

The Surface Array Enhancement of the IceCube Neutrino Observatory

Roxanne Turcotte and Dennis Soldin on behalf of the IceCube Collaboration
TeVPA 2022

picture credit: {Y.Makino}, IceCube/NSF

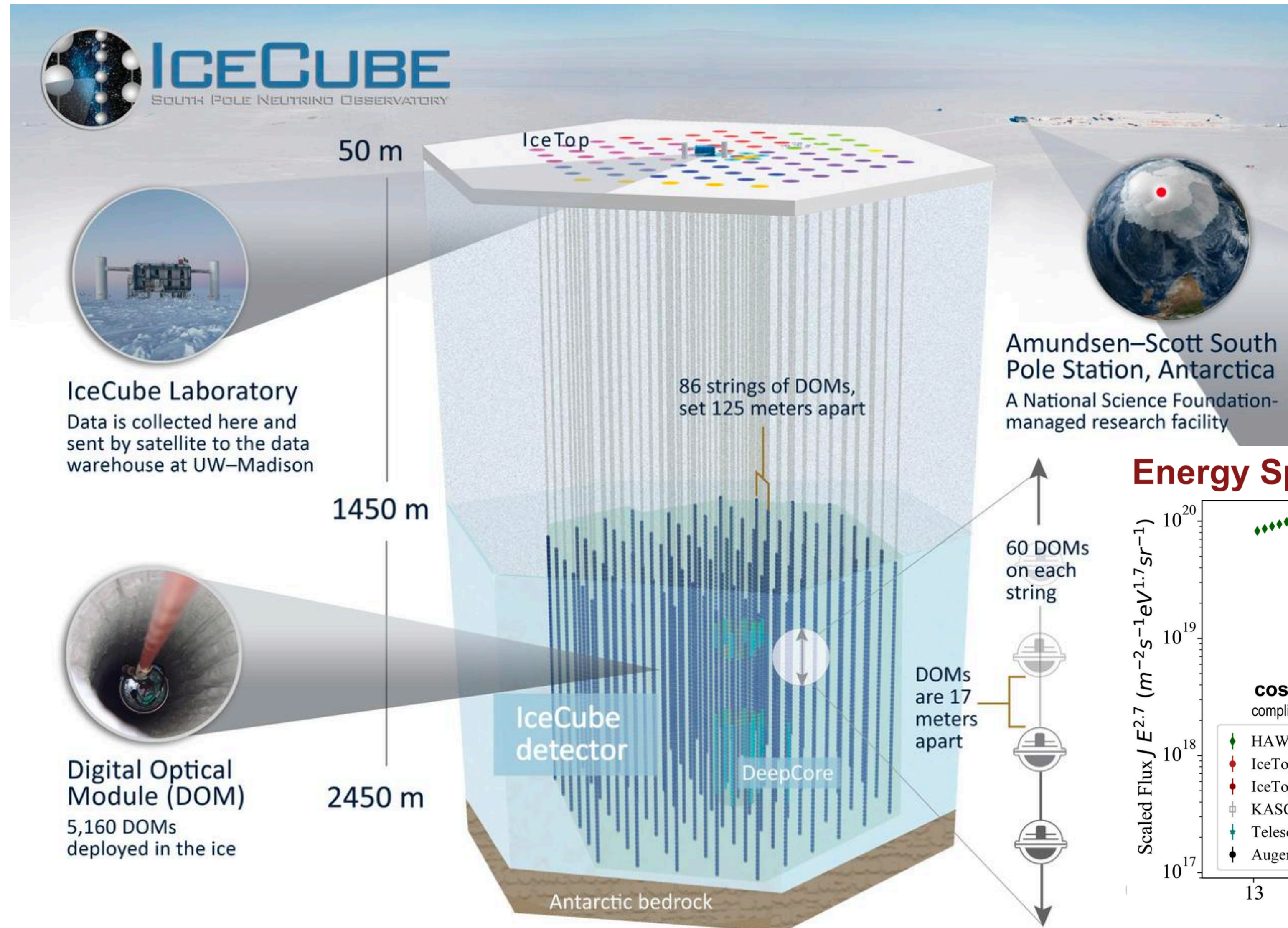
08/10/2022



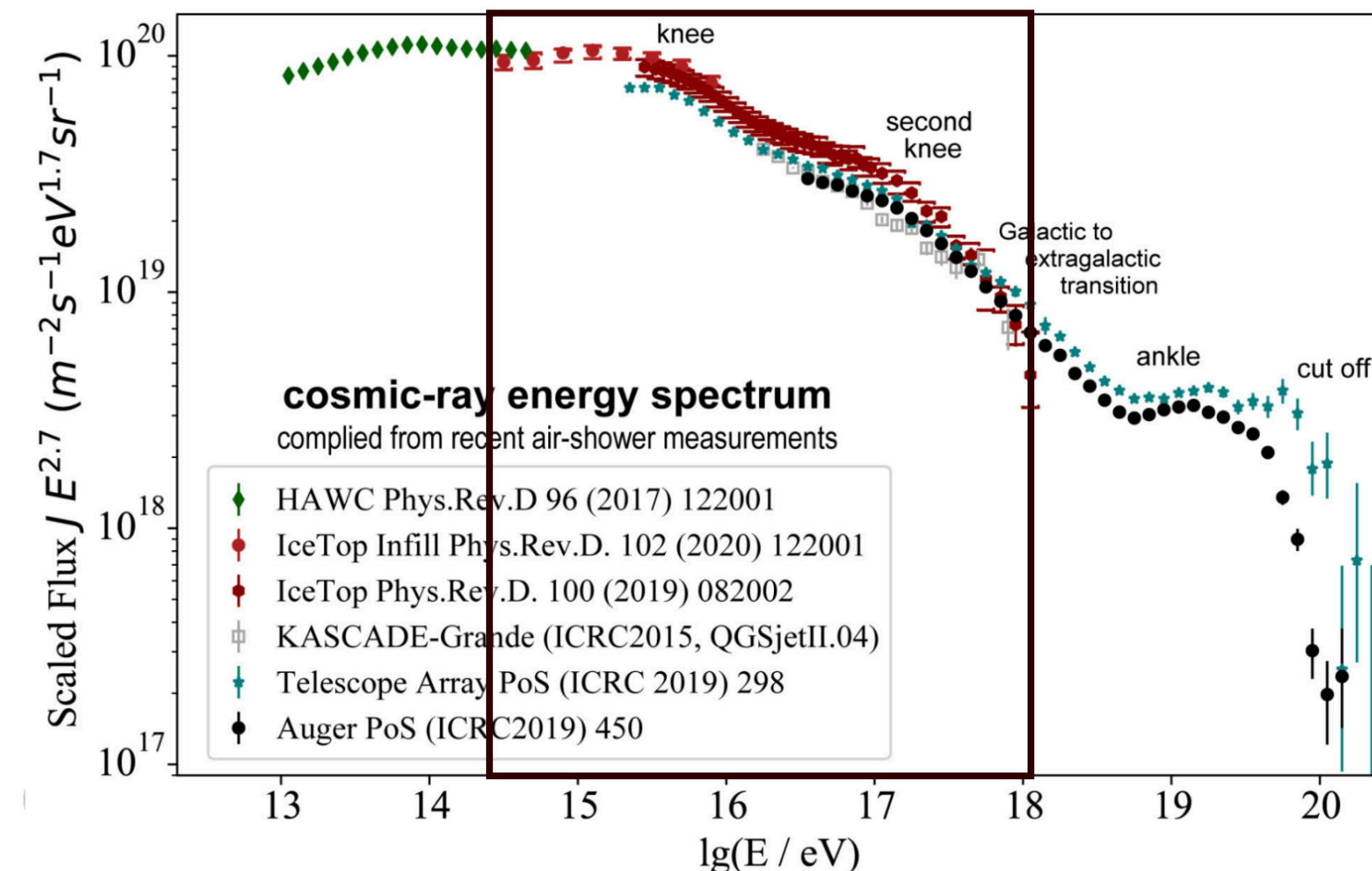
European Research Council
Established by the European Commission

This project has received funding from the European Research Council (ERC) under the European Union's Horizon 2020 research and innovation programme (grant agreement No 802729).

The IceCube Neutrino Observatory



Energy Spectrum of Cosmic Rays

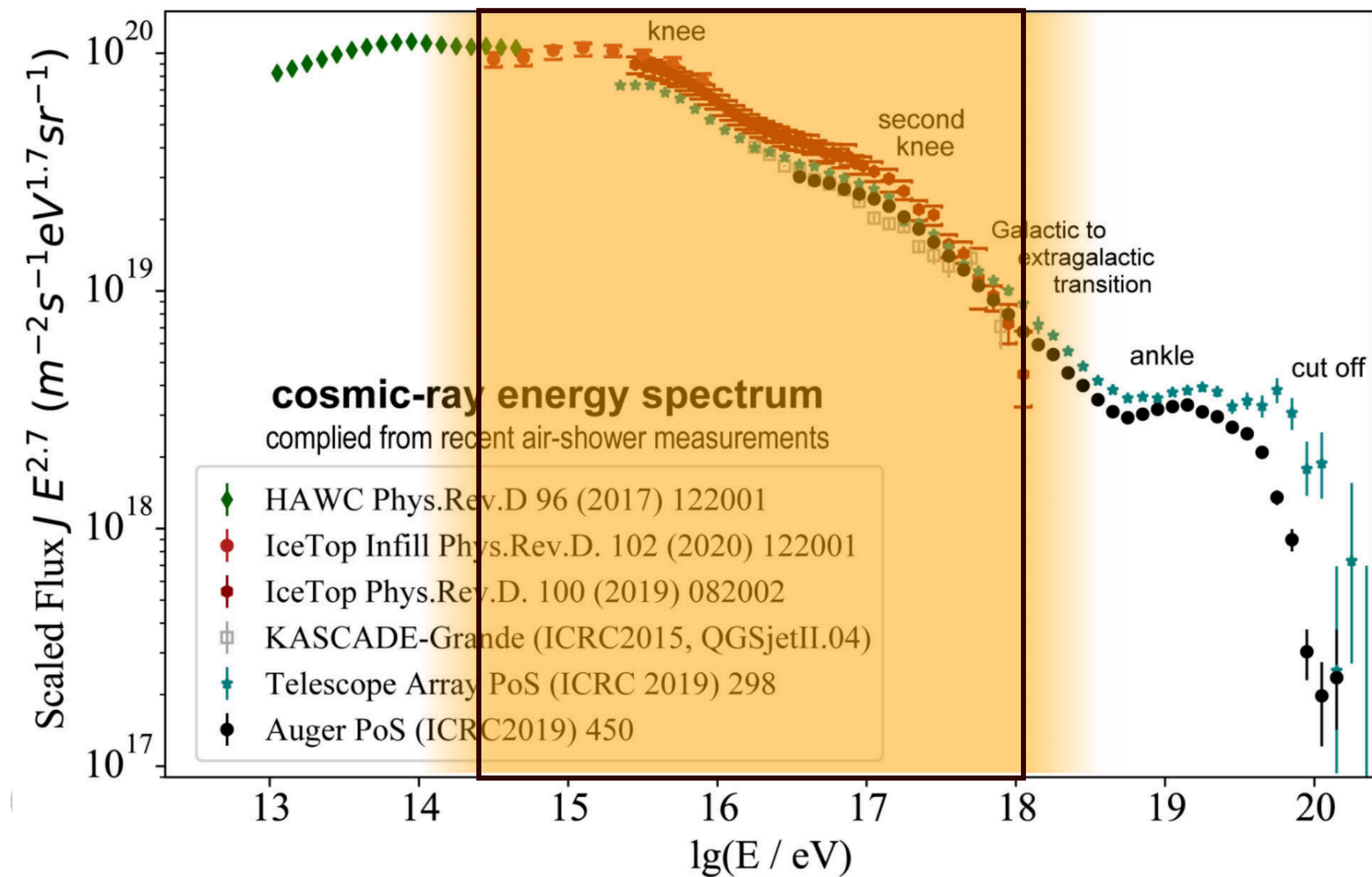


F.G. Schroeder, PoS(ICRC2019)030

Planned Surface Array Enhancement

- Mitigate the **snow accumulation** on the IceTop tanks
- Supporting IceCube's Neutrino mission
 - Improving the **background veto** for IceCube
 - Improving the understanding of **atmospheric background**
- **Cosmic rays physics**
 - Extending the **energy range**
 - Adding independent measurements in the **~ PeV region**
 - Exploring the **galactic to extragalactic transition**
 - Increasing **mass composition** resolution
- **R&D for IceCube-Gen2 (surface)**

Energy Spectrum of Cosmic Rays

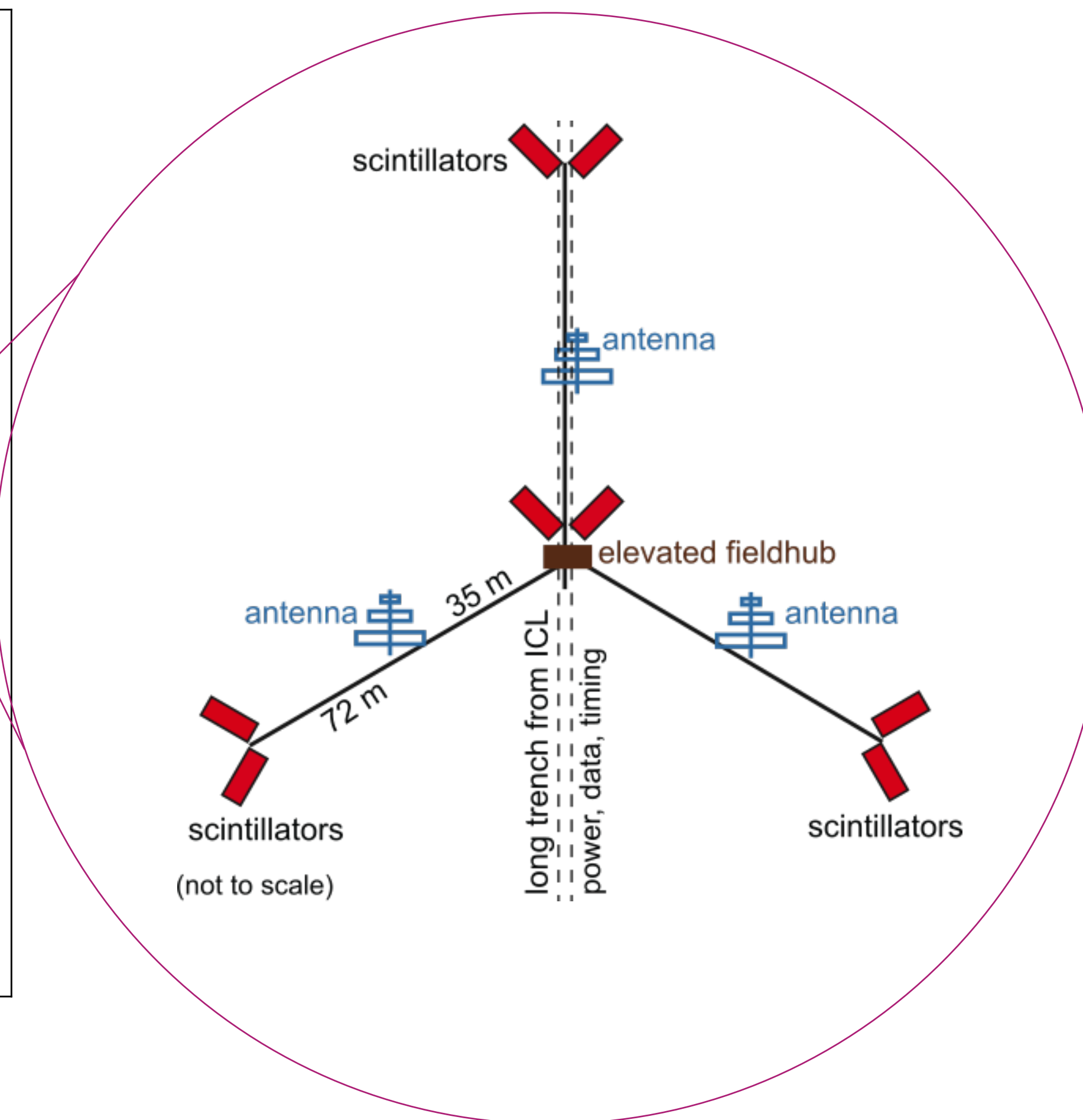
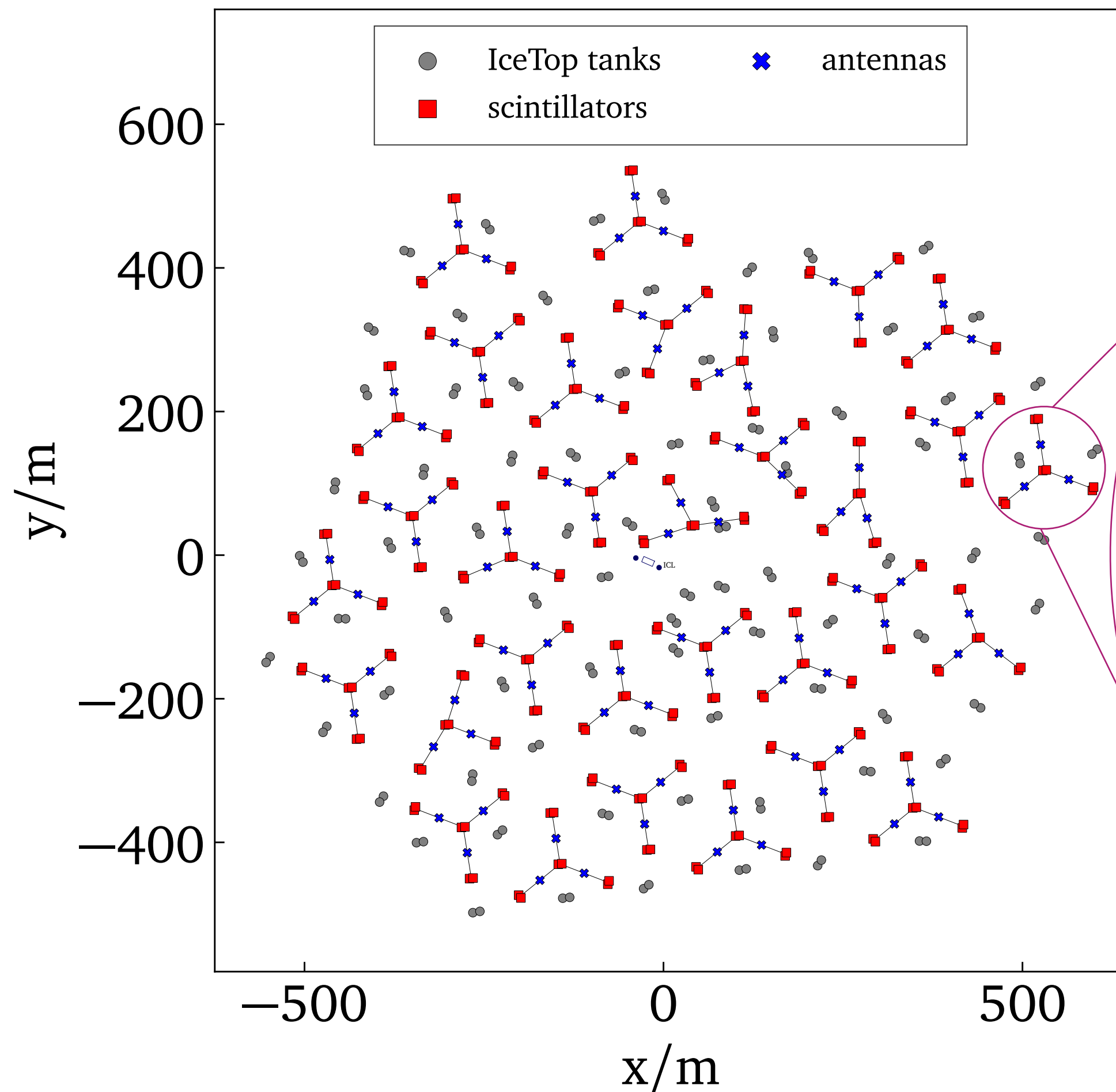


F.G. Schroeder, PoS(ICRC2019)030 (modified)

