

# Long lived particles at FCC

Suchita Kulkarni (she/her)

Junior group leader

 @suchi\_kulkarni

FCC week 2022

# Why, how, where LLPs?

- Answers to a range of unsolved questions can generically lead to LLP

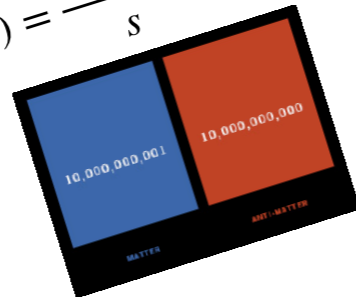
See talk by T. You



$$\theta \sim \frac{g_s^2}{32\pi^2} \tilde{G}_b^{\mu\nu} G_{b\mu\nu}$$

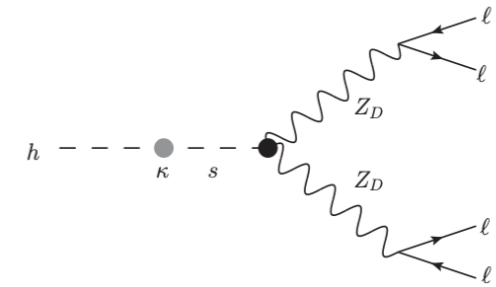


$$\eta(b) = \frac{n_b - n_{\bar{b}}}{s} \sim 10^{-10}$$



$$\mathcal{L} \subset -\frac{1}{4} \hat{B}_{\mu\nu} \hat{B}^{\mu\nu} - \frac{1}{4} \hat{Z}_{D\mu\nu} \hat{Z}_D^{\mu\nu} + \frac{1}{2} \frac{\epsilon}{\cos\theta} \hat{Z}_{D\mu\nu} \hat{B}^{\mu\nu} + \frac{1}{2} m_{D,0}^2 \hat{Z}_D^\mu \hat{Z}_{D\mu}$$

$$V_0(H, S) = -\mu^2 |H|^2 + \lambda |H|^4 - \mu_S^2 |S|^2 + \lambda_S |S|^4 + \kappa |S|^2 |H|^2$$



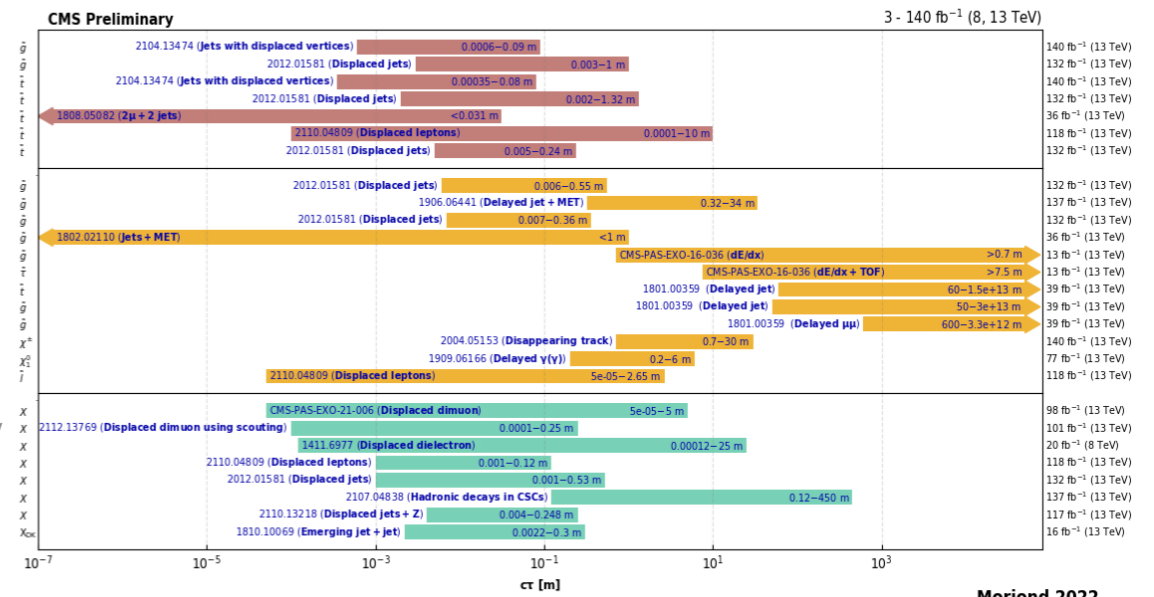
- LHC BSM program: new physics either heavy and/or feebly coupled and/or leads to soft final states

ATLAS Heavy Particle Searches\* - 95% CL Upper Exclusion Limits  
Status: March 2022

Model	$\ell, \gamma$	Jets†	$E_T^{\text{miss}}$	$[\mathcal{L} dt](\text{fb}^{-1})$	Limit	Reference
<b>Extra dimensions</b>	ADD $G_{KK} + g/q$	0, $e, \mu, \tau, \gamma$	1-4 J	Yes	139	$M_{\text{pl}}$ 11.2 TeV
	ADD non-resonant $\gamma\gamma$	2 $\gamma$	2 J	Yes	36.7	$M_{\text{pl}}$ 8.6 TeV
	ADD GBH	2 $\gamma$	2 J	Yes	37.0	$M_{\text{pl}}$ 8.9 TeV
	ADD BH multijet	2 $\gamma$	2 J	Yes	3.6	$M_{\text{pl}}$ 9.55 TeV
	RS1 $G_{KK} \rightarrow \gamma\gamma$	2 $\gamma$	2 J	Yes	139	$G_{KK}$ mass 2.3 TeV
	Bulk RS $G_{KK} \rightarrow WW/ZZ$	multi-channel	2 J	Yes	36.1	$G_{KK}$ mass 4.5 TeV
	Bulk RS $G_{KK} \rightarrow W\gamma$	2 $\gamma$	2 J	Yes	139	$G_{KK}$ mass 2.0 TeV
	Bulk RS $G_{KK} \rightarrow \ell\ell$	2 $\ell$	2 J	Yes	36.1	$G_{KK}$ mass 3.8 TeV
	2UED / RPP	1 $e, \mu$	$\geq 2b, \geq 3J$	Yes	36.1	$G_{KK}$ mass 1.8 TeV
<b>Gauge bosons</b>	SSM $Z' \rightarrow \ell\ell$	2 $e, \mu, \tau, \gamma$	1-4 J	Yes	139	$Z'$ mass 2.4 TeV
	SSM $Z' \rightarrow \tau\tau$	2 $\tau$	2 J	Yes	36.1	$Z'$ mass 2.1 TeV
	Leptophobic $Z' \rightarrow b\bar{b}$	0 $e, \mu, \tau, \gamma$	$\geq 2b, \geq 2J$	Yes	139	$Z'$ mass 4.1 TeV
	Leptophobic $Z' \rightarrow \tau\tau$	0 $e, \mu, \tau, \gamma$	$\geq 2b, \geq 2J$	Yes	139	$Z'$ mass 4.1 TeV
	SSM $W' \rightarrow \ell\nu$	1 $e, \mu, \tau$	1 J	Yes	139	$W'$ mass 6.0 TeV
	SSM $W' \rightarrow \tau\nu$	1 $\tau$	1 J	Yes	139	$W'$ mass 5.0 TeV
	SSM $W' \rightarrow b\bar{b}$	0 $e, \mu, \tau, \gamma$	$\geq 2b, \geq 2J$	Yes	139	$W'$ mass 4.4 TeV
	HVT $W' \rightarrow WZ \rightarrow \ell\nu q\bar{q}$ model B	1 $e, \mu, \tau$	2 J	Yes	139	$W'$ mass 4.3 TeV
	HVT $W' \rightarrow WZ \rightarrow \ell\nu l'l'$ model C	3 $e, \mu, \tau$	2 J	Yes	139	$W'$ mass 340 GeV
	HVT $W' \rightarrow Wl$ model B	0 $e, \mu, \tau, \gamma$	$\geq 2b, \geq 2J$	Yes	139	$W'$ mass 3.2 TeV
	LRSM $W_R \rightarrow \mu N_2$	2 $\mu$	1 J	Yes	80	$W_R$ mass 5.0 TeV
<b>CI</b>	CI $qqqq$	0 $e, \mu, \tau, \gamma$	2 J	Yes	37.0	A 21.8 TeV
	CI $\ell\ell qq$	2 $e, \mu, \tau, \gamma$	2 J	Yes	139	A 35.8 TeV
	CI $e\ell e\ell$	2 $e, \mu, \tau, \gamma$	2 J	Yes	139	A 1.8 TeV
	CI $\mu\mu\mu\mu$	2 $\mu$	1 J	Yes	139	A 2.0 TeV
	CI $t\bar{t}t\bar{t}$	$\geq 1 e, \mu, \tau, \gamma$	$\geq 1b, \geq 1J$	Yes	36.1	A 2.57 TeV
<b>DM</b>	Axial-vector med. (Dirac DM)	0 $e, \mu, \tau, \gamma$	1-4 J	Yes	139	$m_{\text{DM}}$ 378 GeV
	Pseudo-scalar med. (Dirac DM)	0 $e, \mu, \tau, \gamma$	1-4 J	Yes	139	$m_{\text{DM}}$ 378 GeV
	Vector med. $Z'$ -2HDM (Dirac DM)	0 $e, \mu, \tau, \gamma$	2 J	Yes	139	$m_{\text{DM}}$ 560 GeV
	Pseudo-scalar med. 2HDM+A	multi-channel	2 J	Yes	139	$m_{\text{DM}}$ 1.8 TeV
<b>LQ</b>	Scalar LQ 1 <sup>st</sup> gen	2 $e, \mu, \tau, \gamma$	$\geq 2J$	Yes	139	LQ mass 1.7 TeV
	Scalar LQ 2 <sup>nd</sup> gen	2 $\mu, \tau, \gamma$	$\geq 2J$	Yes	139	LQ mass 1.7 TeV
	Scalar LQ 3 <sup>rd</sup> gen	1 $\tau, \gamma$	2 J	Yes	139	LQ mass 1.2 TeV
	Scalar LQ 3 <sup>rd</sup> gen	0 $e, \mu, \tau, \gamma$	$\geq 2J, \geq 2b$	Yes	139	LQ mass 1.24 TeV
	Scalar LQ 3 <sup>rd</sup> gen	$\geq 2 e, \mu, \tau, \gamma$	$\geq 1b, \geq 1J$	Yes	139	LQ mass 1.43 TeV
	Scalar LQ 3 <sup>rd</sup> gen	0 $e, \mu, \tau, \gamma$	$\geq 1b, \geq 1J$	Yes	139	LQ mass 1.85 TeV
	Vector LQ 3 <sup>rd</sup> gen	1 $\tau, 2b$	2 J	Yes	139	LQ mass 1.26 TeV
	Vector LQ 3 <sup>rd</sup> gen	1 $\tau, 2b$	2 J	Yes	139	LQ mass 1.77 TeV
<b>Heavy quarks</b>	VLO $T\bar{T} \rightarrow Zl + X$	2 $e, \mu, \tau, \gamma$	$\geq 1b, \geq 1J$	Yes	139	T mass 1.4 TeV
	VLO $BB \rightarrow WlZb + X$	2 $b, \gamma$	$\geq 1b, \geq 1J$	Yes	139	B mass 1.34 TeV
	VLO $T_{3,1} T_{3,2} T_{3,3} \rightarrow Wl + X$	RSB $\geq 2 e, \mu, \tau, \gamma$	$\geq 1b, \geq 1J$	Yes	36.1	T mass 1.64 TeV
	VLO $T \rightarrow HlZl$	1 $e, \mu, \tau, \gamma$	$\geq 1b, \geq 3J$	Yes	139	T mass 1.8 TeV
	VLO $T \rightarrow Wb$	1 $e, \mu, \tau, \gamma$	$\geq 1b, \geq 1J$	Yes	36.1	T mass 1.85 TeV
	VLO $B \rightarrow Hb$	0 $e, \mu, \tau, \gamma$	$\geq 2b, \geq 1J$	Yes	139	B mass 2.0 TeV
<b>Excited fermions</b>	Excited quark $q^* \rightarrow q\bar{q}$	1 $\gamma$	2 J	Yes	139	$q^*$ mass 6.7 TeV
	Excited quark $q^* \rightarrow q\gamma$	1 $\gamma$	1 J	Yes	36.7	$q^*$ mass 5.3 TeV
	Excited quark $q^* \rightarrow b\bar{g}$	3 $e, \mu, \tau, \gamma$	1 J	Yes	36.1	$q^*$ mass 2.6 TeV
	Excited lepton $\ell^*$	3 $e, \mu, \tau, \gamma$	1 J	Yes	20.3	$\ell^*$ mass 1.6 TeV
	Excited lepton $\nu^*$	3 $e, \mu, \tau, \gamma$	1 J	Yes	20.3	$\nu^*$ mass 1.6 TeV
<b>Other</b>	Type III Seesaw	2, 3, 4 $e, \mu, \tau, \gamma$	$\geq 2J$	Yes	139	$N^c$ mass 910 GeV
	LRSM Majorana	2, 3 $e, \mu, \tau, \gamma$	2 J	Yes	36.1	$N^c$ mass 350 GeV
	Higgs triplet $H^{\pm\pm} \rightarrow W^+W^+$	2, 3, 4 $e, \mu, \tau, \gamma$ (SS)	various	Yes	139	$H^{\pm\pm}$ mass 1.08 TeV
	Higgs triplet $H^{\pm\pm} \rightarrow \ell\ell$	2, 3, 4 $e, \mu, \tau, \gamma$ (SS)	various	Yes	139	$H^{\pm\pm}$ mass 1.08 TeV
	Higgs triplet $H^{\pm\pm} \rightarrow \ell\tau$	2, 3, 4 $e, \mu, \tau, \gamma$	various	Yes	139	$H^{\pm\pm}$ mass 1.08 TeV
	Multi-charged particles	3 $e, \mu, \tau, \gamma$	1 J	Yes	20.3	$H^{\pm\pm}$ mass 400 GeV
	Magnetic monopoles	3 $e, \mu, \tau, \gamma$	1 J	Yes	36.1	multi-charged particles mass 1.22 TeV
		3 $e, \mu, \tau, \gamma$	1 J	Yes	36.1	monopole mass 2.37 TeV

