Third Family Hypercharge

by
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- Anomalies: $b \to s\ell\ell$
- QFT anomalies
- ullet A simple-minded Z' model
- \bullet Z' searches





Cambridge Pheno Working Group

Where data and theory collide



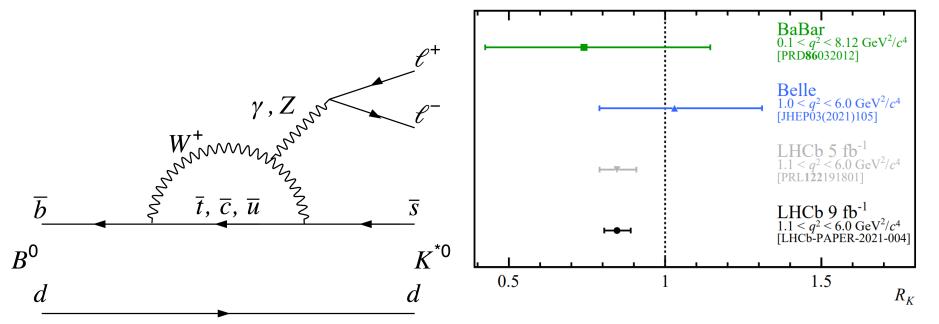
Strange b Activity



$R_K^{(*)}$ in Standard Model

$$R_K = \frac{BR(B \to K\mu^+\mu^-)}{BR(B \to Ke^+e^-)}, \qquad R_{K^*} = \frac{BR(B \to K^*\mu^+\mu^-)}{BR(B \to K^*e^+e^-)}.$$

These are rare decays (each BR $\sim \mathcal{O}(10^{-7})$) because they are absent at tree level in SM+EW+CKM

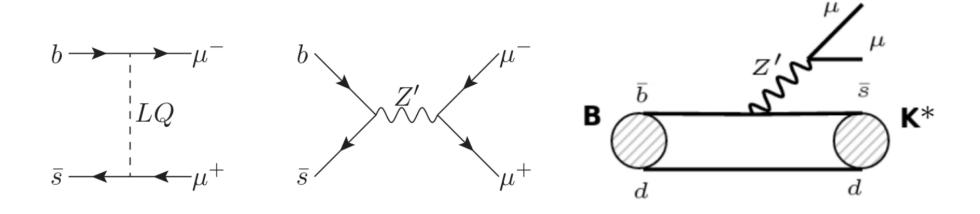




$b \to s \mu \mu$ Simplified Models

A good few $2-4\sigma$ Discrepancies with SM predictions. Computing with look elsewhere effect implies a 4.3σ discrepancy with the SM (conservative theory errors).¹

We have tree-level flavour changing new physics options:



¹Isidori, Lancierini, Owen and Serra, arXiv:2104.05631

Extra u(1) plus SM-singlets

Idea: break $SM \times U(1)_X$ gauge group around a TeV to get Z'. If $U(1)_X$ charges are family non-universal, we should impose quantum field theoretic anomaly cancellation.

- ullet Other uses for Z': dark matter models, axions, fermion masses, ...
- 3 RH neutrinos
- ullet Now, field labels denote the extra u(1) charge
- ACCs become

$$\begin{split} 3^2X: & \ 0 = \sum_{j=1}^3 \left(2Q_j + U_j + D_j\right), \\ 2^2X: & \ 0 = \sum_{j=1}^3 \left(3Q_j + L_j\right), \\ Y^2X: & \ 0 = \sum_{j=1}^3 \left(Q_j + 8U_j + 2D_j + 3L_j + 6E_j\right), \\ \text{grav}^2X: & \ 0 = \sum_{j=1}^3 \left(6Q_j + 3U_j + 3D_j + 2L_j + E_j + N_j\right), \\ YX^2: & \ 0 = \sum_{j=1}^3 \left(Q_j^2 - 2U_j^2 + D_j^2 - L_j^2 + E_j^2\right), \\ X^3: & \ 0 = \sum_{j=1}^3 \left(6Q_j^3 + 3U_j^3 + 3D_j^3 + 2L_j^3 + E_j^3 + N_j^3\right). \end{split}$$

Diophantine Equations

- ullet Since this is u(1), charges are commensurate: looking for compact extensions like the SM
- Thus we are looking for solutions over \mathbb{Z}^{18} .
- Any overall real factor in charge can be absorbed in $u(1)_X$ gauge coupling: $\mathcal{L} \supset -g_X \sum_{\psi} X_{\psi} \overline{\psi} X_{\mu} \gamma^{\mu} \psi$
- General diophantine equations are difficult to solve analytically over the integers
- Number theory state-of-the art for general analytic solution of generic diophantine equations is roughly one cubic in three unknowns

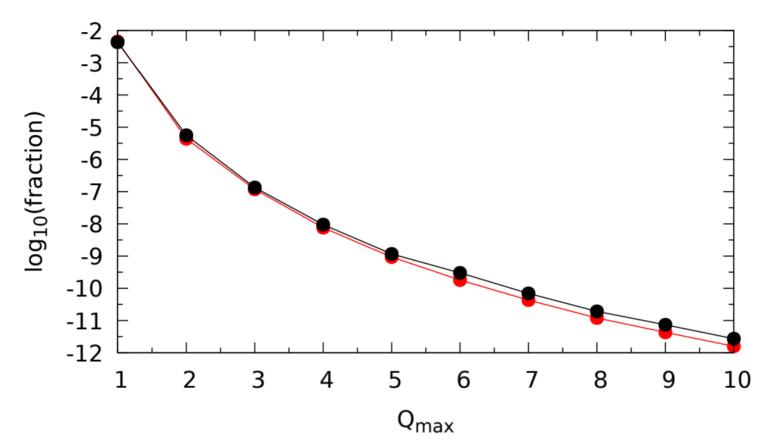
Anomaly-free Atlas

To find solutions for fixed $n \le 3$ and charges between -10 and 10, we did a numerical scan $(21^{18} \sim 10^{24})$: BCA, Davighi, Melville, arXiv:1812.04602.

An Anomaly-Free Atlas is available for public use: http://doi.org/10.5281/zenodo.1478085

Extended to semisimple case (340) in BCA, Gripaios, Tooby-Smith 2104.14555 and MSSM $+3\nu_R \times U(1)_X$ in BCA, Madigan, Tooby-Smith 2107.07926.

Davighi and Tooby-Smith 2206.11271 have investigated which $SM \times U(1)_X$ models fit in semisimple completions.



We begin with 18 charges and 6 anomaly equations reduce these to a 12-dimensional surface of solutions, extending out to infinity, but sparser away from $\mathbf{0}$.

Q_{\max}	Solutions	Symmetry	Quadratics	Cubics	Time/sec
1	38	16	144	38	0.0
2	358	48	31439	2829	0.0
3	4116	154	1571716	69421	0.1
4	$\boldsymbol{24552}$	338	34761022	932736	0.6
5	111152	796	442549238	7993169	6.8
6	435305	1218	3813718154	49541883	56
7	1358388	2332	24616693253	241368652	312
8	3612734	3514	127878976089	978792750	1559
9	9587085	5648	558403872034	3432486128	6584
10	21546920	7540	2117256832910	10687426240	24748

Inequivalent solutions with 3 RH ν

Known Solutions

	Q_1	Q_2	Q_3	U_1	U_2	U_3	D_1	D_2	D_3	L_1	L_2	L_3	E_1	E_2	E_3	N_1	N_2	N_3
A	0	0	1	0	0	-4	0	0	-2	0	0	-3	0	0	6	0	0	0
B	1	1	1	-1	-1	-1	-1	-1	-1	-3	-3	-3	3	3	3	3	3	3
C	-1	0	1	-1	0	1	-1	0	1	-1	0	1	-1	0	1	0	0	0

- *A* is TFHM (BCA, Davighi, arXiv:1809.01158)
- B is B-L, vector-like
- C has inter-family cancellation

Analytic Solution

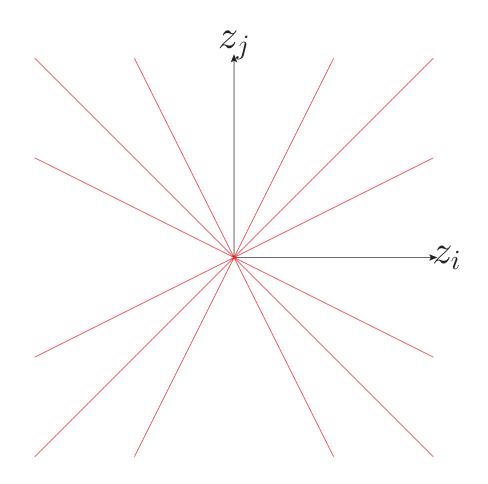
Want a full, general analytic solution for any Q_{\max} .

First step is to convert it into a problem in geometry by noting that solutions over $\mathbb Q$ are equivalent to those over $\mathbb Z$ by clearing all denominators. Since $\mathbb Q$ is a field, you can define geometry on it.

We start with \mathbb{Q}^{18} solution space.

All solutions where charges z_i differ by a common multiple are physically equivalent so we define an equivalence class to obtain $P\mathbb{Q}^{17}$.

