



GELINA at JRC Geel

Arjan Plompen

NSTAPP, CERN, 20y n_TOF, 22 November 2021

EURATOM research infrastructure JRC Geel

GELINA



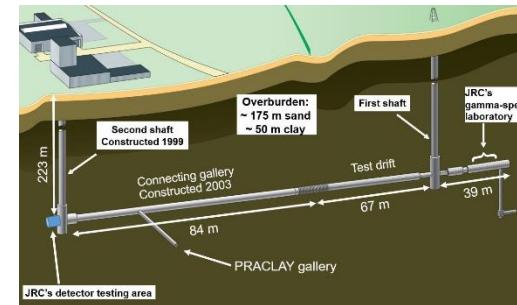
METRO



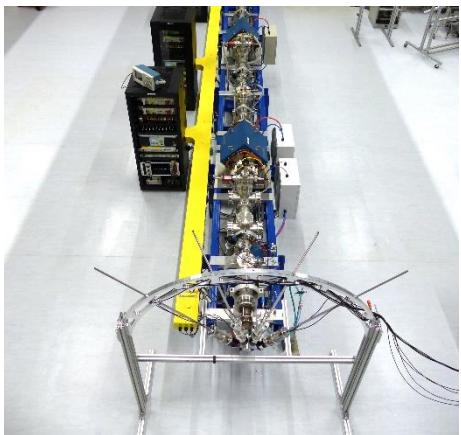
RADMET



HADES



MONNET



TARGET



- Nuclear data
- Metrological tools for nuclear safety, security and safeguards
- Radionuclide metrology

Nuclear data – major use



Source: SKB

Source: Vattenfall.com

800 TWh/year in EU-27 + UK



Nuclear fuel investment of order 10 000 M€/year

Source: World-nuclear.org

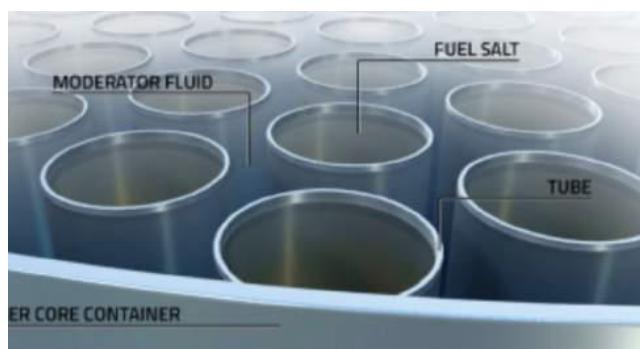
1.500 tons spent nuclear fuel in Switzerland – use of nuclear data can reduce cost (more than 100 M€ for final repository)

50.000 tons spent nuclear fuel in Europe

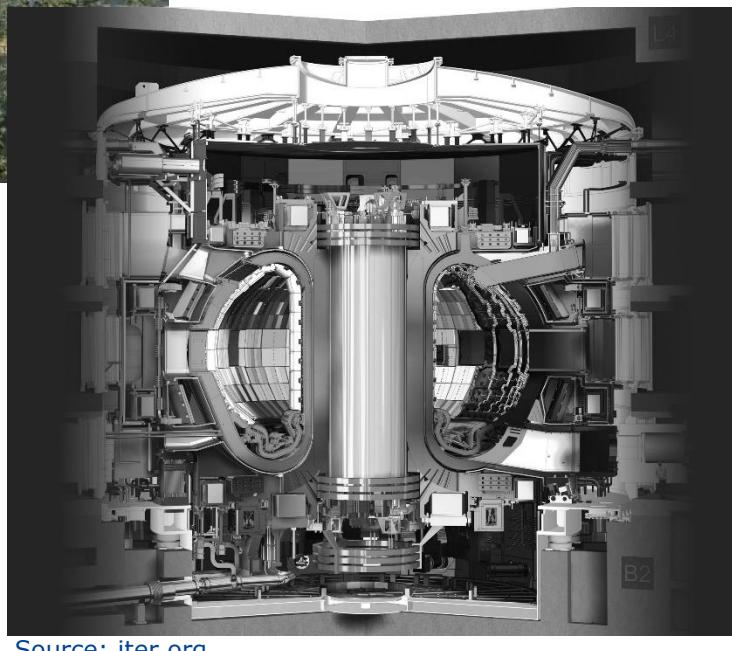
Nuclear data – innovation



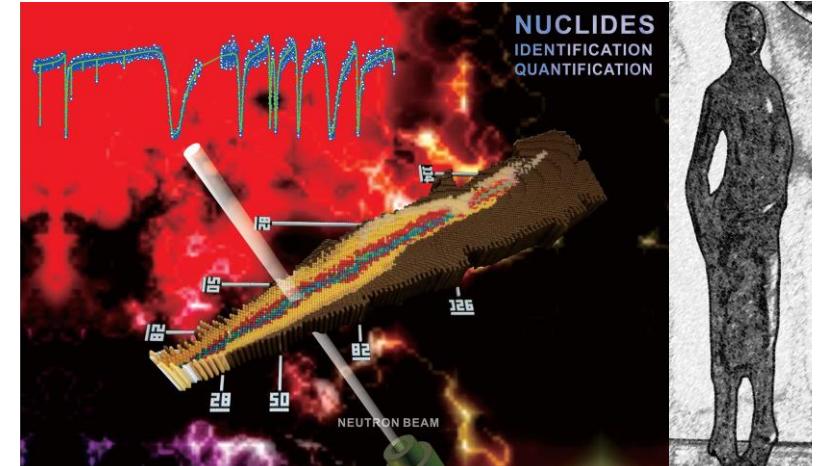
Source: myrrha.be



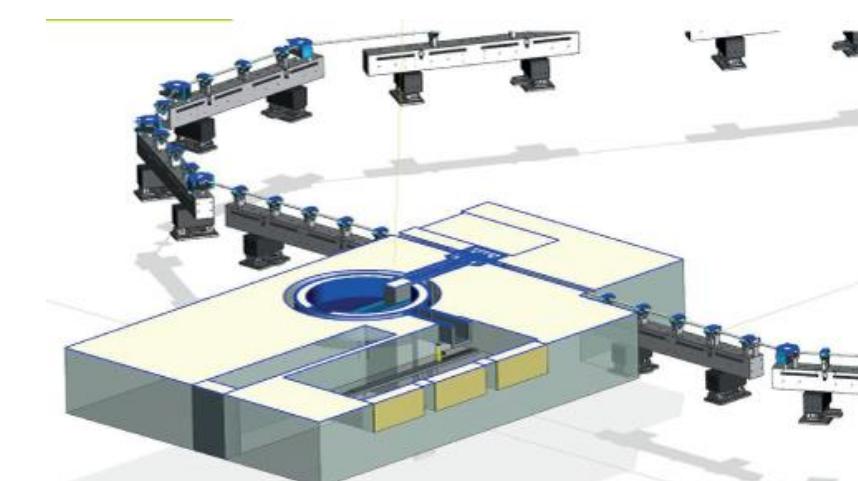
Source: IAEA SMR book 2018



Source: iter.org

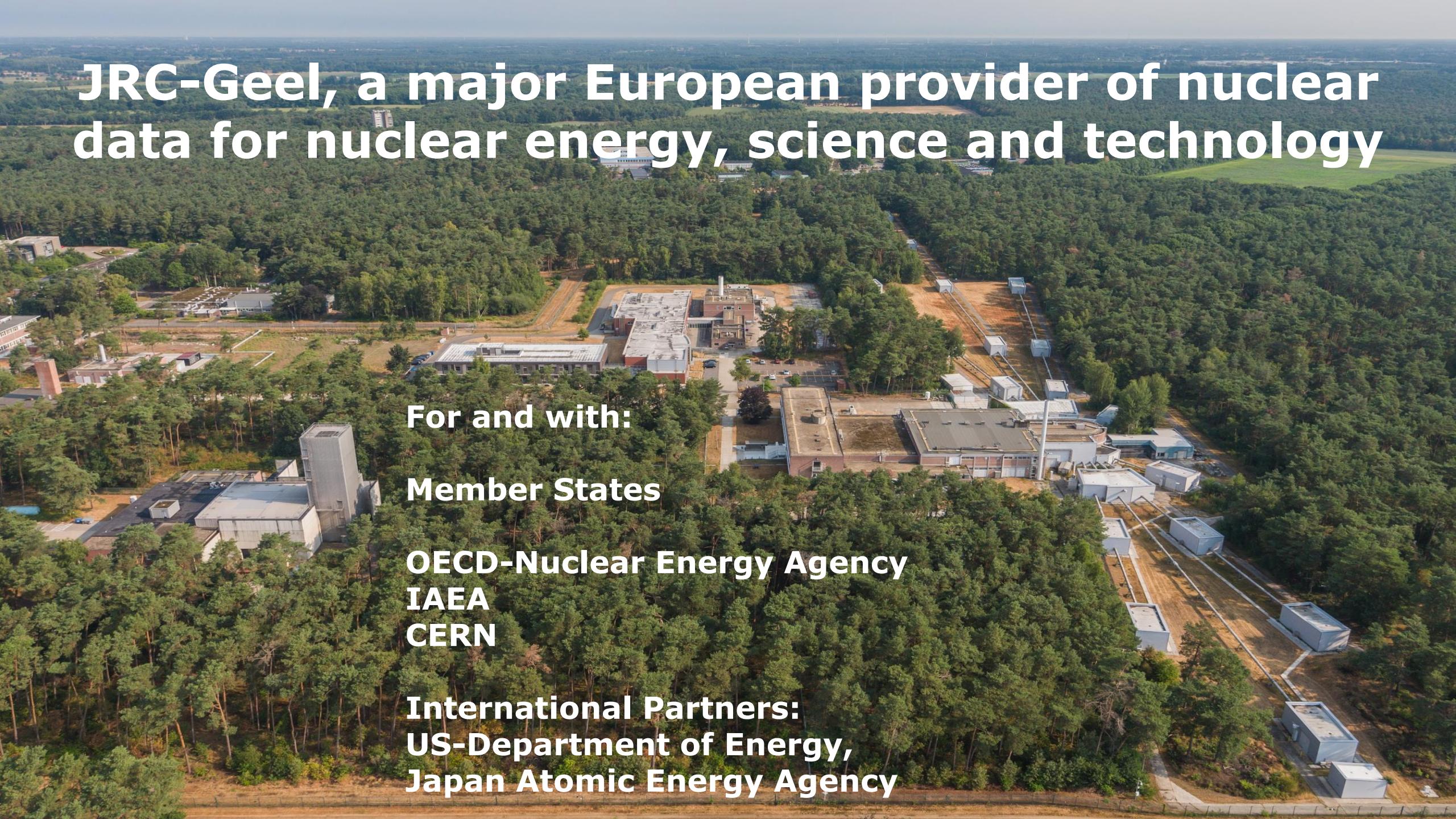


Source: iJAAS / JRC



Source: ire.eu

JRC-Geel, a major European provider of nuclear data for nuclear energy, science and technology



For and with:

