



***HELIOS: a new approach to studying  
transfer reactions in inverse kinematics\****  
*(and potential for the use of a HELIOS-like spectrometer at HIE-ISOLDE)*

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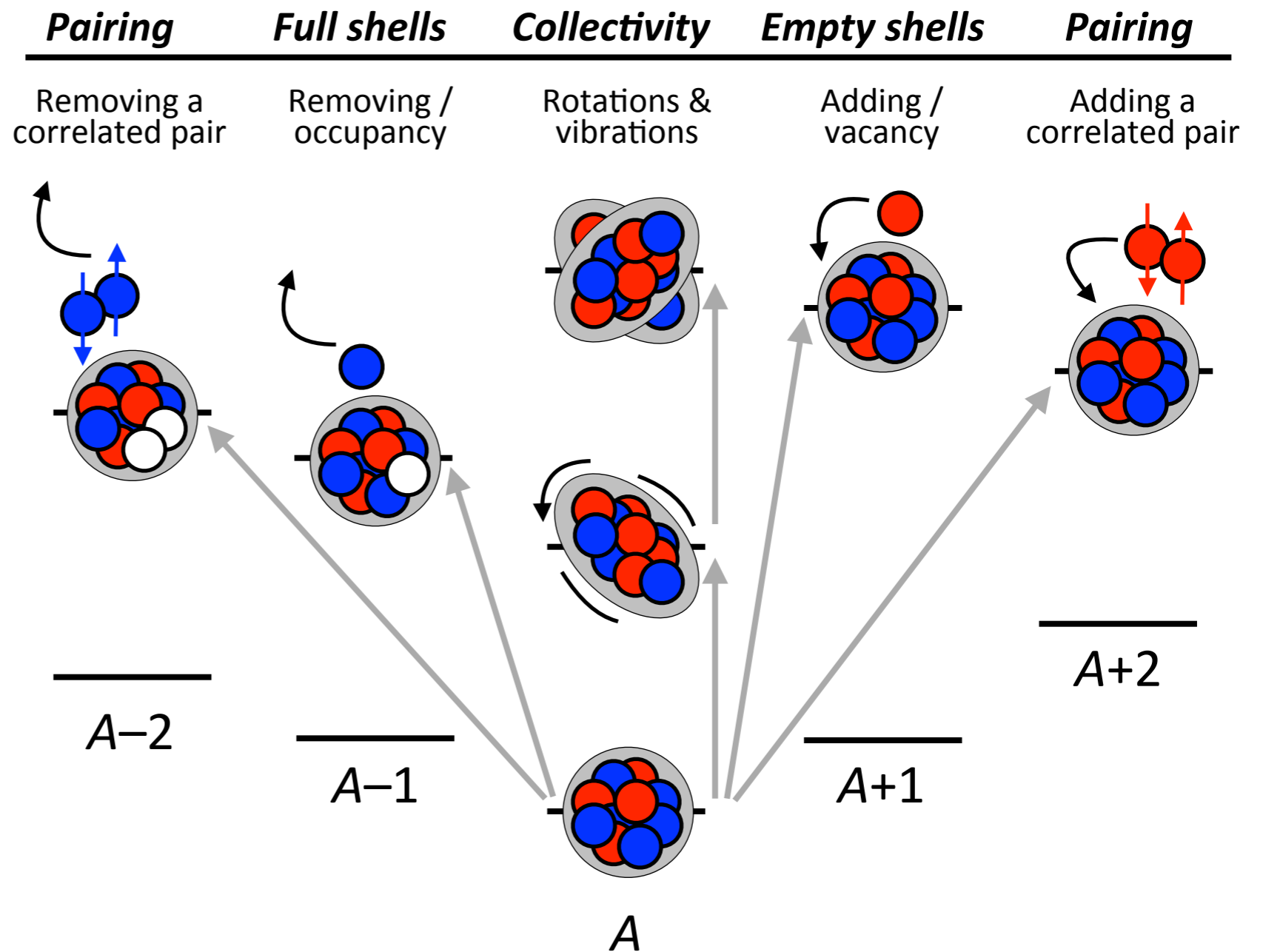
ISOLDE Workshop and Users meeting  
8-10 December 2010

# Motivation – nuclear structure with radioactive beams

(via single- and multi-nucleon transfer reactions)

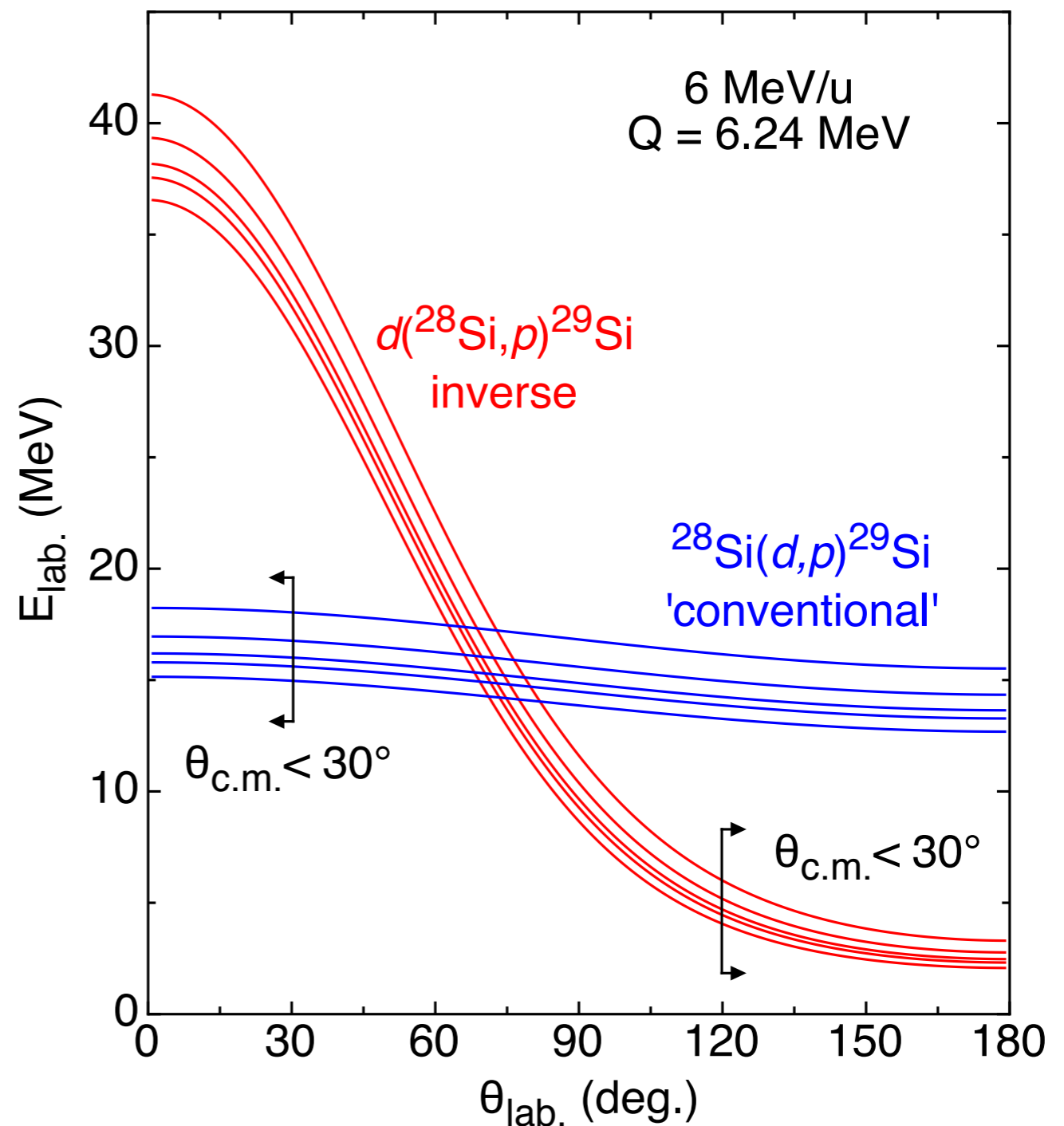
## Explore:

- single-particles states; shell structure evolution,
- pair correlations with two-nucleon transfer e.g.  $(p,t)$ ,  $(t,p)$ ,
- collectivity,  $\beta$  decay, moments, Coulomb excitation,
- Clustering,  $np$  pairing, test ab-initio methods ... etc.



# The problem with inverse kinematics

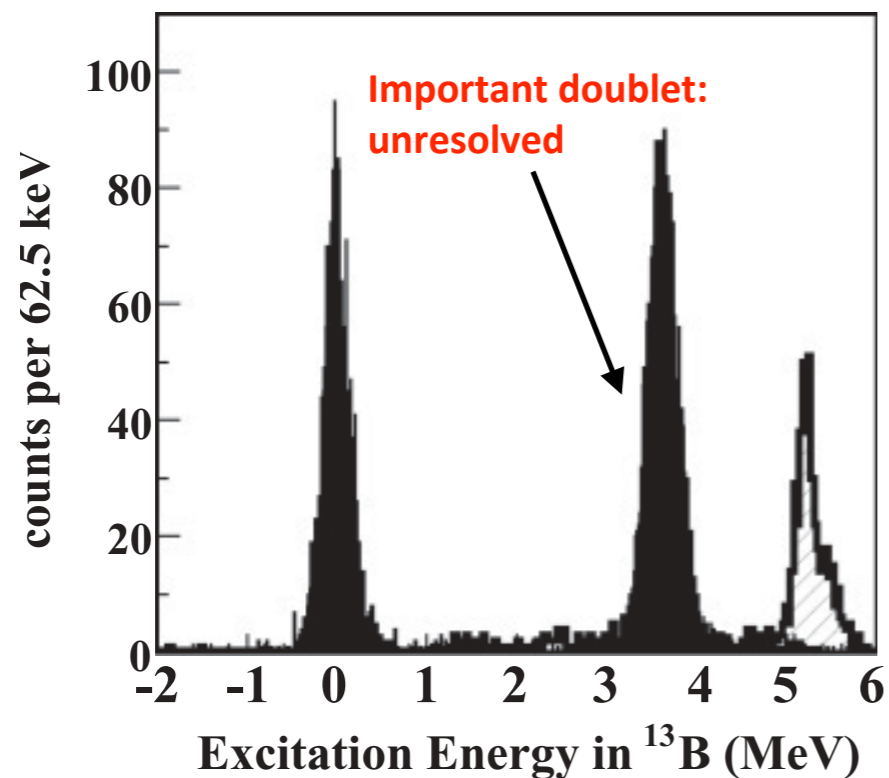
- Particle identification,  $\Delta E$ - $E$  techniques more challenging at **low energies**
- **Strong energy dependence** with respect to laboratory angle
- **Kinematic compression** at forward c.m. angles (in fact nearly all angles)
- Typically leading to **poor resolution** (100s of keV)



# The problem with inverse kinematics

Using the traditional approach of placing a segmented Si detector at a fixed laboratory angle can result in poor excitation-energy resolution, typically of the order of  $\sim 300$  keV (better can be achieved for light nuclei).

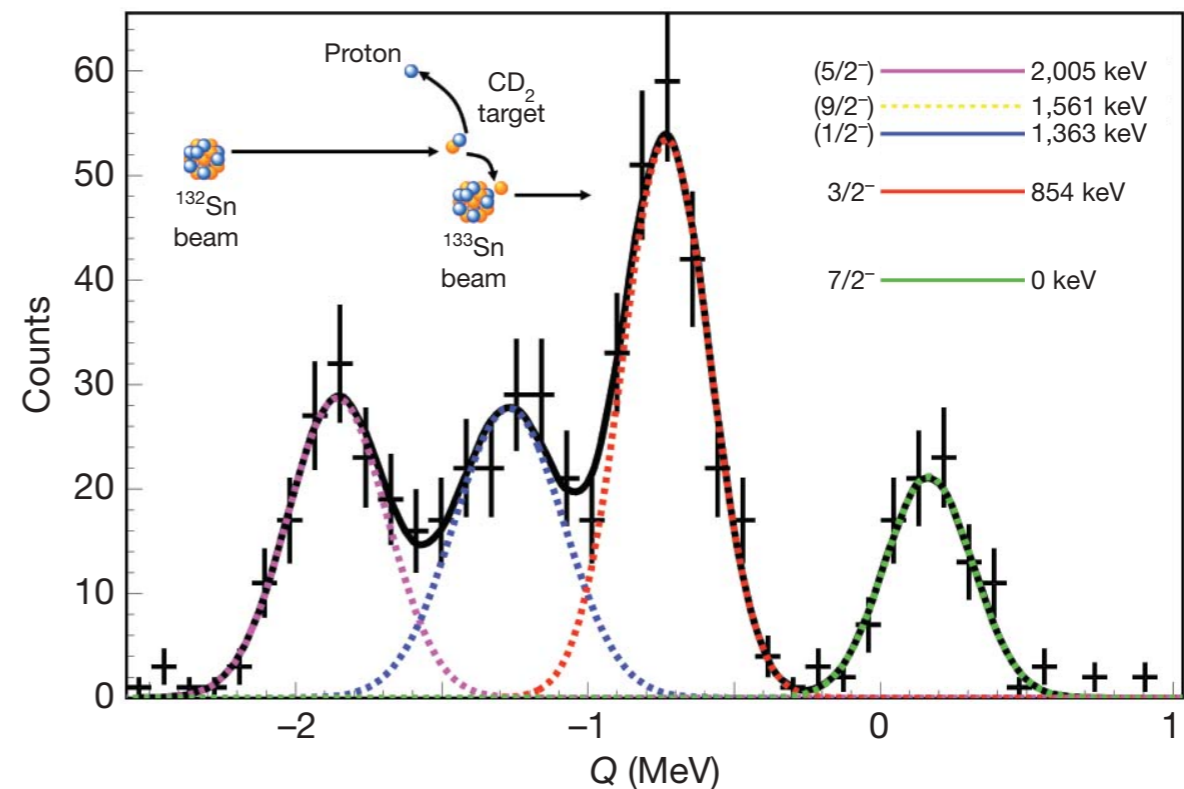
e.g.  $d(^{12}\text{B},p)^{13}\text{B}$ ,  $\sim 250$  keV FWHM