Projective space $P\mathbb{Q}^{17}$



2d surface through origin becomes a line in projective space and a line through origin becomes a point

Preliminaries

4 linear equations restrict $P\mathbb{Q}^{17}$ to a projective subspace isomorphic to $P\mathbb{Q}^{13}$. Within this, we look for the intersection of a quadratic surface

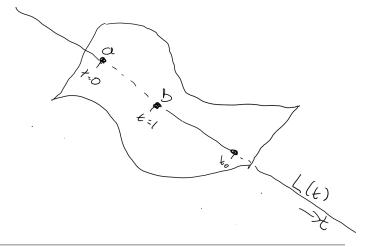
$$0 = \sum_{j=1}^{3} (Q_j^2 - 2U_j^2 + D_j^2 - L_j^2 + E_j^2)$$

and a cubic surface

$$0 = \sum_{j=1}^{3} \left(6Q_j^3 + 3U_j^3 + 3D_j^3 + 2L_j^3 + E_j^3 + N_j^3 \right)$$

The Method of Chords²

"A chord intersecting a rational cubic surface at two known rational points intersects it at 1 other $\mathbb Q$ point" eg Rational cubic $c(z_i) = 0$. Put a line through 2 known intersections a, b: L(t) = a + t(b - a). Along line, $c(L(t)) = kt(t-1)(t-t_0)$, where $k, t_0 \in \mathbb Q$.



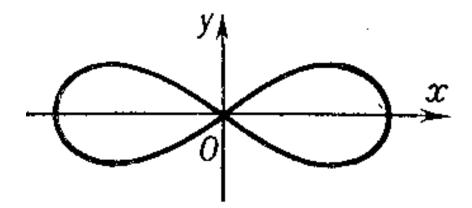
²Newton, Fermat, C17th

Caveat: It is possible that the line lies entirely within the cubic surface, i.e. c(L(t)) = 0 irrespective of t.

Double Points

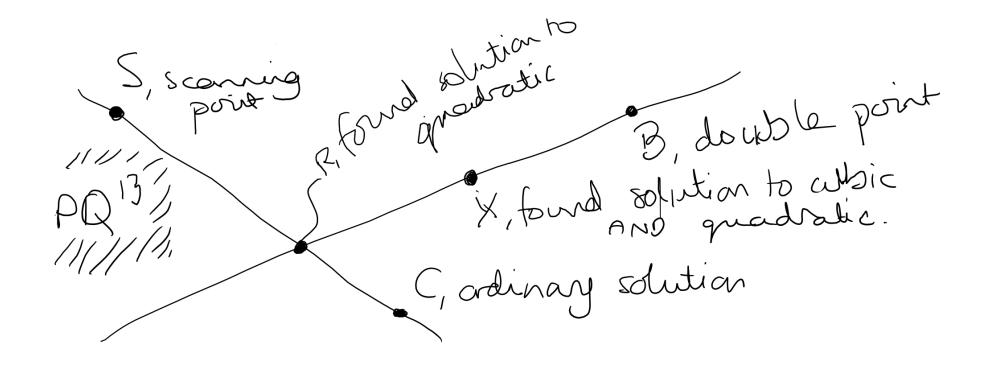
Points which are solutions of multiplicity two. All partial derivatives of the surface vanish there, eg (x,y)=(0,0) of the curve

$$(x^2 + y^2 + a^2)^2 - 4a^2x^2 - a^4 = 0$$



B is a double point of the quadratic and the cubic

Method



- ullet Every solution to quadratic R lies on some line SC
- ullet B-L is double point of quadratic $\Rightarrow RB$ in quadratic
- ullet Every solution X lies on a line between some R and B.

The Nitty-Gritty

$$\begin{split} Q_1 &= \Gamma - \Sigma + \Lambda S_{Q_1}, \\ Q_2 &= \Gamma + \Lambda S_{Q_2}, \\ Q_3 &= \Gamma + \Sigma + \Lambda S_{Q_3}, \\ U_1 &= -\Gamma - \Sigma + \Lambda S_{U_1}, \\ U_2 &= -\Gamma + \Lambda S_{U_2}, \\ U_3 &= -\Gamma + \Sigma + \Lambda S_{U_3}, \\ D_1 &= -\Gamma - \Sigma + \Lambda S_{D_1}, \\ D_2 &= -\Gamma + \Lambda S_{D_2}, \\ D_3 &= -\Gamma + \Sigma + \Lambda S_{D_3}, \\ L_1 &= -3\Gamma - \Sigma + \Lambda S_{L_1}, \\ L_2 &= -3\Gamma + \Lambda S_{L_2}, \\ L_3 &= -3\Gamma + \Sigma + \Lambda S_{L_3}, \\ E_1 &= 3\Gamma - \Sigma + \Lambda S_{E_1}, \\ E_2 &= 3\Gamma + \Lambda S_{E_2}, \\ E_3 &= 3\Gamma + \Sigma + \Lambda S_{E_3}, \\ N_1 &= 3\Gamma + \Lambda S_{N_1}, \\ N_2 &= 3\Gamma + \Lambda S_{N_2}, \\ N_3 &= 3\Gamma + \Lambda S_{N_3}, \end{split}$$

$$\begin{split} &\Gamma = c(R,R,R) + r\delta_{c(B,R,R),0}\delta_{c(R,R,R),0}, \\ &\Sigma = (-3c(B,R,R) + t\delta_{c(B,R,R),0}\delta_{c(R,R,R),0}) \\ &(q(S,S) + a\delta_{q(S,S),0}\delta_{q(C,S),0}), \\ &\Lambda = (-3c(B,R,R) + t\delta_{c(B,R,R),0}\delta_{c(R,R,R),0}) \\ &(-2q(C,S) + b\delta_{q(S,S),0}\delta_{q(C,S),0}). \\ &q(P,P') := \sum_{i=1}^{3} \left(Q_{i}Q'_{i} - 2U_{i}U'_{i}' + D_{i}D'_{i}' - L_{i}L'_{i}' + E_{i}E'_{i}E''_{i}'' + N_{i}N'_{i}N''_{i}''\right). \\ &c(P,P',P'') := \sum_{i=1}^{3} \left(6Q_{i}Q'_{i}Q'_{i}'' + 3U_{i}U'_{i}U''_{i}'' + 3D_{i}D'_{i}D''_{i}'' + 2L_{i}L'_{i}L'_{i}'' + E_{i}E'_{i}E''_{i}'' + N_{i}N'_{i}N''_{i}''\right). \\ &S_{Q_{3}} = \frac{1}{2} \left[-2S_{Q_{1}} - 2S_{Q_{2}} + \sum_{i=1}^{3} (S_{D_{i}} + S_{N_{i}}) \right], \\ &S_{U_{3}} = -\left[S_{U_{1}} + S_{U_{2}} + \sum_{i=1}^{3} (2S_{D_{i}} + S_{N_{i}}) \right], \\ &S_{L_{3}} = -\frac{1}{2} \left[2S_{L_{1}} + 2S_{L_{2}} + 3\sum_{i=1}^{3} (S_{D_{i}} + S_{N_{i}}) \right], \\ &S_{E_{3}} = -S_{E_{1}} - S_{E_{2}} + \sum_{i=1}^{3} (3S_{D_{i}} + 2S_{N_{i}}). \end{split}$$

Solution Space

Is called a projective *variety*, i.e. *not* a manifold (in \mathbb{Q} anyway, but also there are singular cases of lines within planes where the dimensionality decreases).

Over-parameterisation in terms of 18 integers

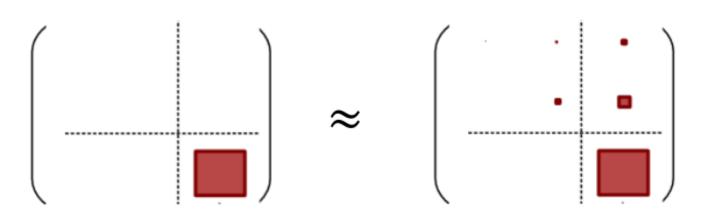
$$S_{Q_1}, S_{Q_2}, S_{U_1}, S_{U_2}, S_{D_1}, S_{D_2}, S_{D_3}, S_{L_1}, S_{L_2}, S_{E_1}, S_{E_2},$$

 $S_{N_1}, S_{N_2}, S_{N_3}, a, b, r, t \in \mathbb{Q}$

It is at most 11-dimensional. $S \cdot C = S \cdot B = 0$. An inverse (S = T, a = 0, b = 1, r = 0, t = 1), was checked against 21 549 920 all Anomaly-free Atlas solns.

