Perturbative QCD at high densities

Tyler Gorda TU Darmstadt

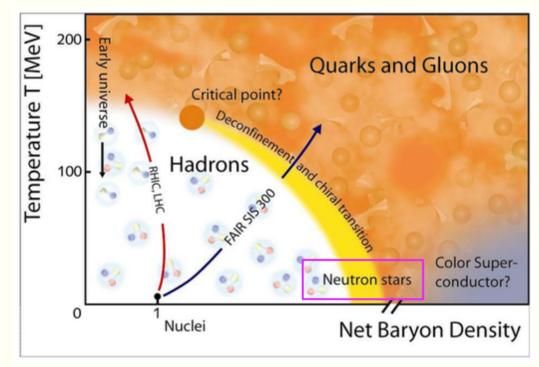
Course on Neutron-Star Physics, IGFAE 17 November 2021



Summary / Overview:

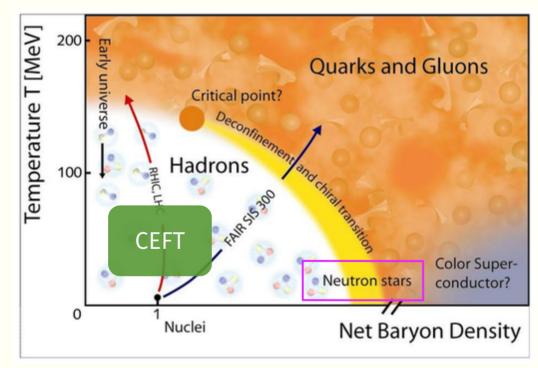
The big picture

NSs probe densities beyond nuclear density, but below pQCD densities



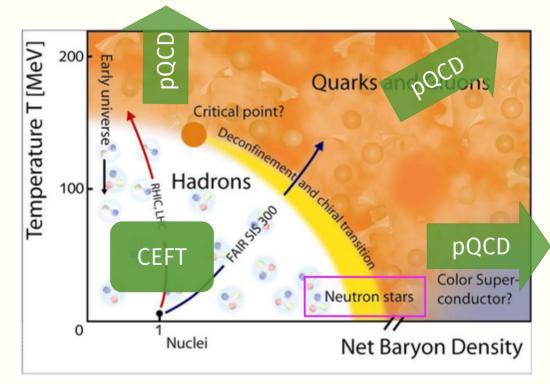
Compressed Baryonic Matter (CBM) experiment

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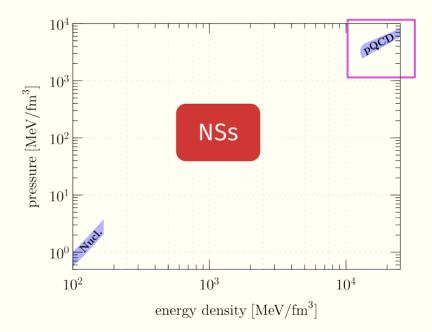
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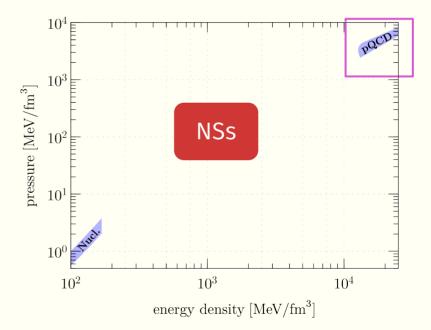
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Want to use the fact that NS matter (EoS) goes to pQCD EoS at high densities as asymptotic limit



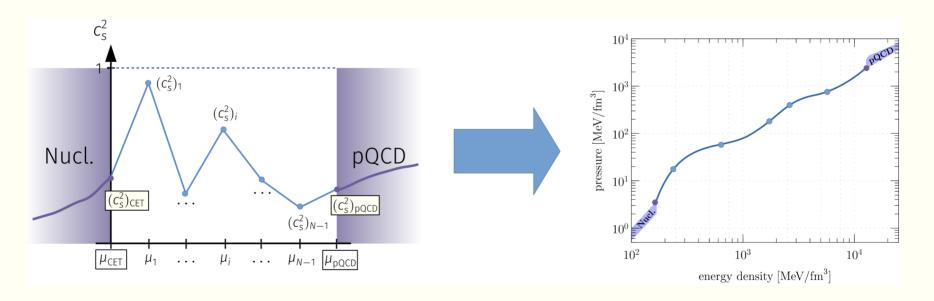
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• Can fill in to lower densities by interpolation



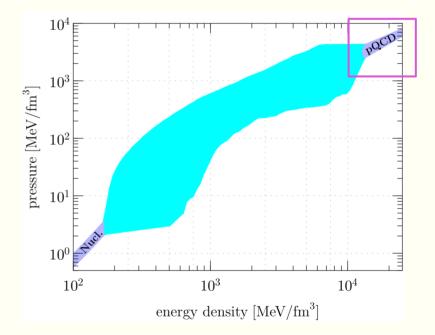
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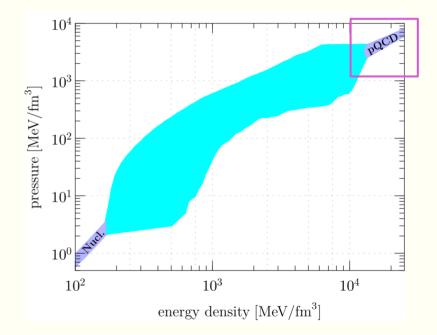
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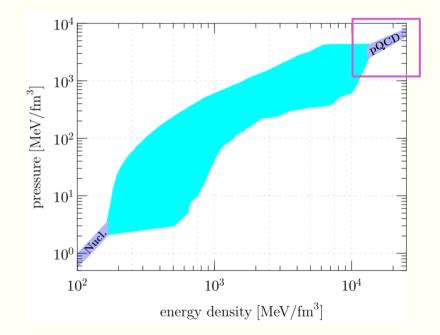
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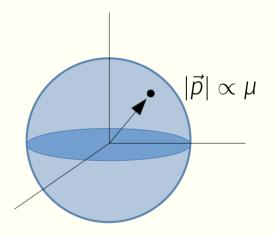
- Can fill in to lower densities by interpolation
- That NS EoS must go to pQCD in (p, ε, n) sets nontrivial constraints when matching Komoltsev, Kurkela, arXiv 2111.05350



