

Machine learning to denoise pulses from a p-type point contact germanium detector

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Overview

1. Introduction

- Rare event searches using HPGe detectors
- Problem statement and motivation
- Autoencoders, denoising autoencoders

2. Methodology

- Overview of our detector
- Datasets (simulated and real)
- Training procedures

3. Results

- Denoising performance on simulations
- Verification with real detector data

4. Conclusions and future work

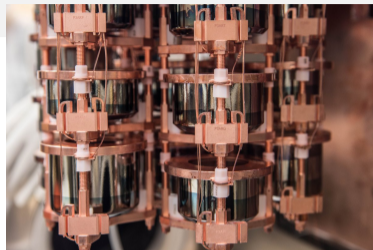
- Broadly applicable to particle astrophysics

Largely a summary of paper:

arXiv:[2204.06655](https://arxiv.org/abs/2204.06655)

Introduction

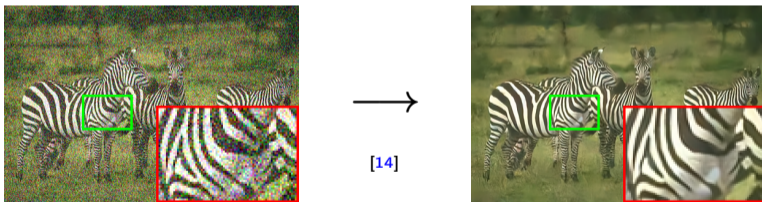
- High-purity germanium (HPGe) detectors widely used in rare event searches
 - Neutrinoless double-beta decay^[2, 3, 4]
 - Dark matter searches^[5, 6]
 - Other beyond Standard Model physics^[7, 8, 9]
- Adaptive denoising techniques are not typically applied to data ^[10]
- Noise removal could help advance searches for certain rare event interactions
 - Identify low-energy signal events that would otherwise be dominated by electronic noise
 - Relevant for solar axions, Pauli Exclusion Principle violation, electron decay, etc.^[11]
 - Improved background rejection based on pulse shapes
 - For example, slow energy-degraded pulses^[12, 13]
 - More accurate measurements of pulse amplitudes → better energy resolution



[10]

Introduction

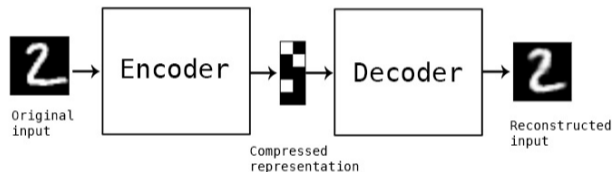
- Deep learning is frequently used to remove noise in other fields
 - Often outperforms traditional denoising methods (e.g., moving average, Savitzky-Golay filtering, wavelet thresholding)
- Most applications are to 2-dimensional images



- Will show that deep learning is also very effective at denoising 1-dimensional electronic signals (demonstrated specifically on HPGe detector pulses)

Autoencoders

- An autoencoder is an algorithm used to learn a useful representation of data
 - Trained to map the inputs to the inputs (with some form of constraint)



- By definition, an autoencoder is *lossy*
 - The goal is to retain as much useful information as possible
- Typically a neural network
- Widely used for dimensionality reduction, anomaly detection, and generative modelling

Autoencoders

- Components include an encoder, f_{θ} , and decoder, $g_{\theta'}$

Internal/latent representation, y , is obtained by applying encoder to input:

$$f_{\theta}(x) = y$$

Input reconstruction, z , is obtained by applying decoder to latent representation:

$$g_{\theta'}(y) = g_{\theta'}(f_{\theta}(x)) = z$$

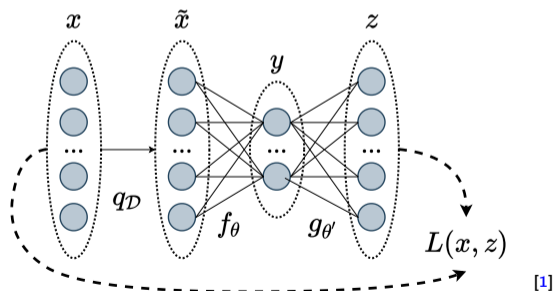
Minimize some **loss function** quantifying the reconstruction of x , $L(x, z)$, to train the autoencoder