H. Y. Lee *et al.*, PRC **81**, 015802 (2010)

**(prohibitive - more on this later)**

e.g.  $d(^{132}\text{Sn},p)^{133}\text{Sn}$ ,  $\sim 300$  keV FWHM



K. L. Jones *et al.*, Nature **465**, 454 (2010)

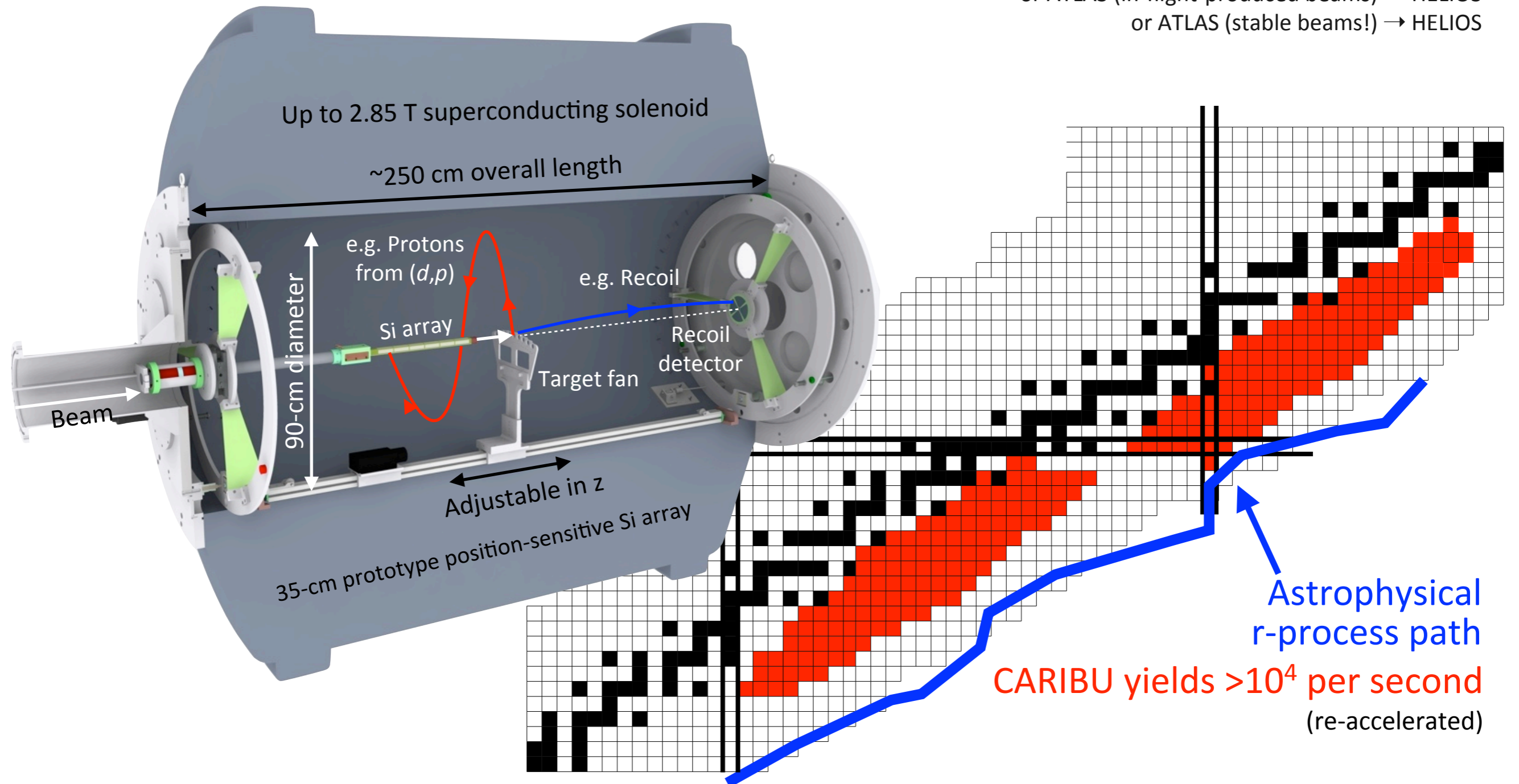
**(not prohibitive - low level density)**

Is there another approach to the problem? ...

# ... HELIOS

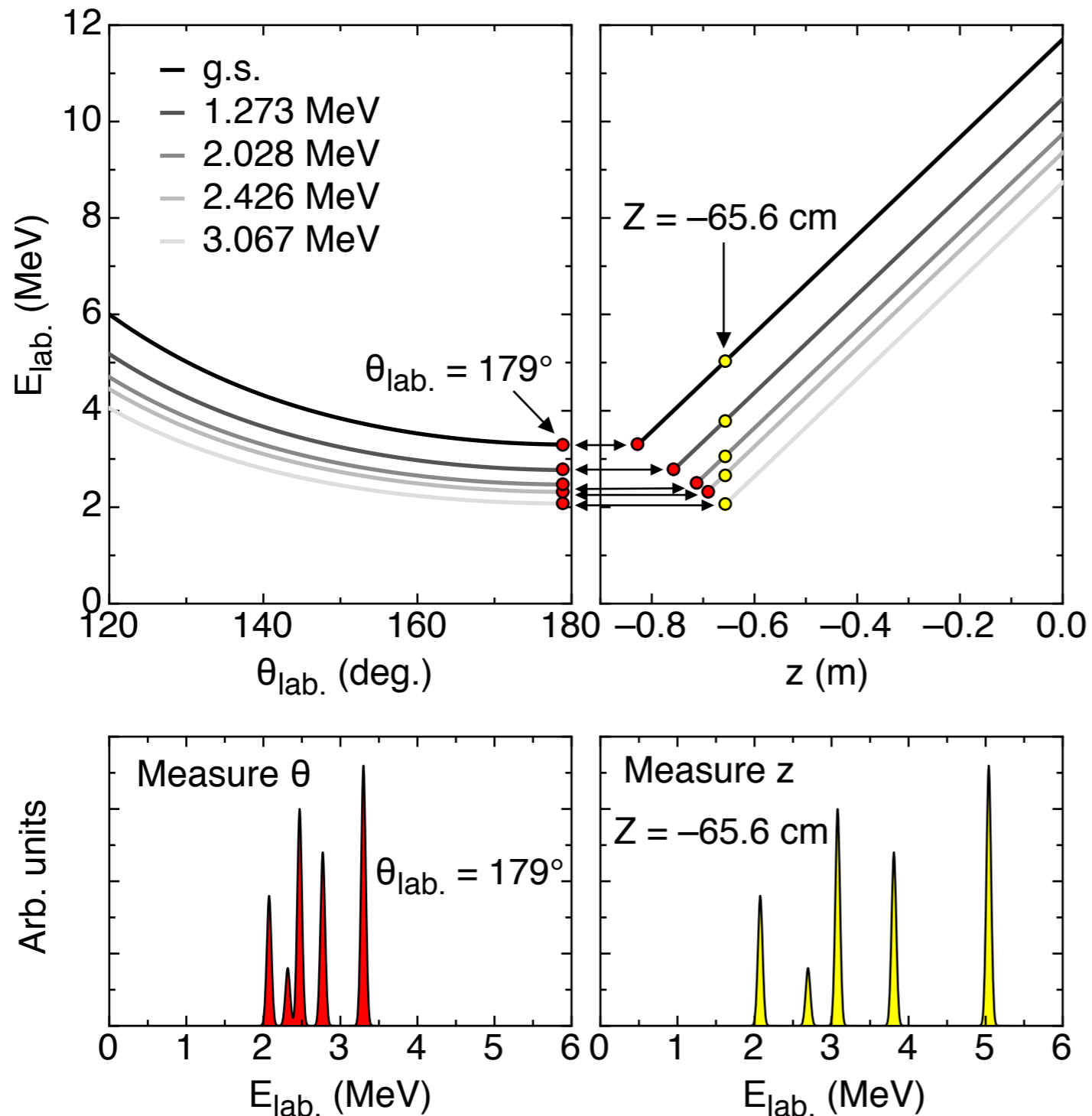
A new type of spectrometer – the helical orbit spectrometer. Born out of the necessity to perform direct reactions in inverse kinematics with good excitation-energy resolution *and* with low intensity beams.

CARIBU → ATLAS → HELIOS  
or ATLAS (in-flight-produced beams) → HELIOS  
or ATLAS (stable beams!) → HELIOS



# The advantages of HELIOS

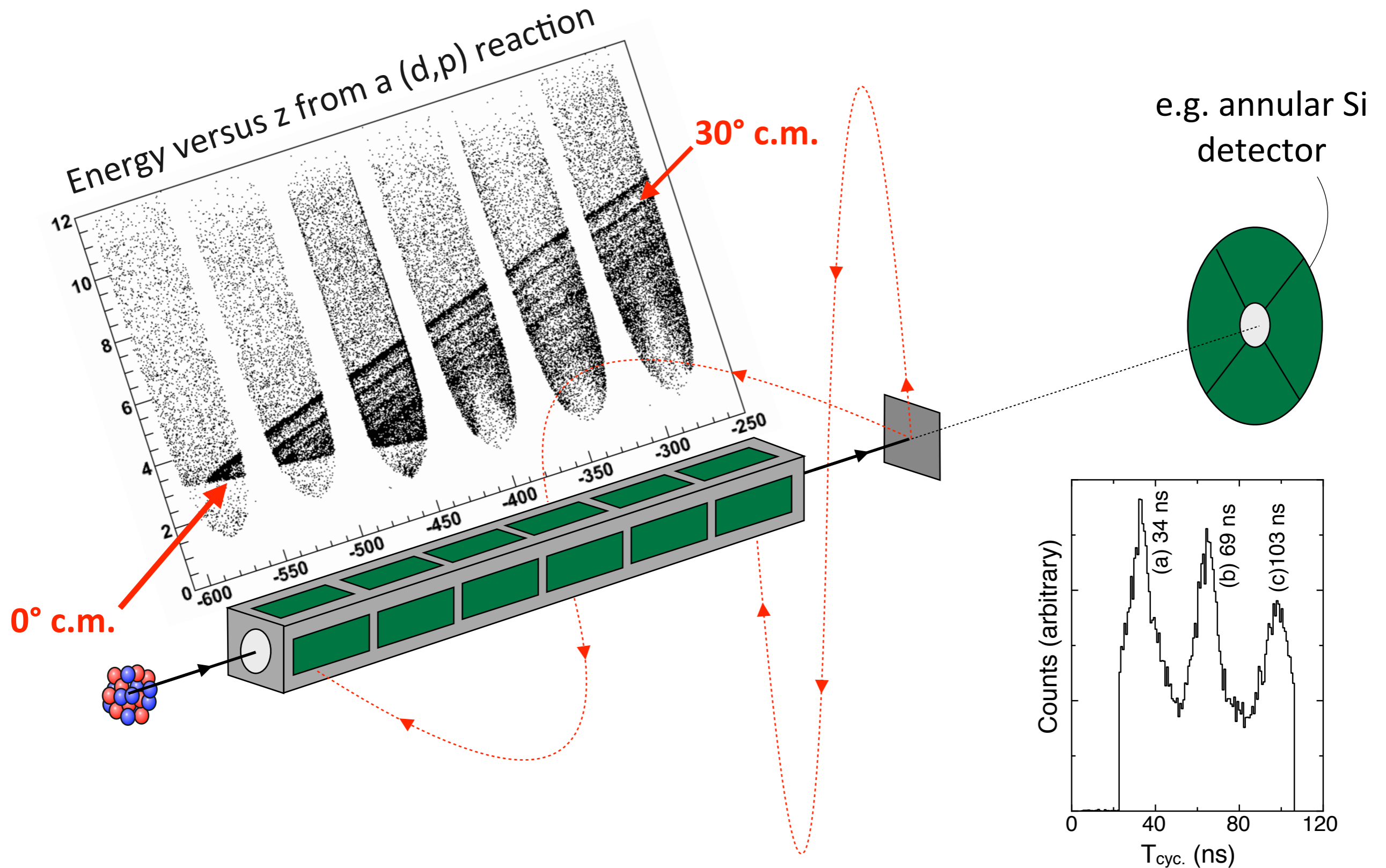
Example:  $d(^{28}\text{Si},p)$  at 6 MeV/u with a 2-T field



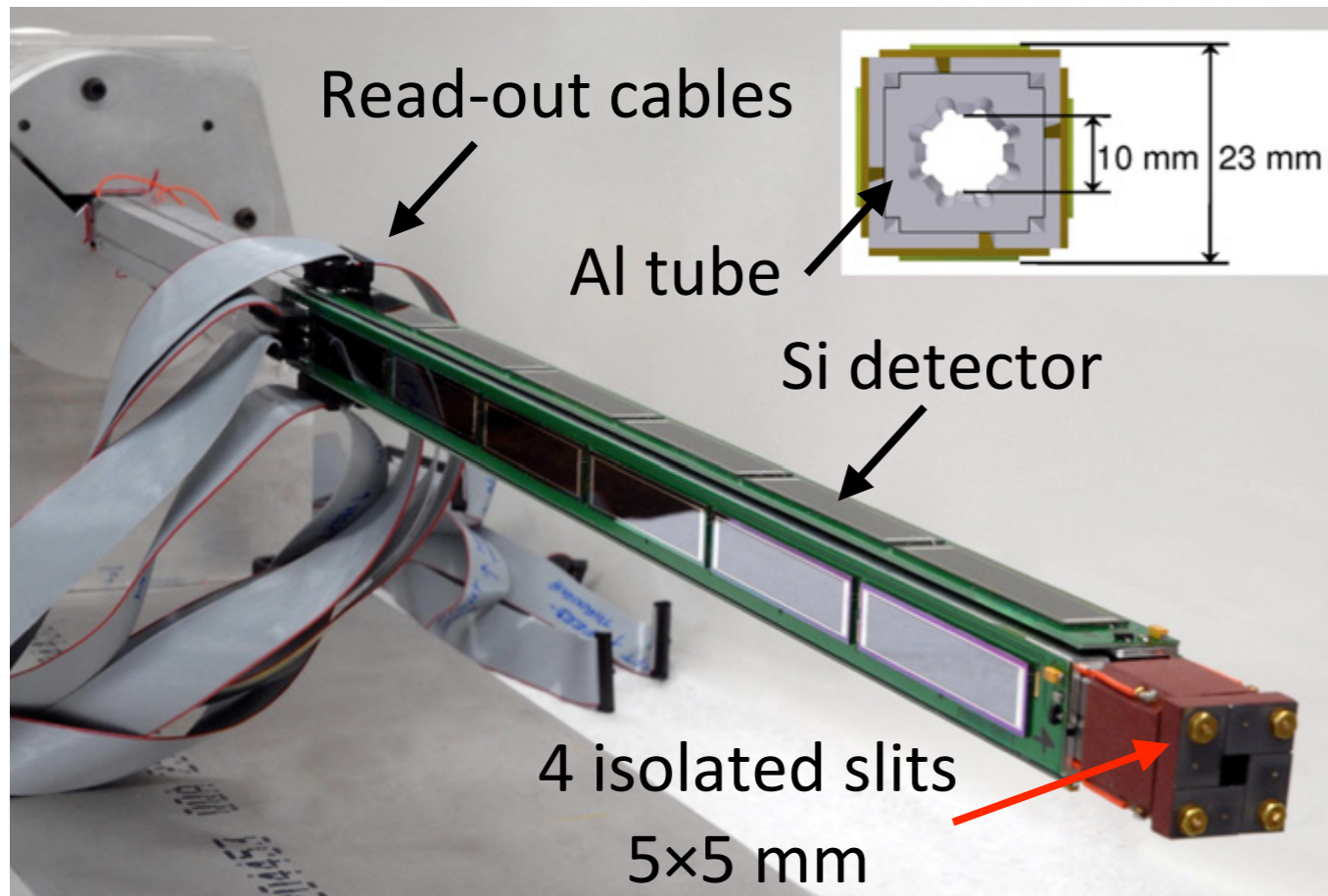
- A simple **linear** relationship between energy and  $z$ , where the energy separation is (nearly) **identical** to the excitation energy in the residual nucleus.
- Avoids the problems of kinematic **compression**.
- Factor of  $\sim 2.4$  improvement in resolution (for this example)



