# Double scaling limits for field theory defects

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> Based on 2202.03472  $\bigoplus$  2206.09935 w/ Jorge Russo



## Studying defects in QFT is interesting for a number of reasons

- of gauge group...)
- May serve as a diagnose the phases of the theory
- Describe impurities coupled to the system

. . .

• Explore all operators in a QFT: extended operators may be sensible to finer details (e.g. topology of space, global properties

• If topological, they give rise to generalized symmetries

- However QFT (with/without defects) is hard...
- One strategy which has proved very successful is to look for small parameters on which one can expand. Celebrated examples include
  - the semiclassical approximation
  - large N
  - large spin
- In the recent past one new item added to the list
  - large charge sectors
- Inspired by this: can we access new information about defects in QFT???



## Contents

- Motivation
- limit
- Taking it over to Wilson lines in the k-symm product
- Defects in Wilson-Fisher
- Final comments

Local operators in N=2 SCFT's in 4d at large charge and a double scaling

## **Correlation functions in N=2** and large charge

- much so as to "trivialize"
- Some of them have holographic duals
- In particular, one can exploit SUSY to compute observables exactly

This includes correlators, defect operators and even the partition function itself (meaningful for 4d N=2)

• N=2 theories are interesting playgrounds to tinker with QFT: they have SUSY enough so as to constrain dynamics to accessible limits but not too

## LOCALIZATION

• The 4d superconformal algebra contains

$$\{\overline{Q}^{a}_{\dot{\alpha}}, \overline{S}^{b}_{\dot{\beta}}\} = \epsilon_{\dot{\alpha}\dot{\beta}}\epsilon^{ab} \left(\Delta - \frac{R}{2}\right) + \epsilon^{ab}M_{\dot{\alpha}\dot{\beta}} + \epsilon_{\dot{\alpha}\dot{\beta}}J^{ab}$$

Hence an interesting shortening condition is

$$[\overline{Q}^a_{\dot{\alpha}}, O] = 0 \rightsquigarrow \Delta_O = \frac{R_O}{2}, \ j_L = s = 0, \ (ar$$

- consequence, they form a ring: the chiral ring
- Their 2-point functions are

$$\langle O_I(0)\overline{O}_{\overline{J}}(x)\rangle = \frac{g_I\overline{J}(\tau^i,\overline{\tau}^i)}{|x|^{2\Delta_I}}\,\delta_{\Delta_I,\Delta_{\overline{J}}}$$

cf. Raffaele's talk!!

nd  $j_R = 0) \longrightarrow$  Chiral Primary Operator (CPO)

• CPO's have a non-singular OPE (not to violate the BPS bound). As a

Endowes the Coulomb branch of a very interesting geometry...but that's another story. See Papadodimas; Baggio, Niarchos & Papadodimas



- The 2-point functions can be mapped to the sphere  $\langle A(x)\overline{B}(0)\rangle = \frac{C_{AB}}{|x|^{2\Delta_A}}\delta_{\Delta_A\Delta_B}$
- To extract C, we can take the large x limit

$$\lim_{|x| \to \infty} |x|^{2\Delta_A} A(x) = 4^{\Delta_A} \lim_{|x| \to \infty} \left( 1 + \frac{|x|^2}{4} \right)^{\Delta_A} A(x)$$

Since  $\bullet$ 

 $ds_{\mathbb{R}^4}^2 = \left( \right.$ 

...it follows that  $4^{\Delta_A} \langle A(N) \overline{B}(S) \rangle_{\mathbb{S}^4} = C_{AB}$  ${\bullet}$ 

$$\rightsquigarrow \langle |x|^{2\Delta_A} A(x)\overline{B}(0) \rangle = C_{AB}\delta_{\Delta_A \Delta_B}$$

$$\left(1 + \frac{|x|^2}{4}\right)^4 ds_{\mathbb{S}^4}^2$$

→ Computed through a matrix integral thanks to localization

$$O_{\Delta}^{\mathbb{R}^4} \to O_{\Delta}^{\mathbb{S}^4} + \frac{\alpha_1}{R^2} O_{\Delta-2}^{\mathbb{S}^4} + \frac{\alpha_2}{R^4} O_{\Delta-4}^{\mathbb{S}^4} + \cdots$$

- Remove this mixing by Gram-Schmidt orthogonalization!

so as to define the double scaling limit (at FIXED N!)

$$n \to \infty$$
,  $g \to$ 

(Gauge instantons truly supressed!)