\*Only a selection of the available mass limits on new states or phenomena is shown.  
† Small-radius (large-radius) jets are denoted by the letter j (J).

Overview of CMS long-lived particle searches

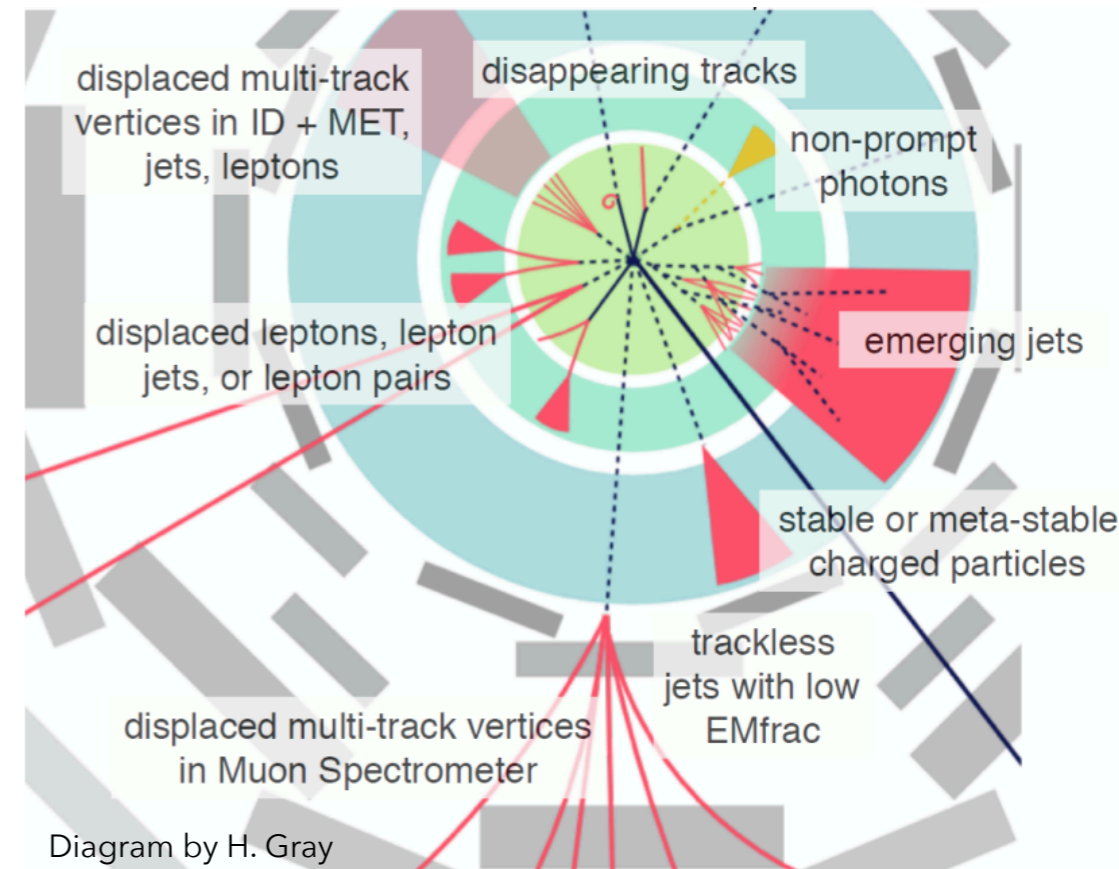


Selection of observed exclusion limits at 95% C.L. (theory uncertainties are not included). The y-axis tick labels indicate the studied long-lived particle.

Moriond 2022

# Why, how, where LLPs?

- Small couplings or small mass differences generically lead to LLPs
- Exact signature depends on LLP quantum numbers and decay modes → many possibilities, necessary to understand experimental sensitivity to these
- FCC-ee: charge, color neutral initial state → direct production of charge, color neutral BSM particles/ final states
  - Light color, electrically charged BSM particles heavily constrained
  - Consider looking for charge, color neutral particle
  - Displaced vertex signature
- FCC-hh: possible to also probe charged/colored BSM particles
  - Disappearing tracks
  - (Heavy) Heavy stable charged particles, stopped particles, R-hadrons as possible signatures
  - These exotic signatures are not covered in this talk



LLPs can also be looked for in MET and prompt final states

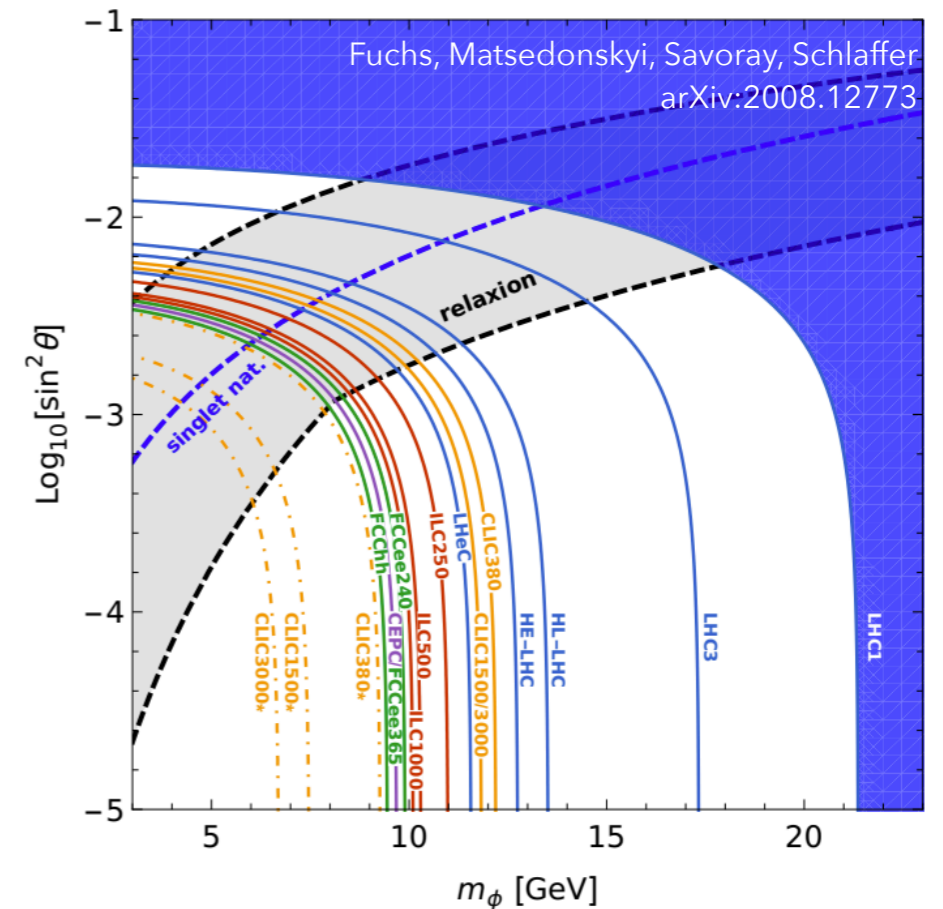
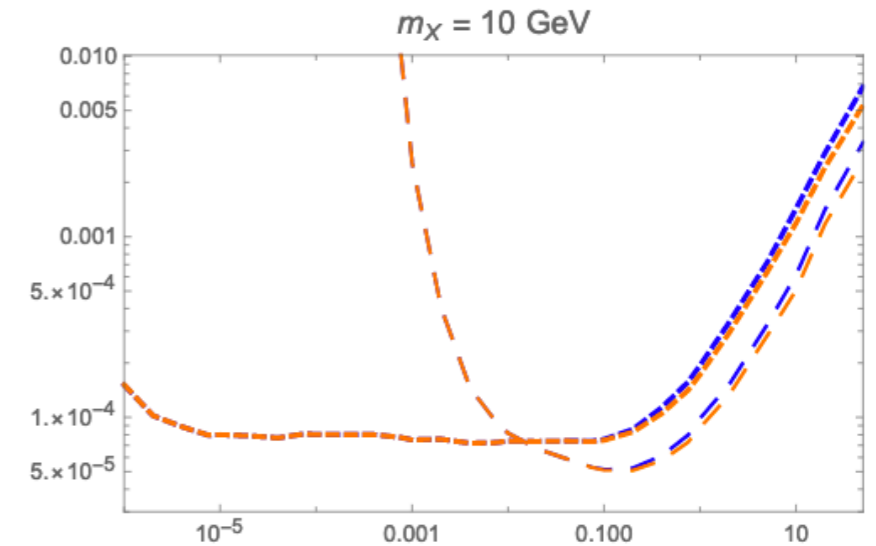
See talk by R. Gonzalez Suarez  
See talks in the morning session

# Higgs exotics

- Higgs precision physics one of the main goals
- Exotic Higgs decays are a window to new physics related to questions from dark matter to naturalness
- $h \rightarrow XX$  a poster child of exotic Higgs benchmarks
  - Advantageous in many models to explore light X
  - Multiple available decay modes and complementarity can lead to excellent coverage
  - Also possible to look at  $h \rightarrow ZX$
  - Mostly considers two body Higgs decays, three body decays are also possible e.g.  $h \rightarrow NN$
- Physics reach can tell us something about neutral naturalness

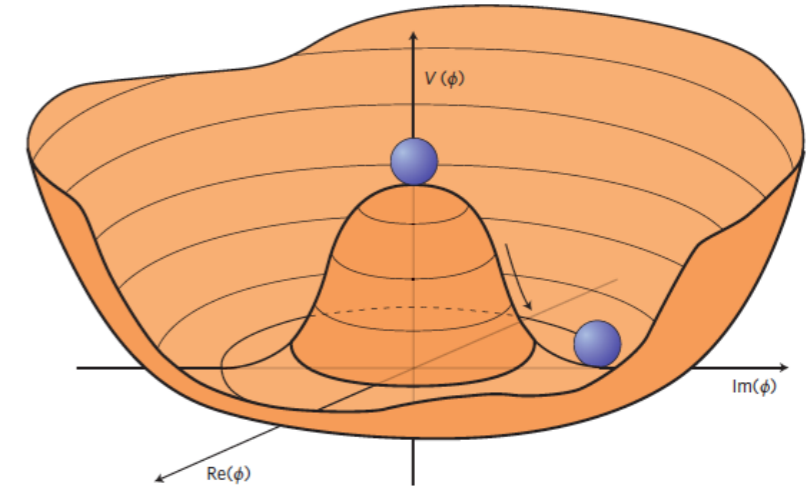
See also [talk by C. Verhaaren](#) in the LLP informal WG meetings

Alipour-Fard, Craig, Jiang, Koren  
arXiv:1812.05588



# Axion/ALPs

- Are there new spontaneously broken global symmetries in nature?
- Search of ALPs i.e. NGBs associated with breaking of such global symmetries
- Examples:
  - $U(1)_{PQ}$ : QCD axion
  - Composite Higgs
  - Supersymmetric pNGB

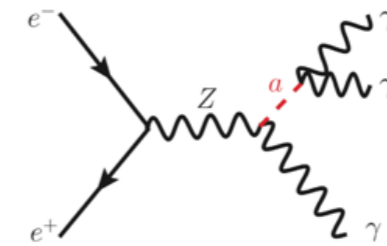


arXiv:9703409, arXiv:0009290, arXiv:1411.3325,  
 arXiv:1504.06084, arXiv:1604.01127,  
 arXiv:1606.03097, arXiv:0902.1483, arXiv:1312.5330,  
 arXiv:1702.02152, arXiv:2104.11064.