Planned Surface Array Enhancement

32 Stations

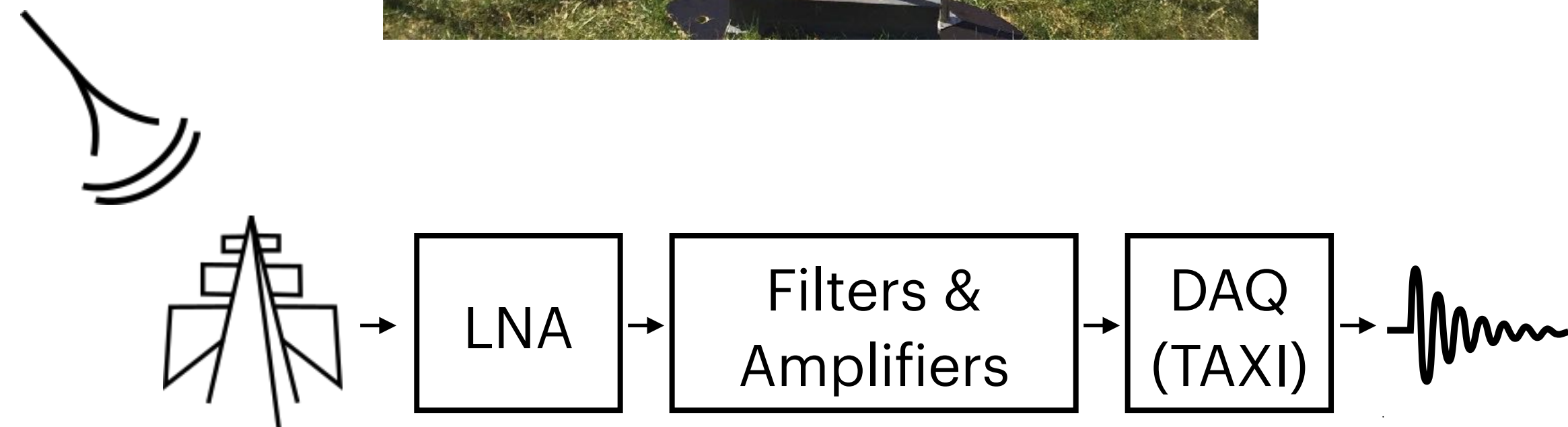
- 8 scintillation panels
- 3 antennas



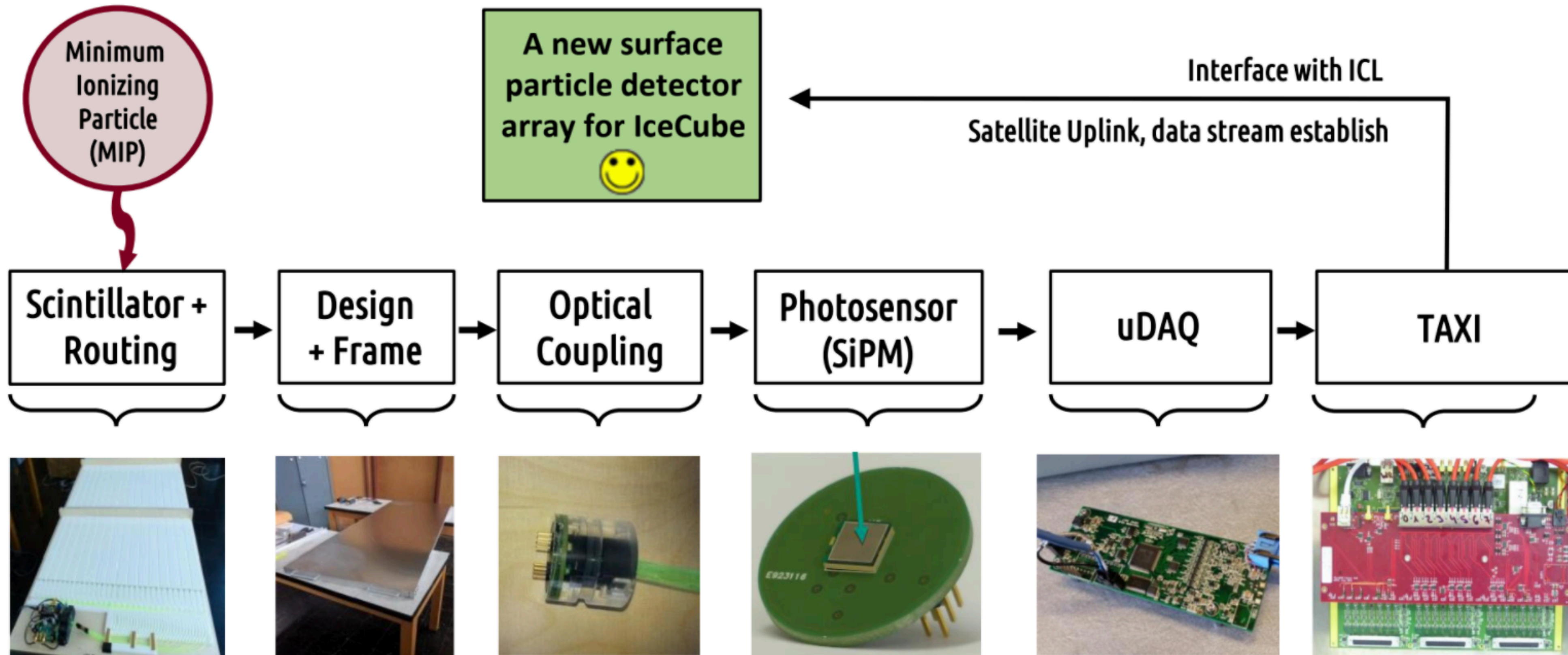
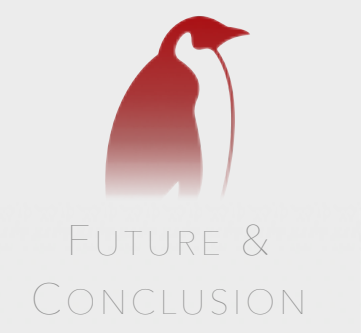
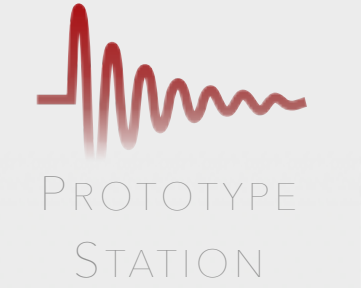
Elevated

Radio Antennas

- SKALA-V2 LPDA Antennas
- Nominal bandwidth : 70MHz to 350MHz
- Integrated LNAs with high amplification and low noise (40K)
- Custom made radio pre-processing board for additional filtering and amplification
- Omni-directional
- Good directivity for $\pm 60^\circ$ from zenith
- Low power
- $\pm 0.5^\circ$ uncertainty on the levelling
- Low cross-polarisation
- Easy assembly in the field



Elevated Scintillation Panels

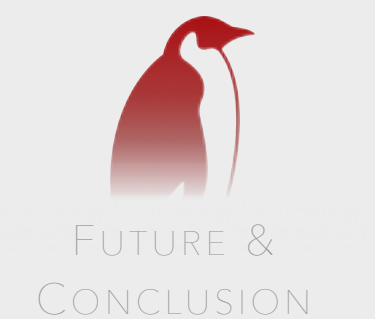
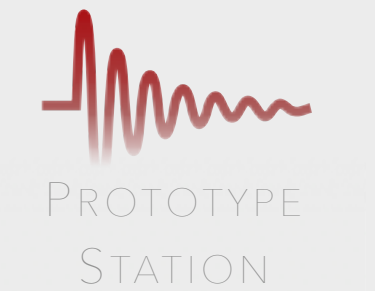
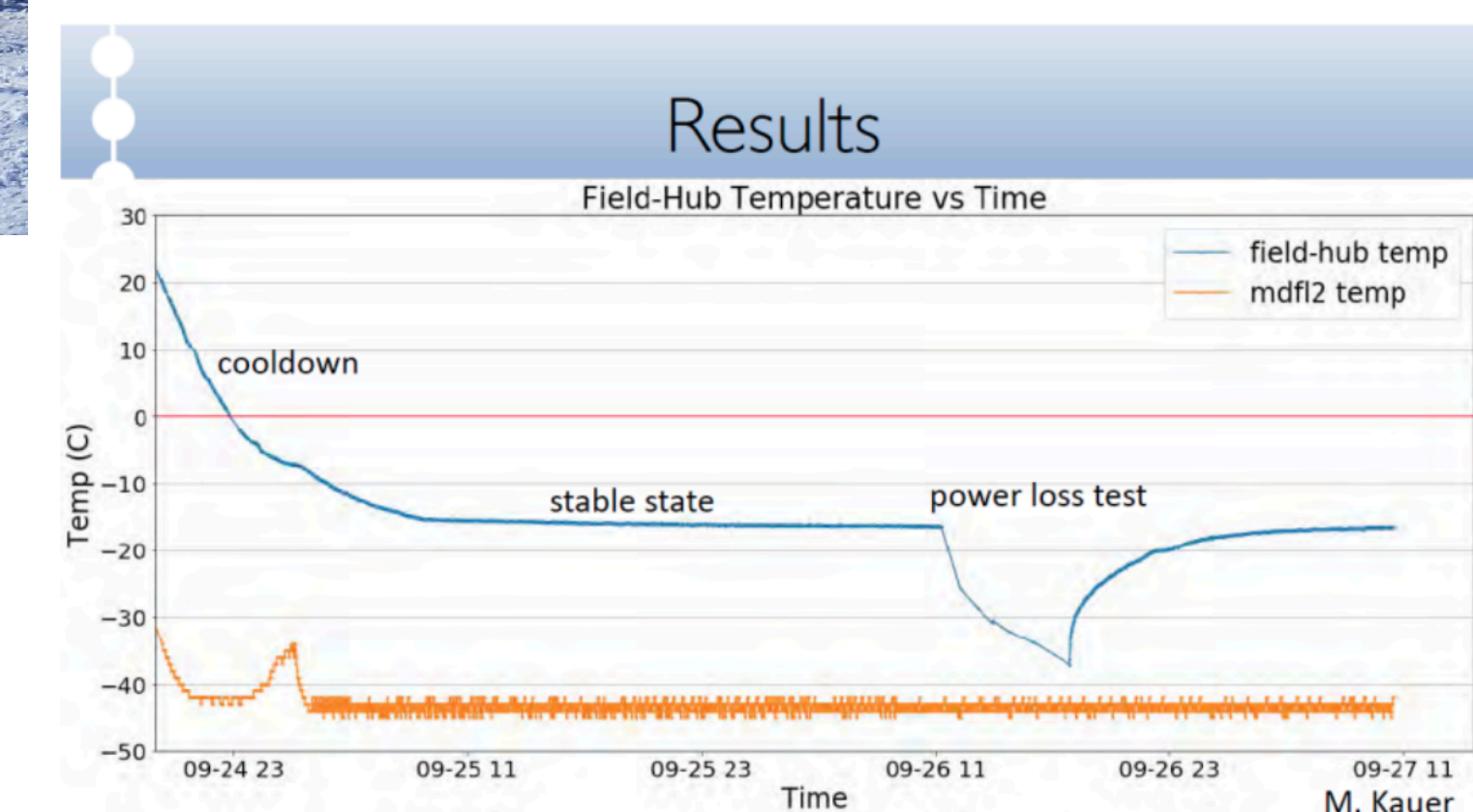
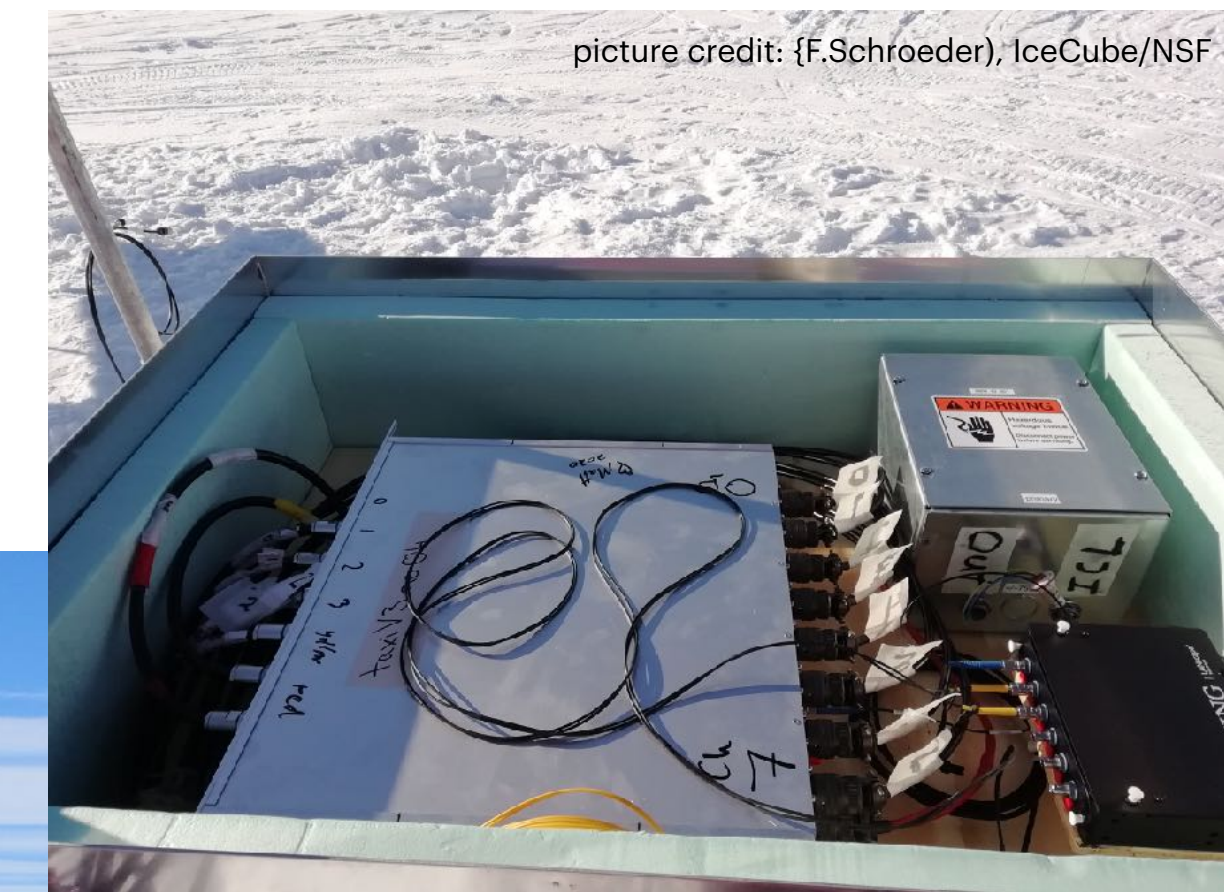


T. Huber

Elevated **Field Hub**

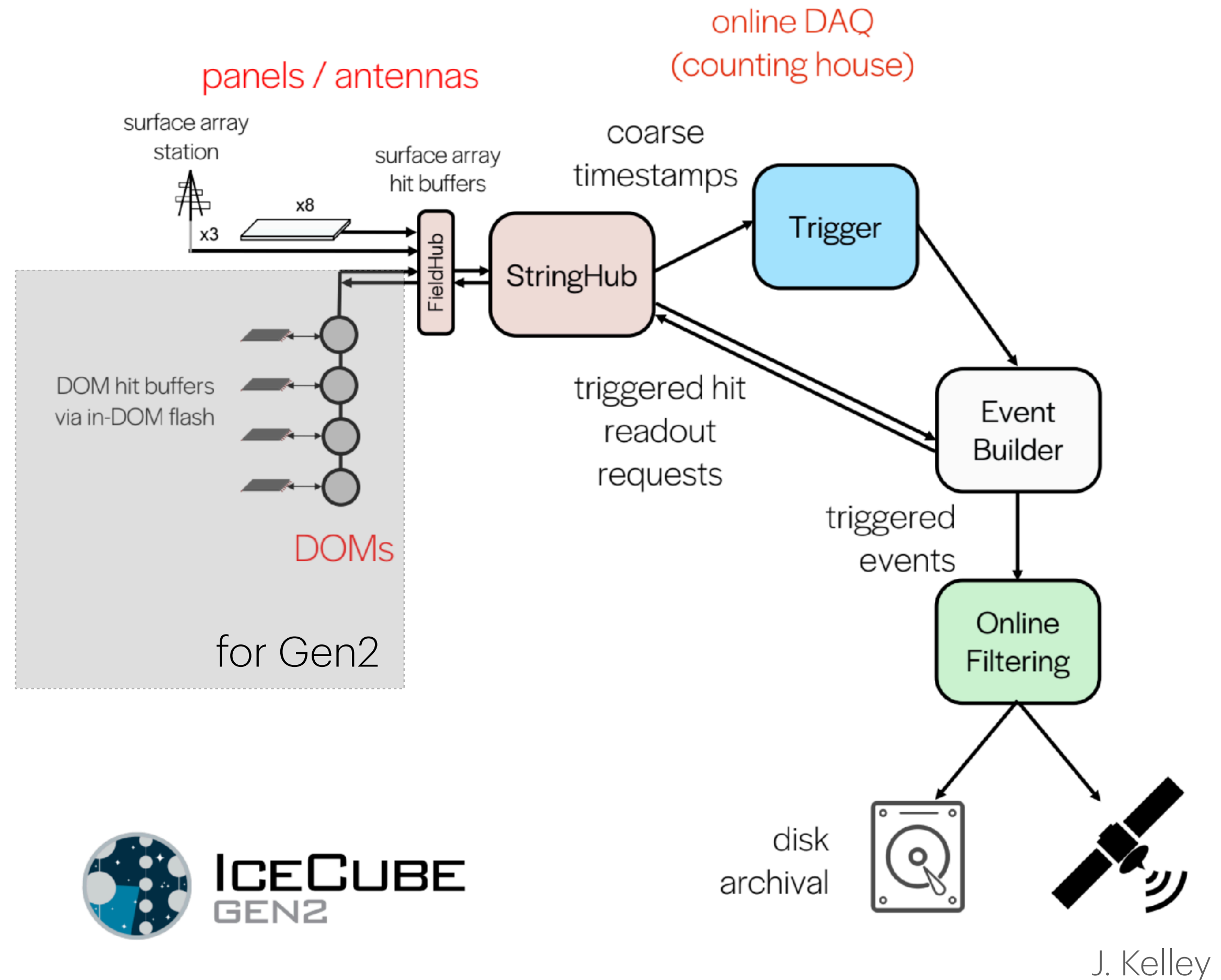
- Insulated container for electronics
 - Ensure a higher and more stable temperature for the electronics. Also provide a slower cool-down in case of power outage
- **Double shielding** to mitigate RFI
- Will have some **storage** for the **cable slack**
- Separated DC/DC converter box for safety

picture credit: {F.Schroeder), IceCube/NSF



Infrastructure

- Communication
 - **Integrated** in the IceCube Infrastructure
 - Use of existing **satellites** for data transfer
- Data transfer and power routed via **cables** from/to ICL
- Future **trigger logic** planned with **in-ice coincidences**
- Power
 - **40-65 W** per station
 - Small fraction of IceCube power budget

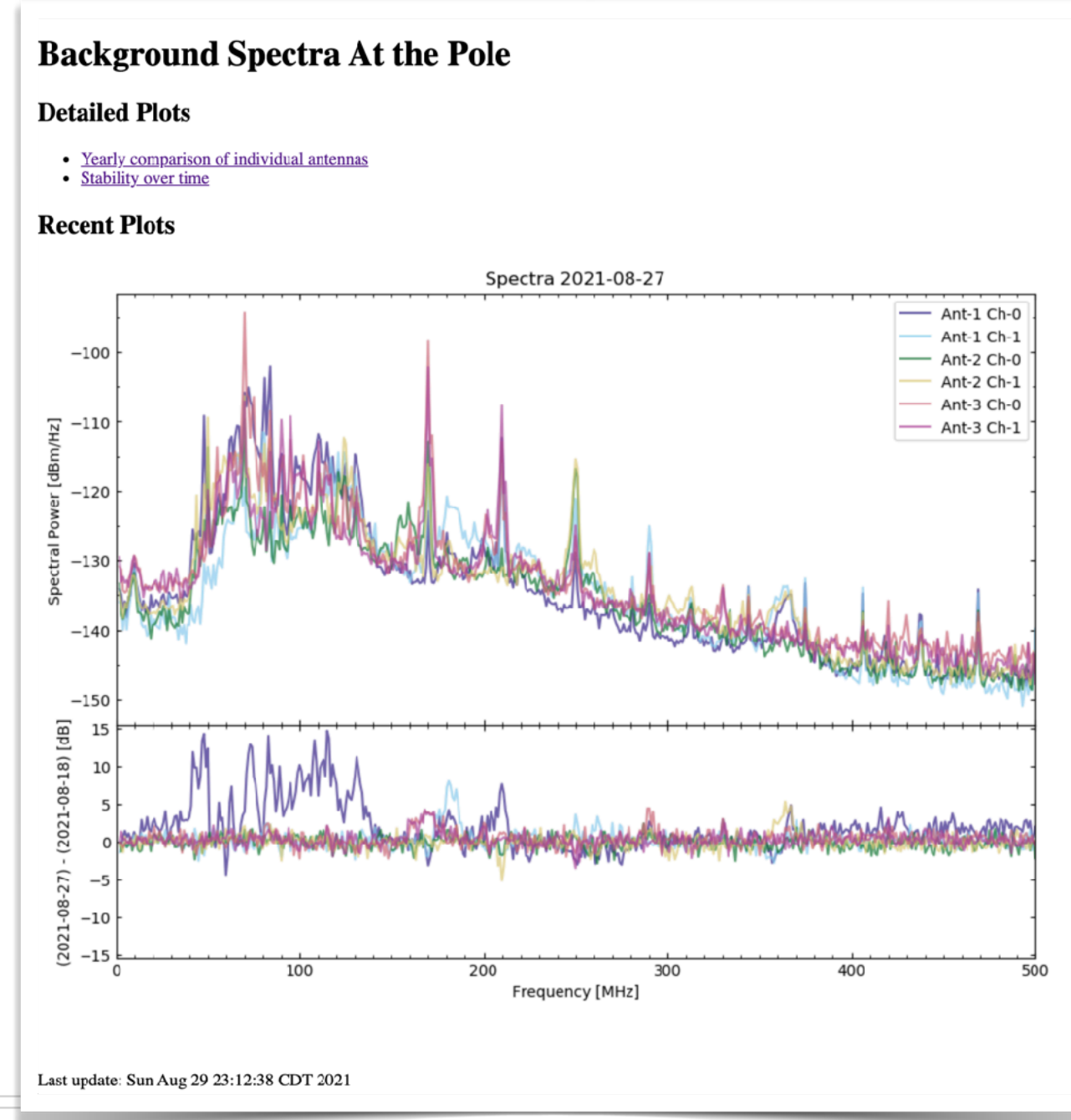
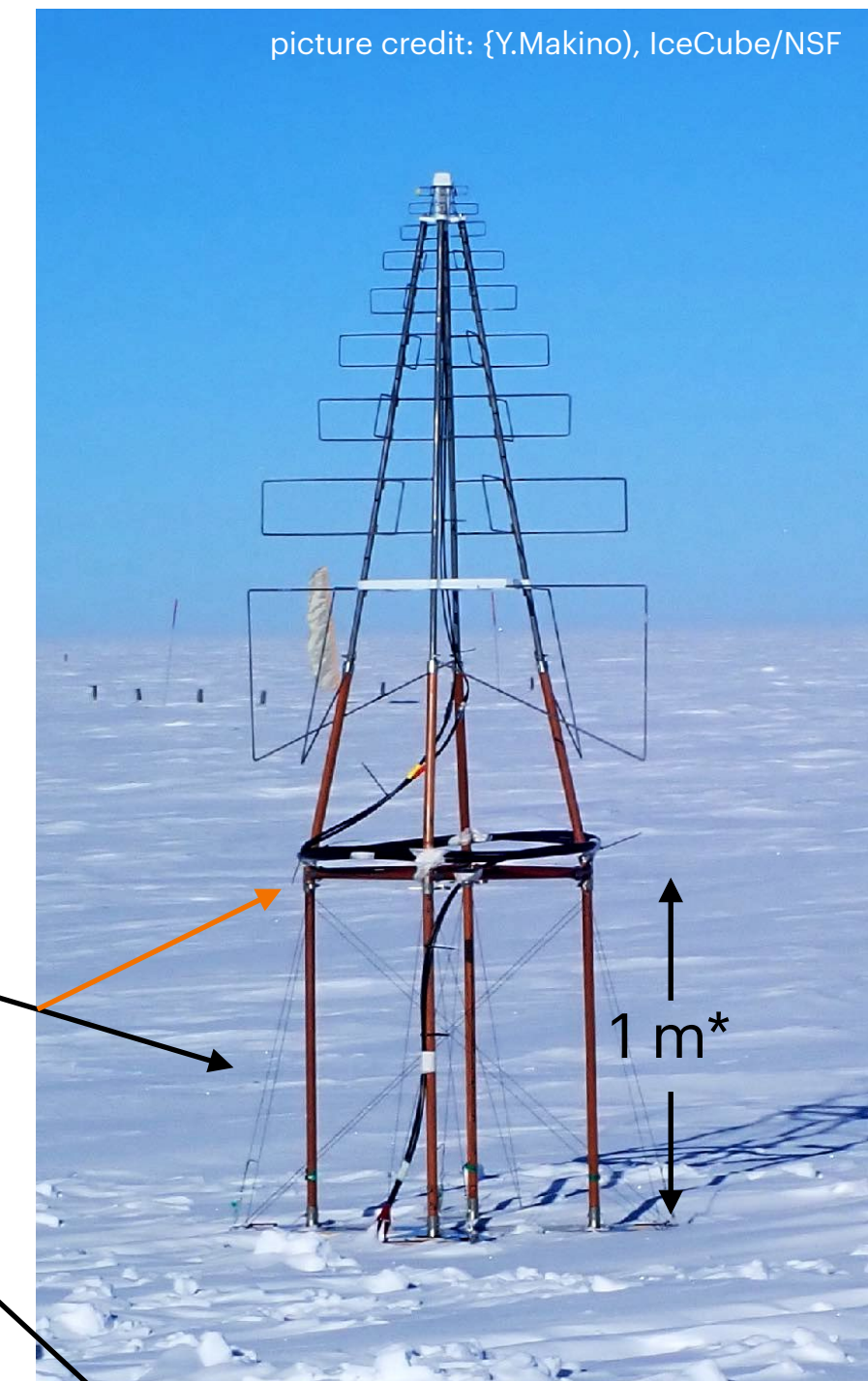


