

# PHOTON-AXION OSCILLATIONS WITH ELMAG



Jonas Tjemsland

PhD candidate

Supervisor: Michael Kachelrieß

Norwegian University of Science and Technology

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# Content

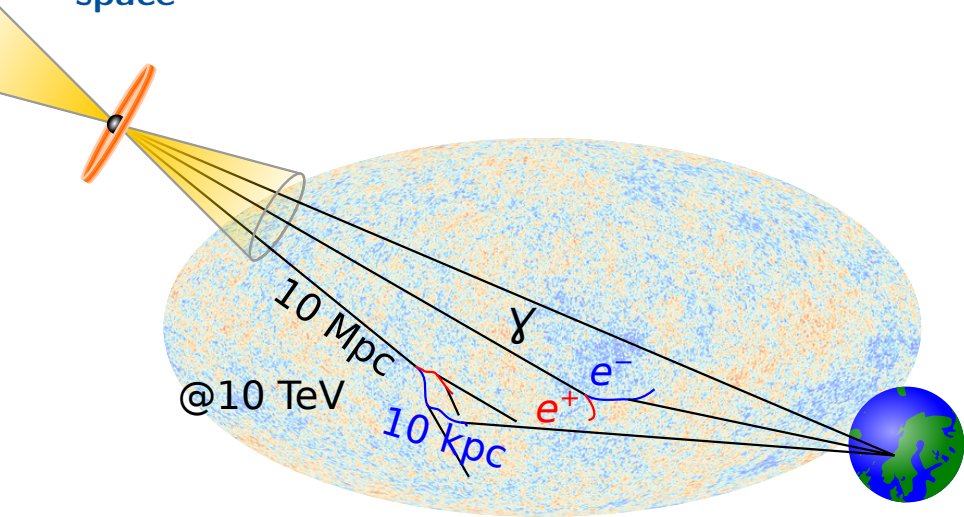
ELMAG and electromagnetic cascades

Axionic dark matter

Axion-photon oscillations in high energy astrophysics

Summary

# Electromagnetic cascades in extragalactic space



## ELMAG [1106.5508, 1909.09210]

*Monte Carlo simulation tool for electromagnetic cascades of high-energy photons and electrons*

- ▶  $\gamma + \gamma_{\text{EBL}} = e^+ + e^-$  (pair production [ $E \lesssim 10^{15}$  eV])
- ▶  $e^\pm + \gamma_{\text{EBL}} = e^\pm + \gamma$  (inverse Compton scattering)
- ▶ Photon-axion oscillations in a magnetic field ( $a \leftrightarrow \gamma$ )

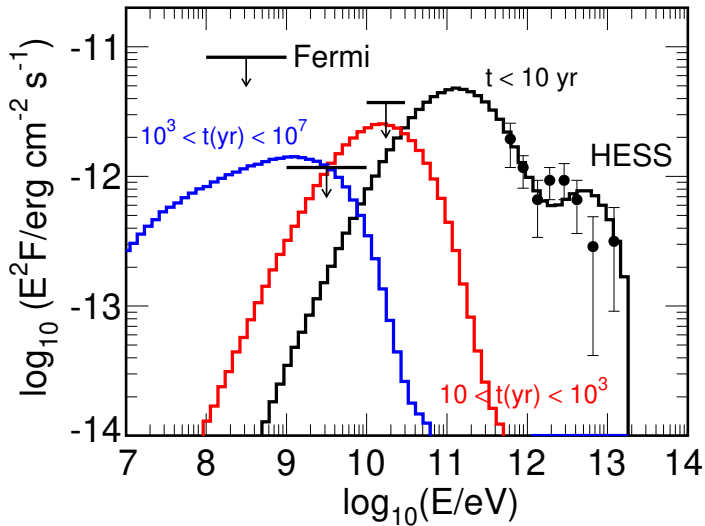
### Features:

- ▶ Deflection in magnetic fields
- ▶ Adiabatic energy losses
- ▶ Synchrotron losses
- ▶ Plasma instabilities
- ▶ Detector properties
- ▶ ...