## Exotic Z decays

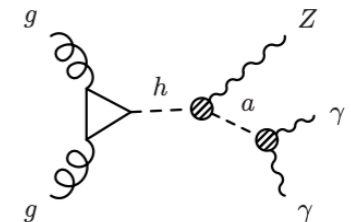
$$\mathcal{L}_{\text{eff}}^{D \leq 5} = \frac{1}{2} (\partial_\mu a)(\partial^\mu a) - \frac{m_{a,0}^2}{2} a^2 + \frac{\partial^\mu a}{f} \sum_F \bar{\psi}_F \mathbf{c}_F \gamma_\mu \psi_F$$

$$+ c_{GG} \frac{\alpha_s}{4\pi} \frac{a}{f} G_{\mu\nu}^a \tilde{G}^{\mu\nu,a} + c_{WW} \frac{\alpha_2}{4\pi} \frac{a}{f} W_{\mu\nu}^A \tilde{W}^{\mu\nu,A} + c_{BB} \frac{\alpha_1}{4\pi} \frac{a}{f} B_{\mu\nu} \tilde{B}^{\mu\nu}$$

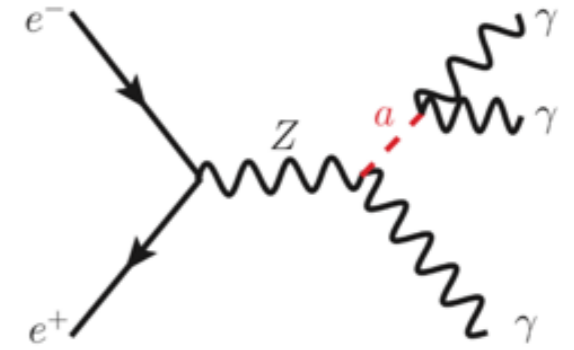
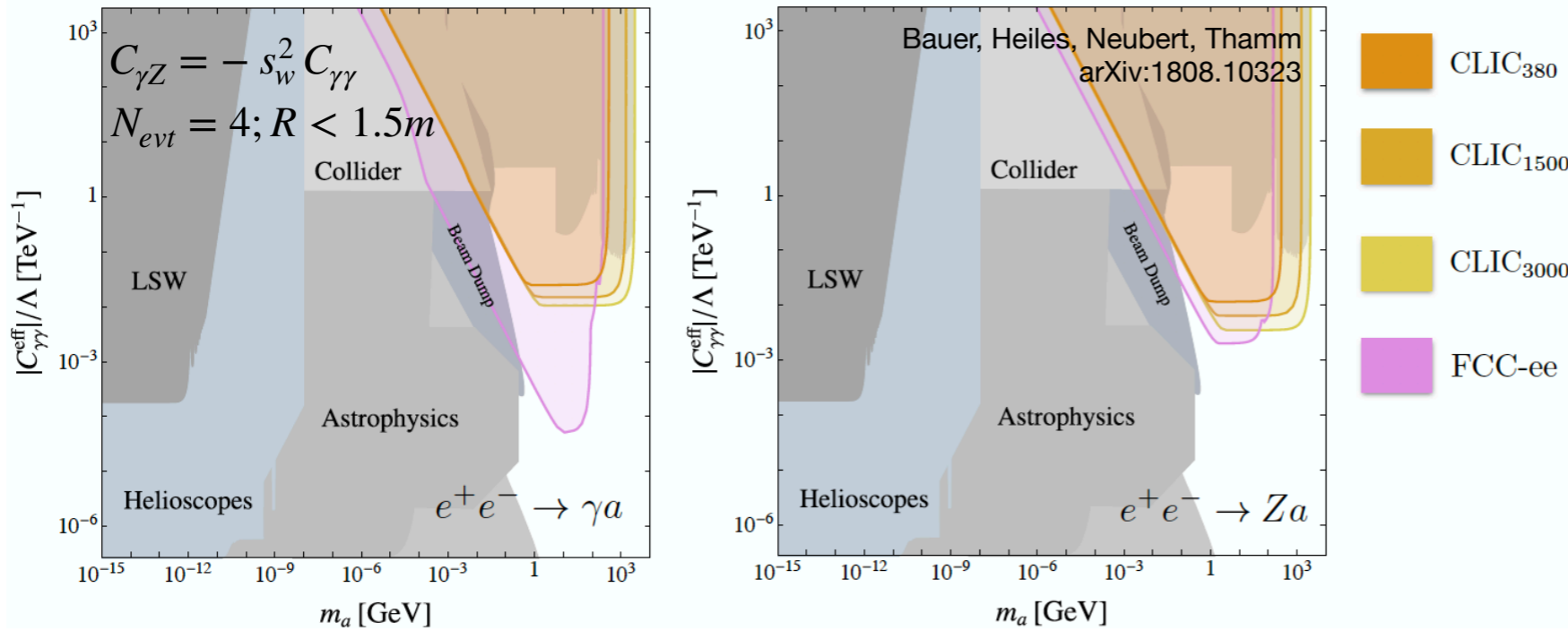


## Exotic h decays

$$\mathcal{L}_{\text{eff}}^{D \geq 6} = \frac{C_{ah}}{\Lambda^2} (\partial_\mu a)(\partial^\mu a) \phi^\dagger \phi + \frac{C_{Zh}}{\Lambda^3} (\partial^\mu a) (\phi^\dagger iD_\mu \phi + \text{h.c.}) \phi^\dagger \phi + \dots$$



# Axion/ALPs portal: FCC-ee



See also e.g. Liu, Wang, Wang, Xue  
arXiv:1712.07237

$$\mathcal{L}_{\text{eff}} \ni e^2 C_{\gamma\gamma} \frac{a}{\Lambda} F_{\mu\nu} \tilde{F}^{\mu\nu} + \frac{2e^2}{s_w c_w} C_{\gamma Z} \frac{a}{\Lambda} F_{\mu\nu} \tilde{Z}^{\mu\nu} + \frac{e^2}{s_w^2 c_w^2} C_{ZZ} \frac{a}{\Lambda} Z_{\mu\nu} \tilde{Z}^{\mu\nu}$$

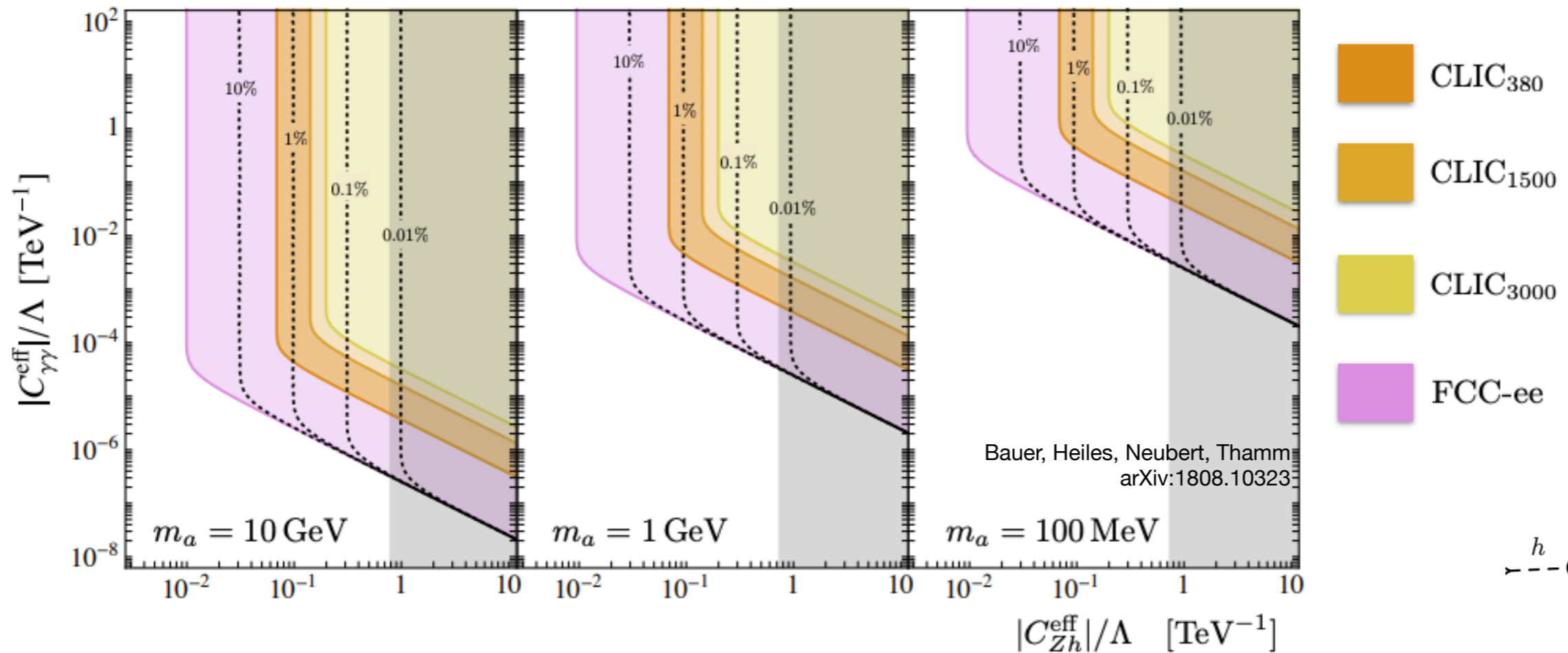
- In broken phase

$$C_{\gamma\gamma} = C_{WW} + C_{BB}; \quad C_{\gamma Z} = c_w^2 C_{WW} - s_w^2 C_{BB}; \quad C_{ZZ} = c_w^4 C_{WW} + s_w^4 C_{BB};$$

- Significantly larger sensitivity for  $e^+e^- \rightarrow \gamma a$  compared to  $e^+e^- \rightarrow Za$
- Low mass axions will lead to collimated photons, will be challenging, necessary to have dedicated studies

