

# CAI and GRICE studies

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Franck Pereira Dos Santos

LNE-SYRTE

*Observatoire de Paris, Université PSL, CNRS, Sorbonne Université,  
61 avenue de l'Observatoire, 75014 Paris, France*

<https://syрте.obsрm.fr/spip/science/iaci/>

# Outline

## I. Introduction

## II. CAI gradiometer study

## III. GRICE Study

Conclusion & Prospects

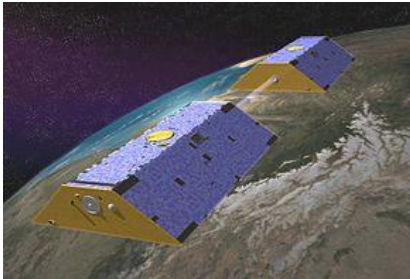
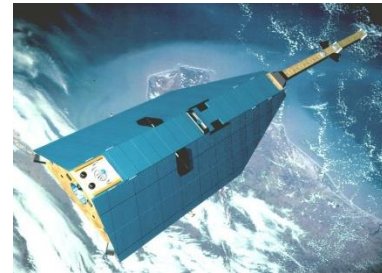
# I. Gravity missions

## General principle:

Precise tracking position of a satellite (in free fall) + measurement of non-inertial acceleration

## Different mission scenarios:

- CHAMP : Orbitography + accelerometry



- GRACE : Twin satellites, Orbitography + SST (Microwave link) + accelerometry

- GOCE : 1 satellite, Orbitography + gradiometry

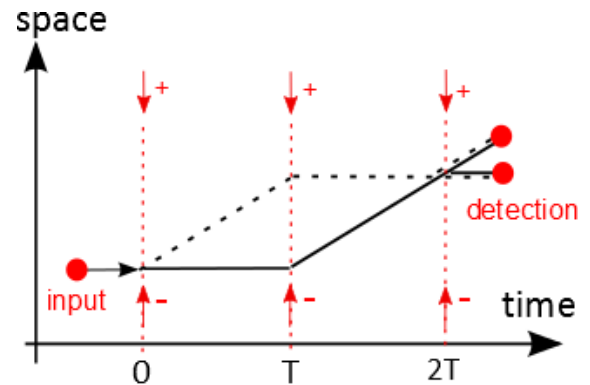


- GRACE FON : GRACE + SST with LRI

All missions rely on electrostatic accelerometers, which suffer from bias and drifts at low frequency

# I. Atomic accelerometer

- Relies on atom interferometry methods
- Sequence of three laser beamsplitters to create an « atomic » Mach Zehnder Interferometer
- Allows creating inertial sensors (accelerometers, gyrometers ...)



The interferometer phase is related to the interaction between atoms and light beam splitters:  
At each pulse, the laser phase gets imprinted onto the atomic wavepacket

$$\begin{cases} \Delta\Phi = \varphi(0) - 2\varphi(T) + \varphi(2T) \\ \varphi(t) = -\vec{k} \cdot \vec{z}(t) \end{cases}$$



$$z(t) = \frac{1}{2} a t^2 \Rightarrow \Delta\Phi = k a T^2$$

$a$  measurement = measure of the relative displacement atoms/laser equiphases

Free falling atoms = « Inertial reference » & Lasers = Very fine, stable, accurate ruler

On Earth : interaction time limited by free fall,  $2T \sim 0.1 \text{ s} \Rightarrow$  sensitivity  $\sim 10^{-8} \text{ m/s}^2/\text{Hz}^{1/2}$   
In Space : interaction time limited by residual expansion of the atoms/instrument size

**$2T = 10 \text{ s} \Rightarrow$  sensitivity in the  $10^{-12} \text{ m/s}^2/\text{Hz}^{1/2}$  range**

# I. Comparison

	<b>Atomic Accelerometer</b>	<b>Electrostatic Accelerometer</b>
Sensitivity	$10^{-12}$ m/s <sup>2</sup> /Hz <sup>1/2</sup> White Noise (potential)	$10^{-12}$ m/s <sup>2</sup> /Hz <sup>1/2</sup> MBW (demonstrated)
Scale factor	Absolute	Calibration required
Stability	No drift	Drift
Measurement capability	Single axis	Three axis
Proof mass motion	Residual velocities ⇒ Coriolis acceleration	-
SWaP	High	Low
TRL	Intermediate	High

