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# Sexaquark: an unexpected candidate for dark matter ?

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on behalf of

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# Outline

**I Introduction**

**II uuddss recent searches**

**III Thermal model results on uuddss**

**IV Summary, Conclusions and Outlook**



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# I Introduction



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# H-uuddss

The uuddss dibaryon state has been initially proposed by Robert Jaffe (MIT) who made a bag model based calculations and predicted the so called H-dibaryon with a mass below  $2m(\Lambda)$  of about 2150 MeV and unstable.

R. L. Jaffe, Phys. Rev. Lett. **38**, 195 (1977)

The H dibaryon has been searched by several experiments without finding it.

B. H. Kim et al. (Belle), Phys. Rev. Lett. **110**, 222002 (2013),  
arXiv:1302.4028 [hep-ex].

J. Badier et al. (NA3), Z. Phys. **C31**, 21 (1986).

R. H. Bernstein, T. K. Shea, B. Weinstein, R. D. Cousins, J. F. Greenhalgh, M. Schwartz, G. J. Bock, D. Hedin, and G. B. Thomson, Phys. Rev. **D37**, 3103 (1988).

J. Belz et al. (BNL-E888), Phys. Rev. Lett. **76**, 3277 (1996),  
[Phys. Rev. C56, 1164 (1997)], arXiv:hep-ex/9603002 [hep-ex].

A. Alavi-Harati et al. (KTeV), Phys. Rev. Lett. **84**, 2593 (2000),  
arXiv:hep-ex/9910030 [hep-ex].

H. Gustafson et al., Phys. Rev. Lett. **37**, 474 (1976).



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# S-uuddss

Recently a new possibility for the uuddss multiquark state has been proposed by Glennys Farrar (NYU) in which the uuddss is a state with small radius (0.1-0.4 fm) and mass below 2 GeV (so called Sexaquark).

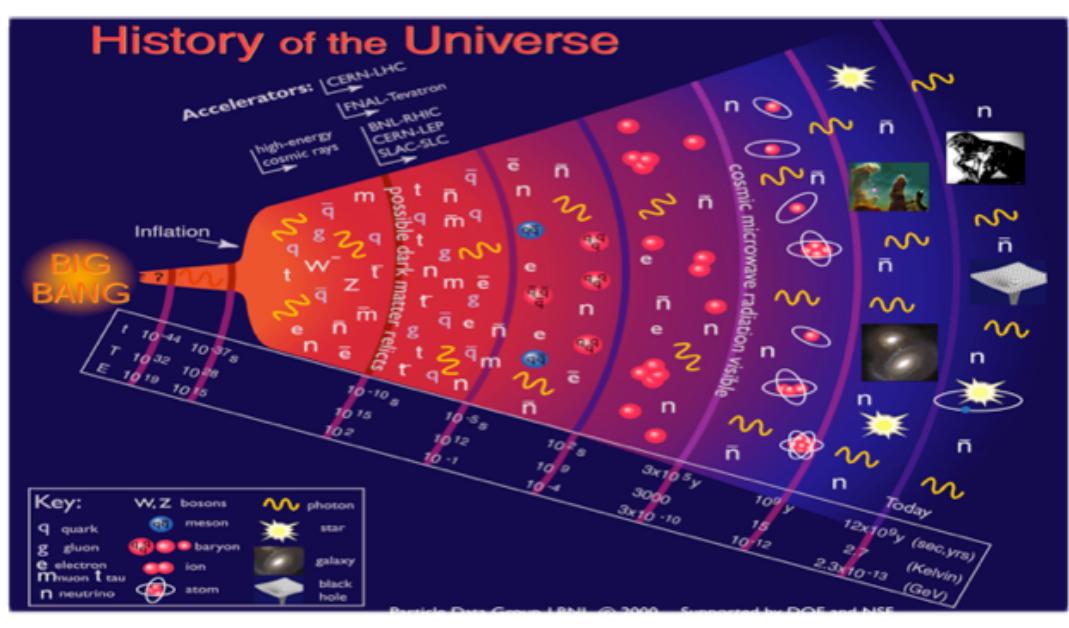
Such a state could exist and could have escaped the experimental searches till now, since (depending on its mass) it can be absolutely stable or have a lifetime of the order of the age of the universe.

Such a stable state could be a candidate for Dark Matter, if produced out of the primordial Quark Gluon Plasma at the QCD phase transition from partons to hadrons.