# Axion/ALPs portal: FCC-ee

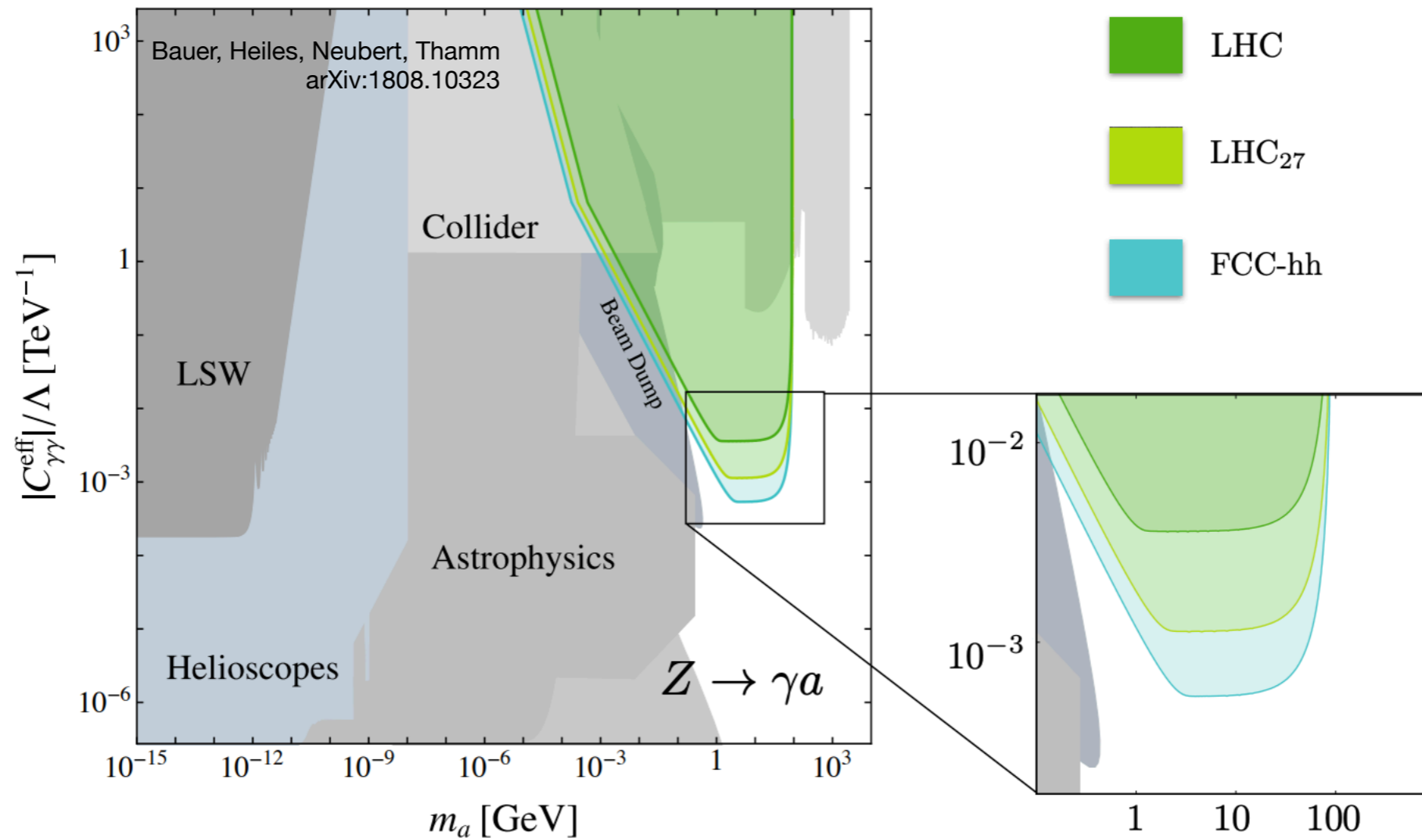
$$e^+e^- \rightarrow ha \rightarrow b\bar{b}\gamma\gamma$$



Bauer, Heiles, Neubert, Thamm  
arXiv:1808.10323

- ALP production and decay governed by two unrelated Wilson coefficients
- Decreasing ALP mass implies increased lifetime, means decreased sensitivity
- Larger coverage for lower  $\sqrt{s}$  because  $\sigma \propto 1/s$
- Contours for 4 events, dotted lines: variation of  $BR(a \rightarrow \gamma\gamma)$  at FCC-ee

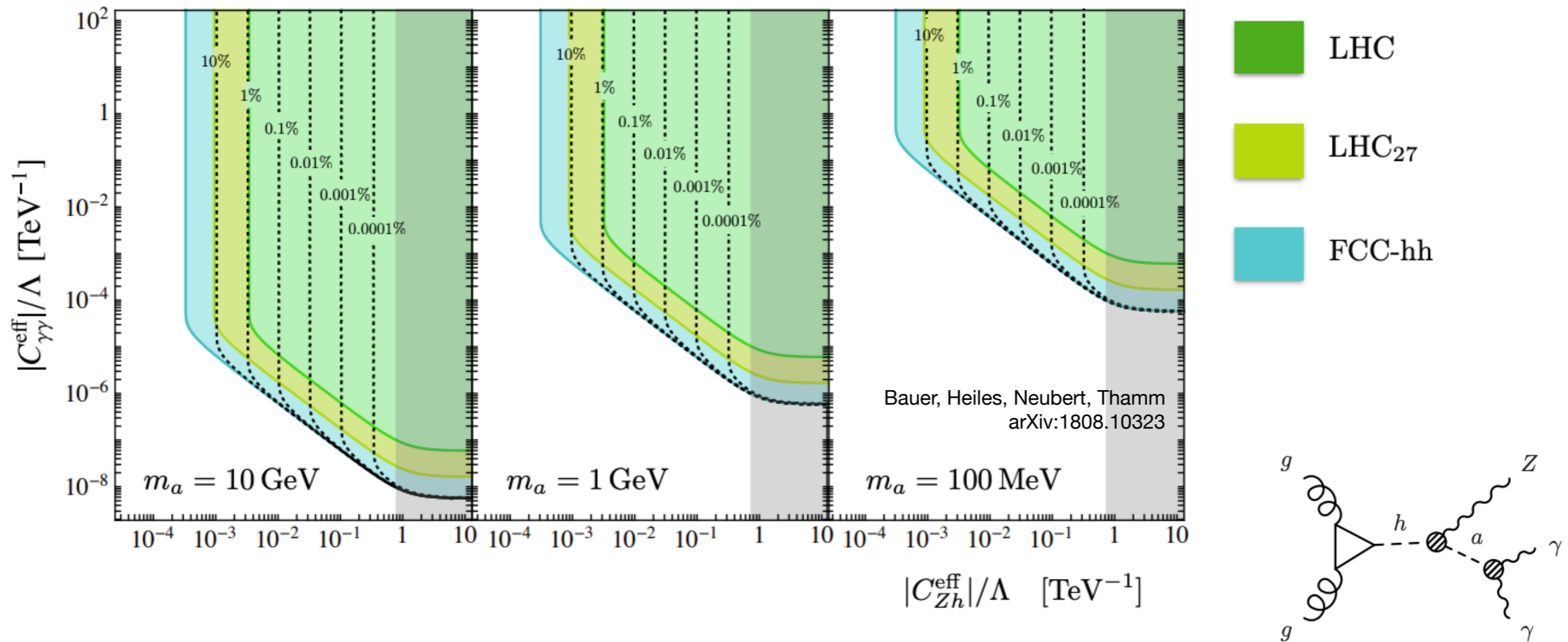
# Axion/ALPs portal: FCC-hh



- At best comparable sensitivity for ALP - photon couplings as FCC-ee has larger luminosity and runs on the Z pole
- FCC-ee may even have higher precision



# Axion/ALPs portal: FCC-hh



- Large gain in production cross section, large sensitivity to couplings
- At FCC-hh the Higgs mode is more promising than the Z mode

# Right handed neutrinos

## Type - I seesaw

$$\mathcal{M}_\nu = \begin{pmatrix} \mathbf{0} & M_D \\ M_D^T & M_N \end{pmatrix}$$

Dirac mass, usually small  $Y_\nu$

Majorana mass, can be heavy, can have  $Y_N \sim 1$ , introduces Lepton number violation

$$\mathcal{L} \supset Y_{ij}^\nu \bar{N}_i L_j \cdot H - \frac{1}{2} M_{ij} \bar{N}_i N_j^c$$

## Minimal inverse seesaw

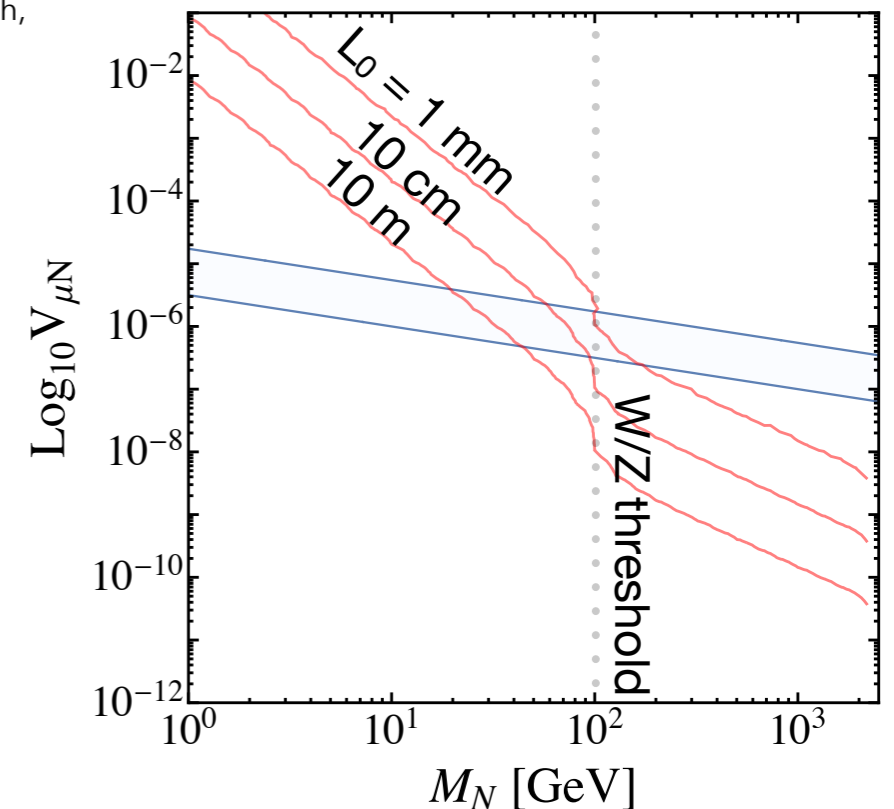
$$\mathcal{M}_\nu = \begin{pmatrix} 0 & m_D & 0 \\ m_D & 0 & m_S \\ 0 & m_S & \mu_S \end{pmatrix}$$

$$m_\nu \approx -\mu_S \frac{m_D^2}{m_S^2}$$

Also possible to generate neutrino masses with e.g. inverse seesaw

$$m_\nu \approx \frac{M_D^2}{M_N} = |V_{\mu N}|^2 \times m_N \quad c\tau_N \propto \left(\frac{10^{-6}}{V_{IN}}\right)^2 \times \left(\frac{100\text{GeV}}{M_N}\right)^5$$