Sol: $X=Y_3+t(B_3-L_3)$, $t\in\mathbb{Q}$

$$\mathcal{L} = Y_t \overline{Q_{3L}'} H t_R' + Y_b \overline{Q_{3L}'} H^c b_R' + Y_\tau \overline{L_{3L}'} H^c \tau_R' + H.c.,$$



A Simple Z' Model

BCA, Davighi, 1809.01158: Add complex SM-singlet scalar 'flavon' $\theta_{X\neq 0}$ which breaks gauged $U(1)_X$:

$$SU(3)\times SU(2)_L\times U(1)_Y\times U(1)_X\\ \langle\theta\rangle\sim \text{Several TeV}\\ SU(3)\times SU(2)_L\times U(1)_Y\\ \langle H\rangle\sim \text{246 GeV}\\ SU(3)\times U(1)_{em}$$

- SM fermion content
- Zero X charges for first two generations
- Expect $SM \times U(1)_X$ to be subsumed in semisimple model
- Still worry about anomaly cancellation

$$\mathcal{L}_{X\psi} = g_{X} \left(\frac{1}{6} \overline{\mathbf{u}_{L}} \Lambda^{(u_{L})} \gamma^{\rho} \mathbf{u}_{L} + \frac{1}{6} \overline{\mathbf{d}_{L}} \Lambda^{(d_{L})} \gamma^{\rho} \mathbf{d}_{L} - \frac{1}{2} \overline{\mathbf{n}_{L}} \Lambda^{(n_{L})} \gamma^{\rho} \mathbf{n}_{L} - \frac{1}{2} \overline{\mathbf{e}_{L}} \Lambda^{(e_{L})} \gamma^{\rho} \mathbf{e}_{L} + \frac{2}{3} \overline{\mathbf{u}_{R}} \Lambda^{(u_{R})} \gamma^{\rho} \mathbf{u}_{R} - \frac{1}{3} \overline{\mathbf{d}_{R}} \Lambda^{(d_{R})} \gamma^{\rho} \mathbf{d}_{R} - \overline{\mathbf{e}_{R}} \Lambda^{(e_{R})} \gamma^{\rho} \mathbf{e}_{R} \right) Z_{\rho}',$$

$$\Lambda^{(I)} \equiv V_{I}^{\dagger} \xi V_{I}, \qquad \xi = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Z' couplings, $I \in \{u_L, d_L, e_L, \nu_L, u_R, d_R, e_R\}$

A simple limiting case

$$V_{u_R} = V_{d_R} = V_{e_R} = 1$$

for simplicity and the ease of passing bounds.

$$V_{d_L} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & -\sin \theta_{23} \\ 0 & \sin \theta_{23} & \cos \theta_{23} \end{pmatrix}, \qquad V_{e_L} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix},$$

$$\Rightarrow V_{u_L} = V_{d_L} V_{CKM}^\dagger$$
 and $V_{
u_L} = V_{e_L} U_{PMNS}^\dagger$.

Important Z' Couplings

$$g_X \begin{bmatrix} \frac{1}{6} (\overline{d_L} \ \overline{s_L} \ \overline{b_L}) \begin{pmatrix} 0 & 0 & 0 \\ 0 & \sin^2 \theta_{23} & \frac{1}{2} \sin 2\theta_{23} \\ 0 & \frac{1}{2} \sin 2\theta_{23} & \cos^2 \theta_{23} \end{pmatrix} Z' \begin{pmatrix} d_L \\ s_L \\ b_L \end{pmatrix} +$$

$$-\frac{1}{2}(\overline{e_L} \ \overline{\mu_L} \ \overline{\tau_L}) \left(\begin{array}{ccc} 0 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 0 \end{array} \right) \not Z' \left(\begin{array}{c} e_L \\ \mu_L \\ \tau_L \end{array} \right) \right|$$

Z-Z' mixing

Because $Y_3(H) = 1/2$, $B - W^3 - X$ bosons mix:

$$\mathcal{M}_{N}^{2} = \frac{1}{4} \begin{pmatrix} g'^{2}v^{2} & -gg'v^{2} & g'g_{X}v^{2} \\ -gg'v^{2} & g^{2}v^{2} & -gg_{X}v^{2} \\ g'g_{X}v^{2} & -gg_{X}v^{2} & 4g_{X}^{2}\langle\theta\rangle^{2} \left(1 + \frac{\epsilon^{2}}{4}\right) \end{pmatrix} \frac{-B_{\mu}}{-(X)_{\mu}}$$

- $v \approx 246$ GeV is SM Higgs VEV,
- $\langle \theta \rangle \sim \text{TeV}$. $M_{Z'} = g_X \langle \theta \rangle$.
- $g_X = U(1)_X$ gauge coupling
- $\epsilon \equiv v/\langle \theta \rangle \ll 1$

Z-Z' mixing angle

$$\sin \alpha_z \approx \frac{g_X}{\sqrt{g^2 + g'^2}} \left(\frac{M_Z}{M_Z'}\right)^2 \ll 1.$$

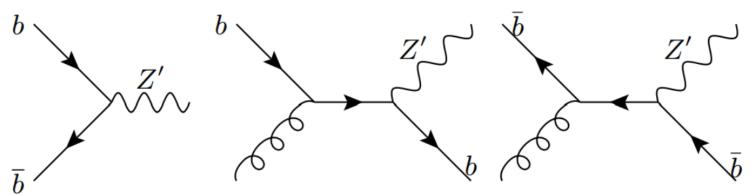
This gives small non-flavour universal couplings to the Z boson propotional to g_X and:

$$Z_{\mu} = \cos \alpha_z \left(-\sin \theta_w B_{\mu} + \cos \theta_w W_{\mu}^3 \right) + \sin \alpha_z X_{\mu},$$

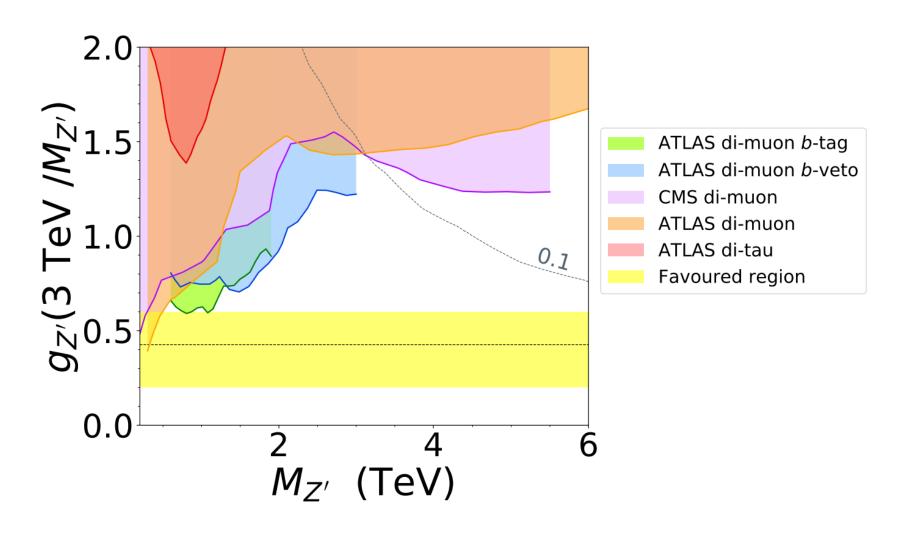
Z' Decay Modes

Mode	BR	Mode	BR	Mode	BR
\overline{t}	0.42	$b\overline{b}$	0.12	$ uar{ u}'$	0.08
$\mu^+\mu^-$	80.0	$\mid au^+ au^- \mid$	0.30	other f_if_j	$\sim \mathcal{O}(10^{-4})$

LHC Z' Production:

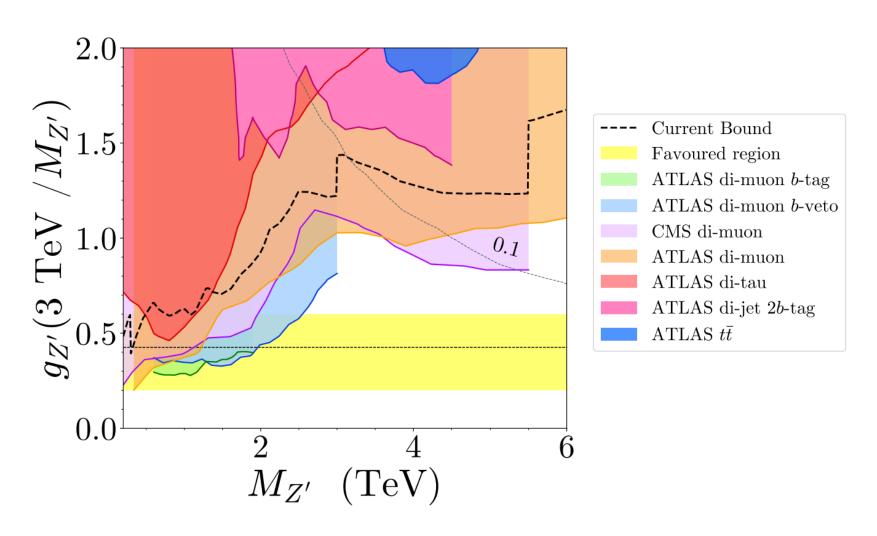


Z' Searches³



³BCA, Banks, 2111.06691

HL-LHC sensitivity⁴



⁴BCA, Banks, 2111.06691

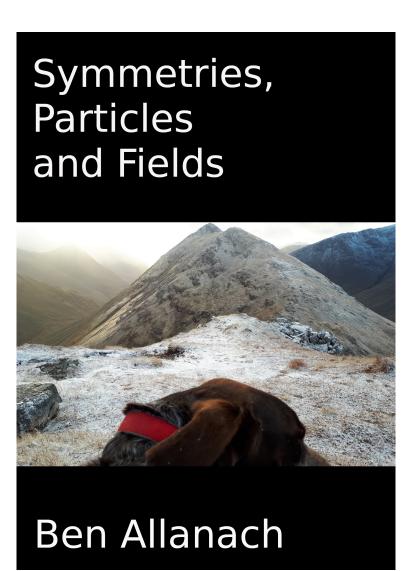
Why $\bar{b}s\mu^+\mu^-$?

If we take these B-anomalies seriously, we may ask: why are we seeing the first BSM flavour changing effects particularly in the $b \to s \mu^+ \mu^-$ transition, not another one?

Perhaps it's because, in hindsight:

- The largest BSM flavour effects are in heavier generations
- ullet We have many more bs than ts, particularly in LHCb
- Leptons in final states are good experimentally but not (yet) τ s: they are too difficult!