G. R. Farrar, (2017), arXiv:1708.08951 [hep-ph] and  
G. R. Farrar, (2018), arXiv:1805.03723 [hep-ph]



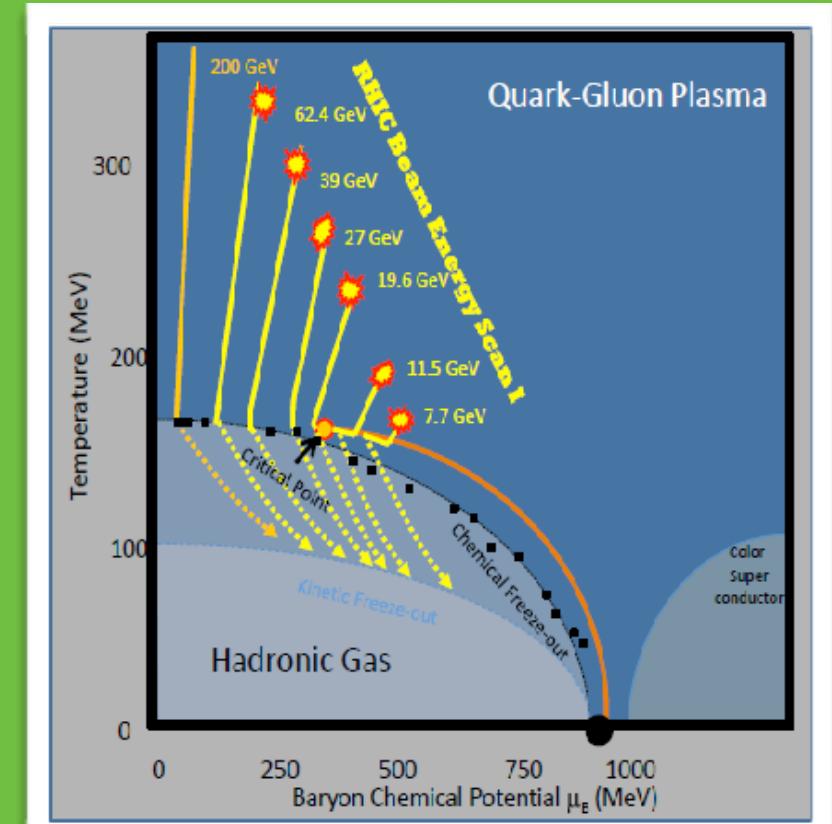
# The QCD phase transition



The transition from quarks and gluons to hadrons is believed that took place few  $10^{-6}$  sec after the Big Bang

Lattice QCD predicts at zero baryon density a crossover with a pseudocritical temperature  $T_c \sim 156.5 \pm 1.5$  MeV

(A. Bazanov et al, Phys. Lett. B **795**, 15-21 (2019).)





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# S-uuddss

The uuddss state can be maximally bound due to its symmetry.

Due to being a flavor singlet it does not bind to pions resulting in a compact configuration.

Assuming it can bind to lightest flavor singlet mesons like the f0 a radius of 0.1-0.3 fm is estimated.

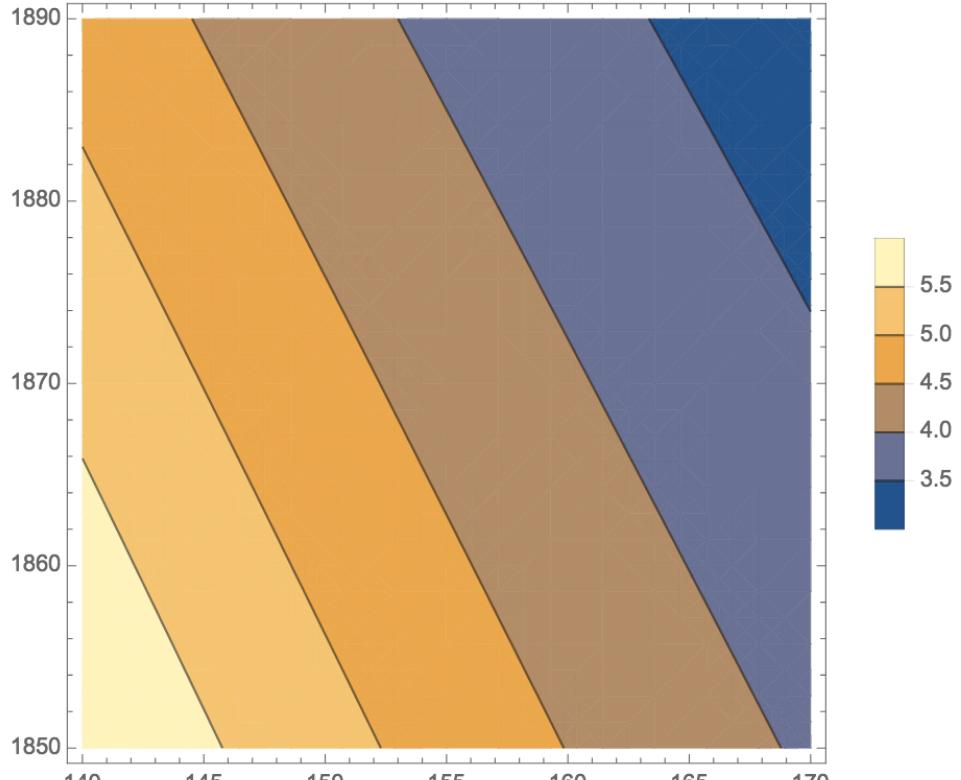
The different size of the S and baryons means that amplitudes involving the S and 2 baryons are strongly suppressed.

Lacking coupling via pions its interaction with matter is lower than that of ordinary hadrons supporting the hypothesis it can be a DM candidate.

G. R. Farrar, (2017), arXiv:1708.08951 [hep-ph] and  
G. R. Farrar, (2018), arXiv:1805.03723 [hep-ph]

# S-uuddss

**Calculation of the ratio of Dark Matter (DM) to Ordinary Matter (OM) assuming production of S(uuddss) from the Quark Gluon Plasma at the early Universe at a temperature of 140-170 MeV is shown in figure below**



y-axis:  $m(\text{Sexaquark})$  in MeV  
x-axis: Freeze out Temperature

The measured value for DM/OM of  $5.3 \pm 0.1$  is in the second from lightest band.

Therefore the assumption of the Sexaquark as DM candidate leads to agreement with the measured DM/OM ratio for masses below 1885 MeV and at freeze out T below 153 MeV.

However all values of the entire plot are within factor of 2 from the measured value



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# S-uuddss

**The previous calculation of the ratio of Omega Dark Matter to Omega Ordinary Matter assuming a Sexaquark as DM candidate is valid for all forms of DM including equal u,d,s parts (like stable sexaquarks, quark nuggets or primordial black holes) or a combination of them.**

**As a result the observed value of  $\Omega_{\text{DM}}/\Omega_{\text{Matter}}$  has been reproduced by assuming that DM is composed of equal number of u,d,s quarks**

**G. R. Farrar, (2017), arXiv:1708.08951 [hep-ph] and G. R. Farrar, (2018), arXiv: 1805.03723 [hep-ph]**



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# II uuddss recent searches



# Methods

G. R. Farrar, (2017), arXiv:1708.08951 [hep-ph]

One proposed method to observe the S-uuddss is via the Upsilon decay in Upsilon factories, namely as missing mass due to the S or anti-S production in association with di-anti-Lambda or di-Lambdas

$$\Upsilon \text{ [} \rightarrow \text{ gluons] } \rightarrow S \bar{\Lambda} \bar{\Lambda} \text{ or } \bar{S} \Lambda \Lambda + \text{ pions and/or } \gamma$$

