

Statistical signal processing for the detection of small/rare features with applications in astrophysics

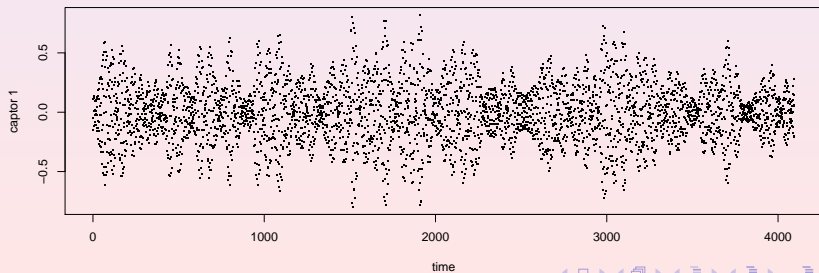
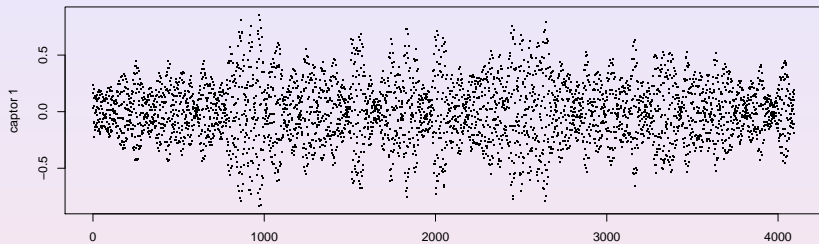
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CERN, PhyStat
9 January 2011

Joint work with Caroline Giacobino, Stefano Foffa, Roberto Terenzi

Gravitational wave detection



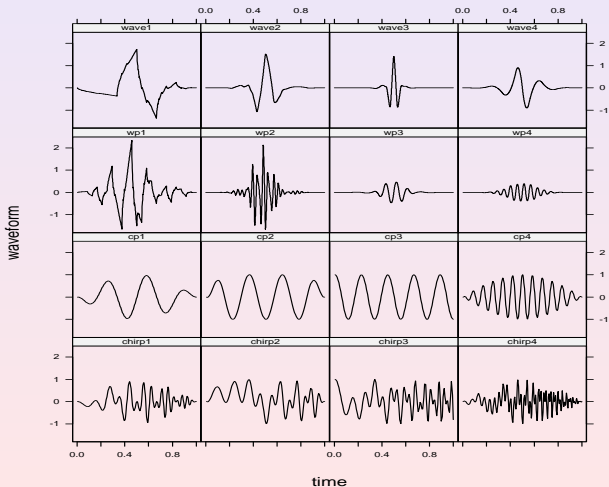
Time series $\{(S_{1,t}, S_{2,t})\}_{t=1,\dots,T}$ for gravitational wave detection:

- ▶ 2 detectors: concomitant (\neq repeated) measurements
- ▶ poor signal f to noise ϵ ratio
- ▶ 5 kHz: 600'000 observations per minute
- ▶ **colored** electronic noise
- ▶ Astrophysics predicts rare arrivals of small wavebursts:

$$f(t) = 0 \quad \text{for most } t.$$

Wavelet representation:

$$f(t) = \sum_i \gamma_i \phi_i(t) + \sum_j^J \sum_{k=1}^{\eta_j} \alpha_{j,k} \psi_{j,k}(t).$$



Sparse wavelet representation:

*With only a few **non-zero** wavelet coefficients $\alpha_{j,k}$,
we can approximate many signals.*

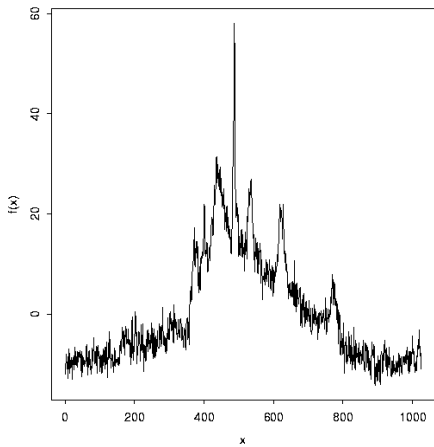
Illustration with an Nuclear Magnetic Resonance noisy signal.

