

Axion dark matter from frictional misalignment

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- 3 DM from frictional misalignment
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Misalignment Mechanism

Assuming a pre-inflationary scenario for the scale of Peccei-Quinn breaking the value of the axion after inflation would be homogeneous $\frac{\partial_i}{f} \equiv \theta_i = \mathcal{O}(1)$ and follows the eom

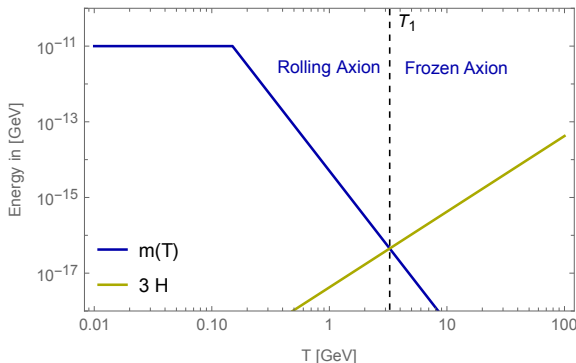
$$\ddot{\theta} + 3H\dot{\theta} + m(T)^2 \sin(\theta) = 0 \quad (1)$$

At early times $3H \gg m(T)$ the axion is frozen at it's initial value

$$\theta(T) = \theta_i \quad (2)$$

At around $3H \sim m(T)$ the axion is released starts oscillating around the bottom of the potential.

Misalignment Mechanism



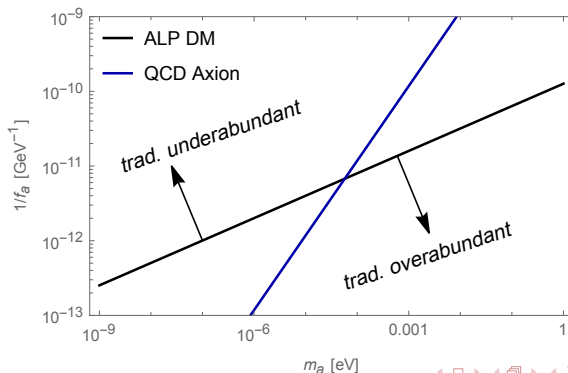
At late times the axion behaves as dark matter

$$A = \frac{a^3 \rho_a}{m(T)} = \text{ct} \rightarrow \rho_a \propto a^{-3} \quad (3)$$

Misalignment Mechanism

Axion dark matter abundance

$$\frac{\rho_{a,0}}{\rho_{\text{DM}}} \simeq 28 \sqrt{\frac{m_a}{\text{eV}}} \sqrt{\frac{m_a}{m_{\text{osc}}}} \left(\frac{\theta_i f_a}{10^{12} \text{ GeV}} \right)^2 \mathcal{F}(T_{\text{osc}}) \quad (4)$$



Axions in a thermal bath

We assume an axion coupled to a dark non-Abelian gauge field which forms a thermal bath of temperature T'

$$\mathcal{L} \supset \frac{\alpha}{8\pi} \theta F_{\mu\nu}^b \tilde{F}^{b\mu\nu}, \quad (5)$$

The effective EOMs for the axion background and gauge field are

$$\ddot{\theta}_a + [3H + \Upsilon(T')] \dot{\theta}_a = -\frac{1}{f_a^2} V'(\theta_a), \quad (6)$$

$$\dot{\rho}_{\text{dr}} + 4H\rho_{\text{dr}} = f_a^2 \Upsilon(T') \dot{\theta}_a^2 \quad (7)$$

Properties of the thermal bath

Friction coefficient for $\alpha < 0.1$

$$\Upsilon(T') = \frac{\Gamma_{\text{sph}}}{2T'f_a^2} \simeq 1.8 \times \frac{N_c^2 - 1}{N_c^2} \frac{(N_c \alpha)^5 T'^3}{2f_a^2} \quad (8)$$

by **McLerran et al.** Recently revived by **Berghaus et al** and applied to inflation, dark energy etc.

Running of the coupling is important

$$\alpha(T') = \frac{4\pi}{\bar{b}_0 N_c} \frac{1}{\ln(T'^2/\Lambda^2)} \quad (9)$$

where $\bar{b}_0 = \frac{11}{3}$ for confinement or $\frac{10}{3}$ spontaneous symmetry breaking.