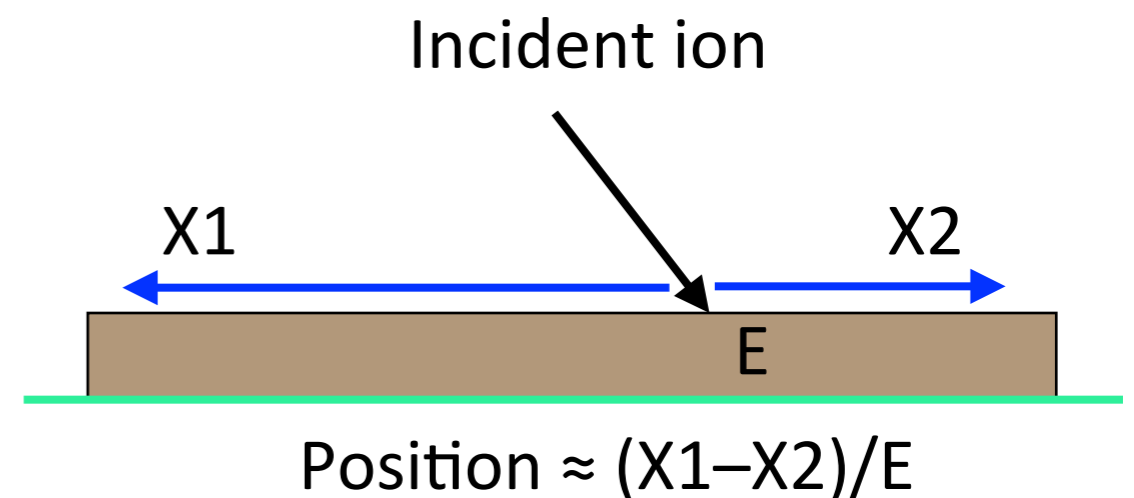
# Motion of ions in HELIOS (cartoon)



# The Si (prototype) array\*

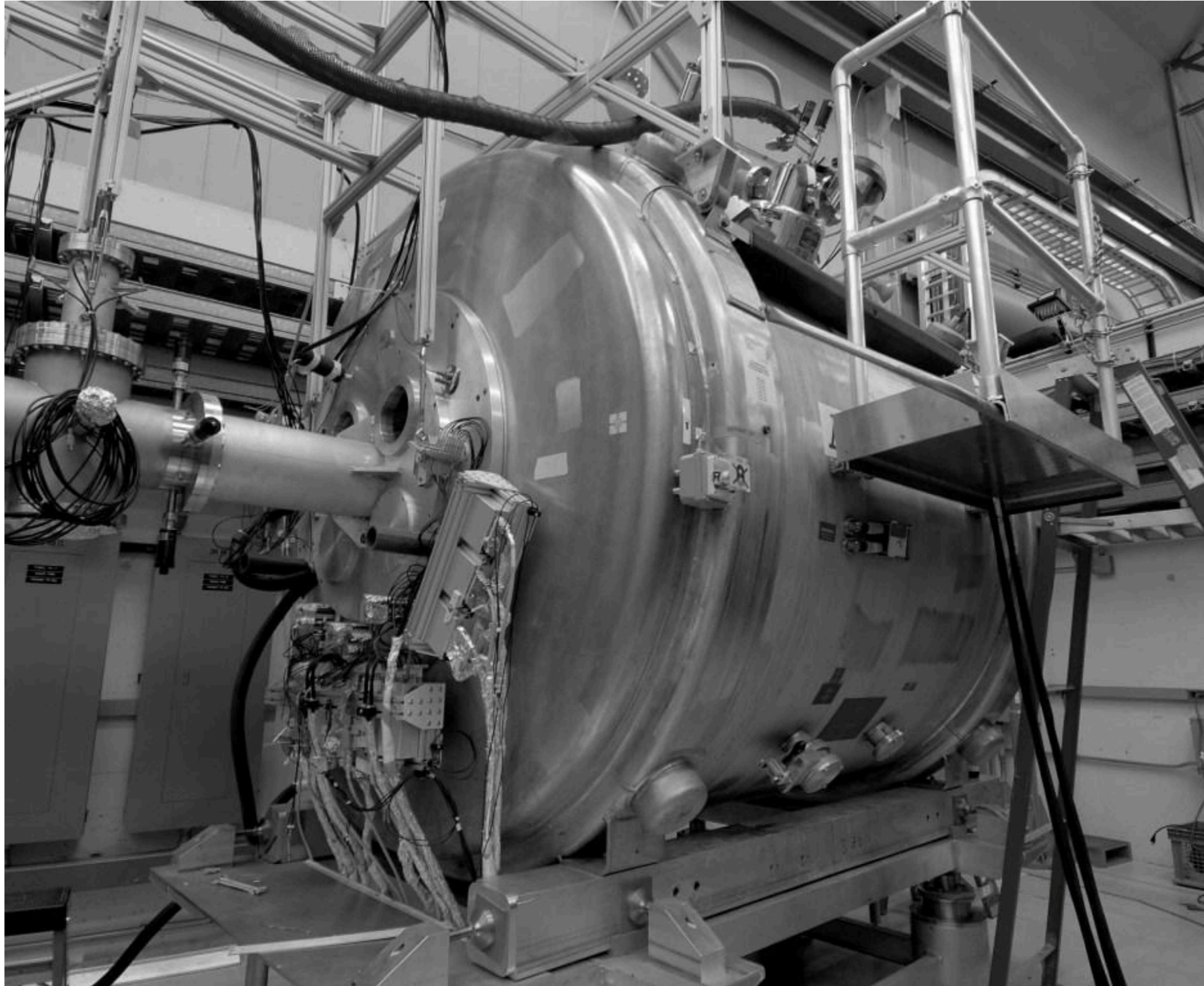


- 4 sides, 6 elements long
- Detector size, 9×50 mm
- 700- $\mu\text{m}$  thick (e.g.  $\sim 10$  MeV protons)
- $\Phi$  coverage, **0.48 of  $2\pi$**
- $\Omega_{\text{element}} = 21 \text{ msr}$
- $\Omega_{\text{array}} = \mathbf{493 \text{ msr}}$





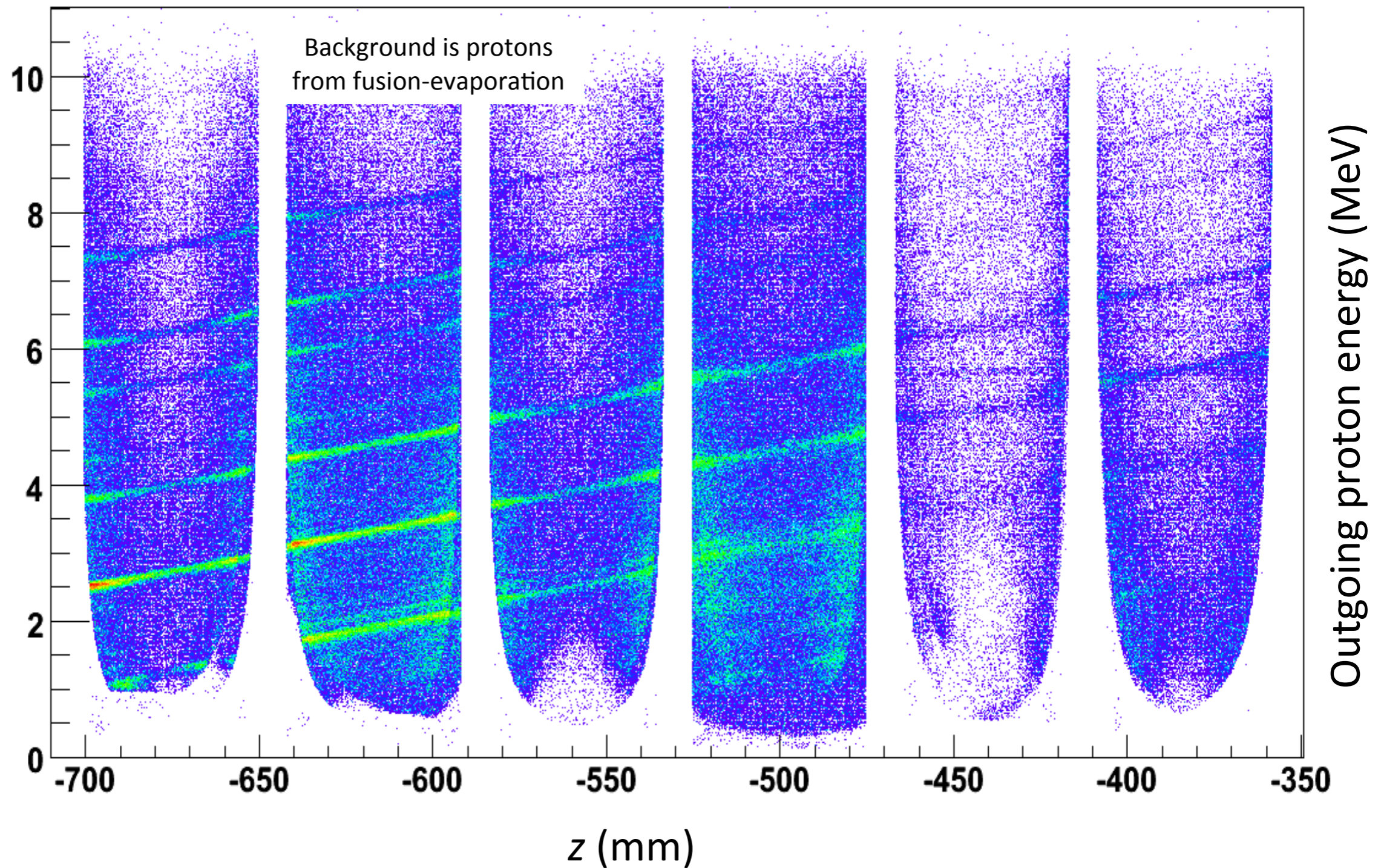
# *Photograph from upstream\**





# *It works as planned, $d(^{28}\text{Si},p)$ at 6 MeV/u, 2 T*

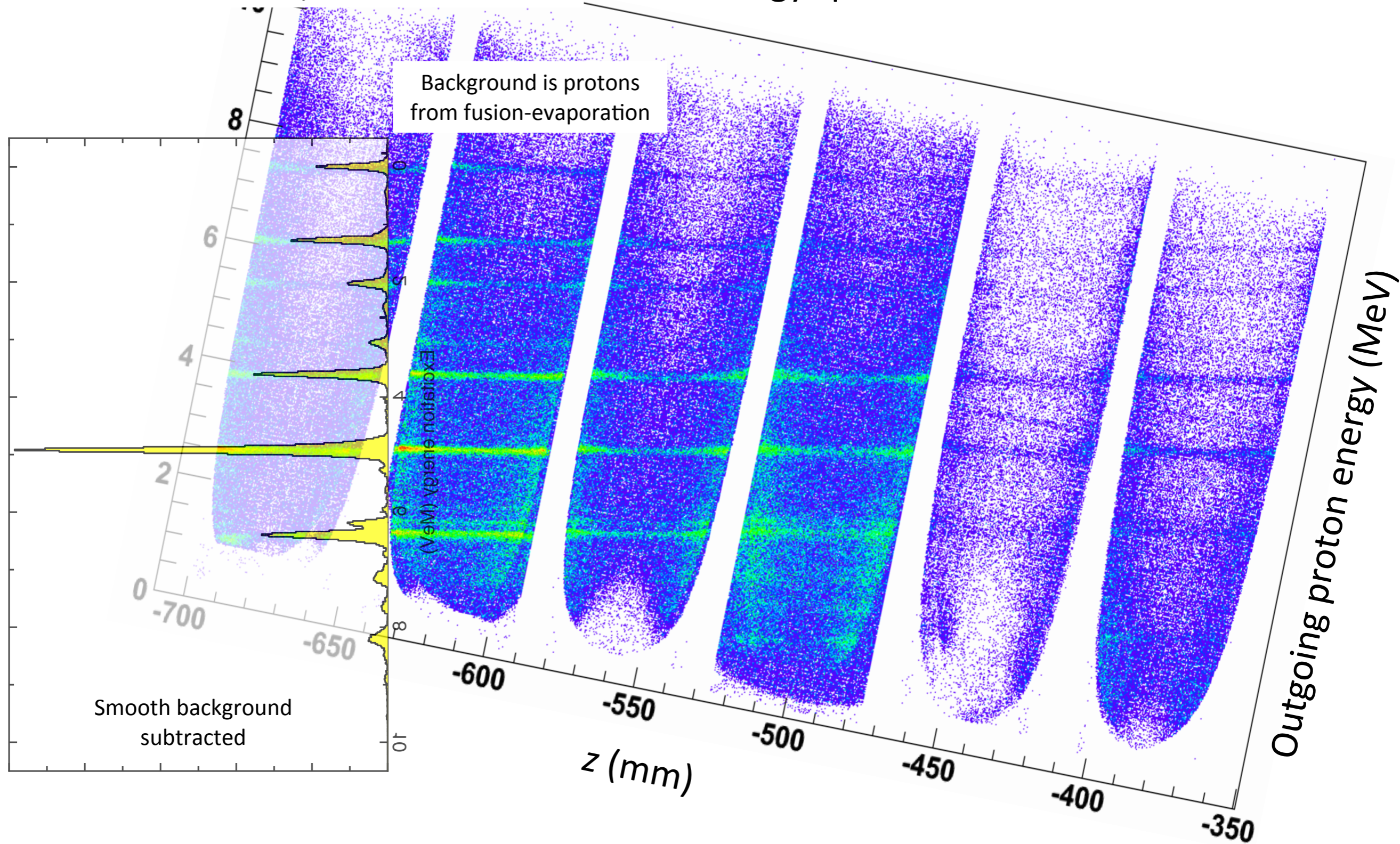
We measure E vs. z, which *is* the excitation-energy spectrum of the residual nucleus



