

NEW RESULTS AND OPPORTUNITIES FOR DETECTING TERRESTRIALLY BOUND DARK MATTER

Harikrishnan Ramani
Stanford Institute for Theoretical Physics



Based on:

arXiv: 2110.06217 HR with A. Berlin, H. Liu, M. Pospelov

arXiv: 2108.05283 HR with D. Budker, C. Smorra, P. Graham, F. Schmidt-Kaler, S. Ulmer

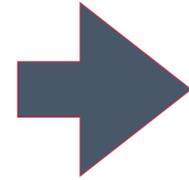
arXiv: 2012.03957 HR with M. Pospelov

OUTLINE

- ◆ Strongly Interacting Dark Relics & Traffic Jams
- ◆ Milli-charge Particle Detection with Ion Traps
- ◆ Dark-Standard Model bound states for a heavy Dark Photon

DARK RELICS

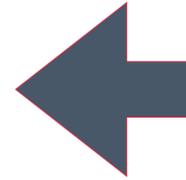
DARK MATTER



STABLE PARTICLE

DARK RELICS

DARK RELIC



STABLE PARTICLE

- ◆ Well motivated stable particles: Monopoles, axions, squarks, heavy quarks (KSVZ), gluinos (SUSY), “**Milli-charge**” Particles (mCPs)
- ◆ Robust prediction for relic fractions $f_\chi = \frac{\rho_\chi}{\rho_{\text{DM}}} \ll 1$
- ◆ The only way to access $M_\chi \gg \text{TeV}$?
- ◆ Use same concept for Detection?

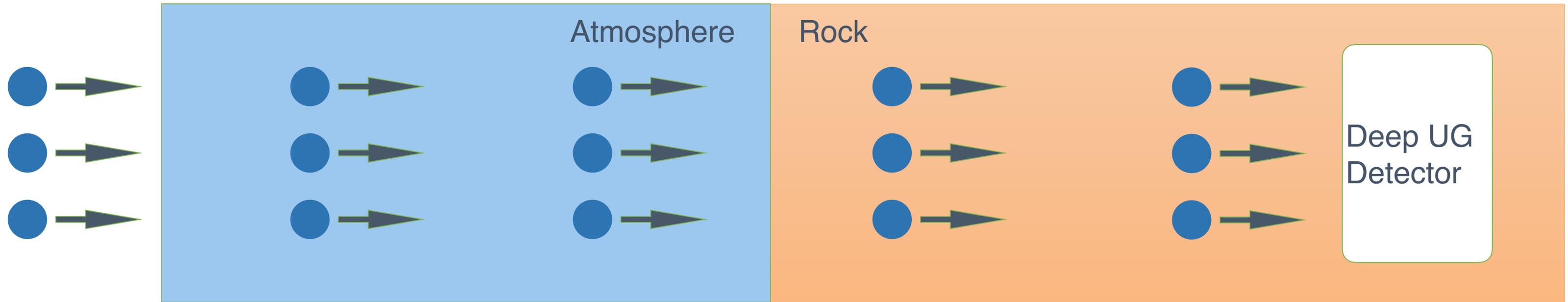
LARGE SM-RELIC INTERACTIONS

◆ Plethora of example models with large interactions with SM

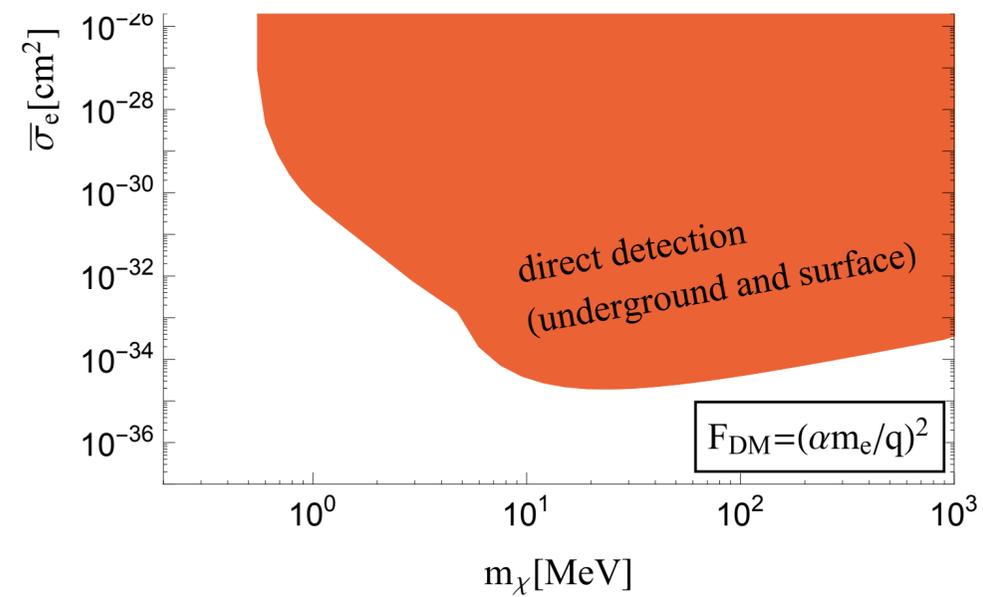
- Gluino/KSVZ/Squark Inspired “hybrid hadrons” (Qg or $Qq\bar{q}'$) [1801.01135](#) Luca, Mitridate, Redi, Smirnov, Strumia
- Sextaquark DM (uuddss bound state) [2007.10378](#) Farrar, Wang, Xu
- Large N Composites [1807.03788](#) Grabowska, Melia, Rajendran
- Milli-charges: Recent interest in explaining EDGES anomaly [1905.06348](#) Emken et al , [1908.06986](#) Liu et al
- Heavy Dark Photon Mediated [2110.06217](#) Berlin, Liu, Pospelov, Ramani