## • There is one subtlety, though: due to the conformal anomaly there can be mixing

Gerchkovitz, Gomis, Ishtiaque, Komargodski & Pufu, 1602.05971

• Let us look to the correlators of the simplest operators  $\mathcal{O}_n = (\text{Tr}\phi^2)^n$  in SU(N) SQCD

• The polynomial in n multiplying each order in the coupling is just the appropriate

 $\lambda \equiv g^2 n = \text{fixed} \,,$ ightarrow 0,

Bourget, R-G & Russo, 1803.00580



 $\lim(\mathfrak{g})$  .



- correlators shows that the limit continues to exist
- correlators: the GS can be recasted as a matrix model
  - derivatives
  - computation of correlators into a matrix model!

$$\det \mathcal{M}_{(n)} = \frac{1}{n!} \int_0^\infty \prod_{j=0}^{n-1} dx_j e^{-4\pi \operatorname{Im} \tau x_j} x_j^{\frac{1}{2}} Z_{1-\operatorname{Loop}} \left(\sqrt{x_j}\right) \prod_{j < k} (x_j - x_k)^2 .$$

speaking, the latter is the weak 't Hooft coupling regime)

### (note that in any case, gauge instantons are safely supressed in this regime)

• Going beyond this tower by explicit computation is hard. The next simplest case is SU(3): there is only one more CPO. Explicitly computing the

Beccaria, 1809.06280 Beccaria, 1810.10483

## • It turns out that the existence of the limit is rooted in the structure of the

• Very sketchy: for SU(2) there is only one CPO, whose sphere correlators are derivatives of Z wrt. the coupling. The flat space correlators are ratios of subdeterminants of the matrix of

• It turns out that each such subdeterminant can be written as a matrix integral: convert the

• The 't Hooft limit of this matrix model is well defined: it is our double scaling limit (strictly

Grassi, Komargodski & Tizziano, 1908.10306 Beccaria, Galvagno & Hasan, 2001.06645

## Wilson loops in the k-fold symmetrized product

representation. The exact formula is

$$\langle W_k \rangle = \frac{1}{N} \frac{1}{Z_N} \int d^N a \prod_{i < j} (a_i)$$

For N=4 both the instanton and 1-loop contributions are 1

Consider now circular Wilson loops in the k-fold symmetrized

$$(-a_j)^2 Z_{1-\text{loop}} Z_{\text{inst}} e^{-\frac{8\pi^2}{g^2} \sum_{i=1}^N a_i^2} W_k,$$

Then

$$W_k = (-1)^{N-1} \sum_{i=1}^N \frac{e^{2\pi(N-1)a_i + 2k\pi a_i}}{\prod_{j \neq i} e^{2\pi a_j} - e^{2\pi a_i}}$$

• Hence in the end

$$\langle W_k \rangle = \frac{(-1)^{N-1}}{Z_{U(N)}} \int d^N a \prod_{k < l} (a_k - a_l)^2 Z_{1-\text{loop}} Z_{\text{inst}} e^{-\frac{8\pi^2}{g^2} \sum_{m=1}^N a_m^2} \frac{e^{2\pi(N-1)a_N + 2k\pi a_N}}{\prod_{j \neq N} e^{2\pi a_j} - e^{2\pi a_N}}$$

• For SU(N) impose

## • The insertion is the character of the k-fold symm rep (of U(N)/SU(N)). This is easy to compute: the generating function is by definition the PE of the fundamental.

$$W_k = \sum_{1 \le i_1 \le i_2 \cdots \le i_k \le N} e^{2\pi a_{i_1} + 2\pi a_{i_2} + \cdots + 2\pi a_{i_k}}$$

$$\sum_{i=1}^{N} a_i = 0$$

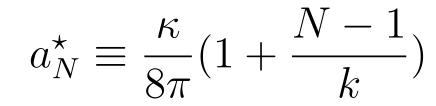
- Introduce now  $\kappa = g^2 k$
- Then (we specify to U(N) N=4)

$$\langle W_k \rangle = \frac{(-1)^{N-1}}{Z_{U(N)}} e^{\frac{k\kappa}{8}(1+\frac{N-1}{k})^2} \int d^N a \prod_{k$$