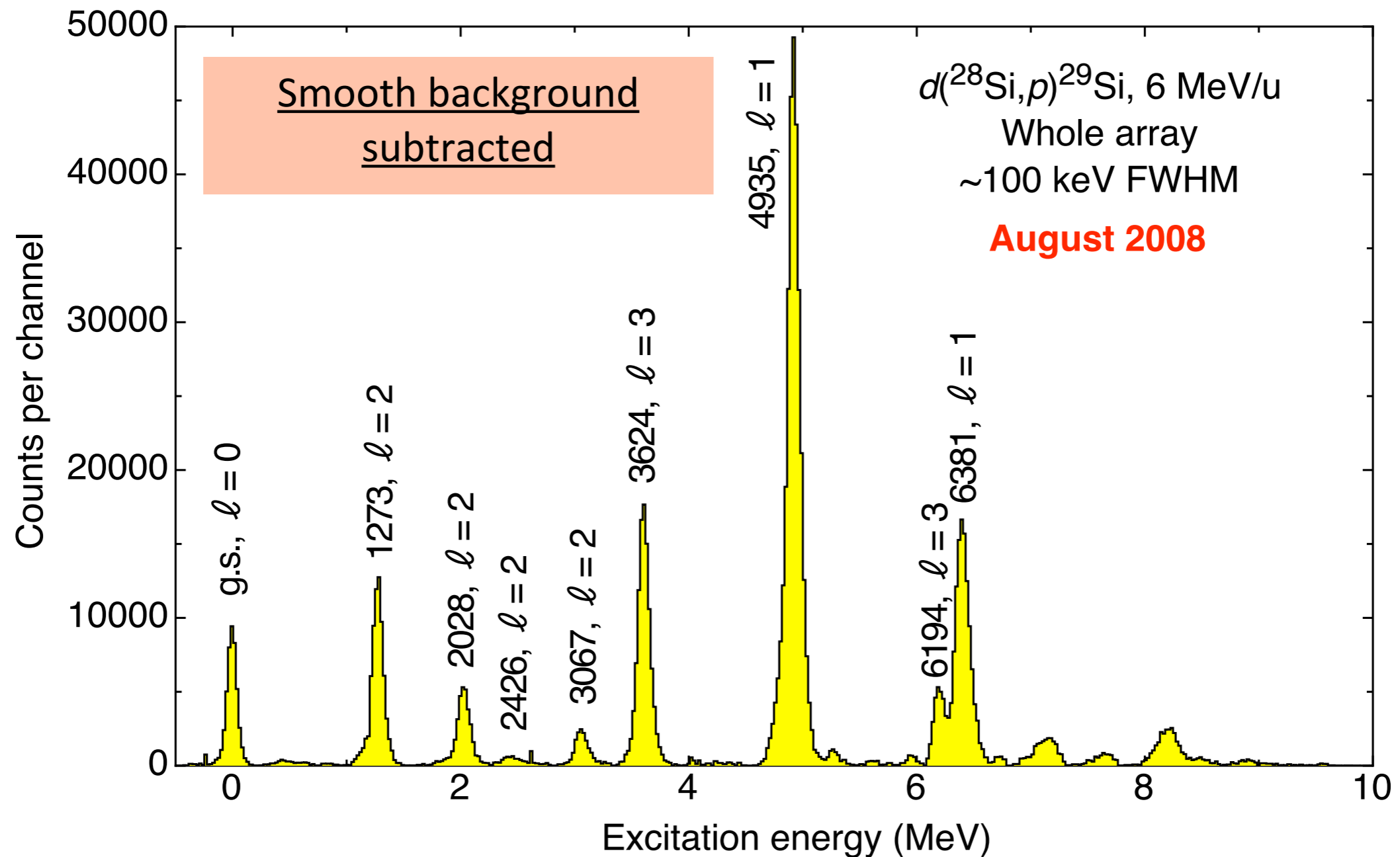


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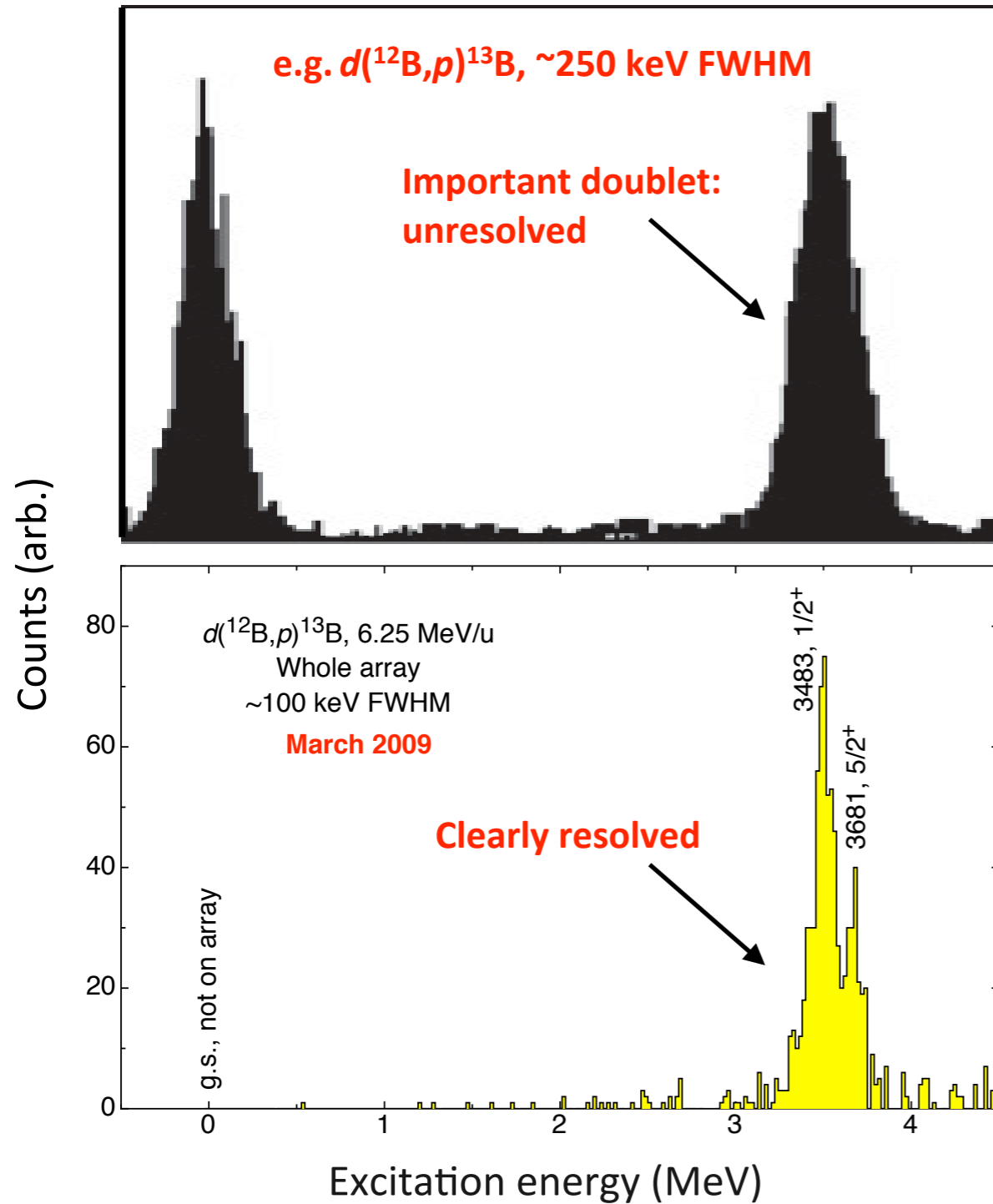


# *It works as planned, $d(^{28}\text{Si},p)$ at 6 MeV/u, 2 T\**



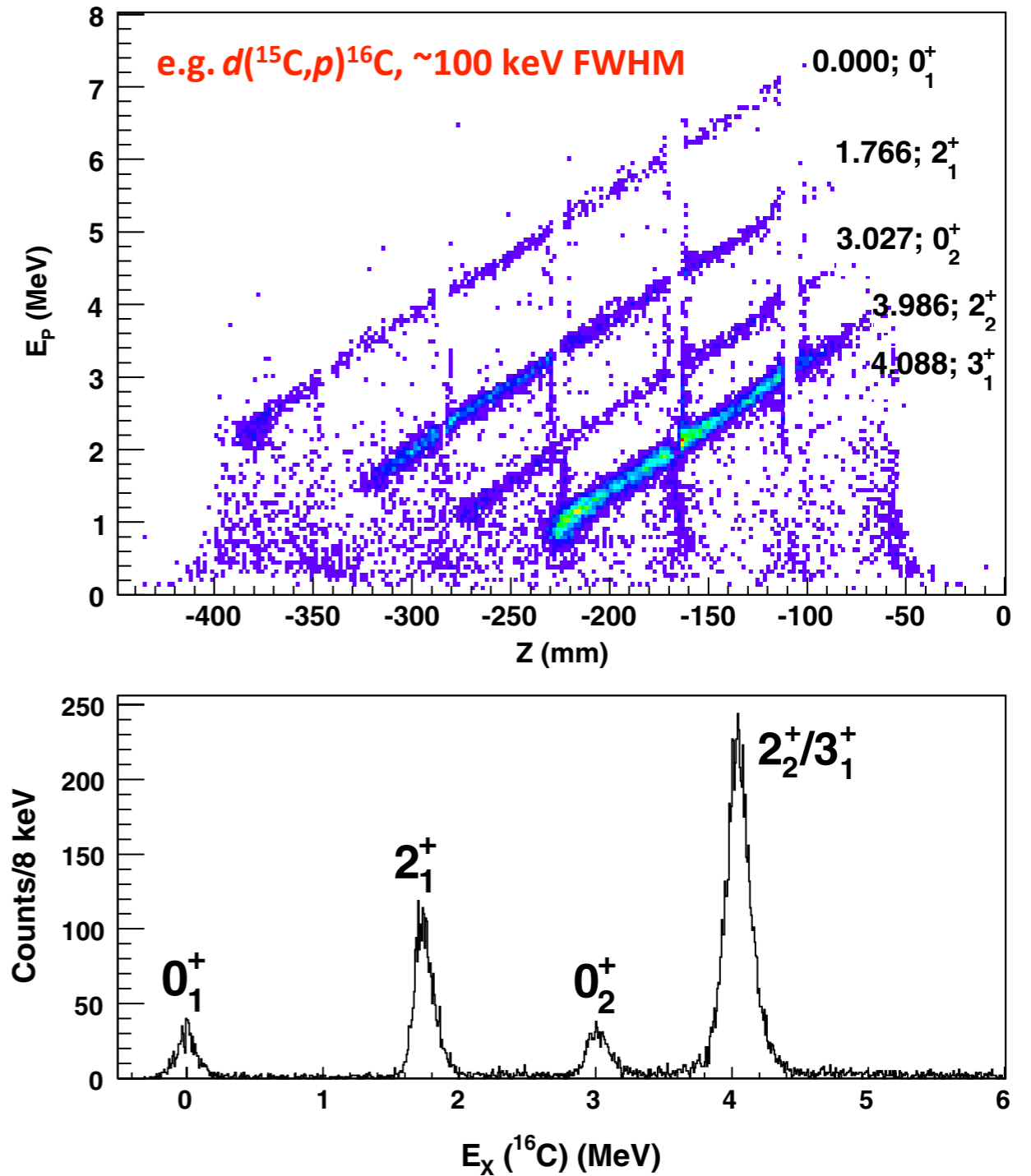


# Other recent measurements



H. Y. Lee *et al.*, Phys. Rev. C **81**, 015802 (2010)

B. B. Back *et al.*, Phys. Rev. Lett. **104**, 132501 (2010)



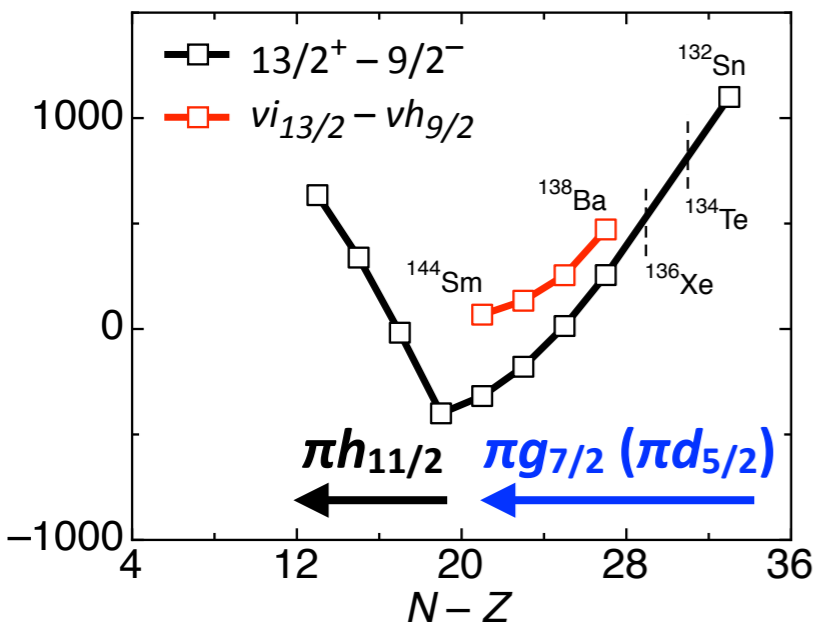
A. H. Wuosmaa *et al.*, Phys. Rev. Lett. **105**, 132501 (2010)

# HELIOS with heavy beams

Keen to see how HELIOS performs with beams of similar mass to those around  $^{132}\text{Sn}$  – *in anticipation of CARIBU beams*

We performed these tests with a bit of physics in mind –  *$d(^{136}\text{Xe}, p)$  in light of single-particle trends in the  $N = 82$  isotones and  $d(^{130}\text{Xe}, p)$  as an early exploration of neutron vacancy (related to  $0\nu 2\beta$  matrix elements)*

BPK *et al.*, Phys. Lett. B658, 216 (2008)