### Results:

- ▶ Energy spectrum
- ▶ Observation angle
- ▶ Time delay
- ▶ Two-dimensional source images

## Example: Spectrum of blazar 1ES 0029+200



[1106.5508]

# The strong CP problem and axions

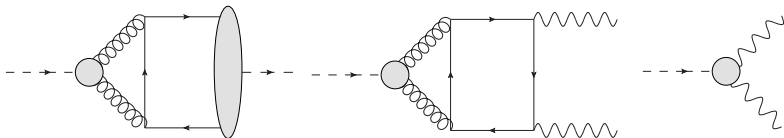
$$\mathcal{L}_{\text{SM}} \supset \frac{\vartheta g^2}{8\pi^2} G \tilde{G}, \quad \vartheta \in [0, 2\pi)$$

- ▶ Measurements of the neutron dipole moment  $\Rightarrow |\vartheta| \lesssim 10^{-10}$
- ▶ Fine tuning: Why not  $\vartheta = \mathcal{O}(1)$ ?!
- ▶ The Peccei-Quinn solution:

(Peccei, Quinn 1977, Wilczek 1978; Weinberg 1978)

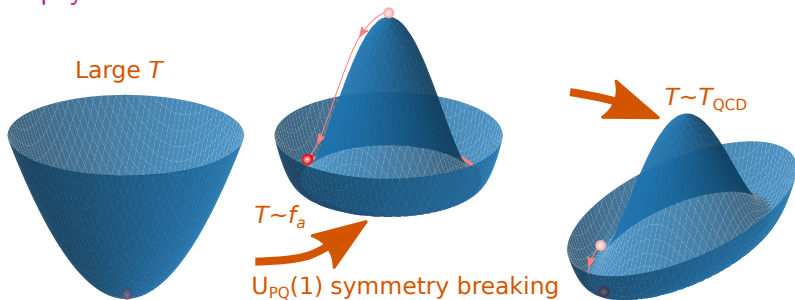
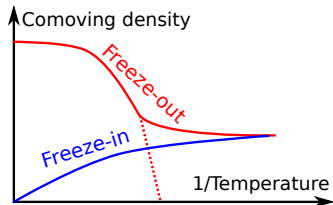
Make  $\vartheta$  dynamically relax to 0 by introducing an U(1) chiral symmetry spontaneously broken at an energy  $f_a$

$$\Rightarrow \text{Axions, } \mathcal{L}_a = \frac{1}{2}(\partial_\mu a)^2 - \frac{g^2}{16\pi^2} \frac{a}{f_a} G \tilde{G}$$

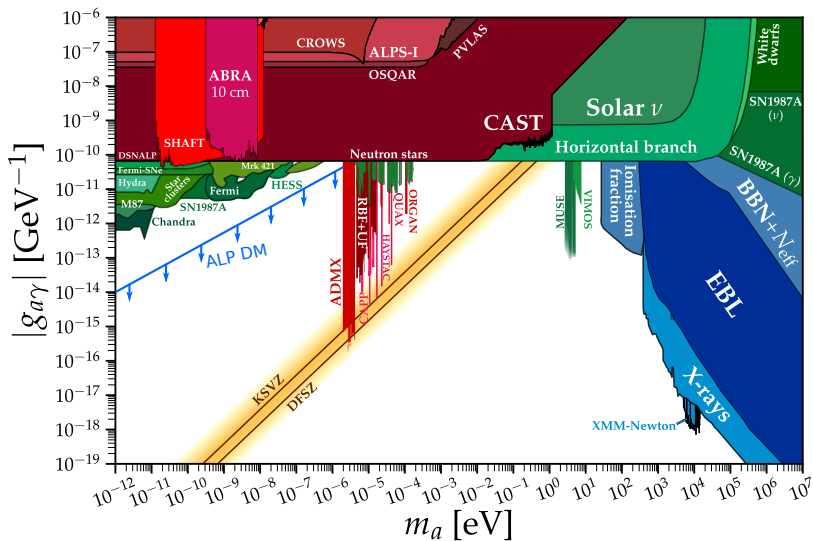


# Axionic dark matter

- ▶ Very light,  $m_a \lesssim \text{eV}$
- ⇒ Thermal production gives **hot dark matter**...
- ▶ **Misalignment mechanism**  
(Preskill, Wise, Wilczek 1983, ++)
- The axion field oscillates coherently and loses energy by **producing physical axions**



