



UNIVERSITY OF  
COPENHAGEN

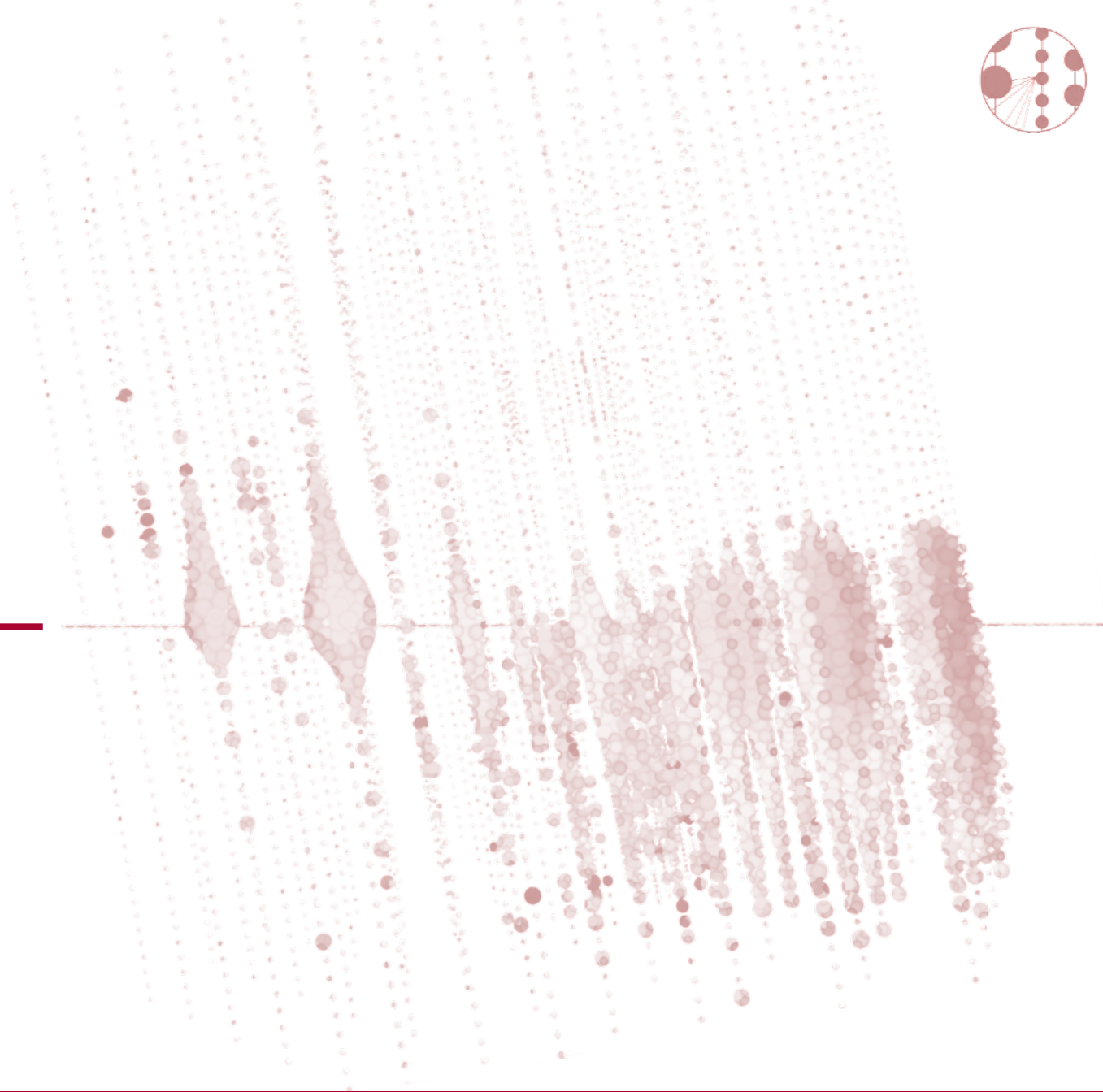


# Neutrino Oscillations with IceCube & the IceCube Upgrade

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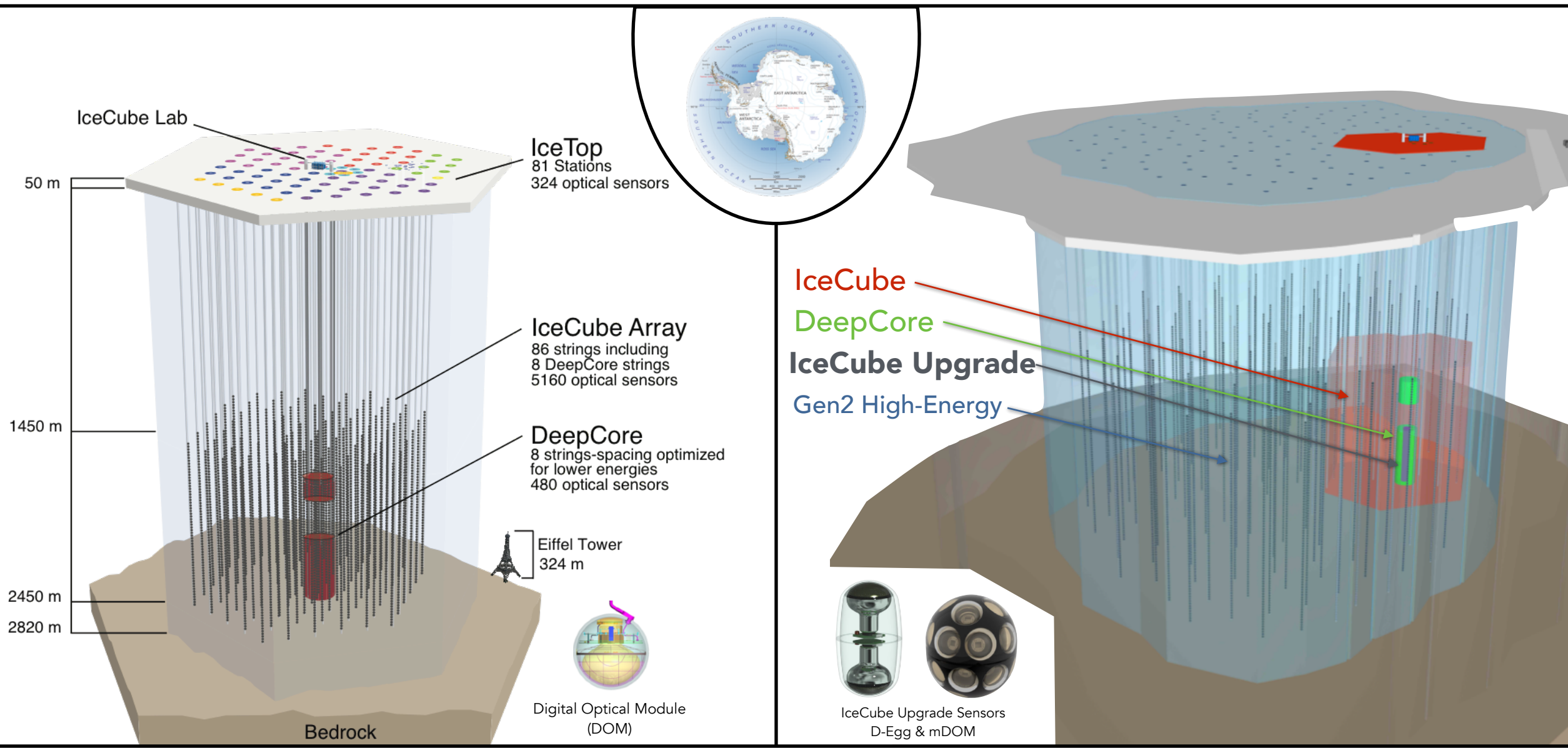
D. Jason Koskinen - Niels Bohr Institute  
for the IceCube Collaboration

TeV Particle Astrophysics (TeVPA 2022)  
August 2022





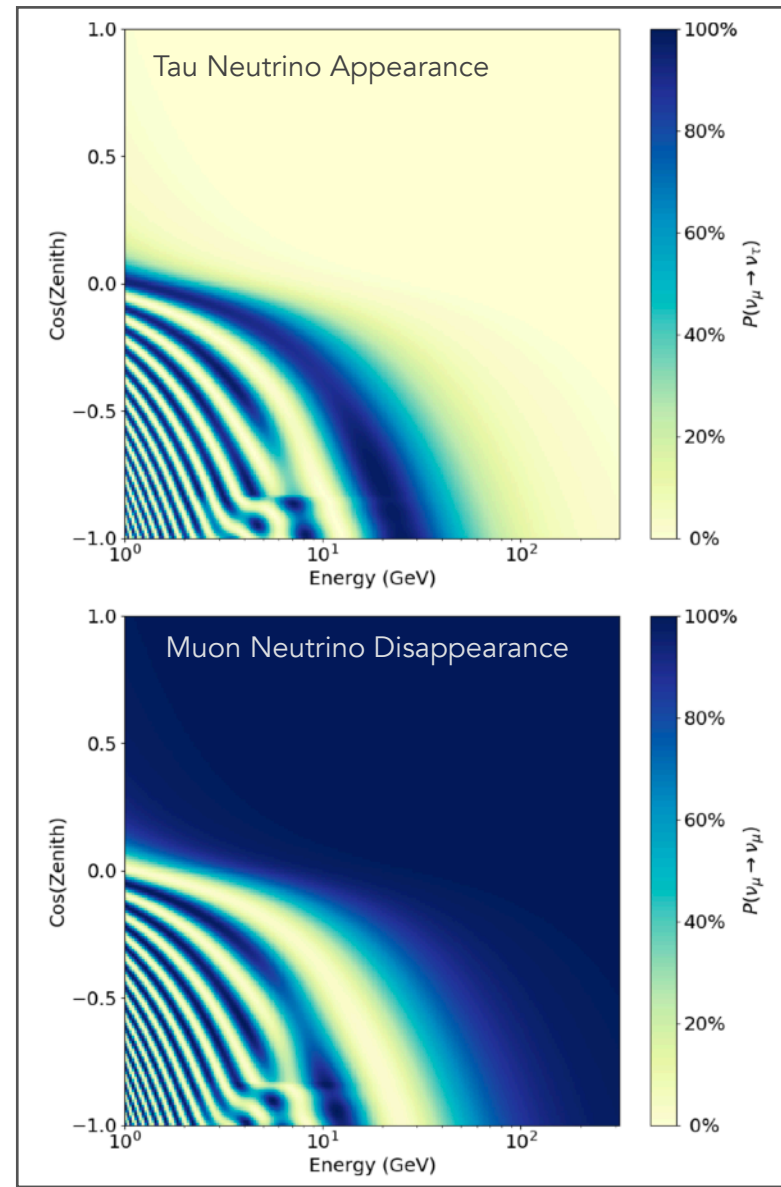
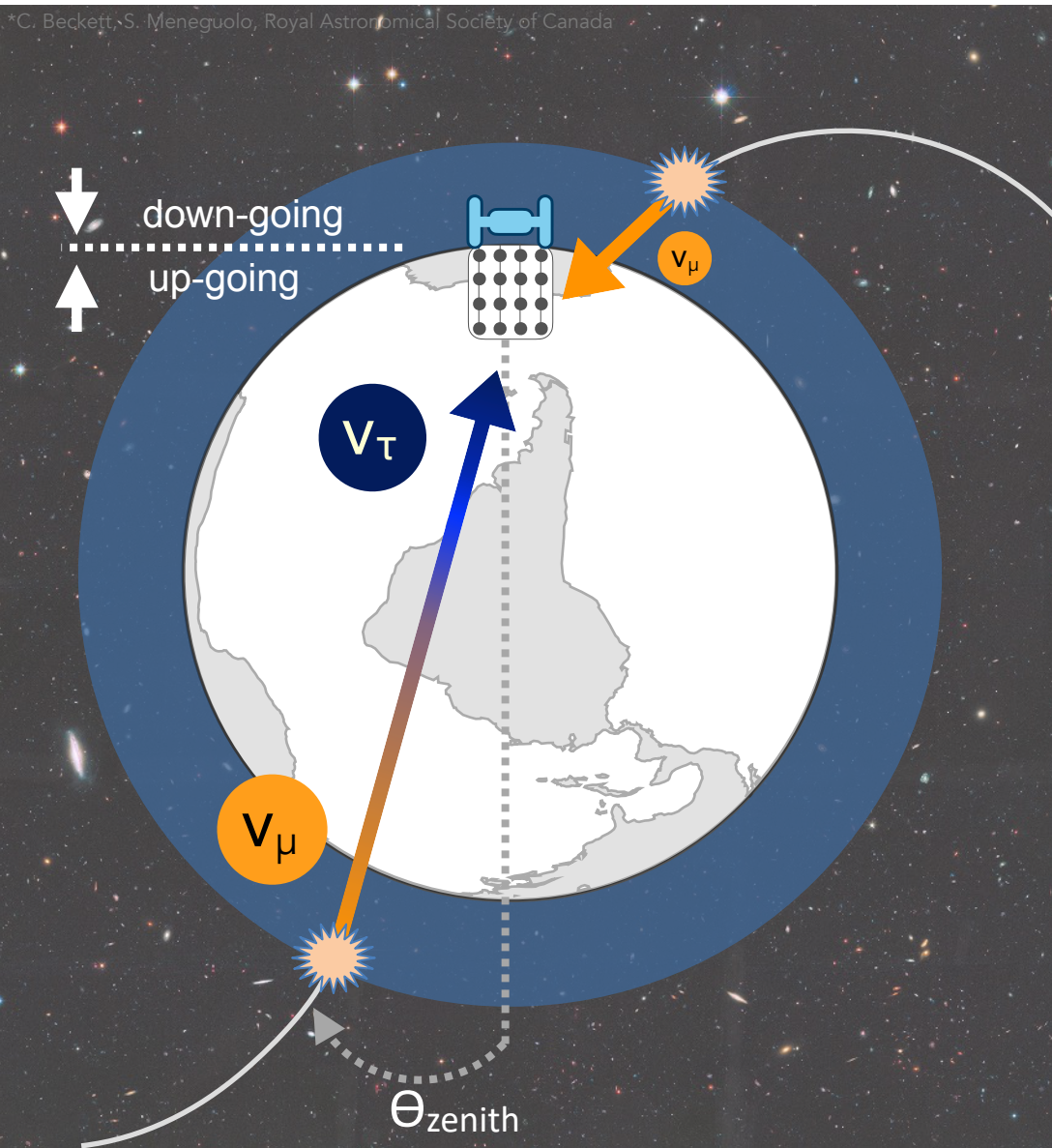
# Detectors