- e.g., mean squared error

$$L(x, z) = \frac{1}{N} \sum_i^N \|z_i - x_i\|_2^2$$

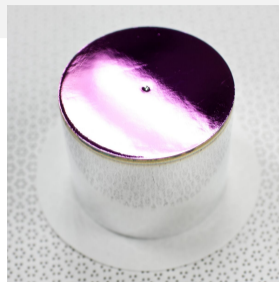
Denosing autoencoders

- Denoising autoencoders impose the constraint that reconstruction must also remove noise
 - Proposed as a method to extract robust features^[16]
 - Purpose was not for denoising, but to learn a better representation for classification tasks
 - Input becomes a corrupted version of x , \tilde{x} , by some process $q_{\mathcal{D}}$

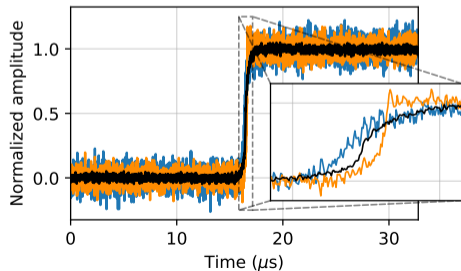


The detector

- Signals are from a 1 kg p-type point contact detector located at Queen's University
 - Cylindrical with a radius of 3 cm and height of 5 cm
 - Manufactured by ORTEC/AMTEK
 - Operated in a PopTop cryostat



[17]



[1]

- Each signal is a sequence of voltages sampled at a fixed interval
 - Recorded with 16-bit digitizer
 - Observed noise levels after preprocessing reflect energy of pulse

Datasets: real detector data

Americium-241 source

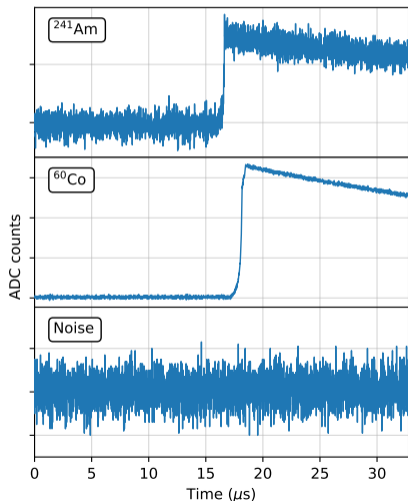
- Produces 60 keV γ s
- Almost entirely single-site events

Cobalt-60 source

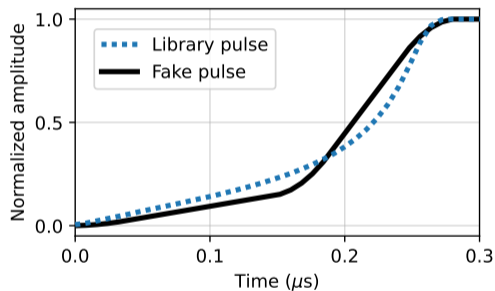
- Produces 1173 keV and 1332 keV γ s
- Numerous multi-site events from Compton scatters

Detector noise

- Collected by randomly triggering the detector
- Signals filtered to remove actual events that occur in the same trigger window



Datasets: simulated data



[1]

Library pulses

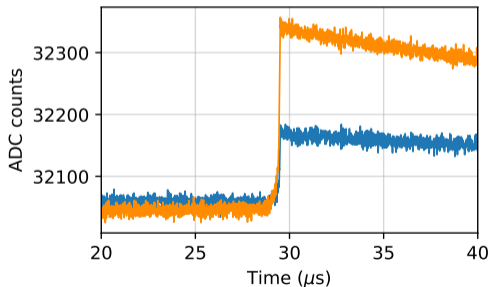
- 1724 simulated “library” pulses^[18]
- Each pulse corresponds to point on $1\text{ mm} \times 1\text{ mm}$ azimuthally symmetric grid
- Created using `siggen` simulation software^[19]
- Used to infer position of real events

Fake pulses

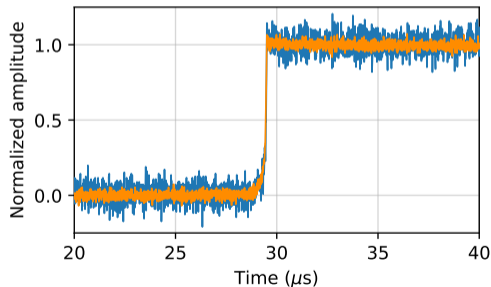
- Generated with piecewise functions
- Mimic the general shape of the library pulses without the requirement of complex physics simulations