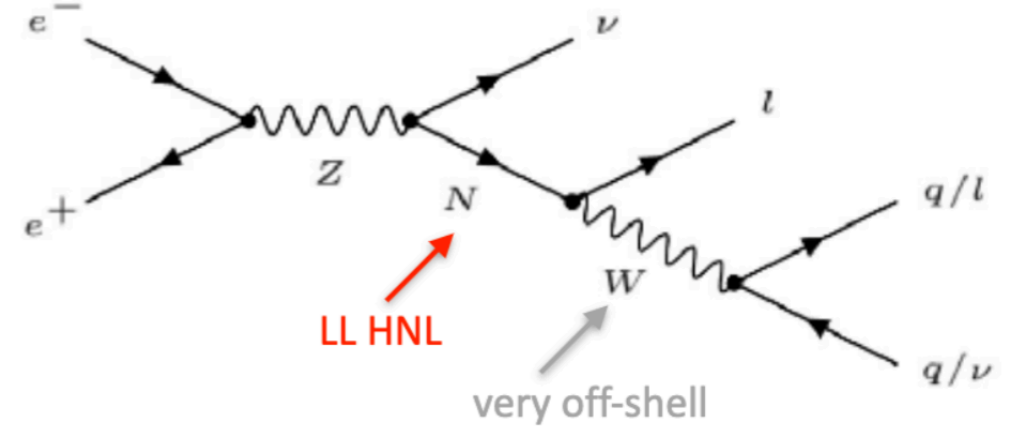
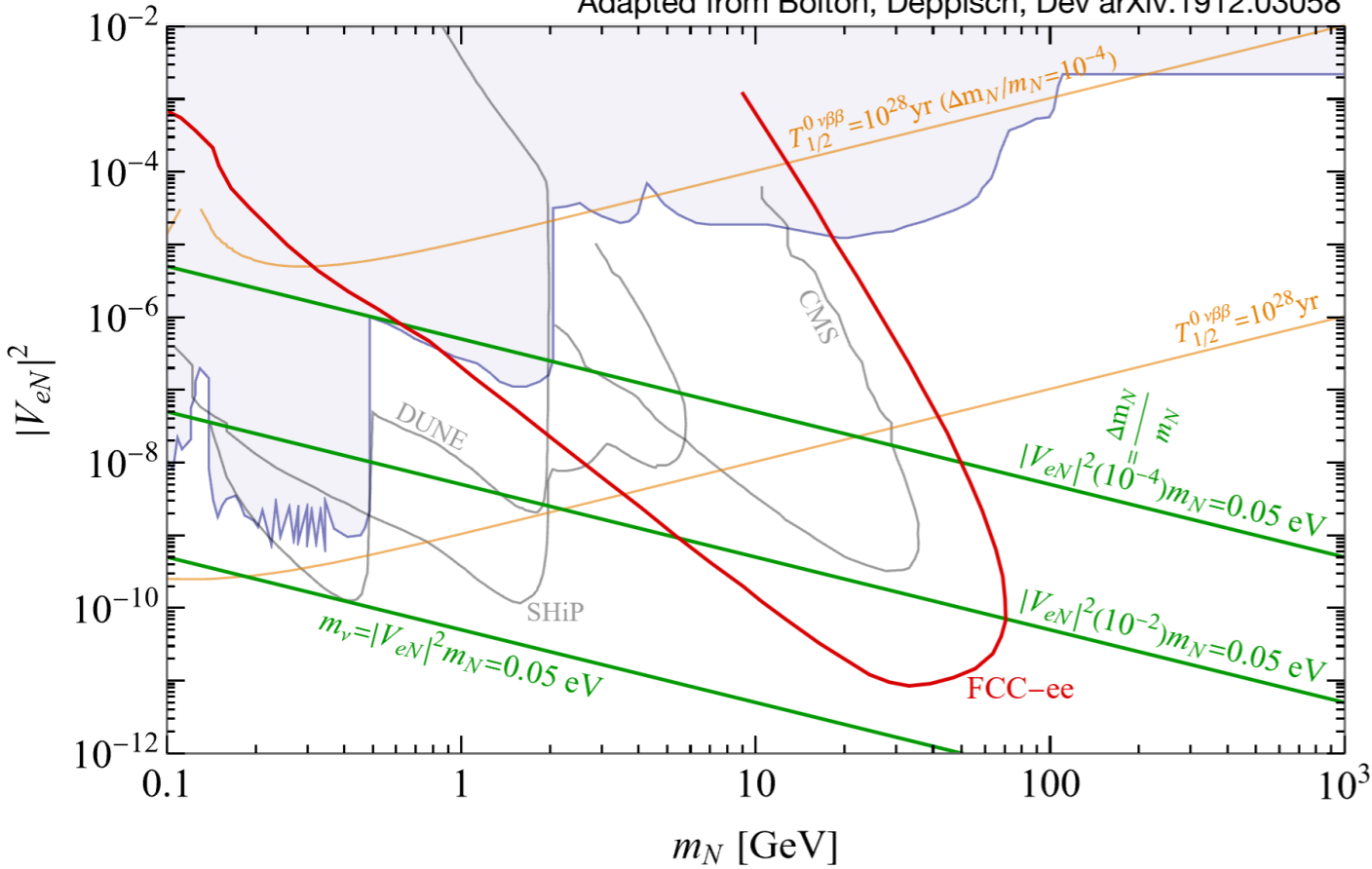
See e.g. Deppisch, New J. Phys. 17 (2015) 075019



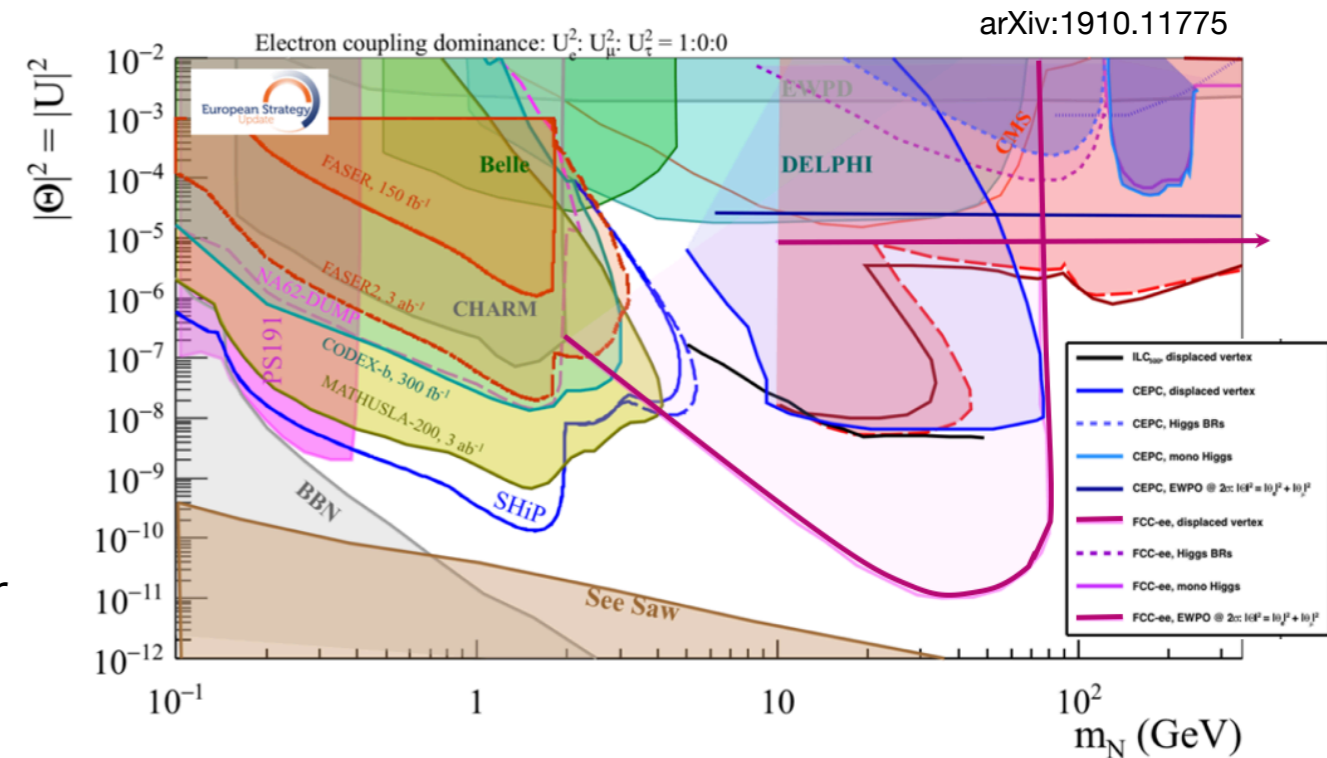
- Are SM neutrinos Dirac or Majorana?
- Are right handed neutrinos responsible for generating SM neutrino masses?
- Are there extra gauge/Higgs bosons featuring in right handed neutrino phenomenology?

# Sterile neutrinos

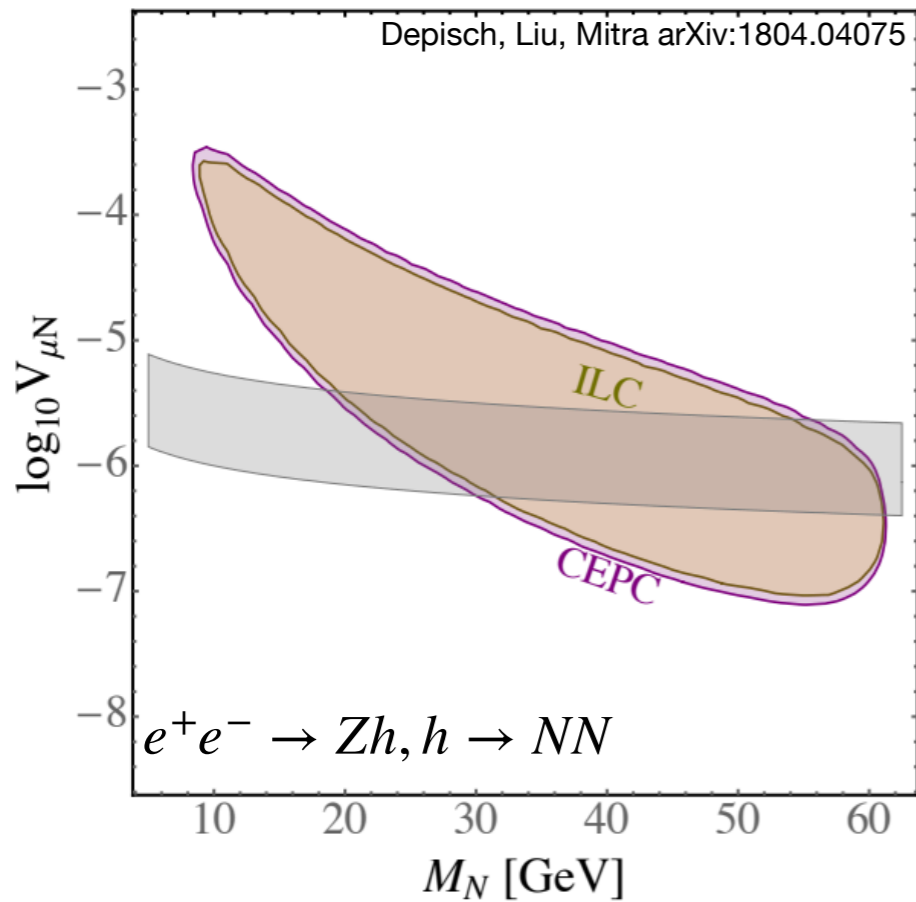
Adapted from Bolton, Deppisch, Dev arXiv:1912.03058



- Simplest production mechanism and assumption one sterile neutrino
- Complementarity with  $0\nu\beta\beta$ , possibility to investigate beyond type-I seesaw mechanisms
- Realistically, more than one right handed neutrinos will be present, necessity to consider beyond 1 flavour scenarios

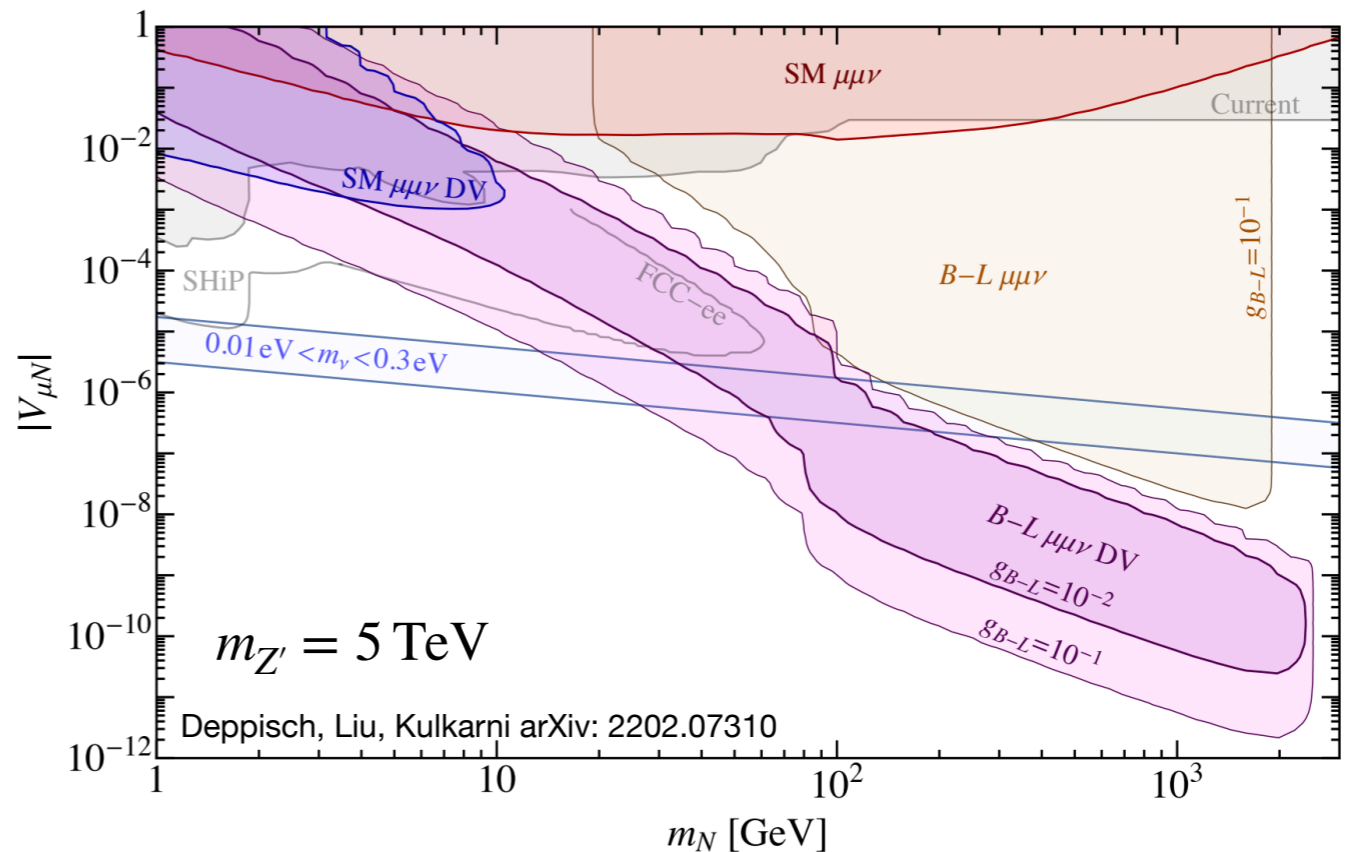


# RH in extended gauge sectors



- Huge complementarity between FCC-ee and FCC-hh
- Access to heavier (model specific) gauge bosons at FCC-hh
- HNL production with large boost at FCC-hh, large coverage of parameter space at the cost of additional model dependence