# Why IceCube/DeepCore/Upgrade?

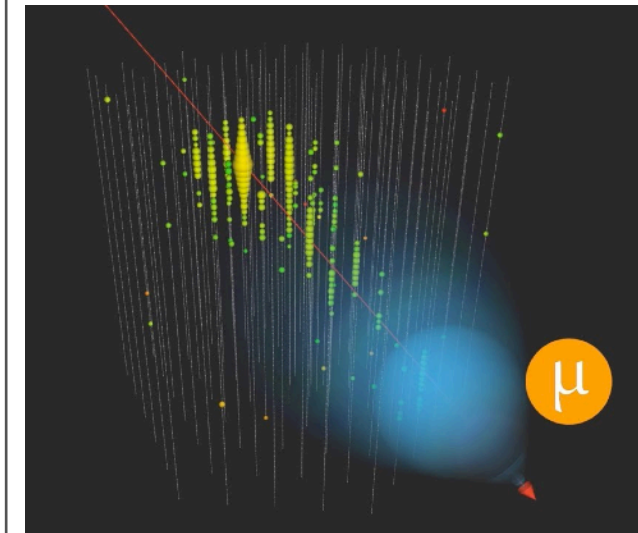


\*C. Beckett, S. Meneguolo, Royal Astronomical Society of Canada



Detection via Cherenkov emission from products of  $\nu - N$  interactions

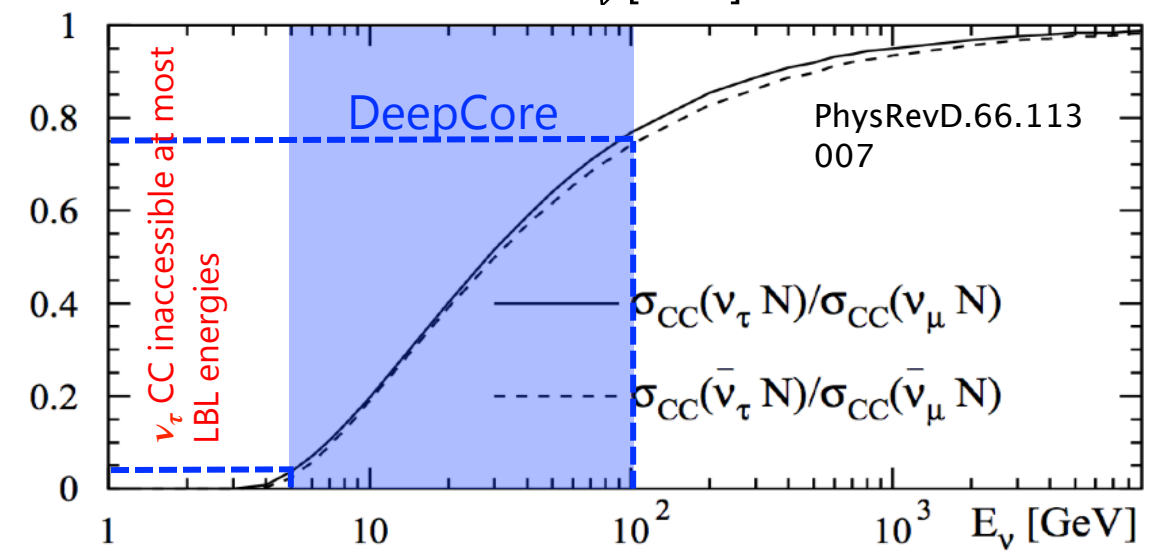
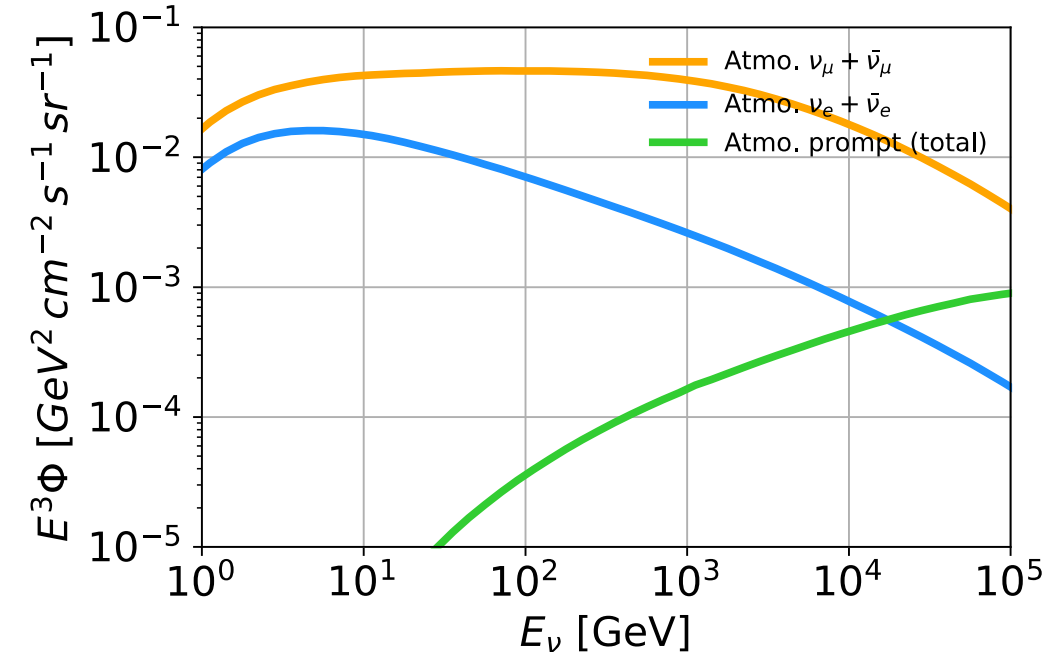
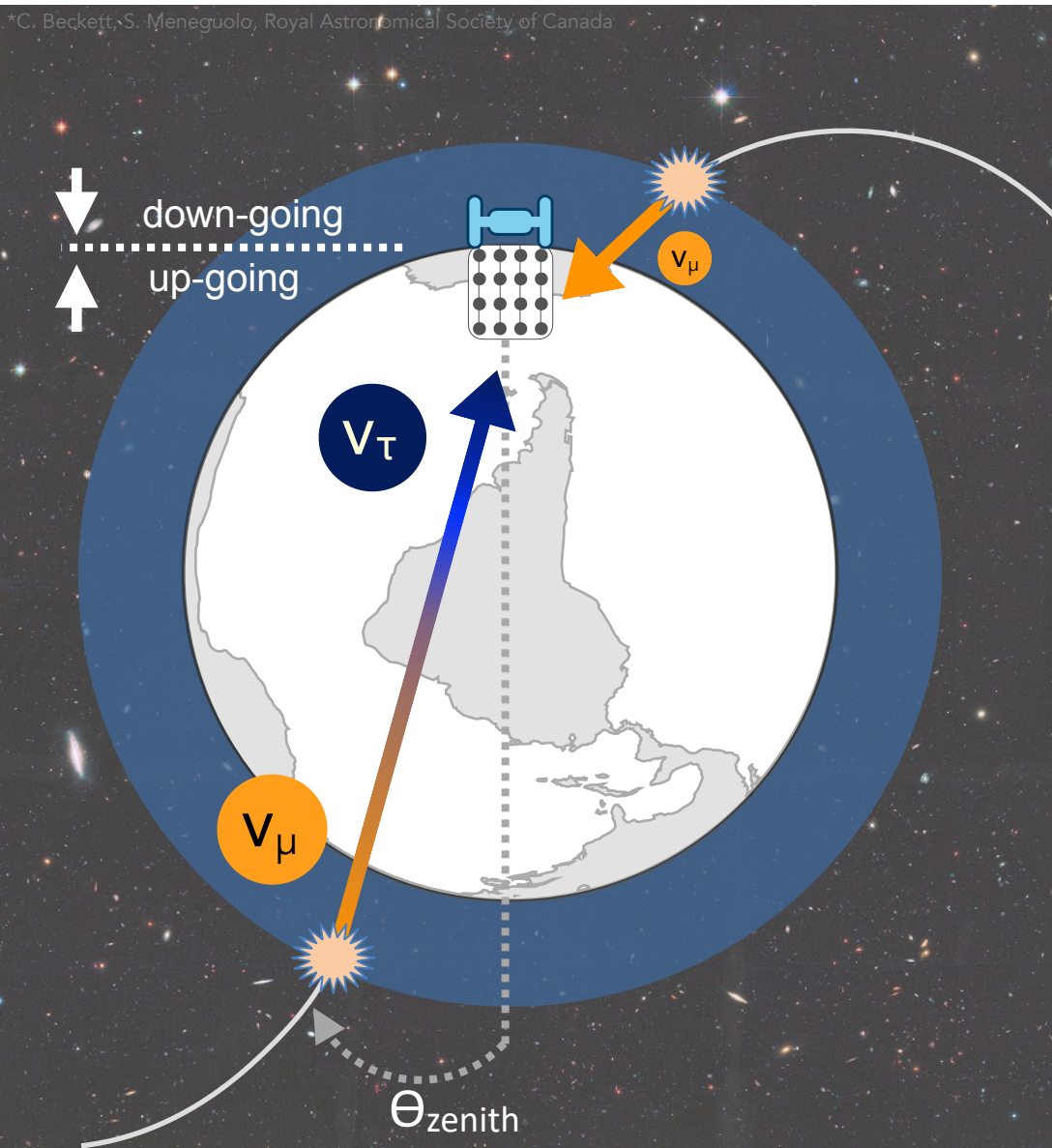
- Predominantly Deep Inelastic Scattering (DIS)
- "Tracks" from secondary  $\mu$
- "Cascades" from secondary  $e, \tau$ , and hadrons