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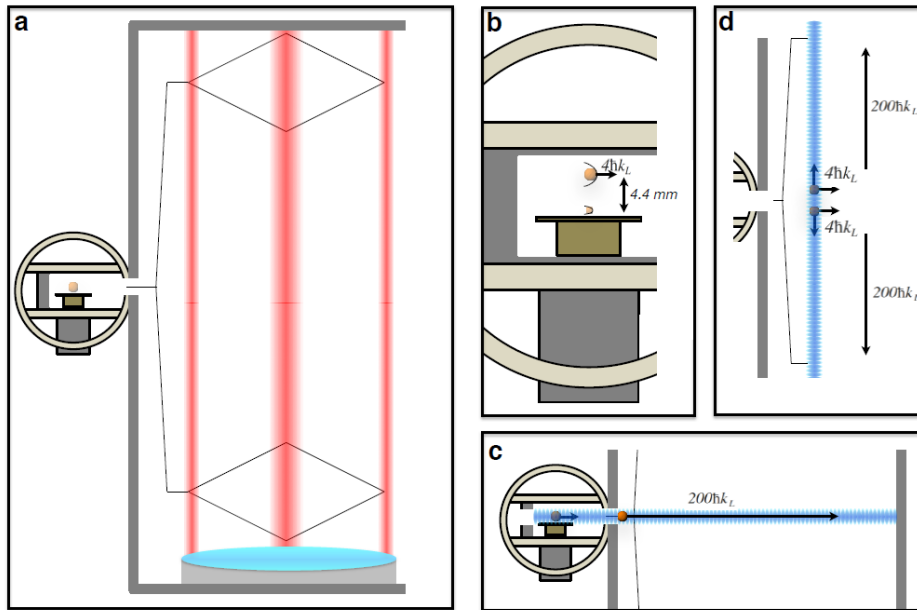
III. GRICE study

Conclusion & Prospects

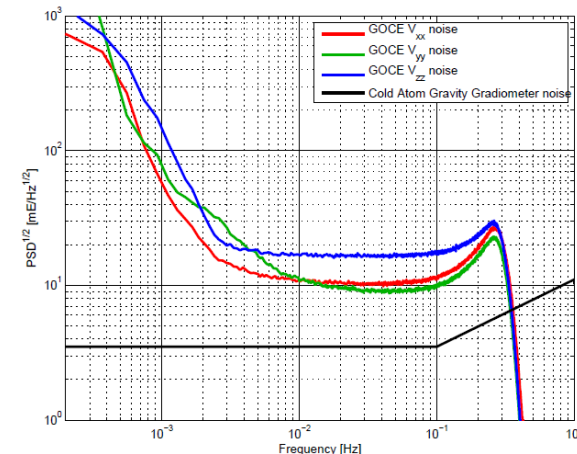
# II. CAI gradiometer Study



Builds on the concept proposed in Carraz et al., Microgravity Science and Technology 26, 139 (2014)



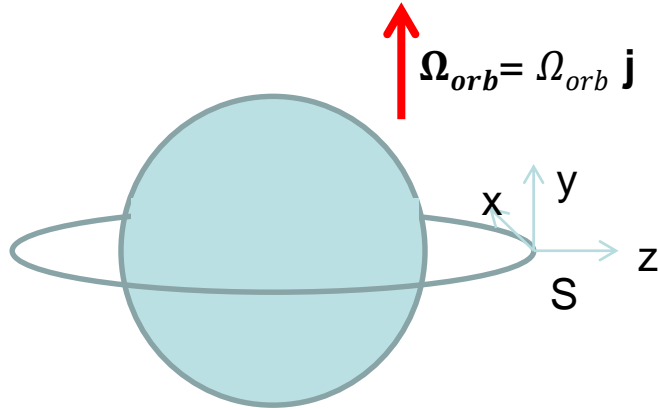
- Atom chip production of ultracold atoms  
 $T_{at}=100$  pK
- Transport and splitting of a single source to feed two interferometer inputs
- Two interferometers separated by 50 cm  
Interferometer duration  $2T = 10$  s