Preprocessing

- Data pulses preprocessed to remove baseline
 - Data pulses have exponential decay removed with pole zero correction
 - Data pulses scaled by amplitude (calculated with a trapezoidal filter)
 - *Simulated pulses do not require this preprocessing* (already amplitude normalized)
- } **Amplitude normalization**



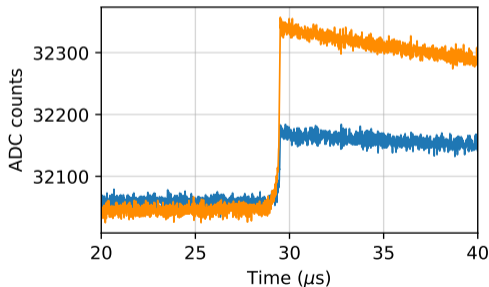
Before preprocessing



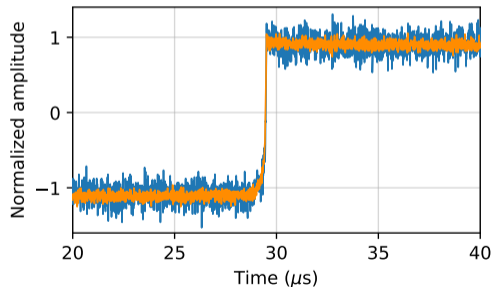
After preprocessing

Preprocessing

- Data pulses preprocessed to remove baseline
 - Data pulses have exponential decay removed with pole zero correction
 - Data pulses are mean-subtracted and scaled to have standard deviation of 0.5
 - *Simulated pulses only require last step*
- } **Standardization**



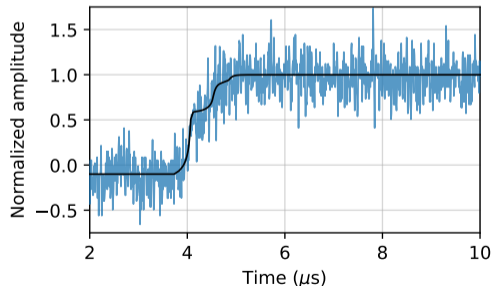
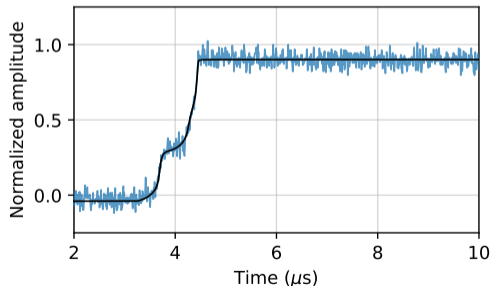
Before preprocessing



After preprocessing

Data augmentation

- From the simulated single-site event pulses, can create a diverse training set
 - Combine single-site simulated pulses to create *artificial multi-site events*
 - Apply random *horizontal shifts*, *vertical shifts*, and *amplitude scales* to each pulse
 - Add *detector noise* to each pulse with a random standard deviation



Training procedures

Regular

- Trained to map the noisy pulse to the corresponding clean underlying pulse
- Must know the true pulse – *only works on simulated data*

Noise2Noise^[20]

- Trained to map noisy pulse to *another* noisy pulse (different noisy realizations of same underlying pulse)
 - An impossible task in practice
 - Model will instead learn to predict the mean, given infinite different noisy realizations
- Can be used with either simulations or real data
 - For detector data, add even more noise to the already noisy pulse
 - Include a *total variation* penalty^[21] to original loss function L_0 to account for the noisy true mean
 - Penalize the absolute difference between given sample (j) and subsequent sample ($j + 1$) in pulse
 - Apply scaling factor λ to control weighting