# Why IceCube/DeepCore/Upgrade?



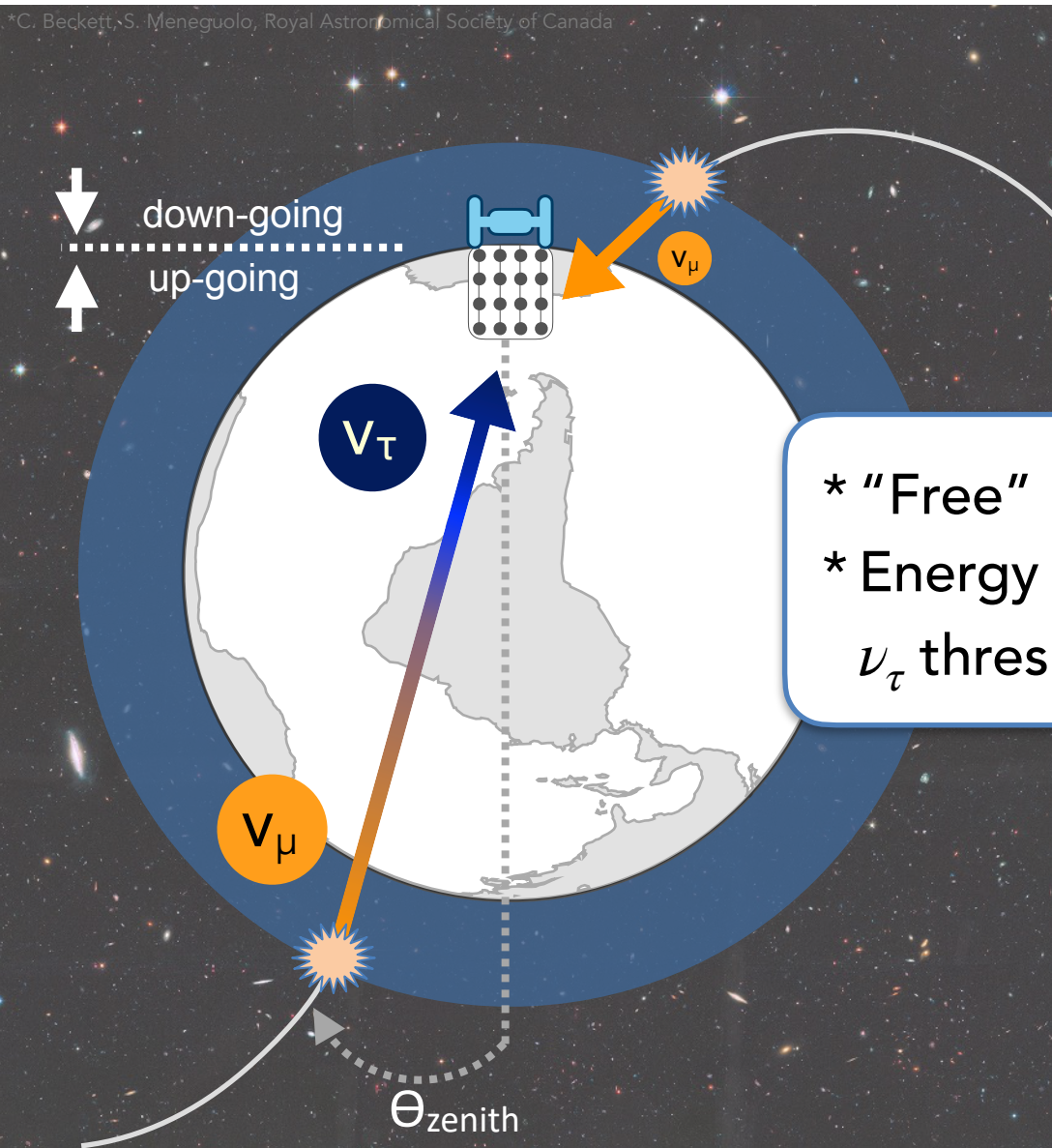
\*C. Beckett, S. Meneguolo, Royal Astronomical Society of Canada



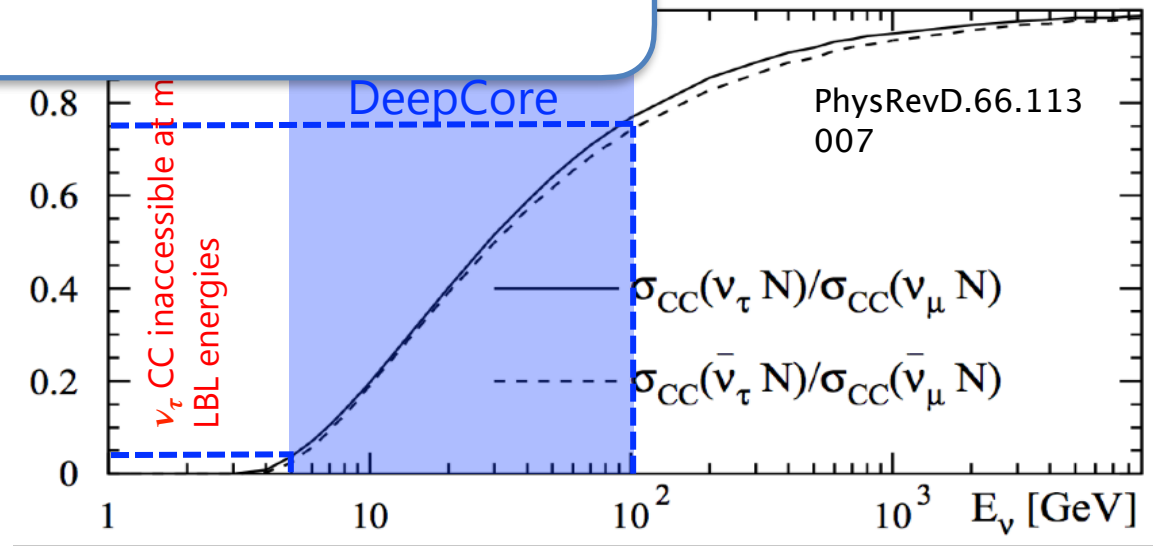
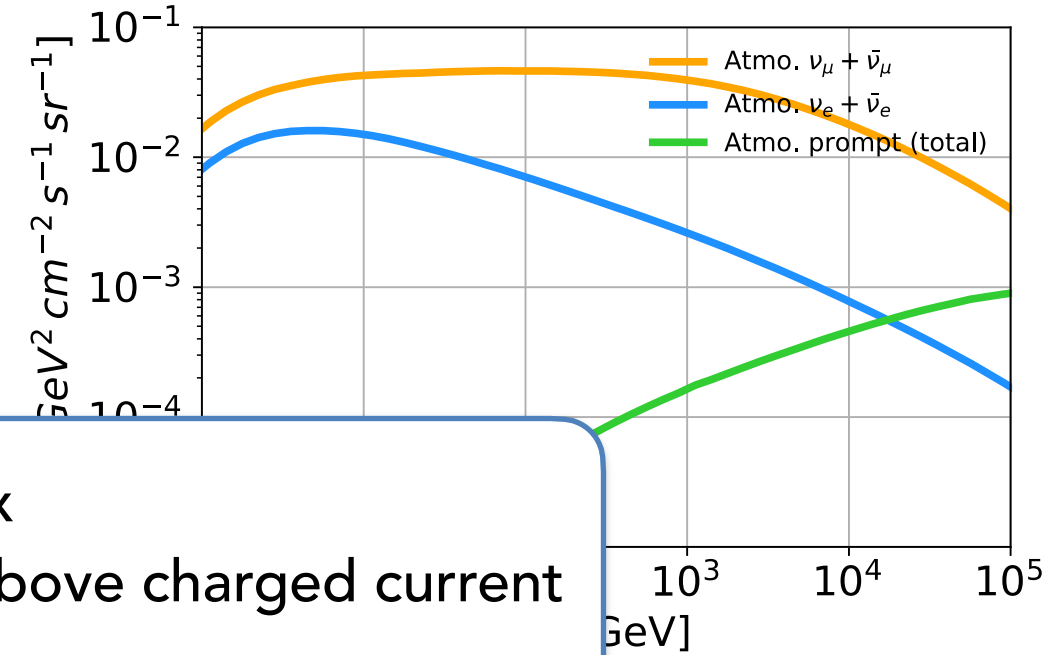
# Why IceCube/DeepCore/Upgrade?



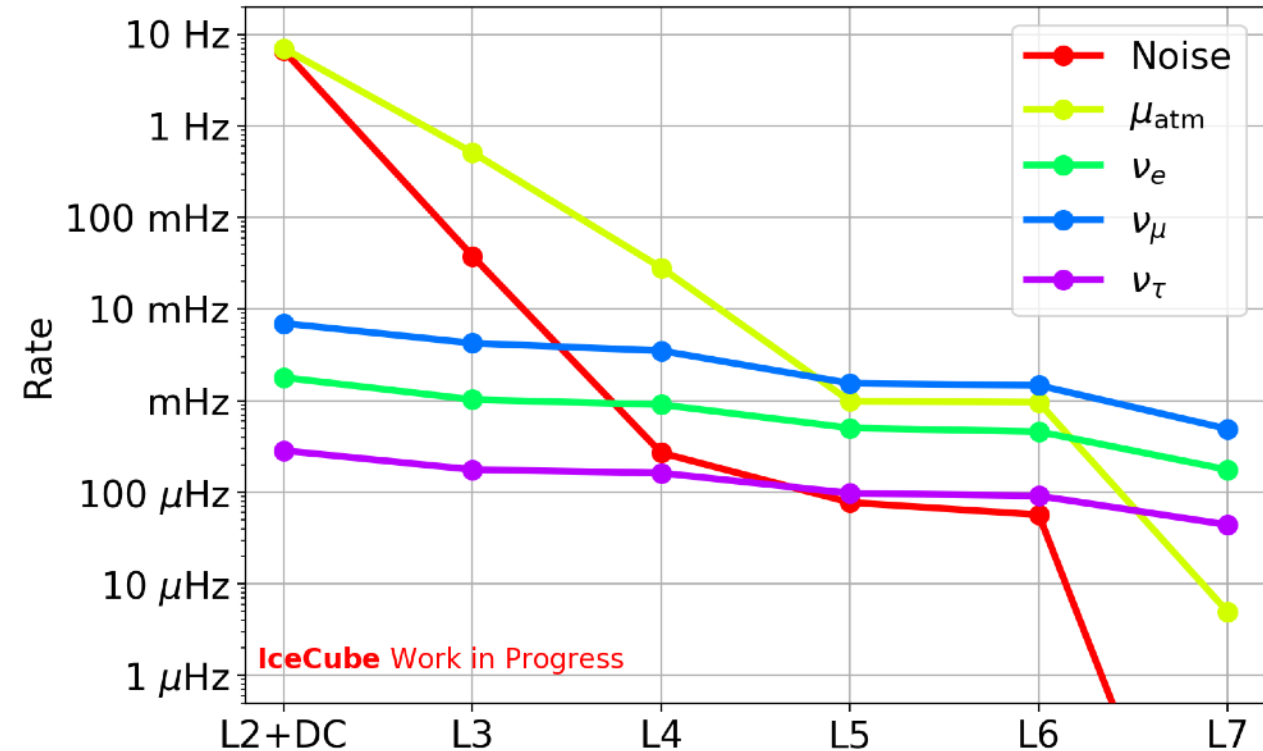
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\* "Free" neutrino flux  
 \* Energy sensitivity above charged current  $\nu_\tau$  threshold