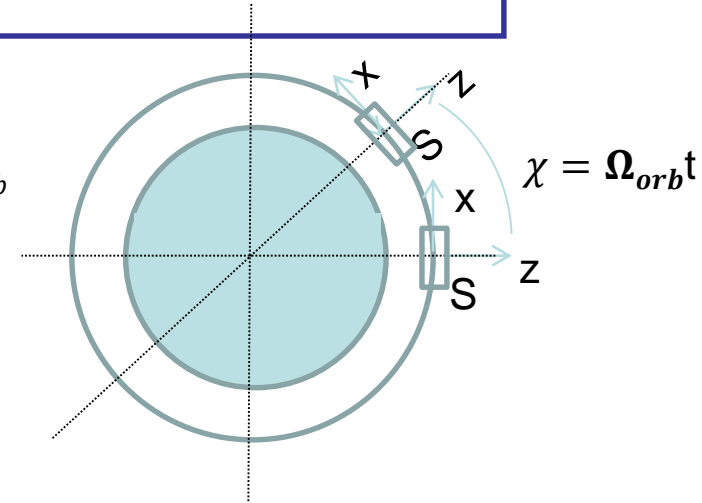
Expected measurement noise :  $3.5 \text{ mE/Hz}^{1/2}$

Competitive with GOCE gradiometers

## II. Nadir mode



Nadir pointing:  $\Omega_y \sim \Omega_{orb}$



### 1) Case where the measurement axis is in the orbital plane (along x or z)

Single interferometers have **additional Sagnac phase** terms :  $k_x v_z \Omega_y T^2$  and  $k_z v_x \Omega_y T^2$

Loss of contrast due to the phase dispersion  $\sigma_{\Phi_S} \leq 1 \text{ rad} \Rightarrow \sigma_v \leq \frac{1}{2k\Omega T^2}$

$$\Omega_y = \Omega_{orb} \sim 10^{-3} \text{ rad/s} \Rightarrow \sigma_v < 0.6 \text{ } \mu\text{m/s} \Leftrightarrow T_{at} < 4 \text{ fK}$$

**$\Rightarrow$  No contrast for measurement axes in the orbital plane (along x or z)**

### 2) Case where the measurement axis is aligned with orbital rotation vector

Single interferometers have no Sagnac phase scaling as  $\Omega_y$

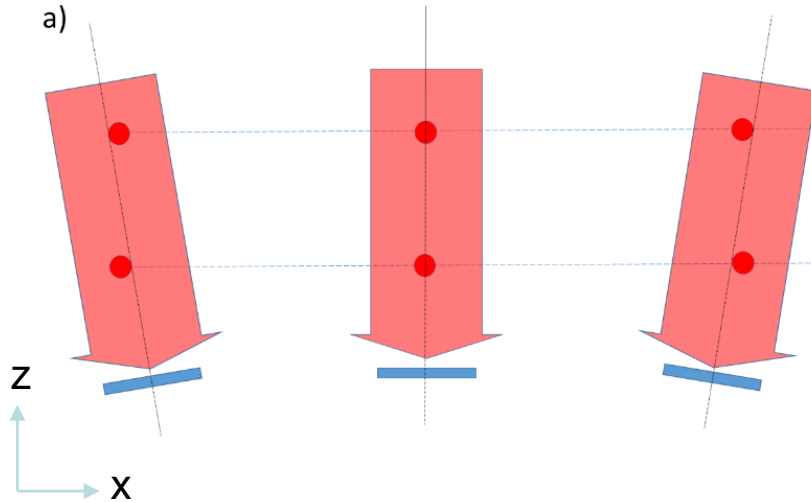
Requirements:  $\Omega_x, \Omega_z < 10^{-6} \text{ rad/s}$



## II. Coping with rotations

**Idea: Tilt the mirrors so as to compensate for the Orbital rotation rate**

Angle given by  $\Omega_m T = \Omega_y T$  where  $\Omega_m$ : equivalent rotation rate of the mirrors



### Eliminates

- Coriolis accelerations and the induced dephasing
- Centrifugal acceleration

Residual centrifugal acceleration still impose a bias on the gravity gradient:  $2 \Omega_y (\Omega_{orb} - \Omega_y)$   
To keep this term below 1mE, one needs a knowledge of  $\Omega_{orb} - \Omega_y$  at the  **$10^{-9}$  rad/s level**, which is challenging