Maintenance

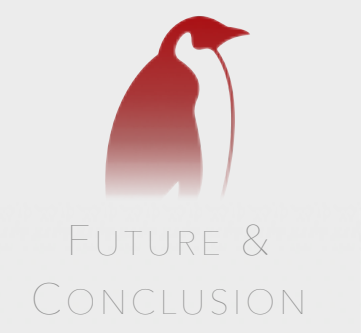
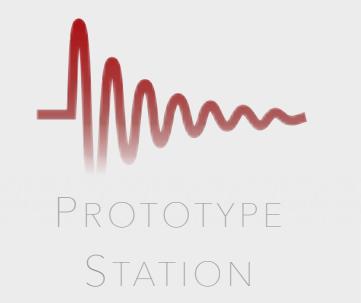
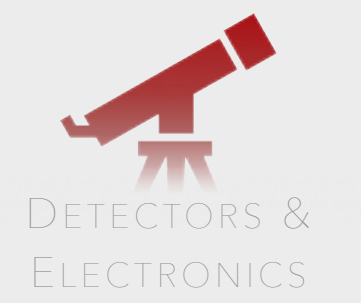
- Average of 20 cm/year of snow accumulation
- Raising the individual detectors every 5 years. The maintenance will naturally be staggered (with deployment)
- About 2 person for 2 weeks every year for maintenance for full array

Monitoring

- Antennas' baseline monitoring already online, daily trigger rates
- Alerts sent when abnormal behavior
- Will be extended (e.g. scintillators' gain, DAQ temperature, etc)

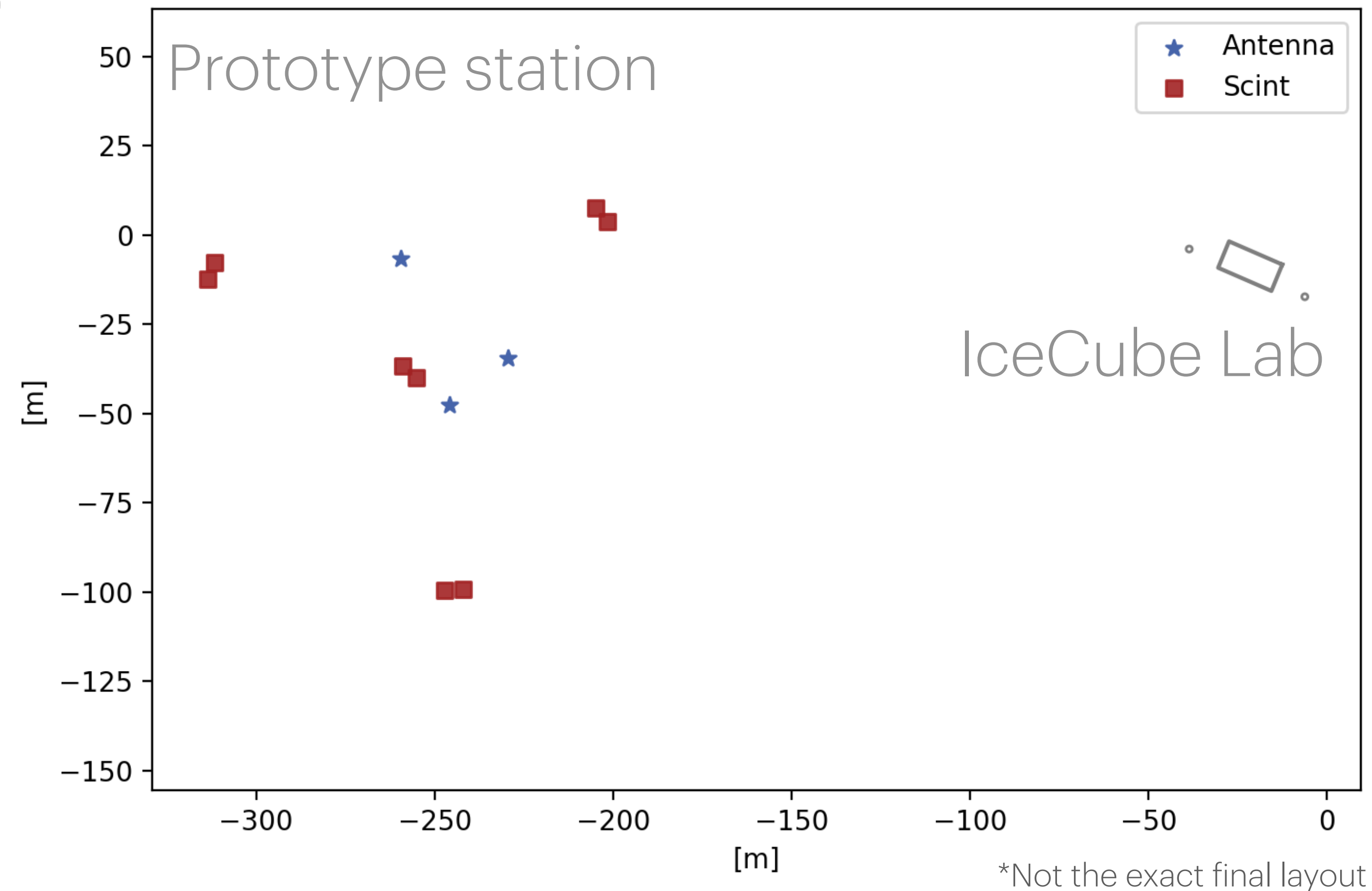


*The final version has 1.20 m



Prototype Station

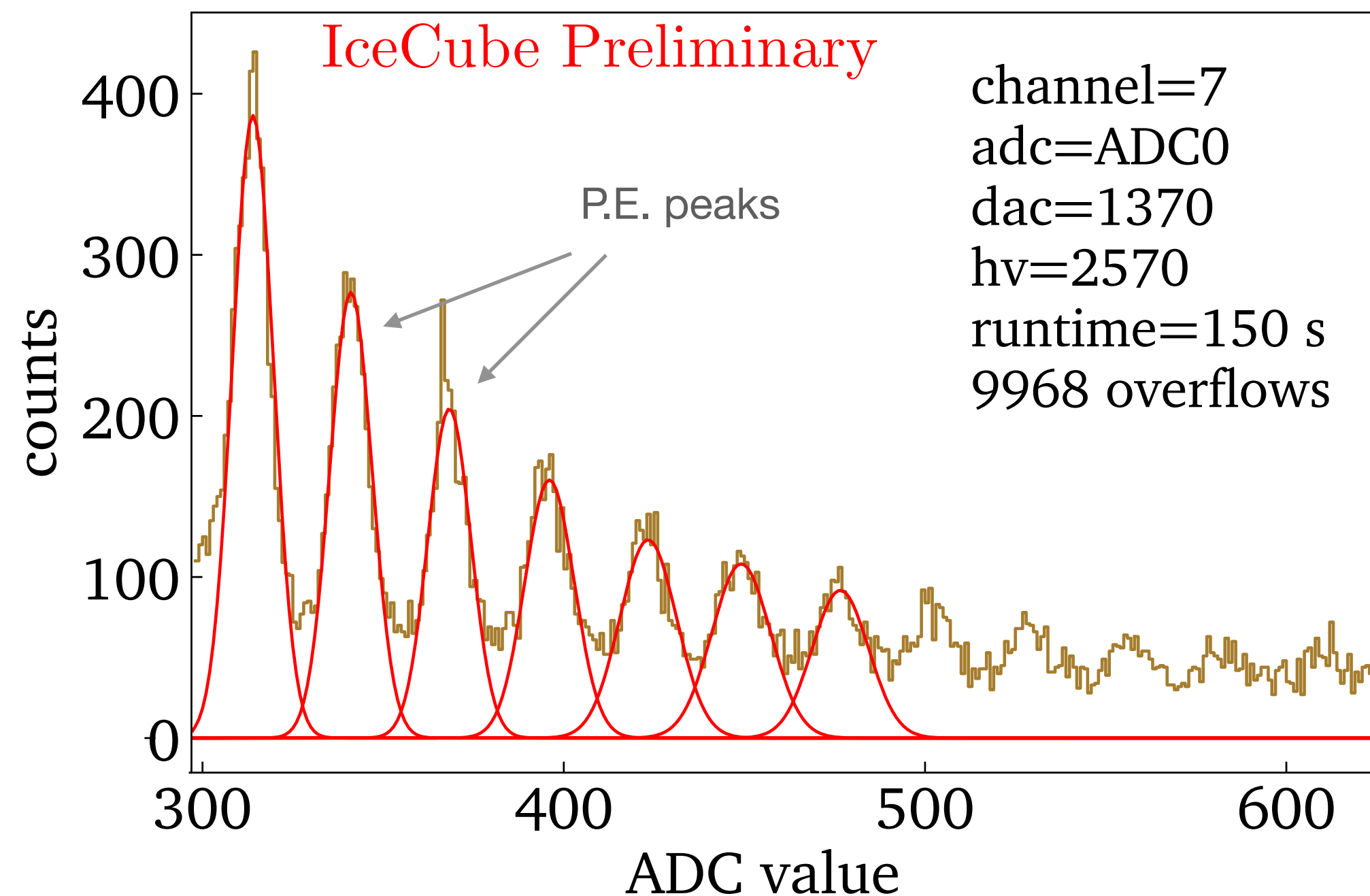
- Deployed in January 2020
 - 8 scintillator panels
 - 3 radio antennas
- **Continuously recording** (background and event data)
- **Air showers recorded!** (next slides...)
- Further analyses on-going



Raw data and background

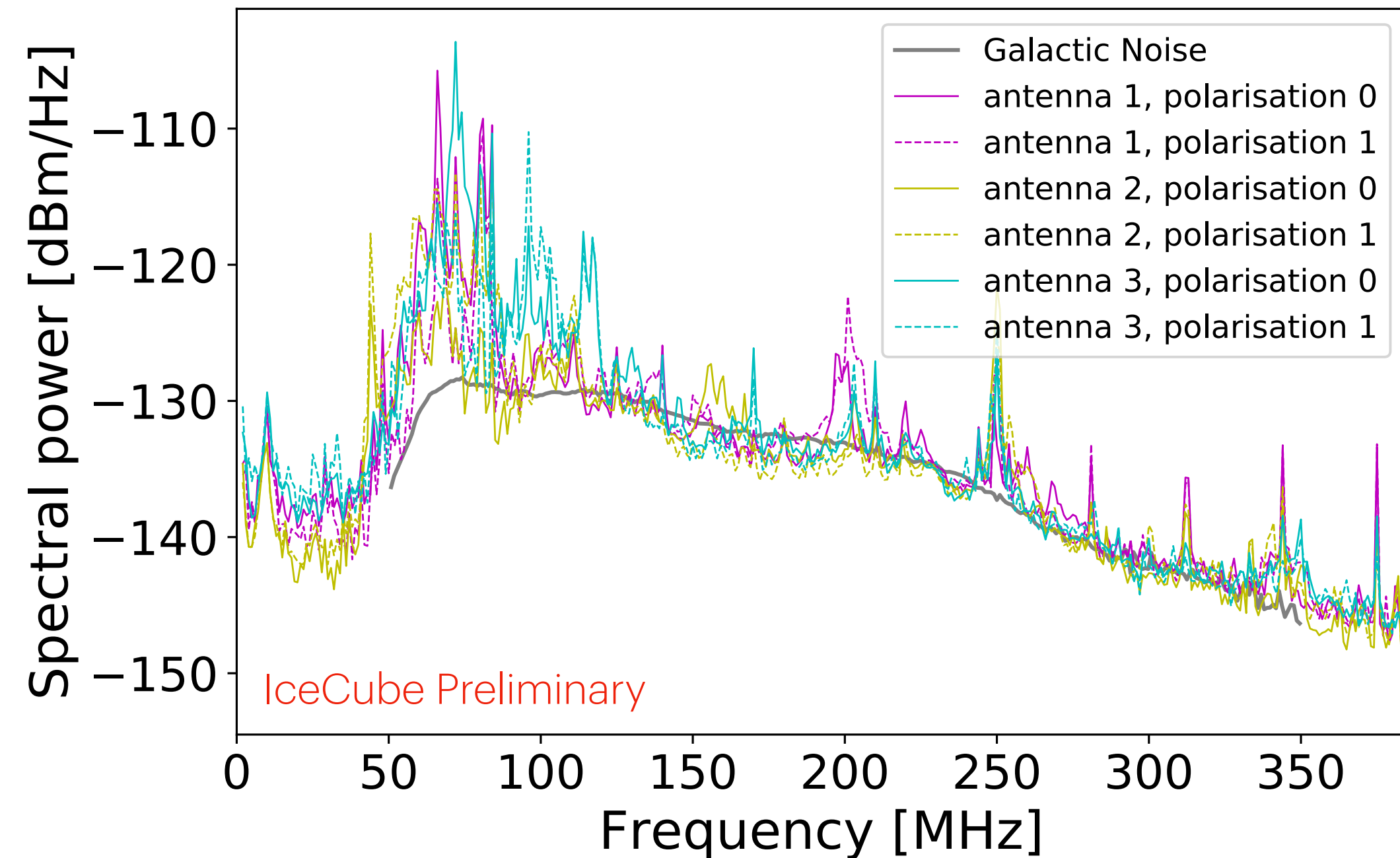
Scintillator panels:

- Example of a charge histogram (high gain)



