Neutrino Physics and Beyond at T2K and Hyper-Kamiokande



Mark Hartz For the Hyper-K Canada Collaboration CAP Congress, PPD Symposium, June 7, 2022



Neutrinos & Their Oscillations

- Neutrinos are neutral and interact via the weak force only
- Flavour and mass states mix, leading to the phenomenon of neutrino oscillations
- From these oscillations, we know that at least two of the neutrinos have non-zero mass



Produce neutrinos as weak eigenstates



Standard Model of Elementary Particles

Interact as weak eigenstates

Three mixing angles: θ_{12} , θ_{13} , θ_{23} Two mass splittings: Δm²₃₂, Δm²₂₁ **One CP violating phase:** δ_{cp}

Potential new source of CP violation!



The T2K Experiment



$$P_{\mu \to \mu} = 1 - \left[\sin^2 2\theta_{23} - \sin^2 \theta_{23} \cos 2\theta_{23} \sin^2 2\theta_{13} \right] \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E_{\nu}} \right) + \frac{1}{2} \left[\frac{\Delta m_{32}^2 L}{4E_{\nu}} \right] + \frac{1}{2} \left[\frac{\Delta m_{3$$

$$P_{\mu \to e} = \overline{\sin^2 \theta_{23} \sin^2 2 \theta_{13}} \sin^2 \left(\frac{\Delta m_{31}^2 L}{4E_{\nu}} \right) \mp \frac{\sin 2 \theta_{12} \sin 2 \theta_{23}}{2 \sin \theta_{13}} \sin^2 2 \theta_{13} \sin \left(\frac{\Delta m_{31}^2 L}{4E_{\nu}} \right)$$

sign flips for antineutrinos







T2K Data



- T2K has been collecting data since 2010 in both neutrino and antineutrino beam modes
- Results shown here correspond to most of the data collected





CP Asymmetry Measurement at T2K



T2K will continue to operate with upgrades to the near detector and beam power

How do we get orders of magnitude more data?



- Much lower antineutrino rate is due to experiment being made of matter (to first order)!
- Statistical errors still dominate, need more data!
- Systematic errors will be become more important with more data











Hyper-K Experiment

- 260 kton detector with fiducial mass is 8x larger than Super-Kamiokande
- Neutrino beam from J-PARC will be 2.5 times more intense (1.3 MW proton beam)
- New photon detectors and near detectors
- 20x the (anit)neutrino rate compared to T2K experiment
- Broad physics program includes
 - Accelerator neutrinos
 - Proton decay searches
 - Supernova neutrino detection
 - Atmospheric neutrino detection
 - Solar neutrino detection
 - Dark matter searches...
- Approved in 2020, planned start of operation in 2027

Hyper-K Detector









Hyper-K Collaboration

20 countries, 99 institutes, ~500 people as of Jan 2022, and growing

▲					Number of Co
Europe	281 members	Asia	149 members	500	
Armenia	3	India	8	450	
Czech	4	Korea	18	400	
France	27	Japan	123	350	212 212
Germany	I	Americae	52 morehour	300	287
Italy	55	Americas	52 members	250	248 (14)
Poland	38	Brazil	3	200	(13 countries)
Russia	22	Canada	32	200	212 231
Spain	35	Mexico	8	150	165
Sweden	5	USA	9	100	
Switzerland	13	Africa	II members	50	
Ukraine	4	Morocco	П	0	2015 2016 2017 201
UK	74				Year

- Hyper-K is a large international and growing collaboration
- Hyper-K Canada is open to new collaborators!





The Hyper-K Canada group was formed in 2018 and is supported by funding from NSERC and CFI

Construction Progress in Japan



Excavation of the detector cavity, one of the world's largest underground cavities, will begin in October of this year.



Proton Decay at Hyper-K

Baryon number violation is generic prediction of GUTs

channels





With its 187 kton fiducial mass, Hyper-K will have the largest mass with sensitivity to most



Supernova Neutrinos at Hyper-K





- There is a background of supernova neutrinos from all past supernovas
- Probes history of heavy element synthesis in stars
- Due to its large mass, Hyper-K will improve on sensitivity of SK-Gd





54k-90k events for 10 kpc distant supernova

- ~10 neutrino events for supernova in Andromeda
- Neutrino-electron scattering introduces pointing capability
- 1.0-1.3 degree accuracy for 10 kpc distant supernova







CP Violation at Hyper-K



- Hyper-K will observe ~2000 electron neutrino and electron antineutrino candidates each
- 3% statistical error on the CP violation measurement is achieved •
- **Controlling systematic errors is critical** •







Systematic Uncertainties

- CP violation measurement becomes systematic error dominated
- Uncertainty on neutrino interaction modeling to be addressed with Intermediate Water Cherenkov **Detector (IWCD)**
- Uncertainty on detector modeling to be addressed with new calibration systems and techniques

- To make precision parameter measurements, also need to know energy scale to 0.5% level
- Current error in Super-K is ~2%