◆ Some really interesting consequences for Direct Detection

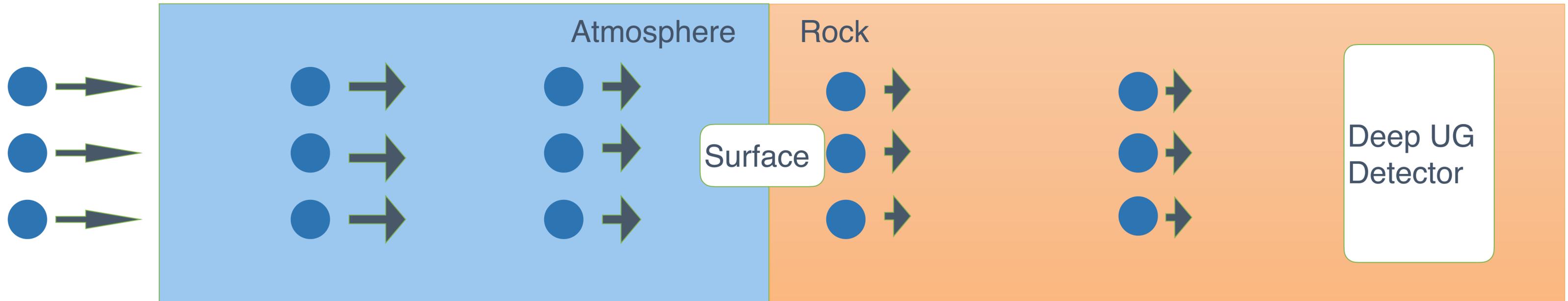
SMALL X-SECTION



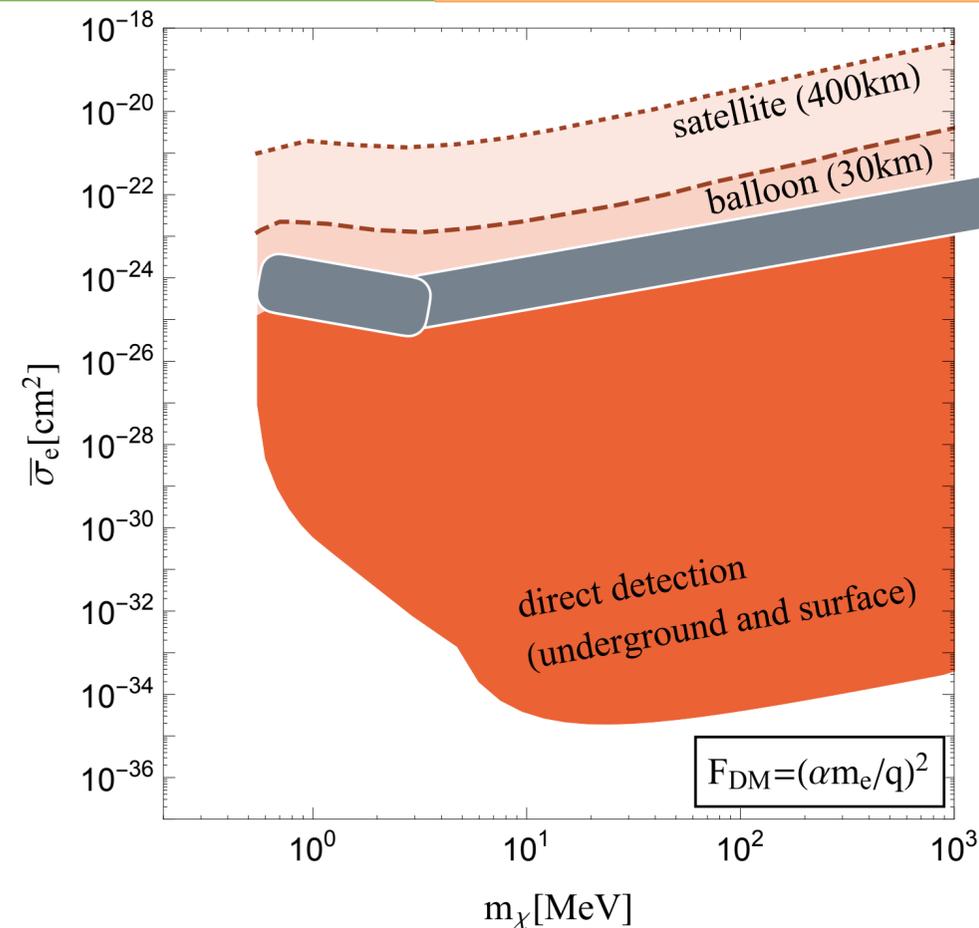
1905.06348 Emken et al



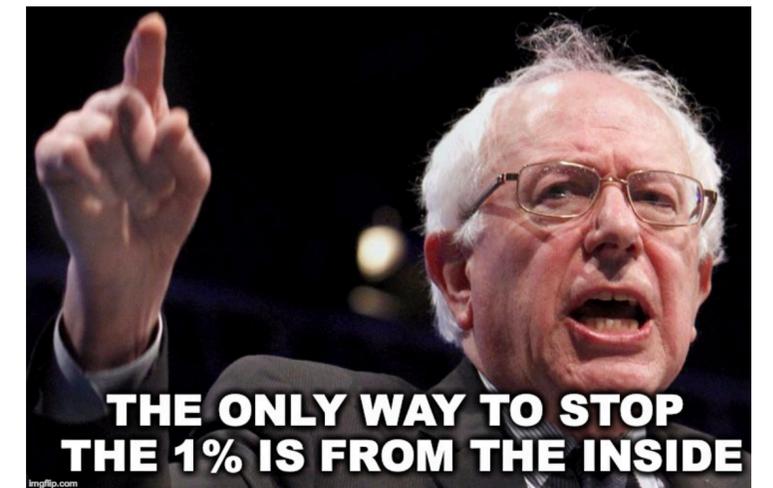
LARGE X-SECTION



- ◆ Reaches detector after thermalizing
- ◆ $KE=300$ Kelvin (26 meV)
- ◆ Current DD threshold : eV