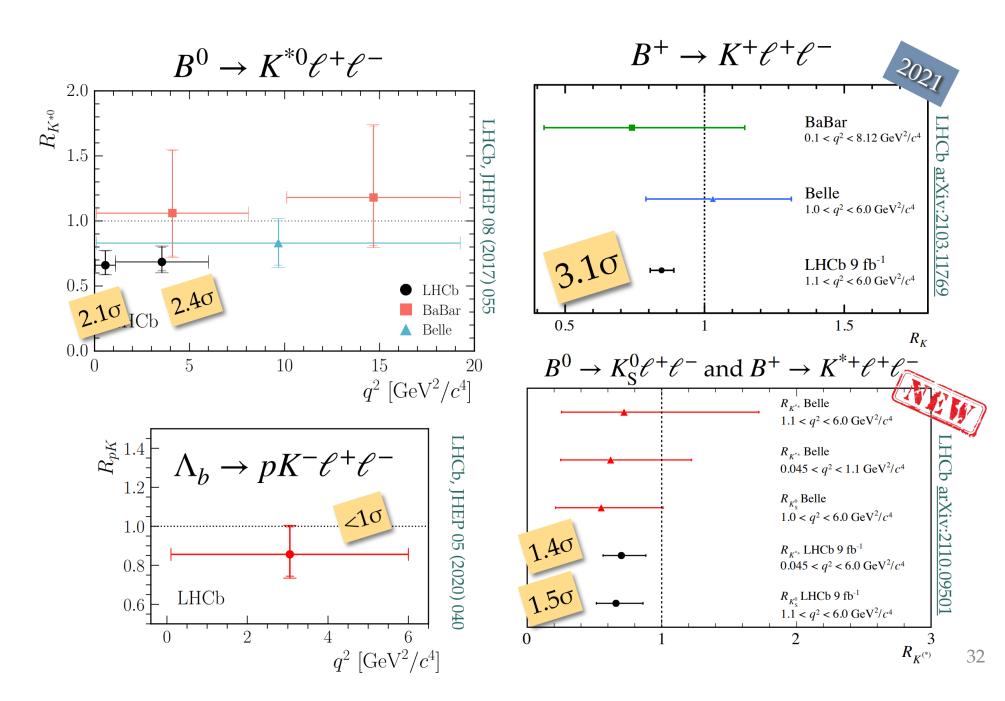
On amazon.com~€20



Summary

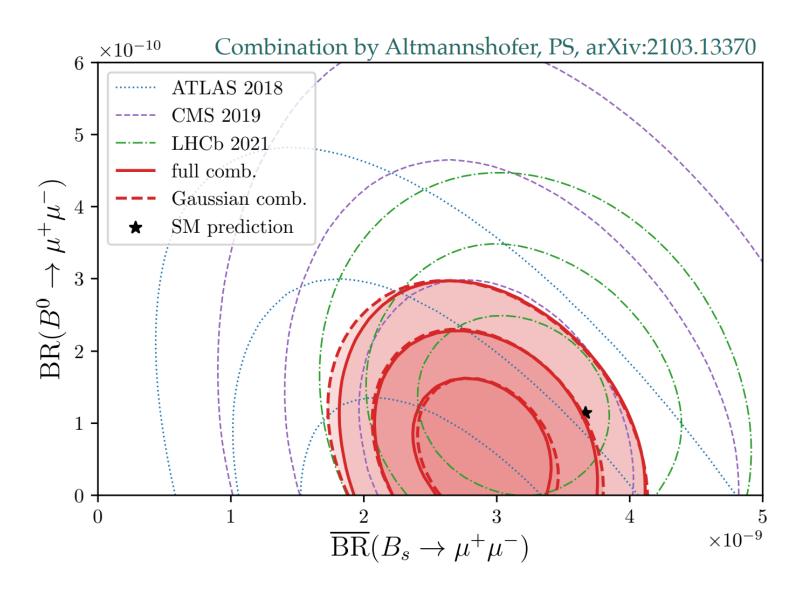
- The $b \to s \mu^+ \mu^-$ anomalies look very interesting from a BSM point of view: a consistent picture is emerging.
- Independent check awaited from Belle II in Japan in the coming three years or so: $e^+e^-(10.58 \text{ GeV}) \rightarrow \Upsilon(4s) \rightarrow$ oodles of B mesons.
- ullet Tree-level explanations: leptoquarks and Z's.
- In case a Z' is found directly, measuring its couplings may give us an experimental handle on the fermion mass puzzle.

Backup

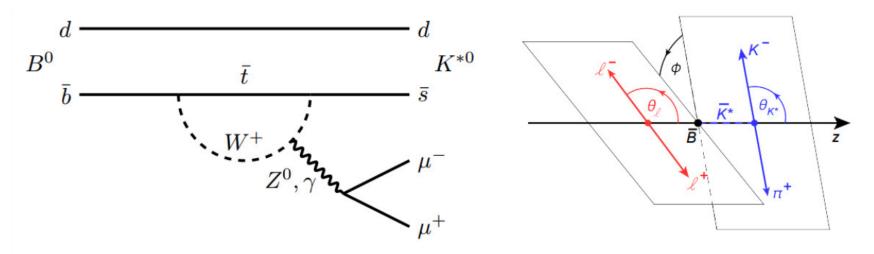


Stolen from Capdevila et al, Flavour Anomaly Workshop '21

$BR(B_s ightarrow \mu^+\mu^-)$: $B_s = (ar{b}s)$, $B^0 = (ar{b}d)$



$B^0 \to K^{*0} (\to K^+ \pi^-) \mu^+ \mu^-$



Decay fully described by three helicity angles $\vec{\Omega}=(\theta_\ell,\theta_K,\phi)$ and $q^2=m_{\mu\mu}^2$

$$\frac{1}{\mathrm{d}(\Gamma + \bar{\Gamma})/\mathrm{d}q^2} \frac{\mathrm{d}^3(\Gamma + \bar{\Gamma})}{\mathrm{d}\bar{\Omega}} = \frac{9}{32\pi} \left[\frac{3}{4} (1 - F_{\mathrm{L}}) \sin^2 \theta_K + F_{\mathrm{L}} \cos^2 \theta_K + \frac{1}{4} (1 - F_{\mathrm{L}}) \sin^2 \theta_K \cos 2\theta_\ell \right.$$

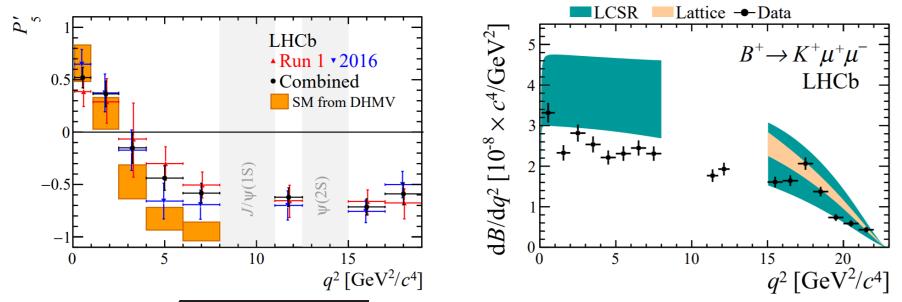
$$- F_{\mathrm{L}} \cos^2 \theta_K \cos 2\theta_\ell + S_3 \sin^2 \theta_K \sin^2 \theta_\ell \cos 2\phi$$

$$+ S_4 \sin 2\theta_K \sin 2\theta_\ell \cos \phi + S_5 \sin 2\theta_K \sin \theta_\ell \cos \phi$$

$$+ \frac{4}{3} A_{\mathrm{FB}} \sin^2 \theta_K \cos \theta_\ell + S_7 \sin 2\theta_K \sin \theta_\ell \sin \phi$$

$$+ S_8 \sin 2\theta_K \sin 2\theta_\ell \sin \phi + S_9 \sin^2 \theta_K \sin^2 \theta_\ell \sin 2\phi \right]$$

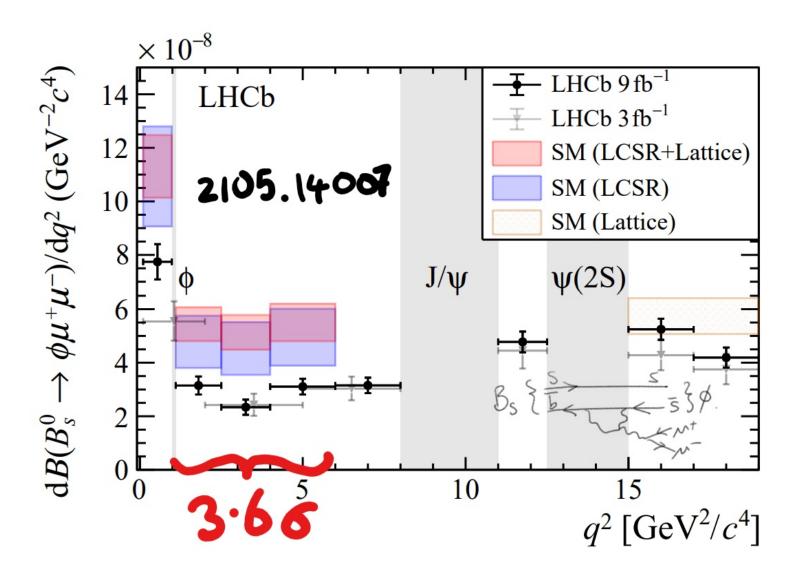
P_5'