This suggests the limit FOR FIXED N

$$g \to 0, \quad k \to \infty, \qquad g^2 k = \kappa = \text{fixed}.$$

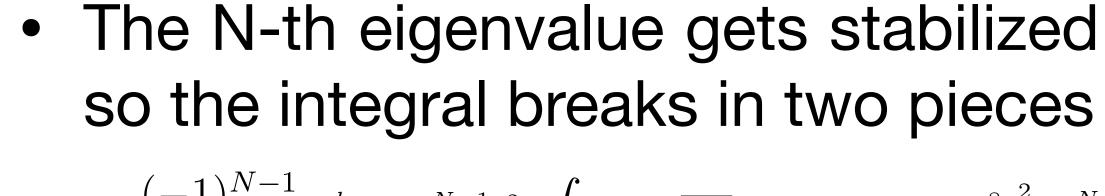
• the large charge limit)



c.f. as well Beccaria, Giombi & Tseytlin Cuomo, Komargodski, Mezei & Raviv-Moshe

Note that in this limit gauge instantons are completely supressed (just like in





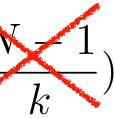
$$\langle W_k \rangle = \frac{(-1)^{N-1}}{Z_{U(N)}} e^{\frac{k\kappa}{8} (1+\frac{N-1}{k})^2} \int d^N a \prod_{k < l} (a_k - a_l)^2 e^{-k\frac{8\pi^2}{\kappa} \sum_{m=1}^{N-1} a_m^2} \left( \frac{e^{-k\frac{8\pi^2}{\kappa} \left(a_N - a_N^*\right)^2}}{\prod_{j \neq N} e^{2\pi a_N} - e^{2\pi a_N}} \right). \qquad a_N^*$$

Doing the last integral (saddle) and putting all factors

$$\langle W_k \rangle = \frac{1}{N!} \left(\frac{k \kappa}{4}\right)^{N-1} e^{\frac{k\kappa}{8} (1 + \frac{N-1}{k})^2} \left(e^{\frac{\kappa}{4}} - 1\right)^{1-N}$$

## • The N-th eigenvalue gets stabilized at a much larger scale than the rest...

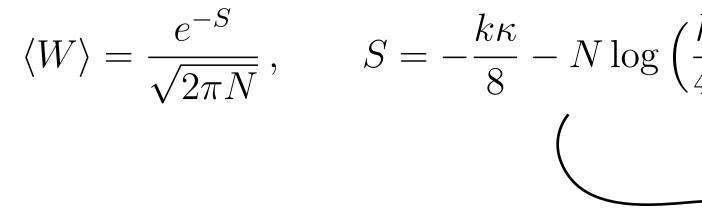
$$a_N^\star \equiv \frac{\kappa}{8\pi} (1 + \frac{1}{2})$$



• Note that

$$\langle W_k \rangle = \frac{(-1)^{N-1}}{Z_{U(N)}} e^{\frac{k\kappa}{8}(1+\frac{N-1}{k})^2} \int d^N a \prod_{k < l} (a_k - a_l)^2 e^{-k\frac{8\pi^2}{\kappa}\sum_{m=1}^{N-1} a_m^2} \left(\frac{e^{-k\frac{8\pi^2}{\kappa}\left(a_N - a_N^\star\right)^2}}{\prod_{j \neq N} e^{2\pi a_j} - e^{2\pi a_N}}\right). \qquad a_N^\star \equiv \frac{\kappa}{8\pi} (1 + \frac{N-1}{k})^2 e^{-k\frac{8\pi^2}{\kappa}\sum_{m=1}^{N-1} a_m^2} \left(\frac{e^{-k\frac{8\pi^2}{\kappa}\left(a_N - a_N^\star\right)^2}}{\prod_{j \neq N} e^{2\pi a_j} - e^{2\pi a_N}}\right).$$

Our result becomes then



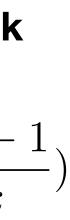
Lets compare with the holographic/matrix model@large N result

$$S_{\rm DF} = -2N \left[ \frac{k\sqrt{\lambda}}{4N} \sqrt{1 + \frac{k^2\lambda}{16N^2}} + \operatorname{arcsinh}\left(\frac{k\sqrt{\lambda}}{4N}\right) \right] \qquad \longrightarrow \qquad S_{DF} \sim -\frac{k\kappa}{8} - N \log\left(\frac{k\kappa}{4N}\right) - N$$

### We could take N large provided it is much smaller than k

$$g\left(\frac{k\,\kappa}{4N}\right) - N + N\,\log\left(1 - e^{-\frac{\kappa}{4}}\right)$$
$$\longrightarrow S = -\frac{k\kappa}{8} - N\log\left(\frac{k\,\kappa}{4N}\right) - N$$