1905.06348 Emken et al

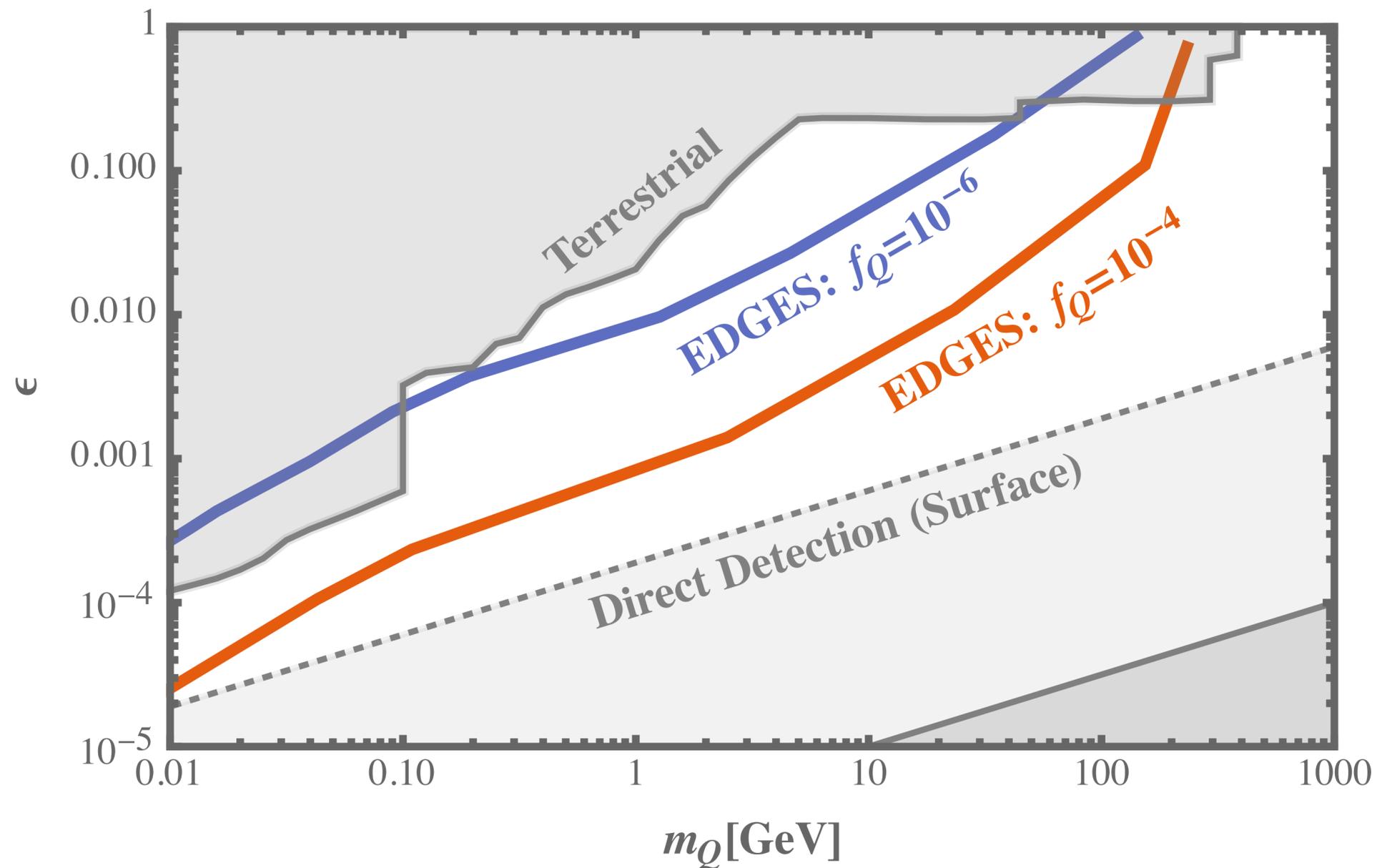


MILLICHARGE PARTICLES

- ◆ Particles with tiny electric charges: ϵe
- ◆ Simple models to write (with or without a dark photon)
- ◆ Charge quantization a century old mystery
- ◆ Predictions of explanation: monopoles and/or GUTs not observed yet
- ◆ Looked for in various experimental programs
- ◆ Recent resurgence due to EDGES anomaly

MCP PARAMETER SPACE

Adapted from: 1908.06986 Liu et al



KE smaller than threshold

Colliders/Terrestrial : no reach for small charge

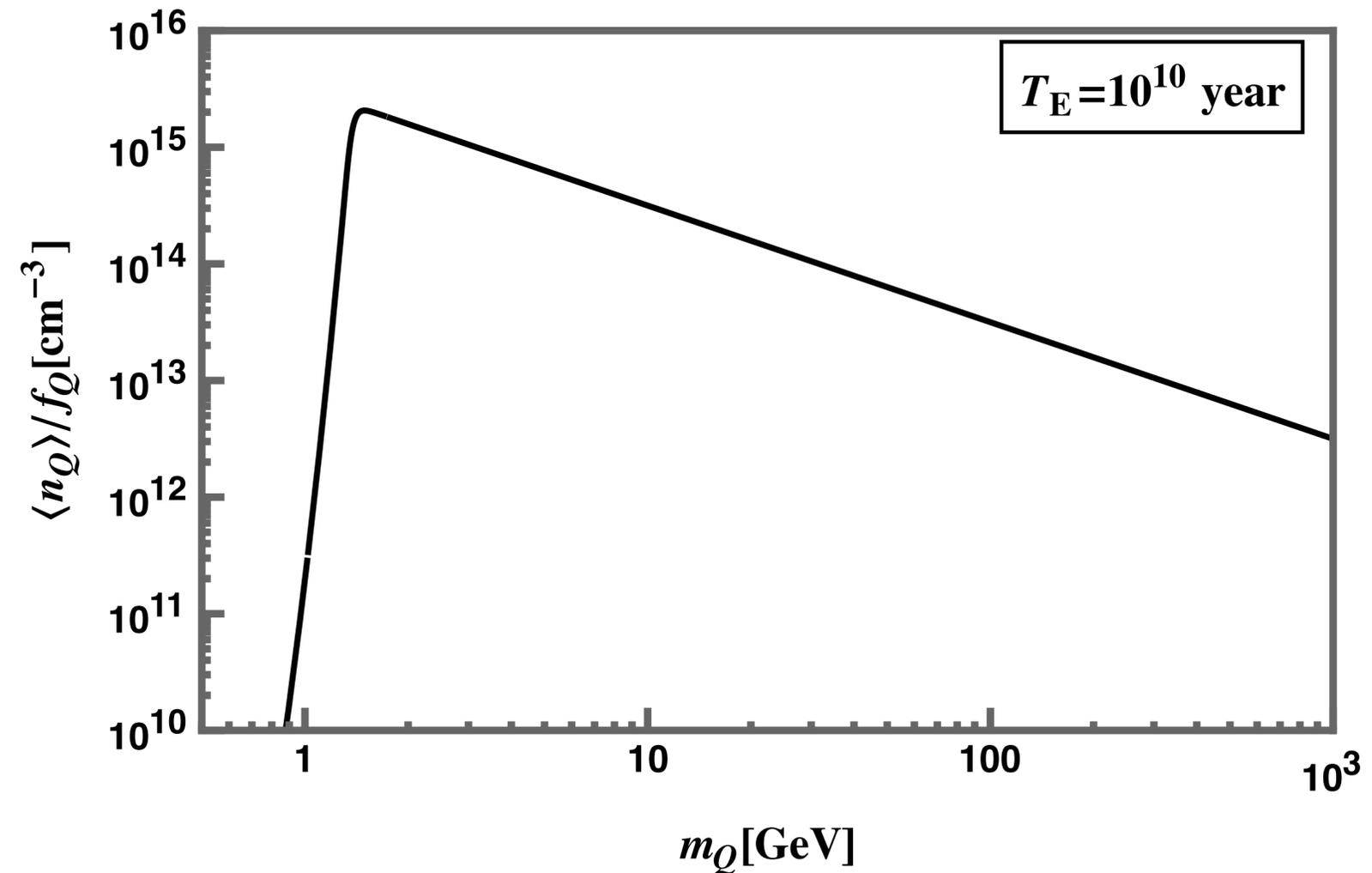
Direct Detection : no reach for large charge (Overburden blocks it)

TERRESTRIAL ABUNDANCE

- ◆ DM thermalizes, but stuck on Earth if $v_{\text{th}} < v_{\text{esc}}$
- ◆ Accumulation over the age of the Earth causes tremendous enhancement