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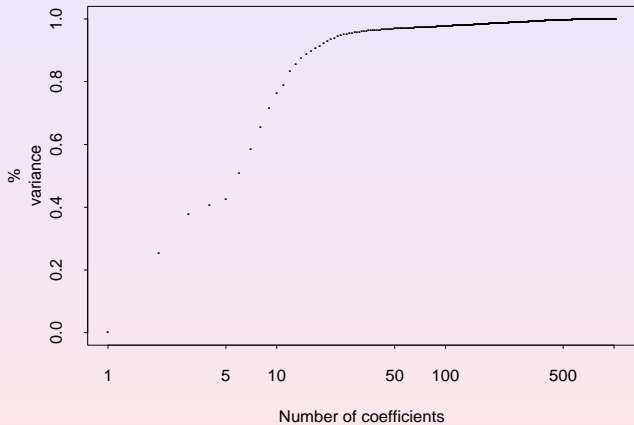
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NMR Spectrum

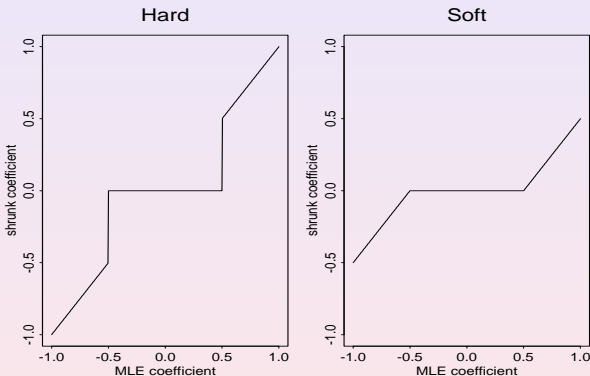


Percentage of variance explained for NMR spectrum (N=1024)



How to enforce sparsity?

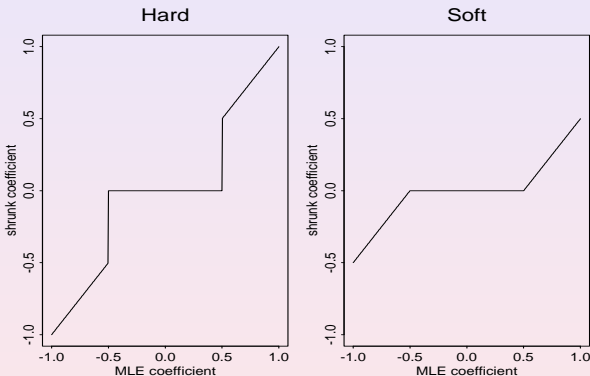
With a non-linear thresholding function with threshold λ :



Anything smaller than λ is treated as noise, and killed.

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1 **captor**: Donoho and Johnstone's 3 easy steps:

1. Analyze

$$\hat{\alpha}^{\text{MLE}} =: \mathbf{Y} = \text{DWT}(\mathbf{S}) \quad (\text{like FFT})$$

2. Threshold

$$\begin{aligned}\hat{\alpha}_{j,k} &= \eta_{\lambda_j}^{(\text{soft})}(Y_{j,k}) \\ &= \text{sign}(Y_{j,k})(|Y_{j,k}| - \lambda_j)_+ \\ &= \left(1 - \frac{\lambda_j}{|Y_{j,k}|}\right)_+ Y_{j,k}\end{aligned}$$

for all levels $j = 1, \dots, J$.

3. Reconstruct $\rightsquigarrow \hat{f}_\lambda$

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Reminiscent of the truncated James-Stein estimator

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~> Generalized James-Stein estimator

$$\hat{\alpha}_{j,k}^{\text{GJS},\lambda_j,\nu_j} = \left(1 - \frac{\lambda_j^{\nu_j}}{\|\mathbf{Y}_{j,k}\|_2^{\nu_j}}\right)_+ \mathbf{Y}_{j,k}, \quad \nu_j > 0$$

is a continuous bridge between soft ($\nu_j = 1$) and hard ($\nu_j \rightarrow \infty$) block-thresholding.

Challenge: selection of λ and ν .