12

The Intermediate Water Cherenkov Detector

- Intermediate detector for Hyper-K
- Located ~1 km from neutrino source
 - Land acquisition is primary focus of host lab at this time
- 600 ton water Cherenkov detector
- Position can be moved to different off-axis angles
- Using new high resolution multi-PMT photon detectors
- Primary physics:
 - Electron (anti)neutrino interactions (with 1% of beam)
 - Measuring (anti)neutrino energy reconstruction

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Canadian led project

See talk by R. Akutsu at W1-1 Neutrino Experiments (PPD)

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Multi-PMT (mPMT) Photosensor

- 19 fast 3-inch diameter PMTs in a single module with readout electronics and HV integrated in the module
- LED light injection will allow mPMTs to be used as light sources in detector calibration
- Electronics development on 125 MSPS 20-channel FADC is well advanced
 - Working firmware for Xilinx FPGA •
 - One more prototype revision planned with final decision on shaping circuit
 - Development of mezzanine care for driving of LED • light injectors

Multi-PMT (mPMT) Photosensor Assembly

- Two methods for assembly being explored:
 - Apply gel to PMTs before assembly of module
 - Challenge achieving good optical contact with limited • tolerance to variation in parts dimensions.
 - Assemble PMTs, dome, support matrix and cylinder first, before applying gel
 - Challenge procedure has to work ~100% of the time since it • is a one-time application for each module
- See Talk by N. Booth in W2-6 Neutrino Experiment and Related Calibrations II (PPD)

See Poster by L. Koerich (G*) (POS-37) Toward a Veto Mechanism to Reduce Background for the Hyper-Kamiokande's Intermediate Water Cherenkov Detector

Prototype for ex-situ gelling at TRIUMF

Prototype for in-situ gelling at Carleton

Photogrammetry Calibration

- Imaging of detector to measure calibration source and photosensor positions
- Standard cameras and lenses with custom-built underwater vessels for fixed cameras
- Ongoing tests at the UBC pool to characterize the optics under water
- Prototype readout system has been developed and being tested at U. Winnipeg
- Photogrammetry simulation of detectors under development - choose the best location of cameras in detectors

See talk by R. Gaur at T4-3 New Directions in **Accelerator-Based Experiments: Future Experiments** - From Collider to neutrinos (PPD)

Prototype camera housing

Measurements at UBC Pool

Camera readout (under development)

Water Quality Monitoring

- Measure light attenuation and scattering in water samples •
- Use low-cost sub-ns LED driver being developed at University of Victoria •
 - LEDs down to ~270 nm allow for study of Raman Scattering •
- Applications for environmental monitoring NSERC alliance grant working with First Nations University of Canada (FNUniv), Cowessess First Nation, Ahtahkakoop Cree Nation, Weyburn Water Treatment Facility, water treatment engineering faculty at U.Regina

Shorter pipe for scattering measurement

LIGHT RECEPTION

Photosensor Calibration

- Ex-situ measurements of photosensor responses using facility at TRIUMF
- Light source can be scanned along PMT photocathode
- Measurements in air and water
- Magnetic field direction and magnitude can be varied
- Measurements of Super-K 20-inch PMTs show where improvements can be made to PMT model

See talk by V. Gousy-Leblanc at W2-6 Neutrino **Experiment and Related Calibrations II (PPD)**

18

Multi-PMT for Hyper-K Detector

- Replace 5 PMTs with LED light injector boards \bullet
- Collimated light with varying wavelengths, including UV to study Raman scattering
 - 278nm Raman scattering
 - 365nm Peak Cherenkov intensity
 - 415nm Peak quantum efficiency
 - 470nm Tail of light distribution
- Sub-ns LED driver+sub-ns PMT resolution+collimated light = reconstruction of scattering position
- Diffuse light used to characterize in-situ angular response of 20-inch PMTs
 - Large 20-inch PMTs are sensitive to residual magnetic field
 - Nearby multi-PMTs provide relative normalization

Hyper-K mPMT Design

Collimated light

Diffuse light

Machine Learning Event Reconstruction

- Machine learning plays a key role in analysis strategy:
 - Use all of the information provided by fine resolution measurements of mPMTs in IWCD
- Very fast event reconstruction allows for quick study of systematic parameter variation in detector calibration
- WatchMal group started by Hyper-K Canada leads studies of • machine learning applications for water Cherenkov detectors across the globe
- Generally seeing significant improvements in areas such as • classification compared to the likelihood-based fiTQun reconstruction

See talk by N. Prouse at W2-1 Machine Learning in **HEP and Novel Reconstruction Tools (PPD)**

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THE UNIVERSITY OF WINNIPEG

Test Experiment at CERN

- Operation of IWCD prototype called WCTE in CERN test beam (T9) in 2024
- Study detector calibration and response with known particle fluxes of 0.2-1.1 GeV/c π , p, e, μ
- Beam test data in 2022 to determine if secondary particle fluxes at low momenta are sufficient
 - Test muon/pion separation with Aerogel Cherenkov Threshold detector (ACT)
 - More data collection in low momentum beam configuration in July
- Have received recommendation from the CERN SPSC and Research Board Approval