Member States

OECD-Nuclear Energy Agency

IAEA

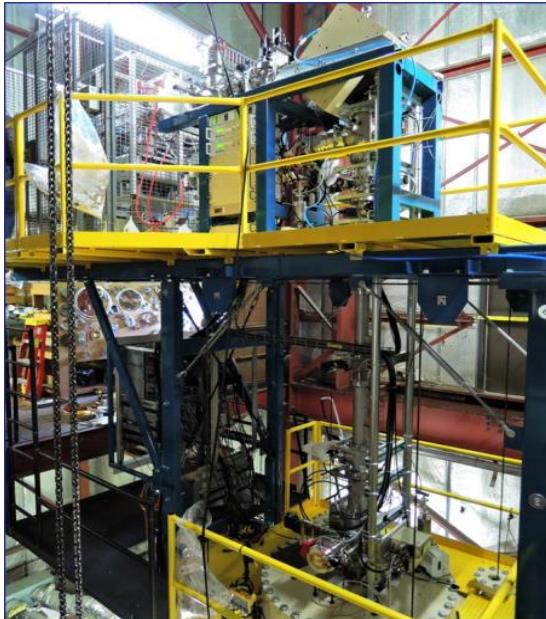
CERN

International Partners:

US-Department of Energy,

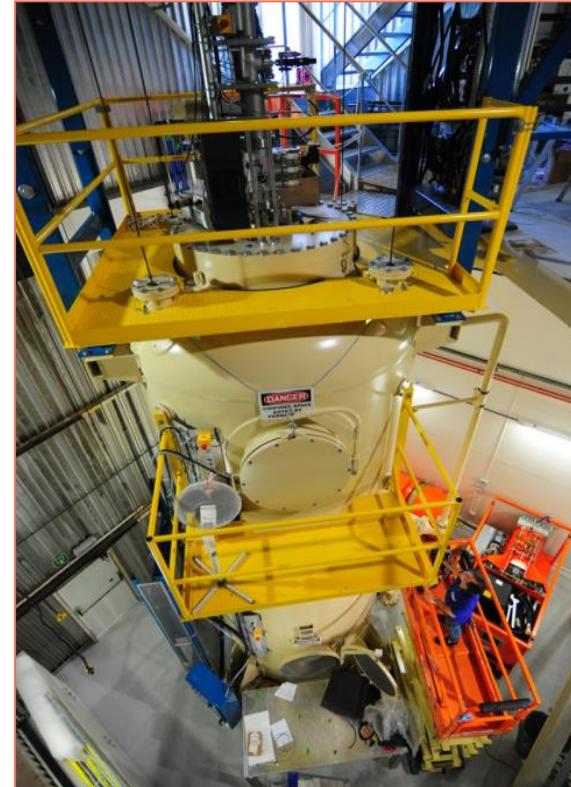
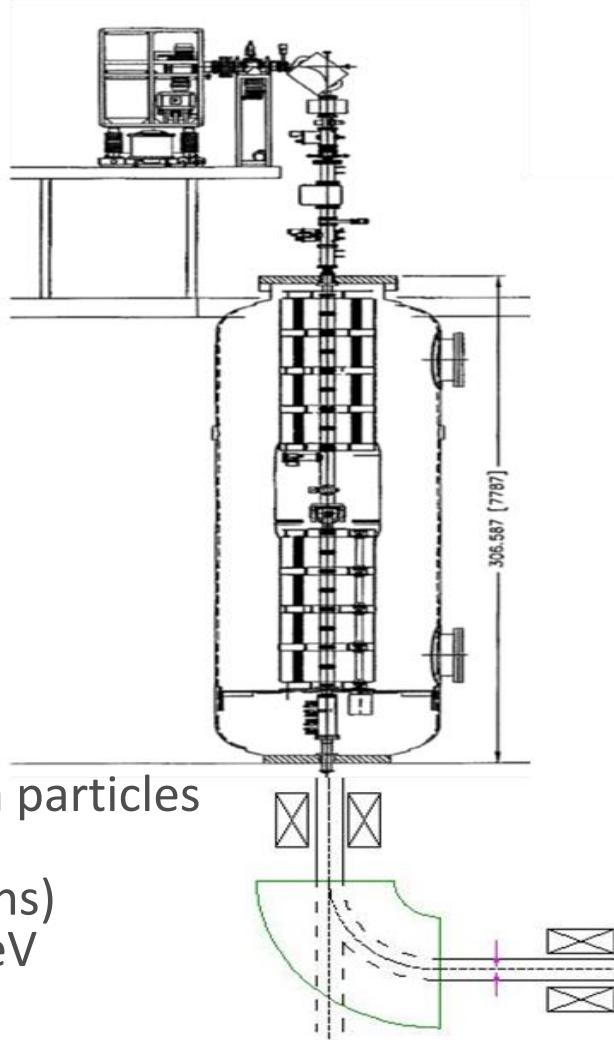
Japan Atomic Energy Agency

MONNET: MOno-energetic NEutrons by Tandem

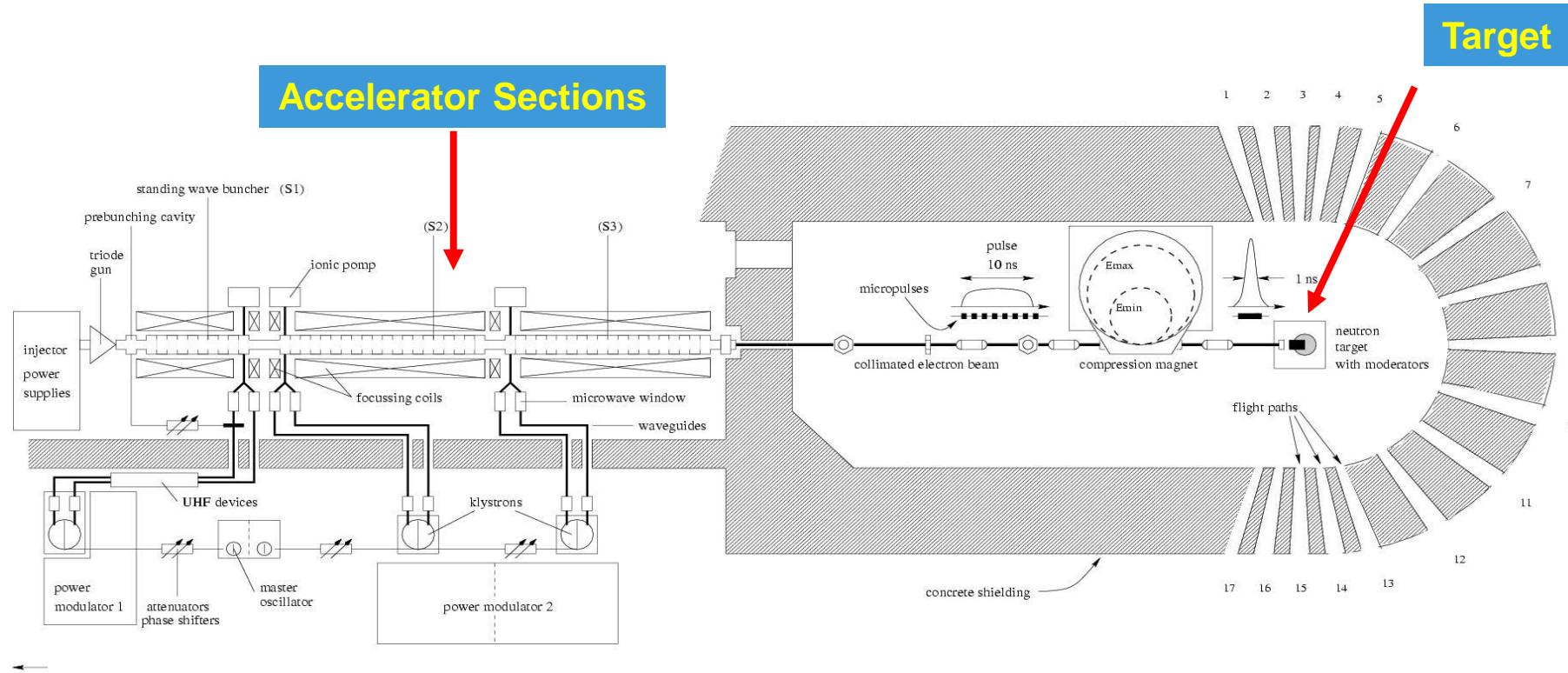


3.5 MV NEC Tandem

- Protons, deuterons and alpha particles
- DC ($p,d < 50 \mu\text{A}$)
- Pulsed beam available (1 – 2 ns)
- Energy range: 200 keV – 7 MeV



GELINA - Electron Linear Accelerator



Normal Operating Parameters

Average Current : $70 \mu\text{A}$
Maximum Electron Energy : 130 MeV
Mean Power : 7 kW

Frequency : up to 800 Hz
Pulse Width : $1-2 \text{ ns}$
Neutron Flux : $2 \times 10^{13} \text{ 1/s}$

GELINA - Experimental set-ups



- Transmission
 - 10 m, 30m, 50 m
- Capture
 - 10 m, 30 m, 60 m
- Elastic sattering
 - 30 m
- In-elastic scattering
 - 30 m, 100 m
- Fission, (n,p) , (n,α) ,
 - 10 m

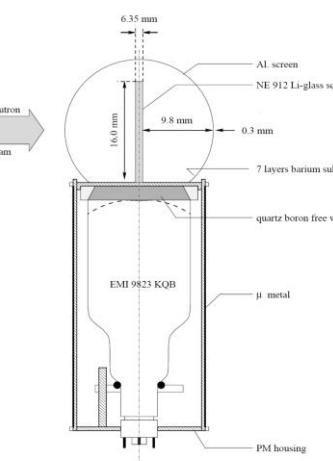
GELINA - Experimental set-ups



Detectors

Low energy :

High energy :