Drukker & Fiol



• What about SU(N)? Do

$$\sum_{i=1}^{N} a_i^2 = \sum_{i=1}^{N} \hat{a}_i^2 + \frac{1}{N} (\sum_{i=1}^{N} a_i)^2, \qquad \hat{a}_i = a_i - \frac{1}{N} \sum_{i=1}^{N} a_i$$

• Then

$$\langle W_k \rangle = \frac{(-1)^{N-1}}{Z_N} \int d^N a \,\prod_{k < l} (\hat{a}_k - \hat{a}_l)^2 \, e^{-\frac{8\pi^2}{g^2} \sum_{m=1}^N \hat{a}_m^2} \, e^{-\frac{8\pi^2 N}{g^2} x^2 - 2\pi kx} \frac{e^{2\pi(N-1)\hat{a}_N + 2k\pi\hat{a}_N}}{\prod_{j \neq N} e^{2\pi\hat{a}_j} - e^{2\pi\hat{a}_N}} \,, \qquad x = \frac{1}{N} \sum_{i=1}^N a_i$$

• The a's sum zero: relax this by introducing a delta

$$\langle W_k \rangle = \frac{(-1)^{N-1}}{Z_N} \int d^N \hat{a} \prod_{k < l} (\hat{a}_k - \hat{a}_l)^2 e^{-\frac{8\pi^2}{g^2} \sum_{m=1}^N \hat{a}_m^2} \frac{e^{2\pi(N-1)\hat{a}_N + 2k\pi\hat{a}_N}}{\prod_{j \neq N} e^{2\pi\hat{a}_j} - e^{2\pi\hat{a}_N}} \,\delta(\sum_{i=1}^N \hat{a}_i) \left(\int dx \, e^{-\frac{8\pi^2 N}{g^2} x^2 - 2\pi kx}\right)$$

### • One recognizes the SU(N) result

$$\langle W_k \rangle_{U(N)} = \left( \frac{Z_{SU(N)}}{Z_{U(N)}} \int dx \, e^{-\frac{8\pi^2 N}{g^2} x^2 - 2\pi kx} \right) \langle W_k \rangle_{SU(N)} \longrightarrow \langle W_k \rangle_{U(N)} = e^{\frac{g^2 k^2}{8N}} \, \langle W_k \rangle_{SU(N)}$$

• The prefactor is a loop for the U(1) part

$$\frac{Z_{SU(N)}}{Z_{U(N)}} \int dx \, e^{-\frac{8\pi^2 N}{g^2} x^2 - 2\pi kx} = \frac{\int da \, e^{-\frac{8\pi^2}{g^2} a^2 - 2\pi \frac{k}{N}a}}{\int da \, e^{-\frac{8\pi^2}{g^2} a^2}}$$

If k>N this is a (leading) contribution: this observable is sensible to U vs SU!!!

- OPE
- The first few such correlators are

• But  $\frac{\langle \operatorname{Tr} \phi^n \rangle}{\langle 1 \rangle} \sim g^n \sim \kappa^{\frac{n}{2}} k^{-n}$ , so in this limit only the ``leading term'' contributes. To compute it

$$\langle \operatorname{Tr} \phi^{n_1} \cdots \operatorname{Tr} \phi^{n_m} W_k \rangle = \frac{1}{Z_{U(N)}} \int d^N a \prod_{k < l} (a_k - a_l)^2 Z_{1-\text{loop}} Z_{\text{inst}} e^{-\frac{8\pi^2}{g^2} \sum_{m=1}^N a_m^2} \\ \times \frac{e^{2\pi (N-1) a_N + 2k\pi a_N}}{\prod_{j \neq N} \left( e^{2\pi a_N} - e^{2\pi a_j} \right)} \left( \sum_{i=1}^N a_i^{n_1} \right) \cdots \left( \sum_{i=1}^N a_i^{n_m} \right).$$