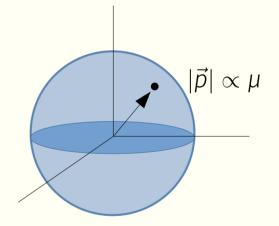
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- Can fill in to lower densities by interpolation
- That NS EoS must go to pQCD in (p, ε, n) sets nontrivial constraints when matching Komoltsev, Kurkela, arXiv 2111.05350
- This is why we are interested in pQCD, and improving pQCD at high density

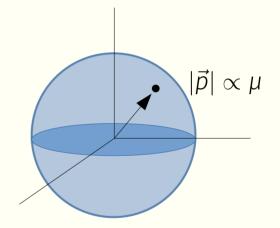




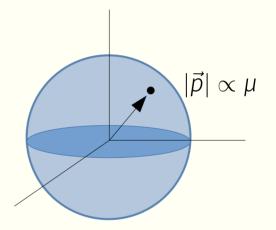
- QM has colored quarks/gluons as DOF
- At high density, $\alpha_s \ll 1$, so quarks/gluons quasiparticles, with quark Fermi sea*



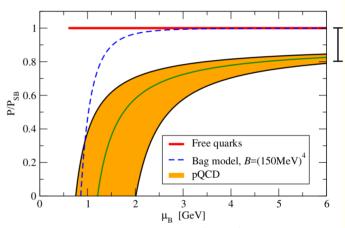
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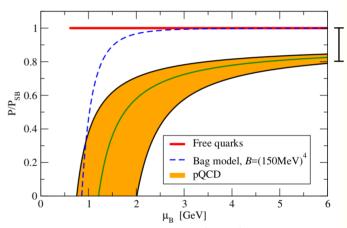
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Fraga+ Astrophys. J. 781 (2014); see also Kurkela+ Phys. Rev. D 81 (2010).

Basic property of cold QM EoS is that it's approximately described by a free electron gas

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So we want to calculate these corrections accurately!

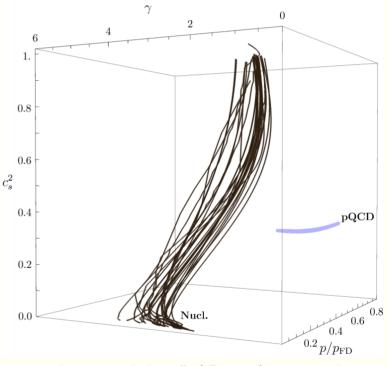
Motivation: what is Quark Matter? Physical properties

• QM has different physical properties than hadronic matter:

Prakash, Astrophys. J. 550 (2001), Gandolfi+ Phys. Rev. C

Fortin+ Phys. Rev. C 94, (2016), Lattimer &

85 (2012) Hadronic Quark C_s^2 increases $\lesssim 1/3$ $\gamma \equiv \frac{d \ln p}{d \ln \epsilon} \approx 2.5 \approx 1$ $p/p_{FD} \approx 0.1 - 0.3 \approx 0.5 - 0.8$



Annala, TG, Kurkela, Nättilä, Vuorinen Nat. Phys. 16 (2020)

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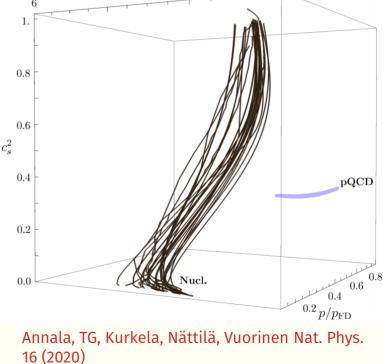
Strategy:

85 (2012)

 $C_{\rm S}^2$

 $\gamma \equiv \frac{d \ln p}{d \ln \varepsilon}$

Identify where EoS changes physical properties from hadronic \rightarrow quark

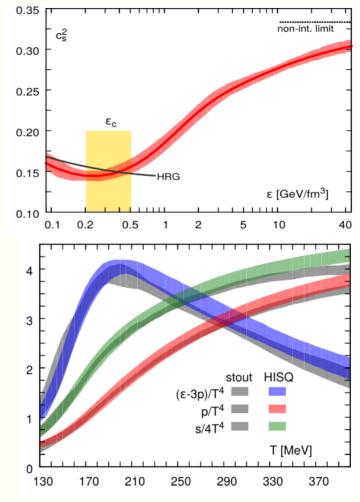


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0

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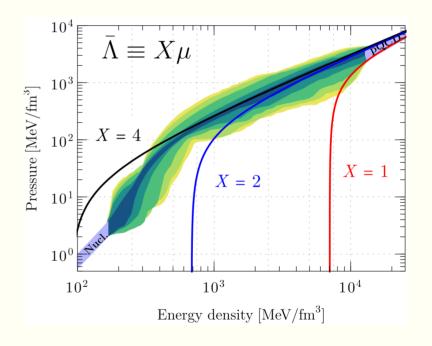
- Similar to looking for change in behavior of lattice results at high *T*.
- Identify change in phase from *change in physical properties* of matter