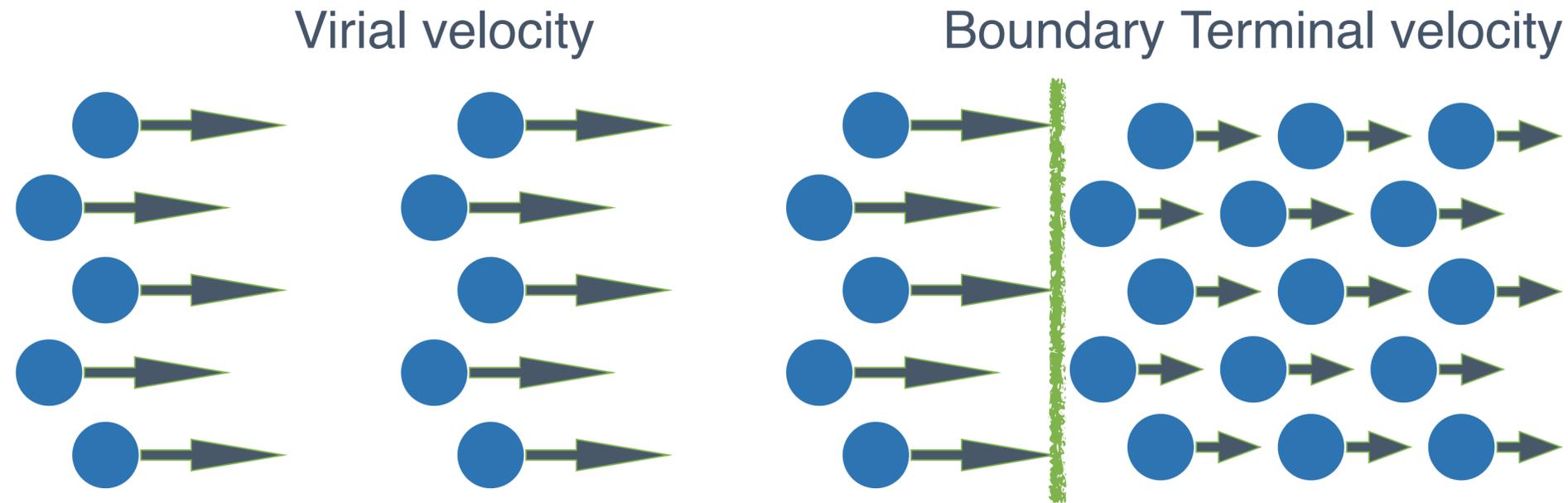
- ◆
$$\eta = \frac{\pi R_E^2 v_{\text{vir}}}{\frac{4}{3} \pi R_E^3} T_E \approx 10^{16}$$

- ◆ DM lighter than GeV evaporates $v_{\text{th}} > v_{\text{esc}}$
- ◆ Heavier than GeV sinks due to gravity



from: 2012.03957 HR M.Pospelov

TRAFFIC JAM



◆ Sinking not immediate.

◆ Downward drift

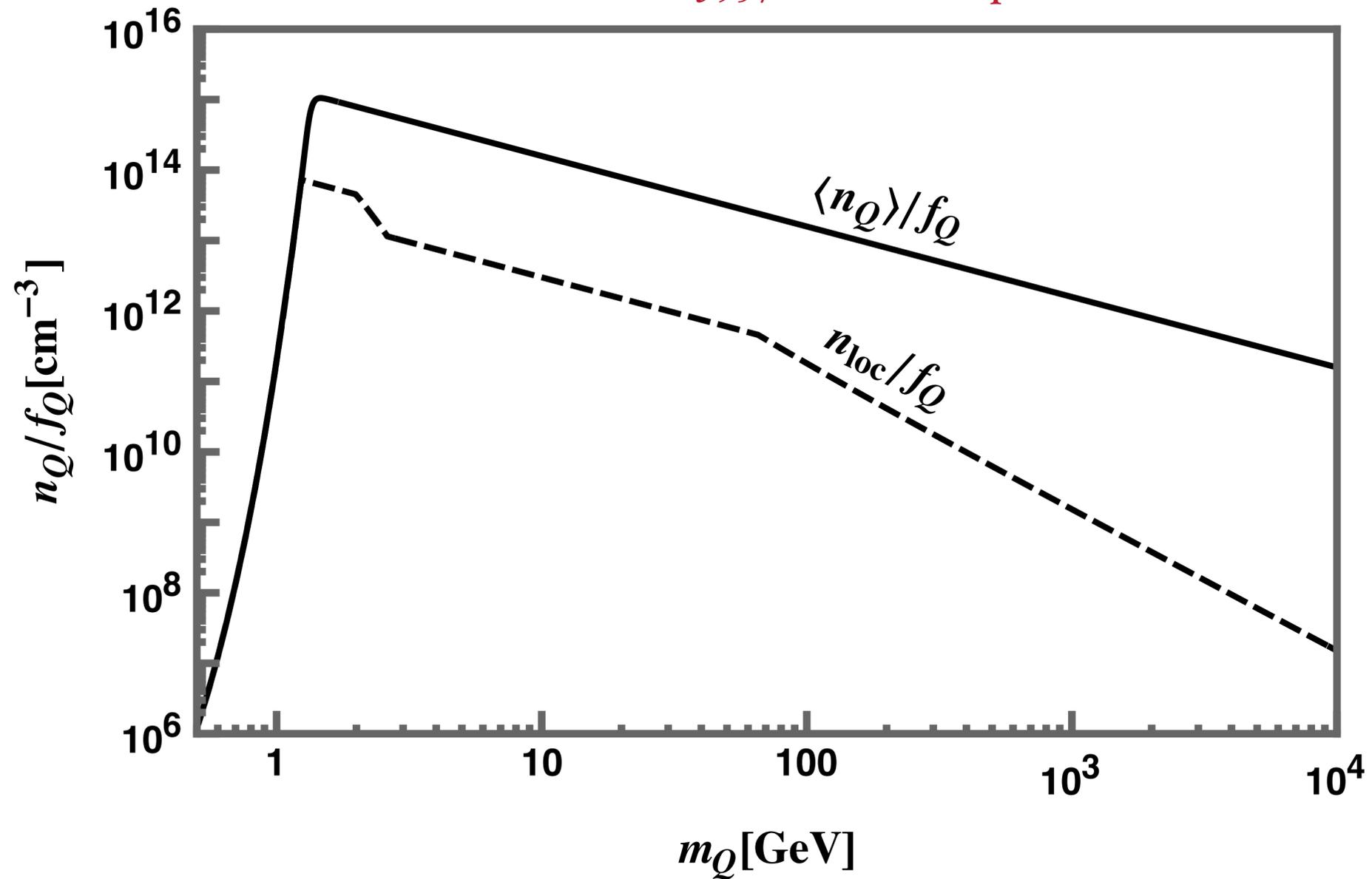
$$V_{\text{term}} \ll v_{\text{th}} \ll v_{\text{vir}}$$

◆ Traffic Jam on the way

$$\eta_{\text{term}} = \frac{n_{\text{lab}}}{n_{\text{vir}}} = \frac{V_{\text{vir}}}{V_{\text{term}}}$$