## One can also compute correlators of loops with CPO's. This has info about the

•••

$$\mathcal{O}_1 = \mathrm{Tr}\phi - \frac{\langle \mathrm{Tr}\phi \rangle}{\langle 1\!\!\!1 \rangle} 1\!\!\!1, \qquad \mathcal{O}_2 = \mathrm{Tr}\phi^2 - \frac{\langle \mathrm{Tr}\phi^2 \rangle}{\langle 1\!\!\!1 \rangle} 1\!\!\!1$$





$$\left\langle \operatorname{Tr}\phi^{n_1}\cdots\operatorname{Tr}\phi^{n_m}W_k\right\rangle = \frac{Z_{U(N-1)}}{Z_{U(N)}} e^{\frac{k\kappa}{8}(1+\frac{N-1}{k})} \int da_N \left(\frac{a_N^2}{e^{2\pi a_N}-1}\right)^{N-1} a_N^{n_1+\dots+n_m} e^{-k\frac{8\pi^2}{\kappa}(a_N-\frac{\kappa}{8\pi})^2}.$$

• Hence

 $\langle \mathrm{Tr}\phi^{n_1}\cdots\mathrm{Tr}\phi^{n_m}W\rangle$ 

• So finally

$$\langle \mathcal{O}_{\Delta} W_k \rangle = \left(\frac{\kappa}{8\pi}\right)^{\Delta} \langle W_k \rangle$$
$$\bigwedge \qquad \langle \hat{\mathcal{O}}_{\Delta} W_k \rangle = \frac{1}{\sqrt{\Delta}} \left(\frac{k\kappa}{16N}\right)$$

### • Just as before in the large k limit the integral can be done via saddle point giving

$$\langle V_k \rangle = \left(\frac{\kappa}{8\pi}\right)^{n_1 + \dots + n_m} \langle W_k \rangle \,.$$

$$\langle \mathcal{O}_{\Delta}(x) \mathcal{O}_{\Delta}(0) \rangle = \frac{C_{\Delta}}{|x|^{2\Delta}}, \qquad C_{\Delta} \equiv \frac{\Delta \lambda^{\Delta}}{(2\pi)^{2\Delta}}$$

 $\langle \rangle^{\frac{\Delta}{2}} \langle W_k \rangle ,$ 

Berenstein, Corrado, Fischler& Maldacena '98 Giombi, Ricci & Trancanelli '06



### A similar double scaling limit holds in general N=2 theories. For instance, for N=2\*

c.f for SU(2) SQCD Cuomo, Komargodski, Mezei & Raviv-Moshe' 22

$$W_k \rangle_{\mathcal{N}=2^*} = \frac{1}{Z_{\mathcal{N}=2^*}} \int d^N a \prod_{i$$

Introducing the same variables

$$\langle W_k \rangle_{\mathcal{N}=2^*} = \frac{e^{\frac{k\kappa}{8}\left(1+\frac{N-1}{k}\right)^2}}{Z_{\mathcal{N}=2^*}} \int d^{N-1}a \prod_{i$$



In the same double scaling limit at large k one then finds

$$\langle W_k \rangle_{\mathcal{N}=2^*} = \left( \frac{H(a_N^*)^2 H(M)^2}{H(a_N^* + M) H(a_N^* - M)} \right)^{N-1} \langle W_k \rangle_{\mathcal{N}=4}$$

finds

$$\log \langle W_k \rangle_{\mathcal{N}=2^*} \approx \log \langle W_k \rangle_{\mathcal{N}=4} + 2(N-1)\log H(a_N^*) - \frac{1}{2}(N-1)R^2 \Big[ 2M^2 \log(MR)^2 - (M-a_N^*)^2 \log(M-a_N^*)^2 R^2 - (M+a_N^*)^2 \log(M+a_N^*)^2 R^2 \Big] + (1-2\gamma)(N-1)(Ra_N^*)^2.$$

## • It is interesting to look to the decompatification limit of large MR. $N=2^*$ undergoes a sequence of phase transitions...what about the large k loop? One

Russo & Zarembo



corresponds to

- ...so this happens "infinitely far away": no phase transitions for this observable
- All in all, in the decompactification limit

 $\log \langle W_k \rangle_{\mathcal{N}=2^*} \longrightarrow \log \langle W_k \rangle_{\mathcal{N}=4} + (N)$ 

### • This suggests a potential non-analytic behavior at a\* of order M. But this

 $\kappa = 8\pi MR$ 

$$-1)\left(2\log H(\frac{\kappa}{8\pi}) + \frac{\kappa^2}{32\pi^2}\left[2 - \gamma + \log\left(MR\right)\right]\right)$$

## **Defects in WF**

 $\bullet$ 6d.

D.R-G.