Other possibility that can be used is the following:

$$\begin{aligned} \bar{S} + N &\rightarrow \bar{\Xi}^{+,0} + X, \text{ with } \bar{\Xi}^{+,0} \rightarrow \bar{\Lambda} \pi^{+,0} \& \bar{\Lambda} \rightarrow \bar{p} \pi^+ \\ \text{or } \bar{S} + N &\rightarrow \bar{\Lambda} + K^{+,0} + X. \end{aligned} \quad (2)$$



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# Kolb and Turner

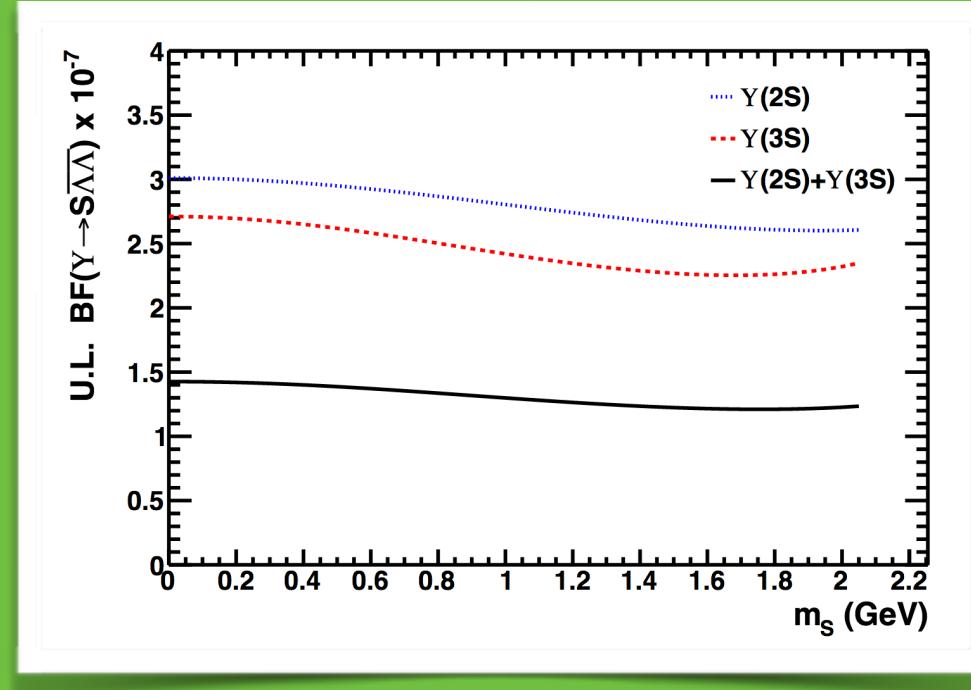
**Kolb and Turner argue that the Sexaquark abundances will freeze out at about Temperature of 10 MeV following decrease in number of Lambdas and very small abundance of S-uuddss remain at that Temperature.**

E. Kolb, M. Turner, Phys. Rev. D99 (2019) no.6, 063519

# Experimental searches

Even though not everyone agrees its possible cosmological implications as DM candidate cannot be excluded and it has been recently searched in the BaBar experiment that set upper limits .

BABAR Coll. J. P. Lees et al, Phys.Rev.Lett. 122 (2019) no.7, 072002



90% conf. level upper limits on the branching fraction  
 $\Upsilon(2S,3S) \rightarrow S + \bar{\Lambda} + \bar{\Lambda}$   
as well as the combined sample



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# Experimental searches

BABAR Coll. J. P. Lees et al., Phys. Rev. Lett. 122 (2019) no.7, 072002

In conclusion, we performed the first search for a stable  $uuddss$  configuration in  $\Upsilon$  decays. No signal is observed, and 90% CL limits on the combined  $\Upsilon(2S, 3S) \rightarrow S\bar{A}\bar{A}$  branching fraction of  $(1.2 - 1.4) \times 10^{-7}$  are derived for  $m_S < 2.05$  GeV. These results set stringent bounds on the existence of a stable, doubly strange six-quark state.

G. R. Farrar, (2017), arXiv:1708.08951 [hep-ph] :

**The predicted inclusive branching fraction of S in Upsilon decays is of the order  $\sim 10^{-7}$ . and the exclusive to inclusive ratio for Upsilon decay is typically  $\sim 10^{-4}$ , so no exclusive signal would have been expected at the level of sensitivity of BABAR which is  $10^{-7}$  in the exclusive channel.**





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# Thermal production of S-uuddss in heavy ion collisions

Thermal models can successfully describe hadron production in heavy ion collisions at high energy.

We use a model that has proven to describe particle production with a thermal assumption very successfully to predict abundance of Sexaquarks in heavy ion collisions at the LHC. The model takes into account the different radius of Sexaquarks assumed in the model of Glennys Farrar.

K. A. Bugaev et al., Nucl. Phys. A970 (2018) 133-155 and references therein,  
K. A. Bugaev et al., Universe 5 (2019) 63.