TRAFFIC JAM DENSITIES

from: 2012.03957 HR M.Pospelov

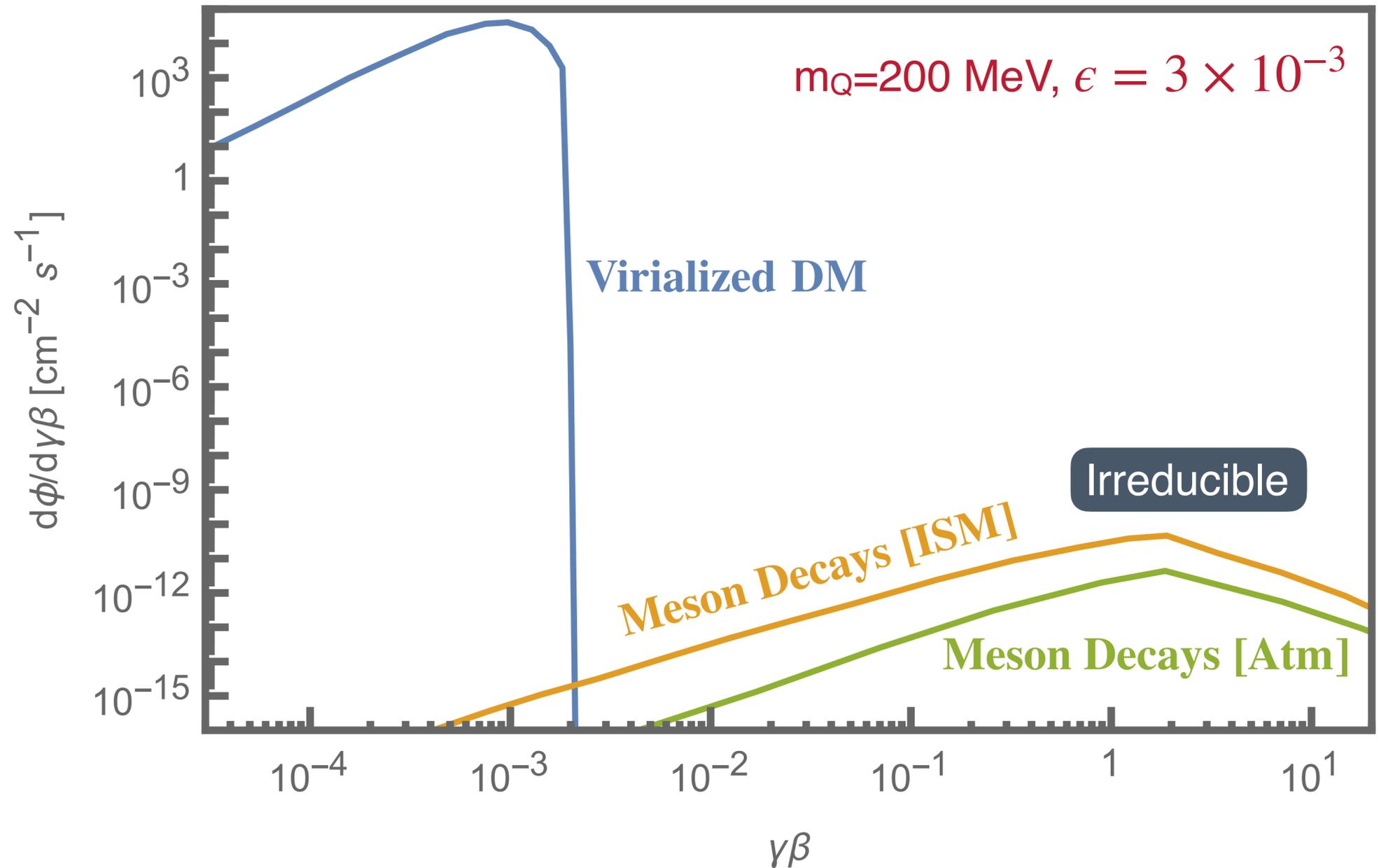
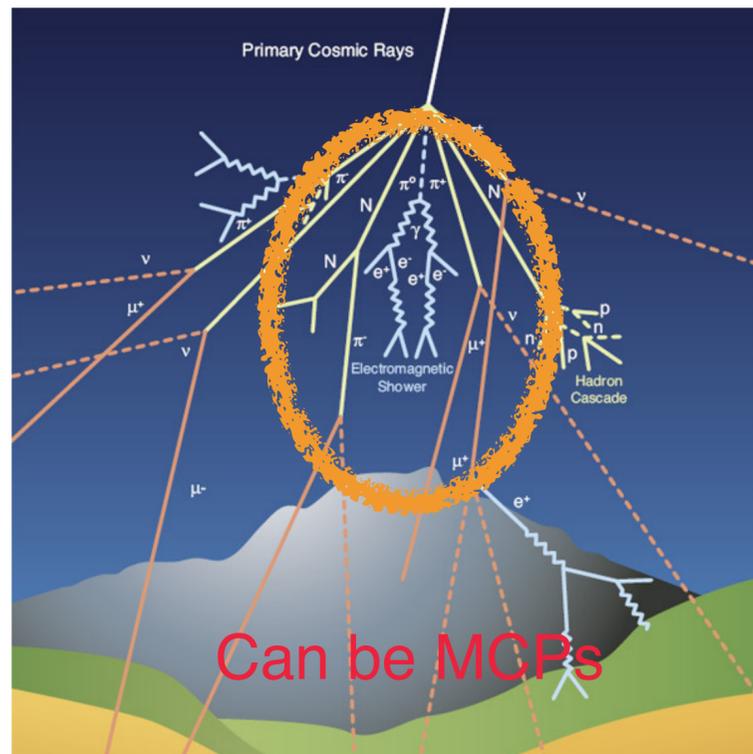


True for charges $\epsilon \gtrsim 10^{-6}$

IRREDUCIBLE MCP POPULATION

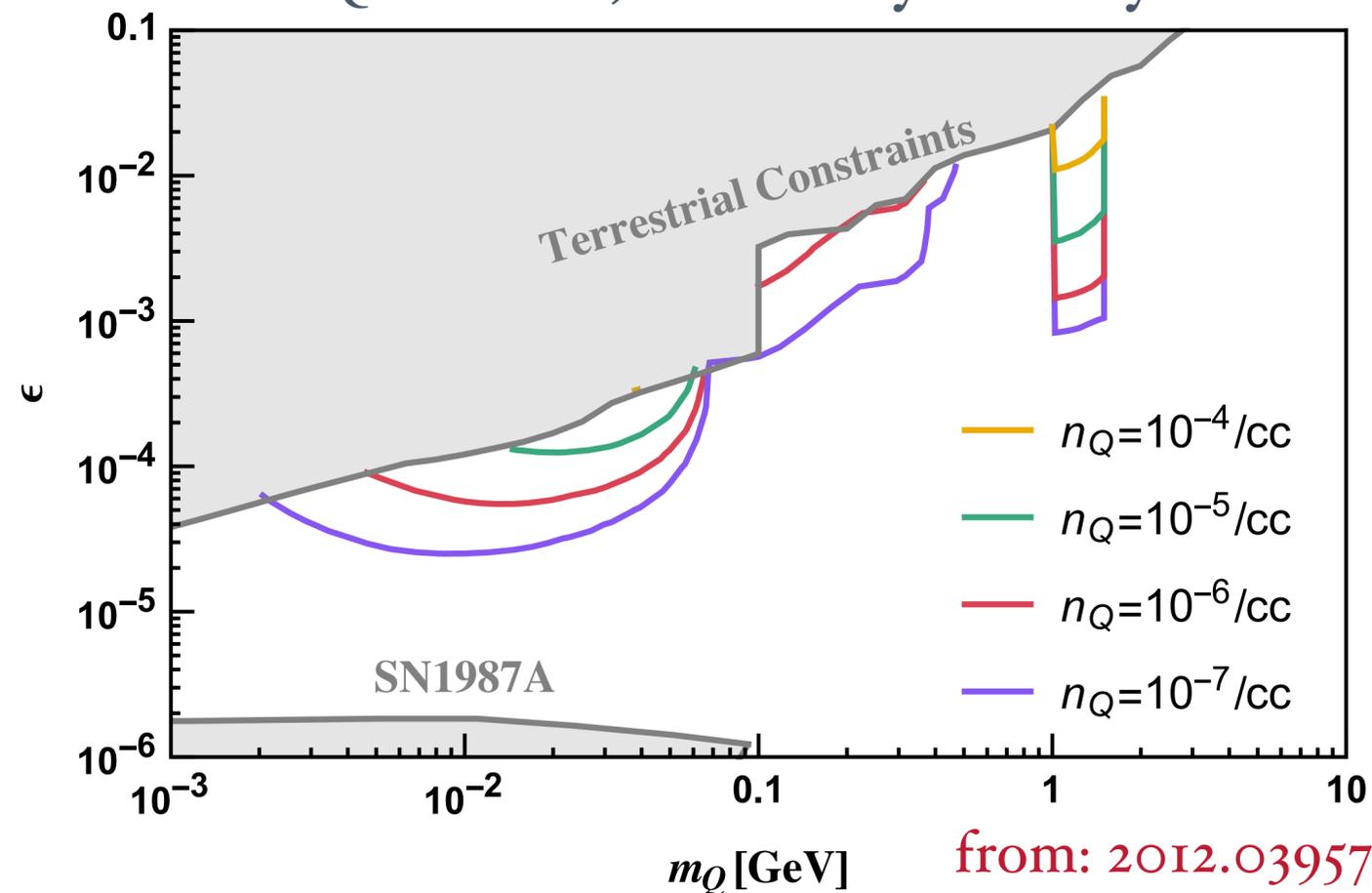
2010.II190 HR, Roni Harnik, Ryan Plestid and Maxim Pospelov