 $P_5' = S_5/\sqrt{F_L(1-F_L)}$, leading form factor uncertainties cancel 5

⁵LHCb, 2003.04831

$B_s \to \phi \mu^+ \mu^-$: $\phi = (s\bar{s})$



Theory: uncertainties

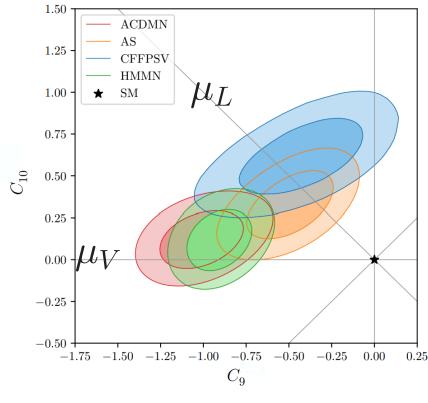
	parametric	form factors	non-local
			MEs
$BR(B \to Mll)$	yes	large	large
angular	no	small	large
$BR(B_s \to ll)$	yes	small	no
LFU	no	tiny	no

- ullet Parametric uncertainties (eg V_{ts}) easy to deal with
- Large theory uncertainties are taken into account in fits,
 but one can argue about them

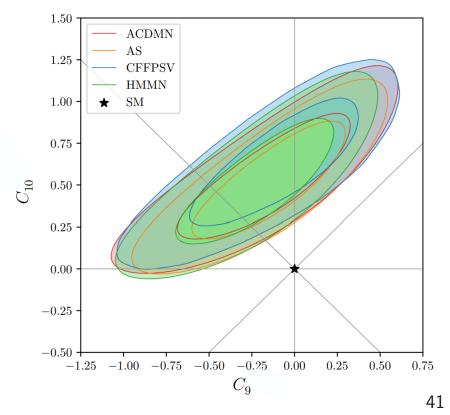
Fits

Alguero et al, 2104.08921; Altmannshofer, Stangl, flavio 2103.13370; Ciuchini et al, HEPfit 2011.01212; Hurth et al, superIso 2104.10058

$$\mathcal{L} = N[\underline{C_9}(\bar{b}_L \gamma^{\mu} s_L)(\bar{\mu} \gamma_{\mu} \mu) + \underline{C_{10}}(\bar{b}_L \gamma^{\mu} s_L)(\bar{\mu} \gamma^5 \gamma_{\mu} \mu)] + H.c.$$



global fit



fit to LFU observables + $B_{\rm S} o \mu \mu$

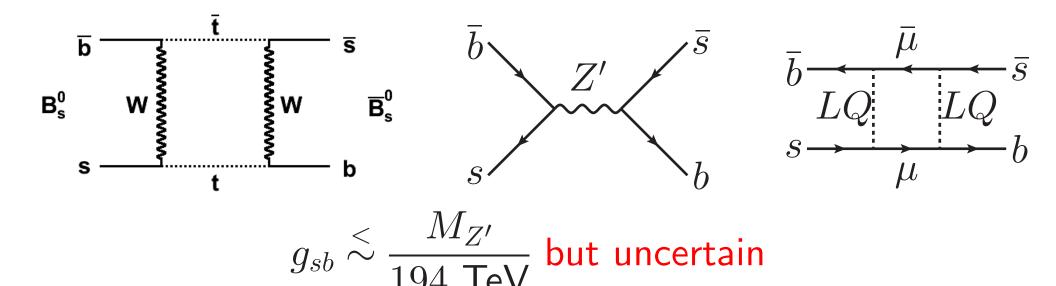
Y_3 Consequences

- Flavour changing TeV-scale Z^\prime to do NCBAs: couples dominantly to EW eigenstates of quarks and leptons of the third family
- First two fermion families massless at renormalisable level
- Their masses and fermion mixings generated by small non-renormalisable operators

This explains the hierarchical heaviness of the third family and small CKM angles

$B_s - \bar{B}_s$ Mixing

Measurement pretty much agrees with SM calculations.



from QCD sum rules and lattice⁶. Weaker on LQs.

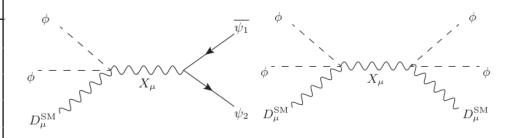
$$M_{Z'}pprox 31~{
m TeV} imes \sqrt{g_{sb}g_{\mu\mu}}, \qquad M_{LQ}pprox 31~{
m TeV} imes \sqrt{g_{s\mu}g_{b\mu}}$$
 $M_{LQ} \approx 31~{
m TeV} imes \sqrt{g_{s\mu}g_{b\mu}}$

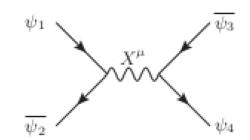
$B/{\sf EW}$ Observables

 $\mathsf{SMEFT}(M_{Z'}) \to \mathsf{smelli} \to \mathsf{WET}(M_W) \to \mathsf{obs}(m_B)$

In units of g_X^2/M_X^2 :

WC	value	WC	value
C_{ll}^{2222}	$-\frac{1}{8}$	$(C_{lq}^{(1)})^{22ij}$	$\frac{1}{12} A_{\xi ij}^{(d_L)}$
$(C_{qq}^{(1)})^{ijkl}$	$\Lambda_{\xi ij}^{(d_L)} \Lambda_{\xi kl}^{(d_L)} \frac{\delta_{ik} \delta_{jl} - 2}{72}$	C_{ee}^{33333}	$-\frac{1}{2}$
C_{uu}^{33333}	$-\frac{2}{9}$	C_{dd}^{3333}	$-\frac{1}{18}$
C_{eu}^{33333}	2 3 2 9	C_{ed}^{33333}	$-\frac{1}{3}$
$(C_{ud}^{(1)})^{3333}$	$\frac{2}{9}$	C_{le}^{2233}	$-\frac{1}{2}$
C_{lu}^{2233}	<u>1</u> 3	C_{ld}^{2233}	$-\frac{1}{6}$
C_{qe}^{ij33}	$\frac{1}{6}\Lambda_{\xi ij}^{(d_L)}$	$(C_{qu}^{(1)})^{ij33}$	$-\frac{1}{9}\Lambda_{\xi \ ij}^{(d_L)}$
$(C_{qd}^{(1)})^{ij33}$	$\frac{1}{18}\Lambda_{\xi ij}^{(d_L)}$	$(C_{\phi l}^{(1)})^{22}$	$\frac{1}{4}$
$(C_{\phi q}^{(1)})^{ij}$	$\begin{array}{c} \frac{1}{6} A_{\xi}^{(d_L)} \\ \frac{1}{18} A_{\xi}^{(d_L)} \\ -\frac{1}{12} A_{\xi}^{(d_L)} \\ ij \end{array}$	$(C_{\phi l}^{(1)})^{22}$ $C_{\phi e}^{33}$	$\frac{1}{2}$
$(C_{\phi q}^{(1)})^{ij}$ $C_{\phi u}^{33}$	$-\frac{1}{3}$	$C_{\phi d}^{33}$	$\frac{1}{6}$
$C_{\phi D}$	$-\frac{1}{2}$	$C_{\phi\Box}$	$-\frac{1}{8}$





smelli observables

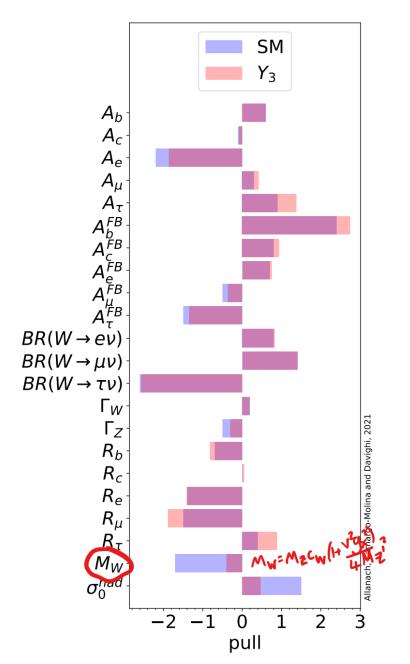
- 167 quarks: P_5' , $BR(B_s \to \mu^+ \mu^-)$ and others with significant theory errors
- 21 LFU FCNCs: $R_K, R_{K^*}, B \rightarrow \text{di-tau decays}$
- 31 EWPOs from LEP not assuming lepton flavour universality

Theory uncertainties modelled as multi-variate Gaussians: approximated to be independent of new physics.