HotQCD Phys.Rev.D 90 (2014), Borsanyi+ Phys. Lett. B 370 (2014)

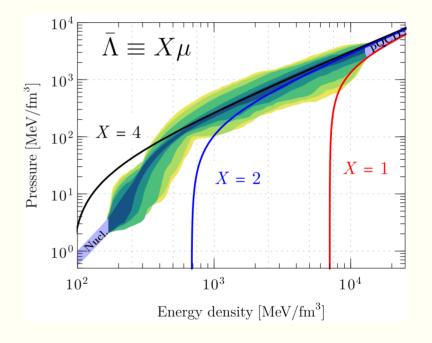
Currently, have *large* renorm.-scale dependence

- Can't use cold QM EoS below $45n_s$; factor of 5-10 higher than is relevant in NSs
- Hope that higher-order pQCD calculations will allow us to fix renorm. scale by, e.g., Principle of minimum sensitivity
 P.M. Stevenson. Nucl. Phys. B 231 (1984)



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- Want to improve cold QM EoS to use down to more relevant densities



Framework for cold QM computations is relativistic thermal QFT.

 Systemmatic framework for calculating corrections in a series expansion in α_s^{*} (*important caveats to come!*)

$$p = \underbrace{p_0}_{} + p_1 \alpha_s + p_2 \alpha_s^2 + \cdots$$

free quark gas

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• Language for this expansion is *Feynman diagrams*

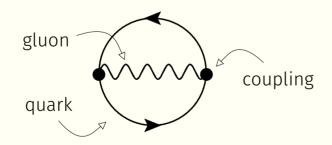
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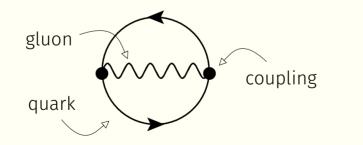
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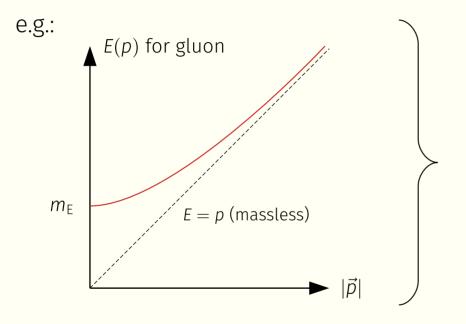
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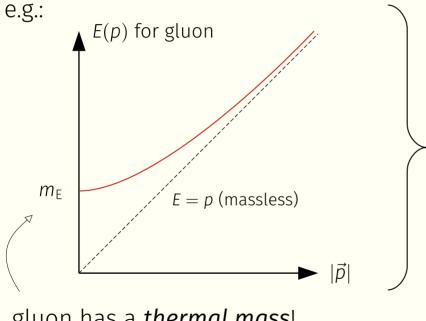


*(no external lines because this is the vacuum with μ >0) *Important caveat* is that TQFT has IR (long-wavelength) differences from what you would expect



$$-E(\vec{p})^2 + \vec{p}^2 + \overbrace{\Pi(E(\vec{p}),\vec{p})}^{\text{"self-energy"}} = 0$$

*describes quantum + statistical corrections to particle propagation *Important caveat* is that TQFT has IR (long-wavelength) differences from what you would expect



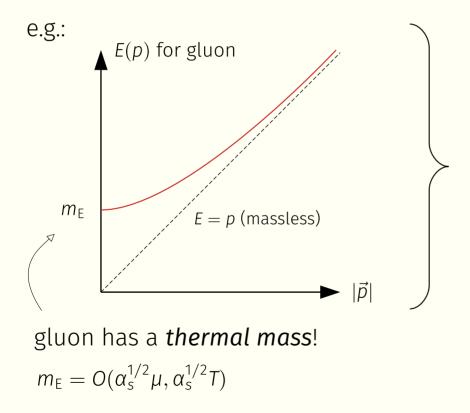
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gluon has a thermal mass!

$$m_{\rm E}=O(\alpha_{\rm s}^{1/2}\mu,\alpha_{\rm s}^{1/2}T)$$

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$$-E(\vec{p})^2 + \vec{p}^2 + \overbrace{\Pi(E(\vec{p}),\vec{p})}^{\text{"self-energy"}} = 0$$

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• Mass screws up naive Feynmandiagram expansions

Loop expansion ≠ coupling expansion

...and their effects

This will modify the naive expectations:

$$p = \underbrace{p_0 + p_1^h \alpha_s + p_2^h \alpha_s^2}_{\text{free quark gas}} \leftarrow \text{scale } |P| \gtrsim \mu$$

$$free \text{ quark gas} + \underbrace{p_2^s \alpha_s^2}_{\text{free soft}} + p_3^s \alpha_s^3 \leftarrow \text{scale } |P| \lesssim m_E$$

$$free \text{ soft} + p_3^m \alpha_s^3 \leftarrow \text{mixed; both scales}$$

$$free \text{ soft} \text{ (screened)}$$

TG+ Phys. Rev. D 104 (2021), Phys. Rev. Lett. 127 (2021); see also TG+ Phys. Rev. Lett. 121 (2018); *O*(α_s²): Freedman & McLerran Phys. Rev. D 16 (1977)

Let's dive in!