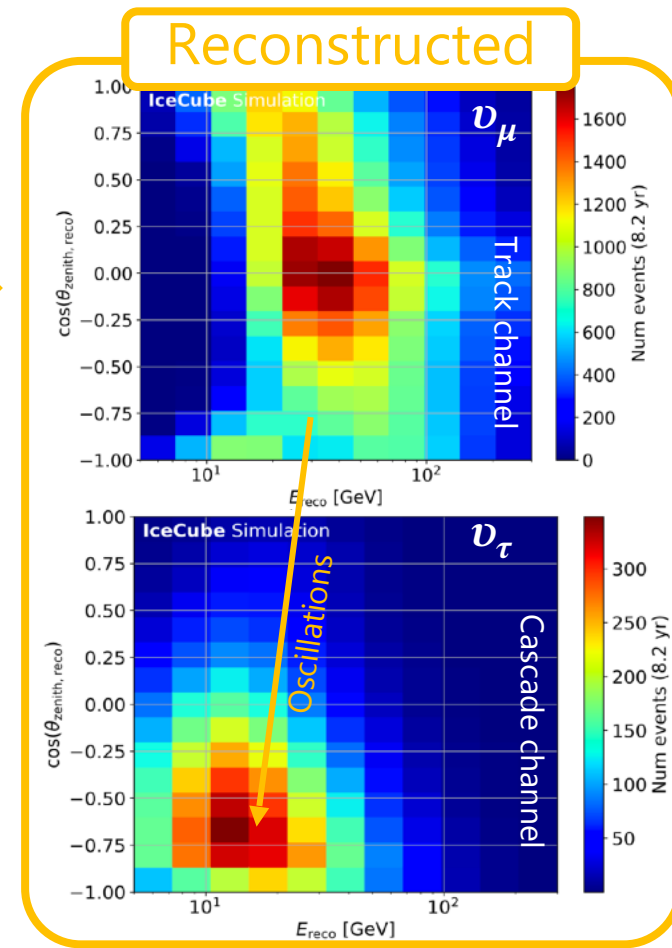
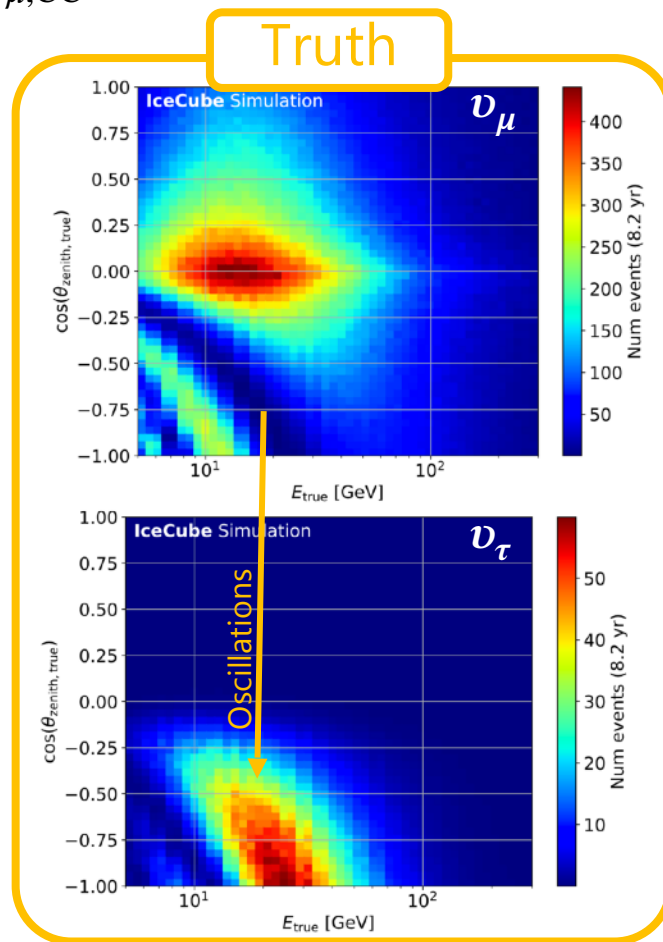
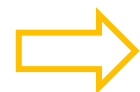
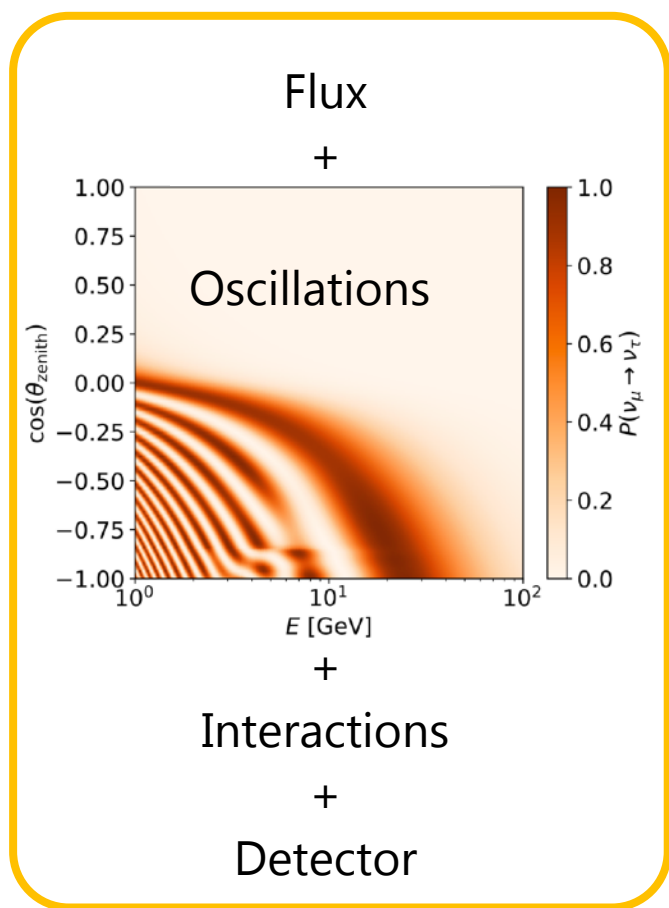
- Iterative approach to event selection
  - Backgrounds reduced by 6+ orders of magnitude\*
    - Detector/electronics pure-noise
    - Atmospheric muons
  - Combination of straight cuts and boosted decision trees
  - Due to high event rates, the high fidelity reconstructions\*\* are introduced near end of the selection chain



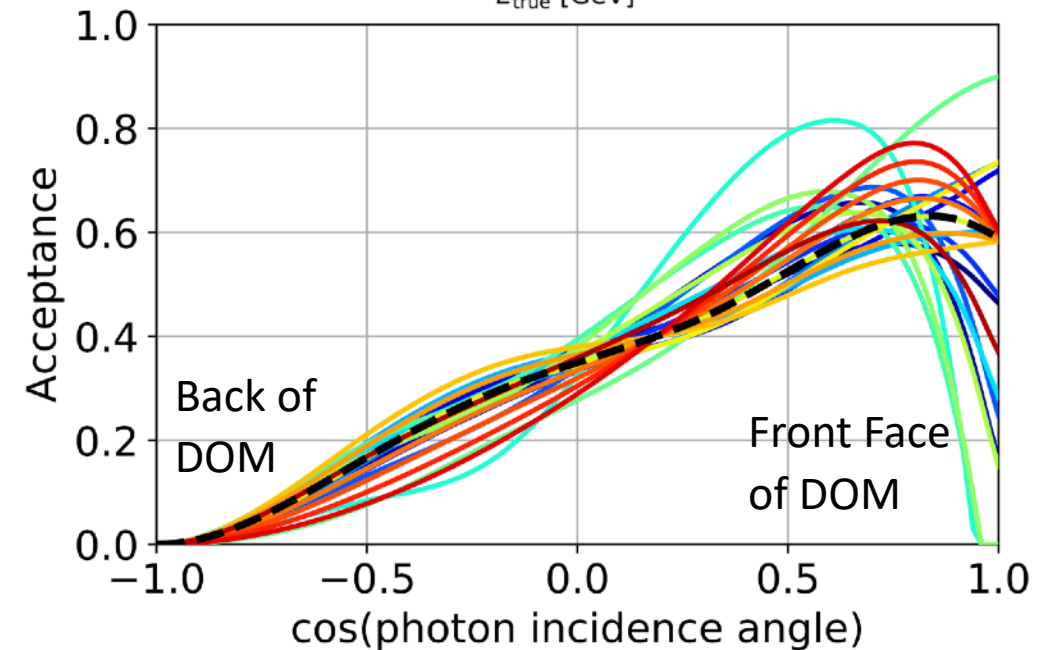
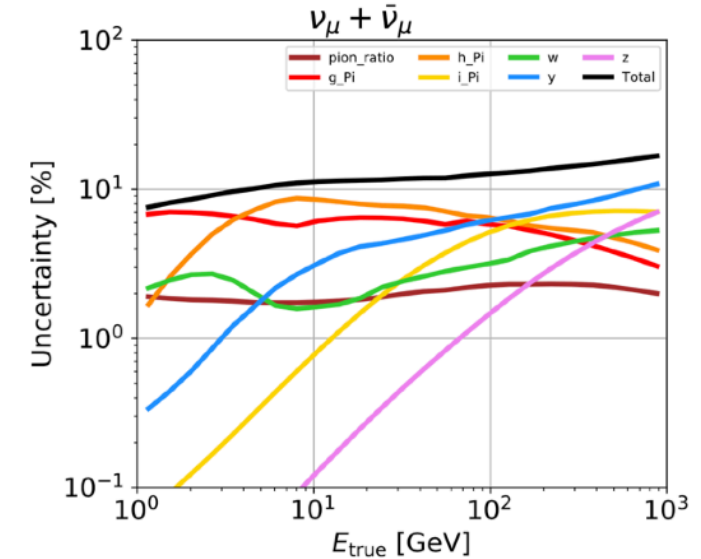
\*upcoming 9.3-year high-statistics event selection

\*\*arXiv:2203.02303

- Measure 3D distortions in reconstructed [energy, zenith, particle type]
  - Robust against systematic uncertainties
  - Particle identification discriminates  $\nu_{\mu,CC}$  interactions vs all other flavors/channels



- Flux
  - Cosmic ray spectrum
  - Pion & Kaon uncertainties from production/decay/(re)interaction
- Cross Sections <small impact>
  - Axial mass uncertainties for QE and RES
  - Transition between GENIE lower energies (approx. 100 GeV) to CSMS\* at higher energies
- Detector & Ice <large impact>
  - Optical property uncertainties from different ice layers and refrozen drill column
  - Angular dependence of photon acceptance on the digital optical module (DOM) photosensor
- ~40 systematic uncertainties are studied and approx. half are used in the final analyses fits





## Sub-sample

High quality events

~20k events

Fast reconstructions

- separate recos for energy and direction/vertex
- can only be applied to certain high-quality events

Results available

## Full Sample

High statistical power

~200k events

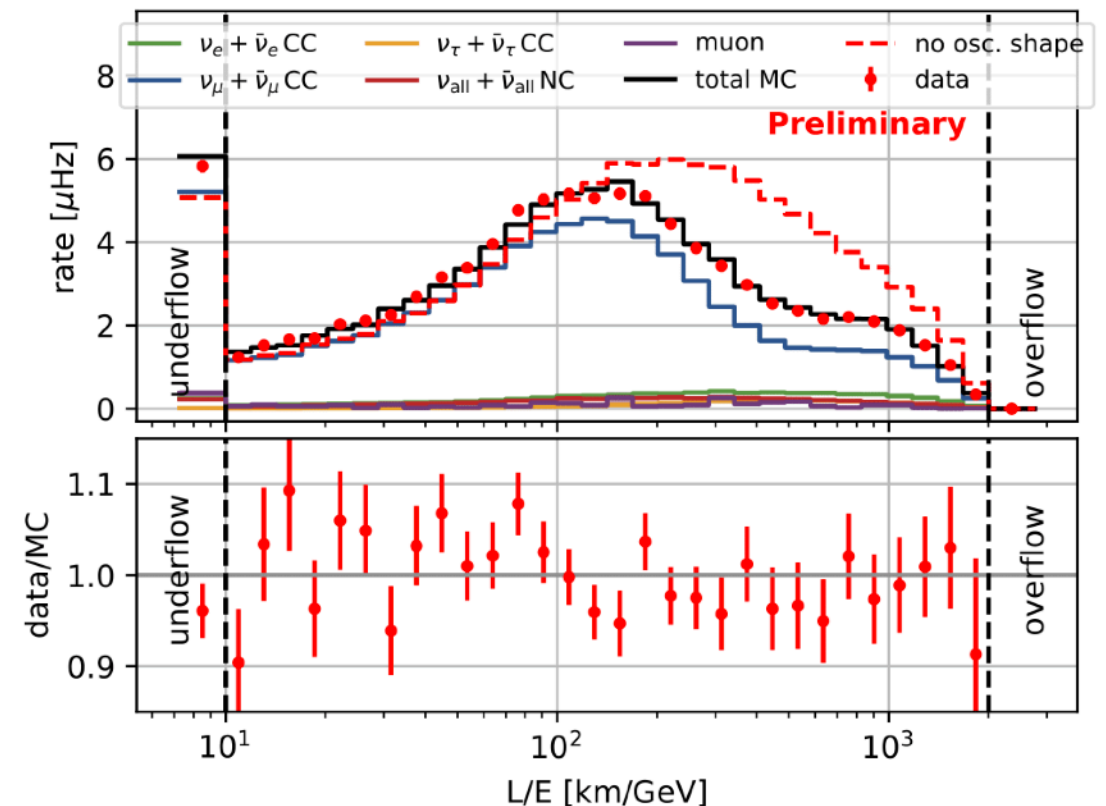
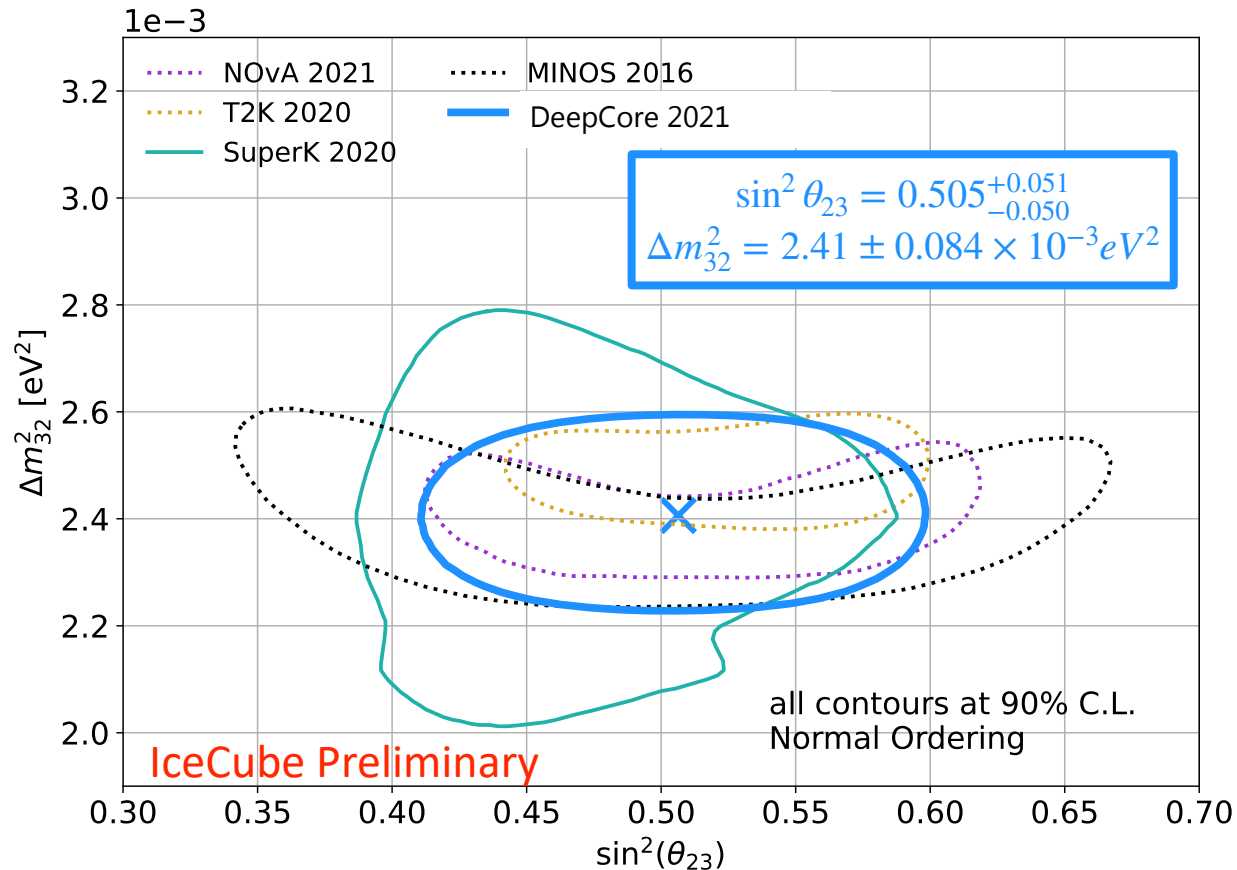
Full 8d reconstruction