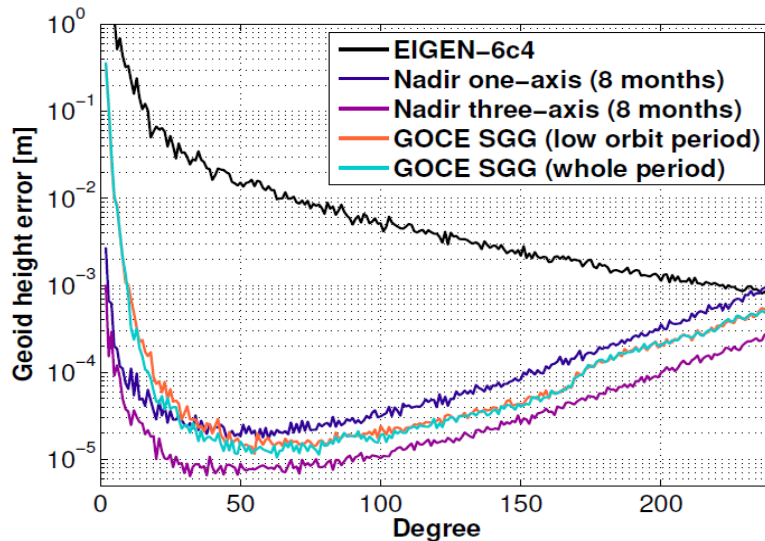
# II. Results

Reconstruction of the Earth gravity field out of synthesized noisy signals

Decomposition on the spherical harmonics (l is the degree, m the order)

$$V(r, \theta, \lambda) = \frac{GM}{R} \sum_{l=0}^{\infty} \left(\frac{R}{r}\right)^{l+1} \sum_{m=-l}^{m=l} K_{lm} Y_{lm}(\theta, \lambda)$$

Simulation performed by IFE, includes both instrument noise and attitude noise



**GOCE**

Whole period :  
47 months

Low orbit period :  
3 months at 239 km  
+ 5 months below 239 km

⇒ Clear improvement with respect to GOCE

A. Trimeche, B. Battelier, D. Becker, A. Bertoldi, P. Bouyer, C. Braxmaier, E. Charron, R. Corgier, M. Cornelius, K. Douch, N. Gaaloul, S. Herrmann, J. Müller, E. Rasel, C. Schubert, H. Wu, F. Pereira dos Santos, "Concept study and preliminary design of a cold atom interferometer for space gravity gradiometry", arXiv:1903.09828

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# III. GRICE study

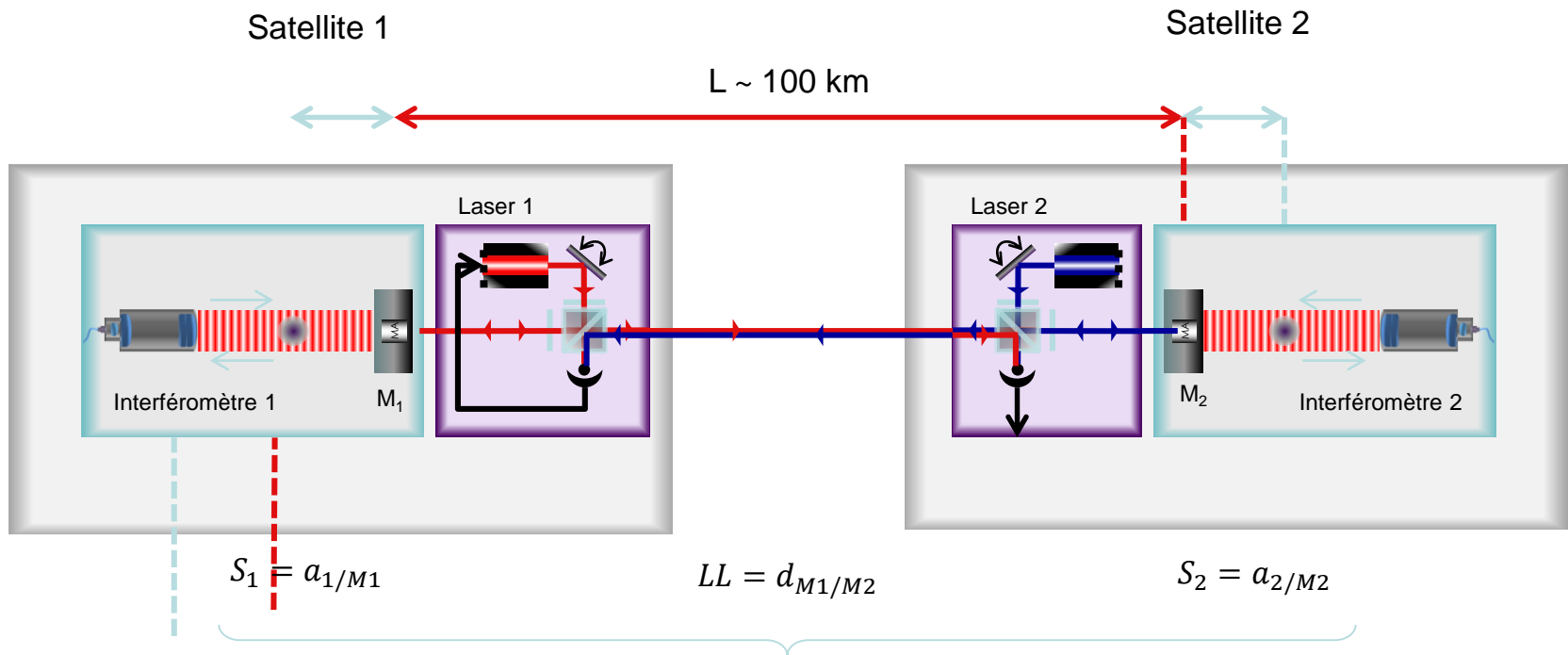
GRICE: a Phase 0 study carried out by CNES, with the participation of (french) scientists  
 End user requirements + Review of mission concepts + Analysis of a chosen concept

