



# The CONNIE experiment: latest results and upgrade



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on behalf of the CONNIE Collaboration

**Magnificent CE $\nu$ NS 2021**  
October 6, 2021

# The CONNIE collaboration

~30 members from 6 countries



**USA**

Fermi National Accelerator Laboratory



**Mexico**

Universidad Nacional Autónoma de México



**Argentina**

Centro Atómico Bariloche  
Universidad de Buenos Aires  
Universidad del Sur / CONICET  
ICAS / ICIFI / UNSAM



**Switzerland**

University of Zurich



**Paraguay**

Universidad Nacional de Asunción



**Brazil**

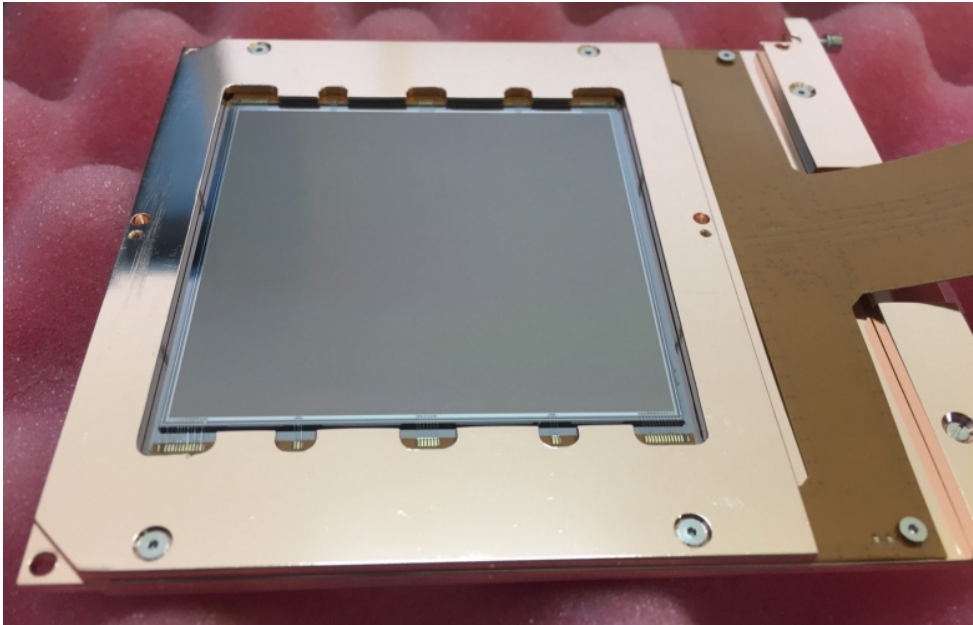
Centro Brasileiro de Pesquisas Físicas  
Universidade Federal do Rio de Janeiro  
CEFET / Angra

# COherent Neutrino Nucleus Interaction Experiment (CONNIE)

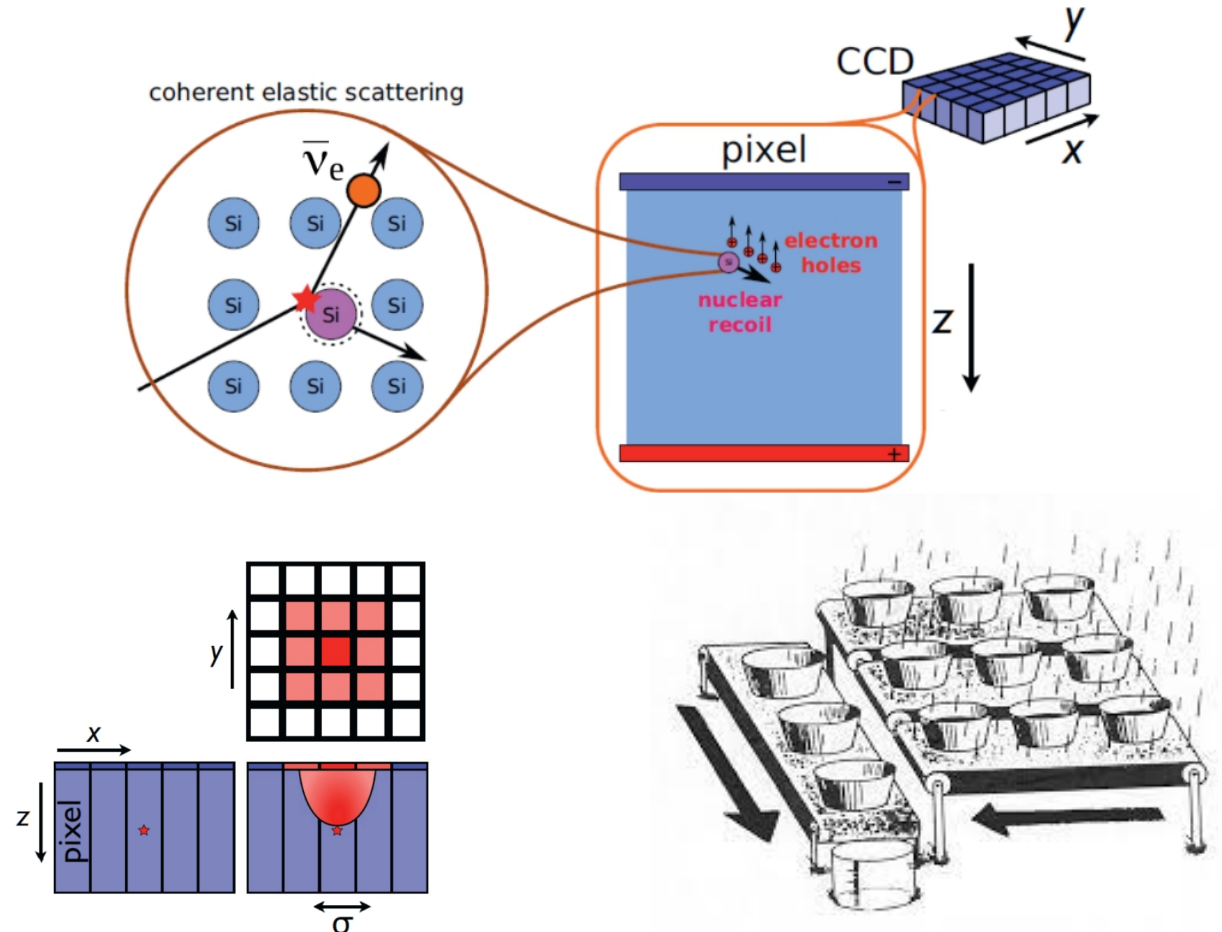
Goal → Measure  $CE\nu NS$  of reactor antineutrinos with Silicon nuclei and probe BSM physics

Detectors → Scientific-grade Charge-Coupled Devices (CCDs) manufactured with high resistivity Si

- ▶ Low radioactive backgrounds
- ▶ Low noise and dark current (DC)
- ▶ Low-energy detection threshold