- energy, direction, vertex fit simultaneously
- can be applied to almost any event

In progress

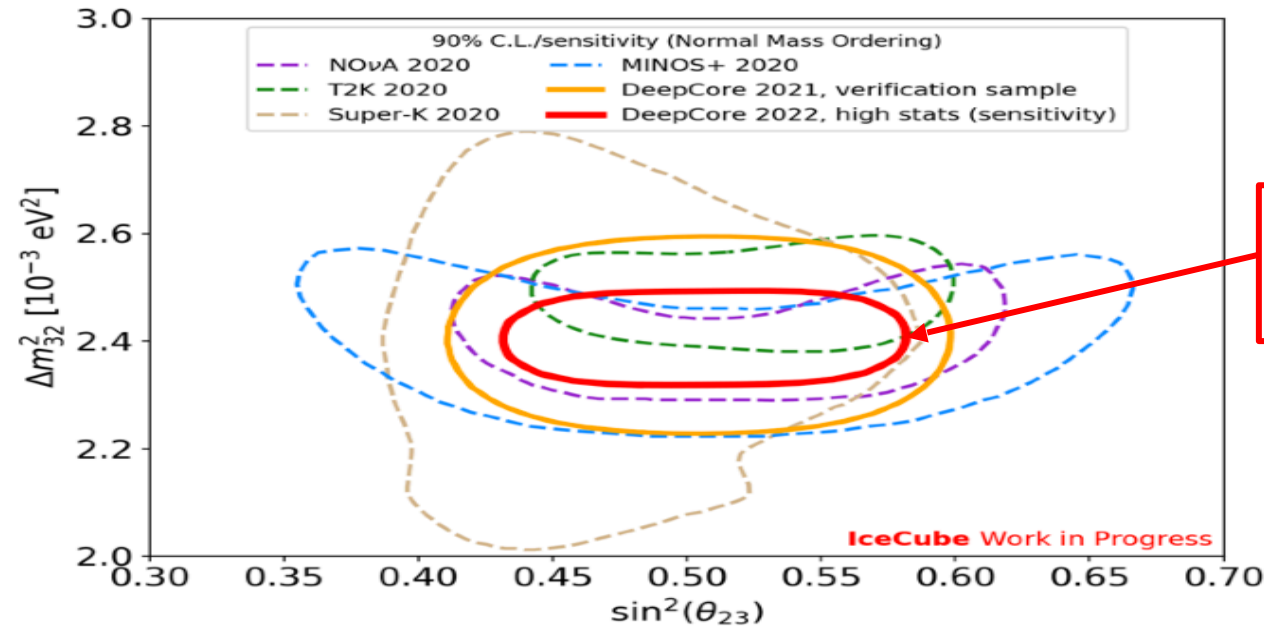


- New measurement of  $\nu_\mu$  disappearance with 8 years of IceCube data
  - Uses a “golden” sub-sample of ~23,000 track-like events (10% of full data set)
  - Clean events with low levels of photon scattering → robust to ice modelling uncertainties



- Suite of analyses underway with a new, high statistics data sample
  - All flavors, state-of-the-art reconstruction and background rejection
- Observe  $\nu_{\mu}$  disappearance and corresponding  $\nu_{\tau}$  appearance

### Atmospheric mixing parameter sensitivity



Sensitivity competitive with LBL accelerators

$\sim 210,000$  neutrinos (0.7% background)  $\rightarrow$  high stats and purity

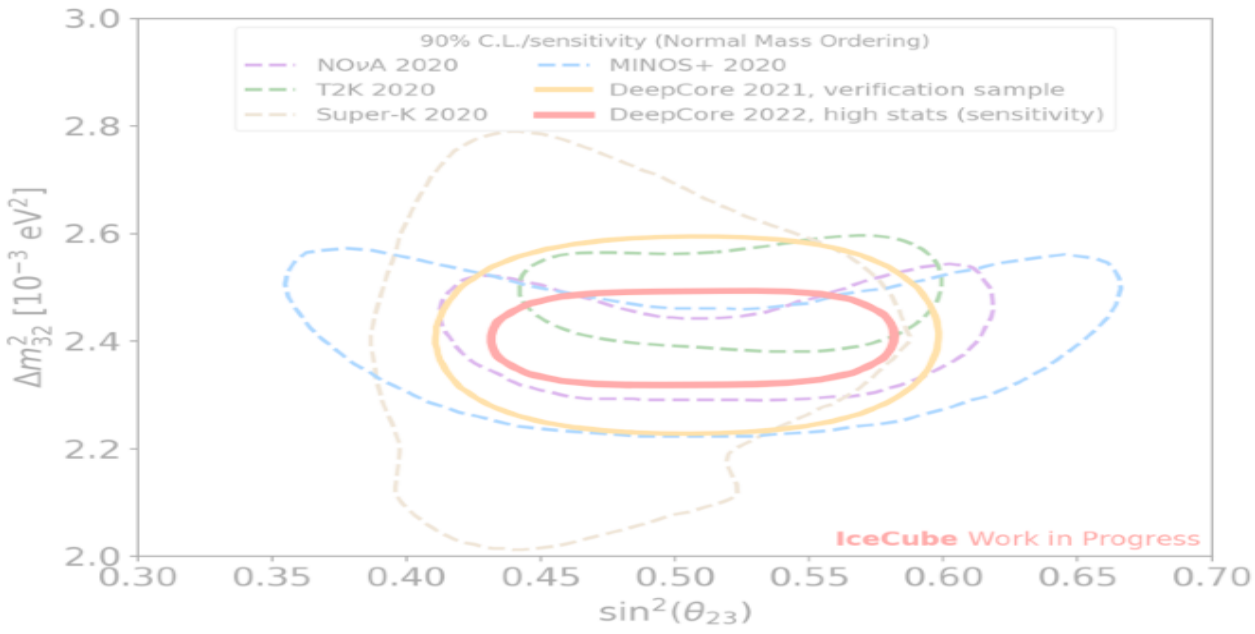
Flavor  
EigenstateMass  
Eigenstate

$$\begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \\ |\nu_\tau\rangle \end{pmatrix} = \mathcal{U}_{\text{PMNS}} \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \\ |\nu_3\rangle \end{pmatrix} \\ = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \\ |\nu_3\rangle \end{pmatrix}$$

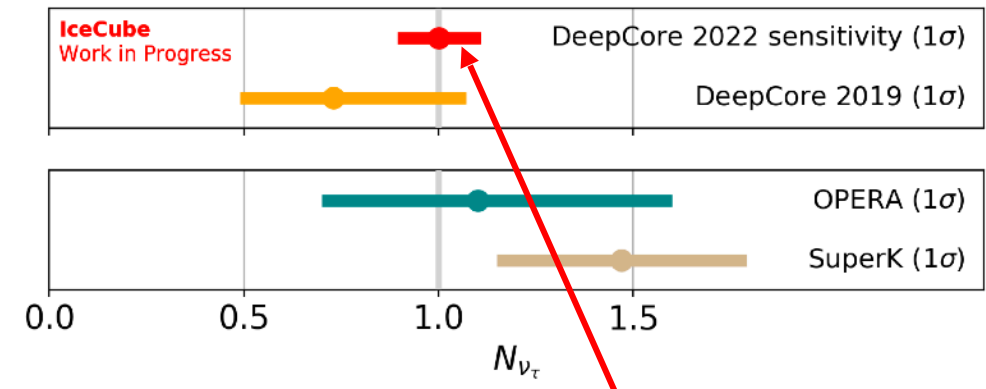
- Neutrino flavor eigenstates are related to mass eigenstates via the PMNS Unitary mixing matrix
- The 3 conventional angles ( $\theta_{12}$ ,  $\theta_{13}$ ,  $\theta_{23}$ ) fully describe the 3x3 mixing matrix **only if** the mixing matrix is unitary
- Tau neutrino appearance ( $\nu_\mu \rightarrow \nu_\tau$ ) explicitly tests  $U_{\tau3}$  and whether the matrix is 3x3 unitary;  
 $|U_{e3}|^2 + |U_{\mu3}|^2 + |U_{\tau3}|^2 \neq 1$

- Suite of analyses underway with a new, high statistics data sample
  - All flavors, state-of-the-art reconstruction and background rejection
- Observe  $\nu_\mu$  disappearance and corresponding  $\nu_\tau$  appearance