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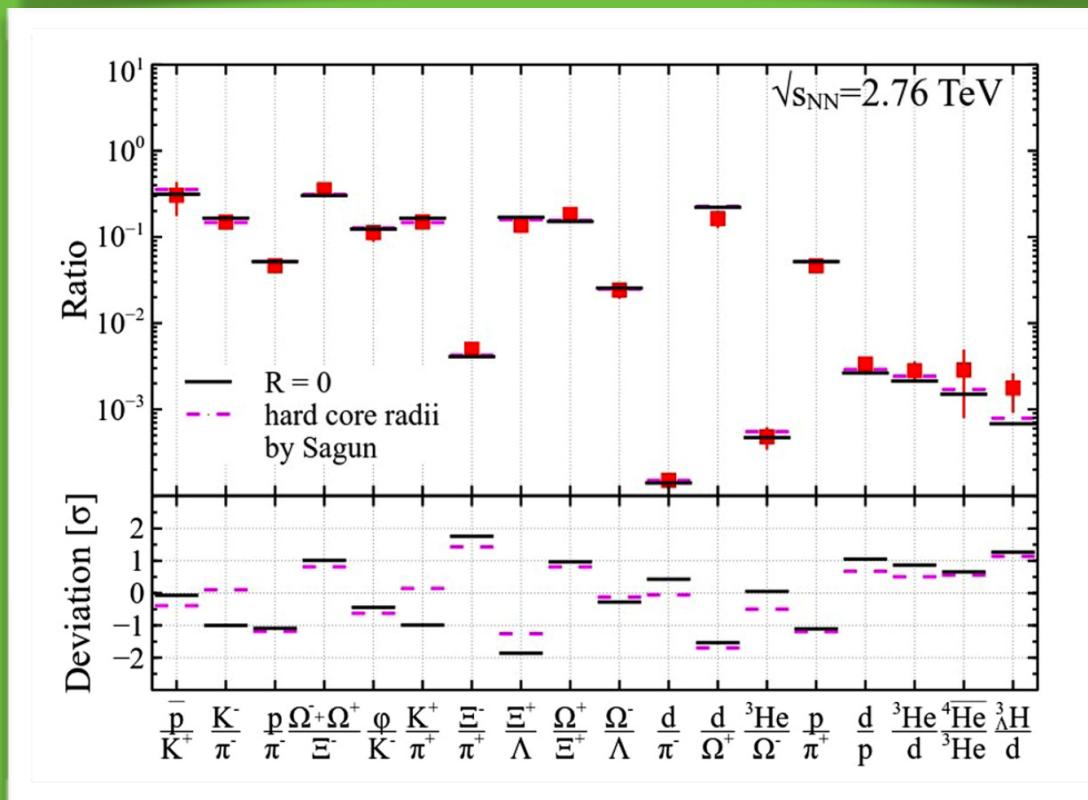


# III Thermal model results on uuddss

# Hadron Resonance Gas Model (HRGM) with multi-component hard-core repulsion (MHRGM)

- Fit gives  $T \sim 151 \pm 7$  MeV and  $\chi^2/\text{DOF} = 0,8$

K.A. Bugaev et al. / Nuclear Physics A 970 (2018) 133–155

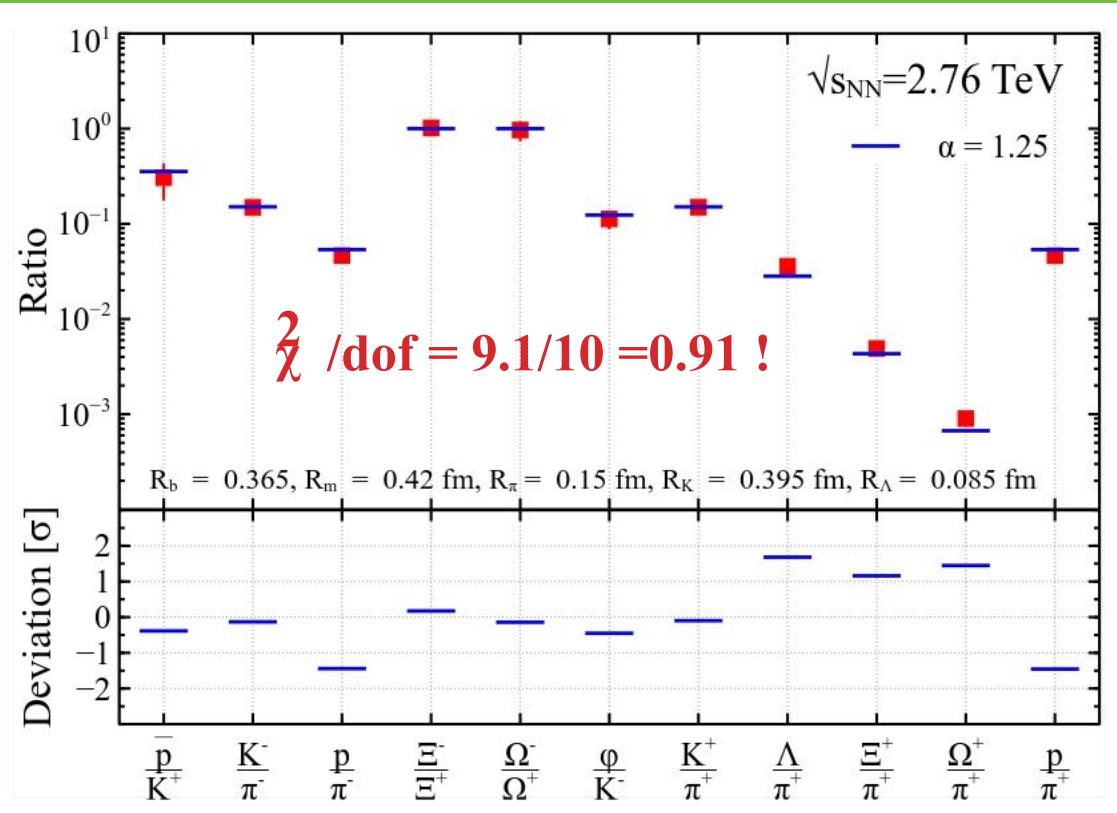


Pb+Pb  
 $\text{sqrt}(s)= 2.76 \text{ TeV}$

with hard-core radii

Fig. 3. The full set of ALICE data (see Table 2) was fitted by the MHRGM with the hard-core radii taken from Ref. [5] with the CFO temperature  $T_{\text{CFO}} \simeq 151 \pm 7$  MeV and  $\chi^2/\text{ndf} \simeq 13.827/17 \simeq 0.8$ . For a comparison the ideal gas fit results are also shown which correspond to  $T_{\text{CFO}} \simeq 148 \pm 7$  MeV and  $\chi^2/\text{ndf} \simeq 19.63/17 \simeq 1.15$ . The upper panel shows the fit of the ratios, while the lower panel shows the deviation between data and theory in units of estimated error.