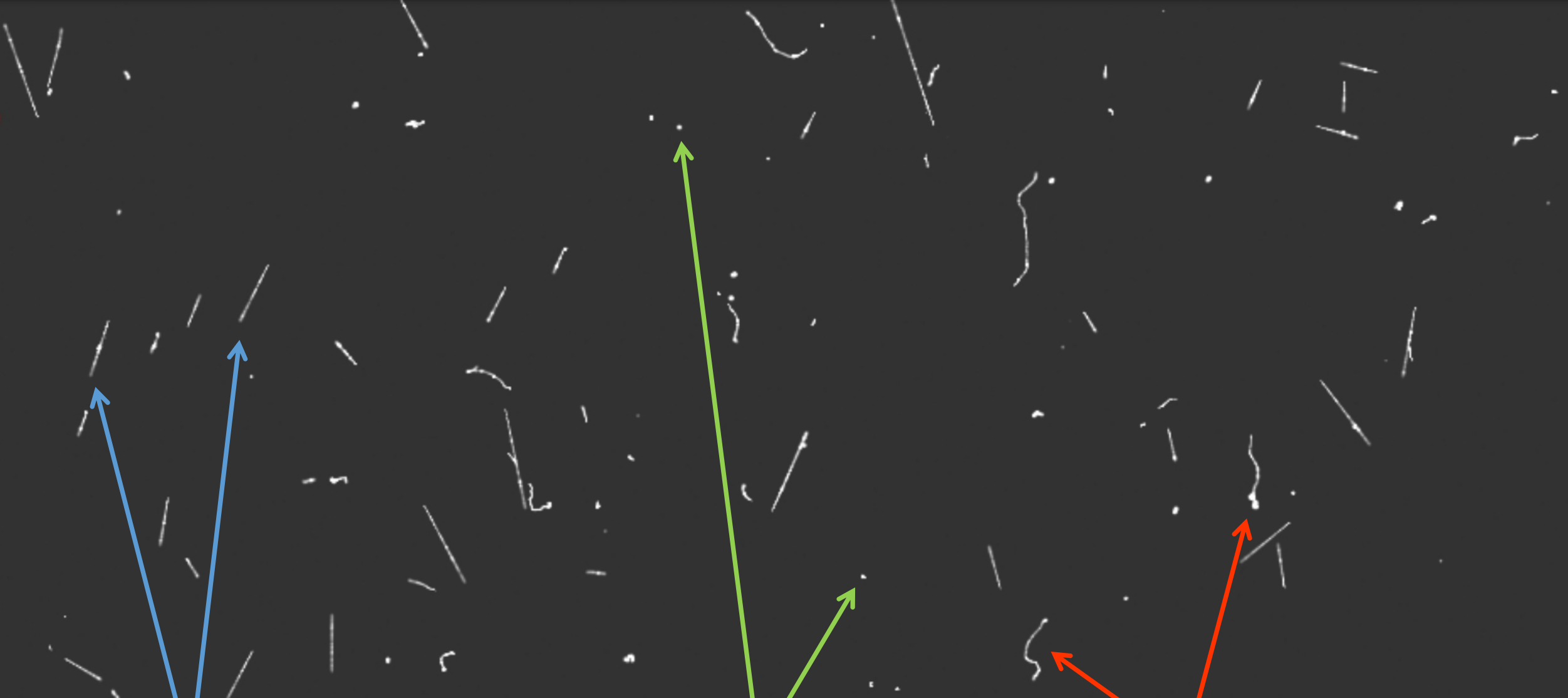


▶ Operating at  $T < 100$  K



Binning improves signal-to-noise ratio

CCDs allow particle identification



muons

diffusion-limited hits

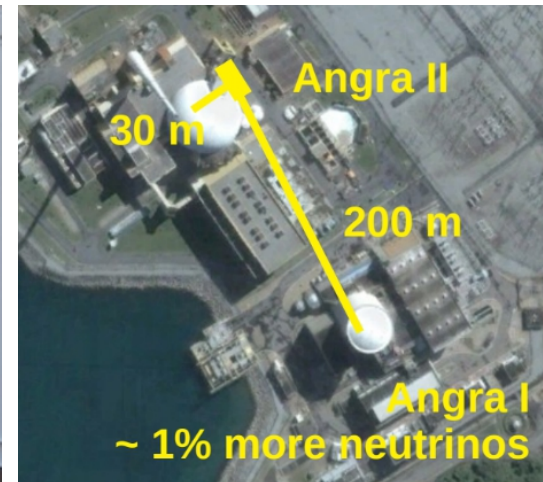
electrons



# COherent Neutrino Nucleus Interaction Experiment (CONNIE)

Site → ~30 m from the 3.8 GW<sub>th</sub> Angra 2 reactor at the Almirante Álvaro Alberto nuclear power plant, in Angra dos Reis, Brazil

- ▶ Flux of  $7.8 \times 10^{12} \tilde{\nu}/s \text{ cm}^2$  at CONNIE
- ▶ Reactor shutdown 1 month every ~13 months (ROFF)





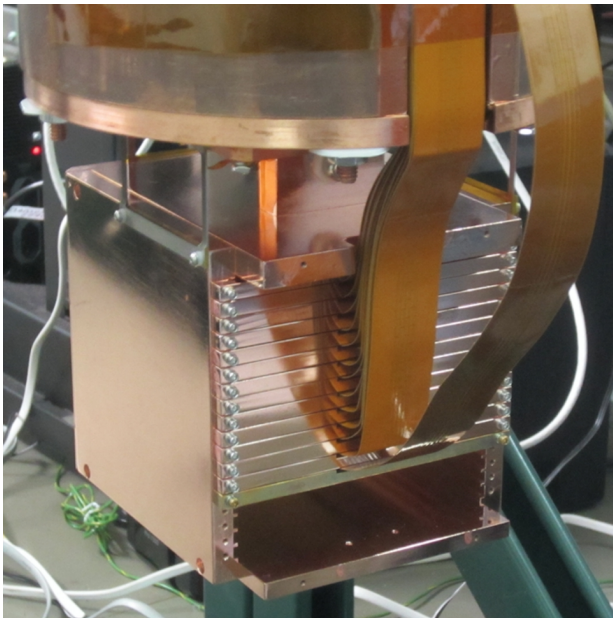
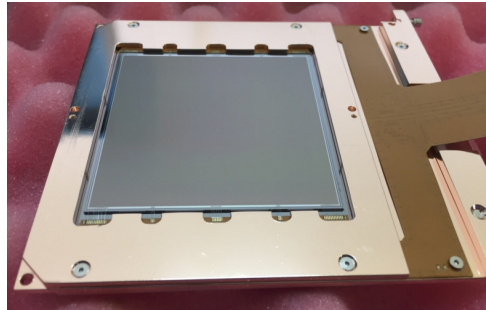
# CONNIE timeline and milestones

Installed at Angra in 2014 (engineering run)

First detector upgrade: Aug 2016 → 14 CCDs (5.95 grams each) developed at LBNL installed at Angra

8 CCDs with stable operation and good quality data running until Dec 2020

- ▶ 4k x 4k pixels  
(15  $\mu\text{m}$  x 15  $\mu\text{m}$  per pix)
- ▶ 675  $\mu\text{m}$  thick