### Atmospheric mixing parameter sensitivity



### $\nu_\tau$ normalization sensitivity

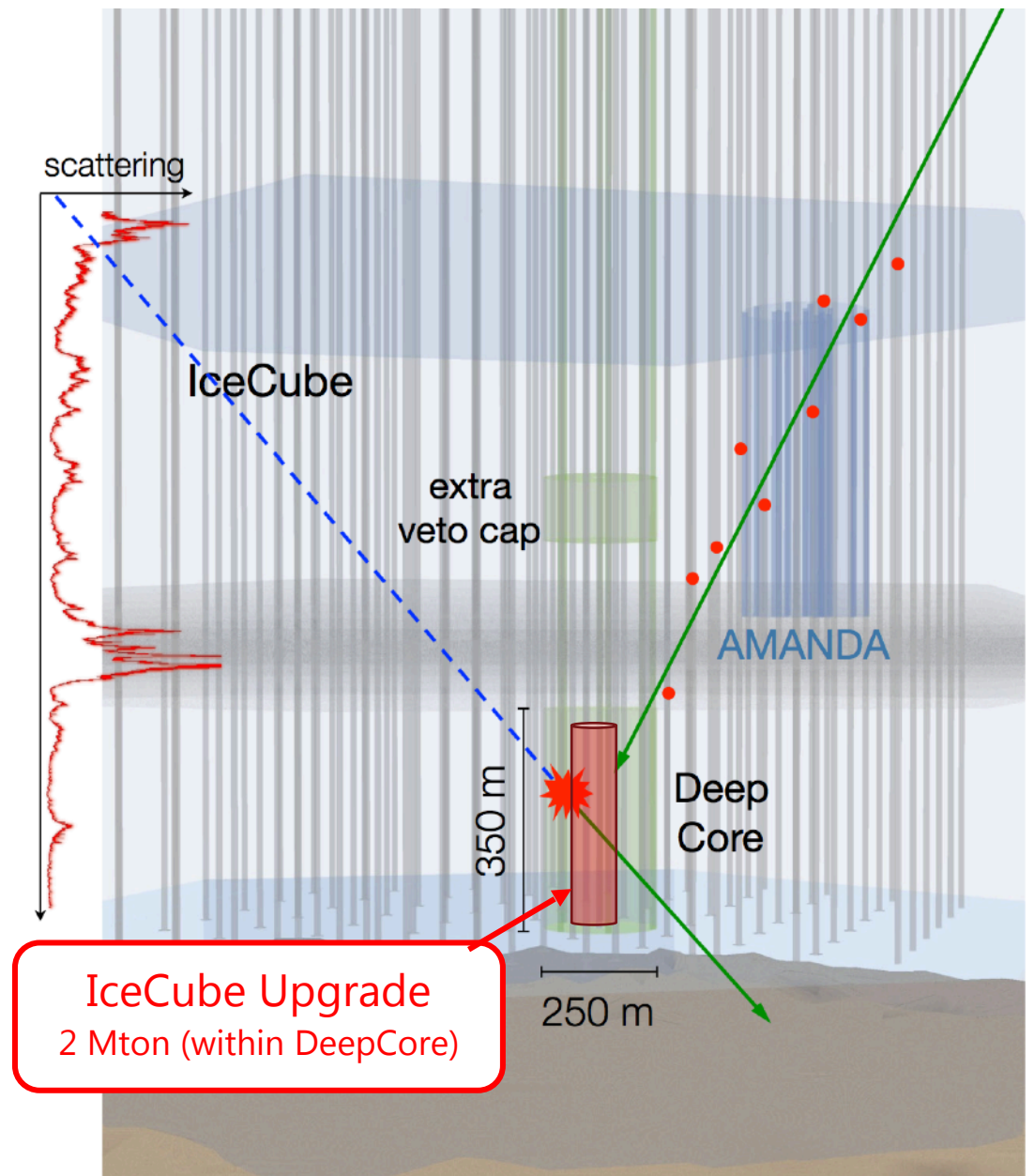
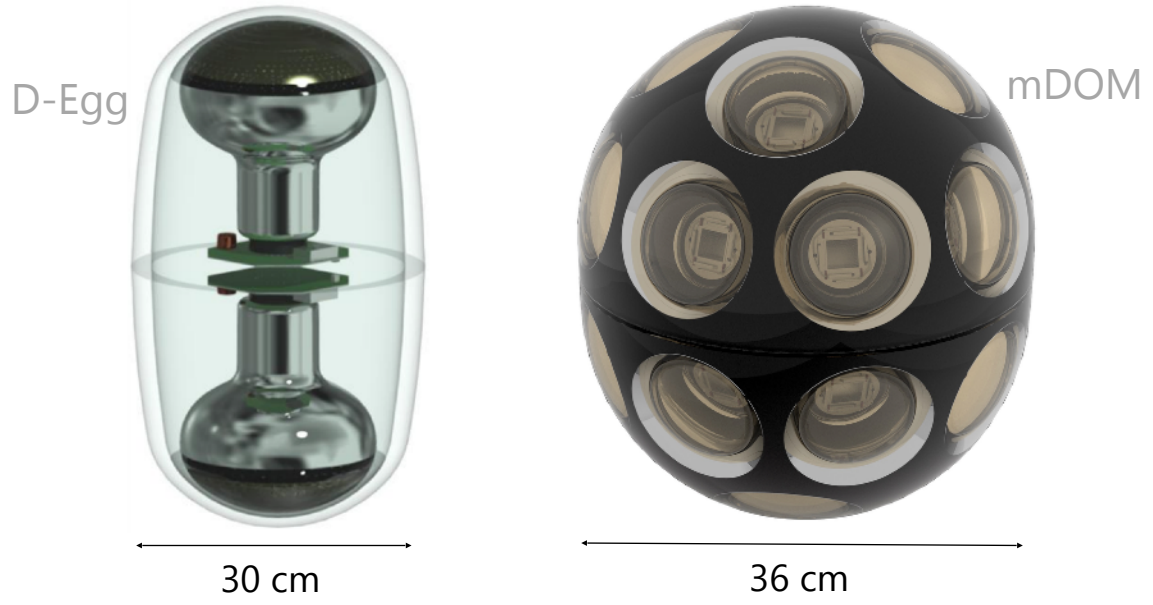


Expecting world-leading 11% precision  
 ~9,700  $\nu_{\tau,CC}$  events expected  
 Signal is statistical excess of upgoing cascades with suppressed cross section  
 Tests PMNS unitarity and  $\nu_{\tau,CC}$  cross section



# THE NEAR FUTURE

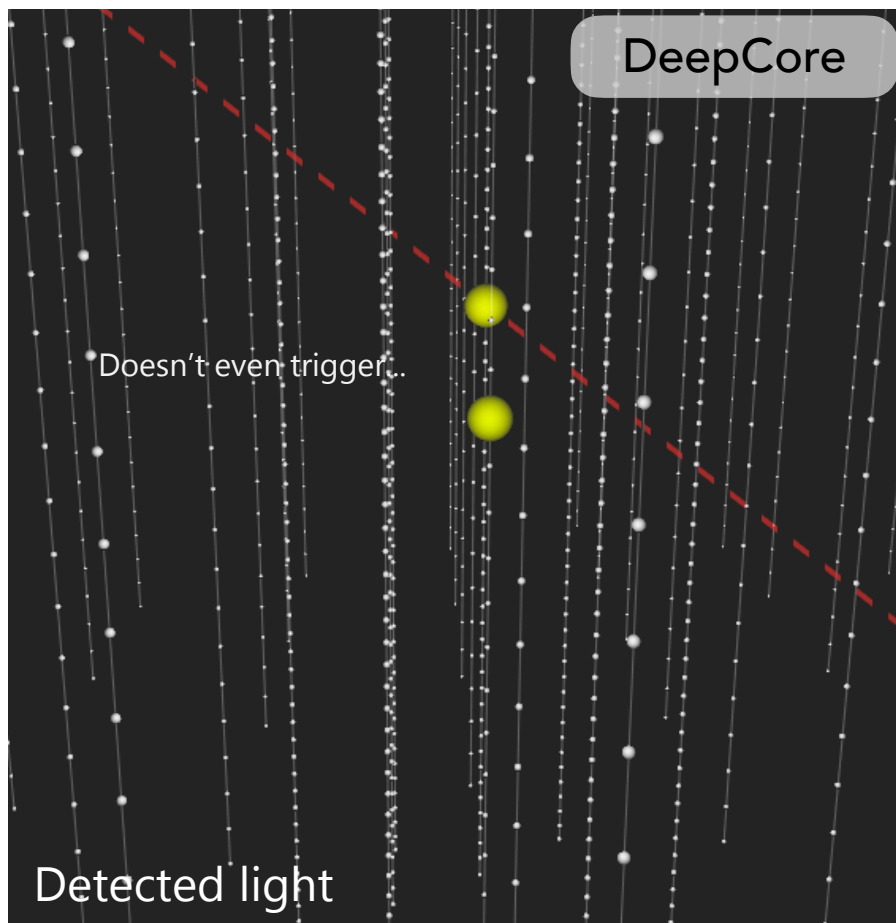
- Low-energy extension to IceCube
  - Deployment in 2025/6
  - Reduce threshold to 1 GeV
- 700 multi-PMT sensors
- Improved detector/ice calibration



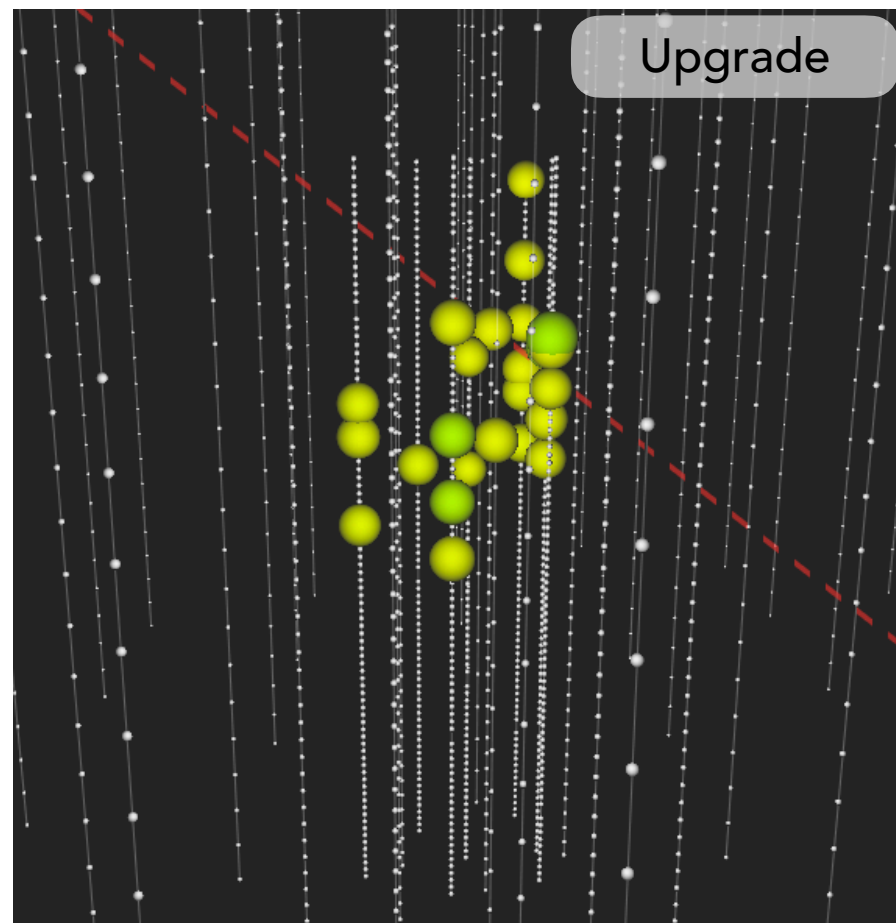
**IceCube Upgrade**  
2 Mton (within DeepCore)

- Dense instrumentation in 2 Mton core

- Large increase in photosensor density  $\rightarrow$  sensitive down to  $\sim 1$  GeV neutrinos



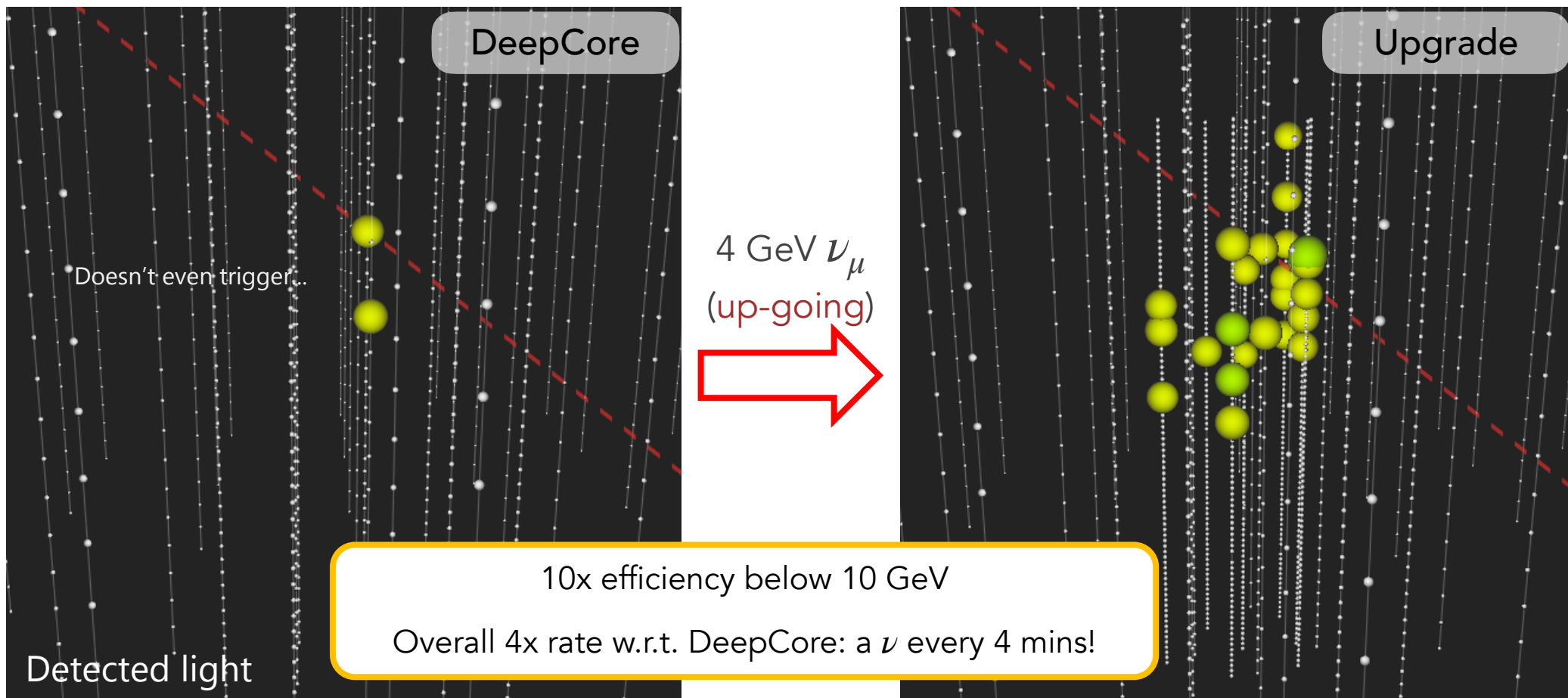
4 GeV  $\nu_{\mu}$   
(up-going)



