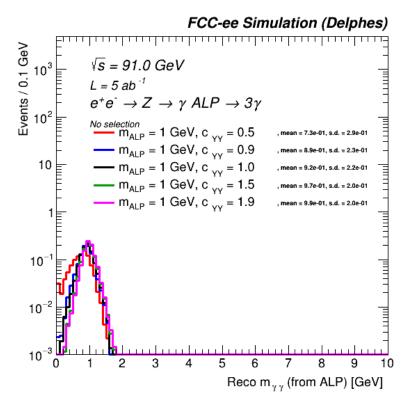
### 2nd Physics Case: LL Axion-Like Particles

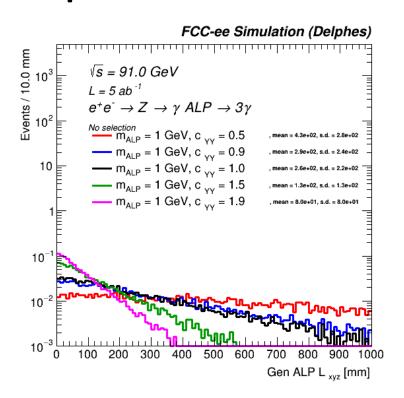
- Axion-like Particles (ALPs) are pseudo Nambu-Goldstone bosons of spontaneously broken global symmetries in BSM scenarios
- Very weakly coupled to the dark sector
- Get long-lived ALPs when couplings and mass are small
- At the FCC-ee:
  - Orders of magnitude of parameter space accessible
  - Especially sensitive to final states with at least 1 photon
- Privately generated ALPs in Madgraph5 v3.2.0 + Pythia8 + Delphes, with the latest IDEA card,  $\sqrt{s}=91~{\rm GeV}$  (arXiv:1808.10323)





### Variables to Explore

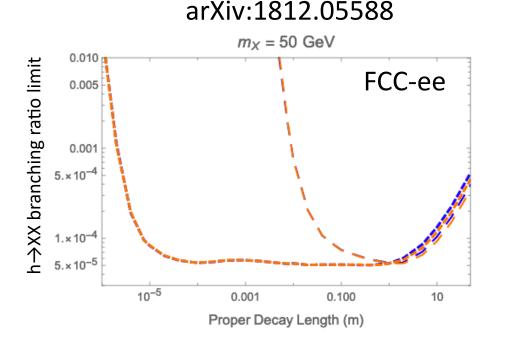


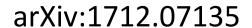


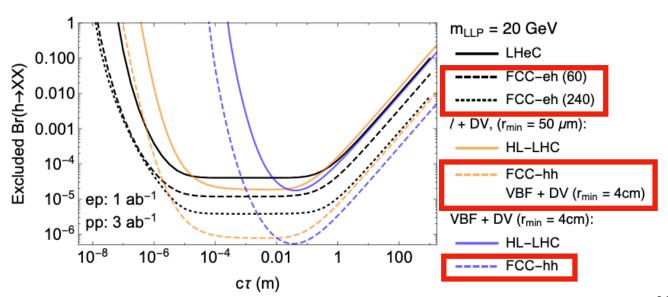
- Started with simulating 1 GeV ALP mass, vary the coupling
- ALP mass confirmed with the reco invariant mass from the 2 photons coming from the ALP
- ALP decay length will also be a nice discriminating variable
- Ready for more personpower to step in and complete an analysis, guidance available!

### 3rd Benchmark: Exotic Higgs decays to LLPs

- Higgs bosons could undergo exotic decays to e.g. scalars that could be long-lived
- Exotic Higgs decays to LLPs could be explored at future colliders
  - Twin Higgs models with displaced exotic Higgs boson decays, Hidden Valley models with Higgs bosons decaying to neutral LLPs (arXiv:1812.05588)
  - LLPs from Higgsinos or exotic Higgs decays (arXiv:1712.07135)
- Another call for personpower: can generate with e.g. this model in Madgraph, then incorporate into FCC framework