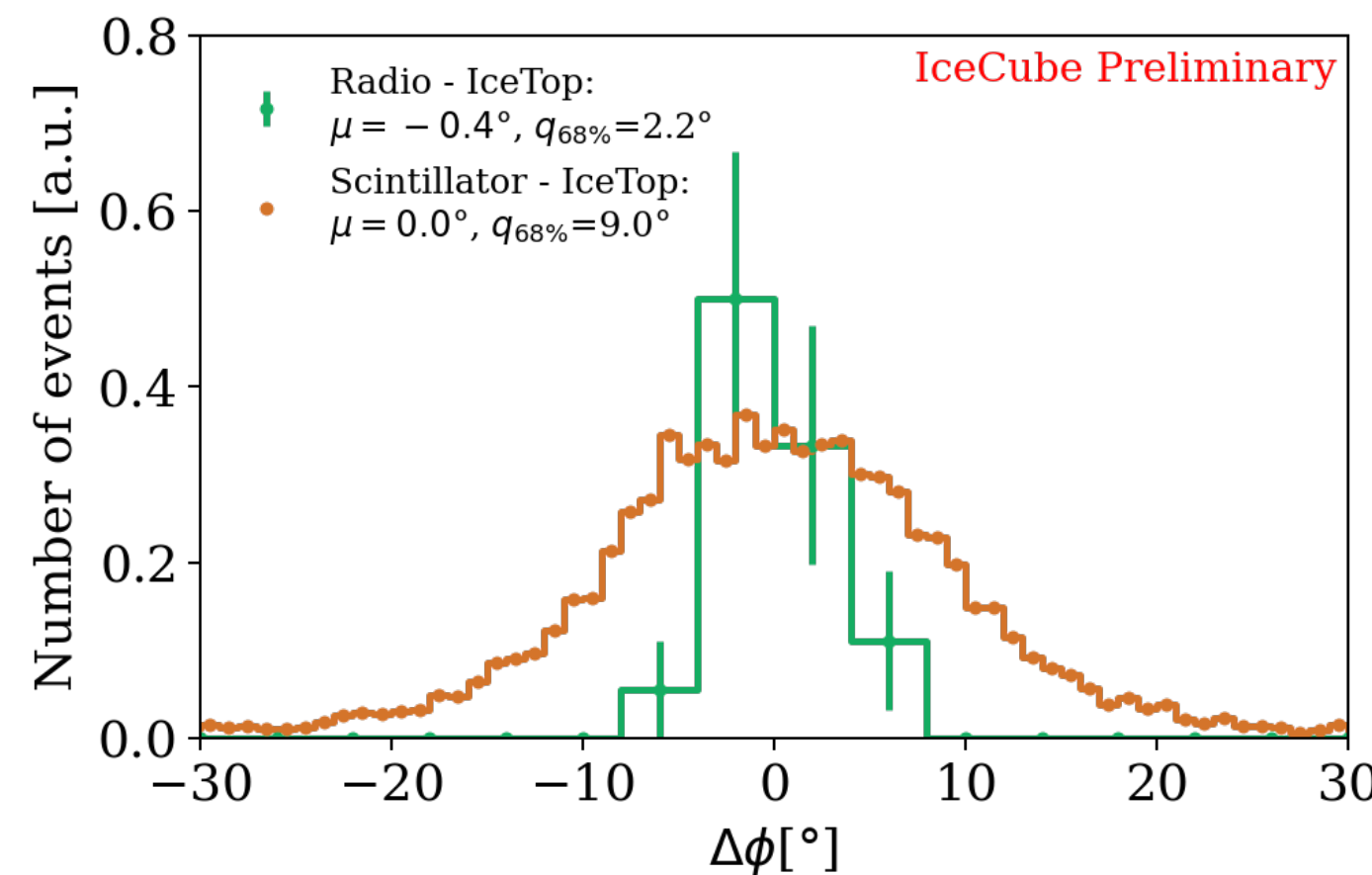
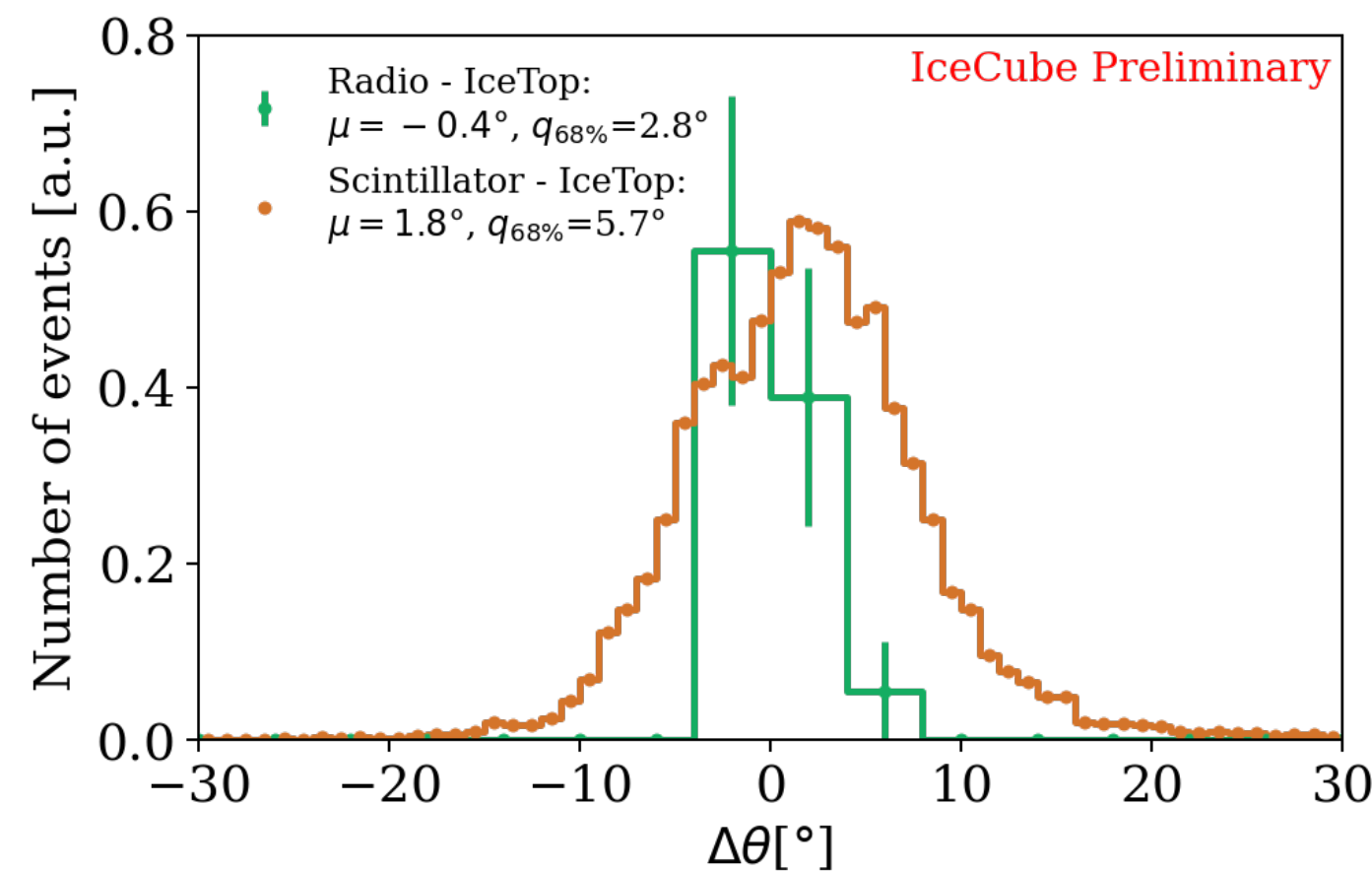
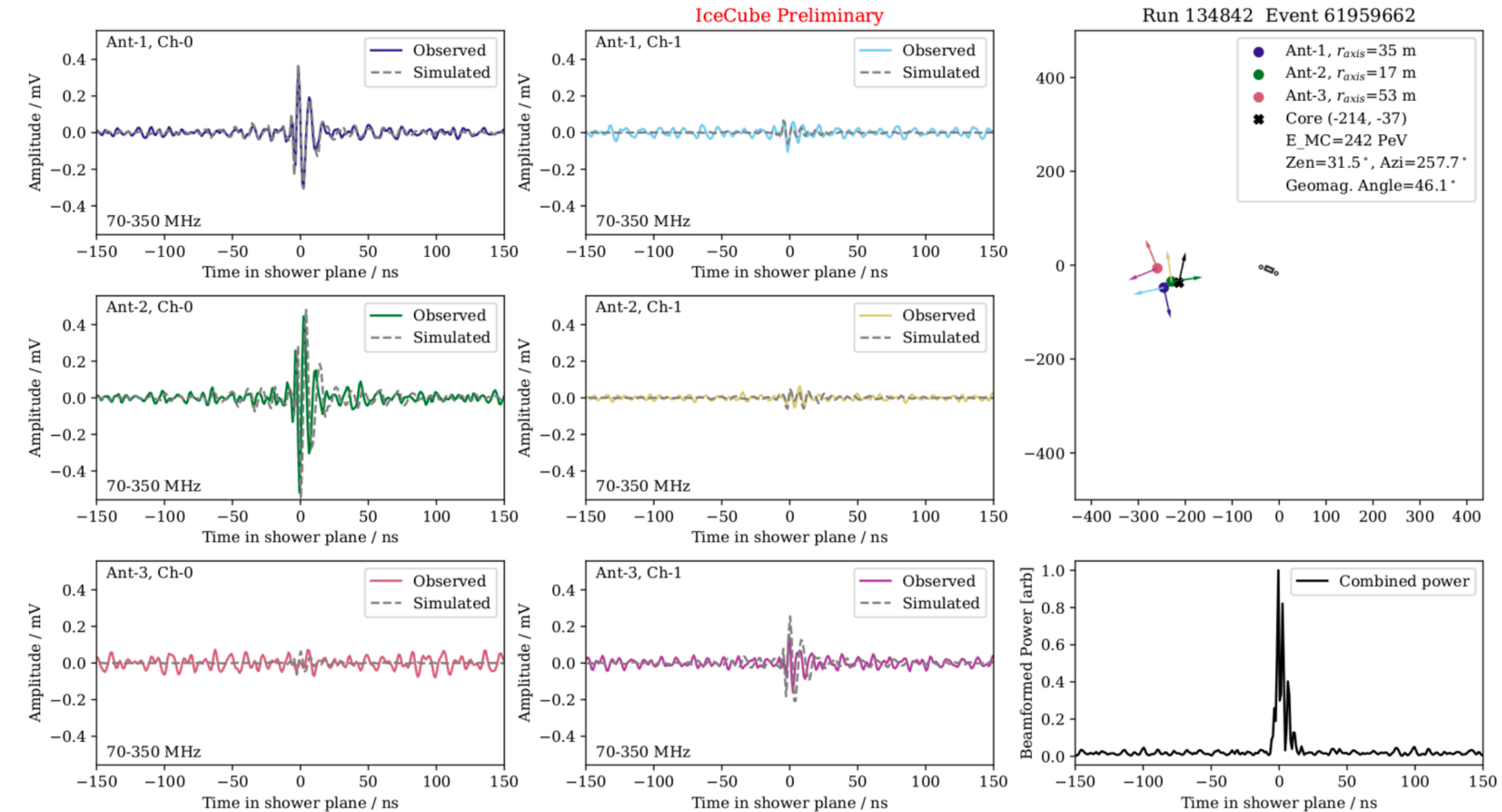
Radio antennas:

- Median spectrum (04/17/2020) compared with galactic noise (Cane model) and thermal noise (40K) in gray



Detection of air showers!

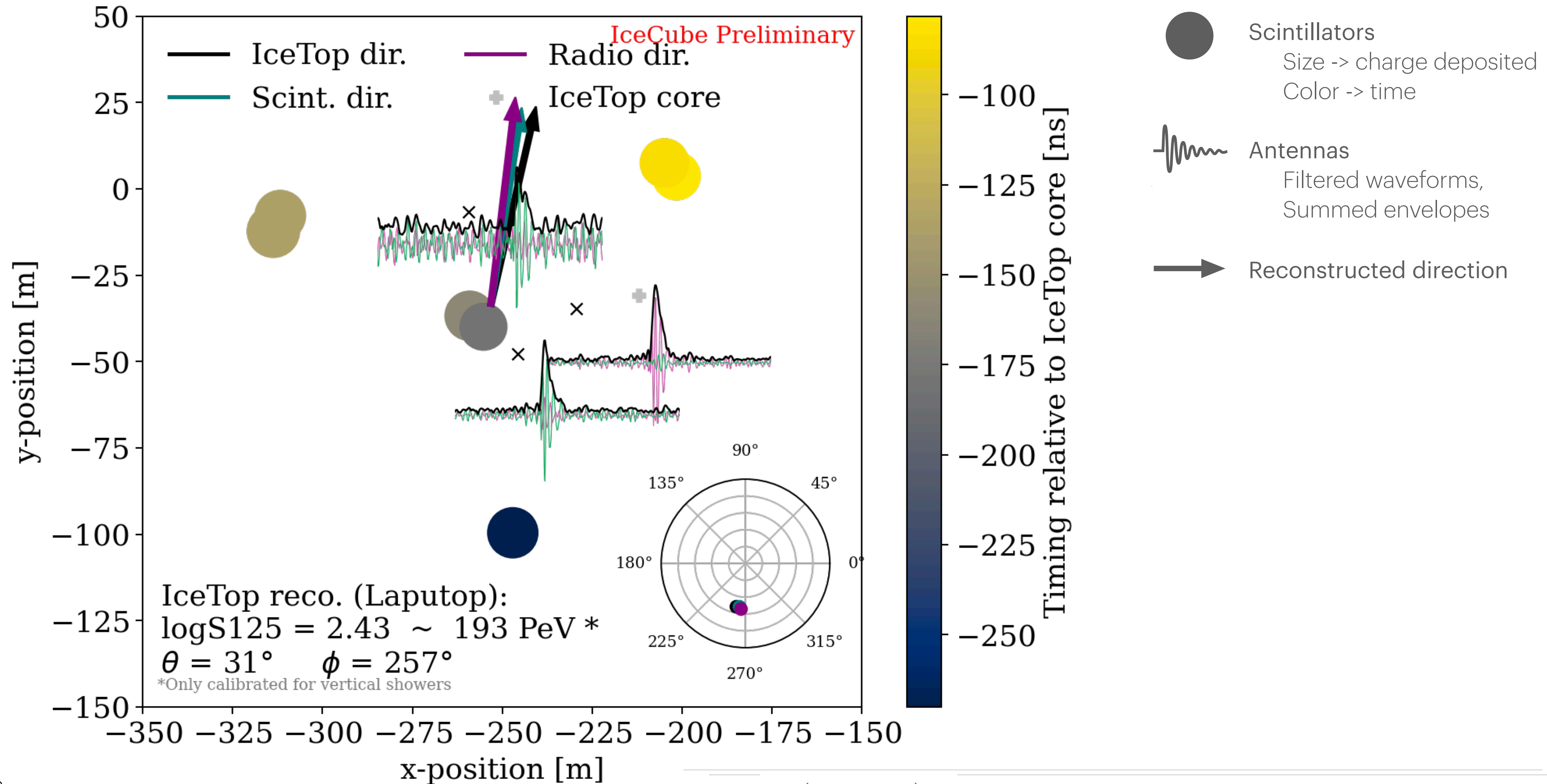
Example of an air-shower recorded (full lines) by the antennas and compared with simulation (dashed lines)



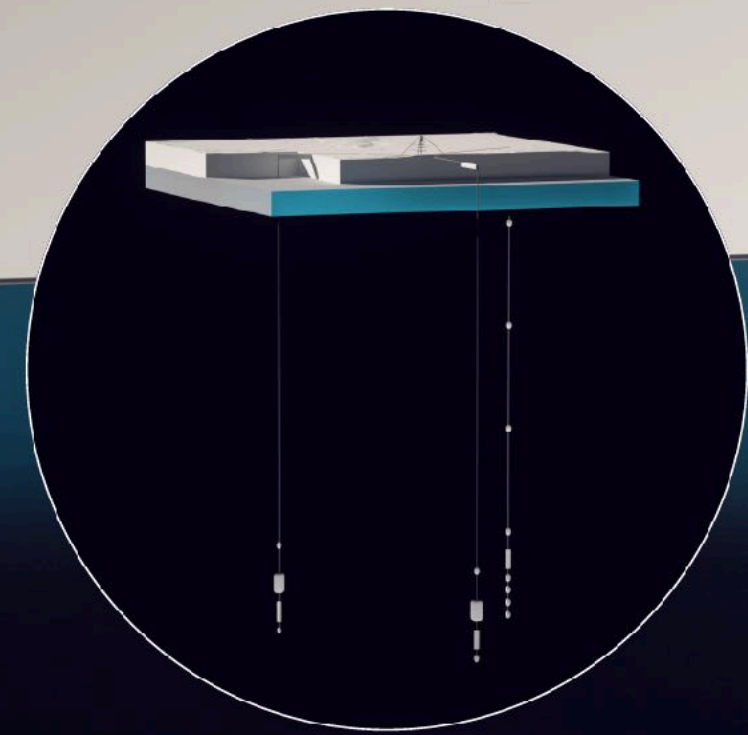
Arrival directions compared to IceTop (plane wavefront assumed)

A. Coleman, [PoS\(ICRC2021\)317](#)

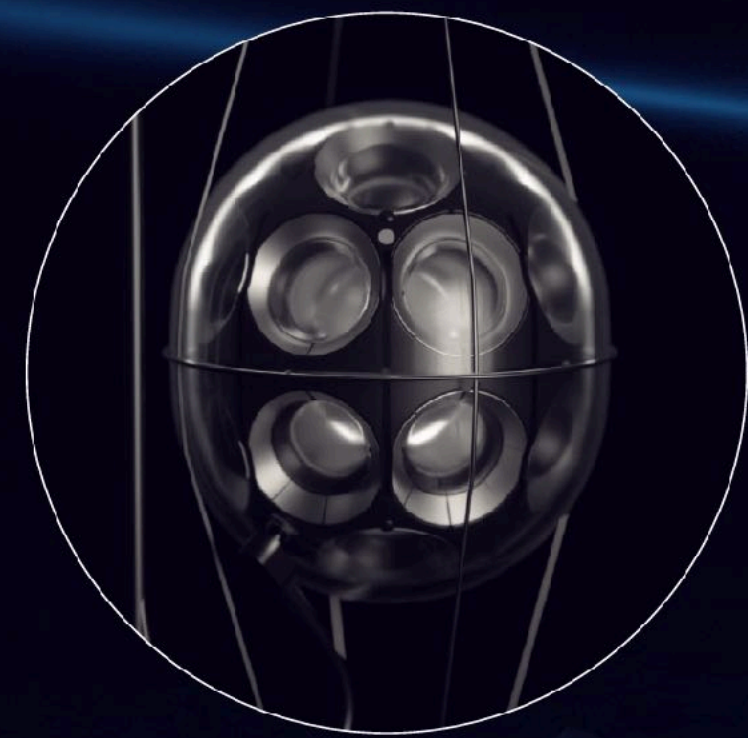
Triple coincidences



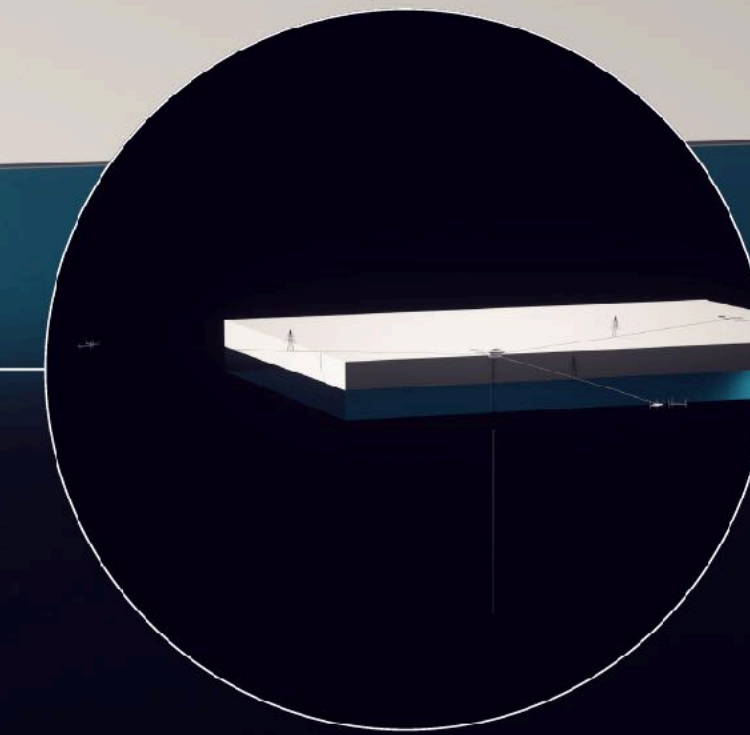
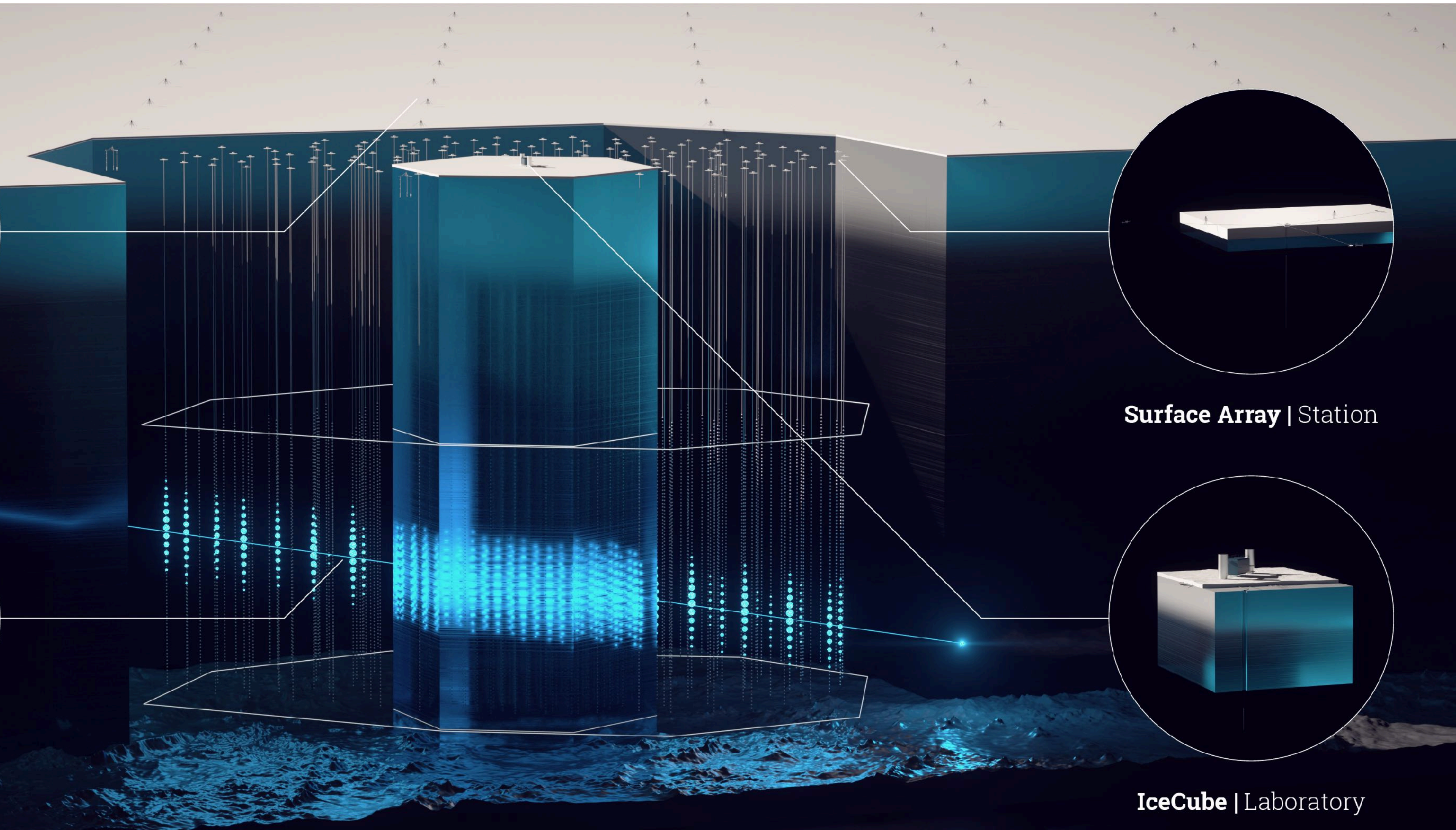
Outlook into the future - IceCube Gen2