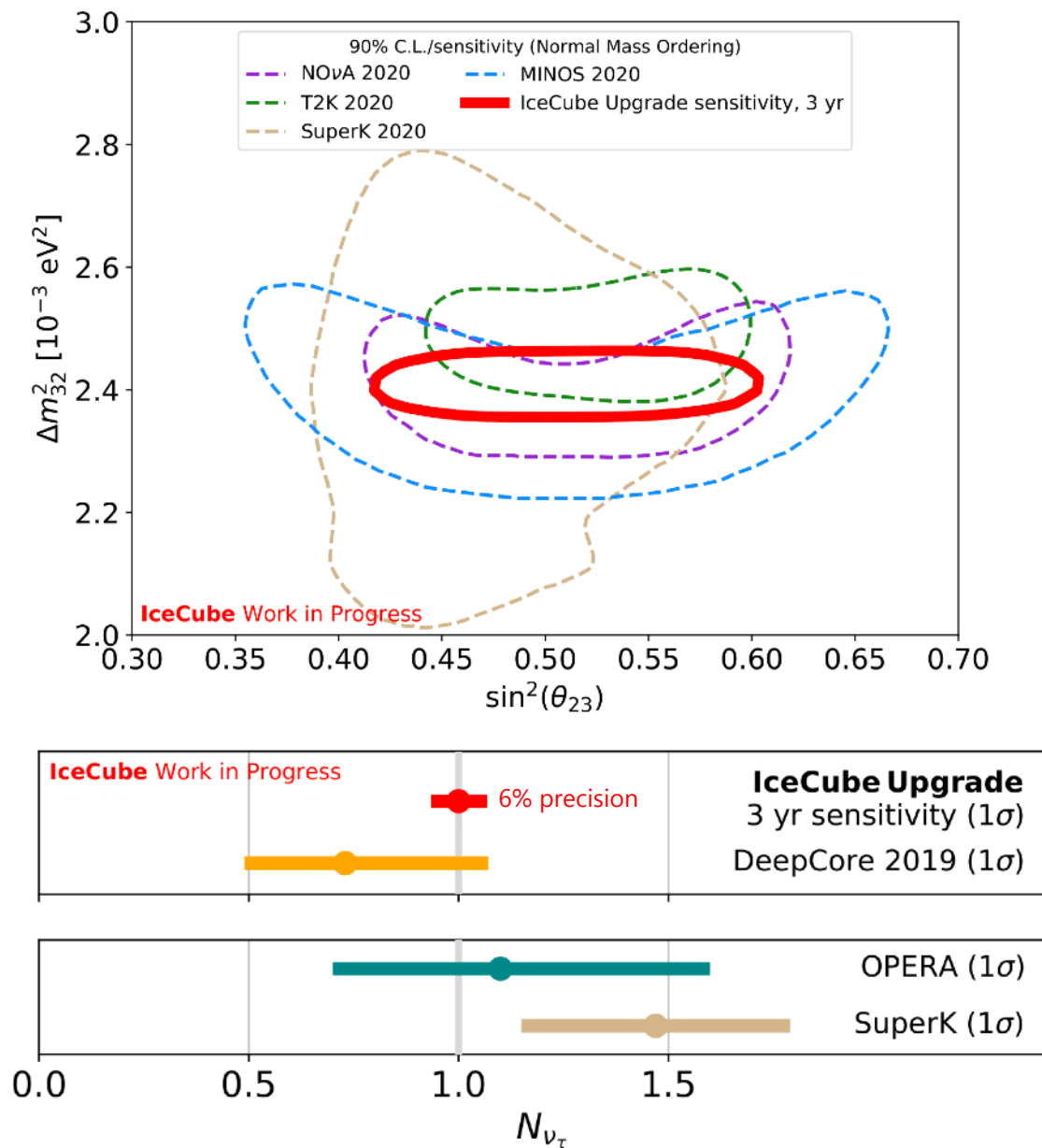


- Dense instrumentation in 2 Mton core

- Large increase in photosensor density  $\rightarrow$  sensitive down to  $\sim 1$  GeV neutrinos



- **Conservative** experimental choices still illustrate potential of IceCube Upgrade for physics
  - $\sim 6\%$   $N_{\nu_\tau}$  resolution with 1-year of data
  - Excludes improvements from new reconstruction techniques, improved detector systematic uncertainties, better flux treatment, and no combination of 10+ years of DeepCore data

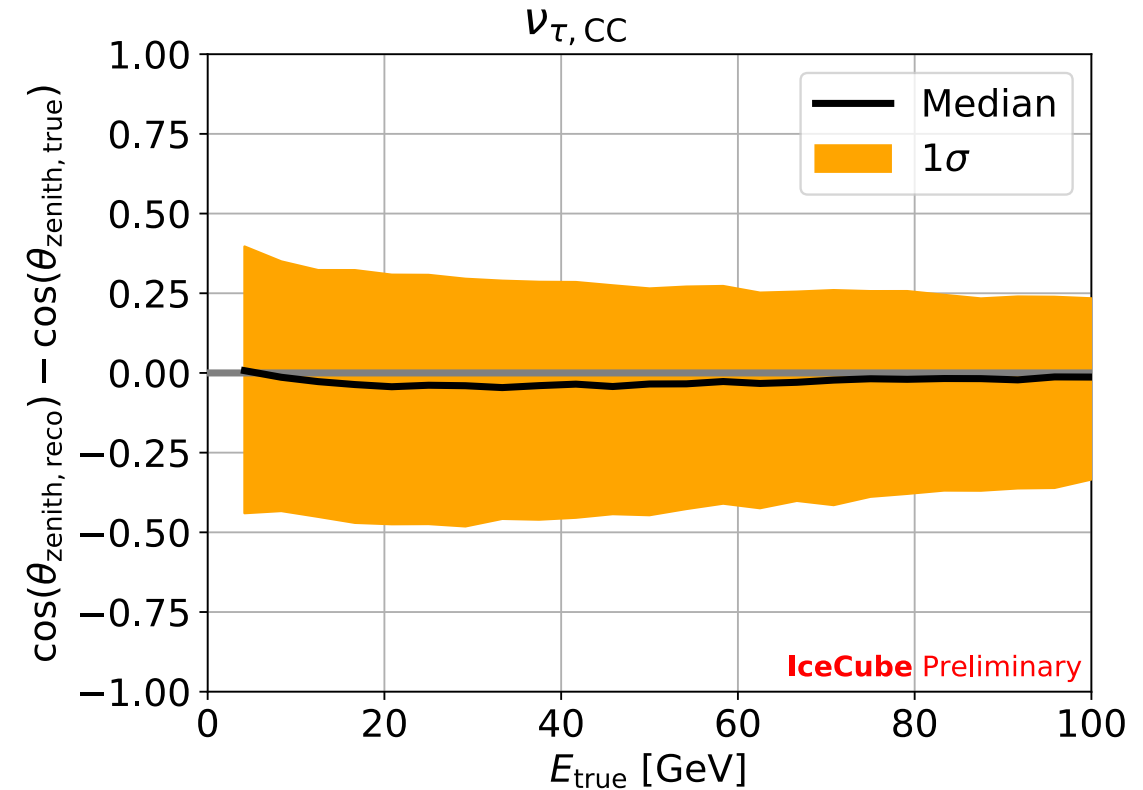


- Unique view of oscillations which complement (and will be competitive) with neutrino beam experiments
- Leading  $\nu_\tau$  appearance measurement(s) & constraints for (non)unitarity through  $N_{\nu_\tau}$
- IceCube Upgrade is on path to improve calibration techniques and expand GeV capabilities



# BACKUP

- Previous analyses - forward light propagation
- Upcoming analysis - backwards light propagation
- Future analysis & IceCube Upgrade - likelihood free inference, convolution neural network, graph neural network
  - Faster reconstructions will provide better event selections... positive feedback loop
  - Better reconstruction resolutions are always better



- ~40 systematic uncertainties evaluated for upcoming analysis

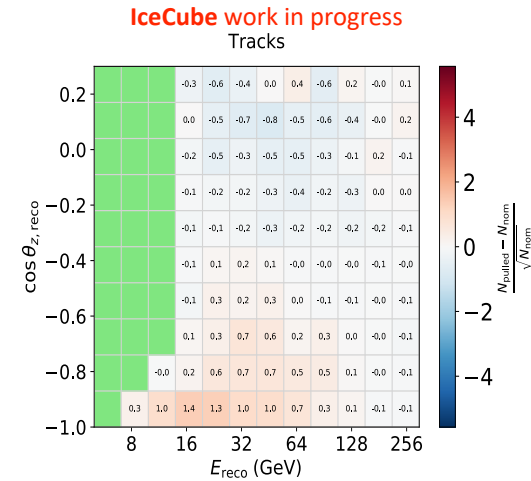
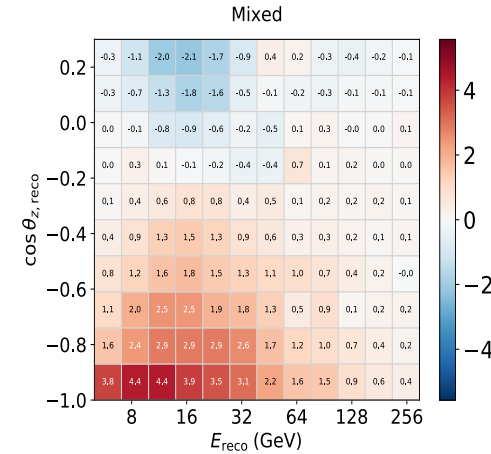
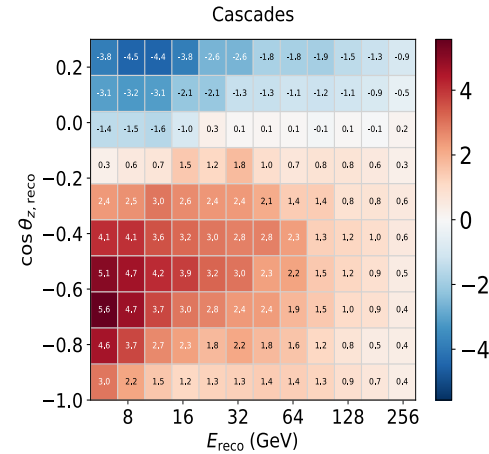
- Many have a negligible impact on oscillation measurements and are dropped

- The non-negligible systematics can be **very** non-negligible

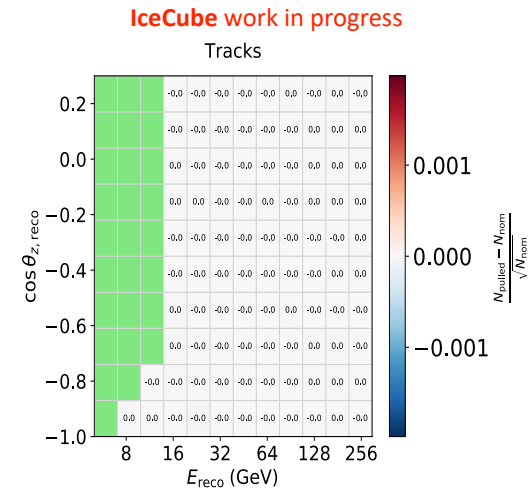
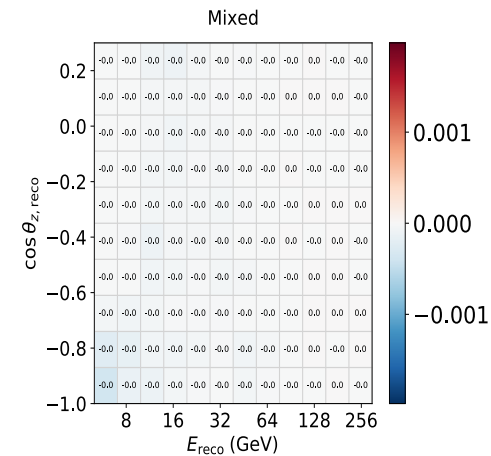
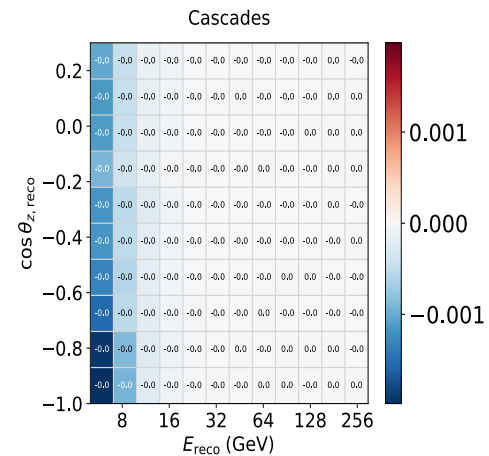
- e.g. DOM efficiency, hole ice photon scattering, spectral index