$$L = L_0 + \frac{\lambda}{N} \sum_i^N \sum_j^{M-1} |z_{i,j+1} - z_{i,j}|$$

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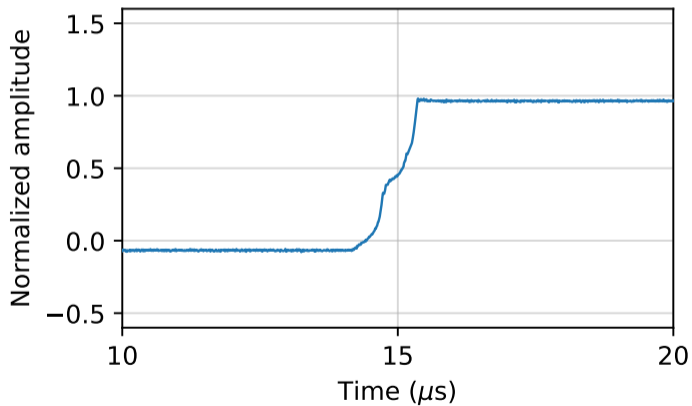
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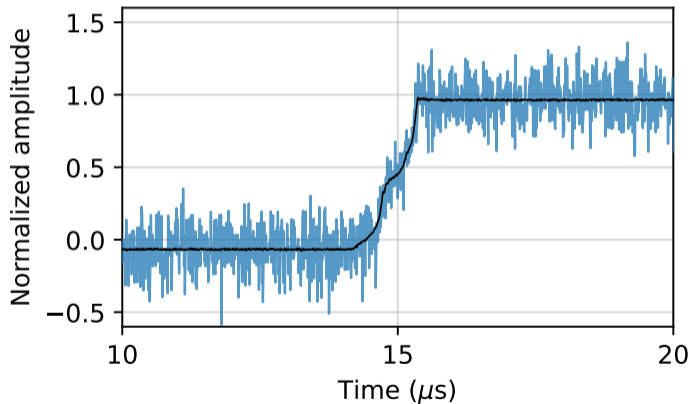
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Original data pulse



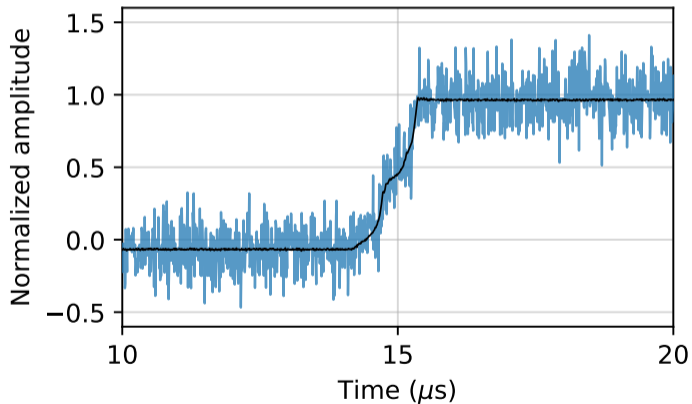
Noise2Noise

Original data pulse with a random noise pulse

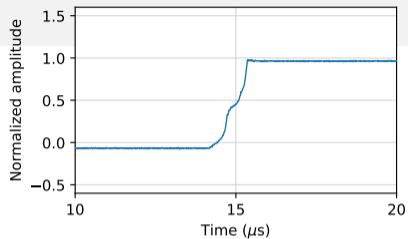


Noise2Noise

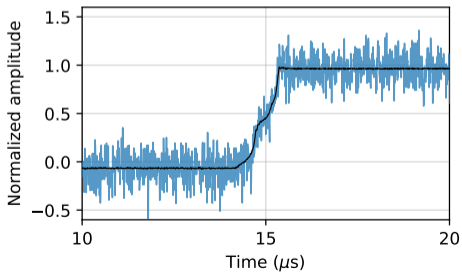
Original data pulse with *another* random noise pulse