$Z < 64$

‘Mainly’  $\pi g_{7/2}$  filling in the  $N = 82$  cores

REPULSIVE effect with  $vi_{13/2}$

ATTRACTIVE effect with  $vh_{9/2}$

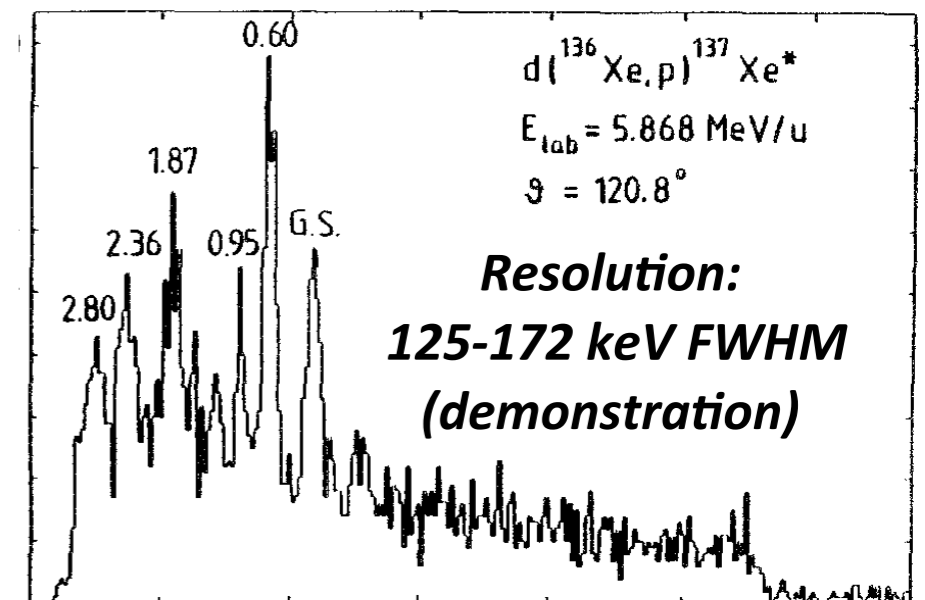
$Z > 64$

$\pi h_{11/2}$  filling in the  $N = 82$  cores

REPULSIVE effect with  $vi_{13/2}$

ATTRACTIVE effect with  $vh_{9/2}$

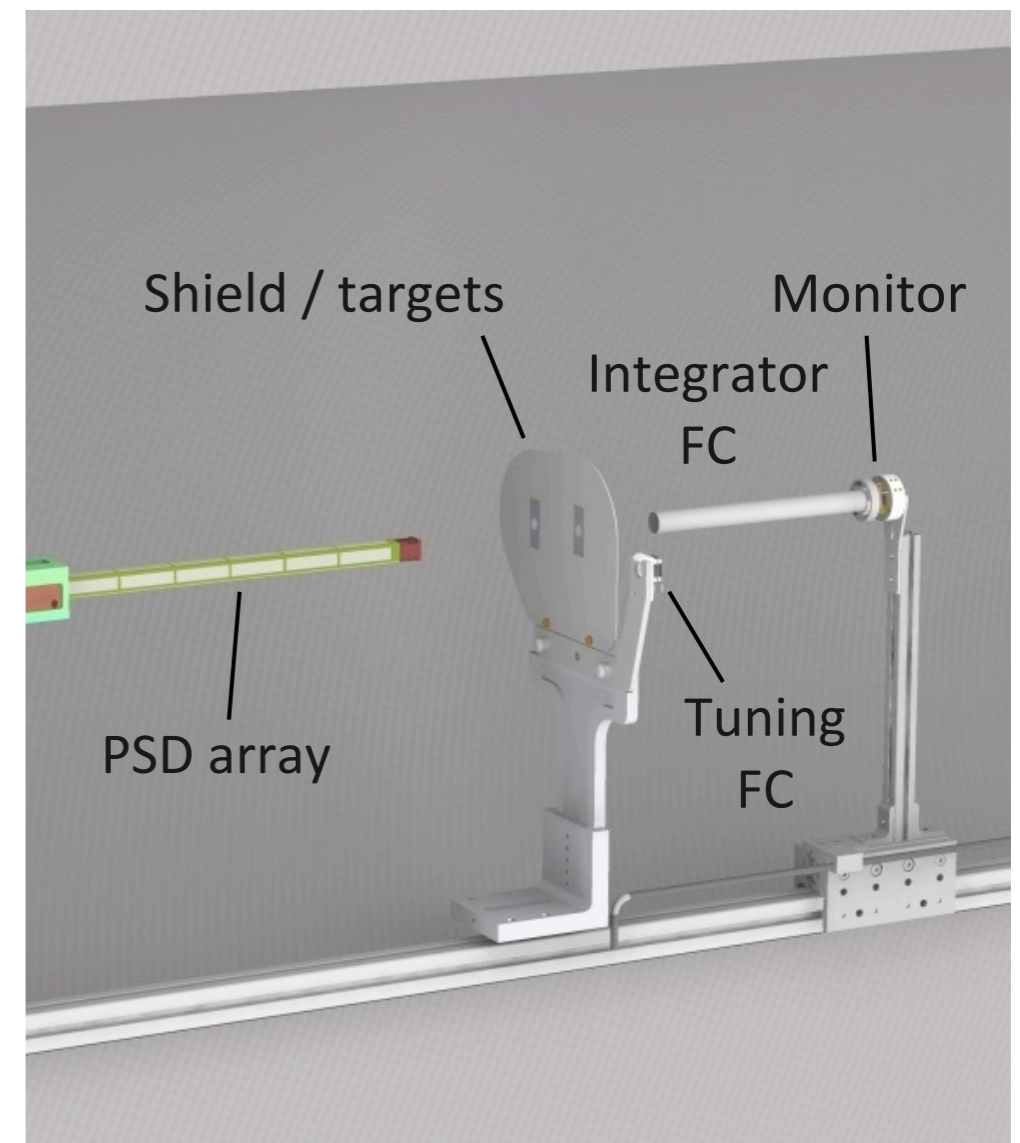
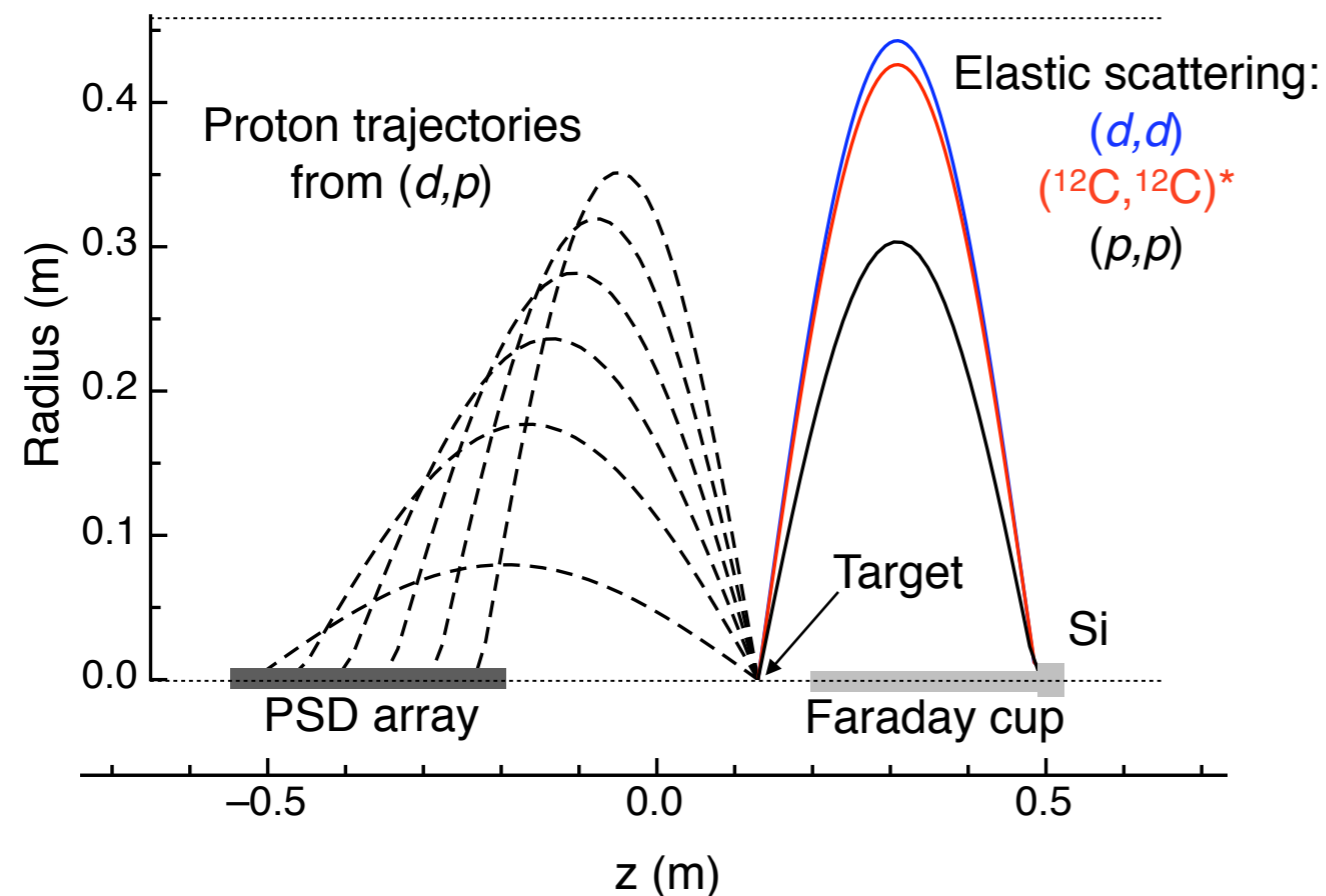
Kraus *et al.*, Z. Phys. A340, 339 (1991)



The first exploration of single-neutron transfer in inverse kinematics – GSI, 1991

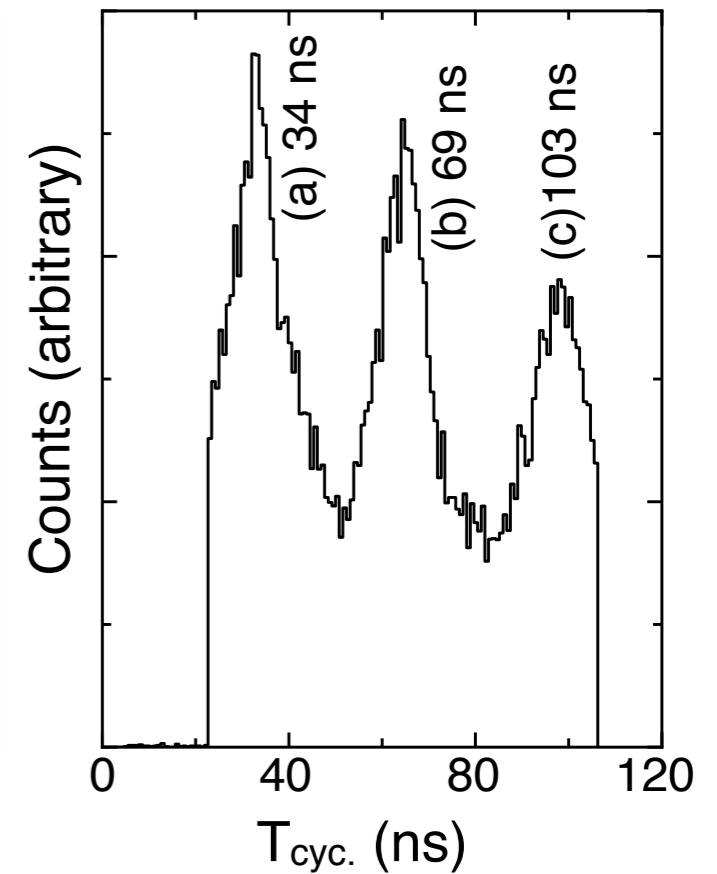
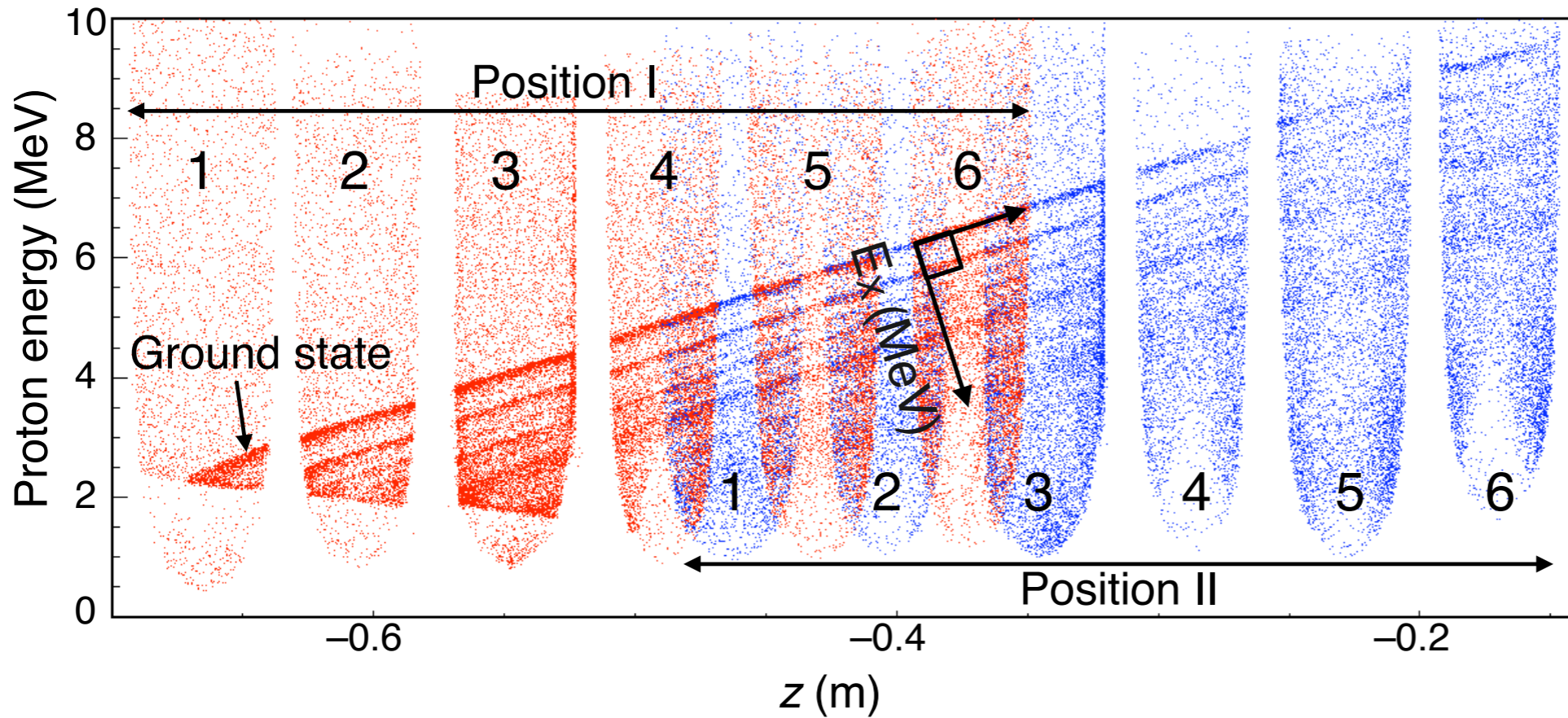
# Absolute cross sections

Desirable to measure absolute cross sections, target thickness, and to understand how the target degrades under beam strike. At 5 MeV/u, elastically scattering target species (deuterons, carbon ions) are within a few percent of **Rutherford scattering** at angles close to 90° in the lab. frame. This can be used to set an absolute cross-section scale – at 10 MeV/u the same set up (though not Rutherford) can be used to monitor the beam luminosity and target degradation.