Defining equations (Minkowski space):

$$\begin{aligned} \mathcal{L}_{\text{QCD}} &= \bar{\psi}(\mathrm{i}\not{D} - m)\psi - \frac{1}{4}F^{a}_{\mu\nu}F^{a\mu\nu} \\ &= \sum_{f}\overline{\psi}^{i}_{f}\Big(\delta_{ij}\big(\mathrm{i}\gamma^{\mu}\partial_{\mu} - m_{f}\big) - g\gamma^{\mu}A^{a}_{\mu}T^{a}_{ij}\Big)\psi^{j}_{f} - \frac{1}{4}F^{a}_{\mu\nu}F^{a\mu\nu}, \end{aligned}$$

with

interactions

$$F^{a}_{\mu\nu} = \partial_{\mu}A^{a}_{\nu} - \partial_{\nu}A^{a}_{\mu} - gf^{abc}A^{b}_{\mu}A^{c}_{\nu}$$

 $T^a = \lambda^a/2$ generators of SU(3) in the fundamental representation.

$$f^{111} = 1, f^{458} = f^{678} = \frac{\sqrt{3}}{2}.$$

$$f^{147} = f^{165} = f^{246} = f^{257} = f^{345} = f^{376} = \frac{1}{2}$$

$$\lambda^{1} = \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \quad \lambda^{2} = \begin{pmatrix} 0 & -i & 0 \\ i & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \quad \lambda^{3} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 0 \end{pmatrix},$$
$$\lambda^{4} = \begin{pmatrix} 0 & 0 & 1 \\ 0 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix}, \quad \lambda^{5} = \begin{pmatrix} 0 & 0 & -i \\ 0 & 0 & 0 \\ i & 0 & 0 \end{pmatrix},$$
$$\lambda^{6} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}, \quad \lambda^{7} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & -i \\ 0 & i & 0 \end{pmatrix}, \quad \lambda^{8} = \frac{1}{\sqrt{3}} \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -2 \end{pmatrix}.$$
$$\mathbf{13} / \mathbf{33}$$

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Usually, the algebraic identities below are more important than the actual matrices: $(T_F = \frac{1}{2}, C_A = 8, C_F = \frac{4}{3})$

$$[T^a, T^b] = i f^{abc} T^c \qquad \operatorname{tr}(T^a T^b) = T_F \delta^{ab}, \qquad (T^a T^a)_{ij} = C_f \delta_{ij}, \qquad f^{acd} f^{bcd} = C_A \delta^{ab},$$

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u - \partial_
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u$

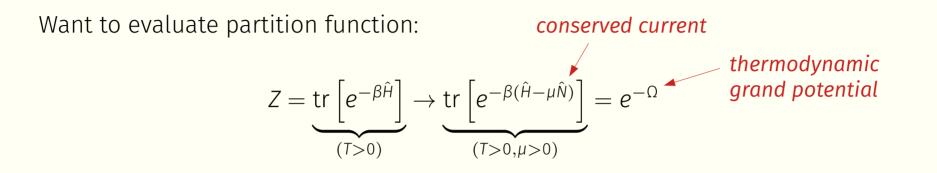
Note that only u,d,s active in dense matter!

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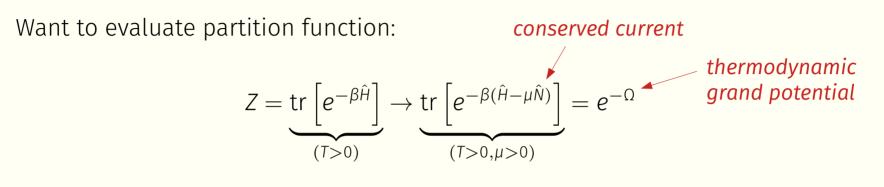
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Thermodynamics of relativistic QFTs: partition function

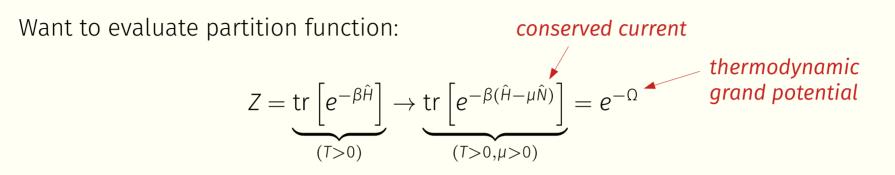


Thermodynamics of relativistic QFTs: partition function



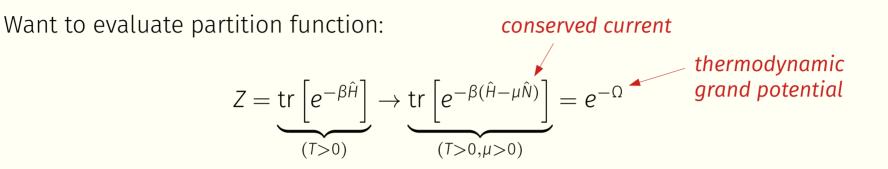
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- Like in normal QFT, simplest to construct a *path-integral* representation of the partition function by dividing up the "time" interval into equal pieces:

$$e^{-\beta(\hat{H}-\mu\hat{N})} = \underbrace{e^{-\Delta\tau(\hat{H}-\mu\hat{N})}e^{-\Delta\tau(\hat{H}-\mu\hat{N})}\cdots e^{-\Delta\tau(\hat{H}-\mu\hat{N})}}_{N \text{ equal pieces}}, \quad \Delta\tau \equiv \frac{\beta}{N}$$