- Heavy neutrinos can also be accommodated via SM gauge extensions, here an example of B-L extension
- Leads to additional operators for SM mediators to RH neutrino decays
- No prompt leptons in the final state, different event topology
- Possibility to learn about B-L breaking scale

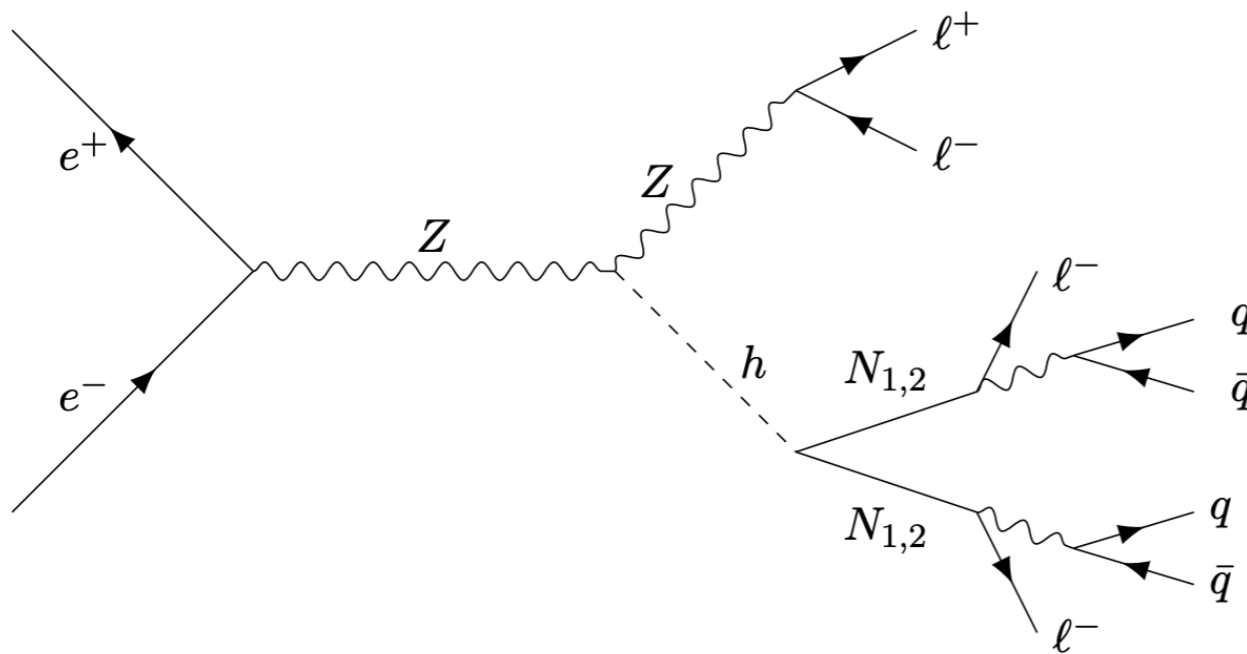


# RH neutrino via effective operators

Barducci, Bertuzzo,  
Caputo, Hernandez, Mele  
arXiv:2011.04725

- HNLs can be incorporated in a variety of SM extensions
- The minimal model is attractive and should certainly be explored
- Alternative possibility: HNL production happens via some effective operators

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \bar{N} \not{\partial} N - \bar{L}_L Y_\nu \tilde{H} N - \frac{1}{2} M_N \bar{N}^c N + \sum_{n>4} \frac{\mathcal{O}^n}{\Lambda^{n-4}} + h.c.$$



$$\mathcal{O}_W = \alpha_W (\bar{L}^c \tilde{H}^*) (\tilde{H}^\dagger L) ,$$

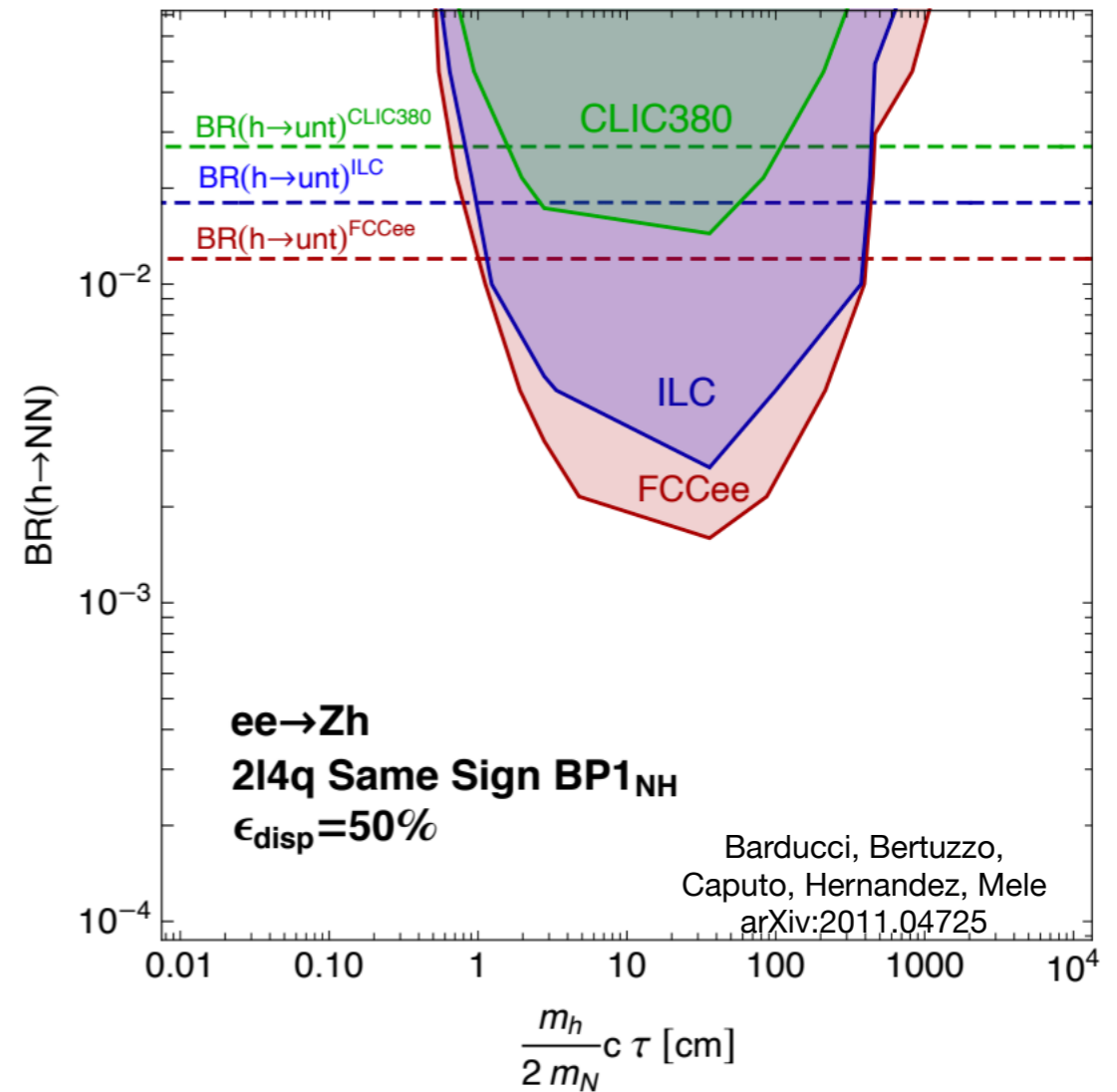
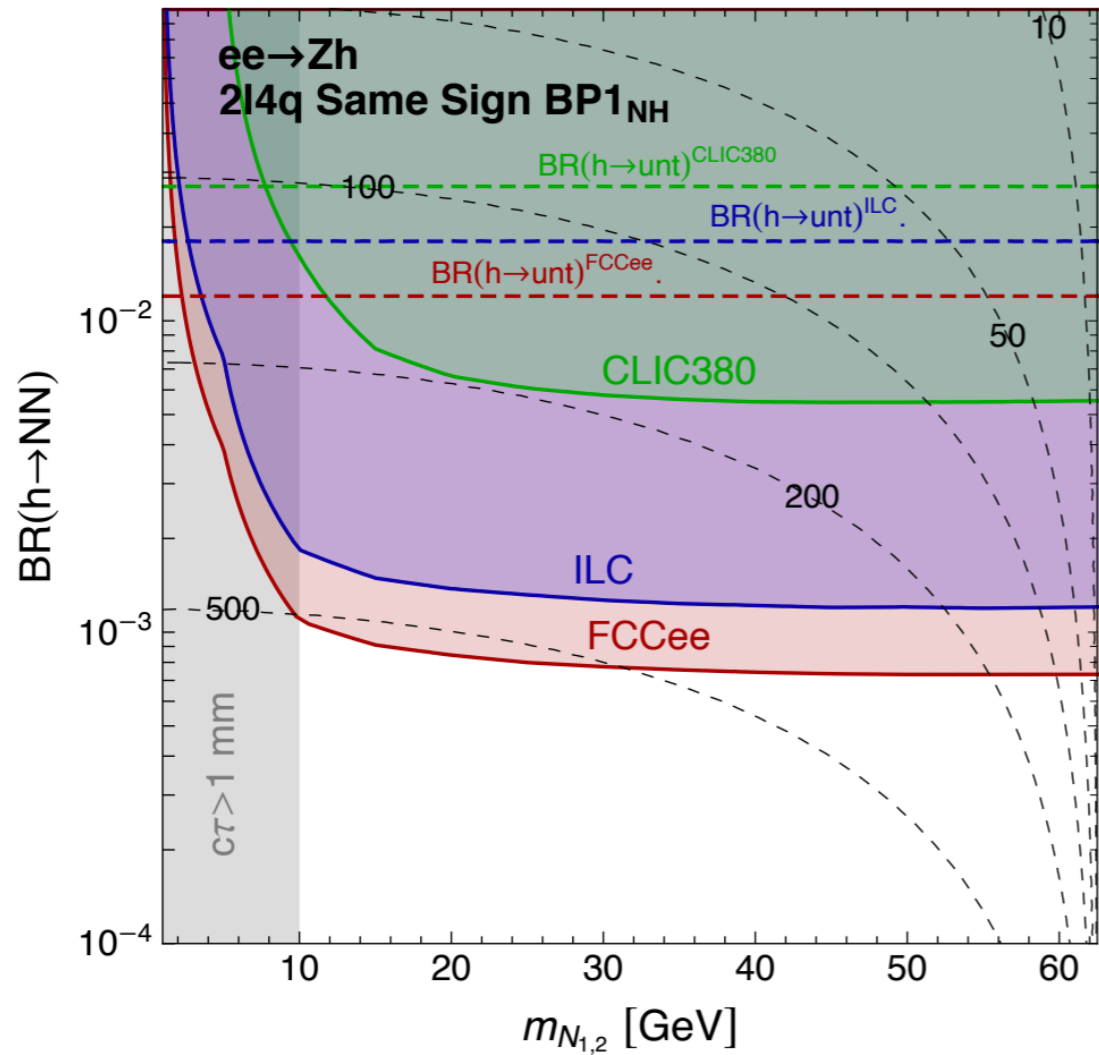
$$\mathcal{O}_{NH} = \alpha_{NH} (\bar{N}^c N) (H^\dagger H) ,$$