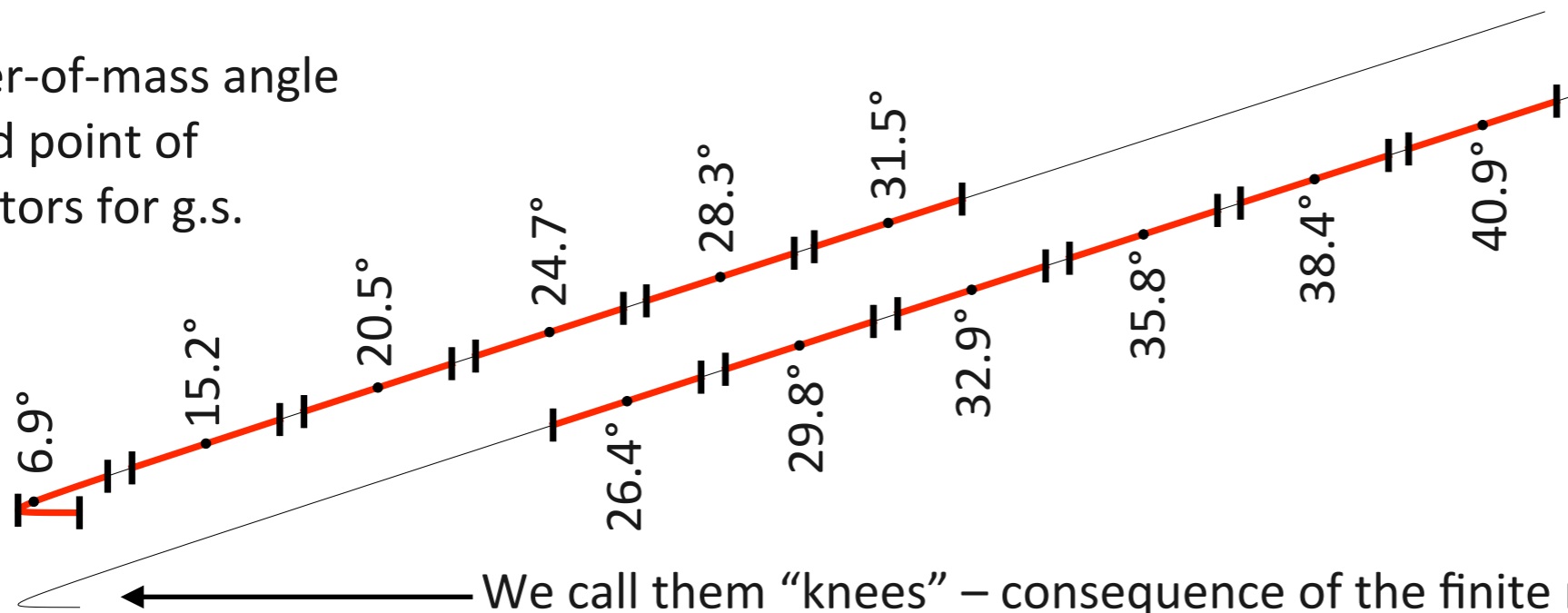


# What we measure

As measured:  $^{136}\text{Xe}(d,p)$  at 10 MeV/u with a 2-T field



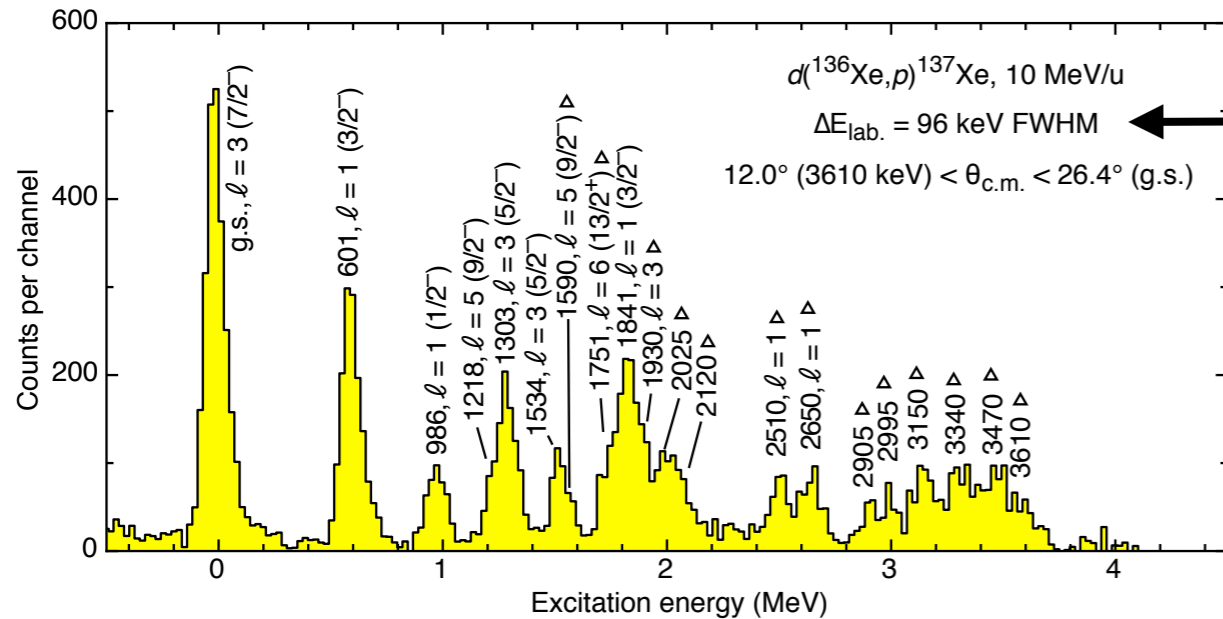
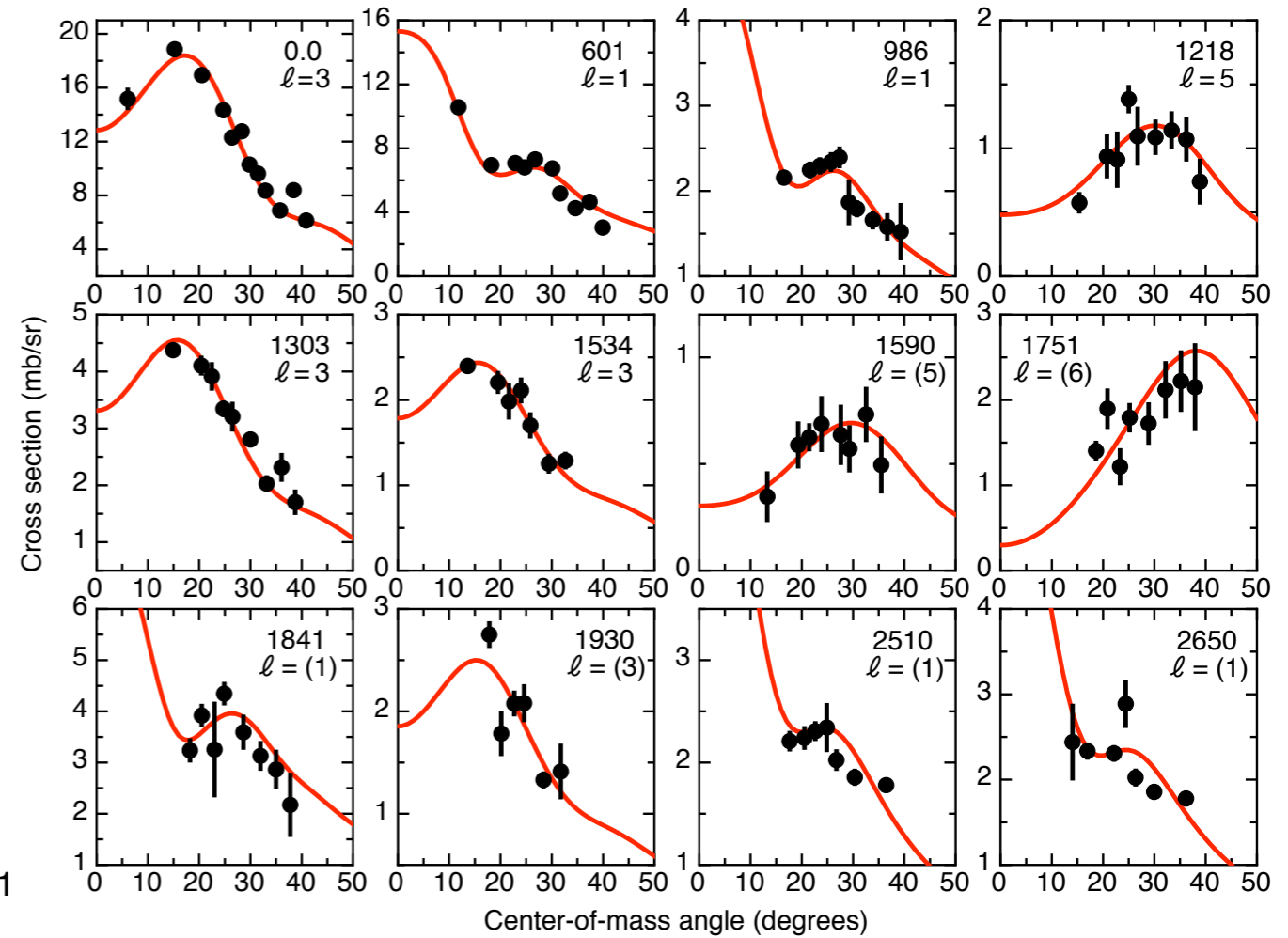
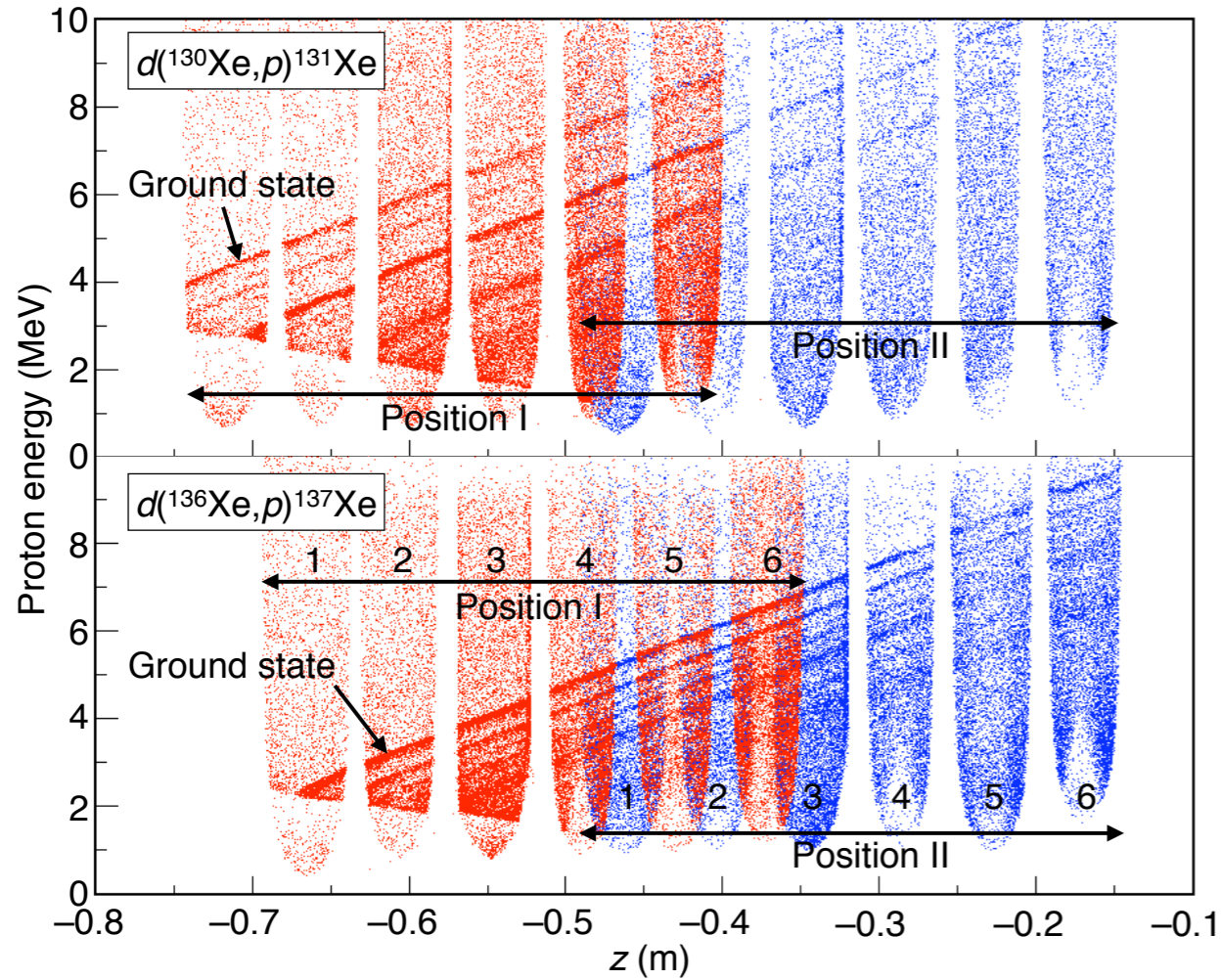
Center-of-mass angle  
at mid point of  
detectors for g.s.



Example signal-RF  
spectrum from  
Position I, detector 6.  
Here we observe  
protons executing 1, 2,  
and 3 complete orbits.



# Detailed spectroscopy of $^{137}\text{Xe}$ ( $^{131}\text{Xe}$ )



Over a factor of 3 improvement compared to other approaches. Allowed for a detailed spectroscopic analysis, such as angular distributions for 12 states, including  $9/2^-$  states and the lowest lying  $13/2^+$  state. Significant addition to prior knowledge of  $^{137}\text{Xe}$ .