# The ALP parameter space

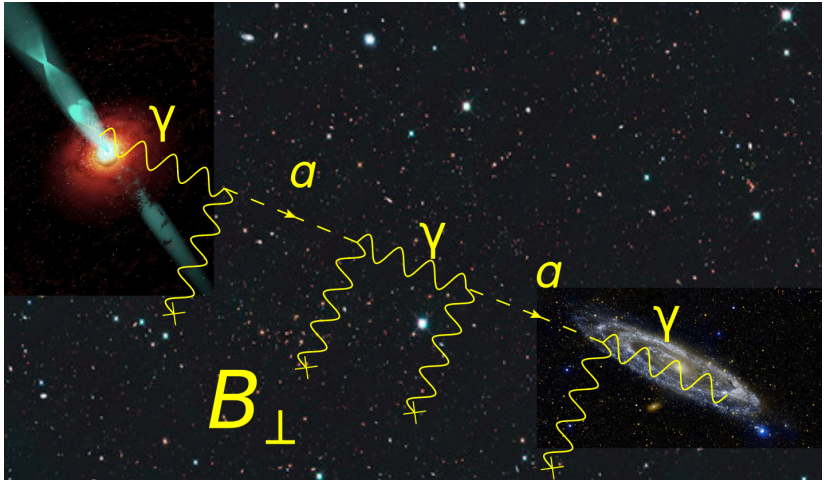


(adapted from [10.5281/zenodo.3932430])



# Photon-axion oscillations in TeV astrophysics

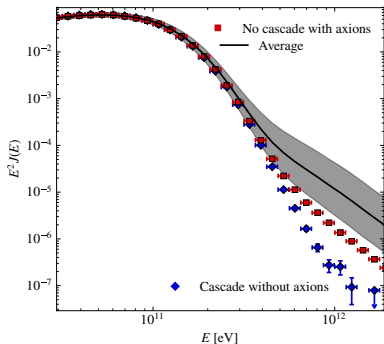
Based on [2111.08303]



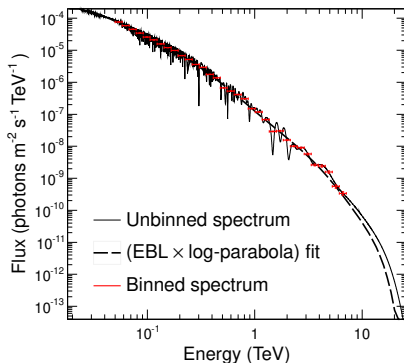
Divergence free Gaussian turbulent field, Kolmogorov spectrum,  
 $B \sim \text{nG}$ ,  $L_{\text{coh}} \sim \text{Mpc}$ ,  $g_{a\gamma} \sim 10^{-20} \text{ eV}^{-1}$

# Observational consequences

## Increased opacity of the Universe



## “Irregularities” in photon spectra



(Adapted from [1205.6428])

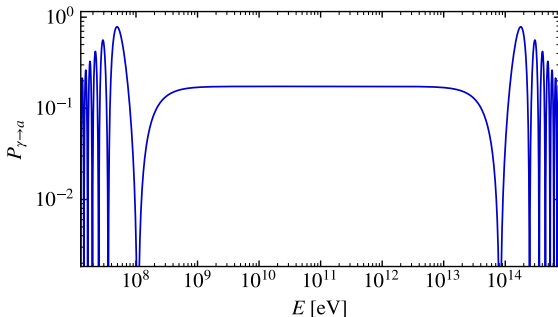
# Photon-axion oscillations

Oscillation due to a mass difference of two mass eigenstates

$$\Rightarrow P_{\gamma \rightarrow a} = |\langle a | \Psi(t) \rangle|^2 = \sin^2(2\vartheta) \sin\left(\frac{L}{2E} (m_1^2 - m_2^2)\right)$$