Path integral for the partition function: Grassman variables

First need a quick summary of how to deal with fermionic (Grassman) variables. ψ is an anticommuting variable. We define it to behave the following way:

Anticommuting:
$$\psi_1\psi_2 = -\psi_2\psi_1$$
 $(\implies \psi^2 = 0)$

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Differentiation:
$$\frac{d}{d\psi}(c) = 0$$
, $\frac{d}{d\psi}\psi = 1$; linear
(Definite) Integration: $\int d\psi = 0$, $\int d\psi \psi = 1$; linear

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(Essentially, all operations defined algebraically)

 $|arphi
angle \equiv e^{\pm arphi \hat{a}^{\dagger}}|0
angle$ so that

$$\begin{split} |\varphi\rangle &\equiv e^{\pm \varphi \hat{a}^{\dagger}} |0\rangle \quad \text{ so that } \quad \hat{a} |\phi\rangle &= [a, e^{\phi \hat{a}^{\dagger}}] |0\rangle \\ &= \sum_{n=1}^{\infty} \frac{\phi^n}{n!} n(\hat{a}^{\dagger})^{n-1} |0\rangle \\ &= \phi e^{\phi \hat{a}^{\dagger}} |0\rangle \\ &= \phi |\phi\rangle \end{split} \qquad bosons$$

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angle$$
 so that $\hat{a}|arphi
angle=arphi|arphi
angle$ (both cases)

$$|\varphi\rangle \equiv e^{\pm \varphi \hat{a}^{\dagger}}|0\rangle$$
 so that $\hat{a}|\varphi\rangle = \varphi|\varphi\rangle$ (both cases)

Need the following relations:

Completeness:
$$\mathrm{Id} = \int d(\varphi^{\dagger}, \varphi) e^{-\varphi^{\dagger}\varphi} |\varphi\rangle \langle \varphi|$$

Overlap:
$$\langle \varphi' | \varphi
angle = \exp \left(\varphi'^{\dagger} \varphi
ight)$$

Measure:
$$d(\varphi^{\dagger}, \varphi) = \frac{d\varphi^{\dagger}d\varphi}{\pi^{(1+\zeta)/2}}$$

$$Z = \operatorname{tr}\left[e^{-\beta(\hat{H}-\mu\hat{N})}\right] = \sum_{n>0} \langle n|e^{-\beta(\hat{H}-\mu\hat{N})}|n\rangle$$
$$= \int d(\varphi^{\dagger},\varphi)e^{-\varphi^{\dagger}\varphi} \sum_{n>0} \underbrace{\langle n|\varphi\rangle}_{n>0} \langle \varphi|e^{-\beta(\hat{H}-\mu\hat{N})}|n\rangle$$

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move to end; exchanges Grassman variables!

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$$= \int d(\varphi^{\dagger}, \varphi) e^{-\varphi^{\dagger}\varphi} \langle \pm \varphi | e^{-\beta(\hat{H} - \mu\hat{N})} | \varphi \rangle$$

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$$= \int d(\varphi^{\dagger}, \varphi) e^{-\varphi^{\dagger}\varphi} \sum_{n>0} \underbrace{\langle n | \varphi \rangle}_{n>0} \langle \varphi | e^{-\beta(\hat{H} - \mu\hat{N})} | n \rangle$$

move to end; exchanges Grassman variables!

$$= \int d(\varphi^{\dagger}, \varphi) e^{-\varphi^{\dagger}\varphi} \sum_{n>0} \langle \pm \varphi | e^{-\beta(\hat{H} - \mu\hat{N})} | n \rangle \langle n | \varphi \rangle$$
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Bosons return to same state; fermions to negative the state!

Now break up into little pieces, inserting identities along the way (following Laine and Vuorinen)

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Now deal with each term:

$$e^{-\varphi_{i}^{\dagger}\varphi_{i}}\langle\varphi_{i}|e^{-\Delta\tau(\hat{H}[\hat{a}^{\dagger},\hat{a}]-\mu\hat{N}[\hat{a}^{\dagger},\hat{a}])}|\varphi_{i-1}\rangle\approx e^{-\varphi_{i}^{\dagger}\varphi_{i}}\langle\varphi_{i}|\varphi_{i-1}\rangle\,e^{-\Delta\tau(H[\varphi_{i}^{\dagger},\varphi_{i-1}]-\mu N[\varphi_{i}^{\dagger},\varphi_{i-1}])}$$

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Taking the limit of a large *N*, gives:

$$Z = \int d(\varphi^{\dagger}, \varphi) e^{-\varphi^{\dagger}\varphi} \langle \pm \varphi | e^{-\beta(\hat{H} - \mu\hat{N})} | \varphi \rangle$$

=
$$\int_{\substack{\varphi^{\dagger}(\beta) = \pm \varphi^{\dagger}(0) \\ \varphi(\beta) = \pm \varphi(0)}} \mathcal{D}\varphi^{\dagger}(\tau) \mathcal{D}\varphi(\tau) \exp\left\{-\int_{0}^{\beta} d\tau \left[\varphi^{\dagger}(\tau) \frac{d\varphi(\tau)}{d\tau} + H[\varphi^{\dagger}(\tau), \varphi(\tau)] - \mu N[\varphi^{\dagger}(\tau), \varphi(\tau)]\right]\right\}$$