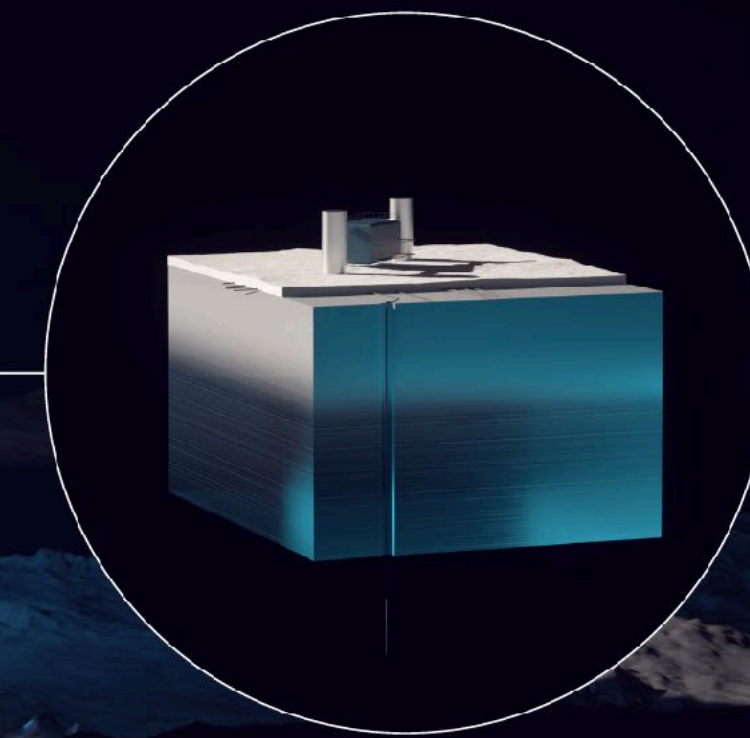
Radio Array | Station



Optical Array | Sensor



Surface Array | Station



IceCube | Laboratory



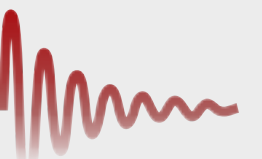
SURFACE ARRAY
ENHANCEMENT



DETECTORS &
ELECTRONICS



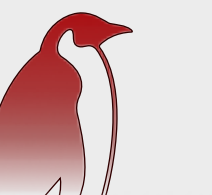
MAINTENANCE
& OPERATION



PROTOTYPE
STATION



RESULTS

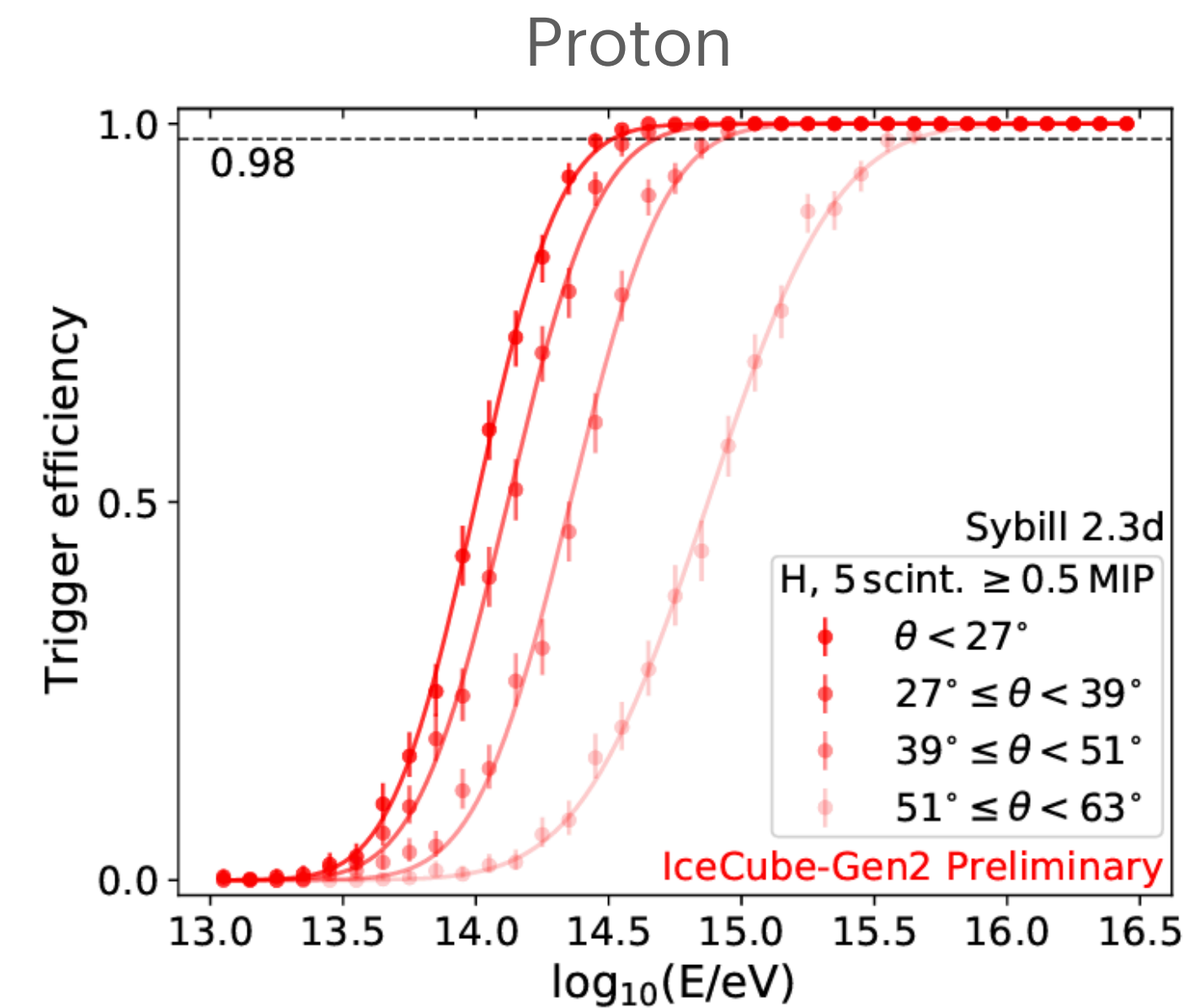
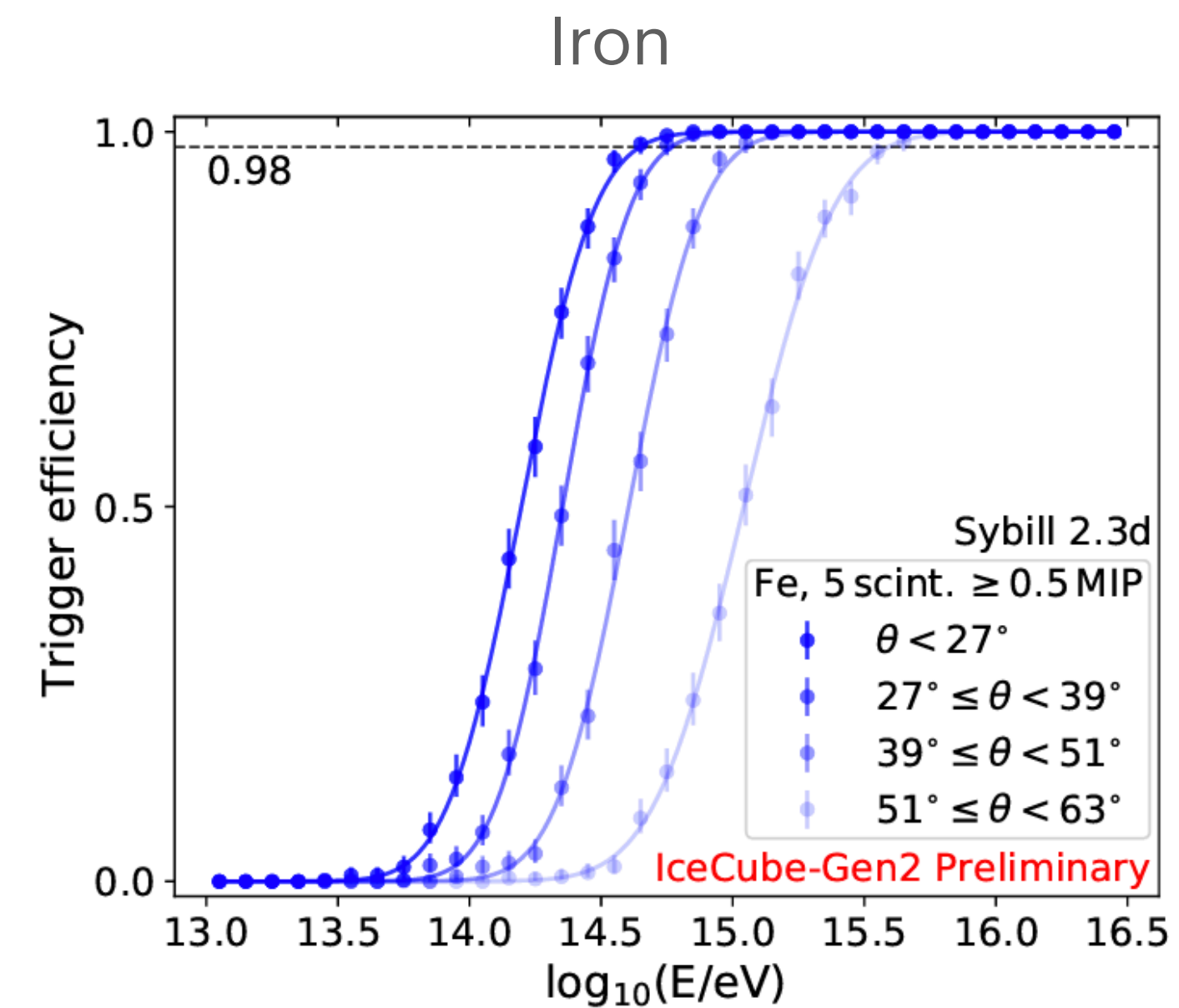


FUTURE &
CONCLUSION



A look into the future

- Veto the background for the in-ice detector (Gen2)
- Extending the energy range by a factor of 3 (fully covering the *predicted* transition region from galactic to extra-galactic CRs)
- 8-10x geometric aperture but 30x for in-ice coincidences aperture
- Observing the transition from conventional to prompt muon (0.5-1 PeV)
- Threshold of ~0.5 PeV



A. Leszczynska, M. Weyrauch, [PoS\(ICRC2021\)411](#)



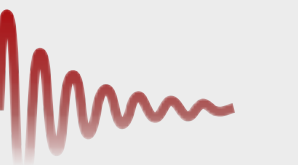
SURFACE ARRAY
ENHANCEMENT



DETECTORS &
ELECTRONICS



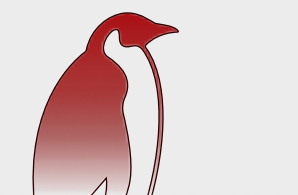
MAINTENANCE
& OPERATION



PROTOTYPE
STATION



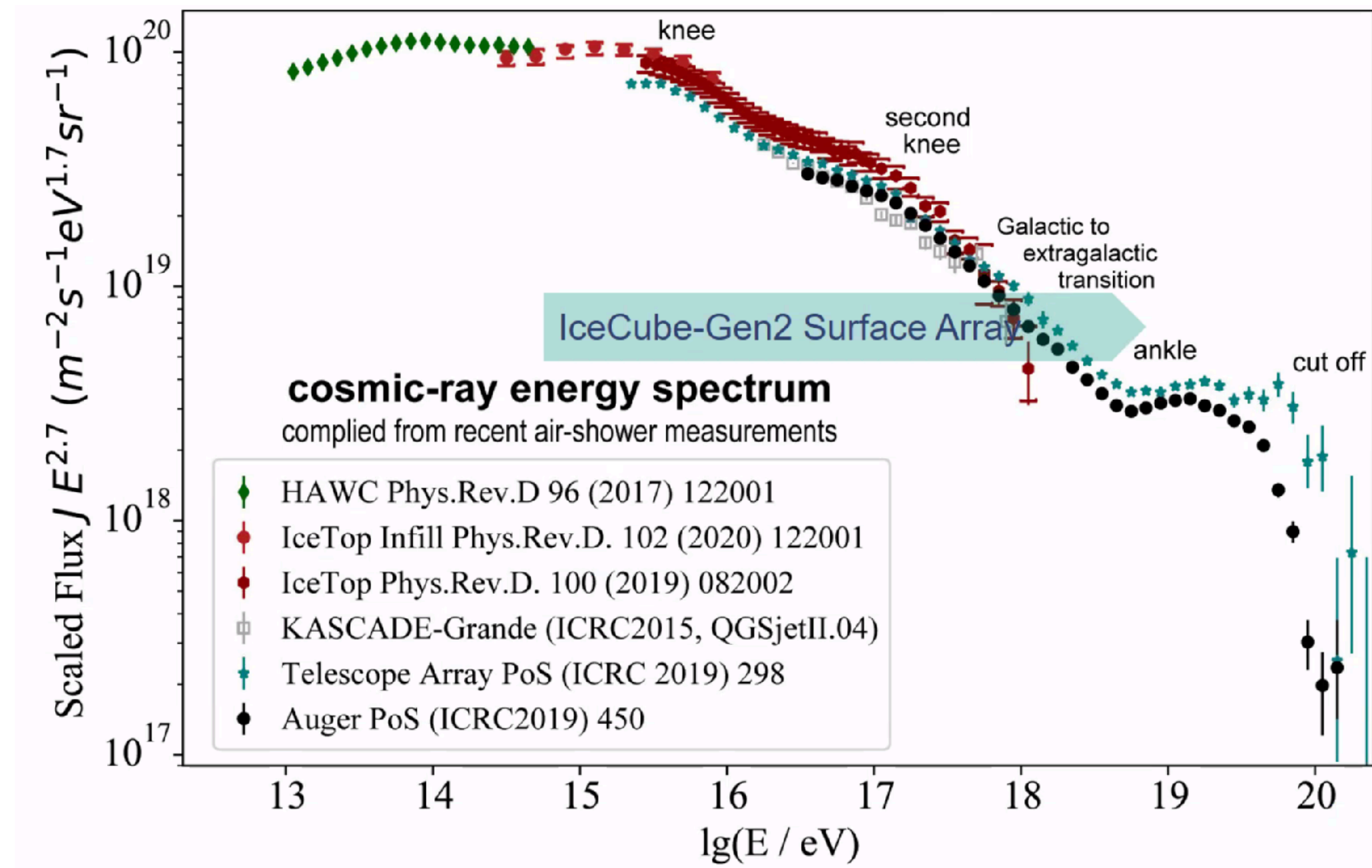
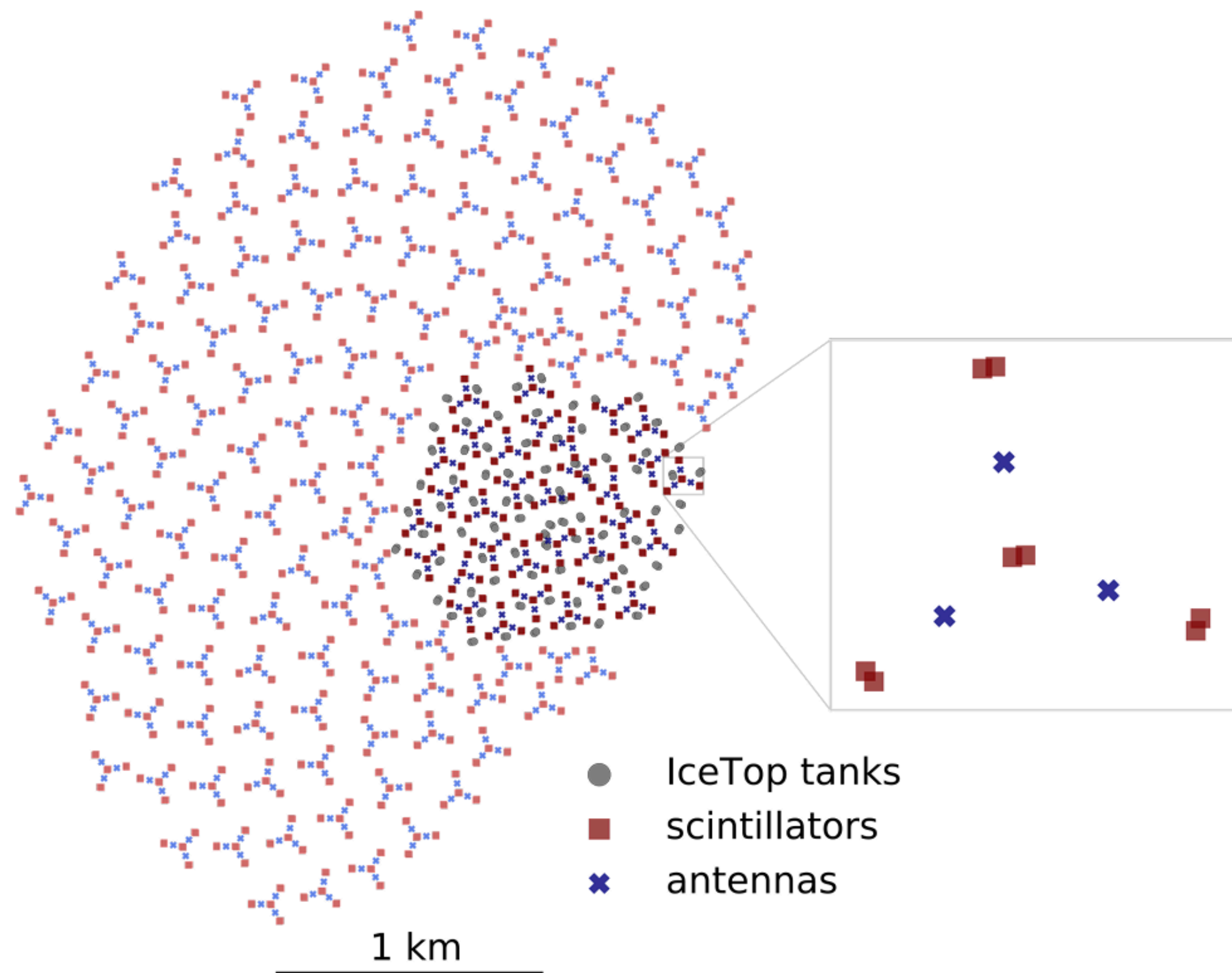
RESULTS



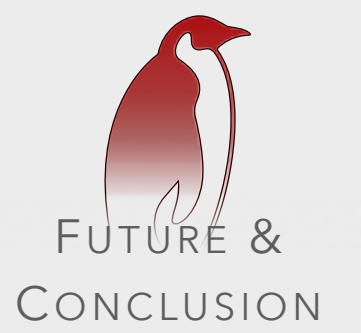
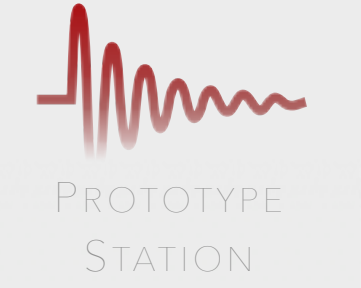
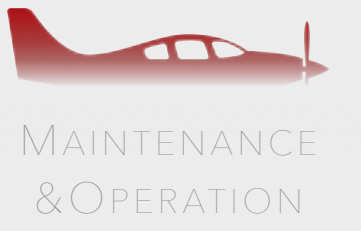
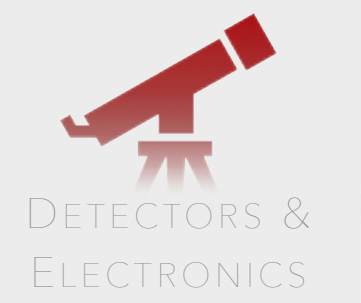
FUTURE &
CONCLUSION



Outlook in the future - IceCube Gen2

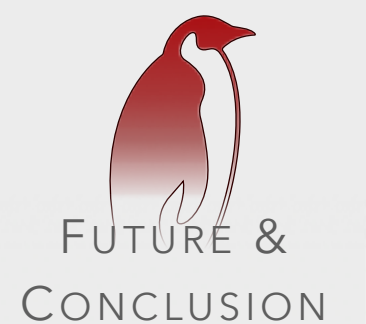
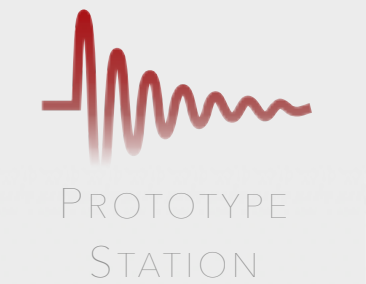
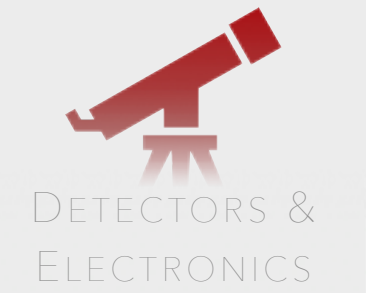


F.G. Schroeder, PoS(ICRC2019)030 (modified)



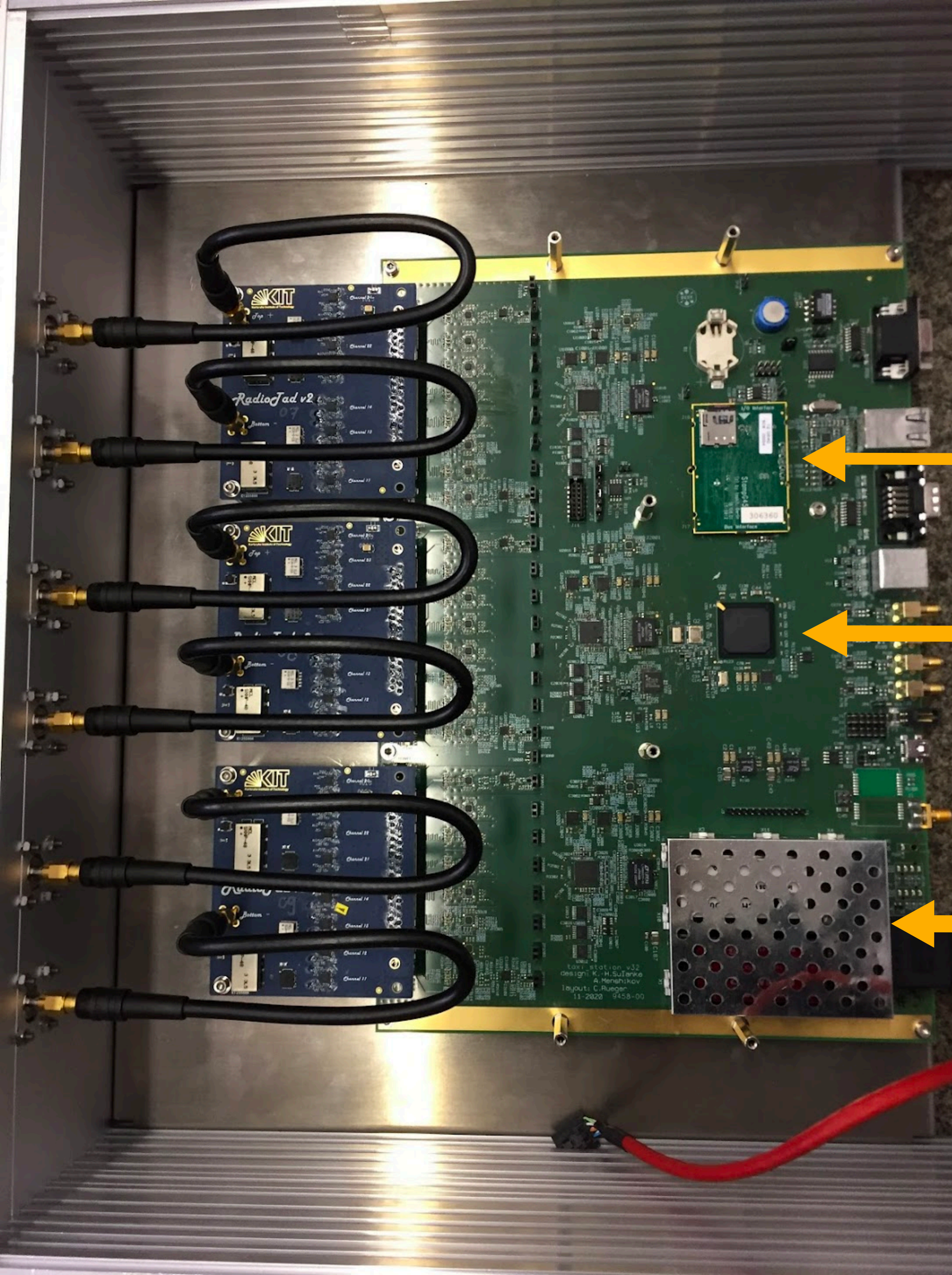
Conclusions

- We developed a **hybrid detector** combining **elevated scintillation detectors** and **elevated radio antennas** to complement the already existing ice-Cherenkov tanks.
 - This will **mitigate** the continuously **increasing detection threshold** of IceTop due to snow.
- A **prototype station** was deployed in January 2020 and as expected **records air-showers** and the **direction reconstruction agrees** between all detector types.
- The **enhanced array** covering **1 km²** planned to be **deployed in the coming years**
- **Planning and development** starting from the experience gained is on-going for the bigger **IceCube-Gen2 surface array**
- Stay tuned for more results of these future cosmic ray detectors