Systematic: cross sections  $\pm 15\%$   
Relative spectroscopic factors  $\pm 15\%$

Previous work

$E$ (keV)	$\ell$	$J^\pi$	$\sigma(\theta)$ (mb/sr)	$C^2S$	$C^2S$ (1)	$C^2S$ (2)	$C^2S$ (3)
0.0	3	$7/2^-$	18.8, 15.2°	1.00	0.70	0.58	0.73
601	1	$3/2^-$	10.6 (11.8°)	0.55	0.41	—	0.40
986	1	$1/2^-, 3/2^-$	2.2 (16.5°)	0.37	0.13	—	0.27
1218	5	$9/2^-$	1.1 (33.3°)	0.46	—	—	—
1303	3	$5/2^-$	4.4 (14.9°)	0.23	—	—	—
1534	3	$5/2^-, 7/2^-$	2.2 (19.2°)	0.13	—	—	—
1590*	(5)	$9/2^-$	0.7 (32.5°)	0.25	—	—	—
1751	(6)	$13/2^+$	2.2 (37.9°)	0.89	—	—	—
1841	(1)	$3/2^-$	3.9 (24.9°)	0.31	—	0.30	0.18
1930*	(3)	$5/2^-, 7/2^-$	2.8 (17.8°)	0.11	—	—	—
2025*	(1,3)?	—	2.1 (19.7°)	0.22 / 0.16	—	—	—
2120*	(1,3)?	—	0.9 (19.4°)	0.10 / 0.06	—	—	—
2510*	(1)	$1/2^-, 3/2^-$	2.0 (22.6°)	0.20	—	0.13	0.15
2650*	(1)	$1/2^-, 3/2^-$	2.1 (22.1°)	0.17	—	—	—
~2900*	(1,3)?	—	0.8 (15.6°)	0.08 / 0.05	—	—	—
~2990*	(1,3)?	—	1.4 (21.1°)	0.17 / 0.05	—	—	—
~3150*	—	—	0.3 (35.1°)	0.12**	—	—	—
~3310*	—	—	0.3 (34.7°)	0.12**	—	—	—
~3470*	—	—	0.5 (34.4°)	0.18**	—	—	—
~3610*	—	—	0.4 (34.1°)	0.14**	—	—	—

\*Determined in this work

\*\*If assumed  $13/2^+$

(1) G. Kraus *et al.*, Z. Phys. A340, 340 (1991)

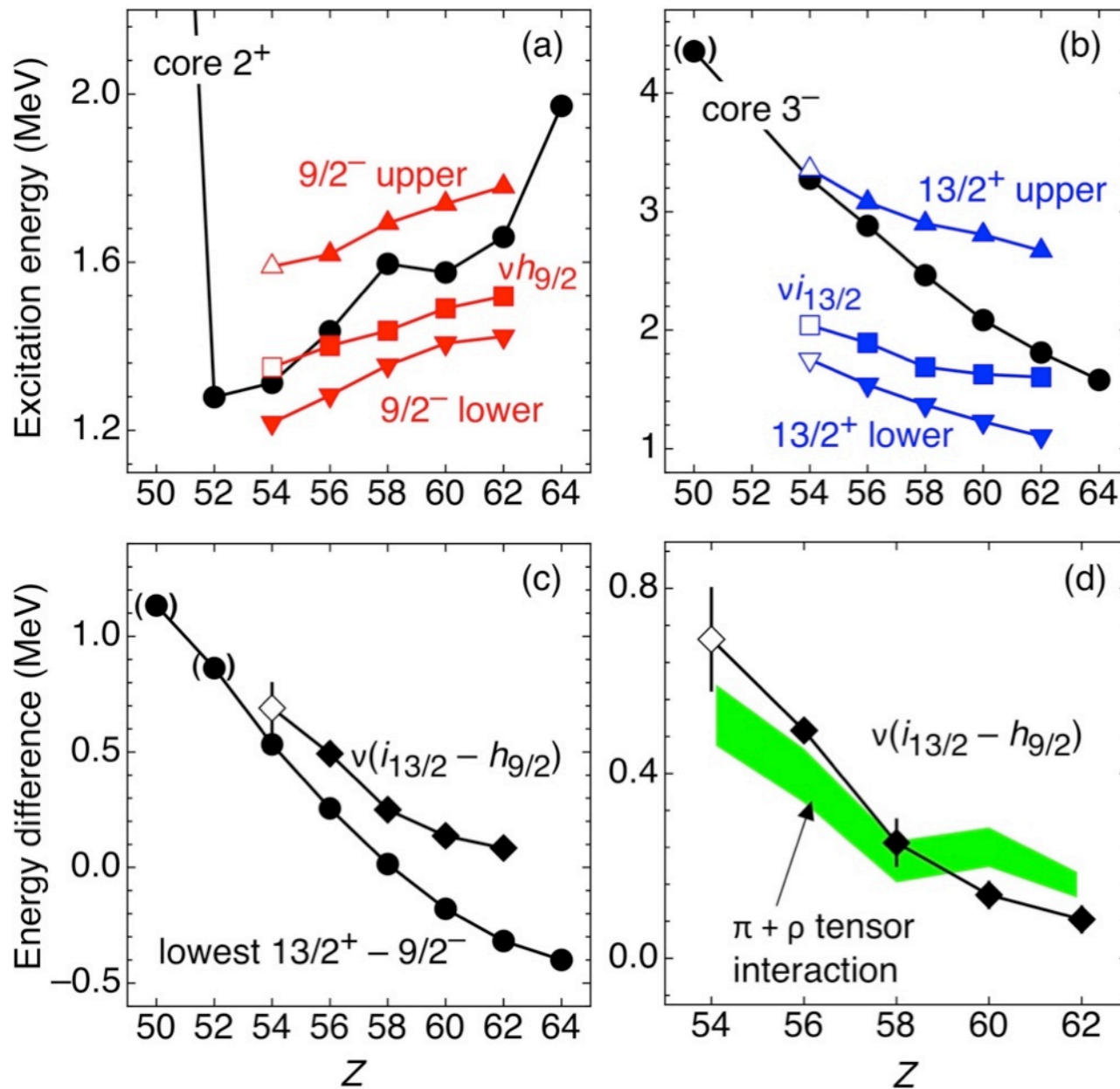
(2) E. J. Schneid and B. Rosner, Phys. Rev. **148**, 1241 (1966)

(3) P. A. Moore *et al.*, Phys. Rev. **175**, 1516 (1968)

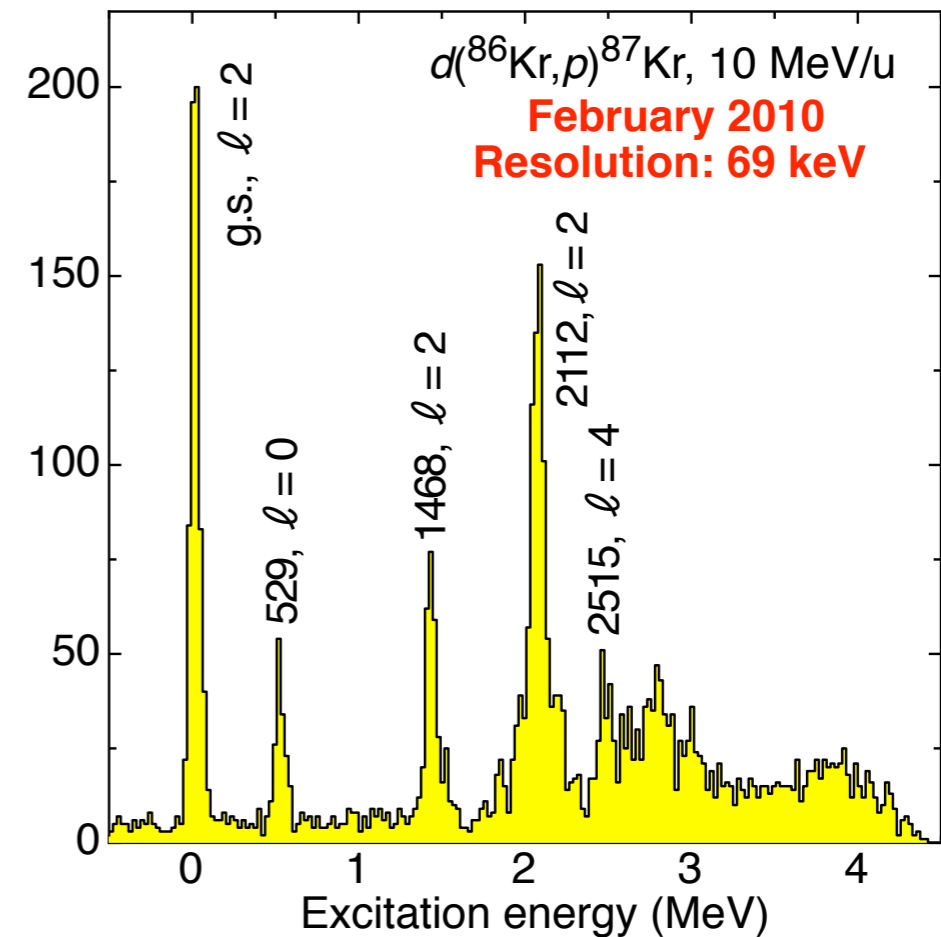
# Latest work ...

Extends our knowledge of single-particle trends outside  $N = 82$ . With CARIBU, measurements with  $^{134}\text{Te}$  and  $^{132}\text{Sn}$  soon.

Achieved a resolution of  $\sim 70$  keV in a recent measurement with a stable, lower  $Z$  beam ...  
[Part of thesis work by *D. K. Sharp*, Manchester]

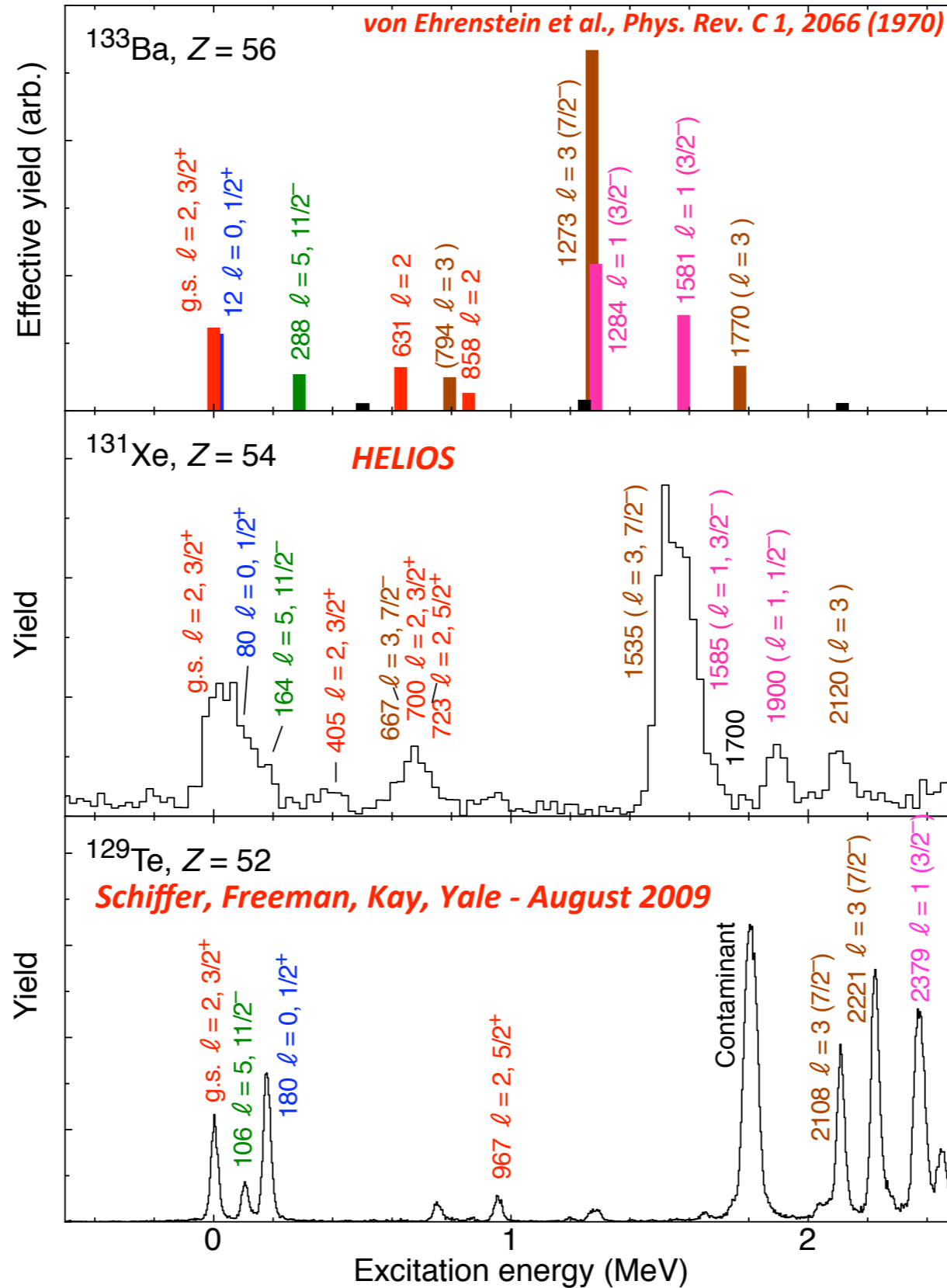


BPK, preliminary. Manuscript in preparation.



Part of an ongoing exploration of s.p. trends outside  $N = 50$ . Limited to the stable targets (beams) at Argonne (*in this region*)

# $d(^{130}\text{Xe}, p)$ and the $N = 76$ isotones





# *Radioactive beams*

CARIBU opens up an interesting region to study many characteristics of nuclei. HELIOS will be an excellent tool to perform measurements with these beams.

Clear we need to go further – facilities such as SPIRAL2 and HIE-ISOLDE offer a *wider range* of beams and potentially *orders of magnitude* more intensity.

So what are the beam and infrastructure requirements for a HELIOS-like device?