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For usual Hamiltonians, Legendre transformation gives *Euclidean Lagrangian* (now go back to fields as well):

$$Z = \int_{\substack{\varphi^{\dagger}(\beta,\vec{x}) = \pm \varphi^{\dagger}(0,\vec{x})\\\varphi(\beta,\vec{x}) = \pm \varphi(0,\vec{x})}} \mathcal{D}\varphi \exp\left\{-\int_{0}^{\beta} d\tau \int d^{3}x \left[\mathcal{L}_{E} - \mu \mathcal{N}\right]\right\}$$

Path integral: Summary

$$Z = \int_{\substack{\varphi^{\dagger}(\beta,\vec{x}) = \pm \varphi^{\dagger}(0,\vec{x})\\\varphi(\beta,\vec{x}) = \pm \varphi(0,\vec{x})}} \mathcal{D}\varphi \exp\left\{-\int_{0}^{\beta} d\tau \int d^{3}x \left[\mathcal{L}_{E} - \mu \mathcal{N}\right]\right\}$$

- Compact "time" integral
- Bosons periodic in imaginary time (energies $\omega_n = 2\pi nT$)
- Fermions anti-periodic in imaginary time [energies $\omega_n = 2\pi(n + \frac{1}{2})T$]
- Path integral with Euclidean Lagrangian $(t \rightarrow -i\tau)$

$$\mathcal{L}_{\text{QCD}}^{\text{E}} = \sum_{f} \overline{\psi}_{f}^{i} \Big(\delta_{ij} \big(\gamma_{\mu}^{\text{E}} \partial_{\mu} + m_{f} \big) - ig \gamma_{\mu}^{\text{E}} A_{\mu}^{a} T_{ij}^{a} \Big) \psi_{f}^{j} + \frac{1}{4} F_{\mu\nu}^{a} F^{a\mu\nu},$$

High density

$$Z = \int_{\substack{\varphi^{\dagger}(\beta,\vec{x}) = \pm \varphi^{\dagger}(0,\vec{x})\\\varphi(\beta,\vec{x}) = \pm \varphi(0,\vec{x})}} \mathcal{D}\varphi \exp\left\{-\int_{0}^{\beta} d\tau \int d^{3}x \left[\mathcal{L}_{E} - \mu \mathcal{N}\right]\right\}$$

• As *T*→0, "time" interval becomes infinite again (periodic/antiperiodic doesn't matter)

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- Chemical potential is like imaginary A^o field in Euclidean Lagrangian

$$\mathcal{L}_{E} \ni \left(\overline{\psi}\partial_{\mu}\psi\right) - \left(\overline{\psi}\gamma_{E}^{0}\psi\right)igA^{0} \quad \longleftrightarrow \quad \left(\overline{\psi}\partial_{\mu}\psi\right) - \left(\overline{\psi}\gamma_{E}^{0}\psi\right)\mu$$
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$$\partial_{0} - igA^{0} \quad \longleftrightarrow \quad \partial_{0} - \mu$$
$$i\omega_{n} \mapsto i\omega_{n} - \mu = i(\omega_{n} + i\mu) \text{ imaginary shift to the frequency!}$$

Perturbation theory

$$e^{-\Omega} = \int \mathcal{D}\varphi e^{-S_0[\varphi] - S_l[\varphi]} \simeq \int \mathcal{D}\varphi e^{-S_0[\varphi]} \left(1 - S_l + \frac{1}{2}S_l^2 + \cdots\right)$$

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$$= \Omega_0 - \ln \left[\left\langle 1 - S_I + \frac{1}{2}S_I^2 + \cdots\right\rangle_0\right] \checkmark \qquad \left\langle O \right\rangle_0 = \frac{\int \mathcal{D}\varphi O e^{-S_0[\varphi]}}{\int \mathcal{D}\varphi e^{-S_0[\varphi]}}$$

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$$= \Omega_0 + \langle S_l \rangle_0 - \frac{1}{2} \underbrace{\left[\langle S_l^2 \rangle_0 - \langle S_l \rangle_0^2\right]}_{\text{connected}} + \cdots$$

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connected
$$\Omega = \Omega_0 + \langle S_l - \frac{1}{2}S_l^2 + \cdots \rangle_{0,c}$$
connected corrections!

Q: How do we compute these connected corrections?

A: Introduce a source and differentiate:

Write $S_0[\varphi] = \int_{x,y} \frac{1}{2} \varphi_x A_{x,y} \varphi_y$,

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Then $\langle \varphi_1 \varphi_2 \cdots \varphi_N \rangle_{0,c} = \frac{1}{Z[0]} \frac{\delta}{\delta J_1} \frac{\delta}{\delta J_2} \cdots \frac{\delta}{\delta J_N} Z[J]\Big|_{J=0}$

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 $= \frac{\delta}{\delta J_1} \frac{\delta}{\delta J_2} \cdots \frac{\delta}{\delta J_N} \left(1 + \int_{a,b} \frac{1}{2} J_a A_{a,b}^{-1} J_b + \frac{1}{2} \int_{a,b} \frac{1}{2} J_a A_{a,b}^{-1} J_b \int_{c,d} \frac{1}{2} J_c A_{c,d}^{-1} J_d + \cdots \right) \Big|_{J=0}$

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An example, from ϕ^4 theory:

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These are *Feynman Diagrams*. This fundamental correlation is called the *propagator*

$$\langle \phi_x \phi_y \rangle_0 = \int_P e^{-P \cdot (X-Y)} \frac{1}{P^2 + m^2}$$