- ◆ Mesons produced in Cosmic ray collisions can decay into mCPs
- ◆ Contribution to irreducible density on Earth

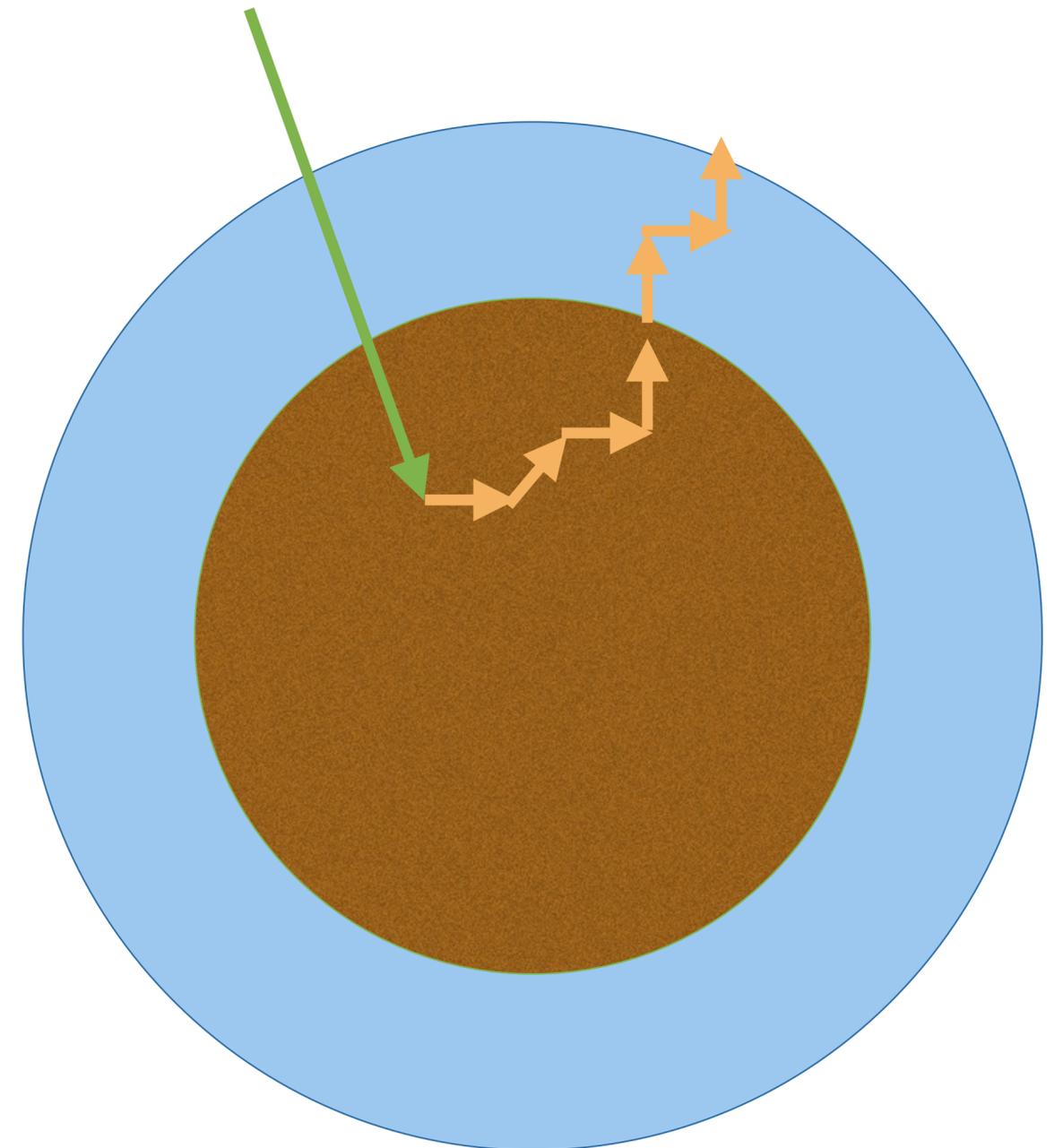


REVERSE TRAFFIC JAM

- ◆ High boost, hence penetrates deep
- ◆ Thermalized mCP, large x-section, (MFP~ micron)
- ◆ Evaporates for $m_Q < \text{GeV}$, but very slowly.