Back-up slides

TAXI



Embedded Linux
(Stamp)

FPGA

Shielded DC/DC
converters

SKALA LNA

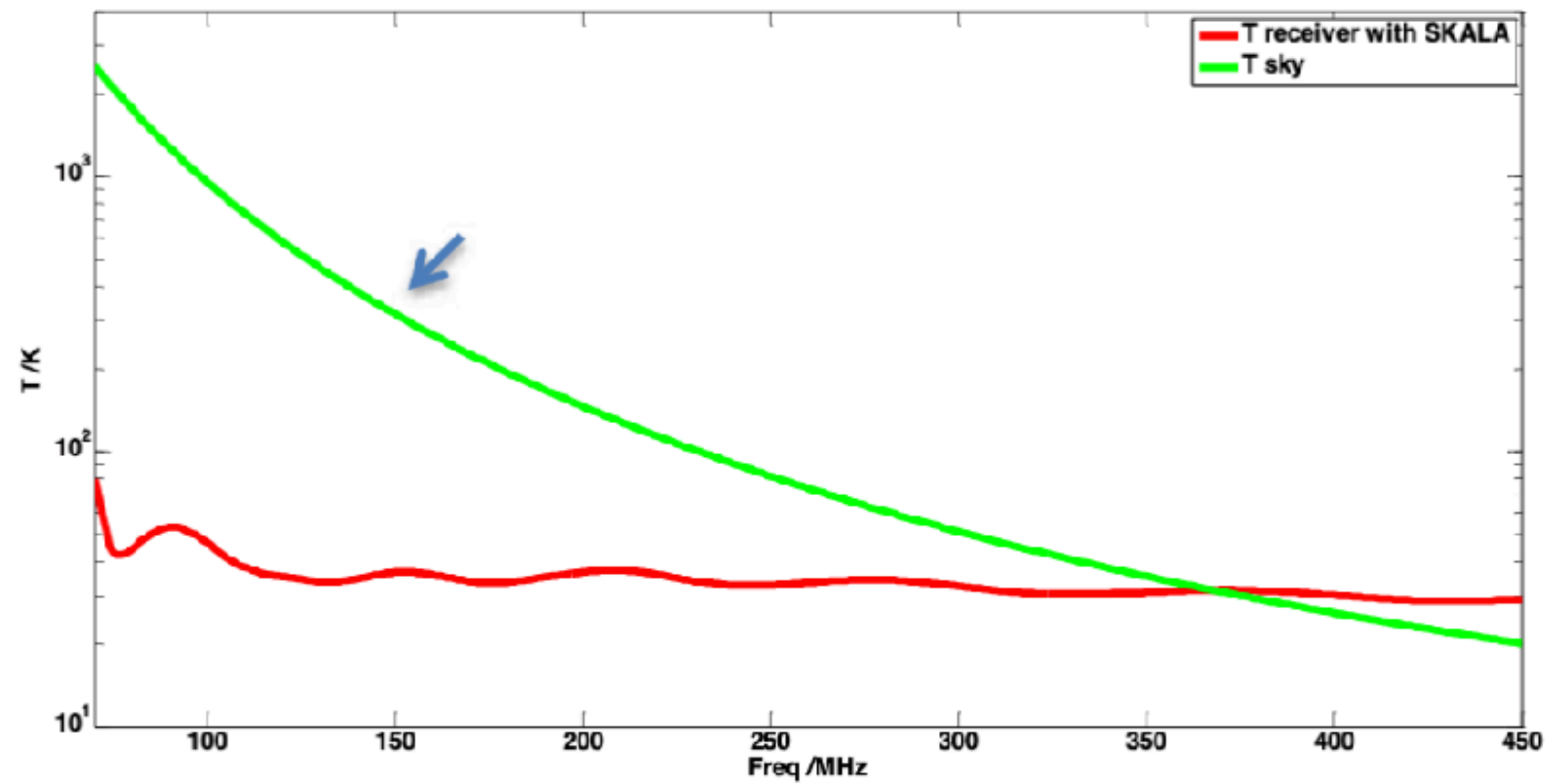
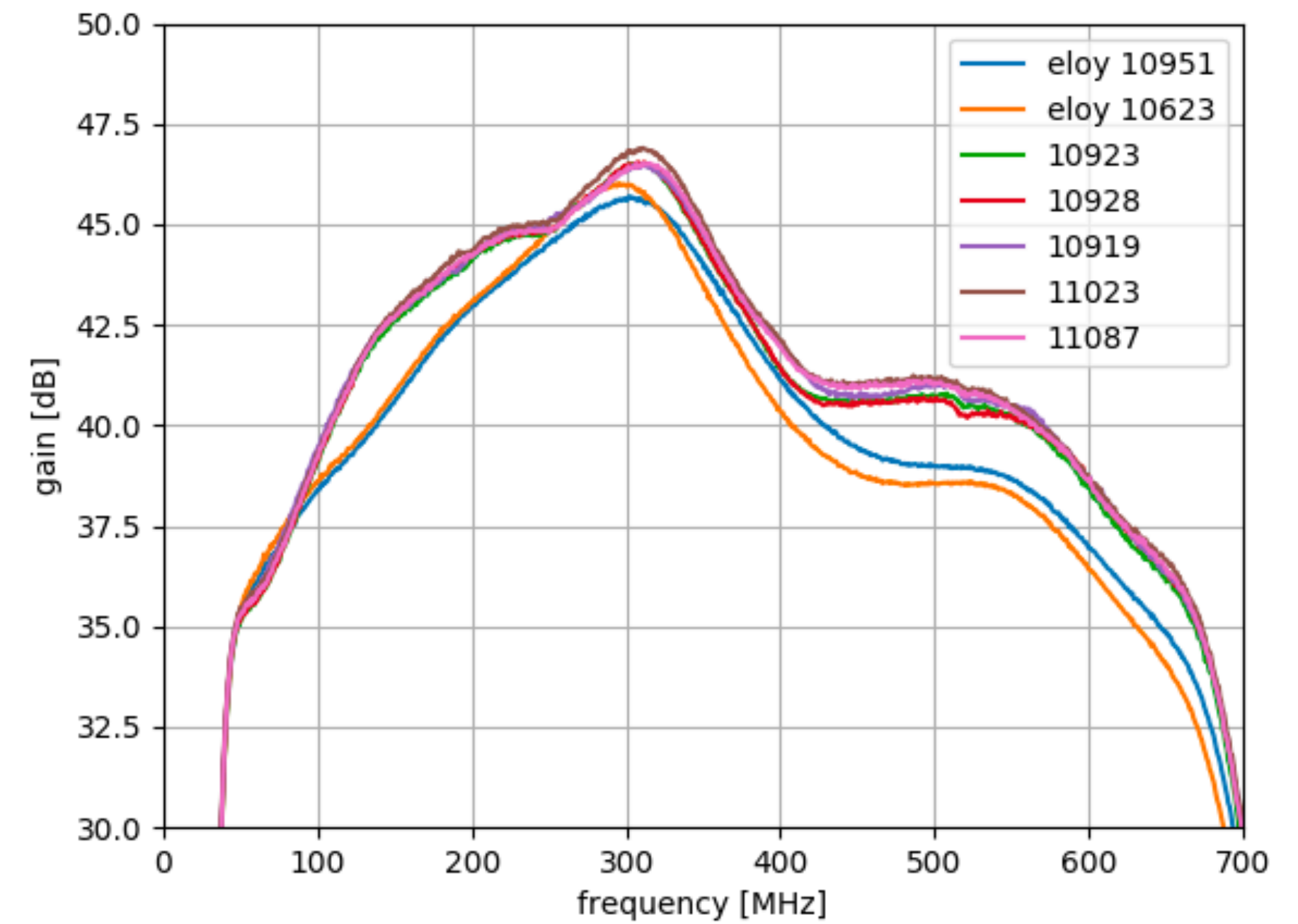


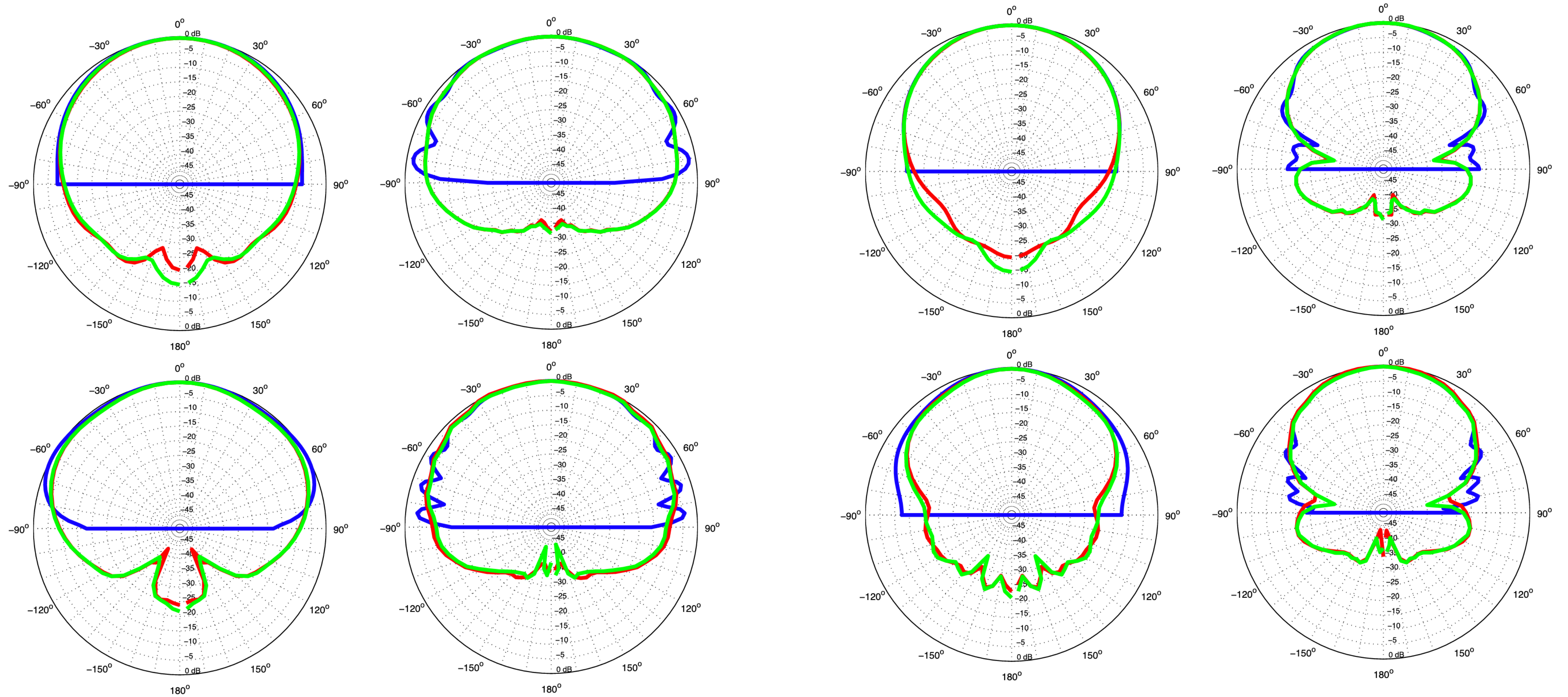
Fig. 9. Receiver noise temperature versus sky noise.



H-plane cut

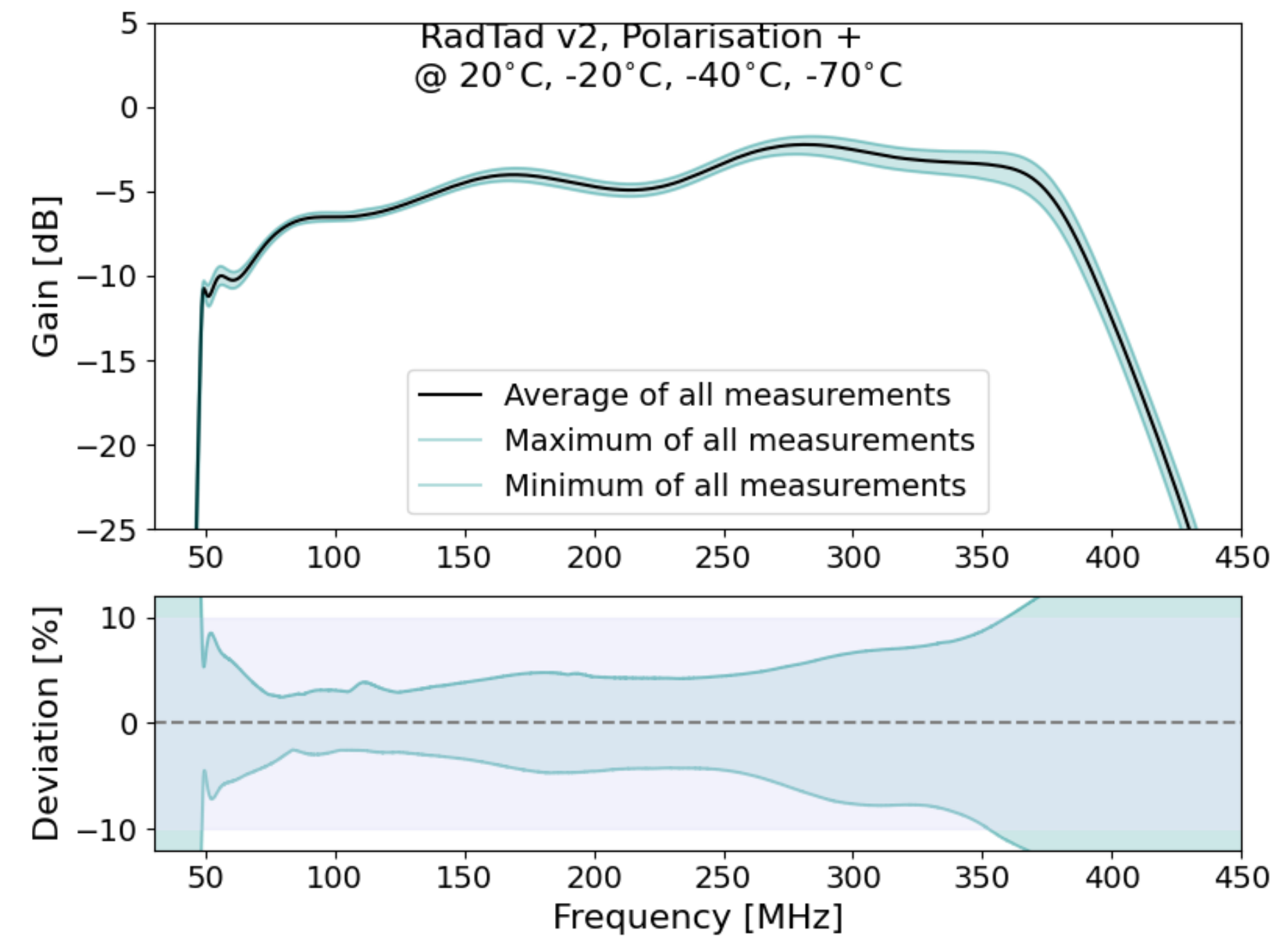
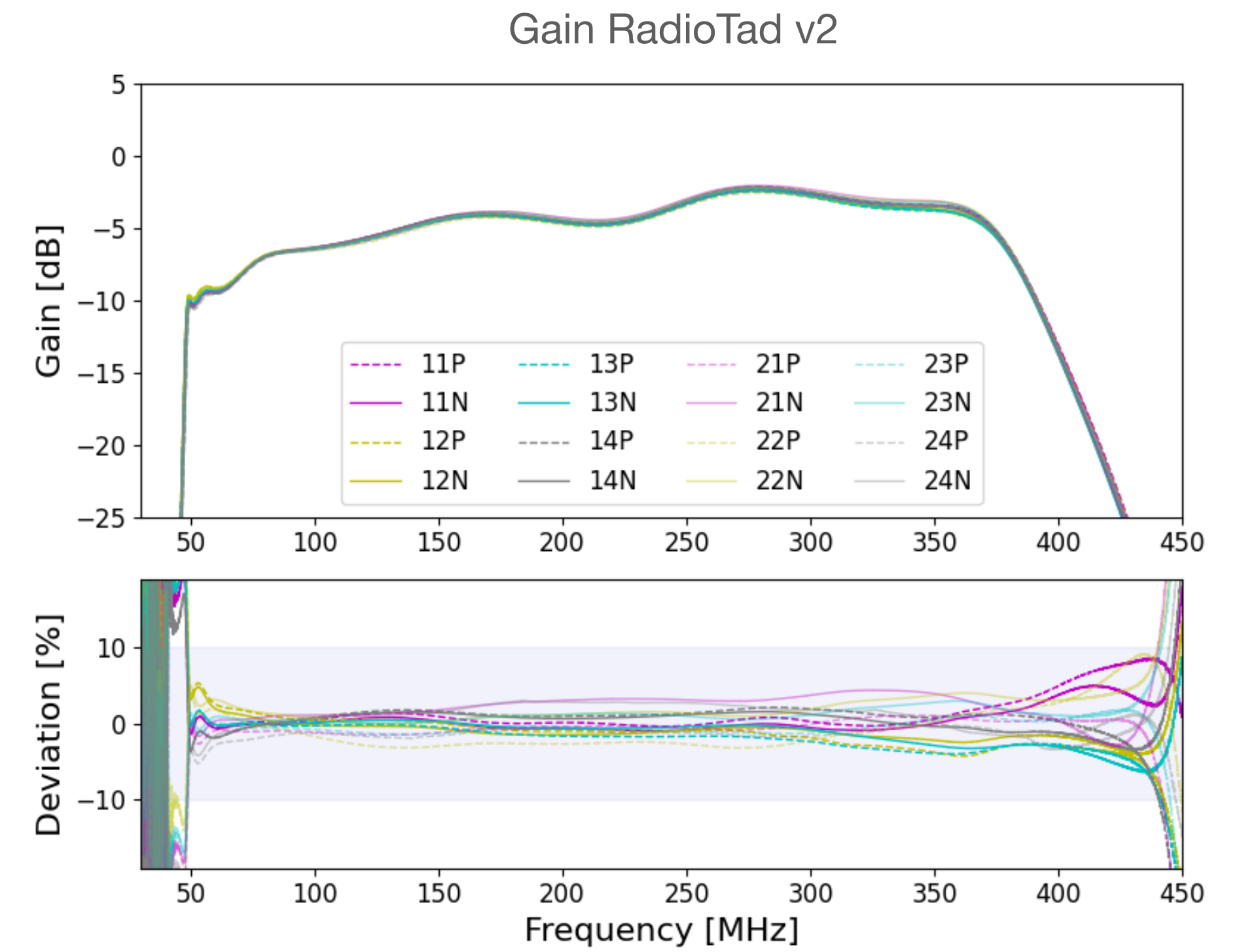
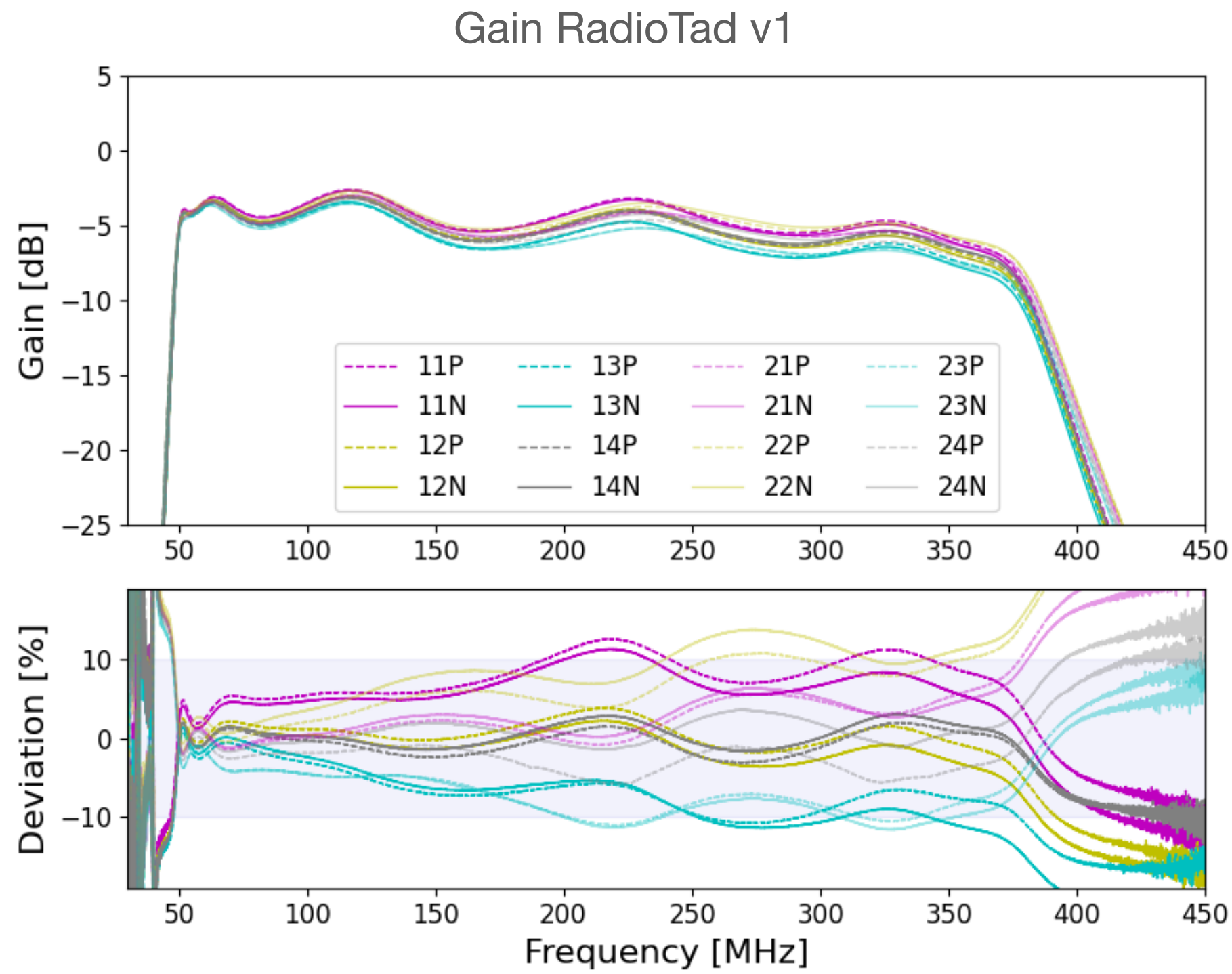
E-plane cut