Mission and instrument concept

*Sheng-wey Chiow, Jason Williams, and Nan Yu, PRA 92, 063613 (2015)*

Twin satellites with onboard quantum accelerometers and laser ranging

Instrument concept: double diffraction Raman interferometer (as in the CAI study, but T lower)

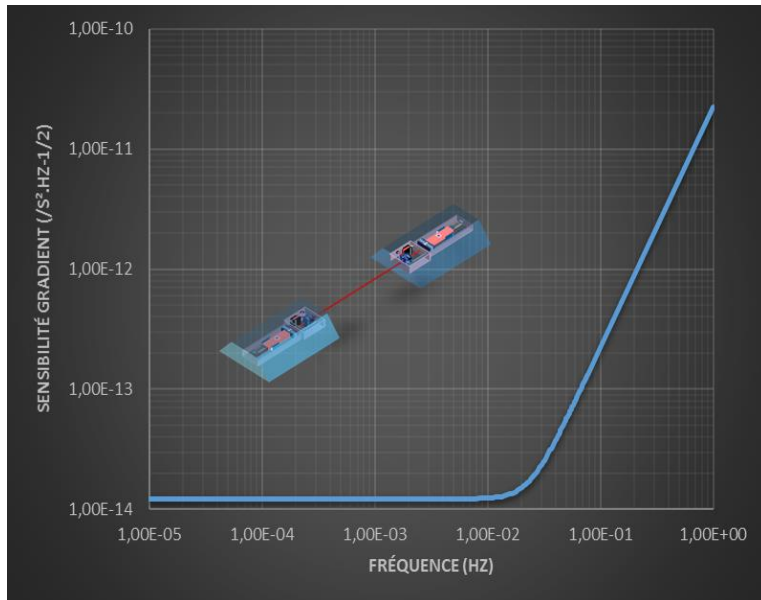


$$\frac{\partial a_x}{\partial x} = \frac{1}{d} (S_1 - S_2 + \ddot{L})$$

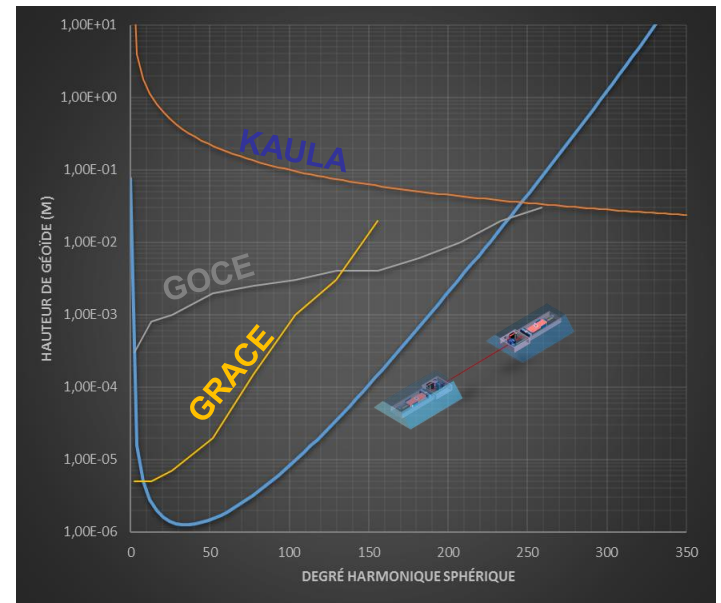
**Sensitivity :  $\sim 10 \mu\text{E} \cdot \text{Hz}^{-1/2}$**