from: 2012.03957 HR M.Pospelov



EXISTING LIMITS

2012.08169 G. Afek, F. Monteiro, J. Wang, B. Siegel, S. Ghosh, D.C. Moore

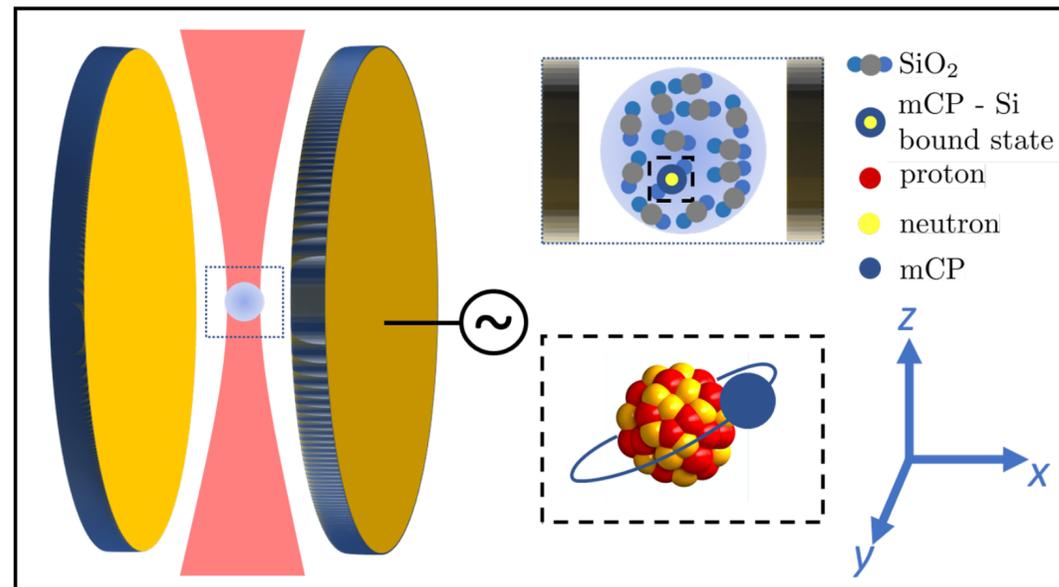
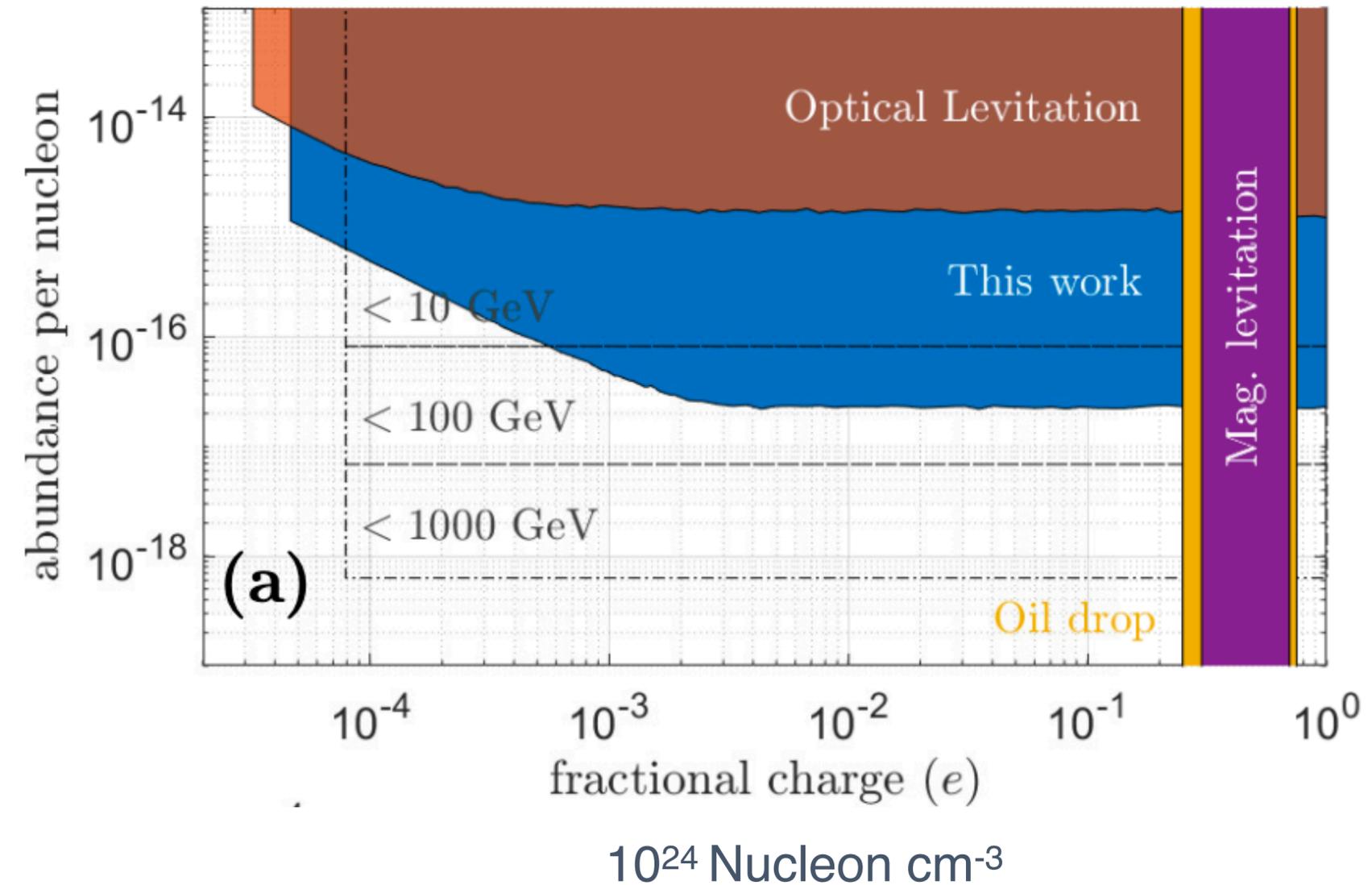
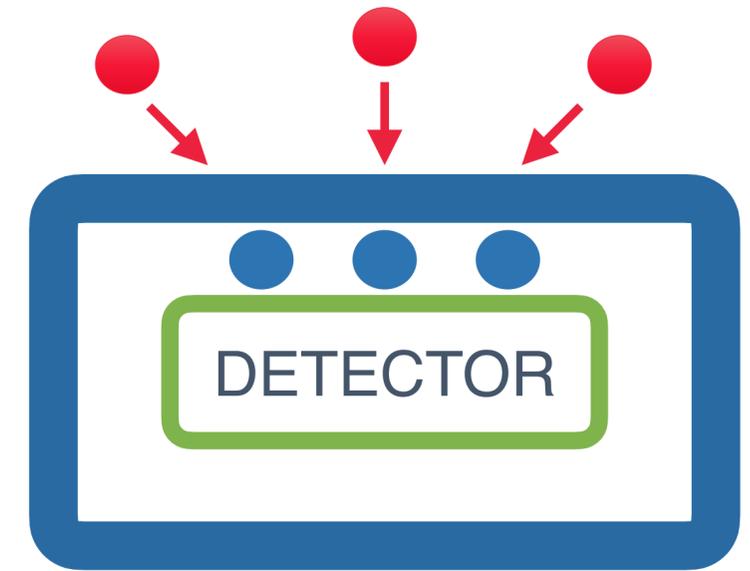


FIG. 1. SiO_2 spheres are levitated in high vacuum between a pair of parallel electrodes to search for a violation of charge neutrality by, *e.g.*, a **mCP electrostatically bound** to a Si or O nucleus in the sphere.