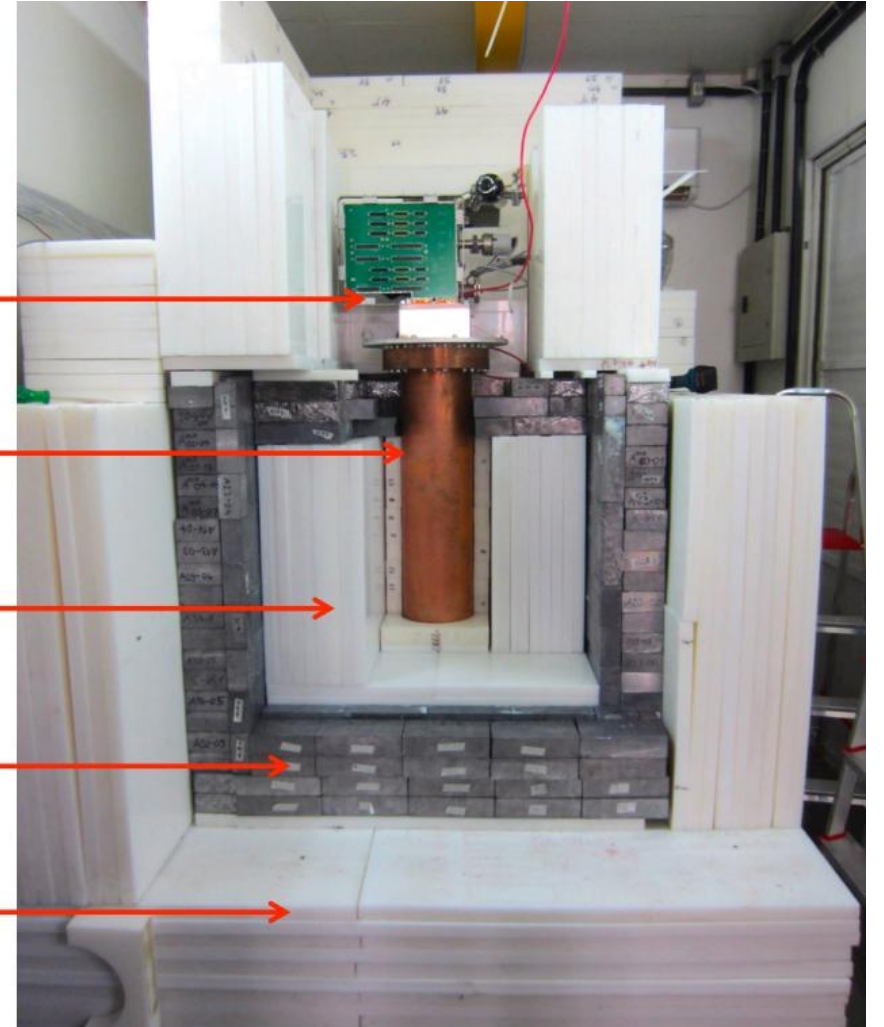
VIB readout board

Dewar in vacuum

Inner polyethylene ~30 cm  
(neutrons shield)

Lead ~15 cm  
(gammas shield)

Outer polyethylene ~30 cm  
(neutrons shield)



# CONNIE timeline and milestones

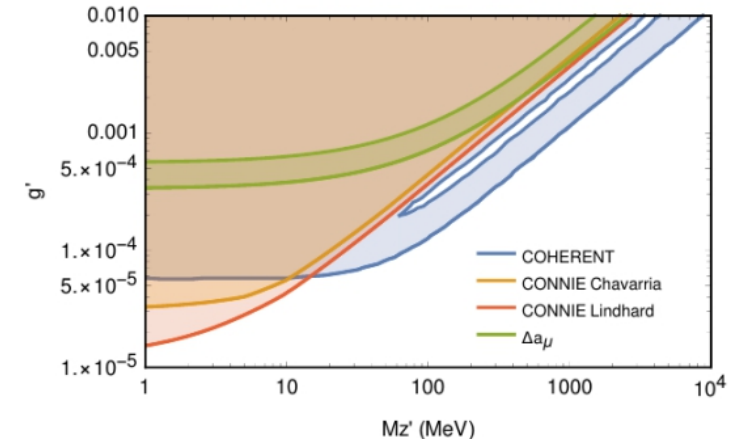
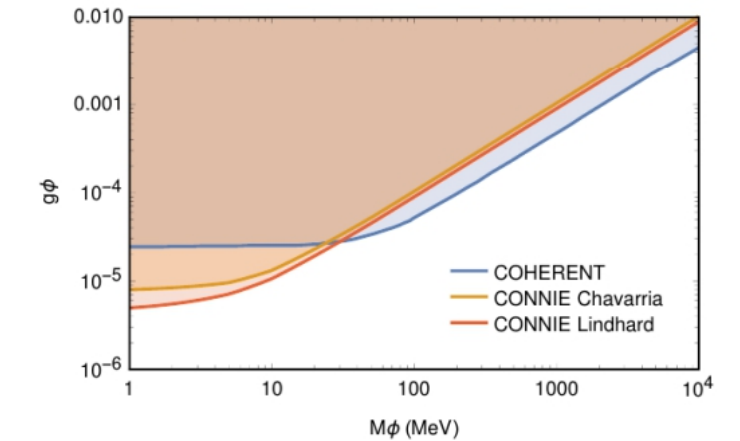
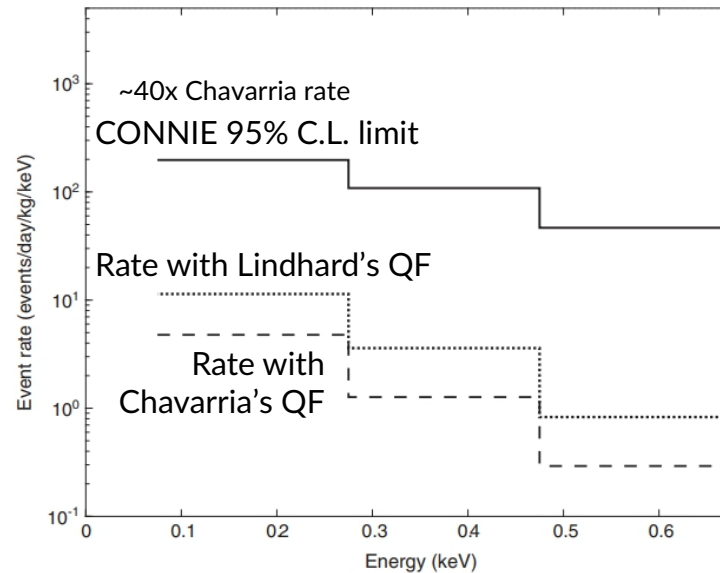
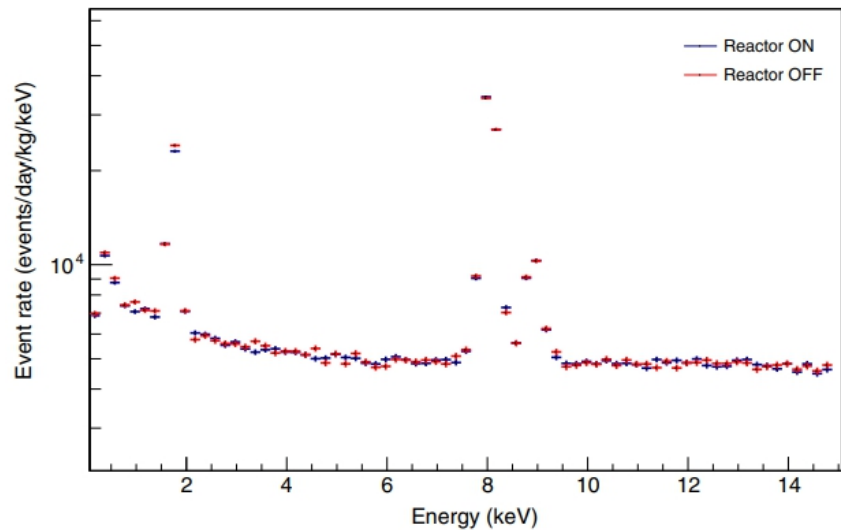
**2016-2018 run data:** 2.1 kg-day with RON / 1.6 kg-day with ROFF → 1x1 binning

Nov 2019 → Published 2016-2018 run data results [**PRD 100 (2019), 092005**]

▶ Establish a model independent limit on  $CE\nu NS$  rate

Apr 2020 → Published limits on simplified SM extensions with light mediators [**JHEP 04 (2020), 054**]