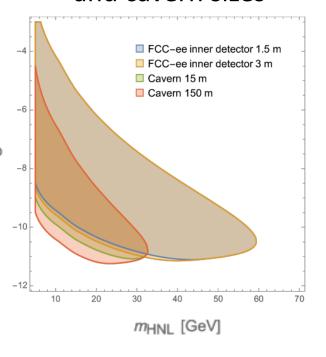




### Other Topics to Explore at the FCC

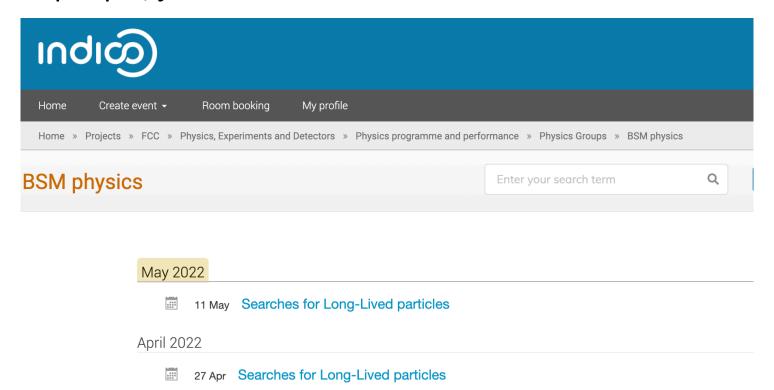
- How well we can distinguish a long-lived HNL/ALP/exotic Higgs decay from SM backgrounds
  - For leptonic decays? For hadronic decays? For decays to photons?
- Vertexing performance of the FCC prototype detectors
- Time-of-flight performance
- Different detector configurations: can we probe a larger/different —
  theory landscape?
  - Bigger tracker? More layers?
- Majorana vs Dirac HNLs
- Not an exhaustive list!

Preliminary study
(arXiv:2106.15459) shows
the sensitivity of HNLs to
different inner detector
and cavern sizes



### LLPs at FCC-ee group

- Informal group with:
  - Meetings: <a href="https://indico.cern.ch/category/5664/">https://indico.cern.ch/category/5664/</a>
  - Mailing list: <LLP-FCCee-informal@cern.ch>
- We welcome new people, join us!

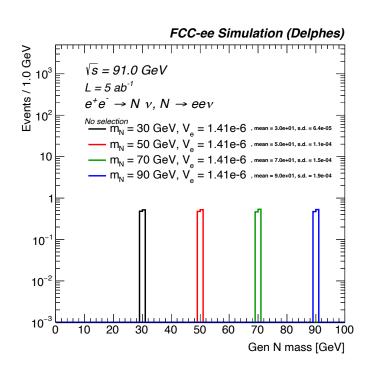


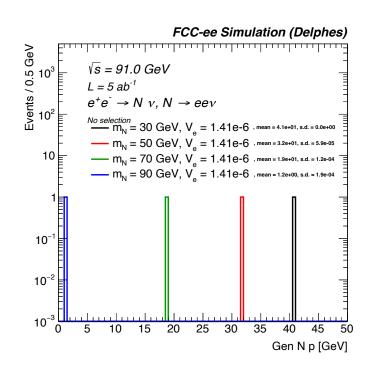
### Summary

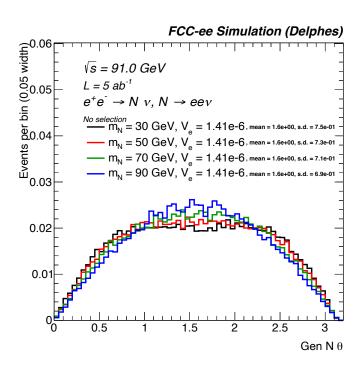
- To discover new phenomena, should look where no one else has looked before
- One way to do this: long-lived particles!
- The FCC will have the ability to uniquely probe LLP areas of phase space, and discovery potential!
- Many interesting signals: Heavy Neutral Leptons, hidden sectors, axion-like particles, exotic Higgs decays, and more
- Showed some brand new results
- We now have the opportunity to design detectors and algorithms with LLPs in mind
- A lot can be learned from Delphes
- See Rebeca's and Suchita's talks for more!
- Also see the ongoing <u>LLP11 workshop</u>
- Plenty of phase space to explore at the FCC! Let's make sure we don't miss new physics!

## Backup

#### **Generated HNL Kinematics**







- At the FCC-ee, should look at total momentum,  $\theta$ , and total missing energy!
- Generator-level distributions look as expected
  - Momentum decreases as HNL mass increases
  - Slightly more central events as HNL mass increases