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Rearranging prev. expression for the one-loop gluon self energy at high density:

$$\langle \phi \phi \rangle_c - \langle \phi \phi \rangle_0 = \langle \phi \phi (e^{-S_i} - 1) \rangle_{0,c} = - \langle \phi \phi \rangle_0 \Pi \langle \phi \phi \rangle_c$$

Rearranging prev. expression for the one-loop gluon self energy at high density:

 $\langle AA \rangle_0 \Pi(P) \langle AA \rangle_0 = \langle A(S_I - \frac{1}{2}S_I^2)A \rangle_{0,c}$

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$$\begin{aligned} \langle AA \rangle_0 \Pi(P) \langle AA \rangle_0 &= \langle A(S_I - \frac{1}{2}S_I^2)A \rangle_{0,c} &= \langle A(-\frac{1}{2}S_I^2)A \rangle_{0,c} & \text{(need fermion}\\ \text{loop)} \\ &= -\frac{(-\mathrm{i}g)^2}{2} \mathrm{tr}(T^a T^b) \langle A(\bar{\psi} A \psi)(\bar{\psi} A \psi)A \rangle_{0,c} \\ &= \langle AA \rangle_0 \left[g^2 T_f \delta^{ab} \langle (\bar{\psi} \gamma^\mu \psi)(\bar{\psi} \gamma^\nu \psi) \rangle_{0,c} \right] \langle AA \rangle_0 \end{aligned}$$

$$\implies \qquad \Pi(P) = g^2 T_f \delta^{ab} \langle (\bar{\psi} \gamma^{\mu} \psi) (\bar{\psi} \gamma^{\nu} \psi) \rangle_{0,c}$$

Rearranging prev. expression for the one-loop gluon self energy at high density:

$$\begin{aligned} \mathsf{AA}\rangle_0 \Pi(P)\langle \mathsf{AA}\rangle_0 &= \langle \mathsf{A}(S_I - \frac{1}{2}S_I^2)\mathsf{A}\rangle_{0,c} \quad = \langle \mathsf{A}(-\frac{1}{2}S_I^2)\mathsf{A}\rangle_{0,c} & \text{(need fermion}\\ \text{loop)} \\ &= -\frac{(-\mathrm{i}g)^2}{2} \mathrm{tr}(T^a T^b) \langle \mathsf{A}(\bar{\psi} \not\!\!A \psi)(\bar{\psi} \not\!\!A \psi)\mathsf{A}\rangle_{0,c} \\ &= \langle \mathsf{AA}\rangle_0 \left[g^2 T_f \delta^{ab} \langle (\bar{\psi} \gamma^\mu \psi)(\bar{\psi} \gamma^\nu \psi)\rangle_{0,c} \right] \langle \mathsf{AA}\rangle_0 \end{aligned}$$

(only connected contraction; reordered the fermions)

Rearranging prev. expression for the one-loop gluon self energy at high density:

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$$\Pi(P) = -g^2 T_f \delta^{ab} \int_Q \operatorname{tr} \left\{ \left[\frac{\mathrm{i} \mathcal{Q}}{Q^2} \right] \gamma^{\mu} \left[\frac{\mathrm{i} (\mathcal{P} + \mathcal{Q})}{(P + Q)^2} \right] \gamma^{\nu} \right\} = g^2 T_f \delta^{ab} \int_Q \frac{\operatorname{tr} \left\{ \mathcal{Q} \gamma^{\mu} (\mathcal{P} + \mathcal{Q}) \gamma^{\nu} \right\}}{Q^2 (P + Q)^2}$$

(Remember $Q^0 \rightarrow Q^0 + i\mu$)

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$$\Pi(P) = g^2 T_f \delta^{ab} \int_Q \frac{\operatorname{tr} \left\{ \not Q \gamma^{\mu} (\not P + \not Q) \gamma^{\nu} \right\}}{Q^2 (P+Q)^2}$$

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When *P*«*Q*, then we are looking at the UV of this integral

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Doing residues cuts of the integral at μ :

 $\Pi(P) \simeq g^2 \mu^2$ for small P

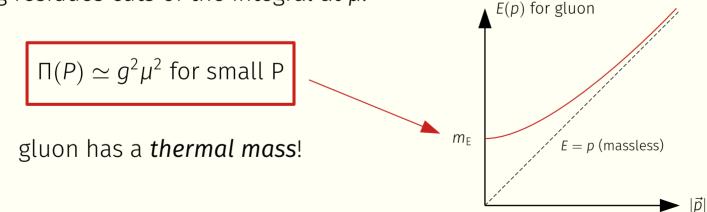
$$\Pi(P) = g^2 T_f \delta^{ab} \int_Q \frac{\operatorname{tr} \left\{ \not Q \gamma^{\mu} (\not P + \not Q) \gamma^{\nu} \right\}}{Q^2 (P+Q)^2}$$

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$$\Pi(P) = g^2 T_f \delta^{ab} \int_Q \frac{\operatorname{tr} \left\{ \mathcal{Q} \gamma^{\mu} (\mathcal{P} + \mathcal{Q}) \gamma^{\nu} \right\}}{Q^2 (P + Q)^2} \longrightarrow g^2 \int^{\mu} dQ Q^3 \frac{Q}{Q^2} \frac{Q}{Q^2} = g^2 \int^{\mu} dQ Q \quad \text{UV dominated}$$