▶ First competitive BSM constraints from  $CE\nu NS$  at reactors



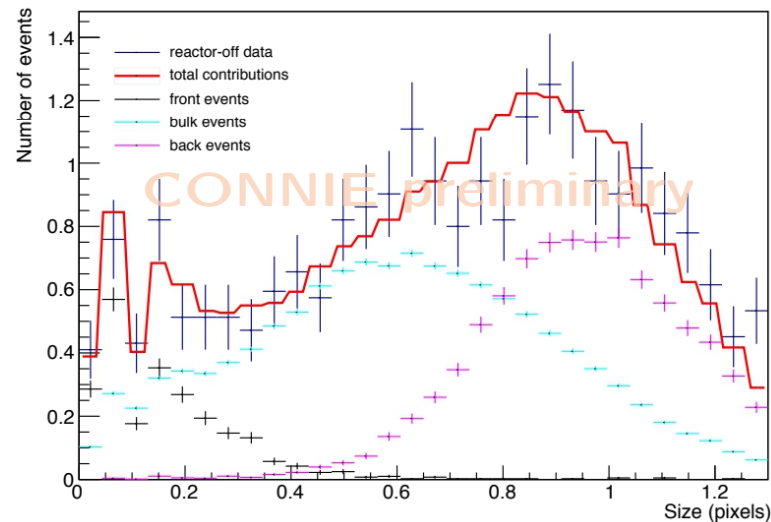
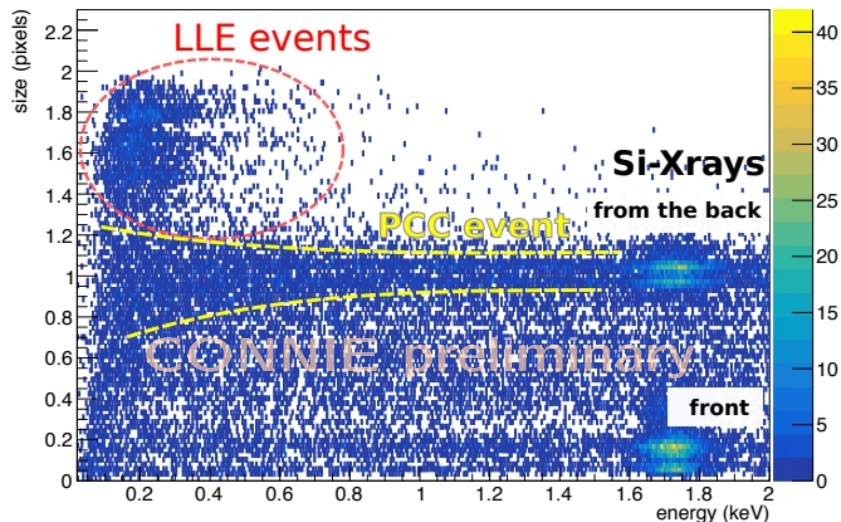
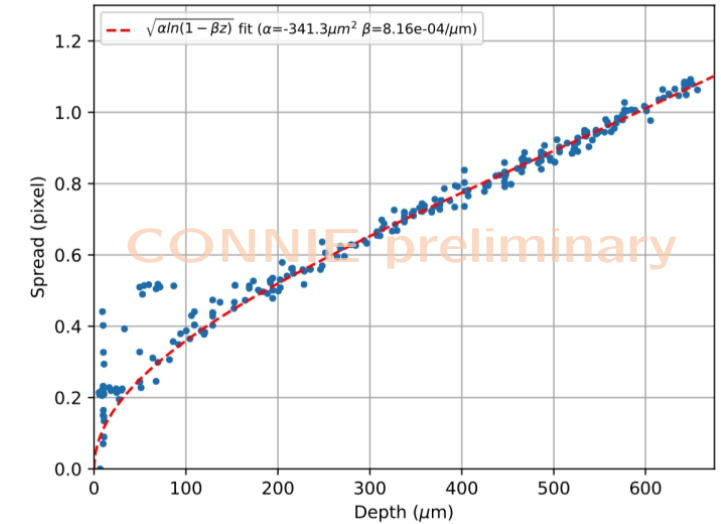
# 2019 run

Paper coming soon!

Jan 2019 → New readout configuration with 1x5 binning (lower noise) and 1 hr exposure (lower DC)

## Analysis:

- ▶ Improved energy and size-depth calibrations
- ▶ Better low-energy background characterization and rejection
  - Large low-energy (LLE) events
  - Partial-charge-collection (PCC) layer events
- ▶ Implemented spatial uniformity check
- ▶ Perform multiple cross-checks



## Blind analysis

- ▶ Freeze analysis parameters with ROFF data
- ▶ Stability checks with mid to high energy RON data
- ▶ Unblind low energy RON data



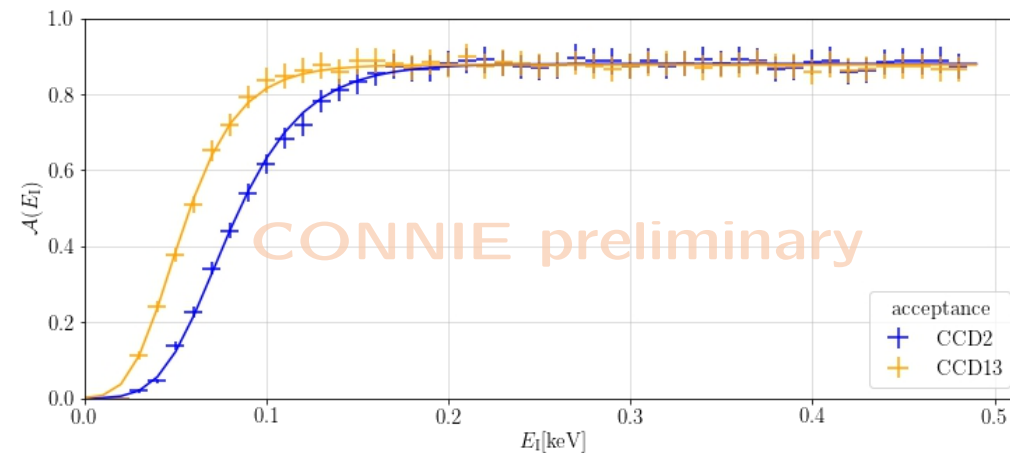
# 2019 run

Improved detector extraction acceptance and selection efficiency at low energies:

Threshold reduced to ~40 eV

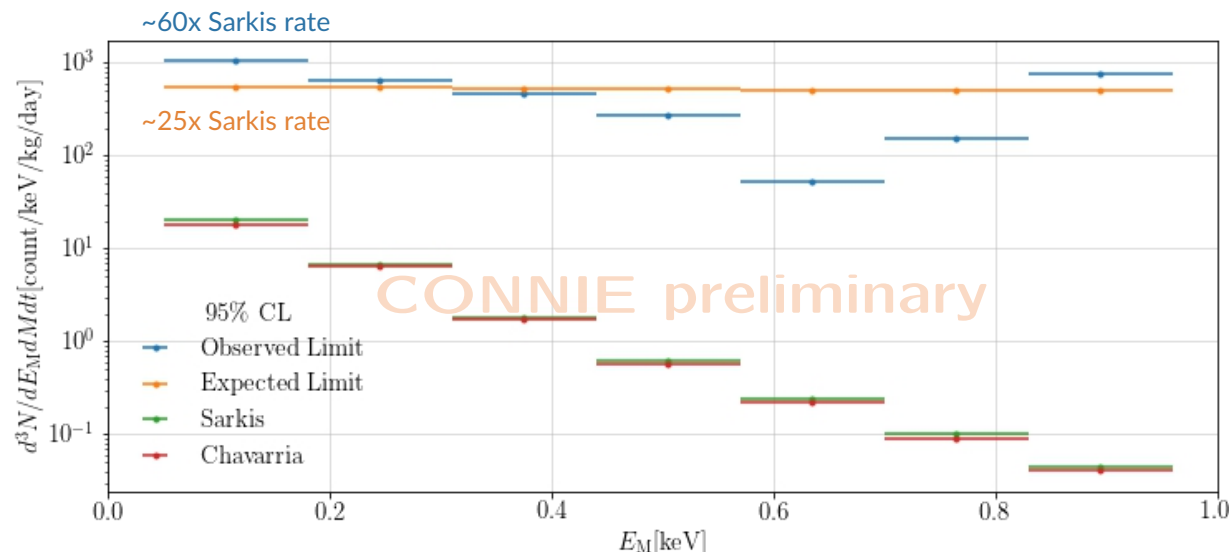
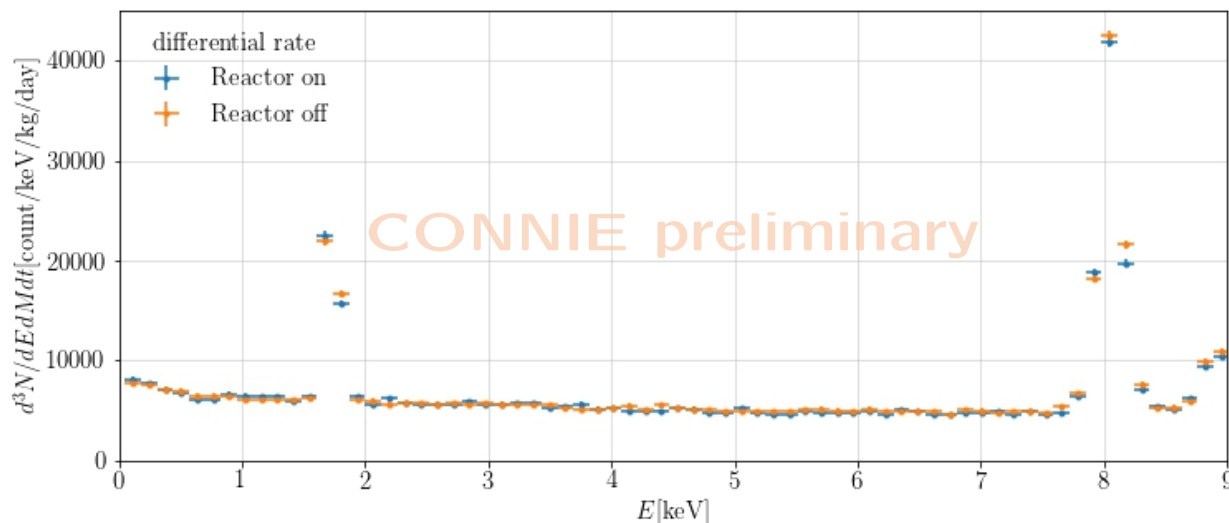
Full acceptance reached at 100-150 eV

Paper coming soon!



Total exposure: 2.7 kg-days (31.85 days with RON and 28.25 days with ROFF)

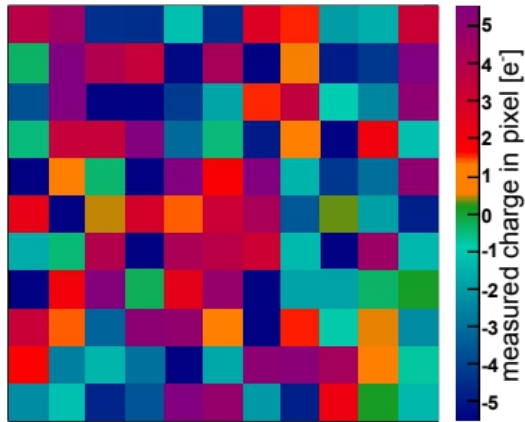
RON-ROFF consistent with zero  $\rightarrow$  95% C.L. limit on observed (expected) CE $\nu$ NS rate



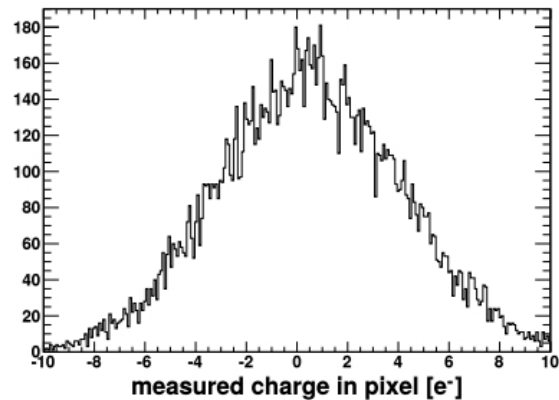
# Skipper-CCD technology

- ▶ Allows multiple sampling of each pixel during data acquisition
- ▶ Reduces readout noise with number of samplings  $\sigma \propto 1/\sqrt{N}$
- ▶ Allows detecting single electrons!
- ▶ Promising for neutrino and dark matter detection

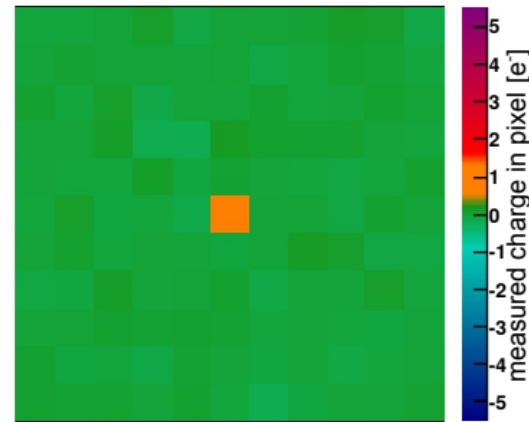
Standard CCD mode: charge in each pixel is measured once



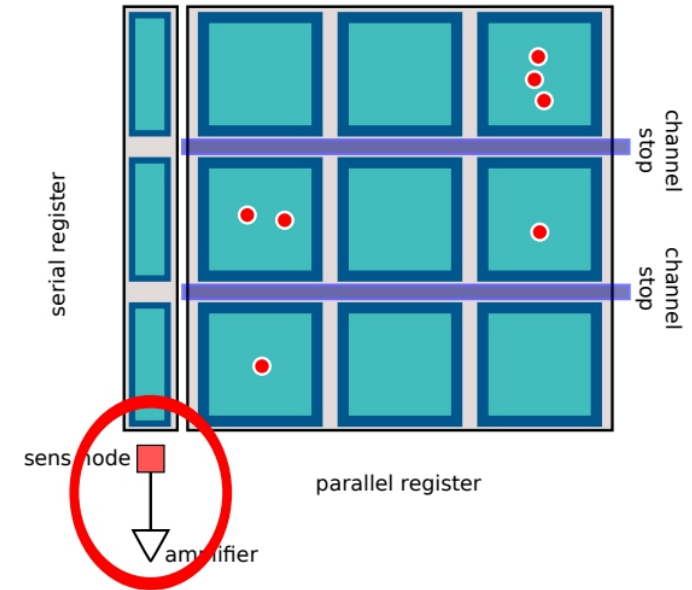
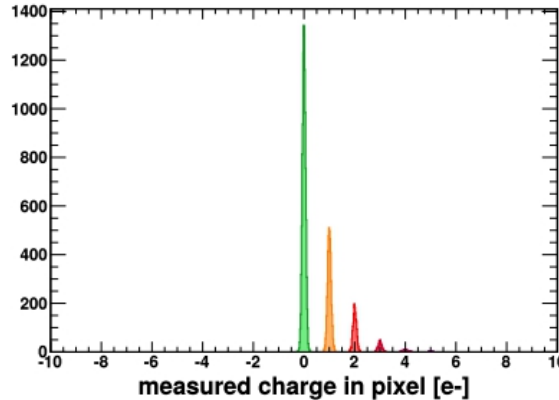
Readout-noise: 3.5 e RMS