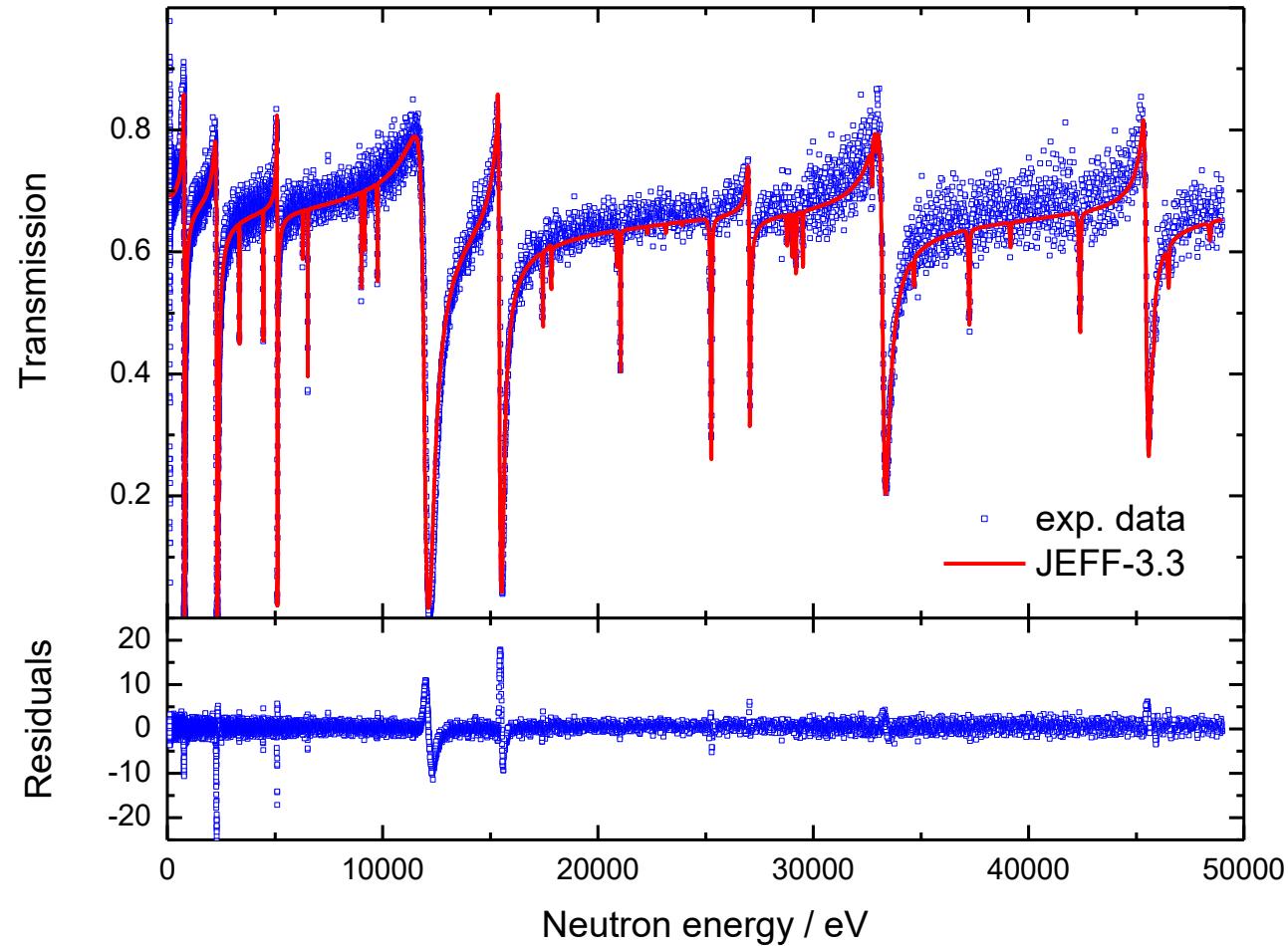


Transmission

- 10 m, 30 m, 50 m



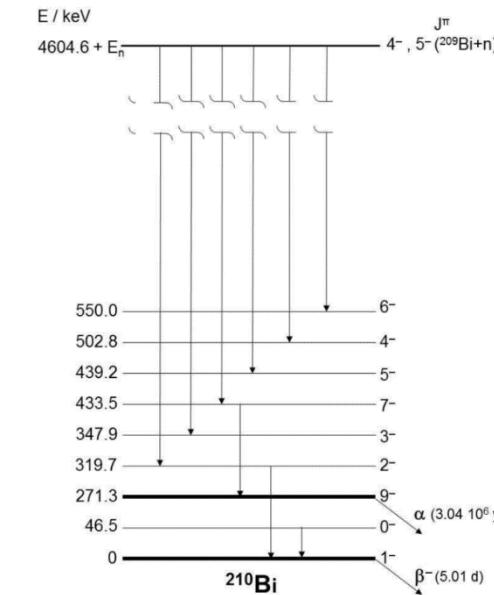
New evaluation for $^{209}\text{Bi} + n$



Project in collaboration with
SCK•CEN and JAEA

- Improve RP for ^{209}Bi
- Capture at GELINA (60m)
- Determine branching ratio

$$\text{J-PARC: } \sigma_{\gamma,g} / \sigma_{\gamma,m}$$

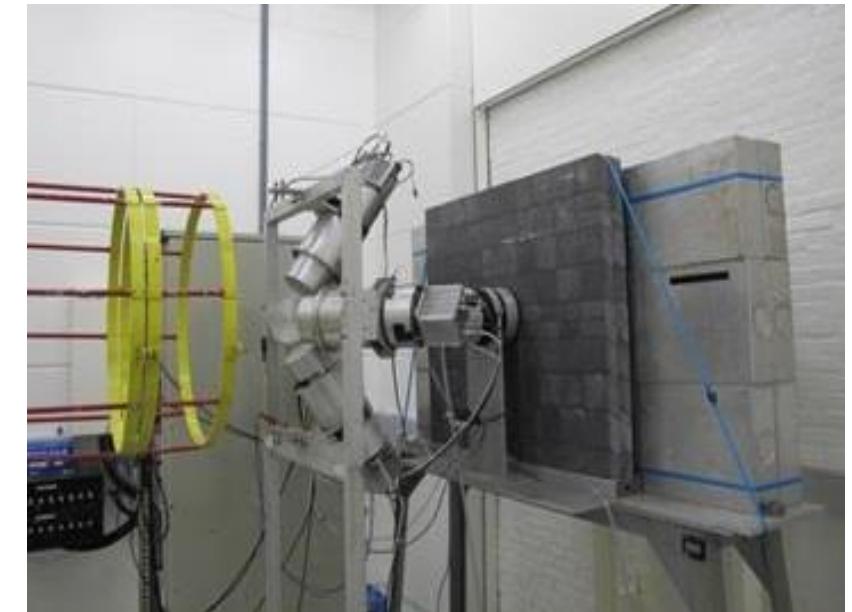


GELINA - Experimental set-ups

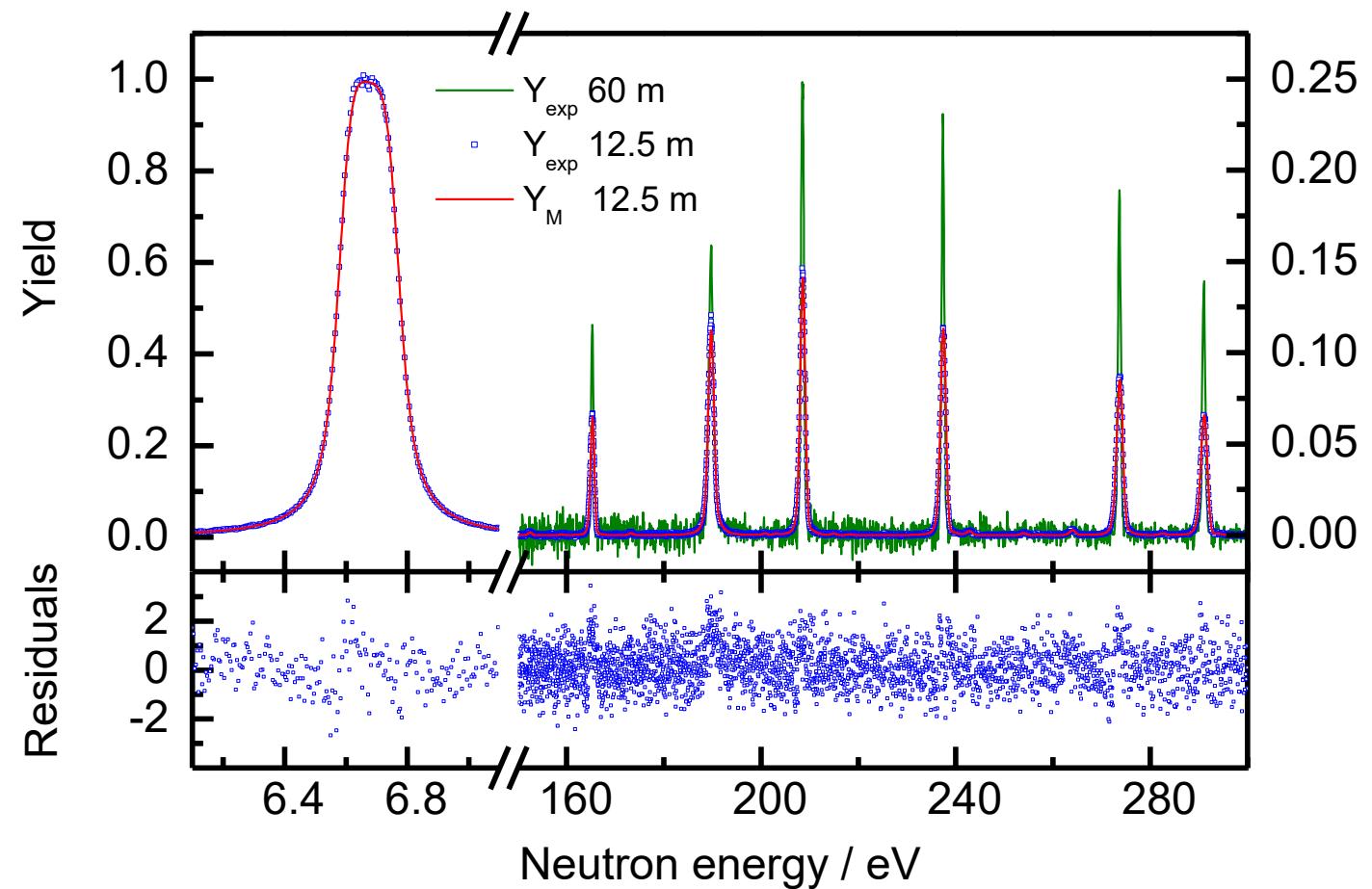


Capture

- 10 m, 30 m, 60 m



$^{238}\text{U}(\text{n},\gamma)$ resonance shape analysis: REFIT



Kim et al., EPJA 52 (2016) 170

GLSQ analysis for $n + {}^{238}\text{U}$: $\sigma(n, \gamma)$

Experimental data from :

- Standard 2007
- Kim et al. (GELINA) ($\sim 2\%$)
- Ullmann et al. (LANSCE)
- Mingrone et al. (n_TOF)
- Wright et al. (n_TOF)

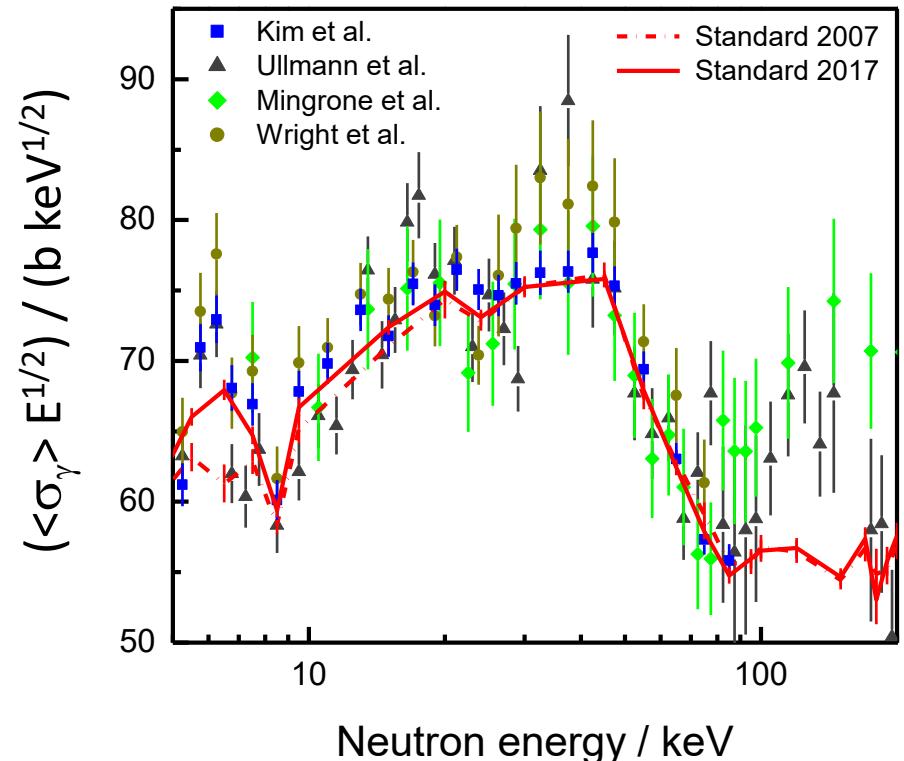
⇒ GMA analysis (GLSQ)