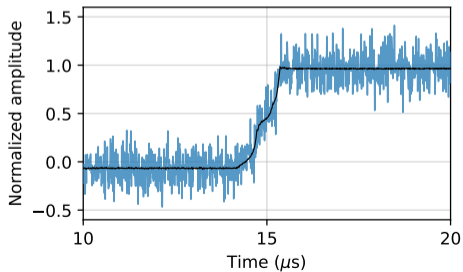
Noise2Noise



Input pulse

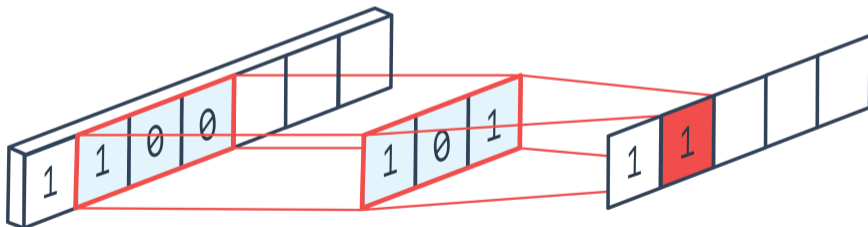


Target pulse



Network architecture

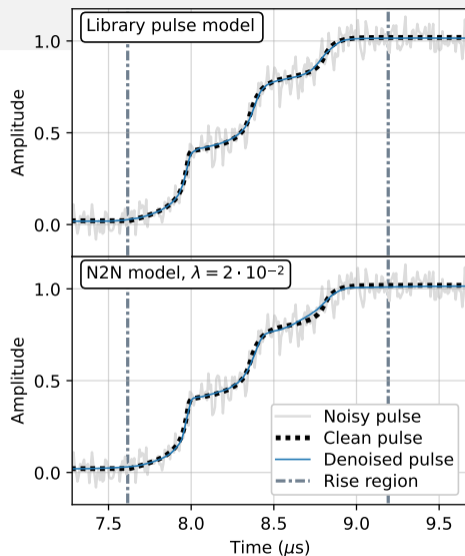
- Fully convolutional autoencoder
 - Weight sharing provides consistent noise removal across pulse
 - Feature locality and shift equivariance
 - Allows for a variable input shape (subject to some restrictions)
 - Significant reduction in the number of trainable parameters



[22]

Results: simulations

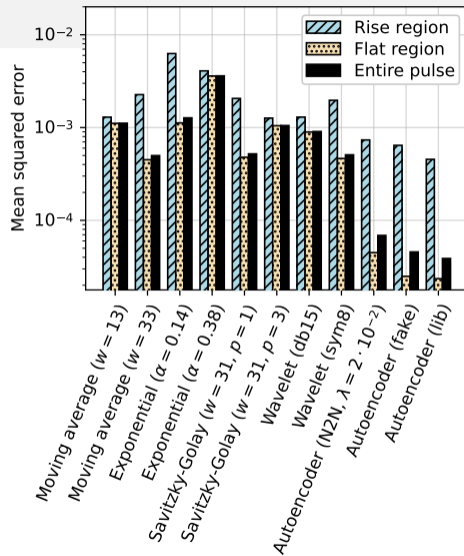
- Qualitatively, denoising with deep learning performs very well on simulations
- Autoencoder is superior to all traditional denoising methods investigated
 - Compared mean squared error on test set containing simulated single- and multi-site events
 - Each method optimized on a separate validation set to select hyperparameters
- Regular training procedure (with simulations) outperforms Noise2Noise procedure (with ^{60}Co data)
 - Still very good performance with Noise2Noise



[1]

Results: simulations

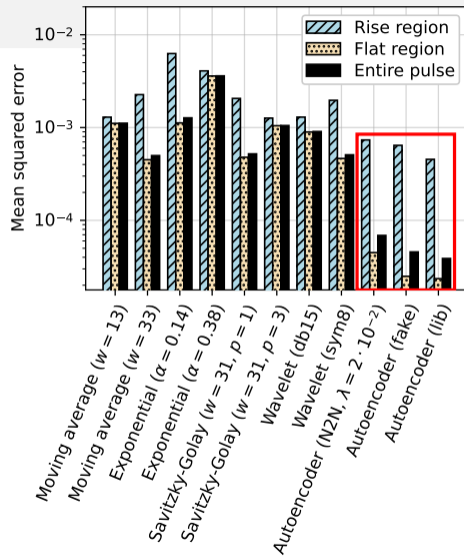
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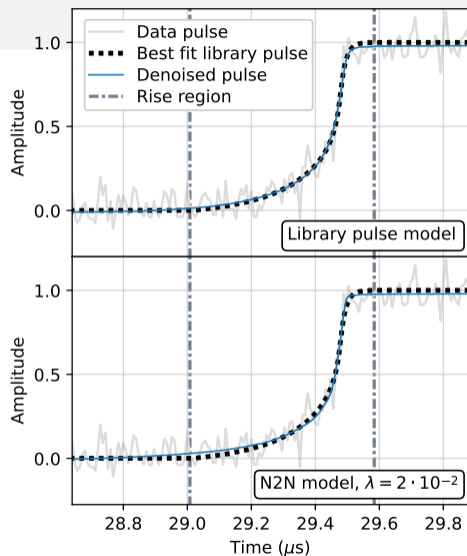
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Results: data