Properties of the thermal bath

$\rho_X \equiv$ energy of dark gauge field and its decay byproducts

$$\Delta N_{\text{eff}} = \frac{8}{7} \left(\frac{11}{4} \right)^{\frac{4}{3}} \frac{\rho_X}{\rho_\gamma} \bigg|_{T=T_{\text{rec}}} < 0.3 \text{ at } 95\% \text{C.L.} \quad (10)$$

We define $\xi \equiv \frac{T'_0}{T_0}$

$$\Delta N_{\text{eff}} = 0.016 \times n^{-1/3} (2N_c^2 - 2)^{4/3} \xi^4, \quad (11)$$

For $SU(3)$ we get $\xi < 0.86$

Properties of the thermal bath

The temperature of the dark thermal bath can be related to the standard model temperature through their respective entropy conservation.

$$T' = \xi \left(\frac{g_{\text{s,SM}}(T) g_{\text{s}}'(T'_0)}{g_{\text{s,SM}}(T_0) g_{\text{s}}'(T')} \right)^{1/3} T \quad (12)$$

Motion of the axion at early times

For $m(T) \equiv m_0 \left(\frac{\Lambda}{T}\right)^\beta$, if $3H \gg \Upsilon(T')$

$$\theta_a(T) \simeq \theta_i e^{-\frac{m_a(T)^2}{6(2+\beta)H(T)^2}}, \quad (13)$$

whereas if $\Upsilon(T') \gg 3H$

$$\theta_a(T) \simeq \theta_i e^{-\frac{m_a(T)^2}{(5+2\beta)\Upsilon(T')H(T)}}, \quad (14)$$

The onset of rolling is given by

$$m_a(T_{\text{osc}}) \simeq \begin{cases} 4H(T_{\text{osc}}) & , 3H > \Upsilon \\ \frac{10\Upsilon(T'_{\text{osc}})H(T_{\text{osc}})}{m_a(T_{\text{osc}})} & , 3H < \Upsilon \end{cases} \quad (15)$$

Motion of the axion at late times

We derive a new adiabatic invariant for generic friction coefficient $\Gamma(T)$

$$A = \frac{\rho_\theta(t)}{\omega(t)} \exp \left[\int^t d\tilde{t} \Gamma(\tilde{t}) \right] = \text{const}, \quad (16)$$

Which recreates the correct result when $\Gamma(T) = 3H(T)$

$$A = \frac{\rho_\theta(T)}{m_a(T)} \exp \left[\int_{t_{\text{osc}}}^t d\tilde{t} 3H(\tilde{t}) \right] = \frac{\rho_\theta a^3}{m_a} = \text{const.} \quad (17)$$

and yields a new result when one considers both Hubble and thermal friction

$$A_{\text{fr}} = \frac{\rho_\theta(T) a^3(T)}{m_a(T)} \exp \left[\int^t d\tilde{t} \Upsilon(\tilde{t}) \right] = \text{const.} \quad (18)$$

DM abundance in the presence of friction

$$\frac{\rho_{a,0}}{\rho_{\text{DM}}} \simeq \underbrace{28 \sqrt{\frac{m_a}{\text{eV}}} \sqrt{\frac{m_a}{m_{\text{osc}}}} \left(\frac{\theta_i f_a}{10^{12} \text{ GeV}} \right)^2 \mathcal{F}}_{\text{standard result}} \underbrace{e^{-D}}_{\text{suppression}} \underbrace{\left(\frac{m_{\text{osc}}}{4 H_{\text{osc}}} \right)^{3/2}}_{\text{enhancement}} \quad (19)$$

where

$$D \simeq 6.3 \left(\frac{10^8 \text{ GeV}}{f_a} \right)^2 \left(\frac{\Lambda}{150 \text{ MeV}} \right) \times \left[\frac{\tau^3 + \tau^2 + 2\tau + 6}{\tau^4} e^\tau - \text{Ei}(\tau) \right]_{\tau_{\text{osc}}}^{\tau_{\text{end}}} \quad (20)$$

and $\tau \equiv \ln \left(\frac{T'}{\Lambda} \right)$

Basic mechanism

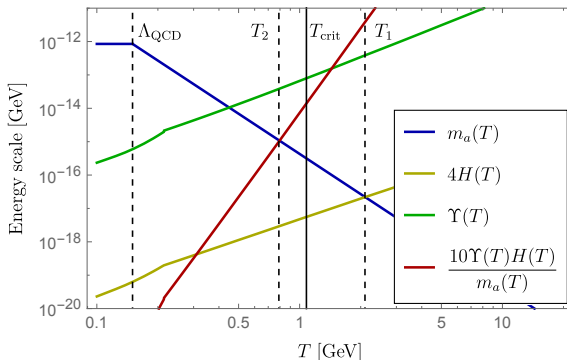


Figure 1: Example for the QCD axion

Minimal ALP scenario

We assume a single gauge group that gives rise to the mass through instanton effects and friction through sphaleron transitions. In that case $m_0 = \frac{\Lambda^2}{f_a}$