• Also a similar double-scaling limit exists

- For instance, consider O(2N+1) near d=4
  - $S = \int \frac{1}{2}$
- $\bullet$

 $\mathcal{D}(\vec{z}) = e^{-h \int d\tau \,\varphi^{2N+1}(\tau,\vec{z})} = e^{-h \int dx \,\varphi^{2N+1} \,\delta_T(\vec{x}-\vec{z})}$ 

## Using the same strategy we can also study lines in the WF fixed point near d=4,

Arias-Tamargo, Russo & R-G, Watanabe, Badel, Cuomo, Monin & Rattazzi. Hellerman, Orlando, Reffert et al.

• Can be interpreted as effective description of large spin impurities in magnets

Cuomo, Komargodski, Mezei & Raviv-Moshe

$$|\partial \vec{\varphi}|^2 + \frac{g}{4} (\vec{\varphi}^2)^2 \,,$$

## One may imagine the trivial line along one direction. It admits a deformation

$$\vec{z} \longrightarrow \langle \mathcal{D}(\vec{z}) \rangle = \int e^{-\int \frac{1}{2} |\partial \vec{\varphi}|^2 + \frac{g}{4} (\vec{\varphi}^2)^2 + h \varphi^{2N+1} \delta_T (\vec{x} - \vec{z}) \rangle$$

- Assume  $g = \frac{\lambda}{n}$ ,  $h = \nu n$
- Then

$$\langle \mathcal{D}(\vec{z}) \rangle = \int e^{-nS_{\text{eff}}}, \qquad S_{\text{eff}} = \int \frac{1}{2} |\partial \vec{\phi}|^2 + \frac{\lambda}{4} (\vec{\phi}^2)^2 + \nu \phi^{2N+1} \delta_T (\vec{x} - \vec{z}).$$

only relevant eq. can be easily solved

$$\partial^2 \phi^{2N+1} - \lambda (\phi^{2N+1})^3 - \nu \,\delta_T (\vec{x} - \vec{z}) = 0 \,.$$

• Finally

$$S_{\text{eff}} = \left( -\frac{\nu^2}{2} + \frac{\lambda\nu^4}{128\pi^2\epsilon} + \frac{\lambda\nu^4}{128\pi^2} (3 - \gamma_E + \log(4\pi)) \right) T \int \frac{d^{d-1}\vec{p}}{(2\pi)^{d-1}} \frac{1}{\vec{p}^2} - \frac{\lambda\nu^4}{128\pi^2} T \int \frac{d^{d-1}\vec{p}}{(2\pi)^{d-1}} \frac{\log|p|^2}{\vec{p}^2}.$$

• So imagine taking large n with everything else fixed: use saddle point. The

• Since 
$$\frac{1}{\langle \mathcal{D} \rangle} \frac{d}{d\nu} \langle \mathcal{D} \rangle = -n \int \langle \phi^{2N+1} \rangle \, \delta_T(\vec{x} - \vec{z})$$
.

• We can use this to define the renormalized coupling

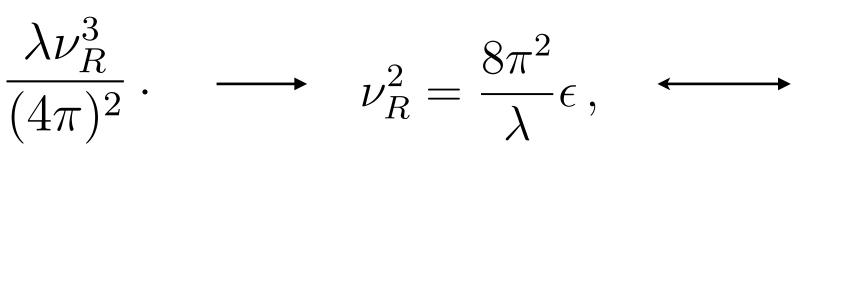
$$\nu = \mu^{\frac{\epsilon}{2}} \left( \nu_R + \frac{\lambda \nu_R^3}{2(4\pi)^2 \epsilon} \right),$$

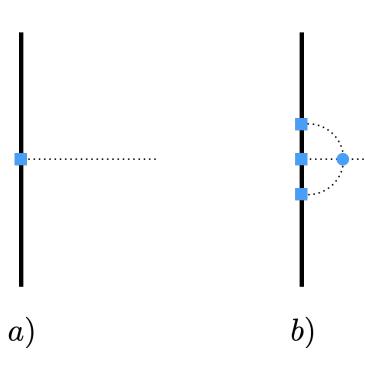
We can compute the beta function, which shows a fixed point