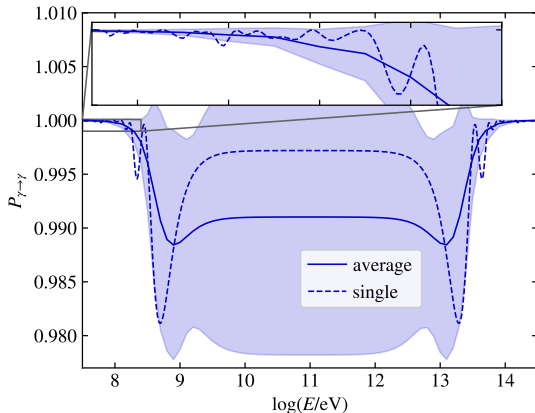
$$\Delta_{\text{osc}}^2 = (\Delta_\gamma - \Delta_a)^2 + 4\Delta_{a\gamma}^2; \quad \Delta_{a\gamma} = \frac{g_{a\gamma} B_\perp}{2}$$

$$\Delta_\gamma = \Delta_\gamma^{\text{CMB}} + \Delta_\gamma^{\text{QED}} + \Delta_\gamma^{\text{pl}}$$



# Energy dependence in a turbulent field

$$\left(-i\frac{\partial}{\partial z} + E + \mathcal{M}\right) \begin{pmatrix} A_{\perp} \\ A_{\parallel} \\ a \end{pmatrix} = 0; \quad \mathcal{M} = \mathcal{M}(B_{\perp}, E)$$



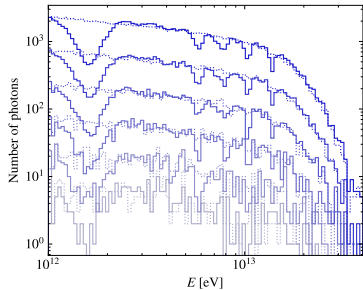
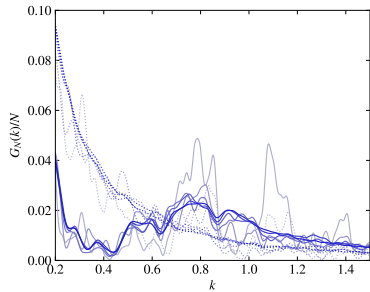
# A direct detection of axion wiggles

- ▶ **Idea:** Use the **energy dependence** of the wiggles as observable

$$G(k) = \left| \int_{\eta_{\min}}^{\eta_{\max}} d\eta q(\eta) e^{i\eta k} \right|^2 \approx \left| \frac{1}{N} \sum_{\text{events}} \exp \{i\eta k\} \right|^2$$

- ▶ Observables:
  - ▶ Peak in  $G(k)$  for  $\eta \sim E$  at “low” energies
  - ▶ Peak in  $G(k)$  for  $\eta \sim E^{-1}$  at “high” energies
  - ▶ No systematic signal otherwise
- ▶ The signal can be used to infer information about the magnetic field

# Example: detecting axion wiggles



**Title:** ELMAG 3.03 [1106.5508, 1909.09210]

**Webpage:** <http://elmag.sourceforge.net>

**Language:** Fortran 90

**Usage:** Monte Carlo simulation of electromagnetic cascades on the extragalactic background light in magnetic fields.

### Photon-axion oscillations with ELMAG

- ▶ Solves the **photon-axion equation of motion** in a Monte Carlo treatment of an electromagnetic cascade
- ▶ Includes cascade photons
- ▶ Can easily consider polarisation effects and any magnetic field configuration
- ▶ Axion wiggles can be detected using the **discrete power spectrum**

Will be made public in a future release of ELMAG.