# Hadron Resonance Gas Model (HRGM) with Induced Surface Tension EOS Results for LHC energy



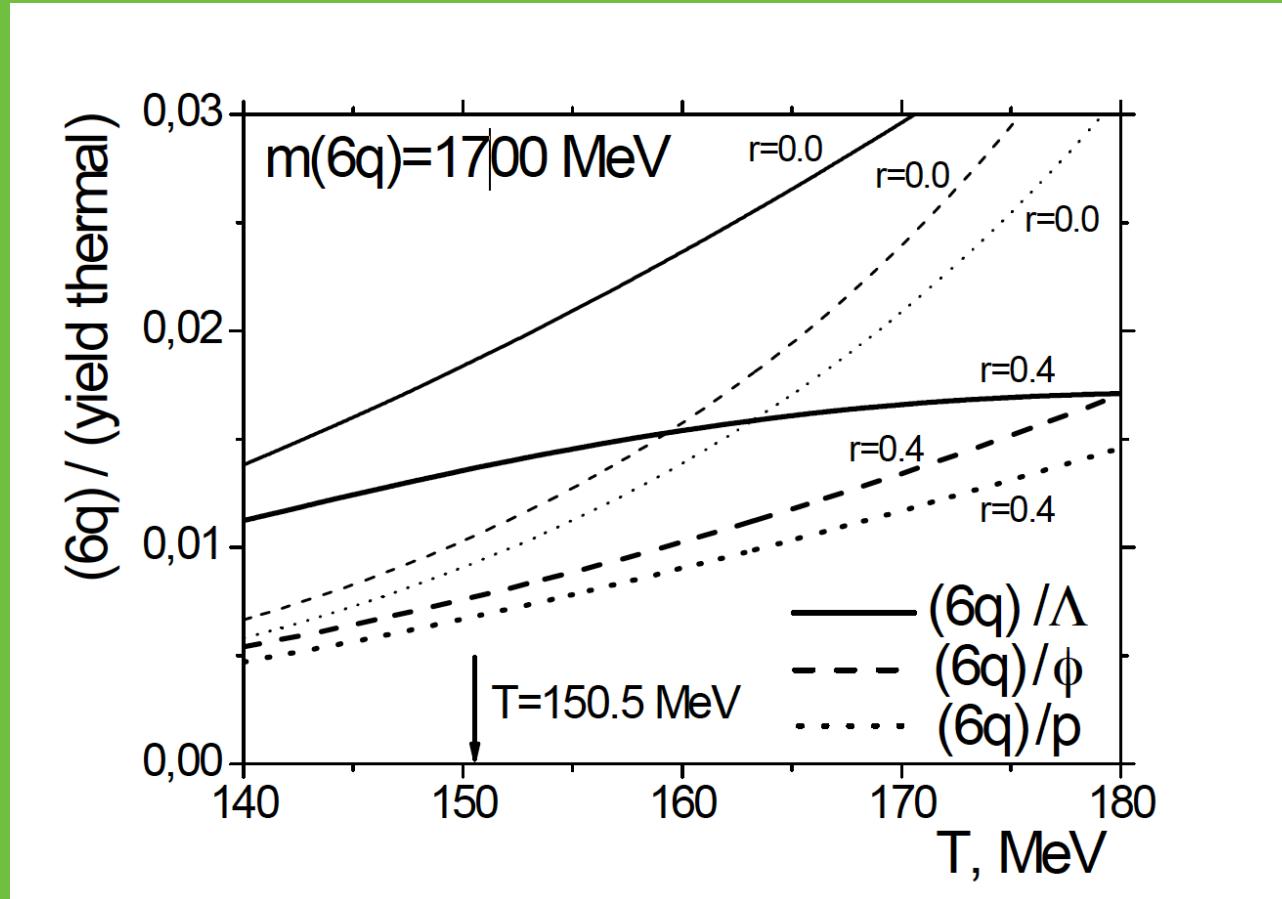
Combined fit of AGS, SPS, RHIC and LHC data