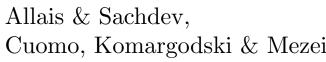
$$\mu \frac{d\nu_R}{d\mu} = -\frac{\epsilon}{2}\nu_R + \frac{1}{2}(1-\frac{1}{2})\nu_R + \frac{1}{2}(1-\frac{1}{2})\nu_R$$

...where

$$\left\langle \phi^{2N+1} \right\rangle = -\nu_R$$







 $\frac{1}{\sqrt{(2\pi)^{d-1}}} \frac{e^{-ip \cdot x}}{|\vec{p}|^{2-\frac{\epsilon}{2}}} \sim \frac{1}{|\vec{x}_T|^{\frac{d-2}{2}}}$ 

themselves. For instance, for the latter

$$\left\langle \mathcal{D}(z_1)\mathcal{D}(z_2)\right\rangle = \int e^{-\int \frac{1}{2}|\partial\vec{\varphi}|^2 + \frac{g}{4}(\vec{\varphi}^2)^2 + h\varphi^{2N+1}\,\delta_T(\vec{x}-\vec{z_1}) + h\varphi^{2N+1}\,\delta_T(\vec{x}-\vec{z_2})}$$

Assuming the same scaling

$$\langle \mathcal{D}(z_1)\mathcal{D}(z_2) \rangle = \int e^{-nS_{\text{eff}}}, \qquad S_{\text{eff}} = \int \frac{1}{2} |\partial \vec{\phi}|^2 + \frac{\lambda}{4} (\vec{\phi}^2)^2 + \nu \phi^{2N+1} \, \delta_T(\vec{x} - \vec{z_1}) + \nu \phi^{2N+1} \, \delta_T(\vec{x} - \vec{z_2}) \, .$$

$$\text{addle point eqs. are}$$

$$(\vec{x} - \vec{z_1}) - \nu \, \delta_T(\vec{x} - \vec{z_2}) = 0. \qquad \longrightarrow \qquad \phi^{2N+1} = \rho_1(\vec{x}) + \rho_2(\vec{x}), \qquad \rho_i(\vec{x}) = -\nu \int dy \, G(x - y) \delta_T(\vec{y} - \vec{z_i}) \, .$$

The sa

 $\partial^2 \phi^{2N+1} - \nu \, \delta_T$ 

So finally

 $\langle \mathcal{D}(z_1)\mathcal{D}(z_2)\rangle = \langle \mathcal{D}(z_1)\rangle \langle \mathcal{D}(z_2)\rangle e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{3\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}(3+\gamma_E+1)\right) \langle \mathcal{D}(z_2)\rangle = \langle \mathcal{D}(z_1)\rangle \langle \mathcal{D}(z_2)\rangle e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{3\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}(3+\gamma_E+1)\right) \langle \mathcal{D}(z_2)\rangle = \langle \mathcal{D}(z_1)\rangle \langle \mathcal{D}(z_2)\rangle e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{3\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}(3+\gamma_E+1)\right) \langle \mathcal{D}(z_2)\rangle = -\left(\frac{\nu_R^2}{4\pi} + \frac{3\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) \langle \mathcal{D}(z_2)\rangle e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{3\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) \langle \mathcal{D}(z_2)\rangle e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{3\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) \langle \mathcal{D}(z_2)\rangle e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{3\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) \langle \mathcal{D}(z_2)\rangle e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{3\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) \langle \mathcal{D}(z_2)\rangle e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{3\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) \langle \mathcal{D}(z_2)\rangle e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) \langle \mathcal{D}(z_2)\rangle e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) \langle \mathcal{D}(z_2)\rangle e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^4}{4\pi} + \frac{\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^4}{4\pi} + \frac{\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}\right) e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^4}{4\pi} + \frac{\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{512\pi}\right) e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^4}{4\pi} + \frac{\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{512\pi}\right) e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^4}{4\pi} + \frac{\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{512\pi}\right) e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^4}{4\pi} + \frac{\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{512\pi}\right) e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^4}{4\pi} + \frac{\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{512\pi}\right) e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^4}{4\pi} + \frac{\lambda\nu_R^4}{512\pi}\right) e^{-nS_{\mathrm{I}}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^4}{4\pi} + \frac{\lambda\nu_R^4}{512\pi}\right) e^$ 