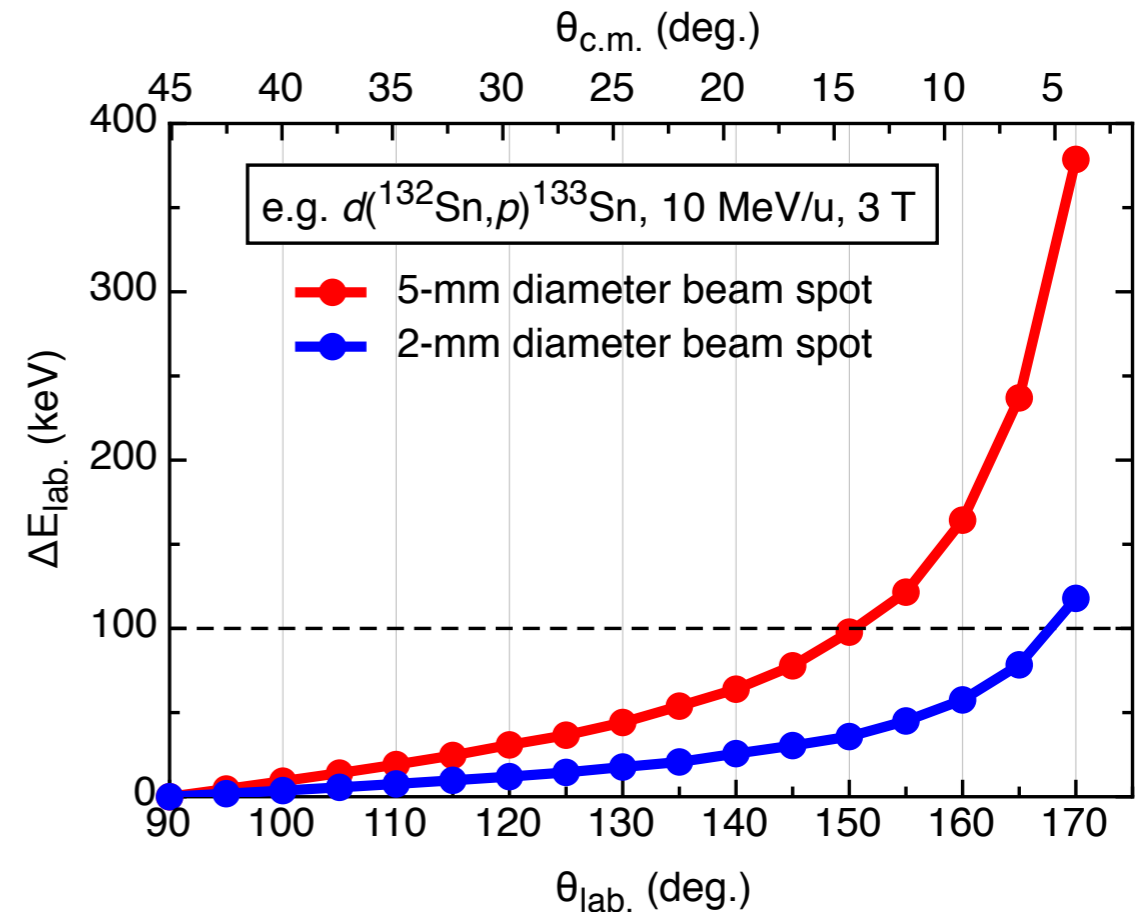
# Beam and infrastructure requirements, I

Q-value resolution set by longitudinal emittance. A 2 MeV energy spread for a 1 GeV mass 100 beam implies 20 keV contribution to q-value resolution.

Beam spot of less than 5 mm contributes less than  $\sim 100$  keV at angles of interest.

Current Si array has an acceptance of  $\sim 10 \pi$ .mm.mrad. For the HIE-LINAC\* a normalised transverse emittance of  $< 0.3 \pi$ .mm.mrad with a  $\beta\gamma \sim 0.15$  for 10 MeV/u beams implies an actual beam emittance of  $\sim 2 \pi$ .mm.mrad. More than sufficient.

Beam purity – with adequate recoil detection (technique depends on beam), impure beams can be tolerated.



# Beam and infrastructure requirements, II

RF period of the beam **needs to be ~100 ns**. (At ATLAS it is 82 ns.)

Time resolution of ~1-2 ns can be tolerated (as demonstrated).

*(many other measurements would benefit from a RF structure on the order of 100 ns e.g. in-beam gamma-ray spectroscopy, RDT, ..., isomer tagging, neutron detection TOF, ...)*

Species	T <sub>cyc.</sub> (ns)		
	1 T	2 T	3 T
p	65.6	32.8	21.9
d	131.2	65.6	43.7 (=2p)
t	196.4	98.2	65.6 (=3p)
<sup>3</sup> He	98.2	49.1	32.7 (=3/2p)
<sup>4</sup> He	131.2	65.6	43.7 (=d,2p)
<sup>12</sup> C (6 <sup>+</sup> )	131.2	65.6	43.7 (=d,2p)

## Other:

The current footprint of the HELIOS setup, including space for recoil detectors is ~6.5 × 4 m<sup>2</sup>

Helium consumption (at Argonne) is about **250 litres per month**. Additional losses when powering up and down (~500 litres for a complete cycle).

# Future challenges

Build on initial successes:

Reactions with the array in a forward position e.g.  $(d, {}^3\text{He})$ ,  $(d, t)$  (*spring 2011*)

Development of a gas-cell target of  $\alpha$  and  ${}^3\text{He}$  induced reactions (*in progress*)

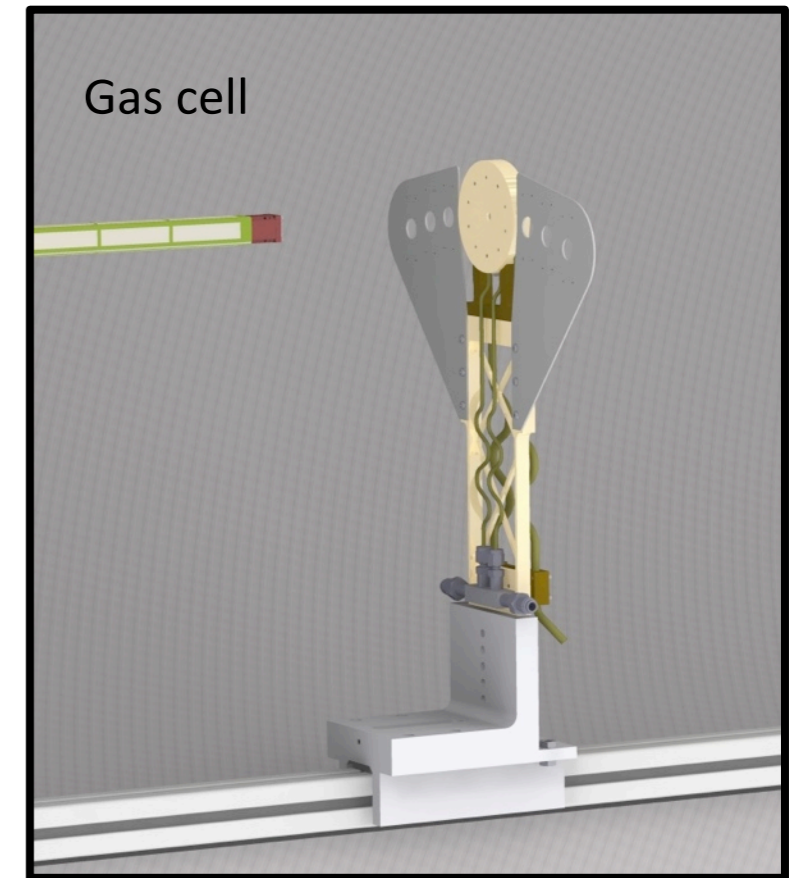
Replace the prototype array with a new, large-acceptance, 6-sided (10-sided) modular position-sensitive Si array (*in progress*)

CARIBU beams – *imminent*

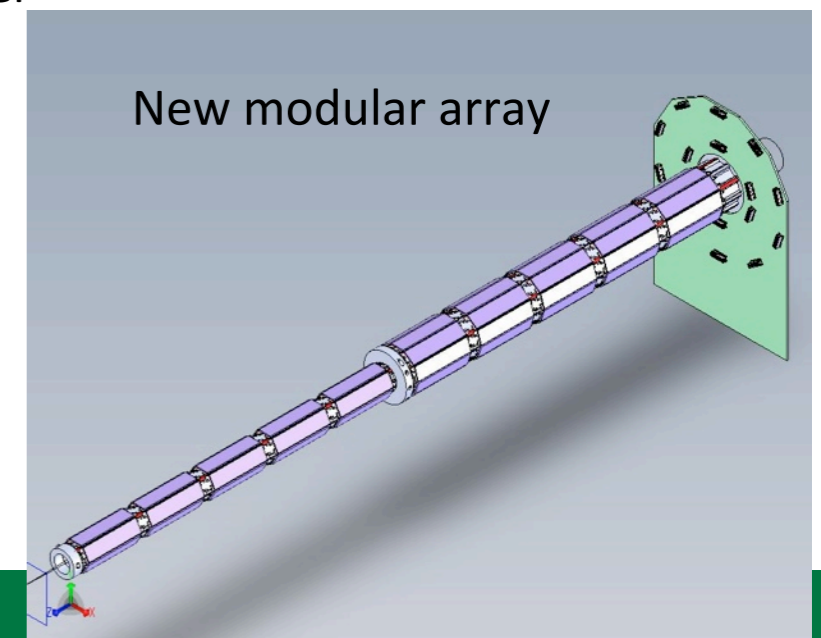
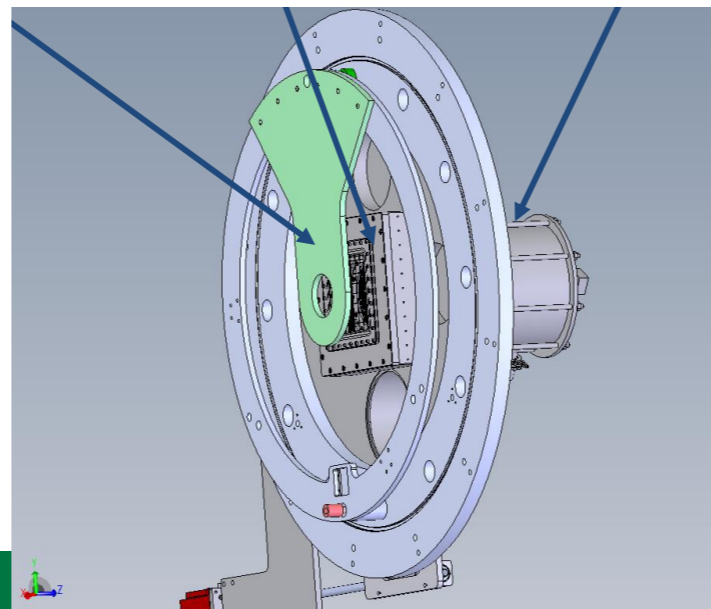
Utilise Manchester-built gas-ionisation chamber for recoil detection

Gamma-ray detection?

... and more ...



Array support PPAC Bragg counter





# Acknowledgements



## *The HELIOS collaboration\**

N. Antler<sup>1</sup>, B. B. Back<sup>1</sup>, S. I. Baker<sup>1</sup>, J. A. Clark<sup>1</sup>,  
C. M. Deibel<sup>1</sup>, B. J. DiGiovine<sup>1</sup>, S. J. Freeman<sup>2</sup>,  
N. J. Goodman<sup>3</sup>, Z. Grelewicz<sup>1</sup>, S. Heimsath<sup>1</sup>,  
C. R. Hoffman<sup>1</sup>, B. P. Kay<sup>1,4</sup>, H. Y. Lee<sup>1</sup>, C. J. Lister<sup>1</sup>,  
S. T. Marley<sup>1,3</sup>, P. Mueller<sup>1</sup>, R. Pardo<sup>1</sup>, K. E. Rehm<sup>1</sup>,  
A. M. Rogers<sup>1</sup>, J. Rohrer<sup>1</sup>, J. P. Schiffer<sup>1</sup>, D. Shetty<sup>3</sup>,  
J. Snyder<sup>3</sup>, M. Syrion<sup>1</sup>, J. C. Lighthall<sup>1,2</sup>, A. Vann<sup>1</sup>,  
J. R. Winkelbauer<sup>1,3</sup>, A. Woodward<sup>1</sup>, A. H. Wuosmaa<sup>3</sup>

<sup>1</sup>*Argonne National Laboratory, USA*

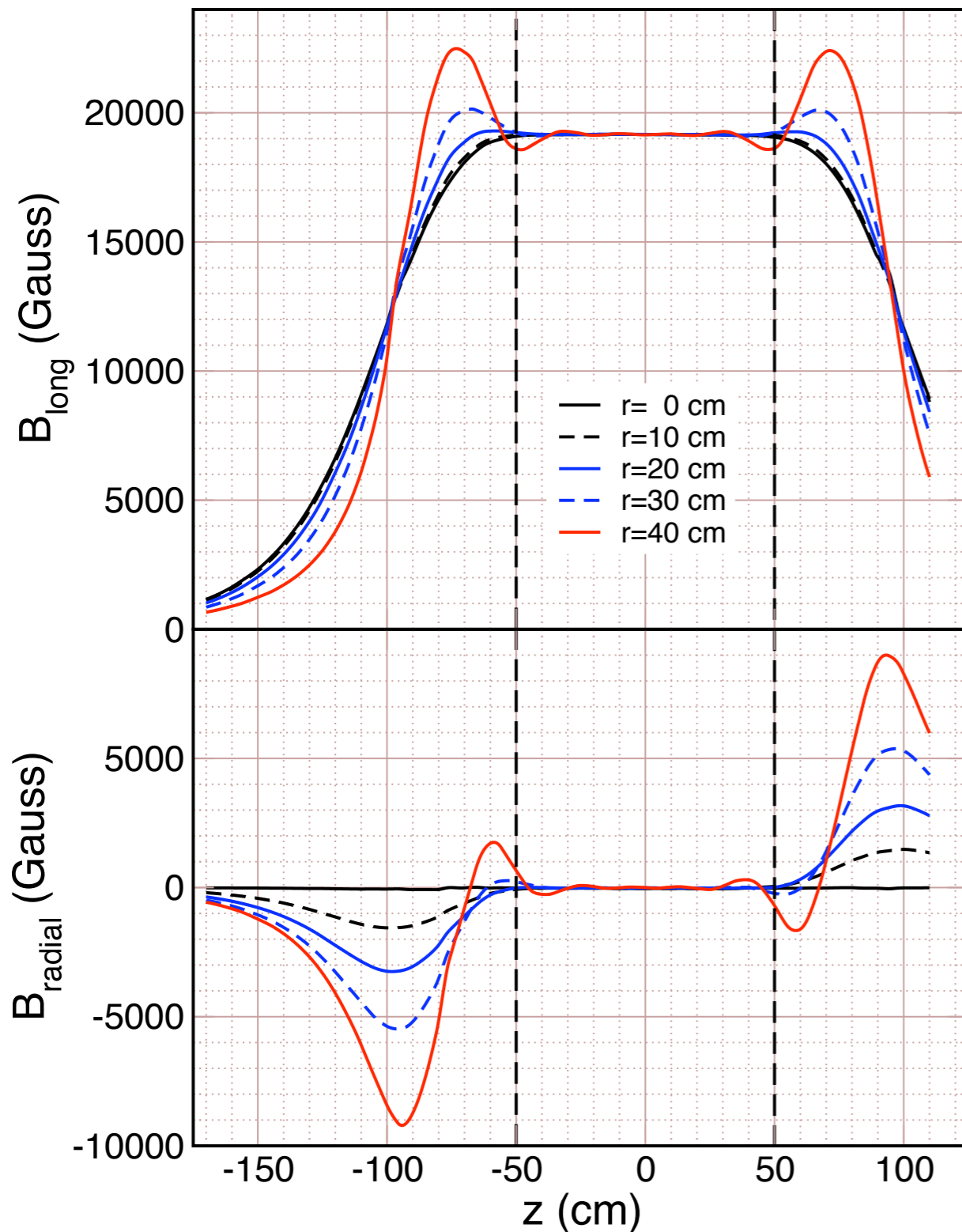
<sup>2</sup>*University of Manchester, UK*

<sup>3</sup>*Western Michigan University, USA*

<sup>4</sup>*University of York, UK*

***Spare***

# HELIOS field map



A 21,240-point field map was measured  
(J. C. Lighthall, J. Winkelbauer)