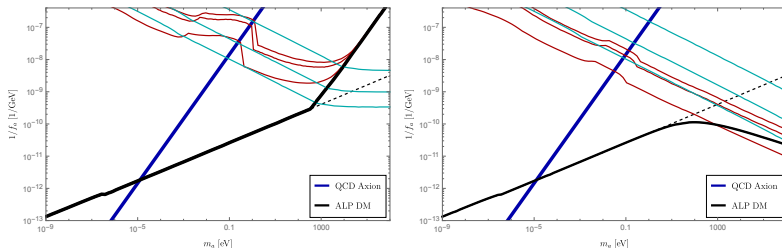


Figure 2: Left panel for $\alpha_{\text{thr}} = 0.2$ and right panel for $\alpha_{\text{thr}} = 0.4$

ALP coupled to two gauge groups

The model under consideration is

$$\mathcal{L}_{\text{int}} = \frac{\alpha_G}{8\pi} \theta_a G_{\mu\nu}^b \tilde{G}^{b\mu\nu} + \lambda \frac{\alpha}{8\pi} \theta_a F_{\mu\nu}^b \tilde{F}^{b\mu\nu}, \quad (21)$$

In this case $m_0 = \frac{\Lambda_G^2}{f_a}$, we define the enhancement parameter

$$\lambda \equiv \text{enhancement parameter} \quad (22)$$

which we assume may be very large. Such largeness can be justified by alignment (Kim et al) or clockwork mechanism (Kaplan et al) in which case $\lambda = 3^N$.

ALP coupled to two gauge groups (underabundant case)

Condition for opening the underabundant regime:

$$T_2 \leq T_{\text{crit}} \quad (23)$$

where $T_{\text{crit}} \simeq 21.6 \text{ GeV} \left(\frac{m_a f_a}{\text{GeV}^2} \right)^{4/7} \frac{\mathcal{F}^{1/7}}{g_{\rho, \text{SM}}(T_{\text{crit}})^{3/28}}$

Simplifies to

$$\frac{\mathcal{F}_a \left(\frac{m_a f_a}{17.0 \text{ GeV}^2} \right)^{10/7} \lambda^2}{\left[1 + 0.17 \left(\ln \left[\mathcal{F}_b \left(\frac{m_b f_b}{\text{GeV}^2} \right)^{1/7} \right] + \ln [\Lambda_G^2 / \Lambda^2] \right) \right]^5} > 1 \quad (24)$$

ALP coupled to two gauge groups (overabundant case)

- We simply demand that the axion dilutes sufficiently enough after it starts to roll so that its abundance matches the observed one.
- The minimum value of λ in this case is when the friction is the minimum possible over the longest possible time

ALP DM parameter space

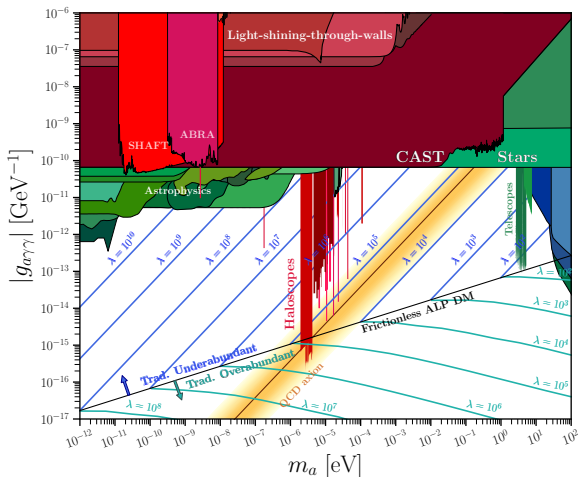


Figure 3: Minimum enhancement parameter λ for ALP DM

What about the QCD axion?

The same results apply for the QCD axion with two important differences

- Our conclusions require a small correction due to the presence of light degrees of freedom charged under QCD.

~~$m_a^{\text{QCD}} f_a = \Lambda_{\text{QCD}}^2$~~ instead

$$m_a^{\text{QCD}} f_a = m_\pi f_\pi \frac{\sqrt{z}}{1+z}, \quad \text{where } z \equiv \frac{m_u}{m_d} \quad (25)$$

- The only possible fate of the hidden sector is spontaneous breaking so that the axion potential is not affected and the strong CP problem is solved.

Conclusions

- We call this mechanism "Frictional misalignment". We hope to add it to a short list of other modifications of the standard mechanism such as "Kinetic misalignment" [Co et al](#), "Trapped misalignment" [Di Luzio et al](#) etc.
- It can open up both the traditional over and underabundant regimes.
- Most of the parameter space requires the clockwork mechanism to justify the large scale hierarchy.
- It mostly evades constraints from axion fragmentation.
- Works for the QCD axion.
- Alters the picture in the minimal model for axion masses greater than 10^2 eV

Thank You