⇒ New evaluation for ${}^{238}\text{U} + n$ in RRR and URR

Dr. K. Kim, Dr. Y.-O. Lee

KAERI, Nuclear Data Center, Daejeon

Kim et al., EPJA 52 (2016) 170
Sirakov et al., EPJA 53 (2017) 199

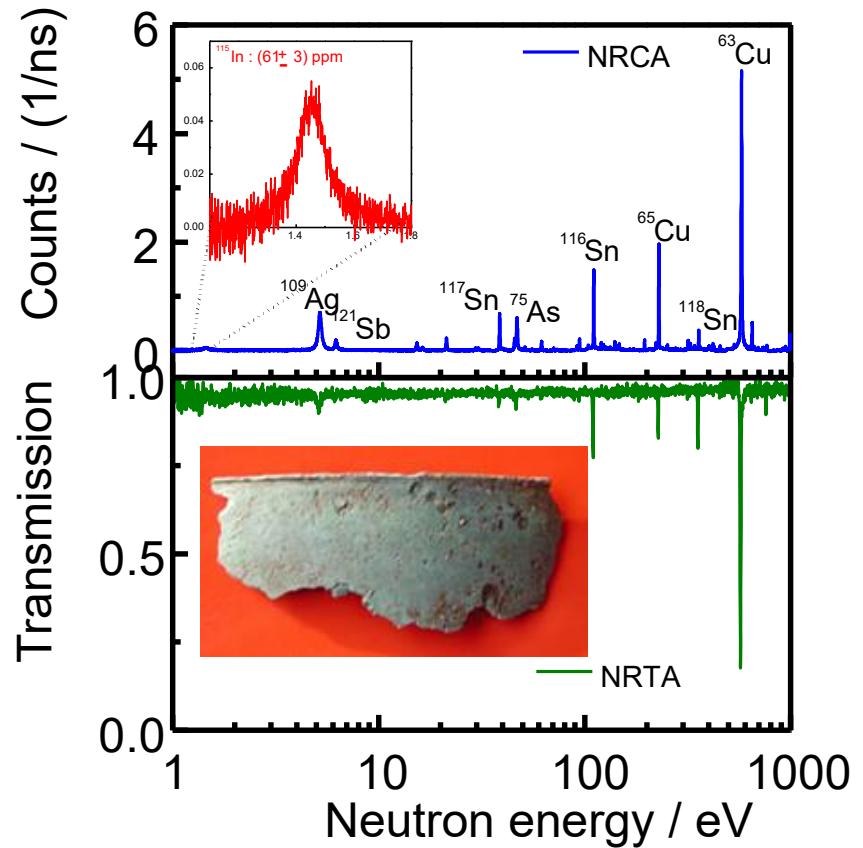


HIGHLIGHTS FROM EUROPEAN JOURNALS

NUCLEAR PHYSICS

From experiment to evaluation,
the case of $n + {}^{238}\text{U}$

NRTA and NRCA



Postma et al., Czech. J. Phys. 53 (2003) A233

Areal density Cu: 0.55 (2) g/cm²

Element	wt%	Isotope	E _r / eV
Cu	77.76 \pm 0.11	^{63}Cu	579.0
		^{65}Cu	230.0
Sn	20.85 \pm 0.10	^{112}Sn	94.8
		^{116}Sn	111.2
		^{117}Sn	38.8
		^{118}Sn	45.7
		^{119}Sn	222.6
		^{120}Sn	427.5
		^{124}Sn	62.0
As	0.34 \pm 0.01	^{75}As	47.0
Sb	0.20 \pm 0.02	^{121}Sb	6.24
		^{123}Sb	21.4
Ag	0.09 \pm 0.01	^{107}Ag	16.3
		^{109}Ag	5.2
Fe	0.77 \pm 0.10	^{56}Fe	1147.4
In	0.0061 \pm 0.0003	^{115}In	1.46

GELINA - Experimental set-ups



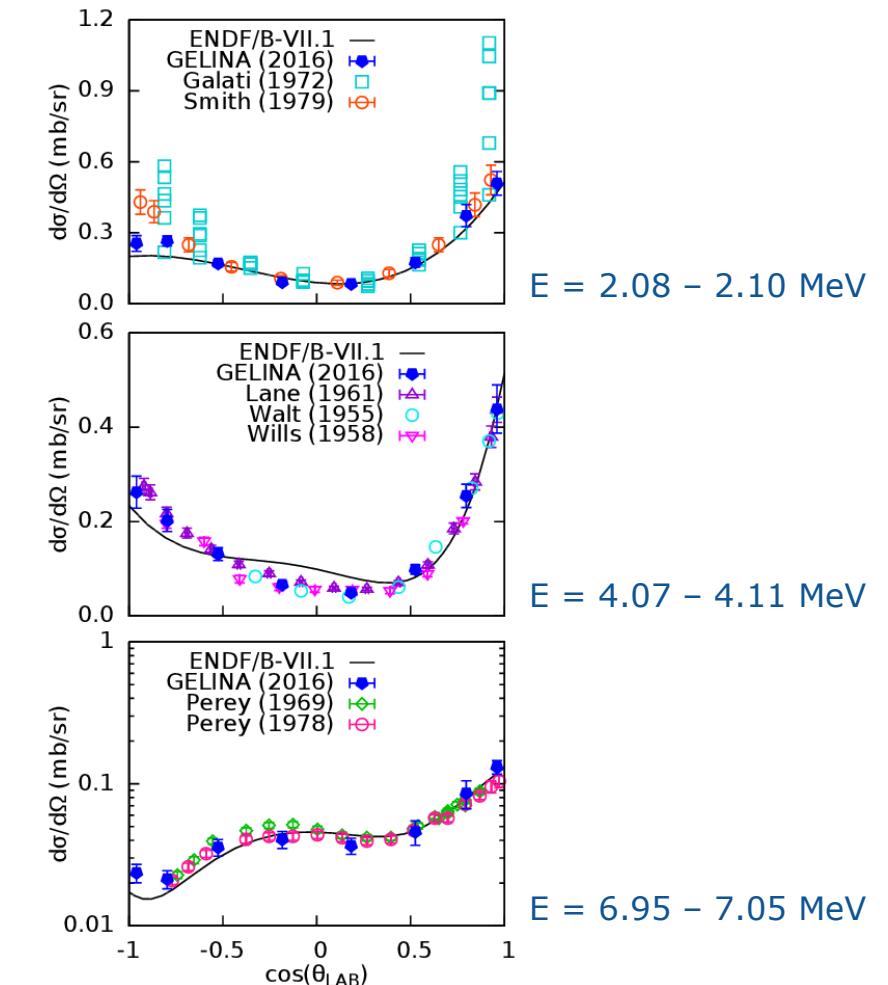
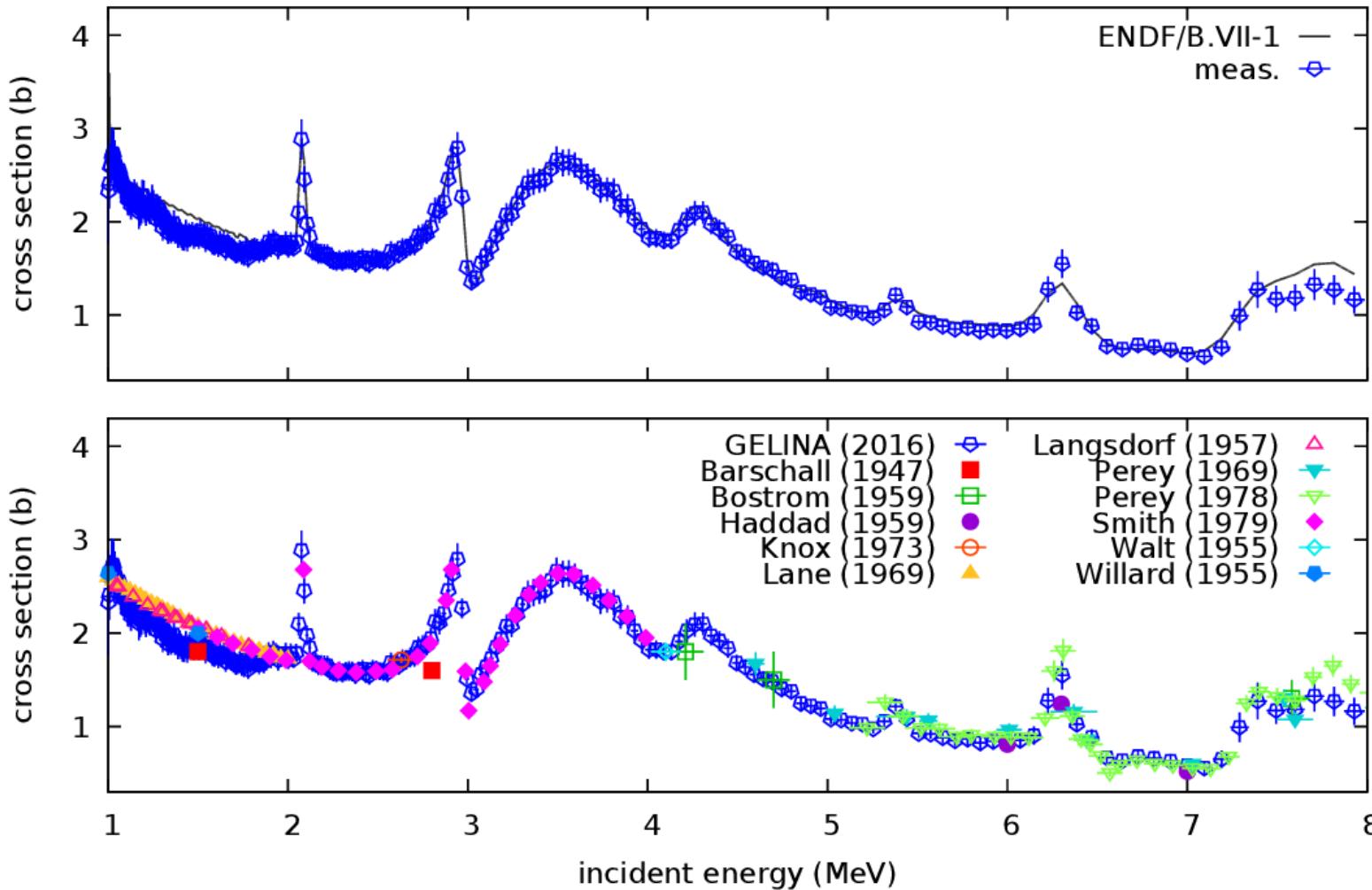
Elastic scattering