# III. GRICE study

## Instrument performances



## Expected mission performances

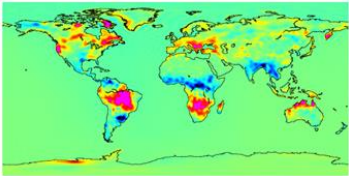


Parameter	Notation	Value	Unit
Noise LL	$\sigma_l$	$40 \times 10^{-9}$	$m/Hz^{1/2}$
Noise AI	$\sigma_a$	$6 \times 10^{-10}$	$m/s^2/Hz^{1/2}$
Baseline	L	100	km

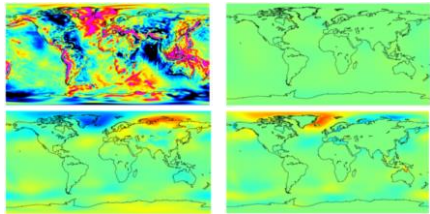
# III. GRICE study

## Mission performances simulation

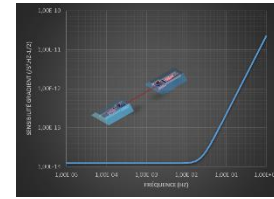
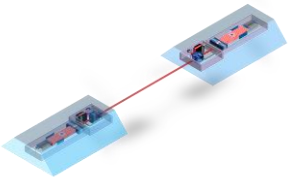
Signal : Hydrology