One possibility is to minimize the risk

$$R(\lambda, \nu) = \mathbb{E} \|\hat{f}_{\lambda,\nu} - f\|_2^2$$

or an unbiased estimate.

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Selection of λ_j and ν_j with Stein Unbiased Risk Estimate SURE

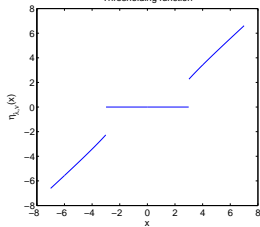
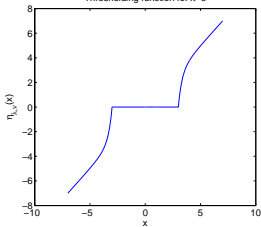
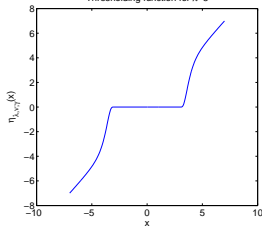
The GJS thresholding function is weakly differentiable in Stein's sense, so we can estimate unbiasedly the risk with

$$\begin{aligned}\hat{R}(\lambda_j, \nu_j) &= \sum_{k=1}^{2^j} \|\mathbf{Y}_{j,k}\|_2^2 \mathbf{1}(\|\mathbf{Y}_{j,k}\|_2 \leq \lambda_j) + \sum_{k=1}^{2^j} \frac{\lambda_j^{2\nu_j} \|\mathbf{Y}_{j,k}\|_2^2}{\|\mathbf{Y}_{j,k}\|_2^{2\nu_j}} \mathbf{1}(\|\mathbf{Y}_{j,k}\|_2 > \lambda_j) \\ &\quad + 2 \sum_{k=1}^{2^j} \left(Q - \frac{\lambda_j^{\nu_j} (Q - \nu_j)}{\|\mathbf{Y}_{j,k}\|_2^{\nu_j}} \right) \mathbf{1}(\|\mathbf{Y}_{j,k}\|_2 > \lambda_j) - 2^j Q. \\ &= \text{RSS}(\lambda_j, \nu_j) + 2 \text{ d.f.}_j\end{aligned}$$

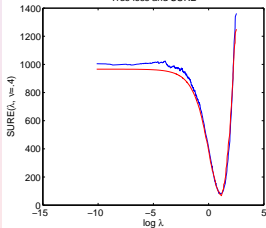
and minimize it over λ_j and ν_j .

Subbotin l_q penalty: $v=4$

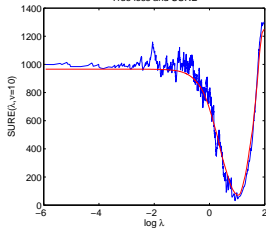
Thresholding function

Generalized James–Stein: $v=10, \gamma=1$ Thresholding function for $\lambda=3$ Smooth James–Stein: $v=10, \gamma=2\log v+1$ Thresholding function for $\lambda=3$ 

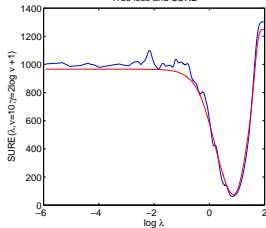
True loss and SURE



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Selection of λ_j by universal rule for Q captors

$$\lambda_{j,N,Q} = \sqrt{2(\log N_j + (Q/2) \log \log N_j - \log \Gamma(Q/2))},$$