— 30 m

ELastic and Inelastic Scattering Array

- 32 liquid organic scintillators
 - 16 EJ301 (NE213)
 - 16 EJ315 (C6D6)
- n/g discrimination
- Time resolution ~1 ns
- Flux by ^{235}U fission chamber
- natC , ^{56}Fe
- Ongoing ^{54}Fe , ^{23}Na , (^{56}Fe enr)

$^{nat}C(n,n)$: test set-up and procedures



GELINA - Experimental set-ups



- 12 HPGe detectors:
 - 4 @ 110°
 - 4 @ 150°
 - 4 @ 125°
- ^{235}U Fission chamber
- Flight Path: 100 m
- Repetition rate: 800 Hz
- Neutron energy 0 – 20 MeV
- Flux \sim 2000 neutrons/cm 2 s

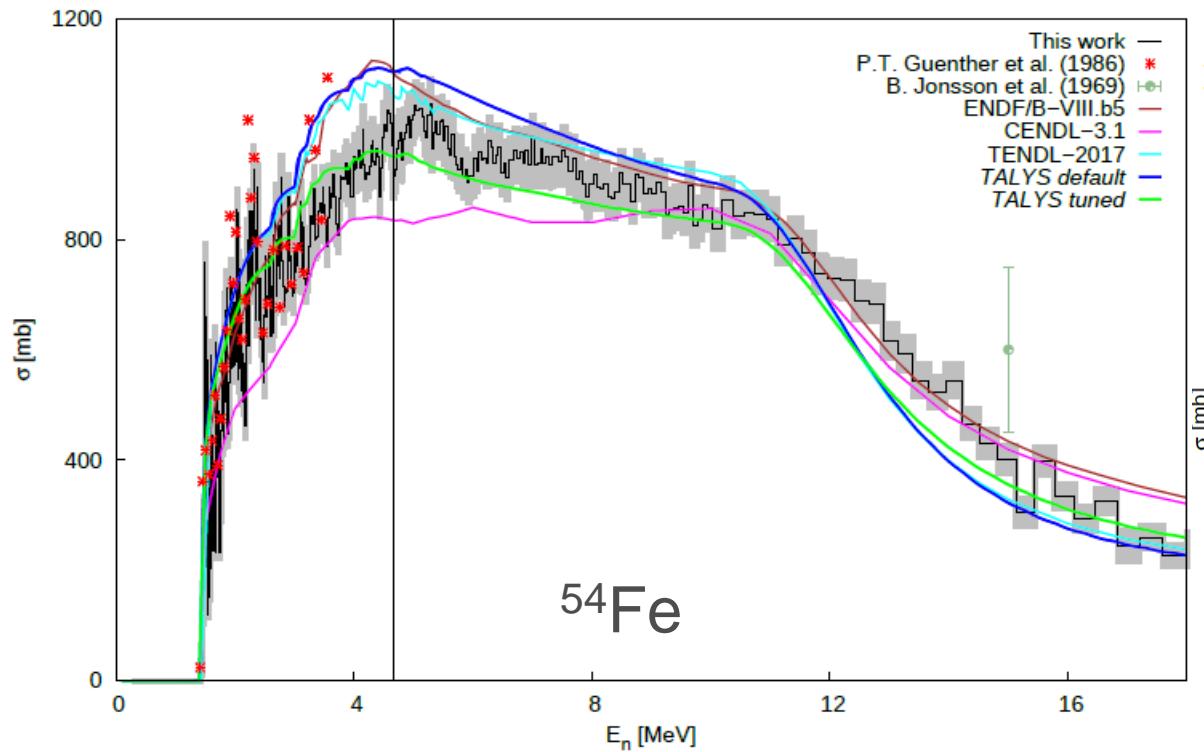
Inelastic scattering

- 100 m, GAINS

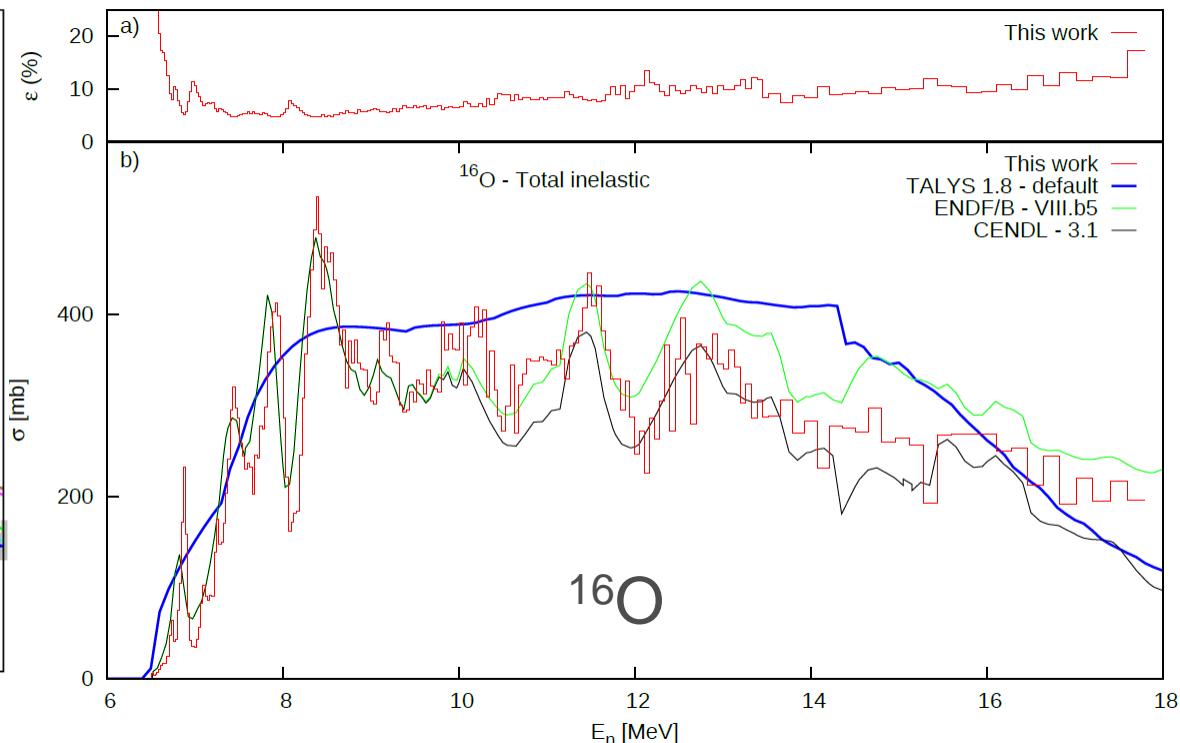


Examples: $^{54}\text{Fe}(n,n'g)$ and $^{16}\text{O}(n,n'g)$

A. Olacel et al., Eur. Phys. Journal A54, 183 (2018)



M. Boromiza et al., Phys. Rev. C101, 024604 (2020)



GELINA - Experimental set-ups

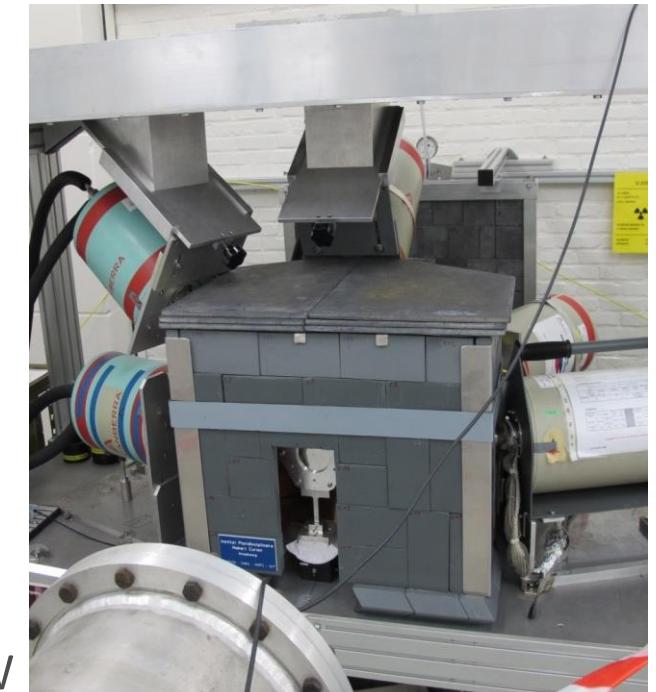


GeRmanium array for Actinides PrEcise MEasurements

- Inelastic scattering set-up
- 5 planar HPGe detectors, one segmented (36 pixels)
- Neutron flux monitoring with a ^{235}U fission chamber
- ^{235}U , ^{238}U , ongoing ^{233}U , ^{232}Th , W

Inelastic scattering

- 30 m, GRAPhEME



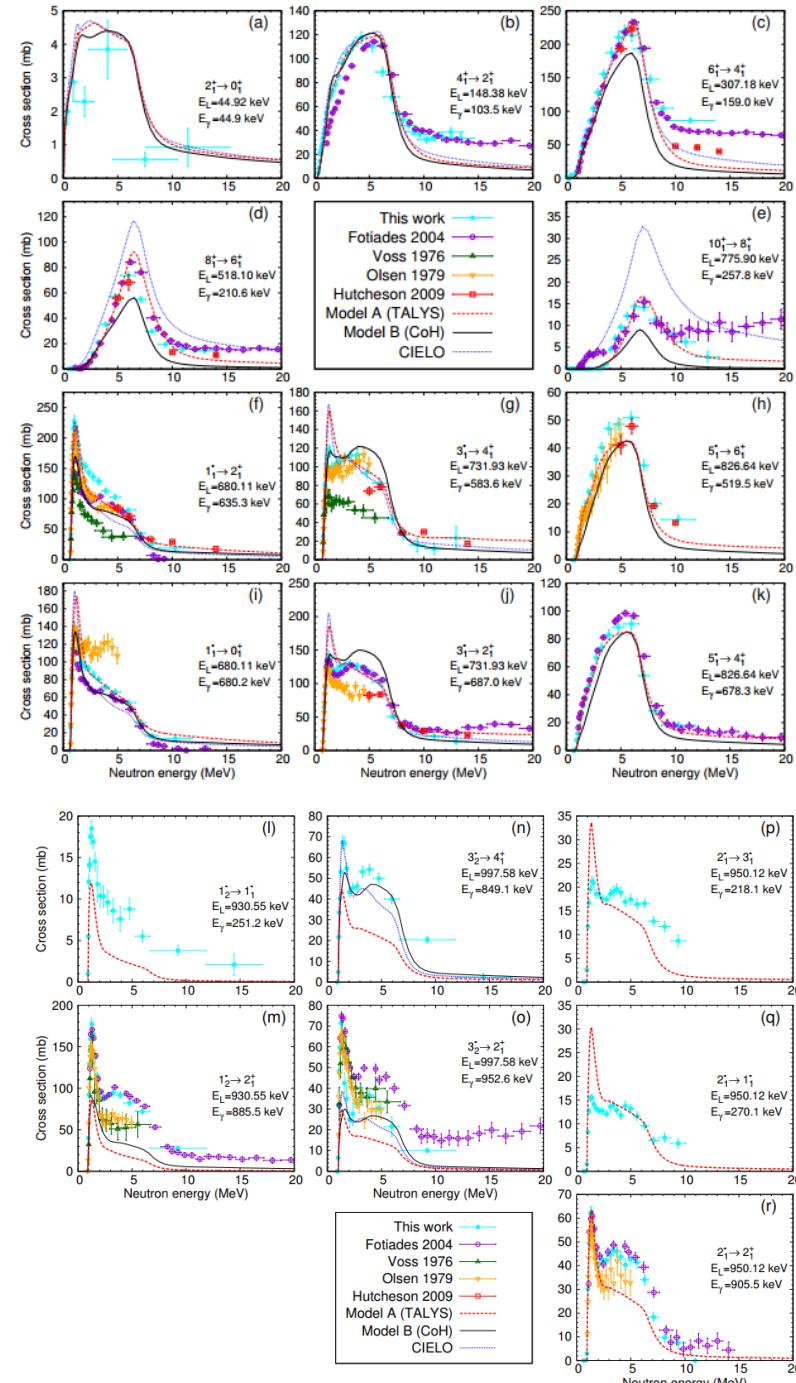
GRAPhEME : ^{238}U measurements

Measurement of $^{238}\text{U}(n, n\gamma)$ cross section data and their impact on reaction models,
M. Kerveno, M. Dupuis *et al.*
Phys. Rev. C **104**, 044605 (2021) DOI :10.1103/PhysRevC.104.044605
- EXFOR entry 22795 (being updated)