SM:

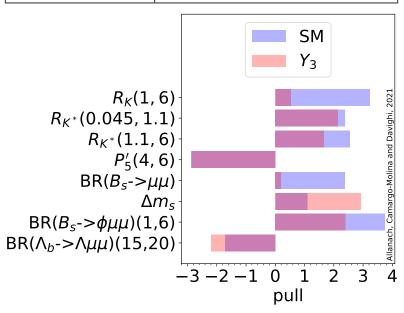
data set	χ^2	n	p-value
quarks	221.6	167	.003
LFU FCNCs	35.3	21	.026
EWPOs	35.7	31	.26
global	292.6	219	.00065

Global Fits $M_{Z^\prime}=3$ TeV

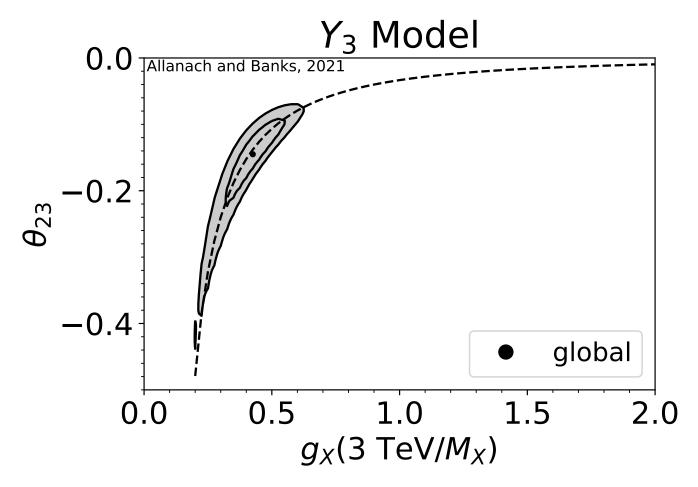


data set	χ^2	n	p-value
quarks	221.6	167	.003
LFU FCNCs	35.3	21	.026
EWPOs	35.7	31	.26
global	292.6	219	.00065

data set	χ^2	n	p-value
quarks	192.5	167	.071
LFU FCNCs	21.0	21	.34
EWPOs	36.0	31	.17
global	249.5	219	.064



TFHM Fit, 95% CL



Relies on: smelli-2.2.0 (Aebischer, Kumar, Stangl, Straub, 1810.07698), flavio-2.2.0 (Straub, 1810.08132), Wilson (Aebischer *et al*, 1712.05298)

$Z' ightarrow \mu \mu$ ATLAS 13 TeV 139 fb $^{-1}$

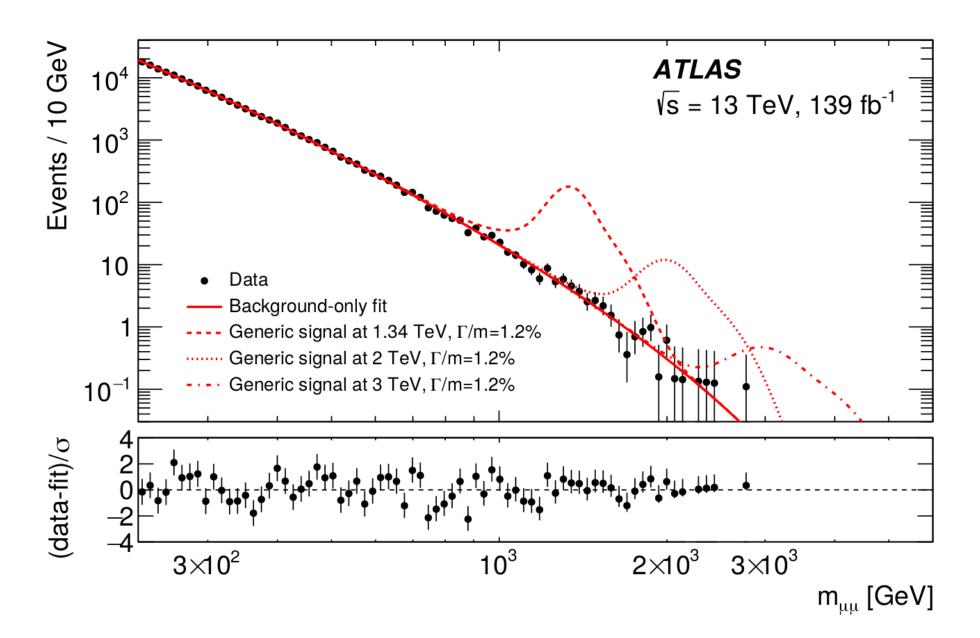
ATLAS analysis: look for two track-based isolated μ , $p_T > 30$ GeV. One reconstructed primary vertex. Keep only highest scalar sum p_T pair⁷

$$m_{\mu_1\mu_2}^2 = (p_1^{\mu} + p_2^{\mu}) (p_{1\mu} + p_{2\mu})$$

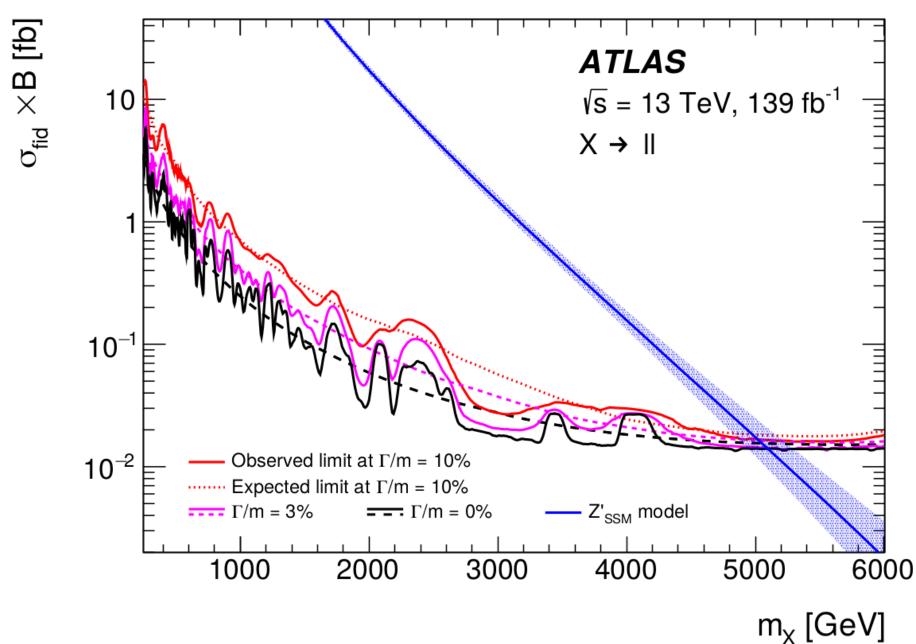
CMS also have released⁸ a 139 fb $^{-1}$ analysis.

⁷1903.06248

^{82103.02708}



ATLAS l^+l^- limits



CDF II M_W

As already noted, Z - Z' mixing implies

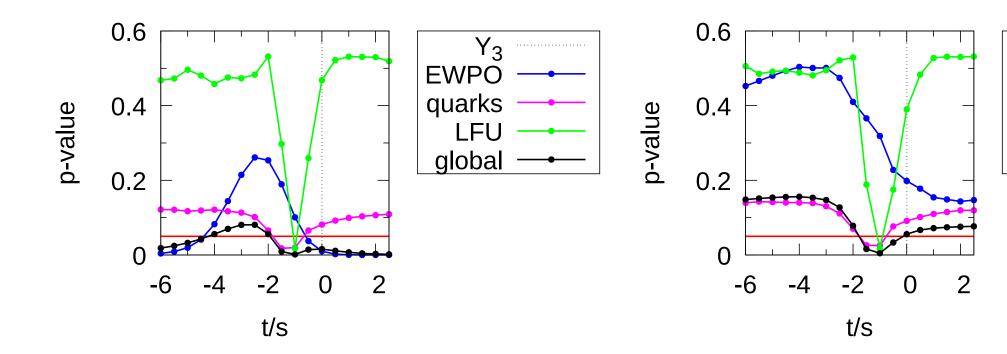
$$M_W = \rho_0 M_Z \cos \hat{\theta}_W$$

where

$$\rho_0(SM) = (1.01019 \pm 0.00009),$$

$$\rho_0(Y_3) \approx 1 + \frac{X_H^2 g_X^2}{g^2 + g'^2} \frac{M_Z^2}{M_{Z'}^2} > 1.$$

$$sY_3 + t(B_3 - L_3)$$

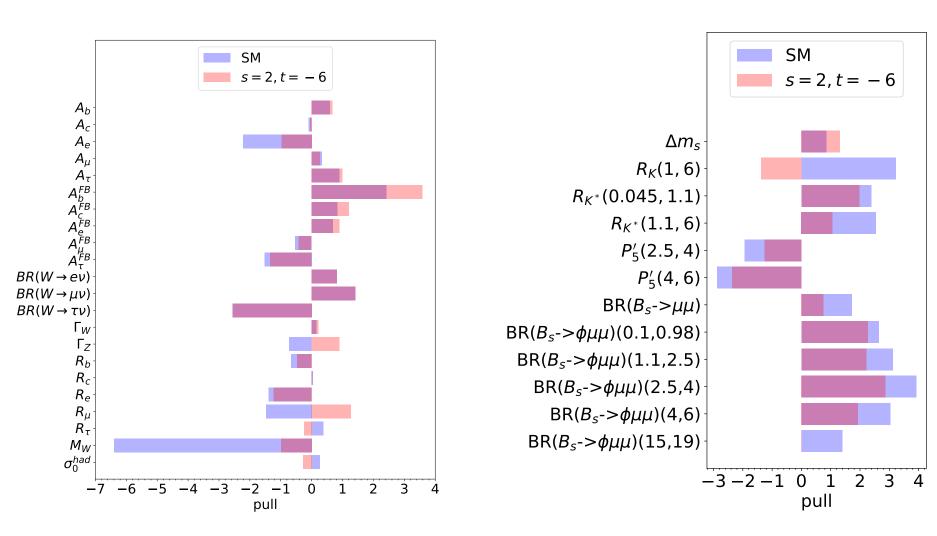


Left incl CDF II M_W , Right excl

BCA, Davighi, 2205.12252

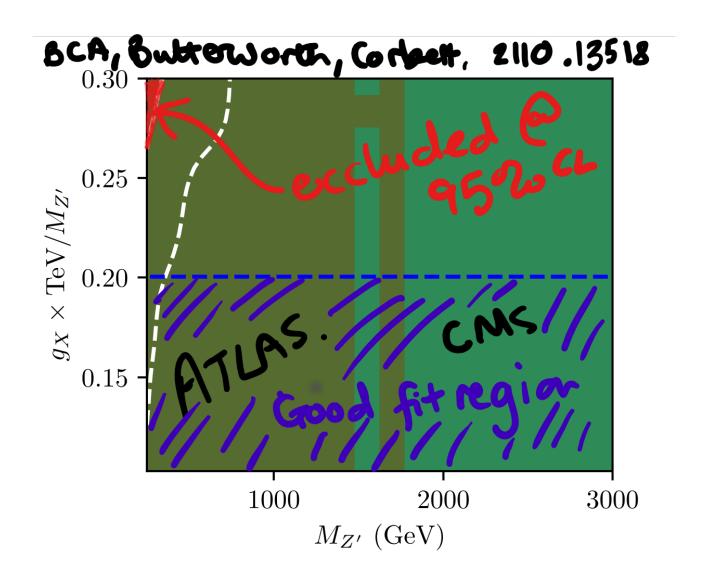
Pick $Y_3 - 3(B_3 - L_3)$ as a well fit example