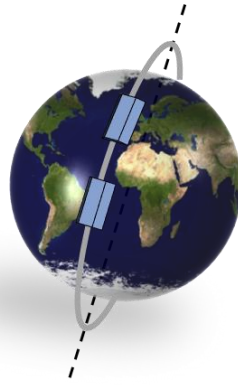
Environmental models



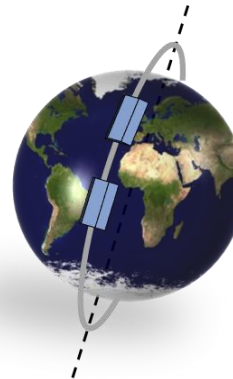
Instrumental measurements



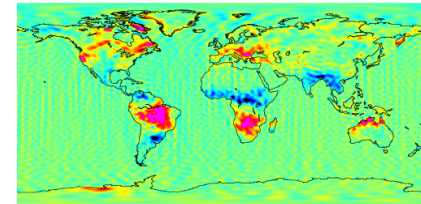
Synthetic satellite data



Noise introduction



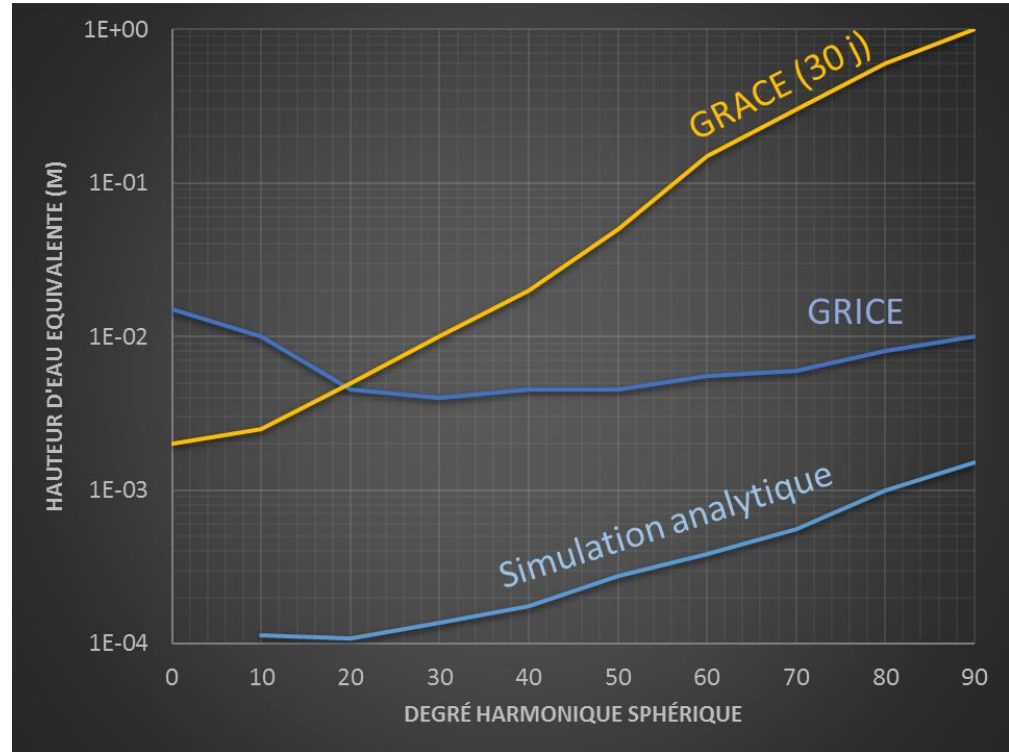
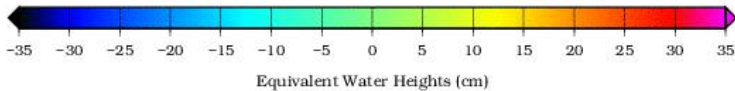
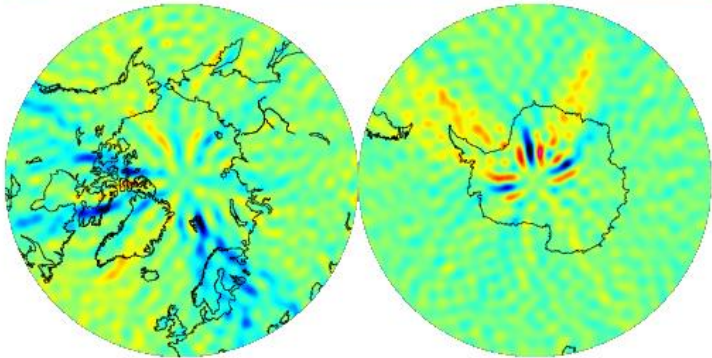
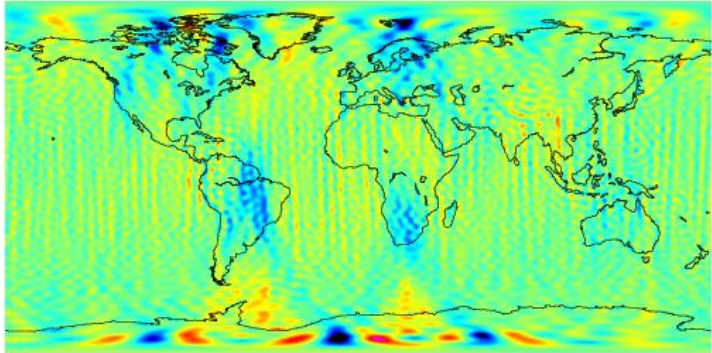
Gravity field restitution



Introducing noise in the models for hydrology and environment introduces additional errors in the restitution