SKALA2 Gain



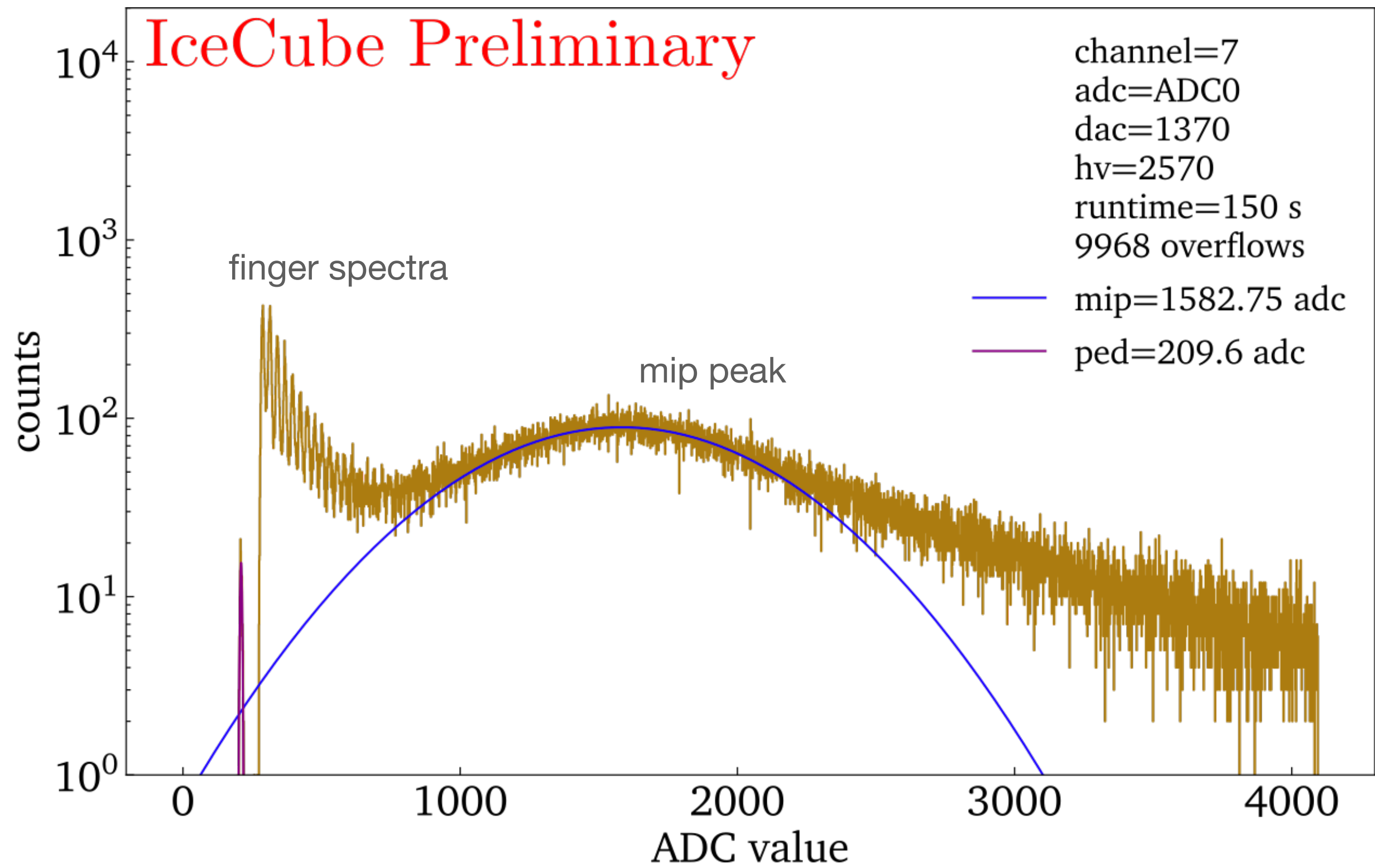
Inf. GND ■ Soil ■ Mesh over Soil ■

RadioTad

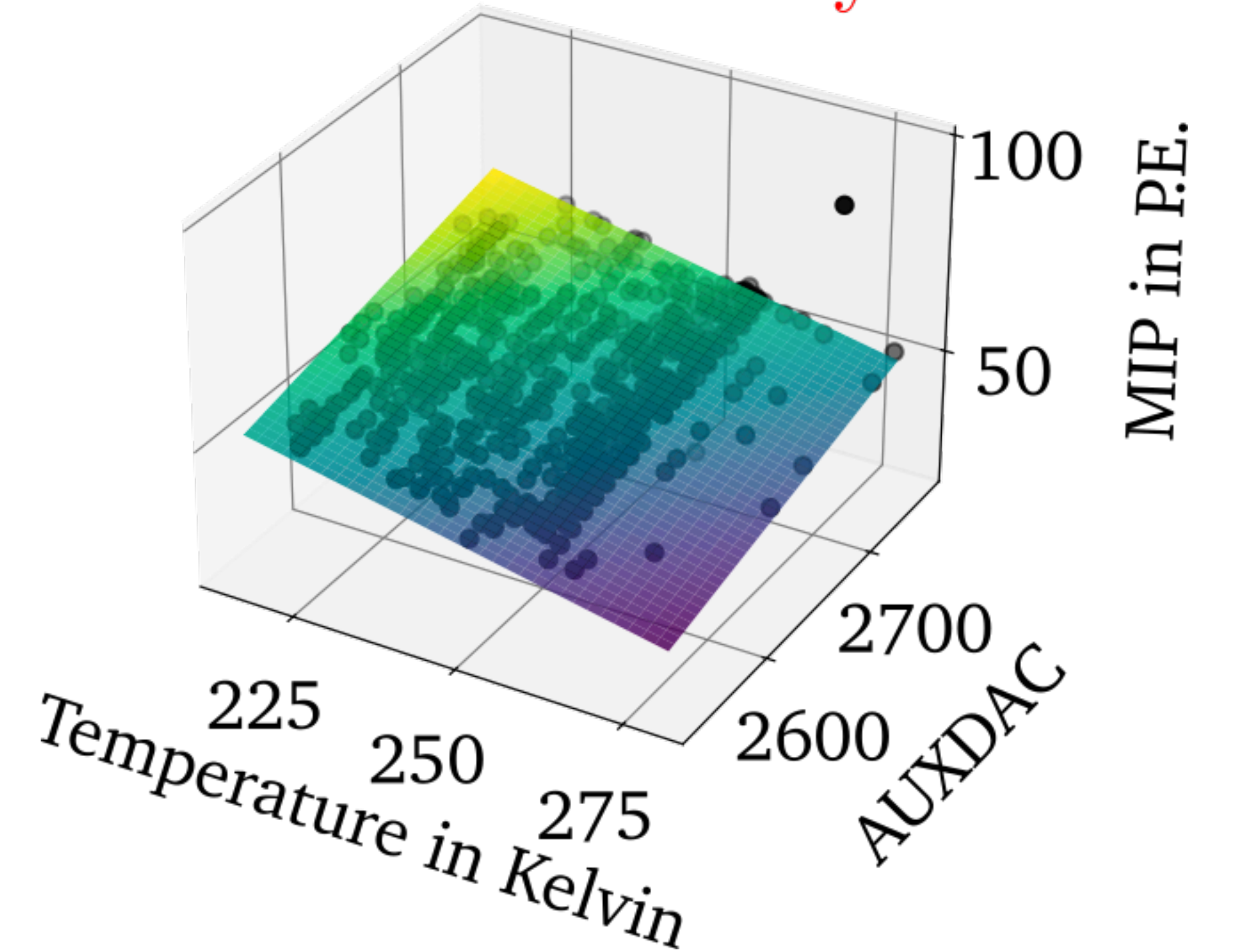


Scintillators

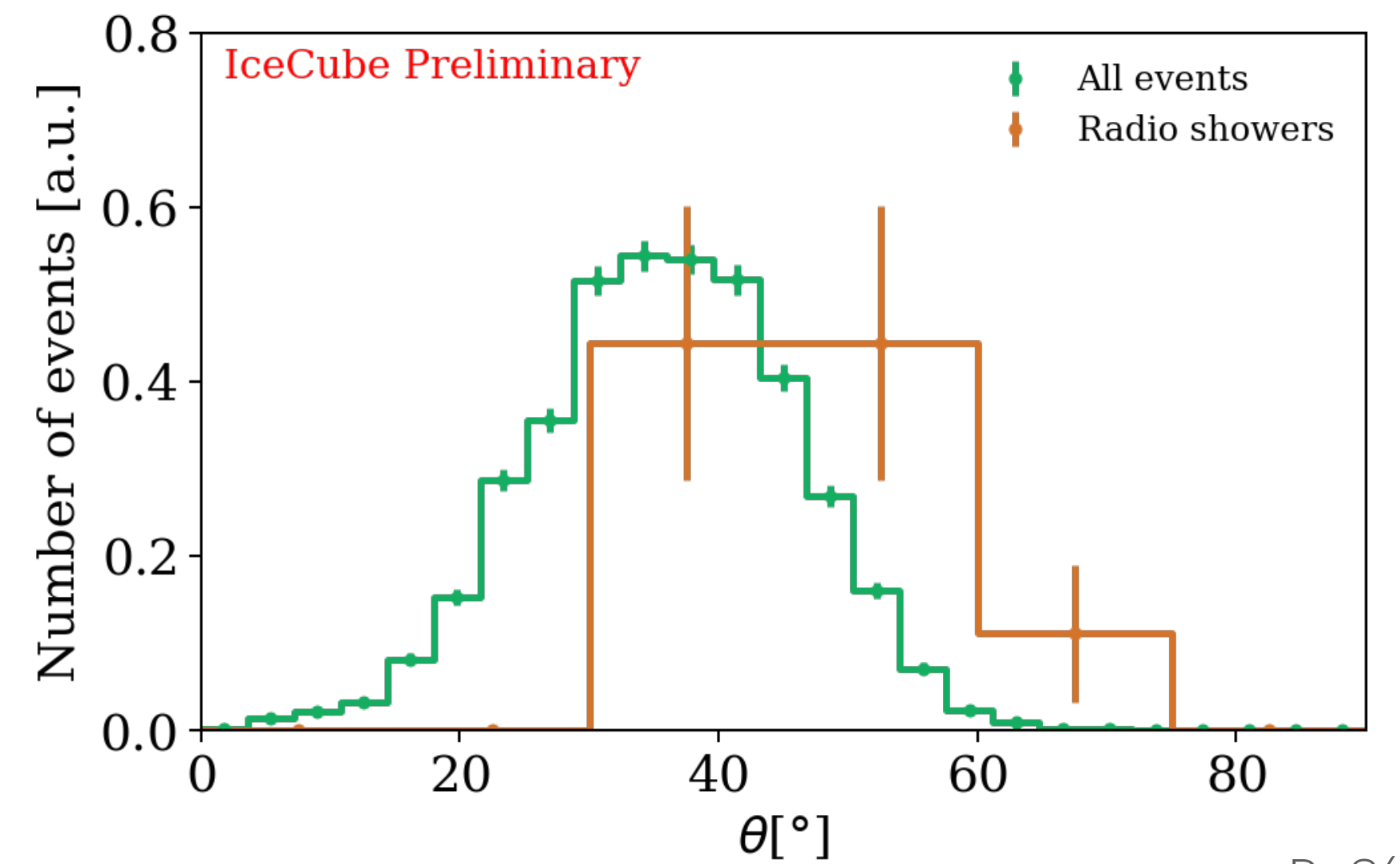
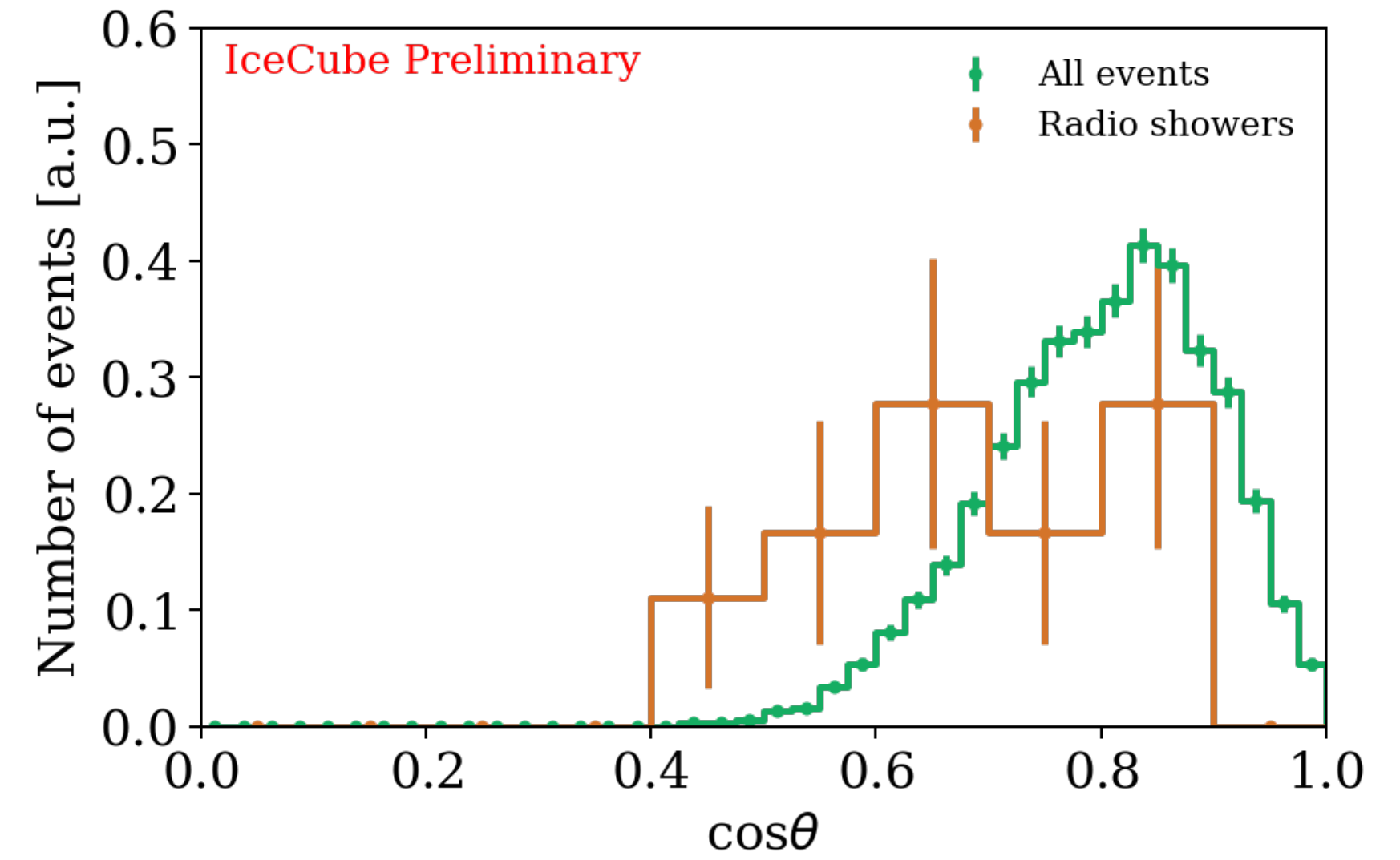
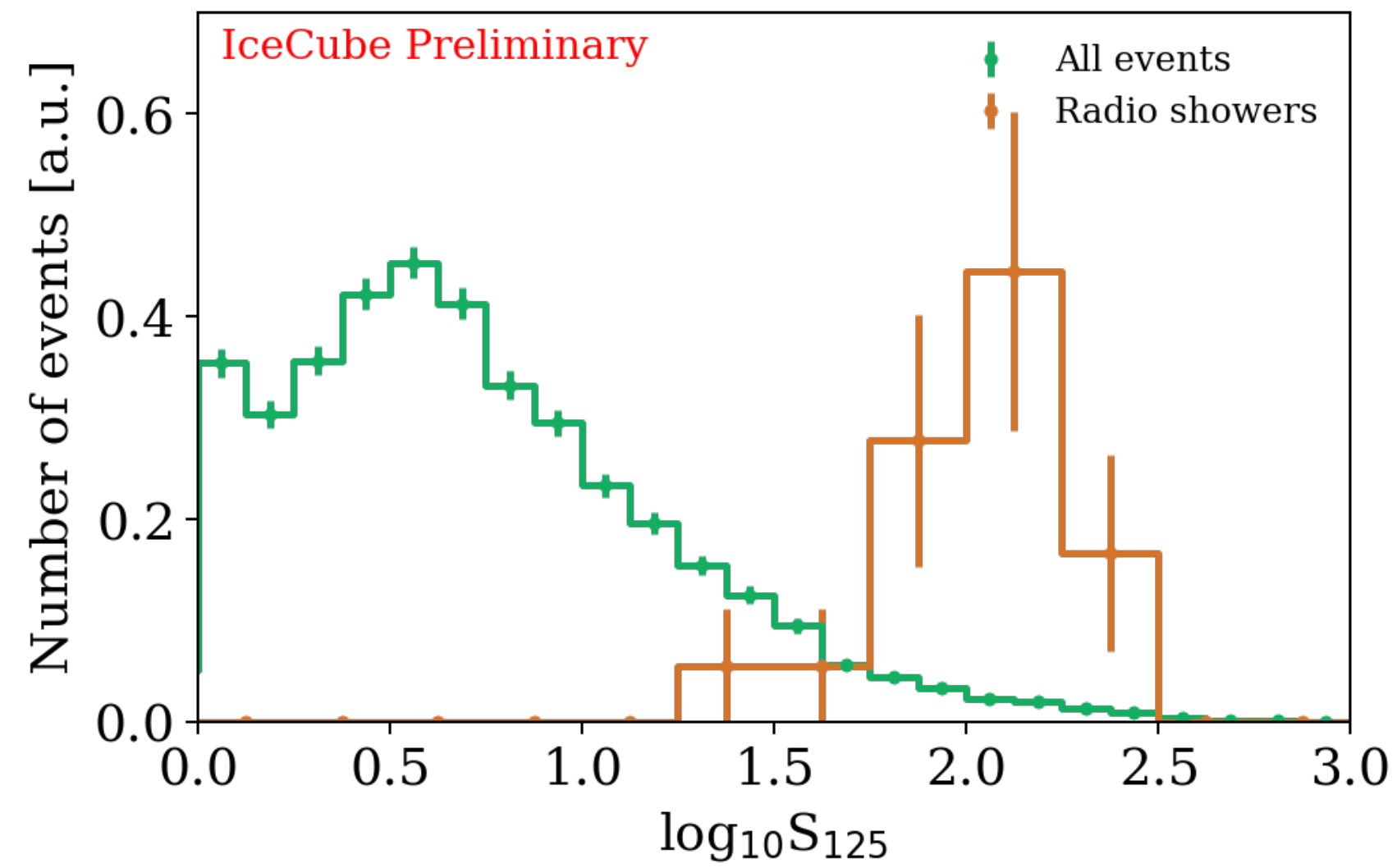
Temperature and voltage versus gain. The gain can be maintained constant by varying the voltage according to the temperature