e.g., with 2 captors

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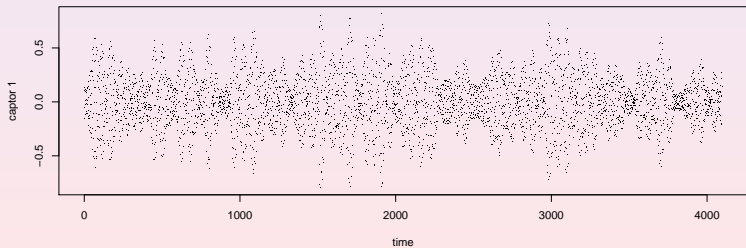
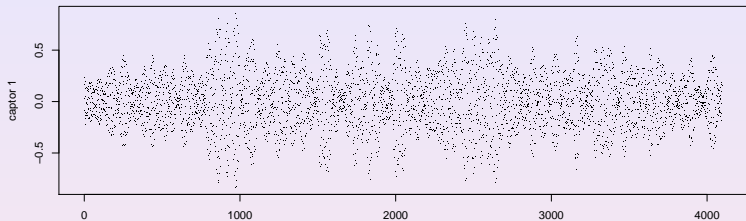
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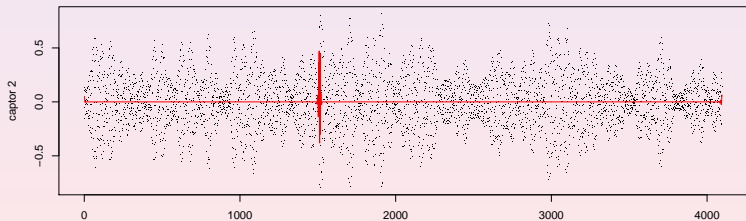
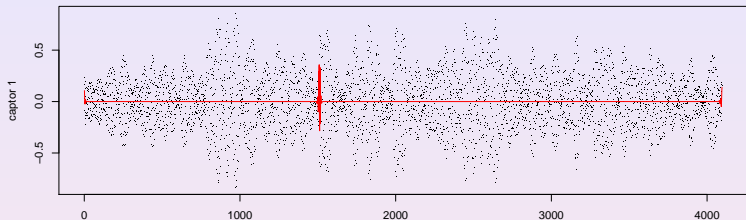
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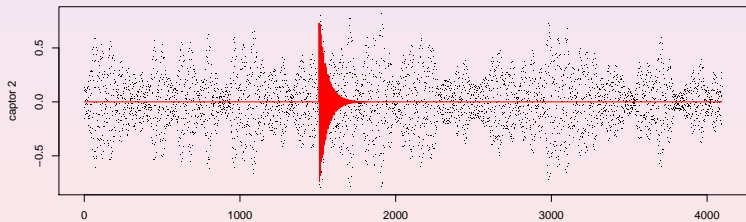
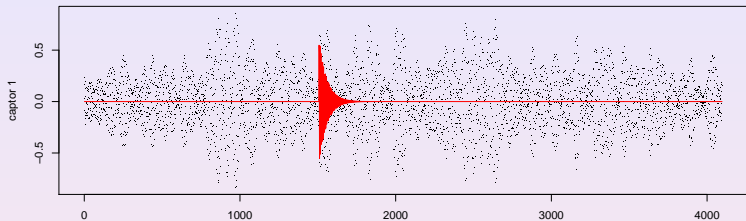
Wavebursts data



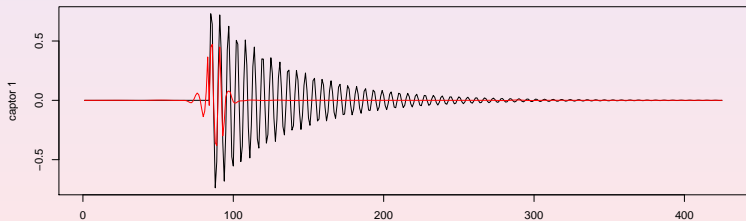
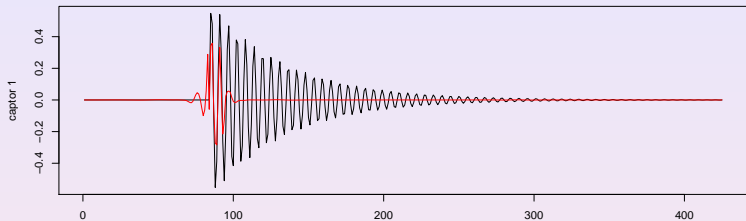
Waveburst detection: SJS with SURE thresholds



Waveburst detection: true “injections”



Zooming on the waveburst



Conclusions

Finding a needle in a haystack is possible.

Future work: compare with existing techniques

- ▶ *Performance of the WaveBurst algorithm on LIGO data*, Klimenko et al. (2004)
- ▶ *Wavelet entropy filter and Crosscorrelation of gravitational wave data*, Terenzi and Sturani (2009)

by looking at Type I, Type II errors and false discovery rates.

Density estimation/New particle or statistical fluctuation (L. Lyons)?