$$\mathcal{O}_{NB} = \alpha_{NB} \bar{N}^c \sigma^{\mu\nu} N B_{\mu\nu} ,$$

$$\Gamma(h \rightarrow \bar{N}_i^c N_i) = \frac{1}{2\pi} \frac{v^2}{\Lambda^2} m_H \beta_N^3 (\alpha_{NH}^{ii})^2 ,$$

$$\beta_N = \sqrt{1 - \frac{4m_N^2}{m_H^2}} .$$

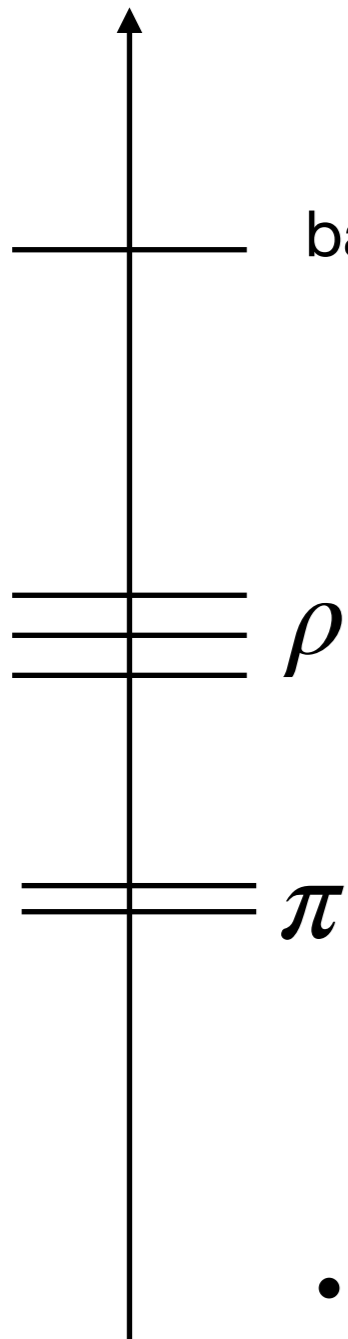
# Sensitivity estimates



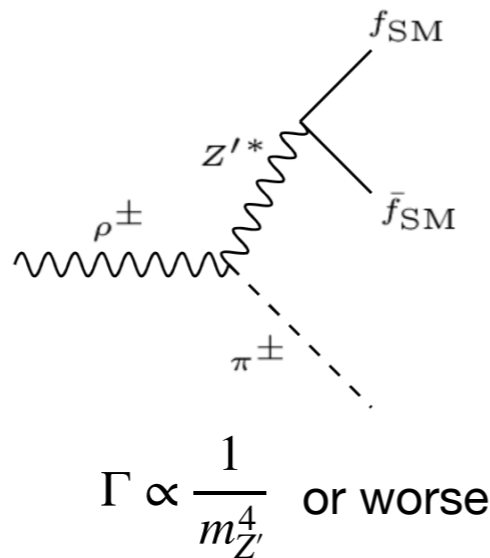
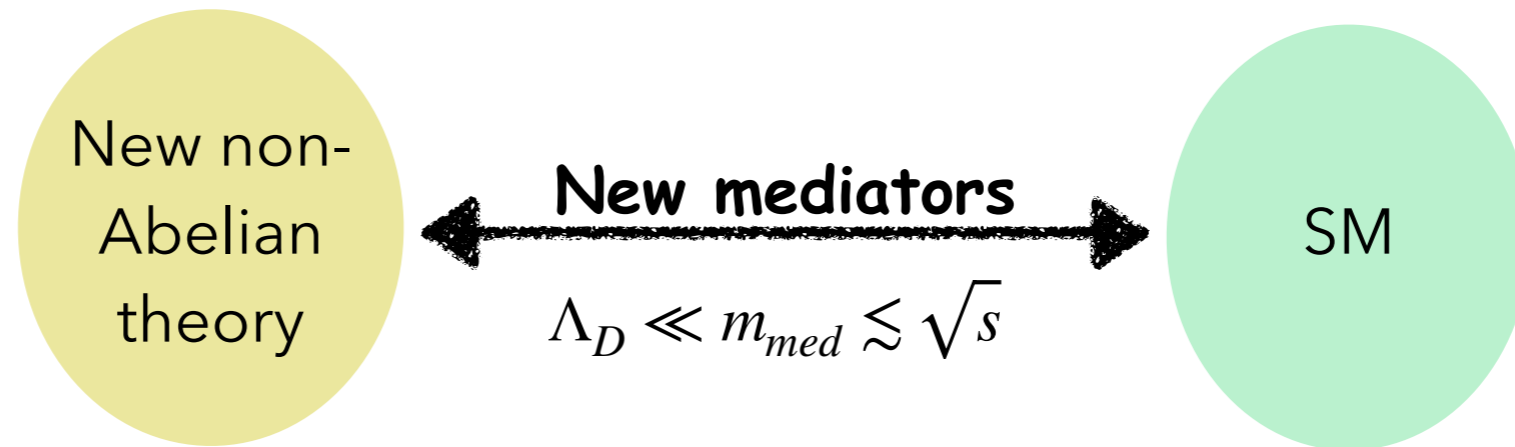
- Two benchmarks corresponding to different SM neutrino flavour mixing patterns
- Dashed lines correspond to the cutoff scale lambda
- Promising limits
- Favourable scenarios for cut-off scale above 500 GeV

# Dark matter

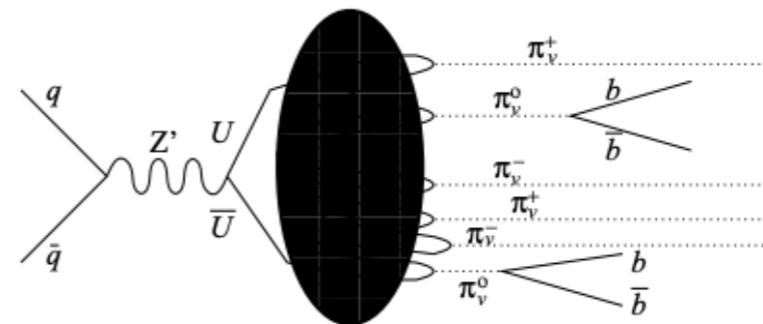
Spectrum



- Dark matter can result from confinement under new non-Abelian sector
- Resulting low energy (IR) spectrum can lead to pions or baryons as dark matter candidates depending on the details



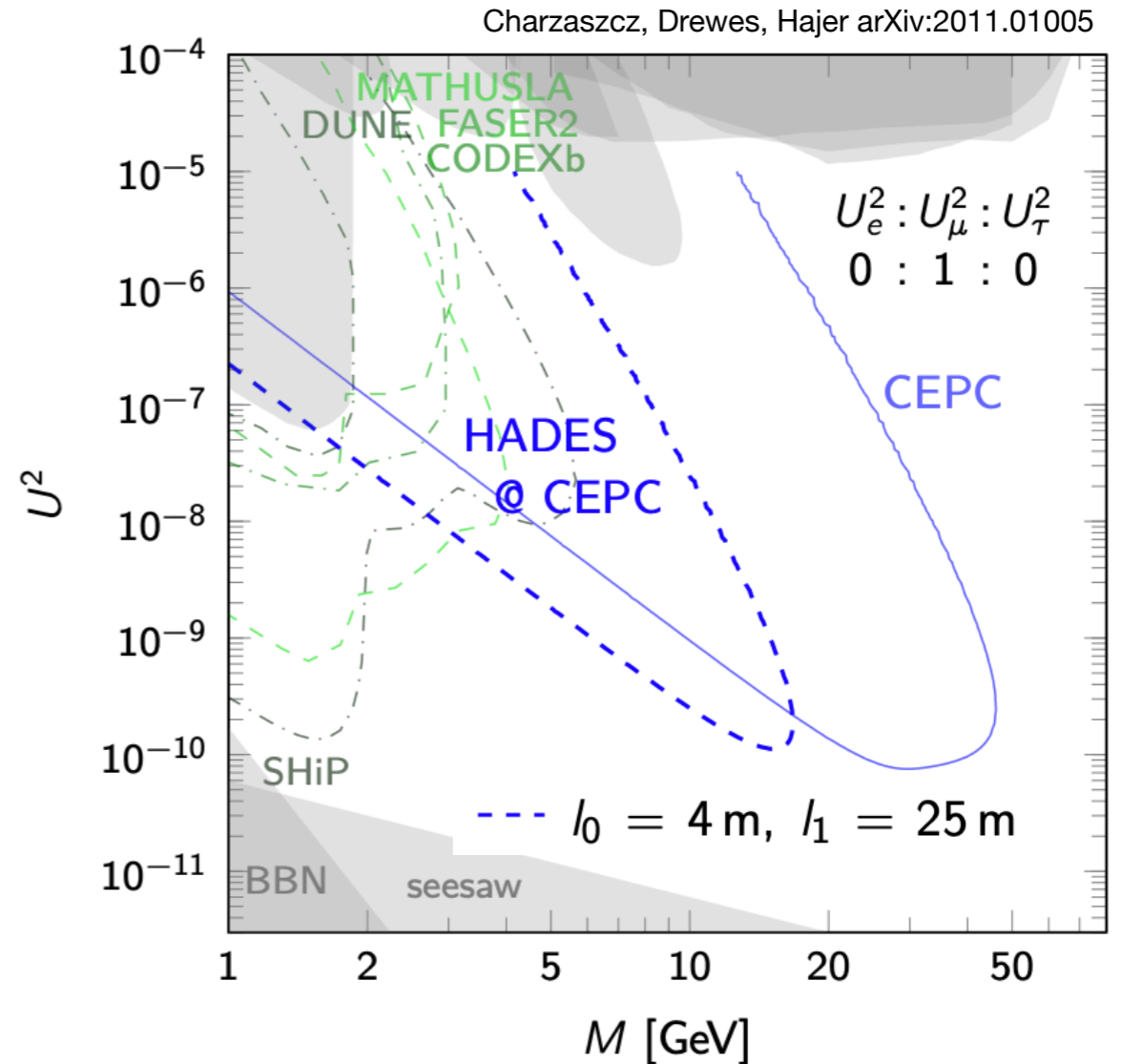
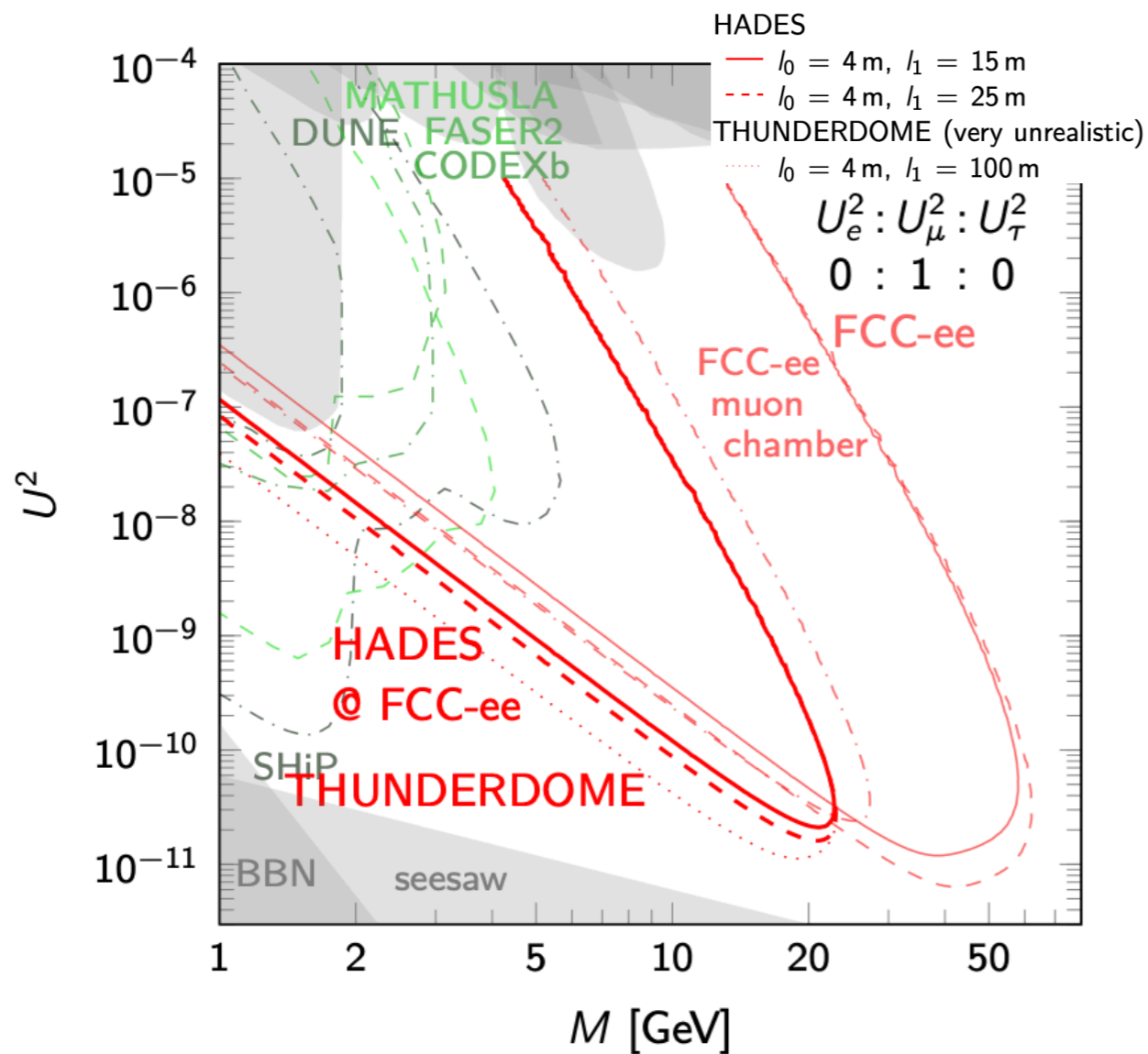
$$\Gamma \propto \frac{1}{m_{Z'}^4} \text{ or worse}$$