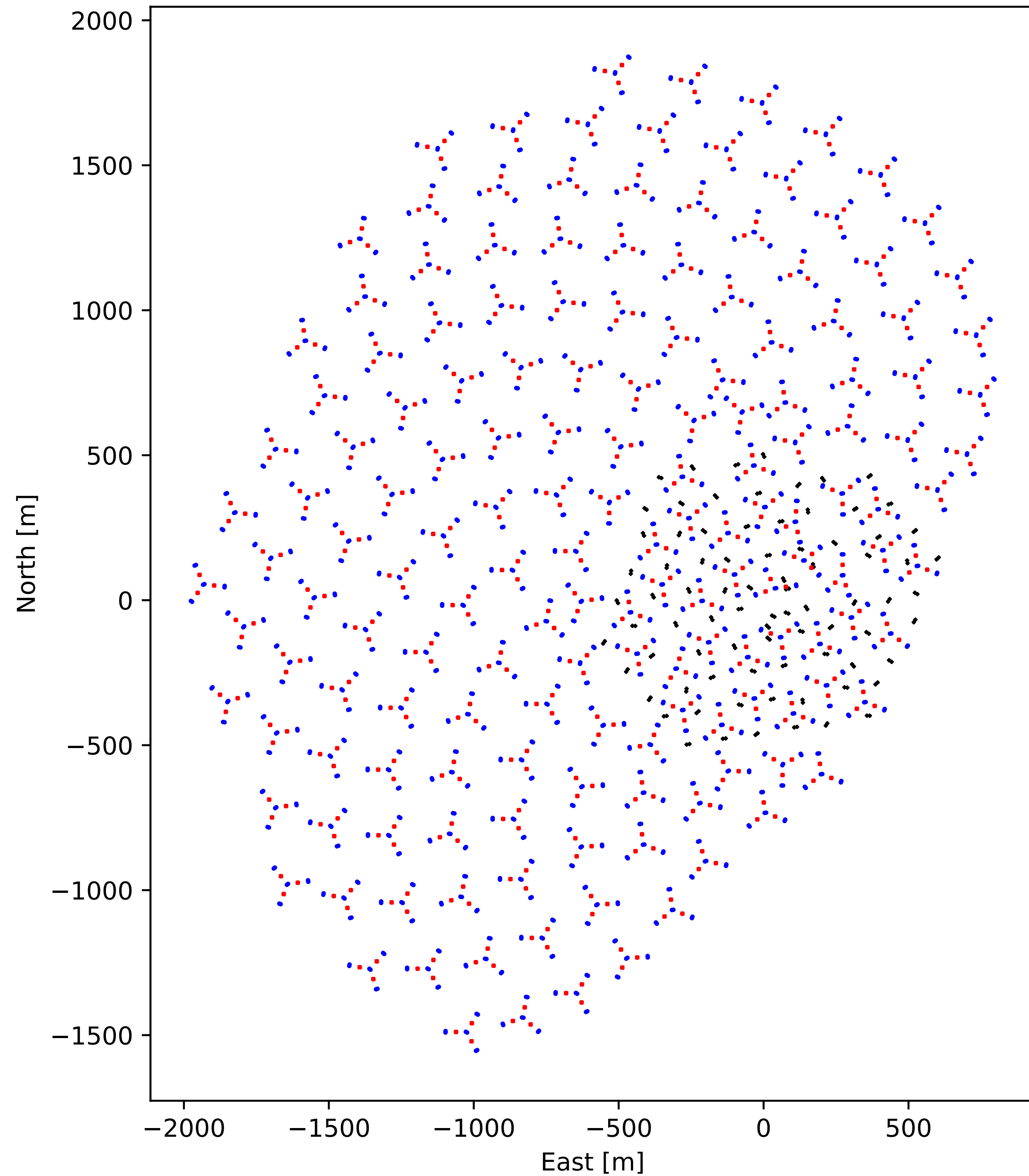
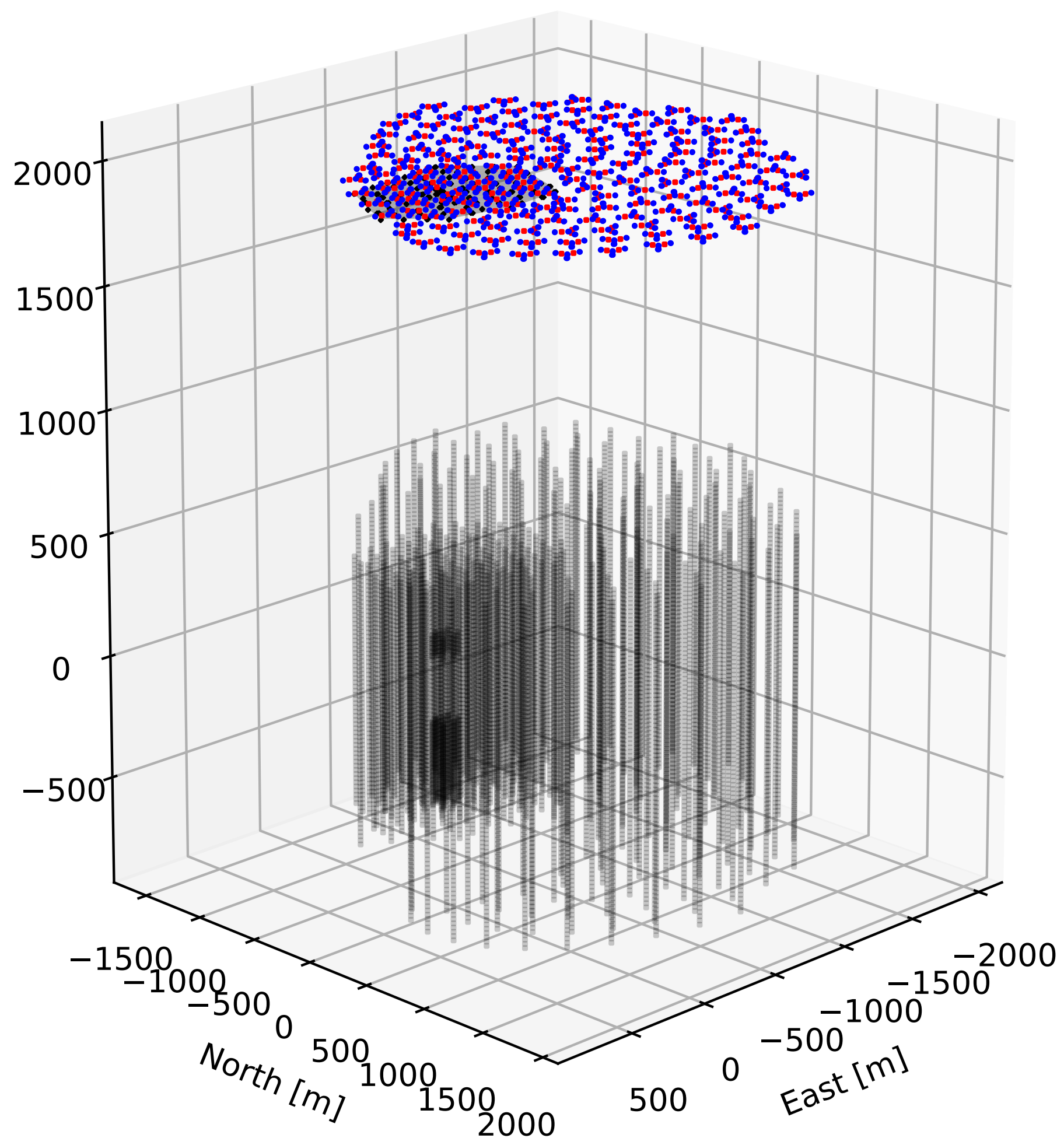


IceCube Preliminary



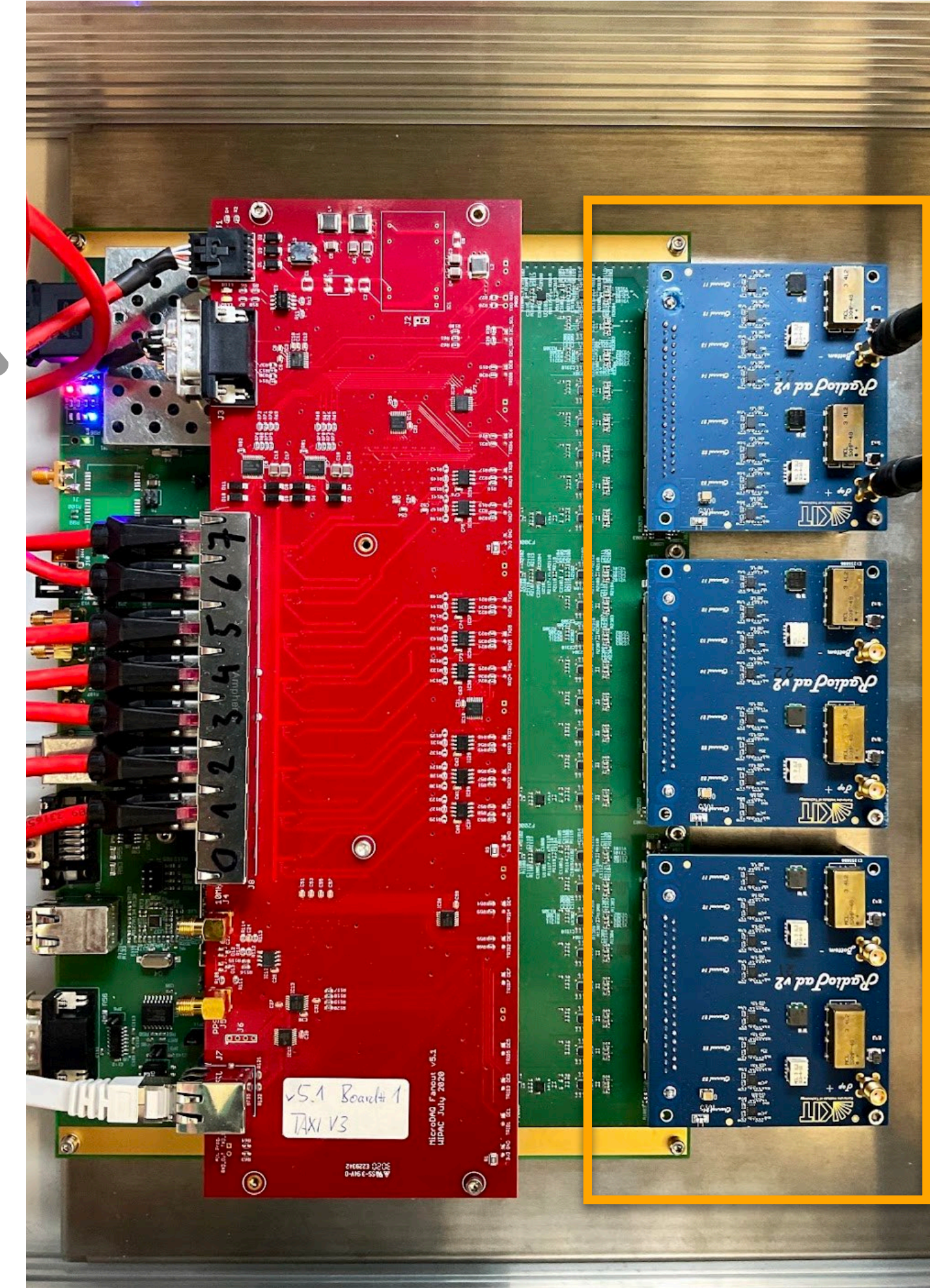
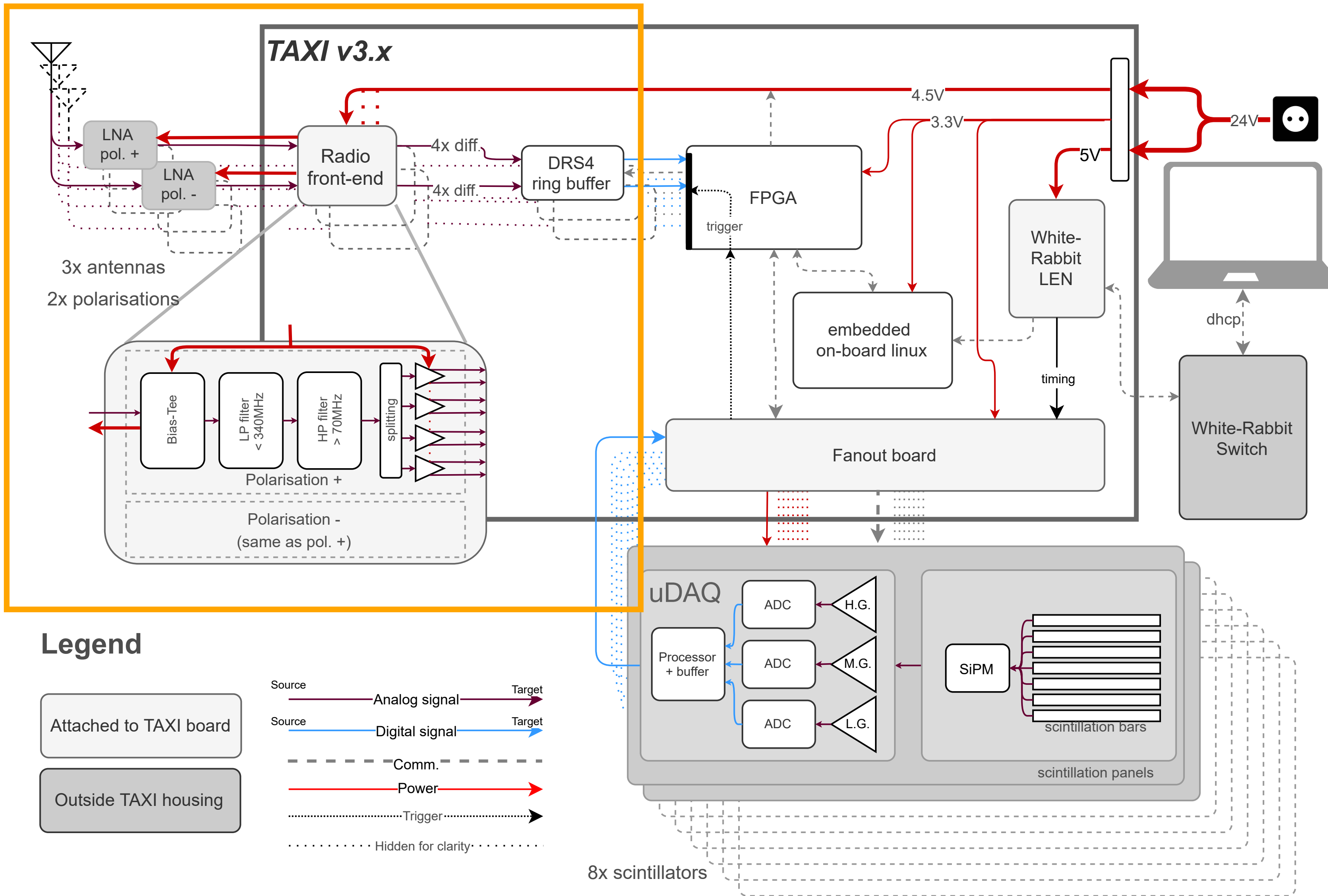
Air showers reconstruction





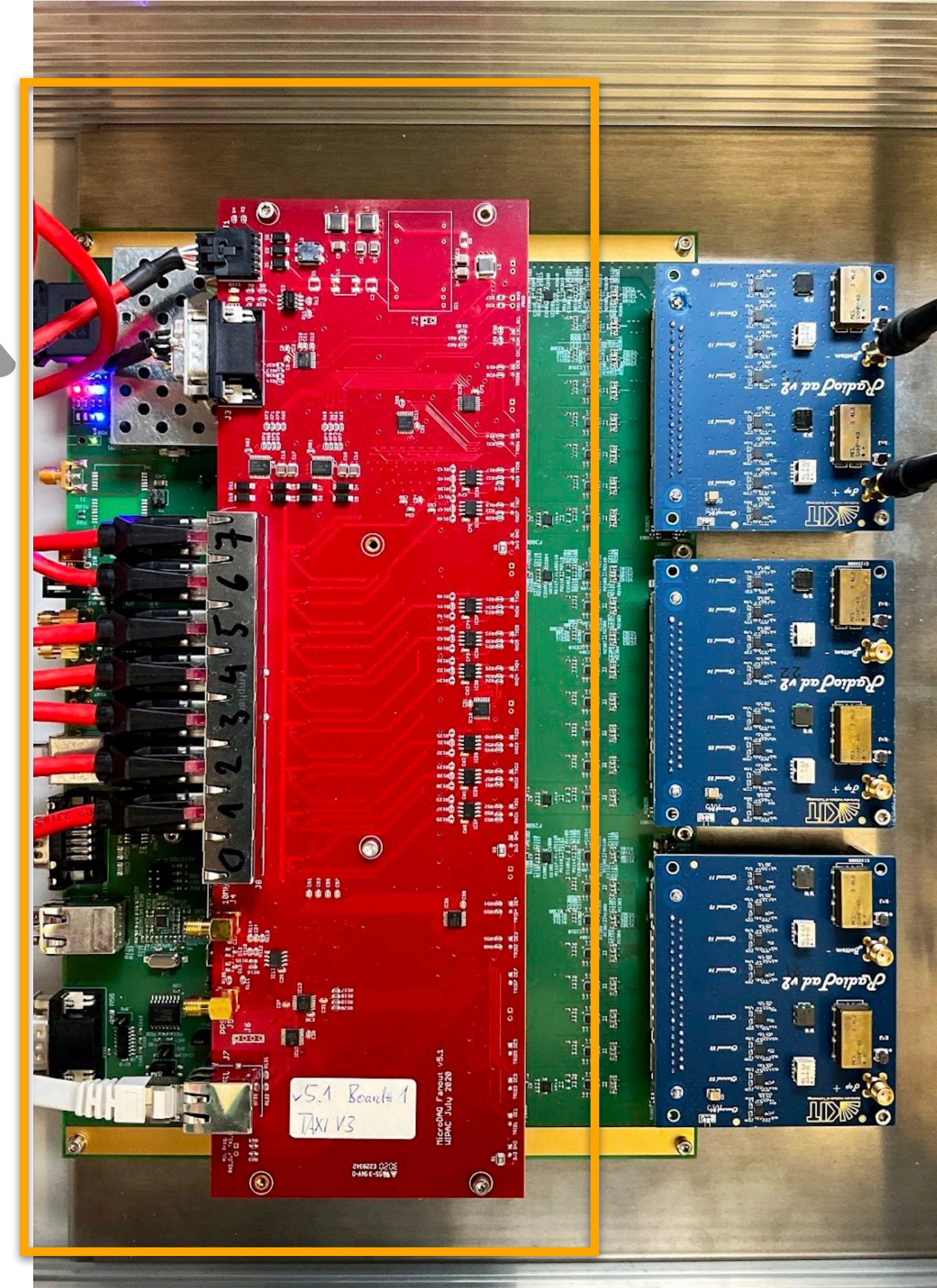
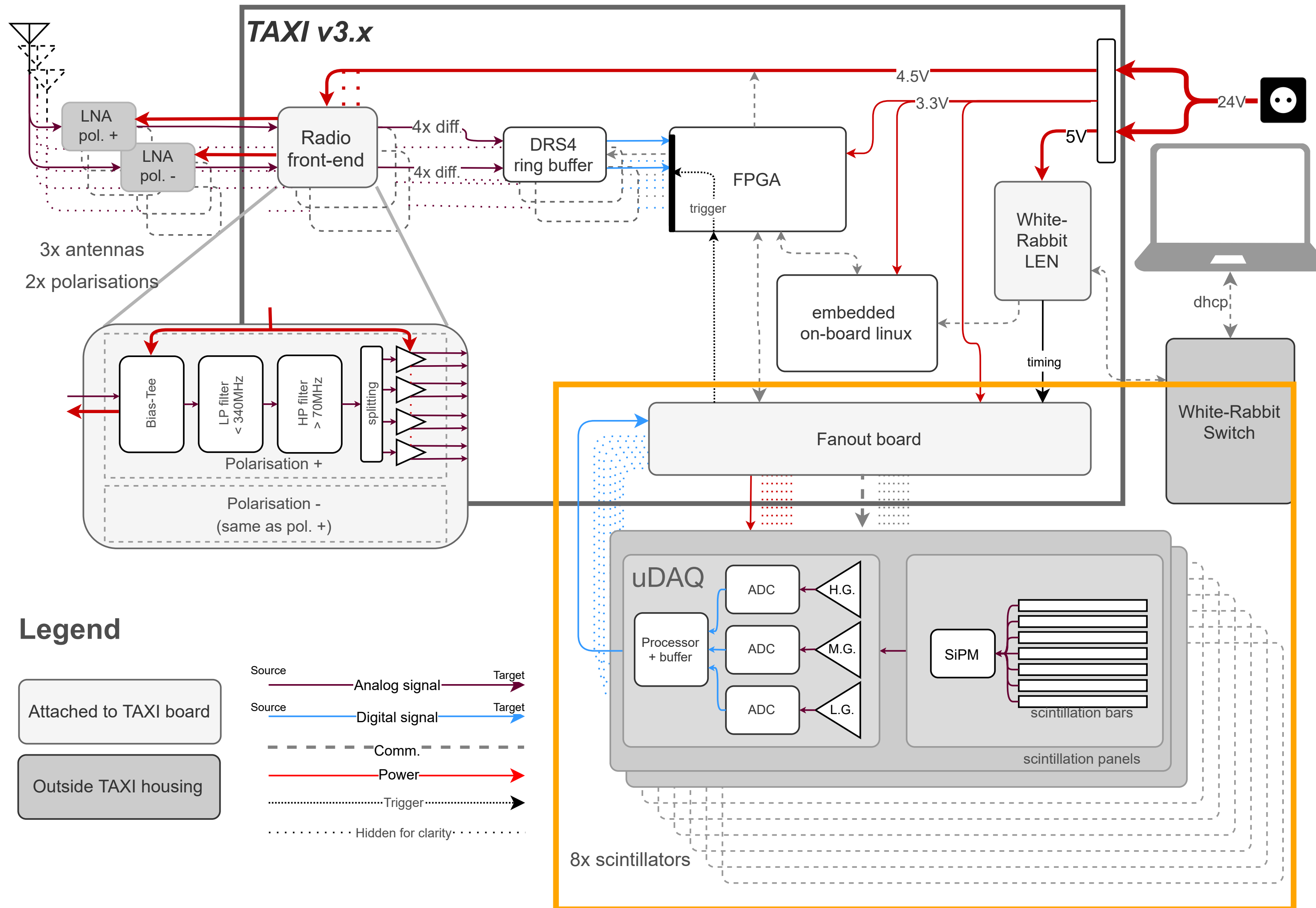
Radio components

DAQ and related components



Scintillator components

DAQ and related components



Communication, control and timing

DAQ and related components