$$\chi^2_{tot}/dof \simeq 64.8/60 \simeq 1.08$$

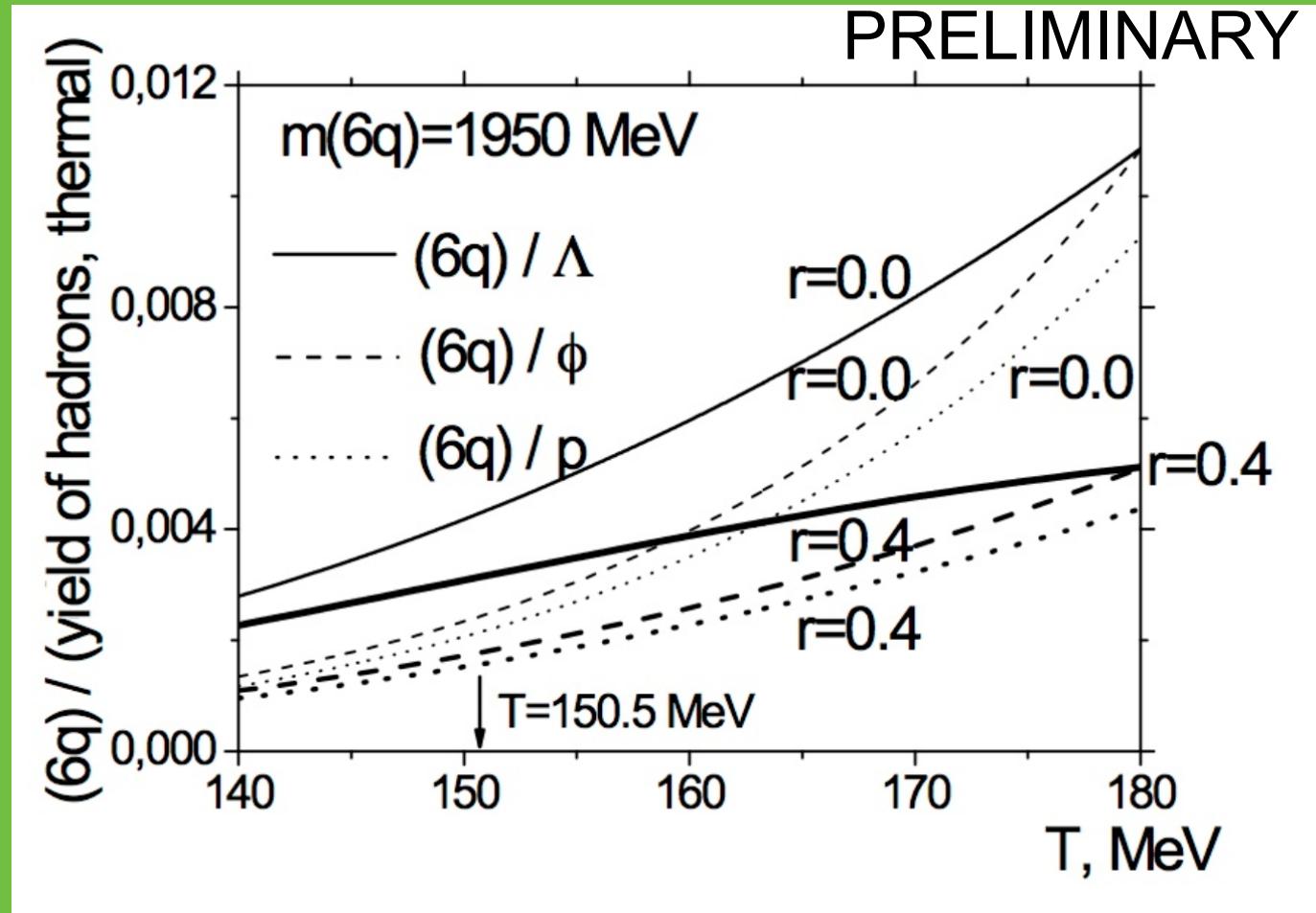
K. A. Bugaev et al., Nucl. Phys. A970 (2018) 133-155 and references therein,  
K. A. Bugaev et al., Universe 5 (2019) 63