PHYSICAL REVIEW C
covering nuclear physics

Main results:

- ▣ Cross section data for **18 γ transitions** (five never measured before) compared to TALYS, EMPIRE and CoH codes.
- ▣ New information about **branching ratio** have been obtained.
- ▣ Microscopic modeling of the preequilibrium emission improved the spin distribution description and then improve the prediction for $(n, n'\gamma)$ transition at high spin level in the GS band.
- ▣ Some of the interband transitions are shown to be very sensitive to the choice of **prescription for discrete levels γ decay** that are not given in the RIPL3 library.
- ▣ Large uncertainties in the **current modeling** are related to the **knowledge of the decay scheme**: poorly known branching ratios and an incomplete information of the discrete states above 1.3 MeV.
- ▣ Other aspects of the modeling, such as **variation of the $E1$ and $M1$ strength functions** that define the γ decay from continuum levels, were **shown to be of importance** and should be studied more in detail in the context of $(n, n\gamma)$ reactions.



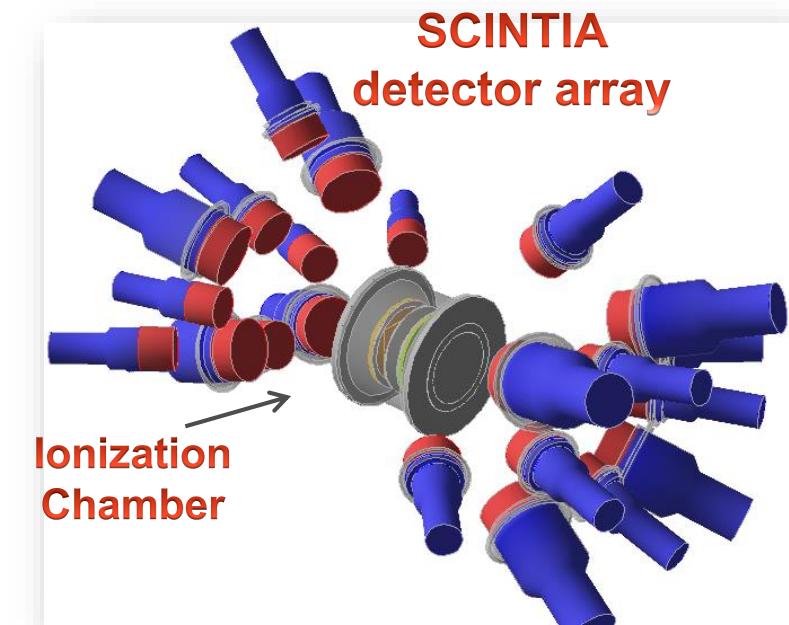
To come next: 20 others ($n, n'\gamma$) XS but also ($n, 2n\gamma$) and ($n, 3n\gamma$) XS

GELINA - Experimental set-ups

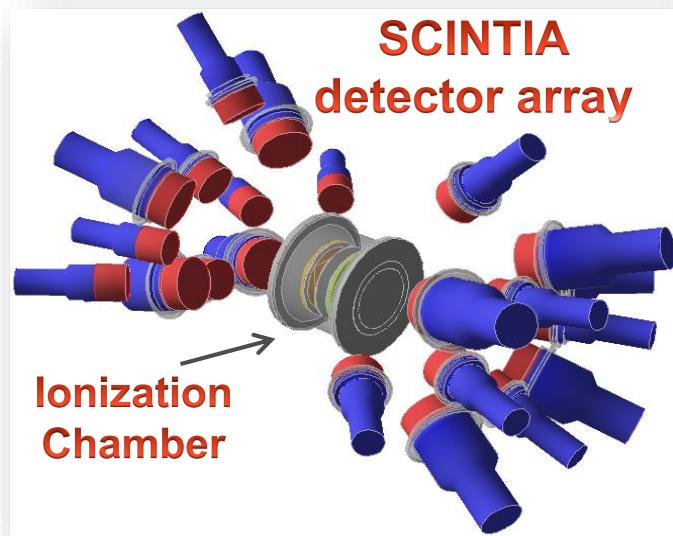


Fission studies

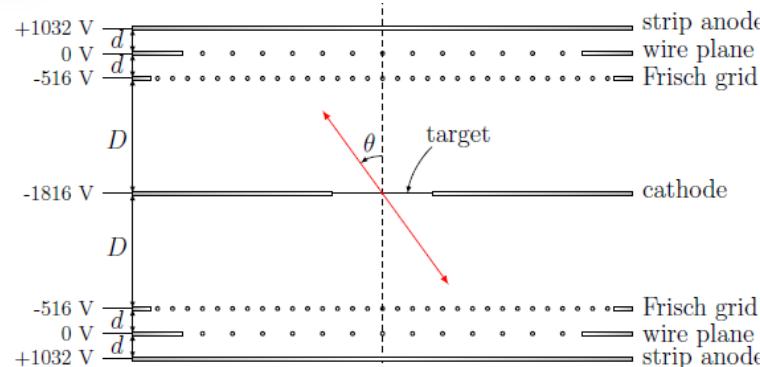
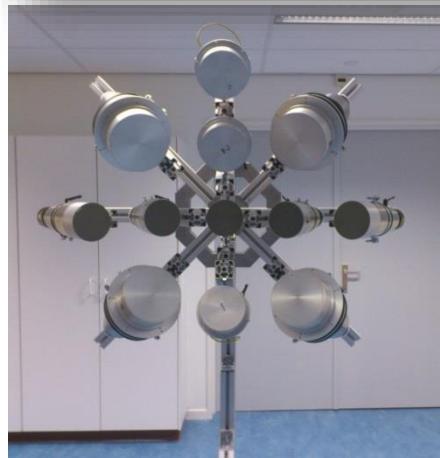
– 10 m



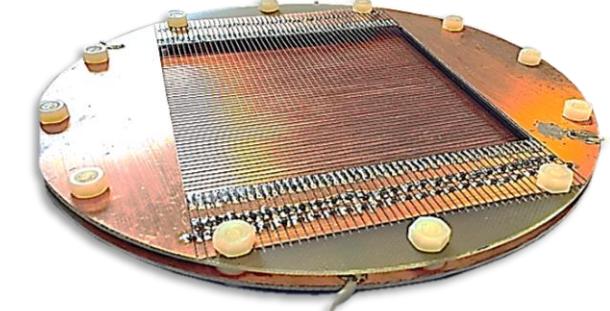
Fission fragment properties and prompt fission neutrons



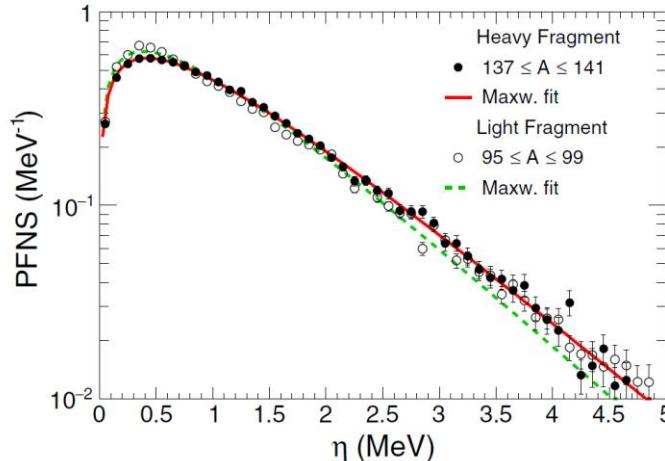
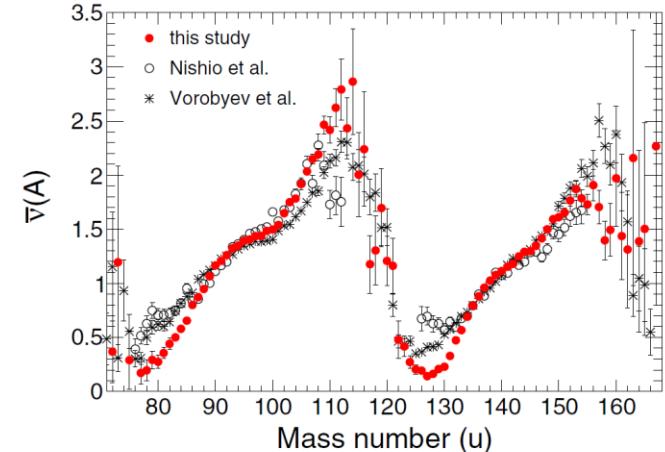
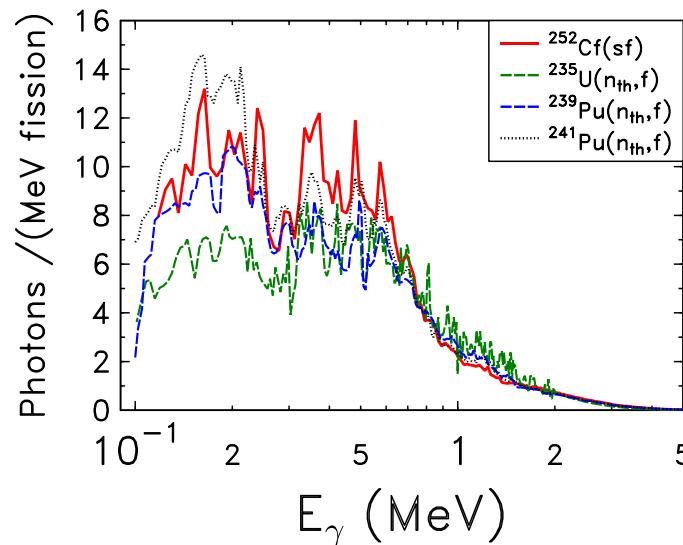
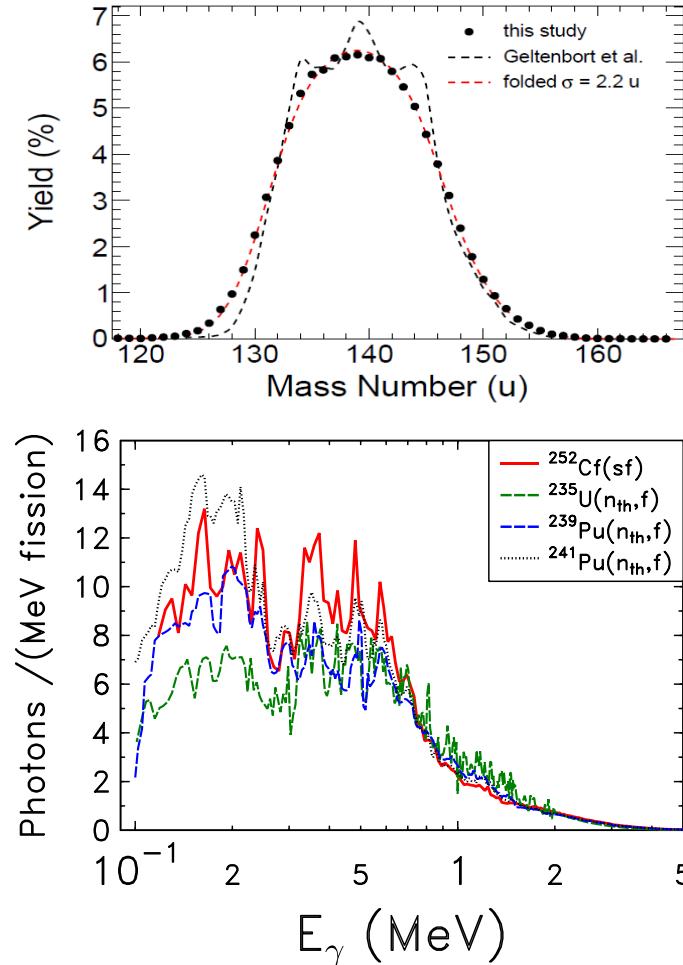
- Fission fragments by **twin position sensitive IC (2PIC)**
 - Fragment energy
 - Fragment masses - 2E-technique
 - Fission axis orientation
- Prompt fission neutrons
 - 22 x Scintillators
 - Energy : time-of-flight