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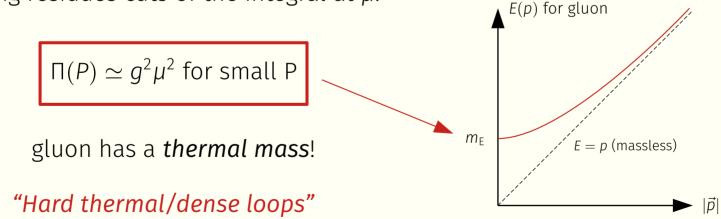
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Doing residues cuts of the integral at μ :



Braaten & Pisarski, Phys. Rev. D 42 (1990), 46 (1992); in cold QM context: Manuel, Phys. Rev. D 53 (1996)

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Nontrivial dependence on $P^0/|\vec{p}|$ in the HTL result (so more than just a thermal mass):

$$\Pi^{\mu\nu}_{ab}(P) = m_{\rm E}^2 \int_{\hat{V}} \left(\delta^{\mu 0} \delta^{\nu 0} - \frac{{\rm i} P^0}{P \cdot V} V^{\mu} V^{\nu} \right)$$

$$m_{\rm E} \equiv \sum_{f} \frac{g^2 \mu_f^2}{2\pi^2}, \quad V^{\mu} \equiv (-i, \hat{v}), \quad \hat{v} \in S^2 \text{ (unit vector in } \mathbb{R}^3), \quad \int_{\hat{v}} \text{normalized to 1}$$

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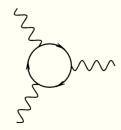
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Similar HTL contributions for *N*-point gluon functions:



Corrections to the EoS from different kinematic regions

Current state-of-the-art: contributions from different kinematic regions

$$p = \underbrace{p_{0} + p_{1}^{h}\alpha_{s} + p_{2}^{h}\alpha_{s}^{2}}_{free \ quark \ gas} \leftarrow scale \ |P| \gtrsim \mu$$

$$free \ quark \ gas \qquad + p_{2}^{s}\alpha_{s}^{2} + p_{3}^{s}\alpha_{s}^{3} \leftarrow scale \ |P| \leq m_{E}$$

$$free \ soft \ + p_{3}^{m}\alpha_{s}^{3} \leftarrow mixed; \ both \ scales$$

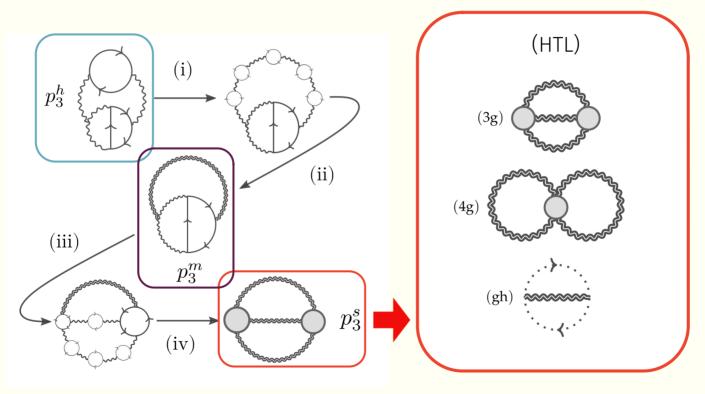
$$free \ soft \ (screened)$$

TG+ Phys. Rev. D 104 (2021), Phys. Rev. Lett. 127 (2021); see also TG+ Phys. Rev. Lett. 121 (2018); *O*(α_s²): Freedman & McLerran Phys. Rev. D 16 (1977)

Corrections to the EoS from different kinematic regions

Current state-of-the-art: have now computed N³LO contributions from *HTL effective theory*

TG, Kurkela, Paatelainen, Säppi, Vuorinen, Phys. Rev. Lett. 127 (2021), Phys. Rev. D 104 (2021)

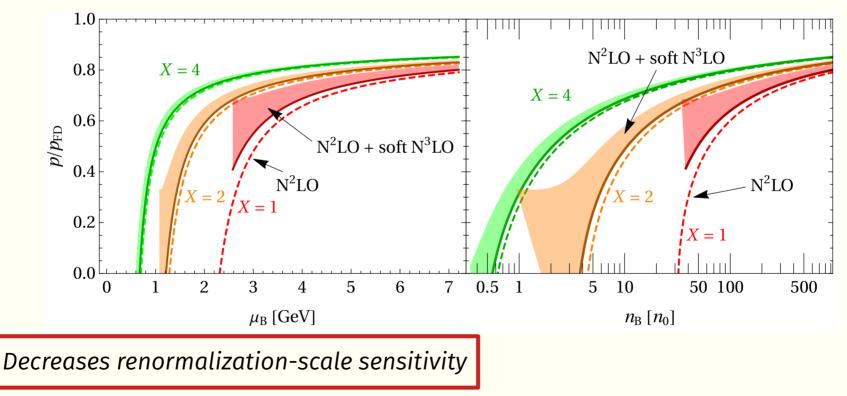


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- TQFT at high density is systematically improvable framework for calculating corrections to thermodynamic properties
- Rich EFT structure, involving multiple scales
- Current state-of-the-art for cold QM calculations are N3LO, and are ongoing, using approach of kinematic regions
- Cold QM EoS restricts NS-matter EoS at lower densities

Two books I referenced, and one I recommend for general QFT:

- A. Altland, B. Simons. Condensed matter field theory. Cambridge, Univ. Press (2006).
- M. Laine, A. Vuorinen. Basics of thermal field theory. Lect. Notes Phys. 925 (2016).
- M. Schwartz. Quantum field theory and the standard model. Cambridge, Univ. Press (2013).