# Thermal production of S-uuddss in heavy ion collisions

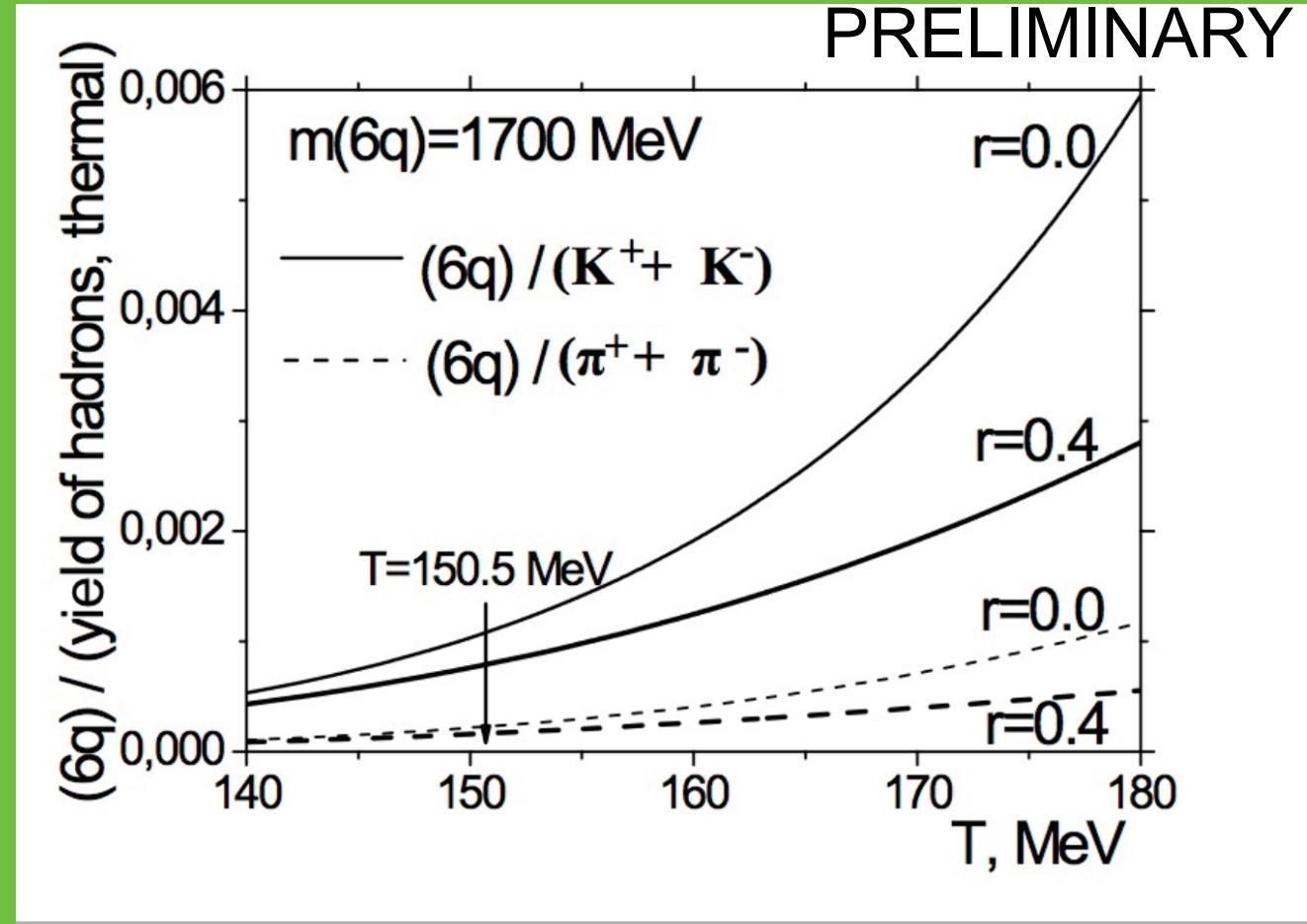
PRELIMINARY



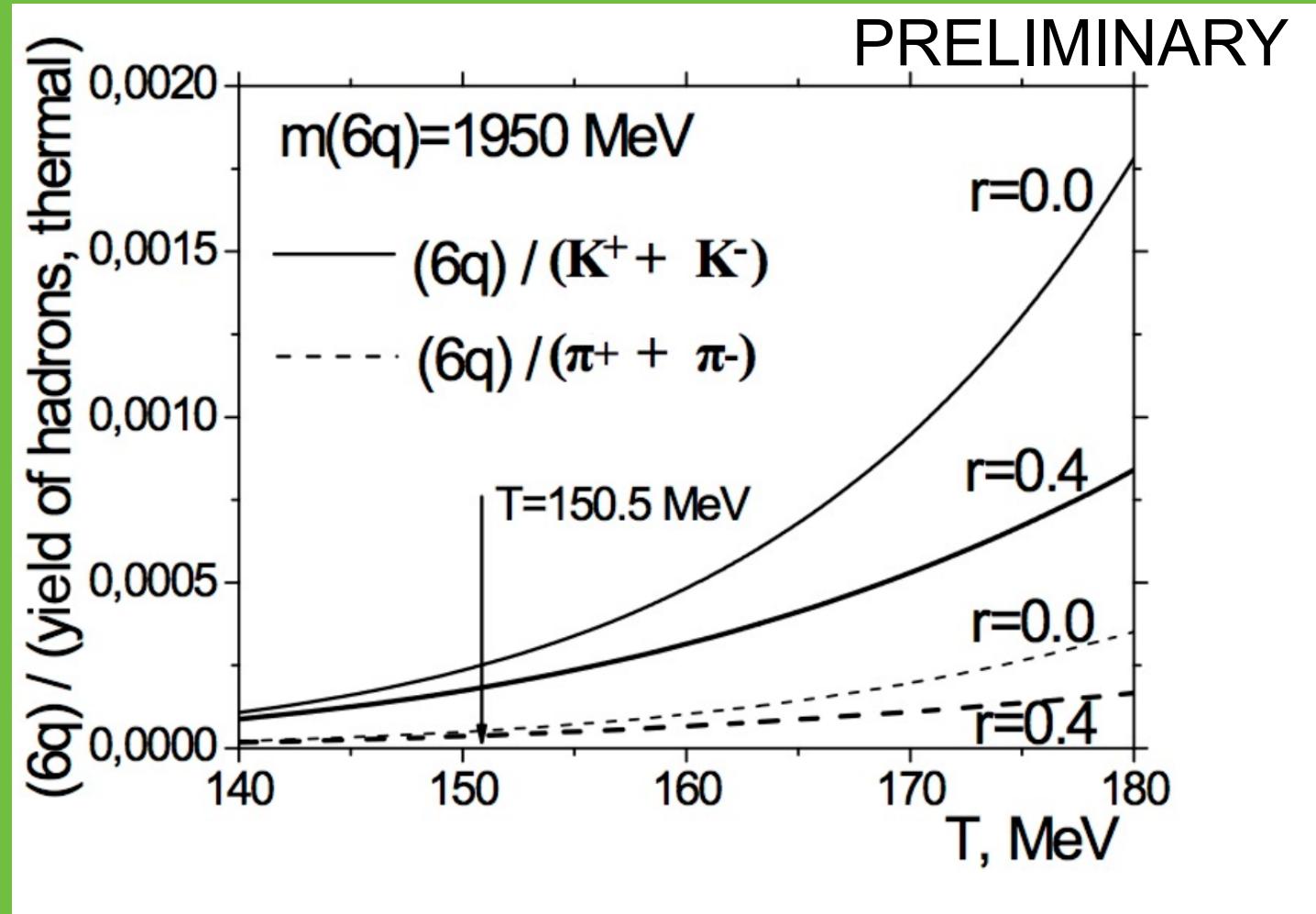
# Thermal production of S-uuddss in heavy ion collisions