hep-ph/0604261;  
arXiv:1503.00009;  
arXiv:1502.05409;  
arXiv:1903.04497;  
arXiv:2203.09503  
For fcc specific study see:  
arXiv:2005.05221,  
arXiv:1501.05310

- LLPs arise generically and easily, 'hide' in SM jets, how to best look for them at FCC-ee and hh?

# Beyond 'traditional' detectors: FCC-ee

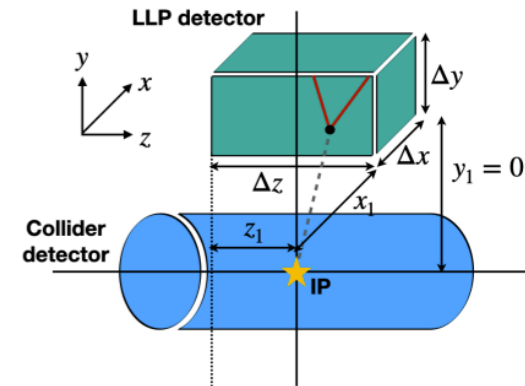
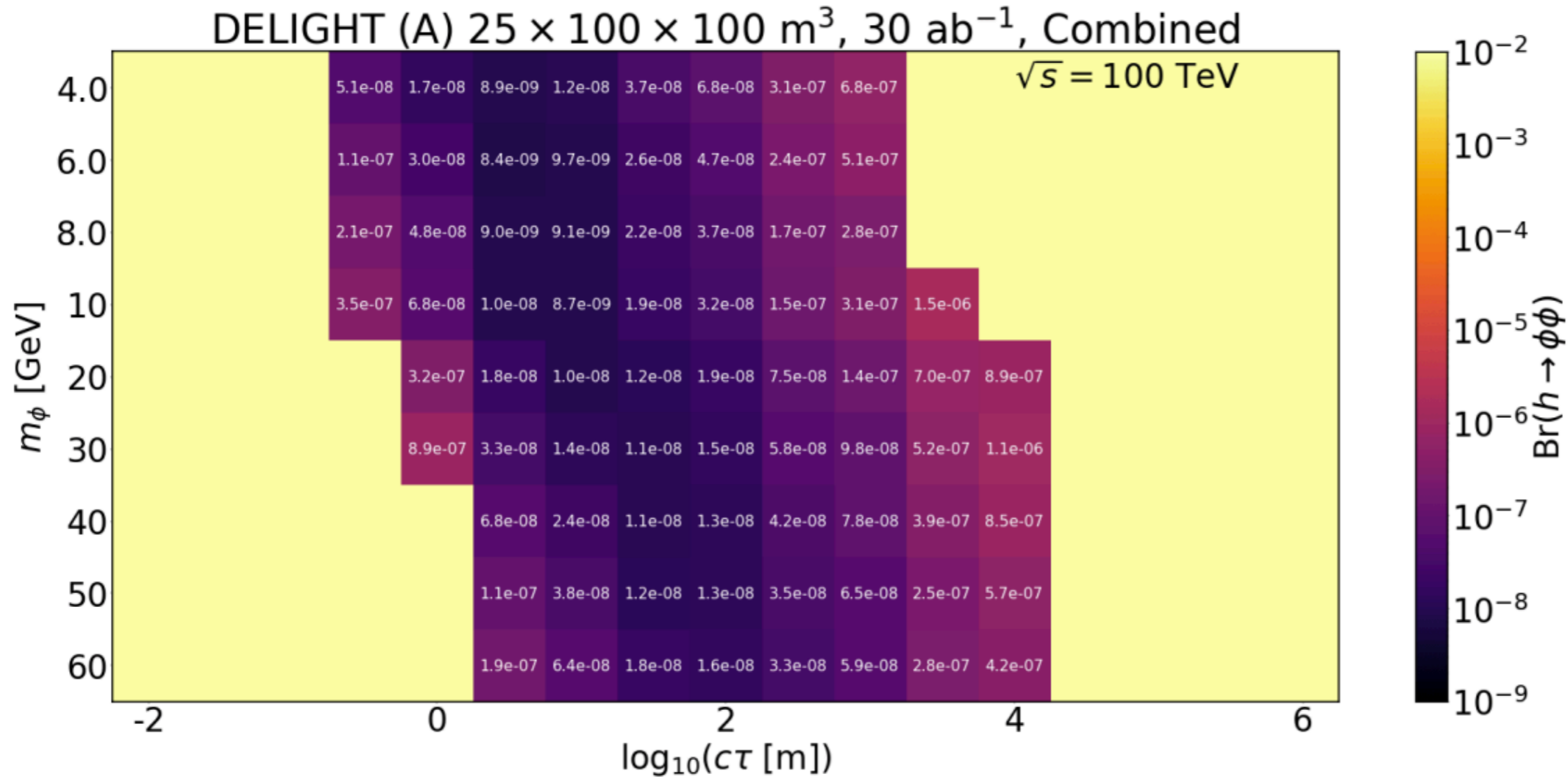


- Could we build/propose lifetime frontier/forward physics detectors at lepton colliders?
- HADES/HECATE detector proposes to use the cavern to build a new lifetime frontier detector
- Unique opportunity for  $4\pi$  solid angle coverage
- Can help improve reach of LLPs, here case study of HNL



# Beyond 'traditional' detectors: FCC-hh

Bhattacharjee, Matsumoto, Sengupta arXiv:2111.02437



$x = 25 \text{ m}; \eta = 0$

See also  
Boyarsky, Mikulenko,  
Ovchynnikov, Shchutska  
arXiv:2204.01622

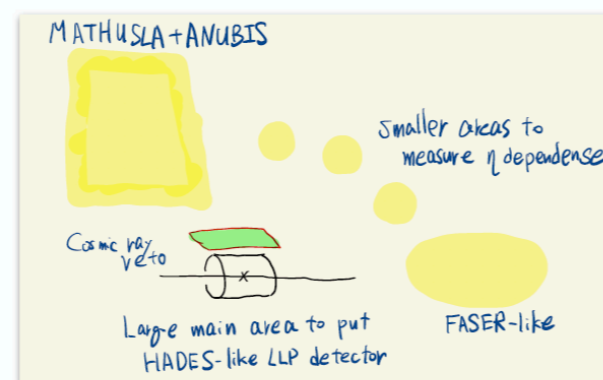
- Detector for long-lived particles at high energy of 100 TeV
- Same dimensions as MATHUSLA ( $25 \times 100 \times 100 \text{ m}^3$ ); transverse detector
- Has ca 500 times more sensitivity; 150 due to increased cross section and luminosity and factor 3-4 due to moving detector closer to IP

# Reality check

- All the studies I showed so far are somehow apples to oranges comparison
- Not all contours are done on same footing, not all studies done with realistic setup
- Good enough to get first estimate, but we need to have realistic expectations
- Can we do better?

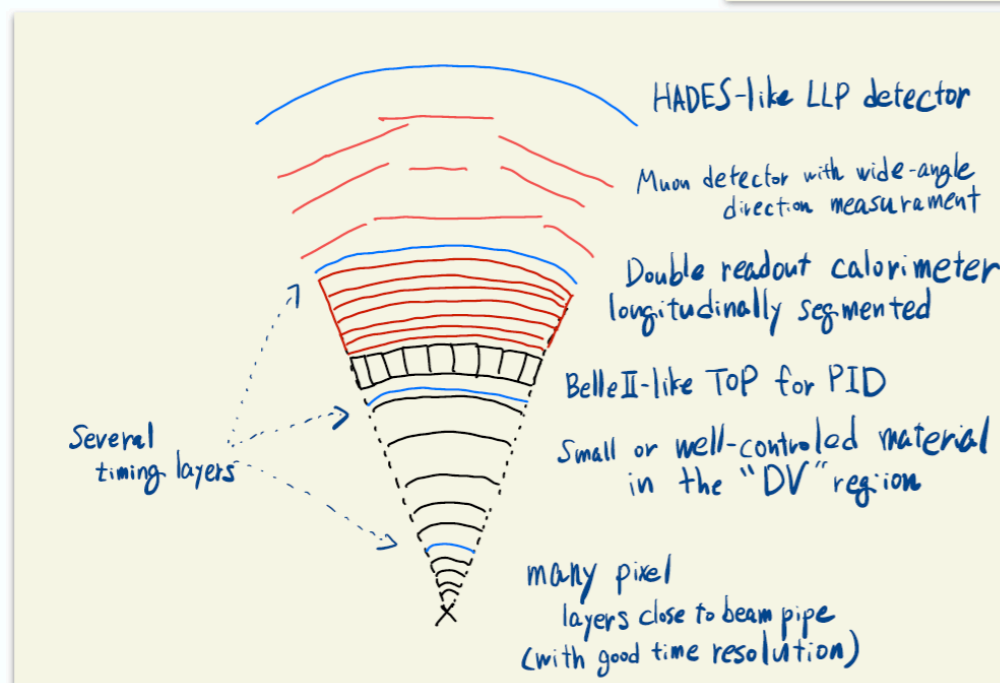
See talk by M. McCullough

## A dream LLP detector?



It is a good time to plan our **dream LLP detectors**, following Ryu Sawada's first example at the LLP workshop in November-2020 ([link](#))

See talk by Juliette Alimena



# Conclusions

- We want to make the best use of the available experimental data to learn about fundamental laws governing the Universe
- Searching for LLPs is one way of chasing that quest
- LLPs arise in a variety of BSM scenarios in solutions of problems from nature of dark matter to the Higgs hierarchy problem
- LLPs lead to a variety of final states, and demand new analysis techniques and probably improved/additional detectors
- FCC-ee and hh offer complementary information and are capable for delivering a more detailed understanding of new physics via LLP
- Similar considerations should also be given for FCC-eh physics (not covered in the talk)
- Most of the studies so far are done under simplistic assumptions, more realistic detector simulations are necessary for setting up our expectations

Thanks for the discussions A. Thamm, R. Gonzalez-Suarez, J. Alimena, M. Drewes, F. F. Deppisch, R. Sengupta