- Qualitatively, denoising with deep learning performs very well on data
- More difficult to quantify denoising
 - No true underlying pulse to compare to
- However, can make a statistical comparison to evaluate the performance



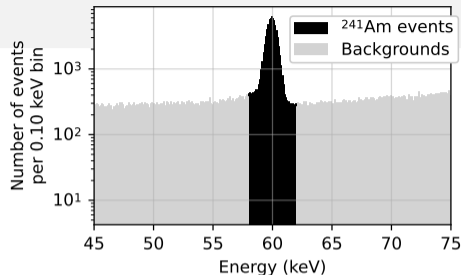
[1]

Results: data (χ^2 comparison)

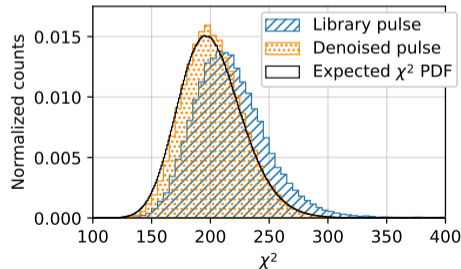
- ^{241}Am dataset contains mostly single-site events from 60 keV γ s
- Use a χ^2 comparison between the original pulse and **denoised pulse**, **best-fit library pulse**

$$\chi^2 = \sum_{i=1}^M \frac{(z_i - x_i)^2}{\sigma_i^2}$$

- χ^2 distribution between noisy and **denoised pulse** is consistent with expected χ^2 distribution of our detector noise
 - Taken over 200 samples, contains rise region



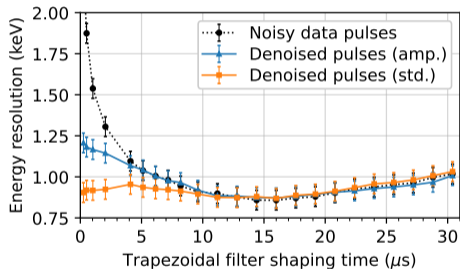
[1]



[1]

Results: data (energy resolution)

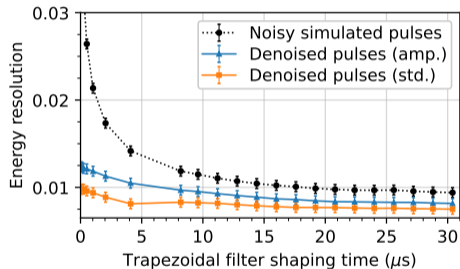
- Can also evaluate the effect of denoising on the energy resolution
 - Energy calculated from the amplitude of a trapezoidal filter
 - FWHM of 60 keV peak is the energy resolution
- Optimal energy resolution is comparable before and after denoising
- Much lower shaping time required to achieve good energy resolution
 - Especially true for **standardized preprocessing**
 - Important for data storage, analysis, etc.



[1]

Results: data (energy resolution)

- Create simulated dataset to mimic ^{241}Am γ s
 - Single-site library events with detector noise
 - Same noise level as ^{241}Am dataset
- Observe same effect as with real data
 - Comparable energy resolution with lower shaping time
 - Best results with **standardized preprocessing**
- Observe different patterns
 - Optimal energy resolution improves
 - Energy resolution continues decreasing up to the limit of the shaping time
- Under ideal conditions, denoising allows for an improvement in energy resolution



[1]

Conclusions

- Deep convolutional autoencoders are effective at removing noise from HPGe detector pulses
 - Outperforms various traditional denoising methods
 - Denoised pulses are statistically consistent with data pulses
 - Can reach optimal energy resolution with a lower shaping time
 - Can improve the optimal energy resolution under some circumstances
- Models can be trained without the need for detailed detector simulations
 - “Fake” pulses are a very rough approximation to library pulses
 - Noise2Noise method requires *only* noisy detector data
 - Results could likely be improved with more data

Conclusions

- Results presented here are focused on HPGe detector data
 - Noise removal is beneficial in many contexts
 - Our group is applying these methods to signals from other detector technologies
 - Gaseous proportional counters, bubble chambers
- The encoder portion of the network can also be used for tasks
 - The denoising autoencoder forces the encoder to learn a robust representation of the data
 - Exploring the use of the encoder output for clipping restoration, peak finding, and single-/multi-site event discrimination
- Work is broadly applicable to the particle astrophysics community
 - Great potential to be expanded on
 - See poster from [Tianai Ye](#), talk from [Noah Rowe](#)

Thank You!