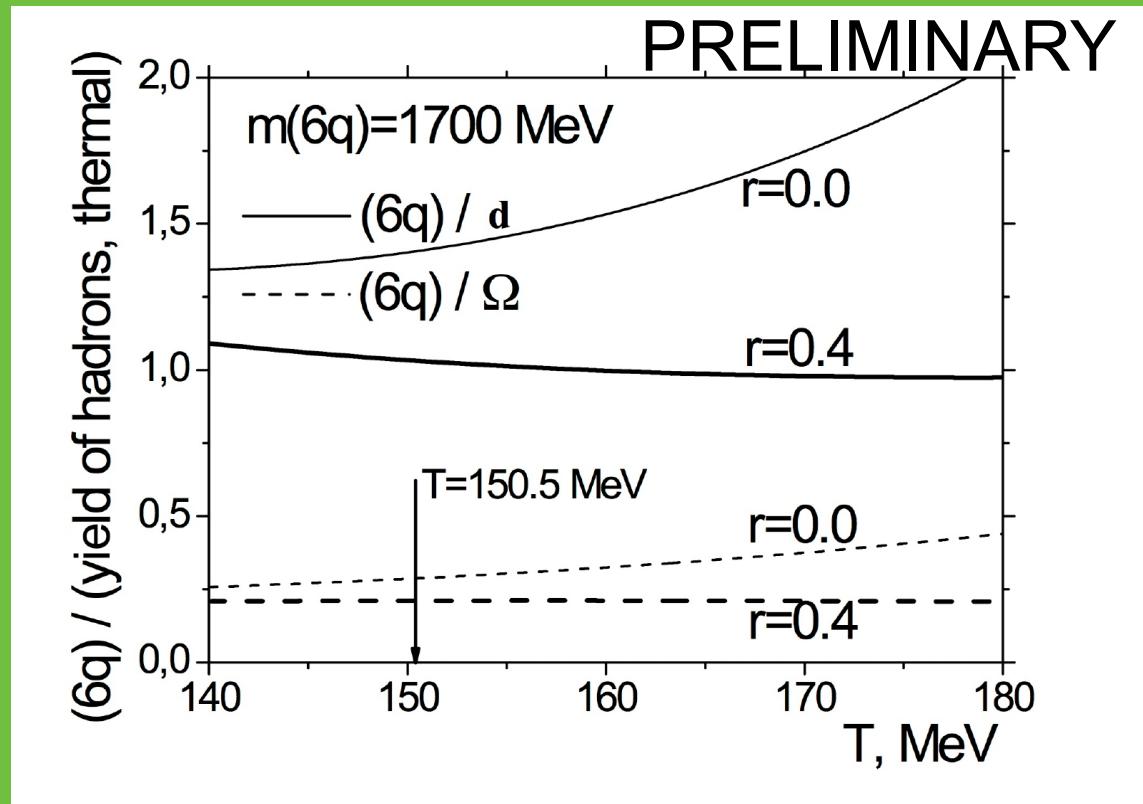
# Thermal production of S-uuddss in heavy ion collisions



# Thermal production of S-uuddss in heavy ion collisions



# Thermal production of S-uuddss in heavy ion collisions



ICNFP 2019, D. Blaschke et al.



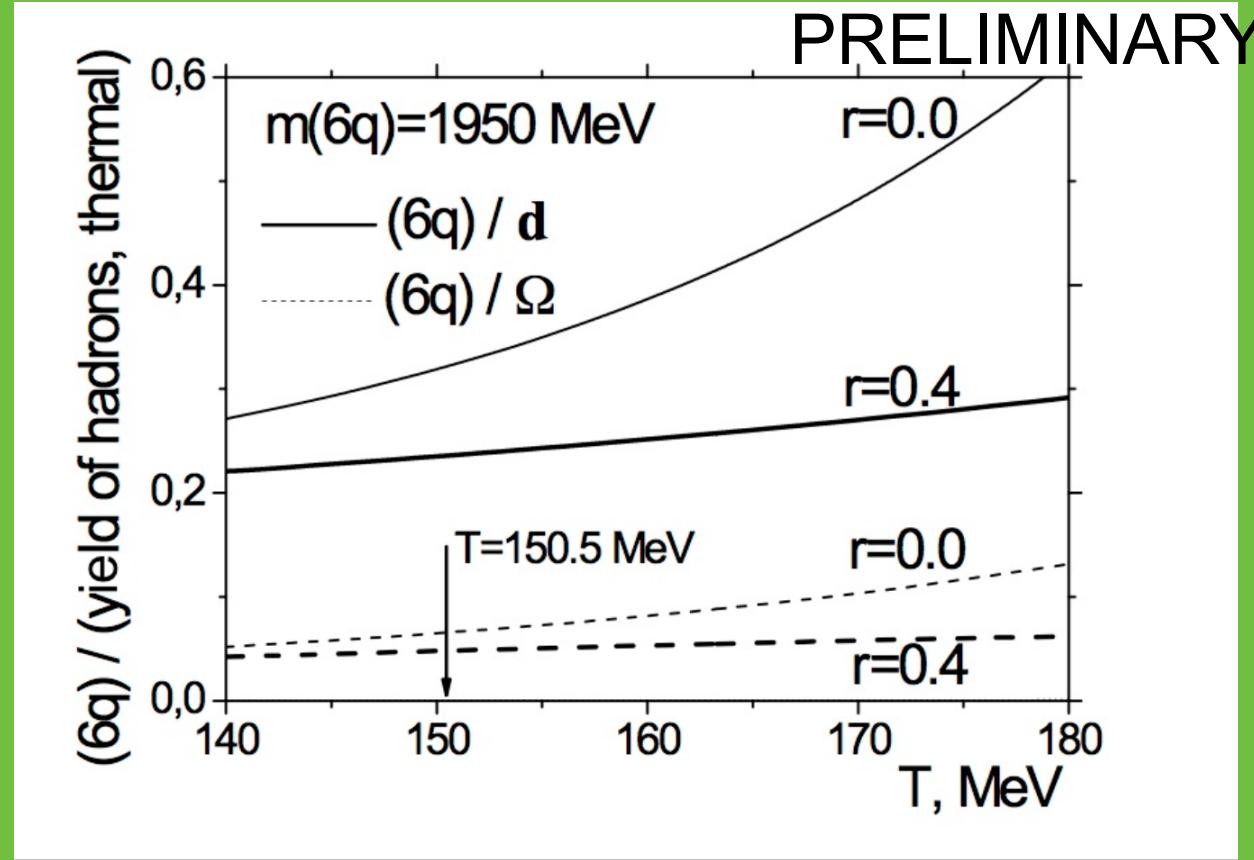
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# Thermal production of S-uuddss in heavy ion collisions

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# IV Summary, Conclusions and Outlook



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## Summary, Conclusions and Outlook

In the framework of the thermal model by K. Bugaev et al, we estimated for the first time ratios of Sexaquarks uuddss to hadrons like proton, kaons, Lambda as well as to the deuteron produced in heavy ion collisions at the LHC

Sextaquarks are produced at relatively high rates, for vanishing and finite radius of 0.4 fm, and for masses of 1700 and 1950 MeV

At T=170 MeV the ratio of thermal Sextaquark with mass 1950 GeV to thermal deuteron is about 0.45

Further studies will elaborate on these results soon.



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**Thank you very much !**