## One can also compute correlators of defect fields as well as correlators of defects

$$\log(4\pi))\Big)\frac{T}{|\vec{z}_1 - \vec{z}_2|^{1 + (-\epsilon + \frac{\lambda\nu_R^2}{8\pi^2})}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{3\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}(3 + \gamma_E + \log(4\pi))\right)\frac{T}{|\vec{z}_1 - \vec{z}_2|^{1 + (-\epsilon + \frac{\lambda\nu_R^2}{8\pi^2})}} \longrightarrow S_{\mathrm{I}} = -\left(\frac{\nu_R^2}{4\pi} + \frac{3\lambda\nu_R^4}{512\pi} - \frac{\lambda\nu_R^4}{64\pi^3}(3 + \gamma_E + \log(4\pi))\right)\frac{T}{|\vec{z}_1 - \vec{z}_2|^{1 + (-\epsilon + \frac{\lambda\nu_R^2}{8\pi^2})}}$$

c.f. Soderberg (free case)

a)





b)



quartic theory above 4d

Fei, Giombi & Klebanov '14 Giombi, Huang, Klebanov, Pufu & Tarnopolski '19

$$S = \int \frac{1}{2} |\partial \vec{\varphi}|^2 + \frac{1}{2} \partial \eta^2 + \frac{g_1}{2} \eta \, |\vec{\varphi}|^2 + \frac{g_2}{6} \eta^3 \, . \qquad g_{1\star} = \sqrt{\frac{6(4\pi)^3}{2N}\epsilon} \left(1 + \mathcal{O}(\frac{1}{N})\right), \qquad g_{2\star} = 6\sqrt{\frac{6(4\pi)^3}{2N}\epsilon} \left(1 + \mathcal{O}(\frac{1}{N})\right)$$

zation this theory is described by

$$S_{\text{quartic}} = \int \partial \vec{\xi}^2 + \sigma \, \vec{\xi}^2 \,,$$

• For the HS field

 $\langle \sigma(x) \, \sigma(0) \rangle$ 

### Near d=6 a similar story holds. Consider the proposed UV completion to the

This is the UV completion of the quartic theory. Upon Hubbar-Stronovich -

$$\rangle \sim \frac{1}{x^4}$$



In the sextic theory one can consider surface defects deformed as

 $\mathcal{D} =$ 

- $\mathcal{D}_{ ext{quart}}$
- Assuming double scaling, we can use saddle point

$$S_{\text{eff}} = n \int \frac{1}{2} |\partial \vec{\phi}|^2 + \frac{1}{2} \partial \rho^2 + \frac{h_1}{2} \rho \, |\vec{\phi}|^2 + \frac{h_2}{6} \rho^3 + \nu \rho \, \delta_T(\vec{x}) \qquad g_i \sqrt{n} = h_i, \ h = \nu \sqrt{n}.$$

$$e^{-h\int d^2x\,\eta}$$

## • This does not break the O(N) symmetry. A natural guess in the quartic theory

$$_{\rm tic} = e^{-\hat{h} \int d^2 x \vec{\xi}^2}$$



we'll leave that for another day

### Proceeding as in the 4d case, one can compute the defect beta function

$$\beta_{\nu} = -\frac{\epsilon}{2}\nu_R - \frac{h_2\nu_R^2}{16\pi^2} \,.$$

$$\nu_R = -\frac{8\pi^2\epsilon}{h_2} \,.$$

One can also compute correlators of defects as well as defect fields...but

## Final comments

- Inspired by the large charge developments, we introduced a double scaling limit for defects
- In the WF fixed near 4/6d we considered lines/surfaces: a fixed point leading to a dCFT exists
- One can compute correlators of defect operators/defects themselves
- We introduced a novel double scaling limit for the k-fold symmetrized Wilson loop
  - Allows to compute exact observables for finite N in gauge theories (free of gauge instantons!)

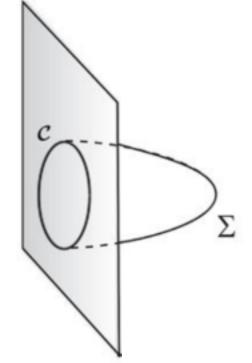
## The loop distinguishes U vs SU...can this be seen holographically?

• U vs SU encoded in a topological BF theory in AdS5...



• ...the fluxed D3 dual to the loop would source the RR 2-form...gives rise to boundary term?

...maybe one needs to do holography "the other way around"
 cf. Shota's talk!!



$$N \int C_2 \wedge dB_2$$

- flows?
  - Perhaps toy models for interesting behaviors (fixed point) annihilation?)
  - Can one study general aspects of RG flows such as entropy extremization?

• The defect on the WF in this limit simplifies...can one study aspects of RG



## Many thanks!!!