- Shown here are a subset of systematics checked

$\Delta(\text{Hole ice, } p_0) = +0.5 : 0.102 \rightarrow 0.602$



$\Delta(\text{Barr, } a_{\pi^+}) = -1\sigma : 0 \rightarrow -0.1$



\*arXiv:astro-ph/0611266

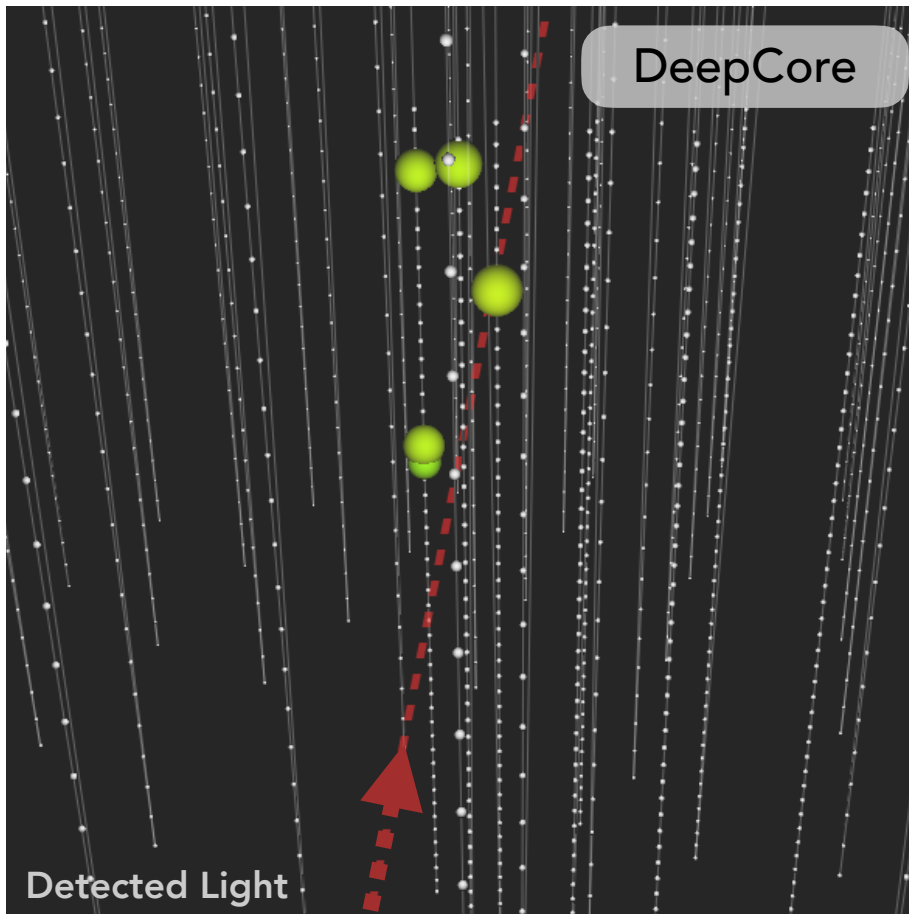
- Using ‘Barr blocks’\* for hadroproduction flux-related uncertainties
  - ~18 total
- Matrix Cascade Equation (MCEq)\*\* for flux
  - Fast and amenable to different cosmic ray spectra & hadronic interaction models
  - Correlated treatment of evolving flux particle interactions & decays

E <sub>i</sub> (GeV)	Pions			Kaons				
	<8	10%		30%		40%		
8–15	30%	10%	30%		40%			
15–30	30	10	5%	10%		30		
30–500	30		15%		40			
>500	30		15%+Energy dep.		40			
	0	0.5	x <sub>LAB</sub>	1	0	0.5	x <sub>LAB</sub>	1

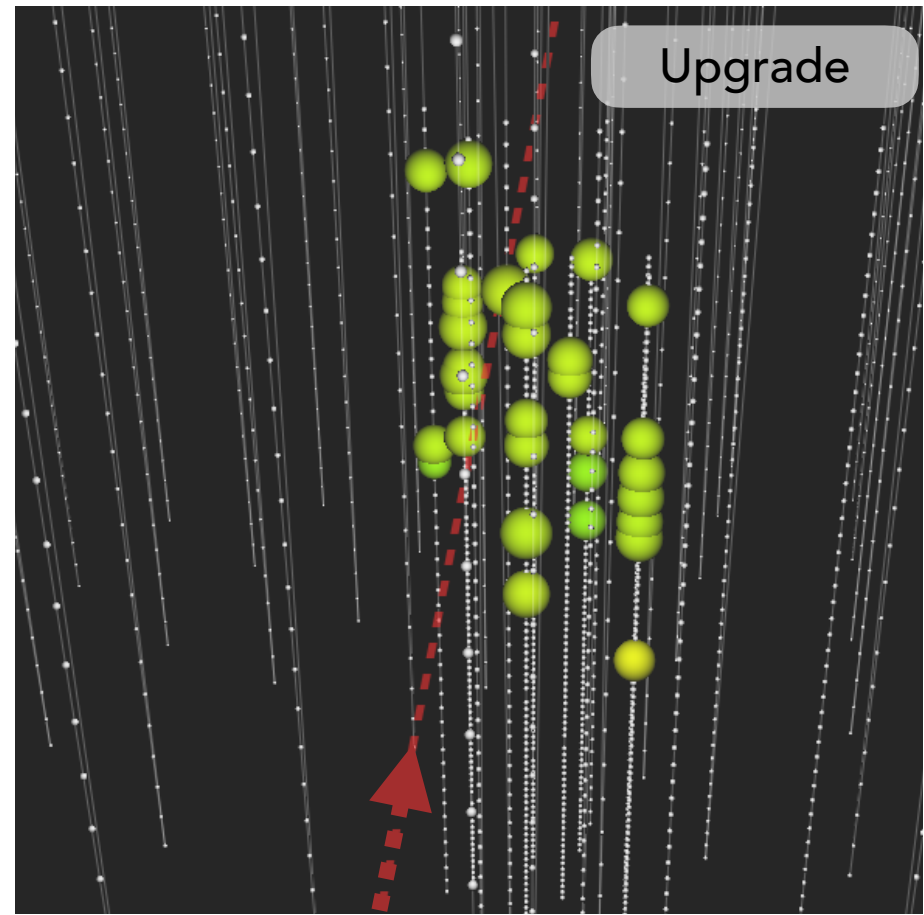
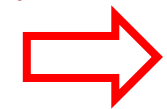
$$\frac{d\Phi}{dX} = [\hat{\Lambda}_{int} + \hat{\Lambda}_{dec}\rho^{-1}]\Phi$$

\*\*arXiv:1503.00544

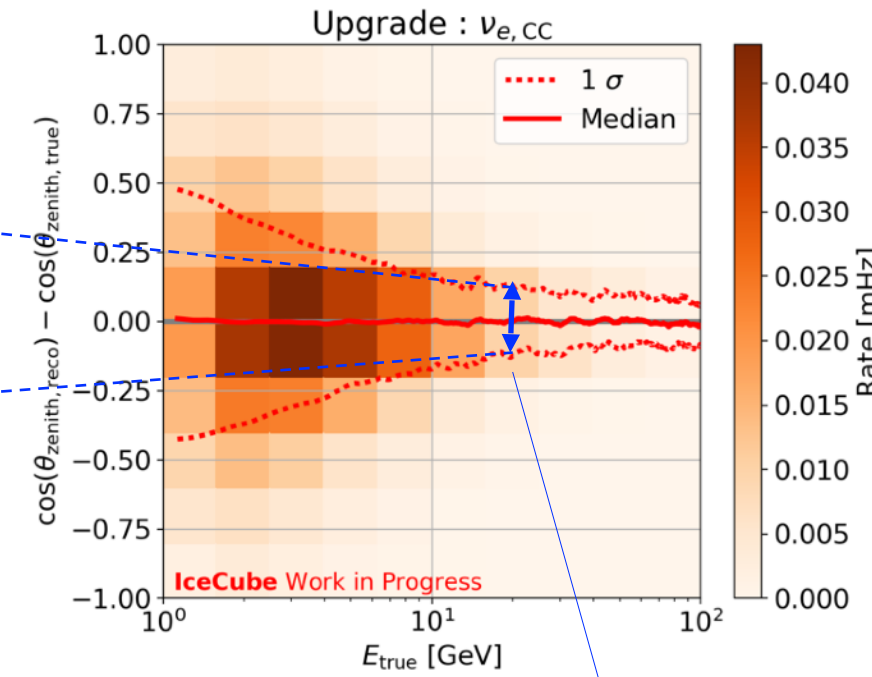
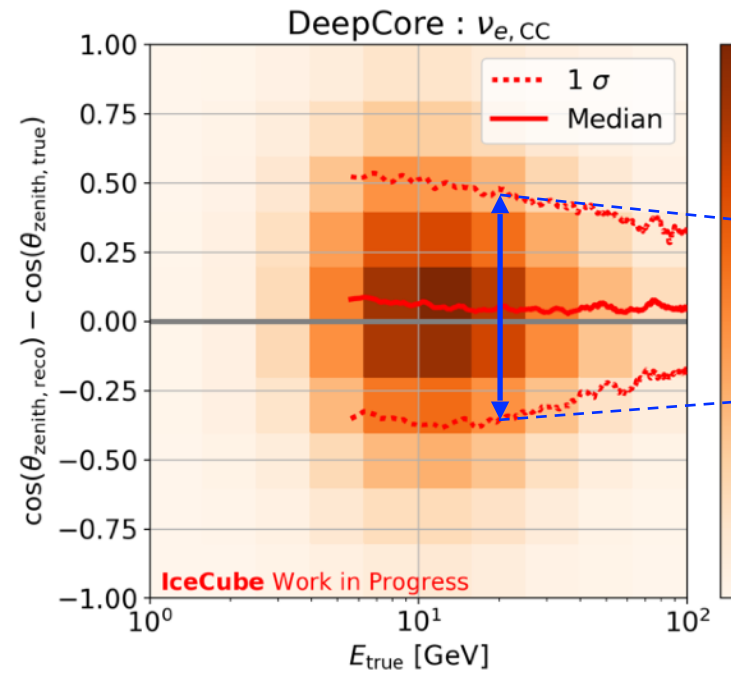
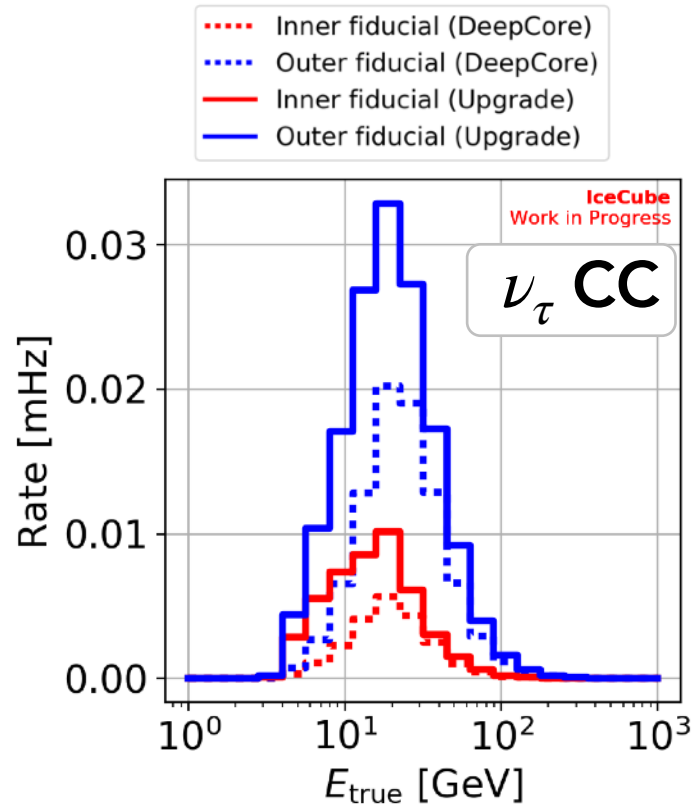
- Dense instrumentation within inner core



29 GeV  $\nu_\tau$   
(up-going)







\* $\nu_e$  CC and  $\nu_\tau$  CC both appear as cascades.  $\nu_e$  are an easier proxy for cascade reconstruction development.

**3x improvement in cascade resolution @  $\nu_\tau$  appearance energies**