Position sensitive electrode

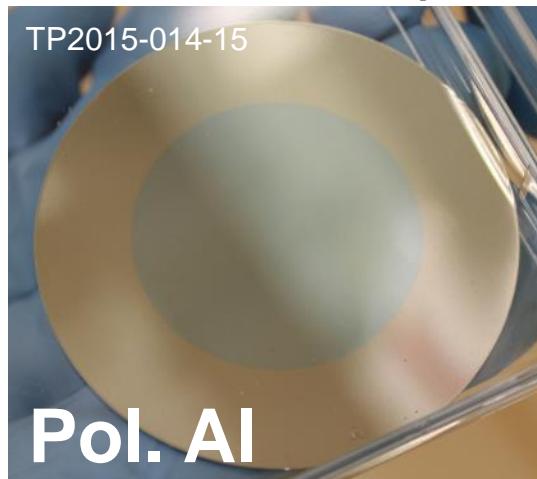


$^{235}\text{U}(\text{n},\text{f})$ FP + prompt fission neutrons + γ -rays



Target Preparation

- Vacuum deposition



material: ^{238}U

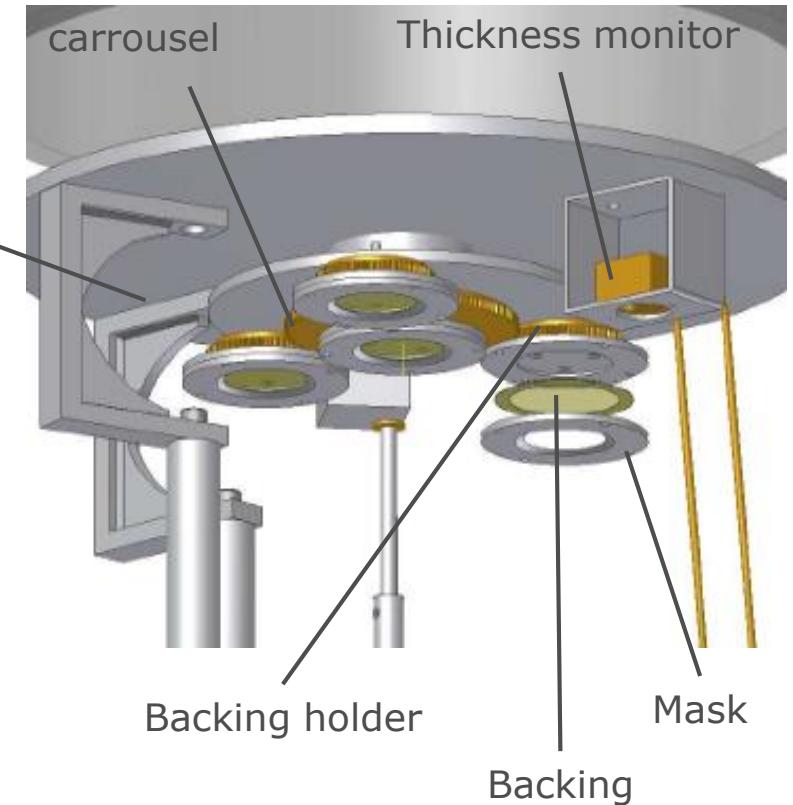
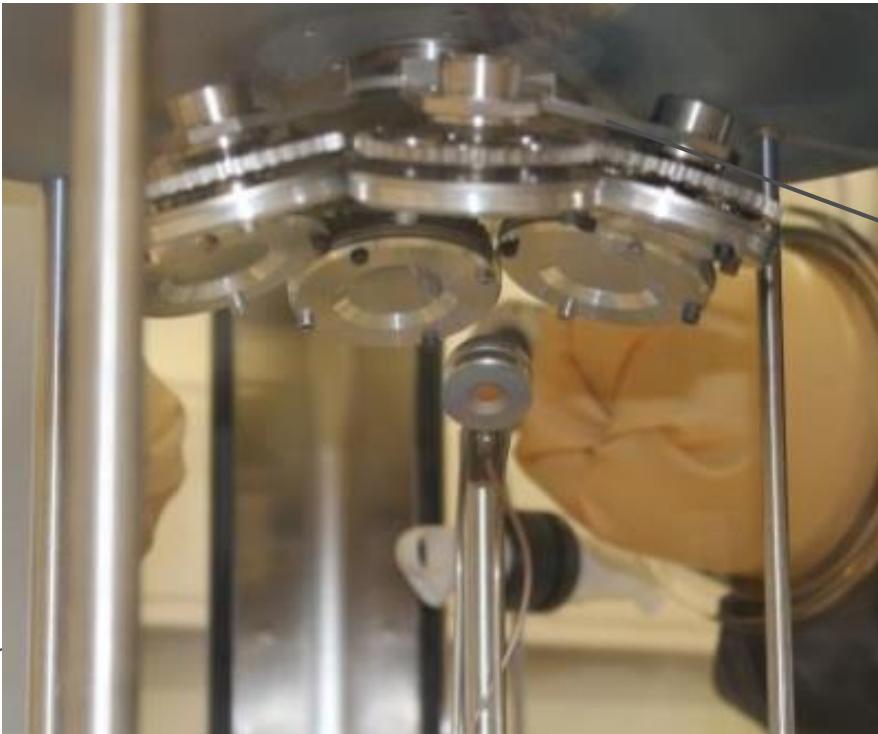
method: vacuum deposition