# III. GRICE study

## Restitution error

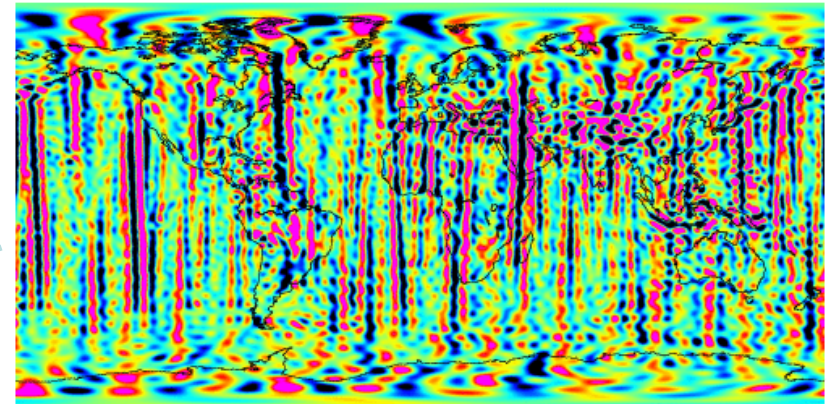
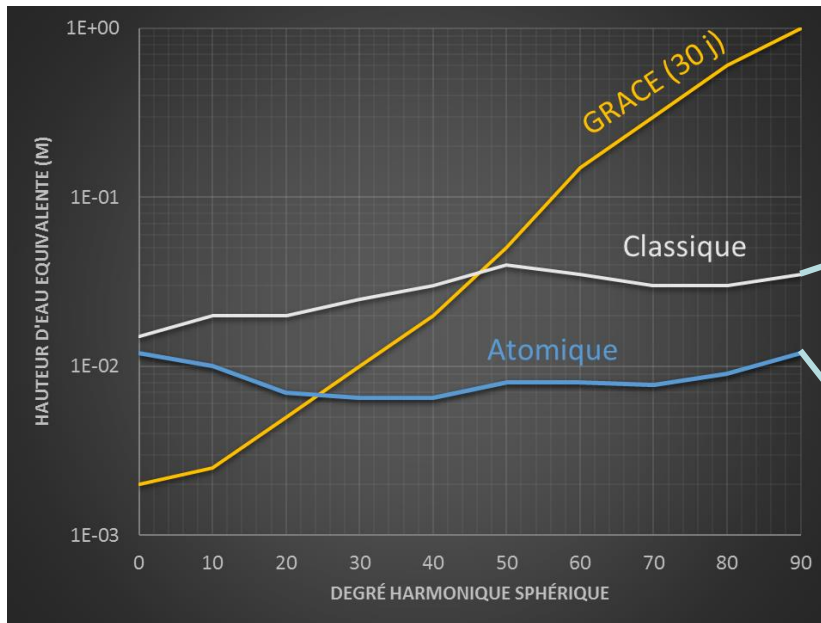


- Aliasing from « background » models degrade the performance
- What is then the impact of the accelerometer performance ?

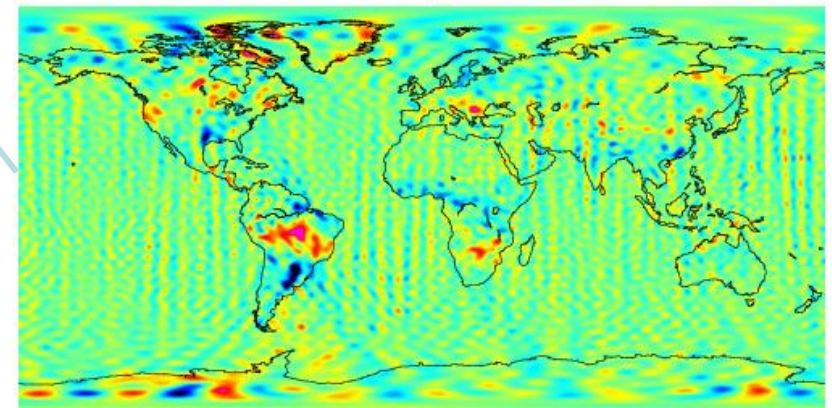
# III. GRICE study

## Quantum vs Classical technologies

Comparison between restitution errors with different accelerometers applying the same data processing (classical accelerometer has a « typical » drift + no empirical parameters to reduce the impact of the drift)



Classical



Quantum

- Less stripes for the quantum sensor
- No (less) need for data reduction that modify the gravity field



# Conclusion and Prospects

- **Atomic accelerometers provide better long term stability**  
(ie lower measurement noise at the lowest frequencies, below 10 mHz)
- **Allows improving the reconstruction of the Earth gravity field**  
(at low degrees especially, if not all degrees)
- **Influence of the variable field to be taken into account**  
Degrades the quality of the reconstruction  
*See also Abrykosov et al., Advances in Space Research 63, 3235 (2019)*
- **But a lower measurement noise will allow for**
  - better accounting for fluctuations/noise arising from aliasing of the variable field
  - improving the modelling of the variable field
- **Atomic sensor technology is under maturation**  
R&D and engineering developments still needed
- **Refined studies are required**