CERN-SPSC-2020-005; SPSC-P-365: http://cds.cern.ch/record/2712416?ln=en

ACT Detector

EMPHATIC

- Collaboration on EMPHATIC hadron production experiment at Fermilab
- Pilot run in 2018 to measure p+C forward scattering
 - Analysis led by postdoc M. Pavin, submitted to PRC (arXiv:2106.15723)
- Initial Phase-1 operation in January, but Canada could not participate due to COVID situation
- Second Phase-1 run in June with Canadian Aerogel Ring Imaging Cherenkov (ARICH) detector for PID

See Poster by B. Ferrazzi (G*) (POS-32) A mirror study in an ARICH detector for a hadron production experiment

Model fit to

forward

scattering

data

Phase-1 Experimental Configuration

ARICH detector

Summary

- T2K has produced world-leading neutrino oscillation measurements
- The Hyper-Kamiokande experiment will allow us to probe neutrino properties with unprecedented precision amongst a broad physics program
- Construction of the experiment is proceeding on schedule in Japan. Start of detector cavity excavation soon!
- Canada is focussed on contributions to Hyper-K in order to meet the requirements for systematic uncertainty reduction
- Many areas of development in Canada to realize multi-PMT, photogrammetry calibration, water quality monitoring systems
- Exciting time for the Hyper-K project in Canada and overall. We are always looking for opportunities to grow the effort in Canada.

Thank you

Kamioka Water Cherenkov Experiments

- **Hyper-Kamiokande**
- ~2027 onwards

2015

<u>Super-Kamiokande</u>

2002

- 1996 onwards X 20
 - 50 kton (22.5 kton FV)

<u>Kamiokande</u>

- 1983 1996
- 3 kton

Backup Slides

Atmospheric Neutrinos

- neutrinos in the analysis
- Can maintain 5σ and 3σ sensitivity to CP violation discovery
 - unknown

If mass ordering is not determined before HK, HK will have sensitivity by including atmospheric

Combination of accelerator and atmospheric gives slightly improved sensitivity, even if MO is

Role of Near(Intermediate) Detectors

- of neutrino interactions
- Systematic errors don't cancel perfectly in the near-to-far extrapolation due to differences in the neutrino spectrum/flavour
- Important sources of systematic error:

Detect electron (anti)neutrinos at far detector. Uncertainty on the relative interaction cross section must be controlled

Must accurately model resolution function for energy reconstruction, including nuclear effects in neutrino-nucleus scattering.

Near detectors measure the neutrino rate/spectrum of neutrino production and properties

 $\sigma(\mathbf{v}_e)/\sigma(\mathbf{v}_\mu)$ $\sigma(\bar{\mathbf{v}}_{e})/\sigma(\bar{\mathbf{v}}_{\mu})$

Neutrino Detection

- Detect charged current scattering of (anti)neutrinos on nuclei
- Electron or muon in final state identifies flavour of parent (anti)neutrino
- In water Cherenkov detector, below threshold hadrons are not tracked - no full energy reconstruction
- Need accurate modeling of (anti)neutrino interactions but scattering on nuclei is difficult to model
- Hyper-K will have a suite of near/intermediate detectors to measure:
 - Properties of the (anti)neutrino beam
 - Properties of the (anti)neutrino scattering on nuclei

Challenges to Overcome

- Can measure electron (anti)neutrino cross section with 1% contamination in beam
 - Challenge: large **background** from beam induced external high energy gamma conversions (see T2K results to right)
 - Need to reduce this background
- Energy spectrum at near detector is different than far **detector** due to oscillations
 - Can't extrapolate near detector measurements perfectly
 - Nuclear effects -> large energy reconstruction error
 - Events with large energy mis-reconstruction can dominate some measurements (right)
 - Need direct measurements

T2K: Phys. Rev. Lett. 113, 241803 (2014) Data v_e CC interactions γ background μ background Other background 3.5 4.5

Reconstructed p_{e} (GeV/c)

29

NuPRISM Concept

high and low energy tails

Obtain very narrow spectrum

Due to pion decay properties, neutrino spectrum varies with offaxis angle

Measurements at different off-axis angle can subtract

Measure reconstructed energy of events

5% measurement precision on events with large misreconstruction

E_v (GeV)

Nuclear Effects in Neutrino Interactions

• We select neutrino interactions in the charged current interaction mode:

- Final state includes hadrons that may not be detected:
 - Water Cherenkov: most protons below Cherenkov threshold
 - Tracking detectors: neutrons may not be detected
- Need neutrino energy to calculate $P(\theta_{23}, \Delta m_{32}^2, ..., E_v)$
 - Since we don't fully reconstruct the event, we rely on models to infer energy

Detected charged lepton

Final state p, π, n, ...