New Skipper CCD: charge in each pixel is measured multiple times



Readout-noise: 0.06 e RMS



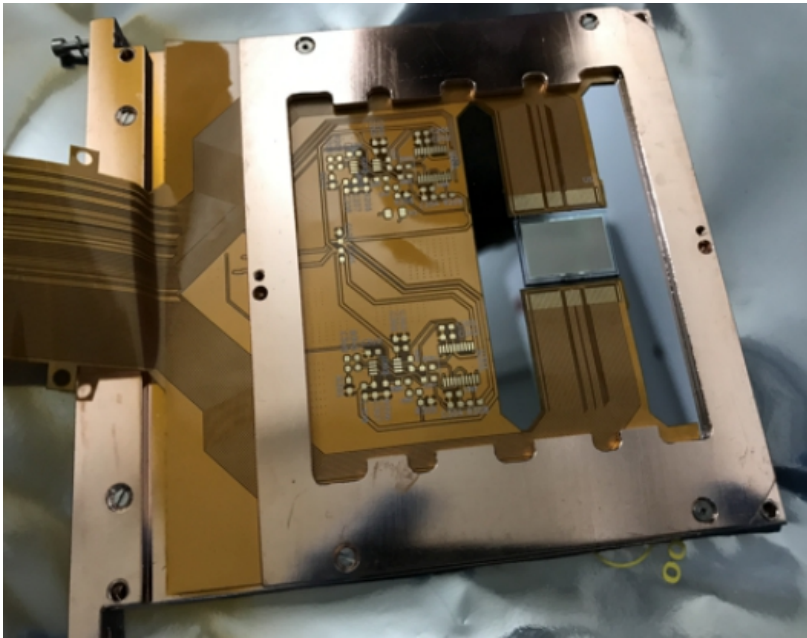
**Only the readout stage is modified**

Skipper technology first demonstrated in 2017 [PRL 119, 131802] using a detector designed by Stephen Holland (LBNL)  
SENSEI DM experiment currently using skipper CCDs [PRL 125, 171802]

# CONNIE upgrade with skipper-CCDs

Jul 2021 → 2 skipper CCDs (768 x 1024 pixels each) installed at Angra

- ▶ New Low Threshold Acquisition readout electronics [JATIS 7 (2021), 1 015001]
- ▶ New dedicated Vacuum Interface Board



Goals (towards next generation experiments)

- ▶ Study skipper-CCDs performance and background at sea level
- ▶ Test LTA electronics

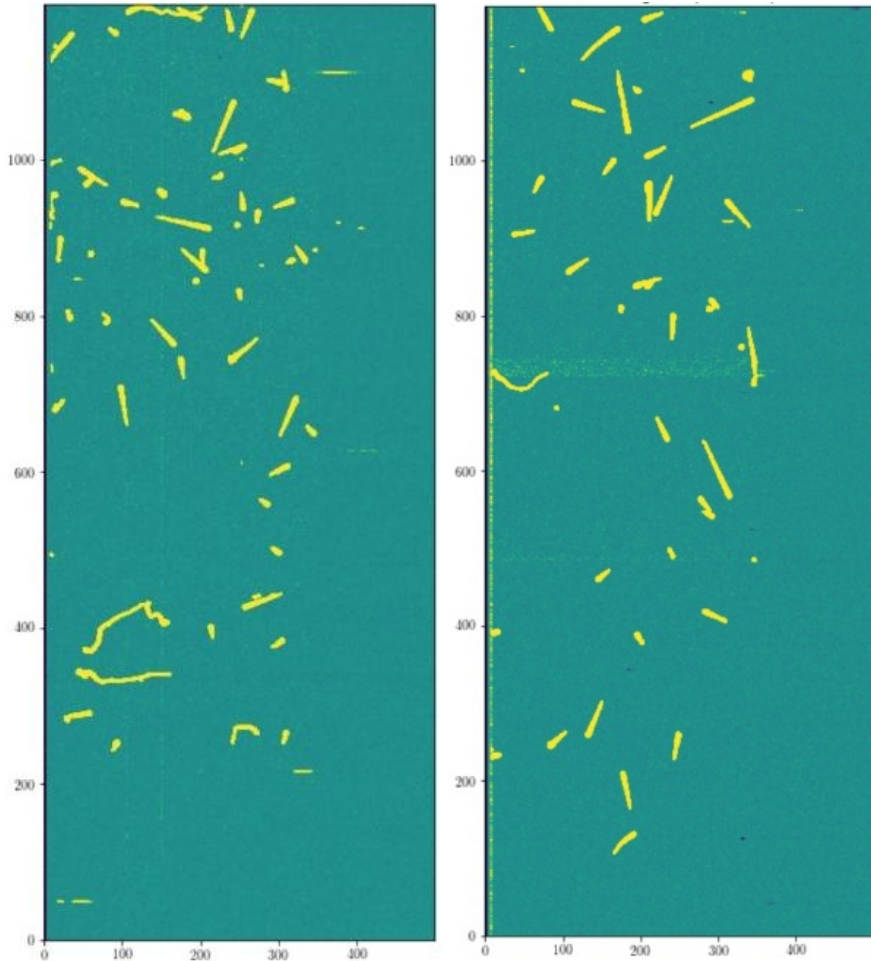
More about installing skipper CCDs at reactors  
See Guillermo Fernandez Moroni talk tomorrow!

# CONNIE upgrade with skipper-CCDs

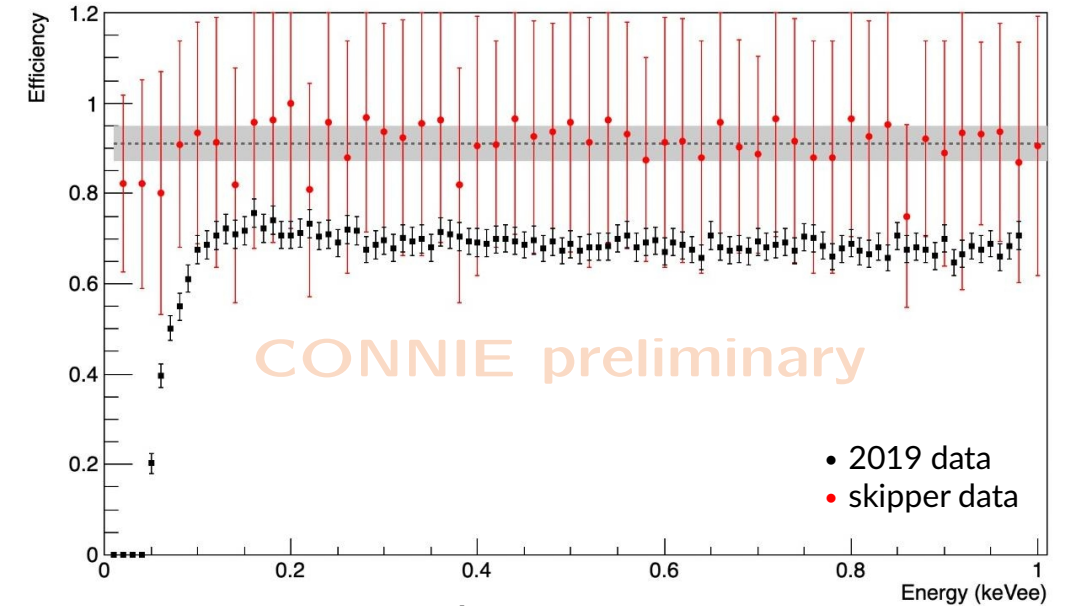
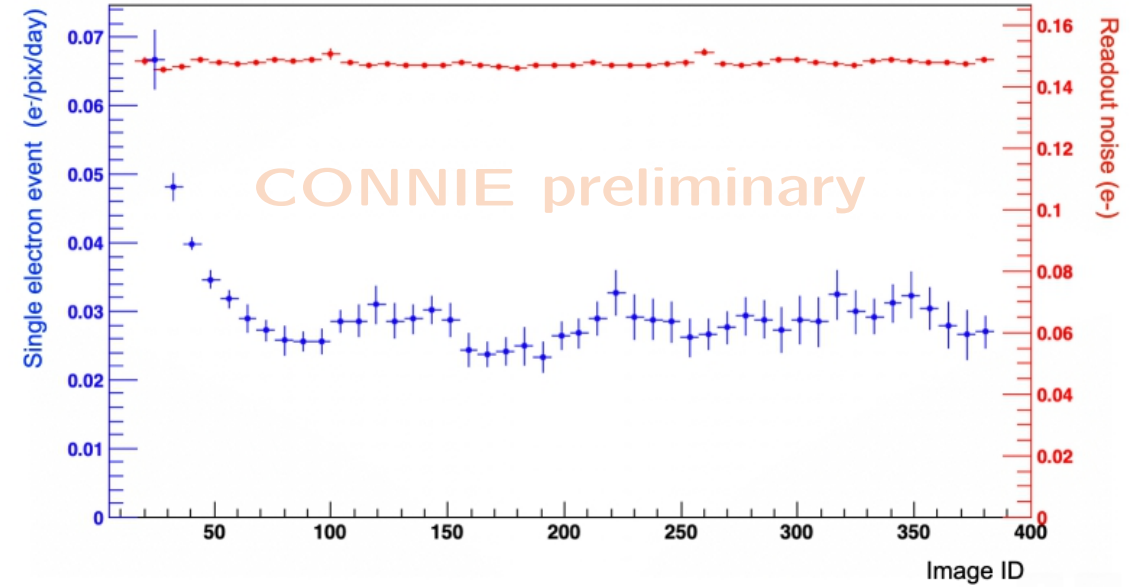
Ongoing data taking