deposit diameter: 30 mm

substrate: 0.25 mm thick, Ø 50 mm

mass ^{238}U : 0.3 mg

areal density ^{238}U : 48 $\mu\text{g cm}^{-2}$



Target Preparation

- Molecular plating



^{238}U layer

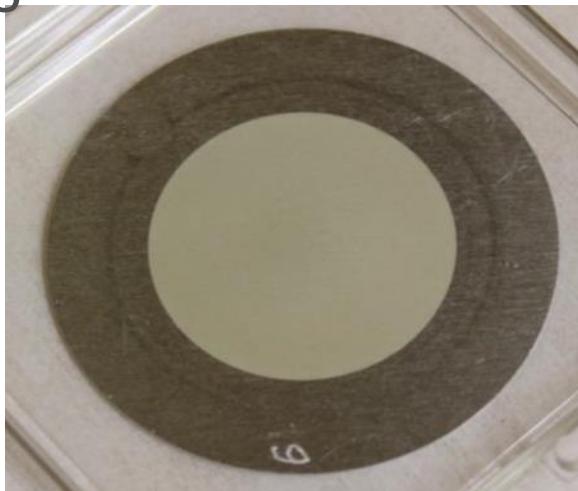
Material: 99.999% ^{238}U

Deposit Ø: 30 mm

Backing: 0.25 mm Al Ø 50 mm

Mass ^{238}U : 1.9 mg

Areal density ^{238}U : 265 $\mu\text{g}\cdot\text{cm}^{-2}$



^{237}Np layer

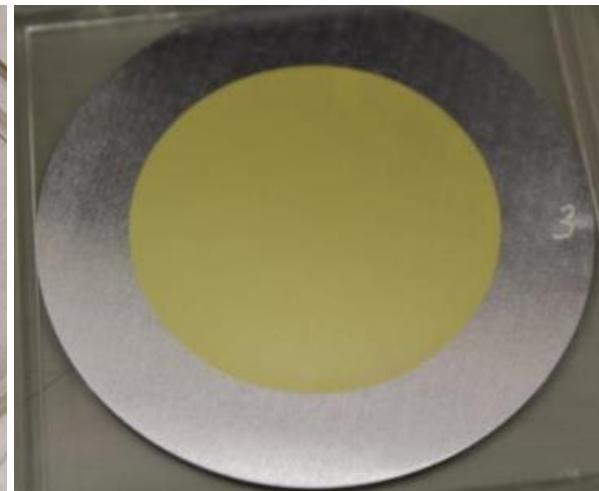
Material: ^{237}Np

Deposit Ø: 30 mm

Backing: 0.25 mm Al Ø 50 mm

Mass ^{237}Np : 1.4 mg

Areal density ^{237}Np : 204 $\mu\text{g}\cdot\text{cm}^{-2}$



^{235}U layer

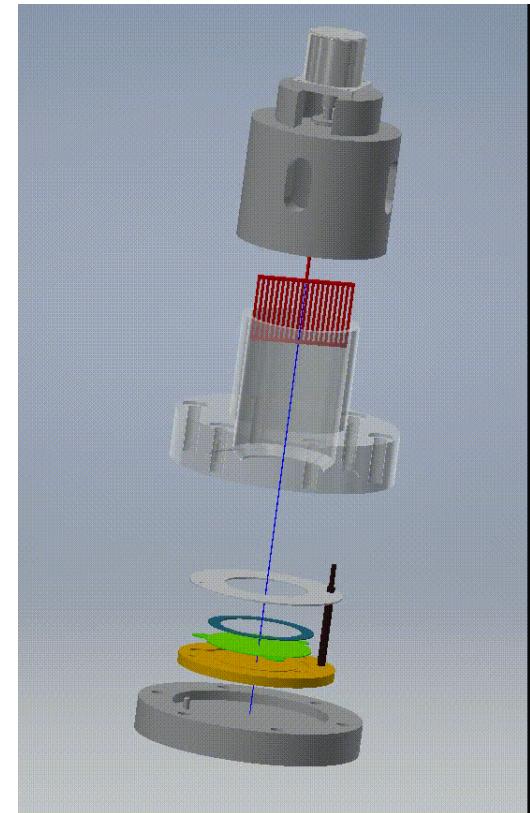
Material: 99.934% ^{235}U

Deposit diameter: 40 mm

Backing: 0.25 mm Al Ø 60 mm

Mass ^{235}U : 3.5 mg

Areal density ^{235}U : 279 $\mu\text{g}\cdot\text{cm}^{-2}$



JRC – CERN collaboration on nuclear data

- **Collaboration**
 - JRC is member of the n_TOF collaboration, collaboration board and editorial board
 - JRC – Geel staff participate in experiments at n_TOF
 - JRC – Geel target preparation facilities used for experiments at n_TOF
 - nTOF collaboration members perform detector tests at GELINA
- **Coordination of nuclear data measurement capabilities and expertise**
- **Pooling and preparation of high quality nuclear targets for nuclear data measurements**
- **Supporting international nuclear data projects (e.g. EXFOR, CIELO, JEFF)**
- **Strengthen partnership and role in European nuclear data networks, promoting infrastructures and transnational access (CHANDA, SANDA)**
- **Knowledge transfer and training**
 - Neutron Resonance Analysis Schools (2008, 2011, 2014, 2017)
 - n_TOF winter school (2018)
 - Mobility grants: PhD students from CERN at JRC Geel

Scientific output

Physics Letters B 804 (2020) 135405

Contents lists available at ScienceDirect

Physics Letters B

www.elsevier.com/locate/physletb

Measurement of the $^{154}\text{Gd}(\text{n},\gamma)$ cross section and its astrophysical implications



Cross section measurements at the n_TOF facility at CERN have been complemented by transmission measurements at GELINA to

- improve the accuracy
- reduce the uncertainty

ABSTRACT

The neutron capture cross section of ^{154}Gd was measured from 1 eV to 300 keV in the experimental area located 185 m from the CERN n_TOF neutron spallation source, using a metallic sample of gadolinium, enriched to 67% in ^{154}Gd . The capture measurement, performed with four C_6D_6 scintillation detectors, has been complemented by a transmission measurement performed at the GELINA time-of-flight facility (JRC-Geel), thus minimising the uncertainty related to sample composition. An accurate Maxwellian averaged capture cross section (MACS) was deduced over the temperature range of interest for s process nucleosynthesis modelling. We report a value of $880(50)$ mb for the MACS at $kT = 30$ keV, significantly lower compared to values available in literature. The new adopted $^{154}\text{Gd}(\text{n},\gamma)$ cross section reduces the discrepancy between observed and calculated solar s-only isotopic abundances predicted by s-process nucleosynthesis models.

MEDIA INAF
Il notiziario online dell'Istituto nazionale di astrofisica

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MISURATE IN LABORATORIO LE PROPRIETÀ DELL'ISOTOPO

Ricetta stellare per il gadolinio-154

Per studiare una discrepanza fra l'abbondanza solare del gadolinio-154 rispetto a quanto previsto dai modelli astrofisici, due astronomi dell'Inaf d'Abruzzo – Sergio Cristallo e Luciano Piersanti – hanno chiesto aiuto ai loro colleghi fisici nucleari sperimentali, fra i quali alcuni ricercatori dell'Infn, per compiere nuove misure. I risultati, ottenuti con gli esperimenti n_Tof e Gelina, hanno risolto la discrepanza

Redazione Media Inaf 10/04/2020



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In questo periodo di forzata permanenza in casa, uno sguardo dalla finestra ci fa apprezzare l'immobile tranquillità e laboriosità della natura. Ovviamente anche il Sole, così lontano da noi, prosegue indisturbato la sua opera di trasformazione di elementi chimici leggeri in elementi più pesanti. E lo stesso fanno le altre stelle nell'universo. Anche gli astrofisici, che si occupano di questi processi fisici, non hanno interrotto le loro ricerche e hanno continuato a interrogarsi sul funzionamen-



Cross sections for neutron induced capture reactions are of primary importance for a better understanding of the origin of elements heavier than iron through stellar nucleosynthesis modelling.

Scientific output

PHYSICAL REVIEW C **104**, L032803 (2021)

Letter

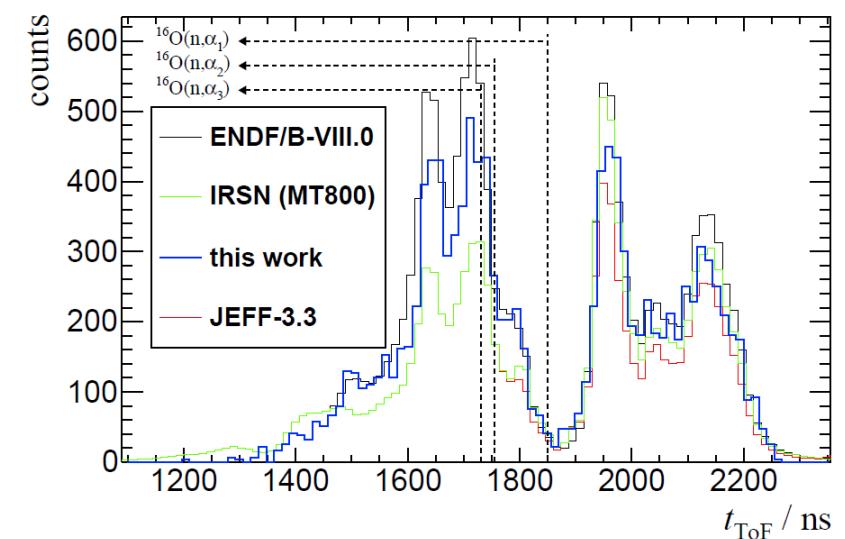
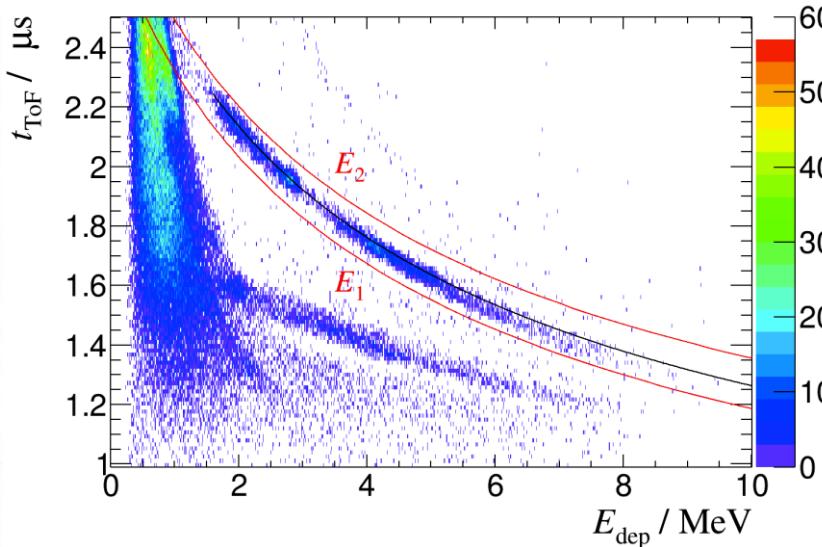
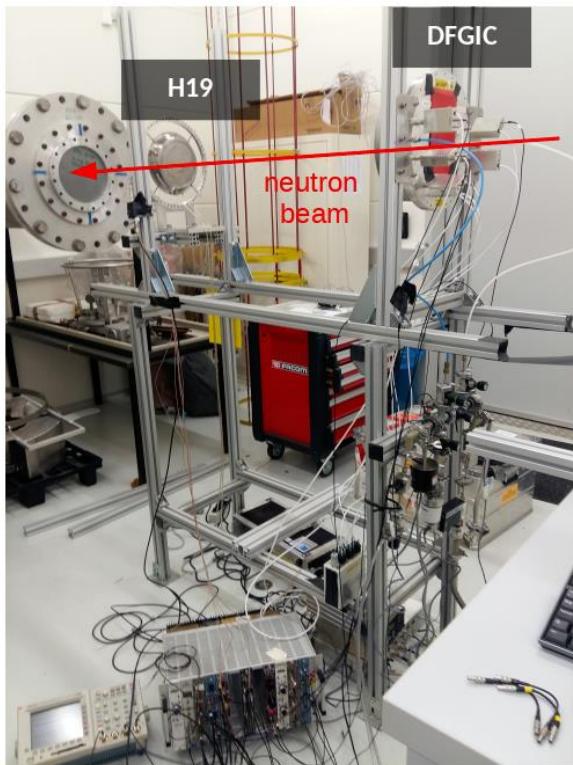
Destruction of the cosmic γ -ray emitter ^{26}Al in massive stars: Study of the key $^{26}\text{Al}(n, \alpha)$ reaction

C. Lederer-Woods,^{1,*} P. J. Woods,¹ T. Davinson,¹ A. Estrade,^{1,†} J. Heyse,² D. Kahl,^{1,‡} S. J. Lonsdale,¹ C. Paradela,² P. Schillebeeckx,² O. Aberle,³ S. Amaducci,^{4,5} J. Andrzejewski,⁶ L. Audouin,⁷ M. Bacak,^{8,3,9} J. Balibrea,¹⁰ M. Barbagallo,¹¹

Neutron destruction reactions of the cosmic γ -ray emitter ^{26}Al are of importance to determine the amount of ^{26}Al ejected into our galaxy by supernova explosions and for ^{26}Al production in asymptotic giant branch stars. We performed a new measurement of the $^{26}\text{Al}(n, \alpha)$ reaction up to 160-keV neutron energy at the neutron time-of-flight facilities n_TOF at CERN and GELINA at EC-JRC. We provide strengths for ten resonances, six of them for the first time. We use our data to calculate astrophysical reactivities for stellar temperatures up to 0.7 GK. Our results resolve a discrepancy between the two previous direct measurements of this reaction, and indicate higher stellar destruction rates than the most recently recommended reactivity.

$^{16}\text{O}(\text{n},\text{a})^{13}\text{C}$

- Sebastian Urlass, CERN, HZDR



S. Urlass¹, A. Plompen², A. Junghans¹, S. Kopecky², R. Beyer¹,
T. Kögler¹, A. Göök², M. Nyman², C. Paradela², P. Schillebeeckx² and
L. Tassan-Got³.

Summary

- JRC – Geel infrastructure for neutron data measurements:
 - GELINA neutron time-of-flight facility
 - MONNET quasi mono-energetic neutron facility
 - Target preparation laboratory
- Support to nuclear data applications – evaluated nuclear data libraries (JEFF)
 - Neutron induced reactions
 - Fission observables
 - Detector development
 - Target preparation
 - Validation measurements
 - Evaluations
- Emphasis on collaborations with stakeholders
- Open Access (EUFRAT)

<https://ec.europa.eu/jrc/en/research-facility/open-access>