Best-fit point: incl CDF M_W

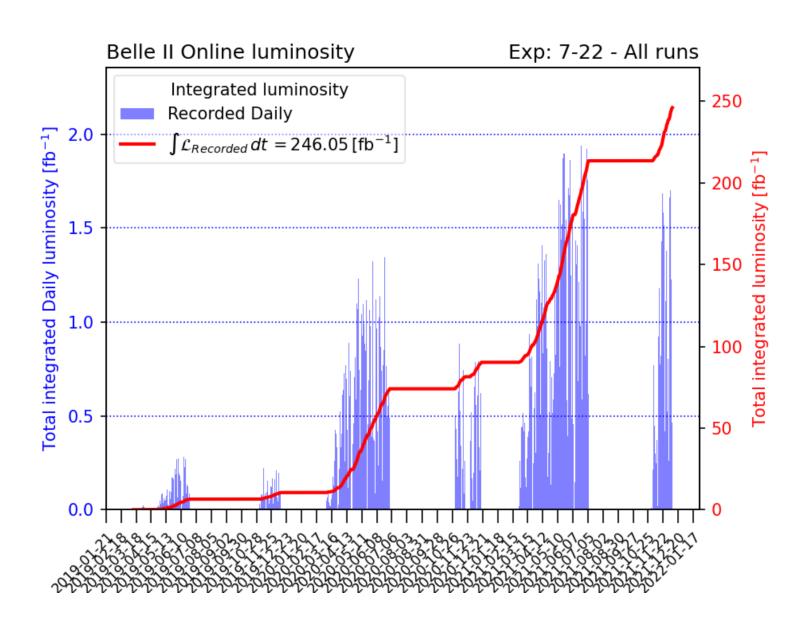


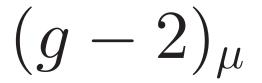
$$g_X = 0.021 \times 1 \; {\rm TeV}/M_{Z'}$$
, $\theta_{23} = -0.0191$, $p = .08$

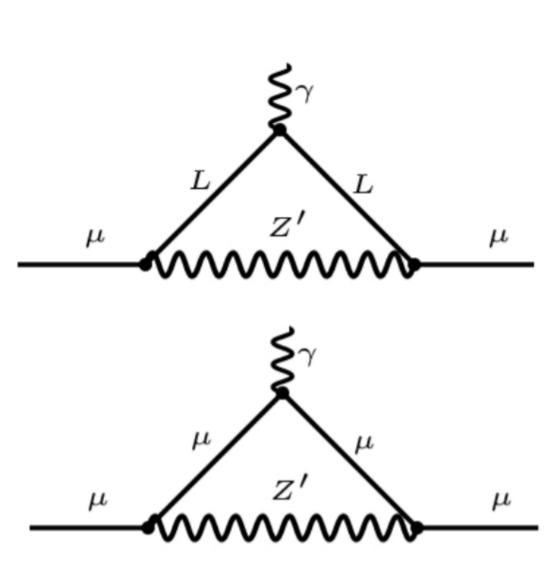
TFHM $Z' \to \mu^+ \mu^- + SM$ obs



1 fb⁻¹ $\approx 10^6 B\bar{B}$







Trident Neutrino Process

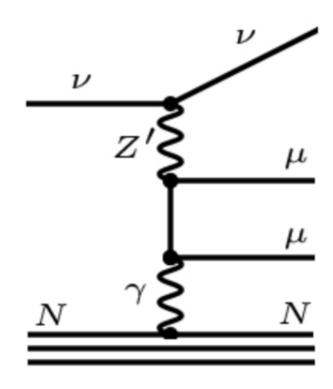
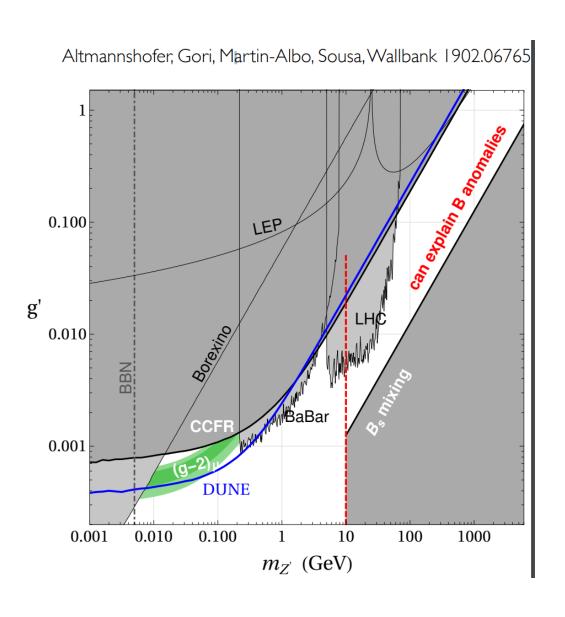
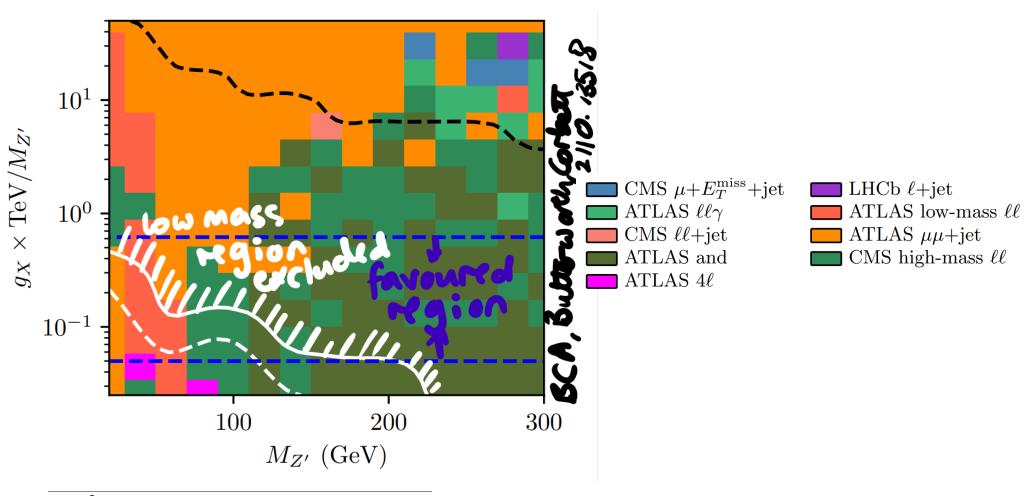


FIG. 10. Neutrino trident process that leads to constraints on the Z^{μ} coupling strength to neutrinos-muons, namely $M_{Z'}/g_{v\mu} \gtrsim 750 \text{ GeV}.$

Light Z' for $(g-2)_{\mu}$: $L_{\mu}-L_{ au}$



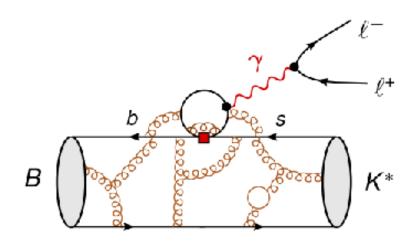
B_3-L_2 model's $^9Z^\prime$



⁹Bonilla, Modak, Srivastava, Valle, 1705.00915, Alonso, Cox, Han, Yanagida 1705.03858

Hadronic Uncertainties

► Hadronic effects like charm loop are photon-mediated ⇒ vector-like coupling to leptons just like C₉



- ► How to disentangle NP ↔ QCD?
 - ► Hadronic effect can have different q² dependence
 - ▶ Hadronic effect is lepton flavour universal ($\rightarrow R_K!$)

Wilson Coefficients c_{ij}^l

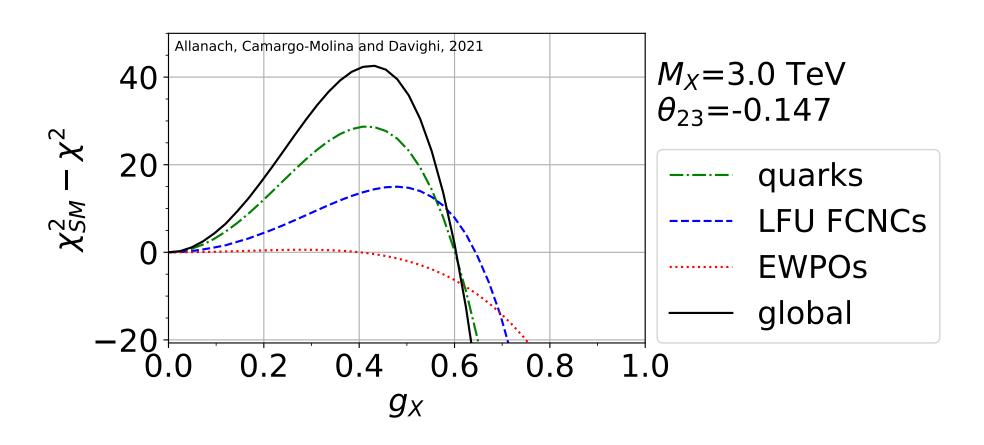
In SM, can form an EFT since $m_B \ll M_W$:

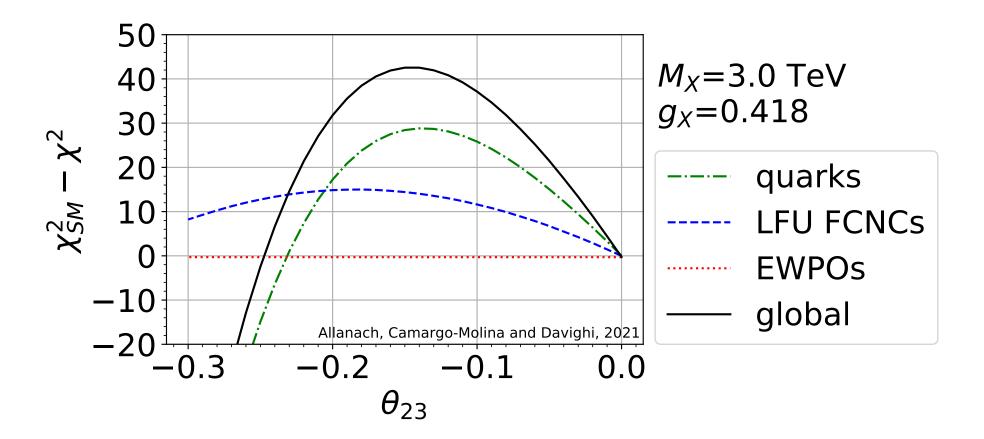
$$\mathcal{L}_{\text{eff}} = \frac{1}{(36 \text{ TeV})^2} c_{ij}^l (\bar{s}\gamma^\mu P_i b) (\bar{l}\gamma_\mu P_j l) \tag{1}$$

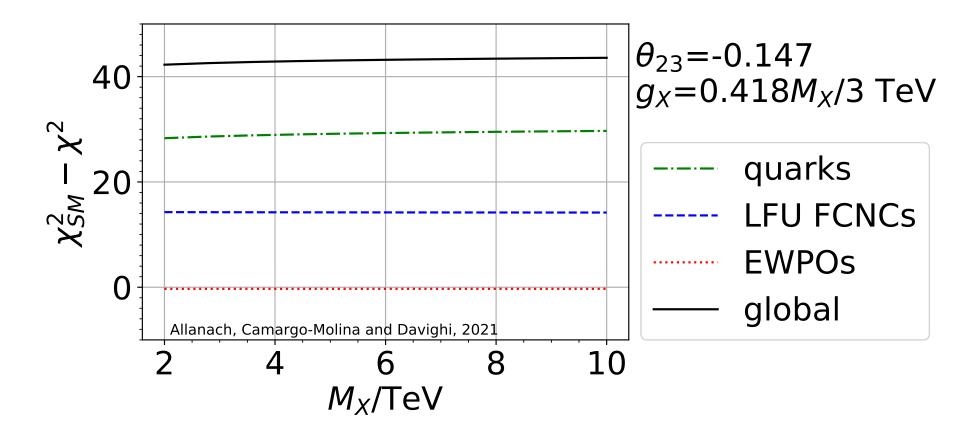
One loop weak interactions give $c_{ij}^l \sim \pm \mathcal{O}(1)$ in SM. $(1/36 \text{ TeV})^2 = V_{tb}V_{ts}^*\alpha/(4\pi v^2)$.

From now on, c_{ij}^l refer to beyond SM contribution.

TFHM Near best-fit point







Which Ones Work?

Options for a single BSM operator:

- c_{ij}^e operators fine for $R_{K^{(*)}}$ but are disfavoured by global fits including other observables.
- ullet c_{LR}^{μ} disfavoured: predicts enhancement in both R_K and R_{K^*}
- c_{RR}^{μ} , c_{RL}^{μ} disfavoured: they pull R_K and R_{K^*} in *opposite* directions.
- $c_{LL}^{\mu}=-1.06$ fits well globally¹⁰.

¹⁰D'Amico et al, 1704.05438; Aebischer et al 1903.10434.

Invisible Width of Z Boson

 $\Gamma_{\rm inv}^{\rm (exp)} = 499.0 \pm 1.5 \,\, {\rm MeV}, \, {\rm whereas} \,\, \Gamma_{\rm inv}^{\rm (SM)} = 501.44 \,\, {\rm MeV}.$

$$\Rightarrow \Delta\Gamma^{(\rm exp)} = \Gamma^{(\rm exp)}_{\rm inv} - \Gamma^{(\rm SM)}_{\rm inv} = -2.5 \pm 1.5 \ {\rm MeV}.$$

$$\mathcal{L}_{\bar{\nu}\nu Z} = -\frac{g}{2\cos\theta_w} \overline{\nu'_{Le}} \mathbb{Z} P_L \nu'_{Le}$$

$$-\overline{\nu'_{L\mu}} \left(\frac{g}{2\cos\theta_w} + \frac{5}{6} g_F \sin\alpha_z \right) \mathbb{Z} \nu'_{L\mu}$$

$$-\overline{\nu'_{L\tau}} \left(\frac{g}{2\cos\theta_w} - \frac{8}{6} g_F \sin\alpha_z \right) \mathbb{Z} \nu'_{L\tau}.$$

Deformed TFHM

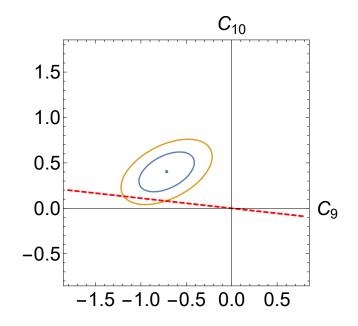
$$F_{Q'_{i}} = 0 F_{u_{R'_{i}}} = 0 F_{d_{R'_{i}}} = 0 F_{H} = -1/2$$

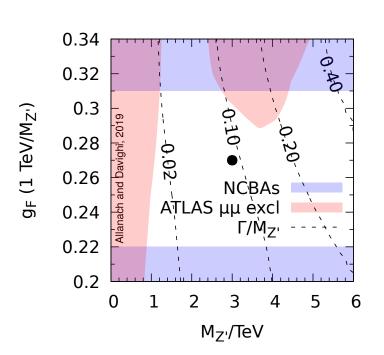
$$F_{e_{R'_{1}}} = 0 F_{e_{R'_{2}}} = 2/3 F_{e_{R'_{3}}} = -5/3$$

$$F_{L'_{1}} = 0 F_{L'_{2}} = 5/6 F_{L'_{3}} = -4/3$$

$$F_{Q'_{3}} = 1/6 F_{u'_{R3}} = 2/3 F_{d'_{R3}} = -1/3 F_{\theta} \neq 0$$

$$\mathcal{L} = Y_t \overline{Q_{3L}'} H t_R' + Y_b \overline{Q_{3L}'} H^c b_R' + H.c.,$$





Neutrino Masses

At dimension 5:

$$\mathcal{L}_{SS} = \frac{1}{2M} (L_3'^T H^c) (L_3' H^c),$$

but if we add RH neutrinos, then integrate them out

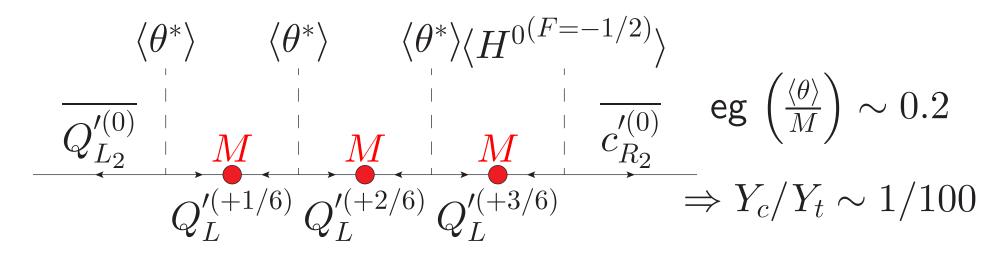
$$\mathcal{L}_{SS} = 1/2 \sum_{ij} (L_i' H^c) (M^{-1})_{ij} (L_j' H^c),$$

where now $(M^{-1})_{ij}$ may well have a non-trivial structure. If $(M^{-1})_{ij}$ are of same order, large PMNS mixing results.

Froggatt Neilsen Mechanism¹¹

A means of generating the non-renormalisable Yukawa terms, e.g. $X_{\theta} = 1/6$:

$$Y_c \overline{Q_{L2}^{\prime}}^{(F=0)} H^{(F=-1/2)} c_R^{\prime}^{(F=0)} \sim \mathcal{O}\left[\left(\frac{\langle \theta \rangle}{M}\right)^3 \overline{Q_{L2}^{\prime}} H c_R^{\prime}\right]$$



¹¹C Froggatt and H Neilsen, NPB**147** (1979) 277