Running stable → Noise:  $\sim 0.15$  e-

DC:  $\sim 0.03$  e-/pix day



Images with 400 samples/pix



Event selection efficiency

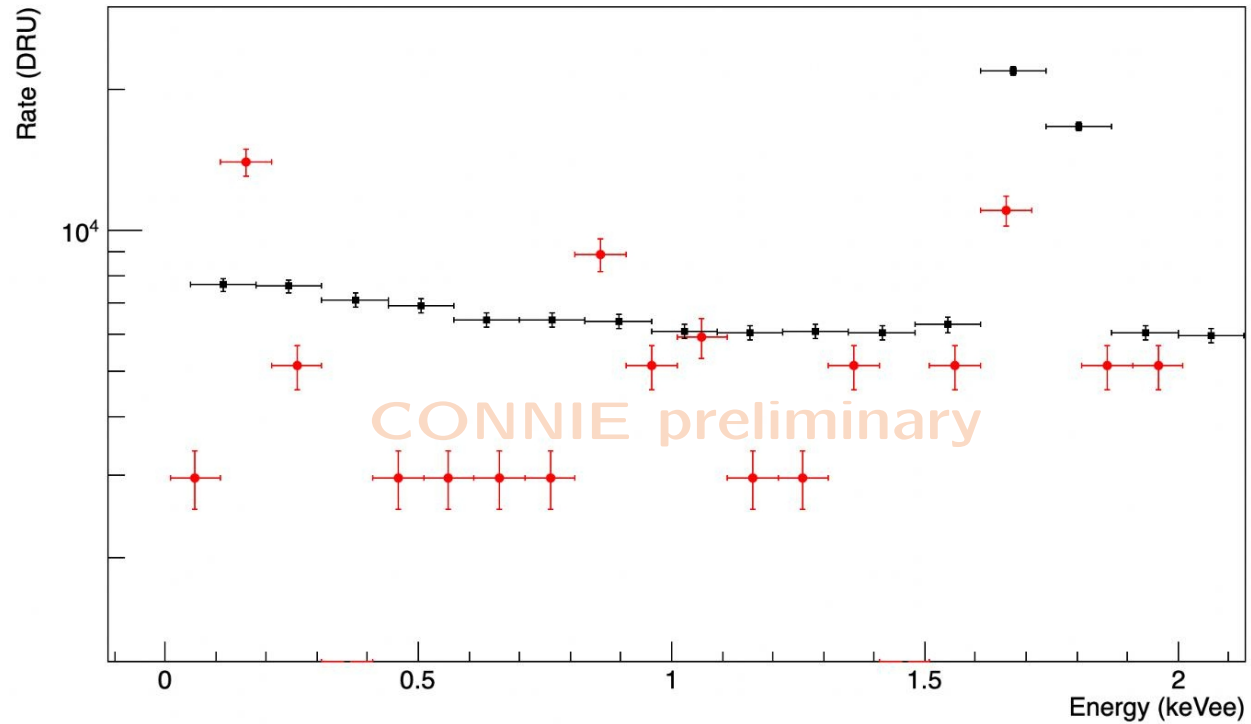
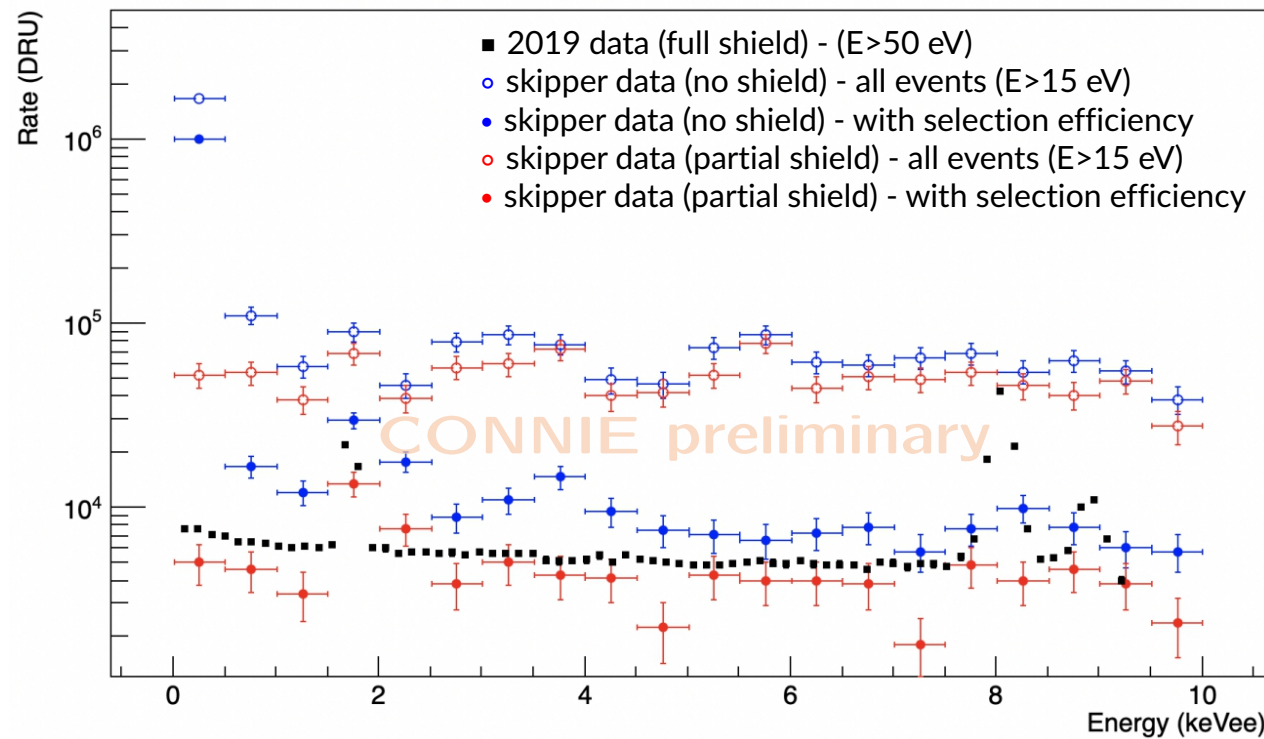


# CONNIE upgrade with skipper-CCDs

Current exposure:

Without shield  $\rightarrow 4.28 \times 10^{-4}$  kg-day

With partial shield ( $\sim 30$  cm inner poly &  $\sim 5$  cm lead)  $\rightarrow 1.69 \times 10^{-3}$  kg-day



Next steps:

Collect more data / Run with full shield / Install more CCDs to increase mass to  $\sim 50$  g

# Summary

- ▶ CCDs are a promising technology for detecting  $CE\nu NS$  at low energies
- ▶ CONNIE has demonstrated to be competitive constraining BSM physics
- ▶ In 2019 run data analysis we achieved better sensitivity due to binning and improved analysis (paper coming soon!)
- ▶ Skipper CCDs allow to improve greatly the low-energy sensitivity
- ▶ The first skipper data at a reactor are encouraging → stable, low noise and DC, rate with partial shield competitive with CONNIE 2019 rate
- ▶ Characterization of skipper CCDs at sea-level background will help prepare for a future larger-mass skipper CCD experiment\*

\*Started discussions for installing skipper CCDs inside the dome of the reactor at Angra (~17 m away from the core)

Thank you!