More details contained in the paper. Check it out!

arXiv:[2204.06655](https://arxiv.org/abs/2204.06655)

Acknowledgements

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 - V. Basu, R. D. Martin, C. Z. Reed, N. J. Rowe, M. Shafiee, T. Ye
- This work was supported by the Natural Sciences and Engineering Research Council of Canada (NSERC)
- The author held a Walter C. Sumner Memorial Fellowship



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Accessed: 2022-05-19.

Additional Slides

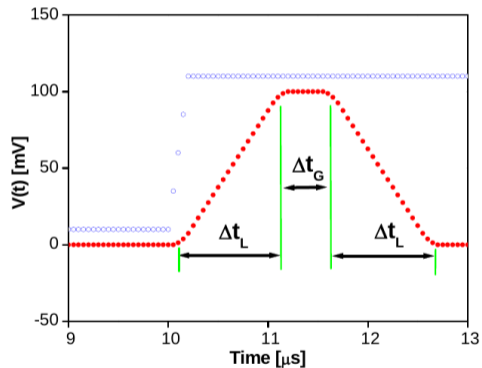
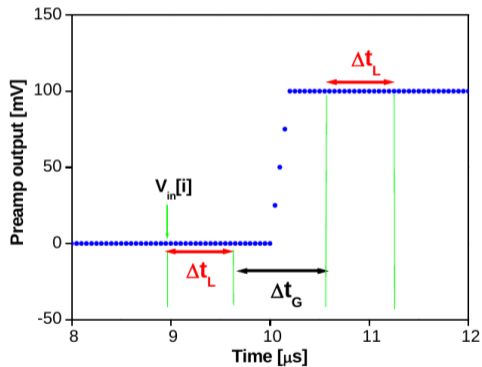
Model architecture

Layer	Stride	Window	Output
Input			4096, 1
Convolution	1	1	4096, 8
Convolution	1	9	4088, 16
Average Pooling	2	2	2044, 16
Convolution	1	17	2028, 32
Average Pooling	2	2	1014, 32
Convolution	1	33	982, 64
Average Pooling	2	2	491, 64
Convolution	1	33	459, 32
Transpose Convolution	1	33	491, 32
Upsampling	2	2	982, 64
Transpose Convolution	1	33	1014, 64
Upsampling	2	2	2028, 64
Transpose Convolution	1	17	2044, 32
Upsampling	2	2	4088, 32
Transpose Convolution	1	9	4096, 16
Convolution (output)	1	1	4096, 1

Results on simulations

Training procedure and data			Mean squared error ($\times 10^{-5}$)			
			Gaussian noise		Detector noise	
Procedure	Data	Noise	Lib	Fake	Lib	Fake
Regular	Library	Detector	4.37	4.94	3.88	4.26
Regular	Library	Gaussian	3.41	3.85	4.49	4.80
Regular	Fake	Detector	5.43	4.55	4.57	3.83
Regular	Fake	Gaussian	3.85	3.35	4.92	4.29
N2N ($\lambda = 0$)	Library	Detector	4.13	4.65	4.03	4.41
N2N ($\lambda = 0$)	Library	Gaussian	3.45	3.85	4.52	4.79
N2N ($\lambda = 0$)	Fake	Detector	4.87	4.08	4.62	3.87
N2N ($\lambda = 0$)	Fake	Gaussian	3.84	3.39	4.92	4.33
N2N ($\lambda = 0$)	Detector	Detector	9.28	9.10	10.02	9.97
N2N ($\lambda = 2 \cdot 10^{-2}$)	Detector	Detector	5.81	6.05	6.88	6.99

Trapezoidal filter



[23]

Noise curve