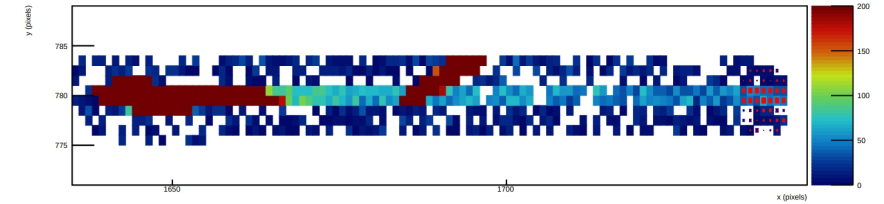
The background features two large, overlapping circles. The left circle is a light red color and contains a white outline of a person's head and shoulders. The right circle is a light blue color and contains a white outline of a person's head and shoulders. A large, semi-transparent white letter 'M' is centered over the intersection of the two circles. The text 'Back up slides' is written in a dark blue, sans-serif font across the middle of the 'M' and the overlapping area.

Back up slides

# LLE and PCC events in low background

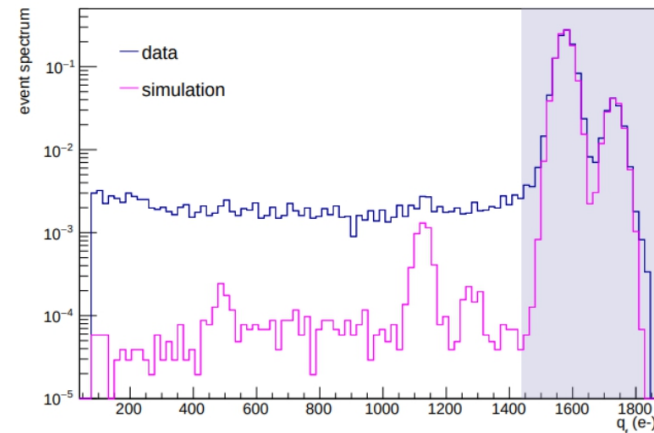
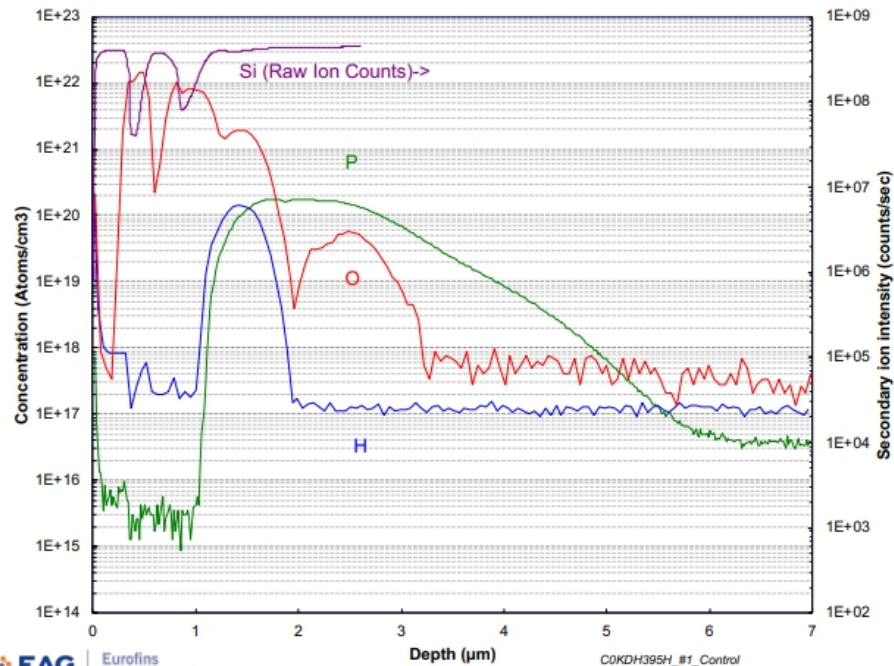
## LLE events

- ▶ Tails of very energetic events that do not follow CTI process
- ▶ Charge deposition in the inactive volume of the sensor (SR)

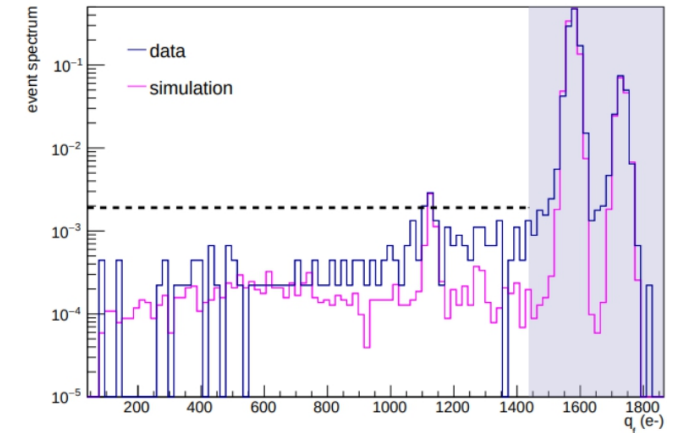


## PCC events

- ▶ CONNIE CCDs have a  $\sim 5 \mu\text{m}$  layer in the back of the sensors where charge partially recombines because of a gradient in the P concentration ( $10^{20} \rightarrow 10^{11} \text{ P atoms/cm}^3$ )



CCD without back treatment



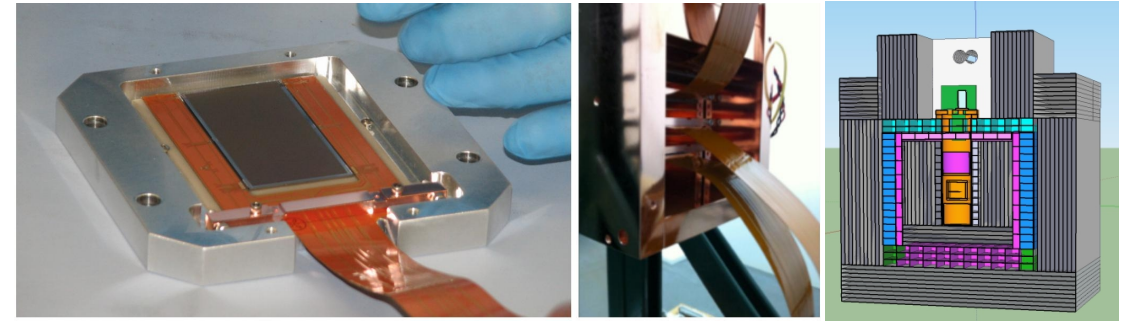
Back-treated CCD



# CONNIE timeline and milestones

## Engineering run

Dec 2014 → 4 CCDs (1 gram each) installed at Angra  
Aug 2015 → Full shield installed



**Engineering run data:** 1 month with RON / 1 month with ROFF → 1x1 binning  
Apr 2016 → Published engineering run data results [**JINST 11 (2016), P07024**]  
▶ Demonstrated remote operations, low noise (<2 e-) and stable bkgd rates