DETECTION NIGHTMARE

- ◆ Despite large number density & cross-section
- ◆ Small energy deposit: $300 \text{ Kelvin} \approx 26 \text{ meV}$
- ◆ Small momentum transfers: See neutral atom
- ◆ Low threshold detectors have low temperature walls to reduce background
- ◆ Small MFP \sim micron, rapidly thermalize with walls



WISHLIST

- ◆ Detect Small energy deposit: 300 Kelvin \approx 26 meV
- ◆ If target charged, then huge Rutherford x-sections

$$\frac{d\sigma}{dq^2} \propto \frac{1}{q^4} \text{ at small momentum transfer}$$

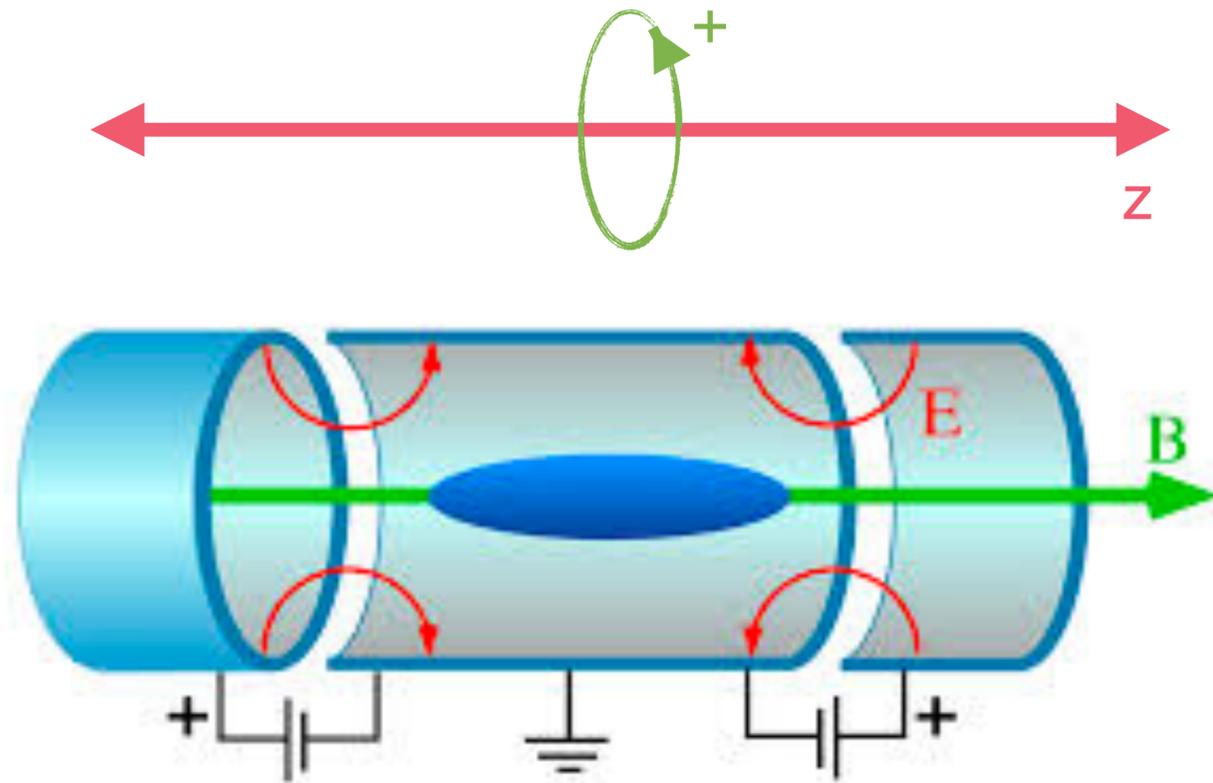
- ◆ $T_{\text{wall}} \gg E_{\text{thr}}$
- ◆ Large number of targets ... not required



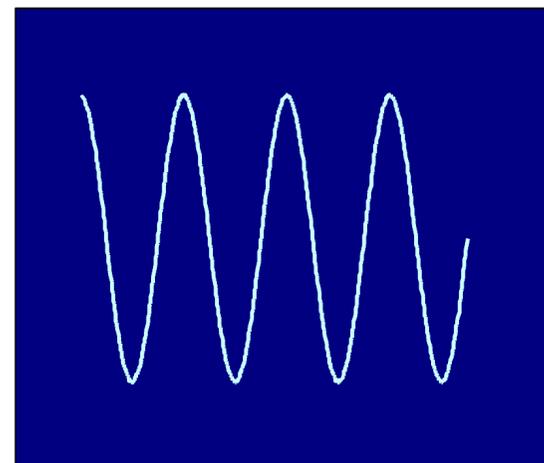
OUTLINE

- ◆ Strongly Interacting Dark Relics & Traffic Jams
- ◆ **Milli-charge Particle Detection with Ion Traps**
- ◆ Dark-Standard Model bound states for a heavy Dark Photon

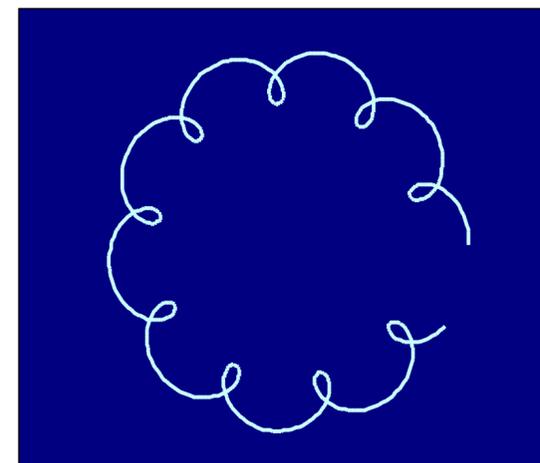
IONS IN COLD TRAPS



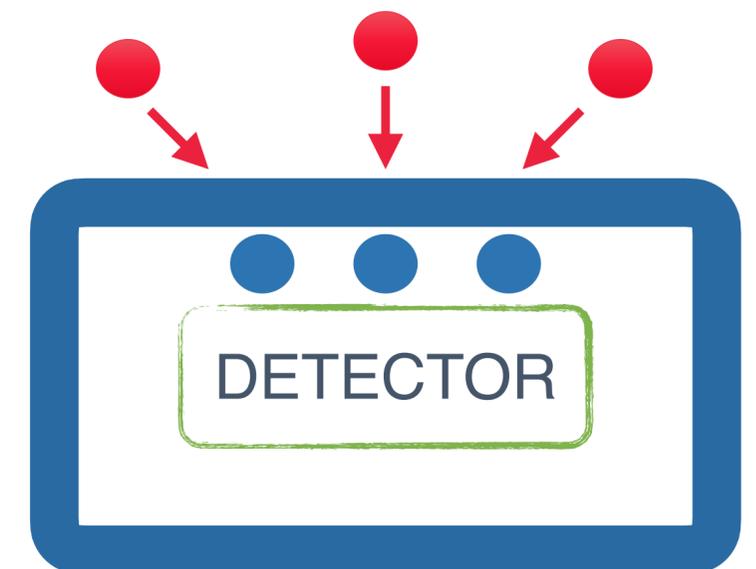
- ◆ Single Ion trapped in Electromagnetic fields
- ◆ Used in Quantum Metrology/ Quantum Computing
- ◆ Stable in trap for O(year)
- ◆ $T_{\text{wall}} \gg T_{\text{trap}}$



Axial motion



Radial motion



DATA SUMMARY

Experiment	Type	Ion	V_z	T_{wall}	ω_p [neV]	T_{ion} [neV]	Heating Rate (neV/s)
Hite et al, 2012 [40]	Paul	${}^9\text{Be}^+$	0.1 V	300 K	$\omega_z = 14.8$	14.8	640
Goodwin et al, 2016 [43]	Penning	${}^{40}\text{Ca}^+$	175 V	300 K	$\omega_z = 1.24$	1.24	0.37
Borchert et al, 2019 [44]	Penning	\bar{p}	0.633 V	5.6 K	$\omega_+ = 77.4$ $\omega_- = 0.050$	7240	0.13

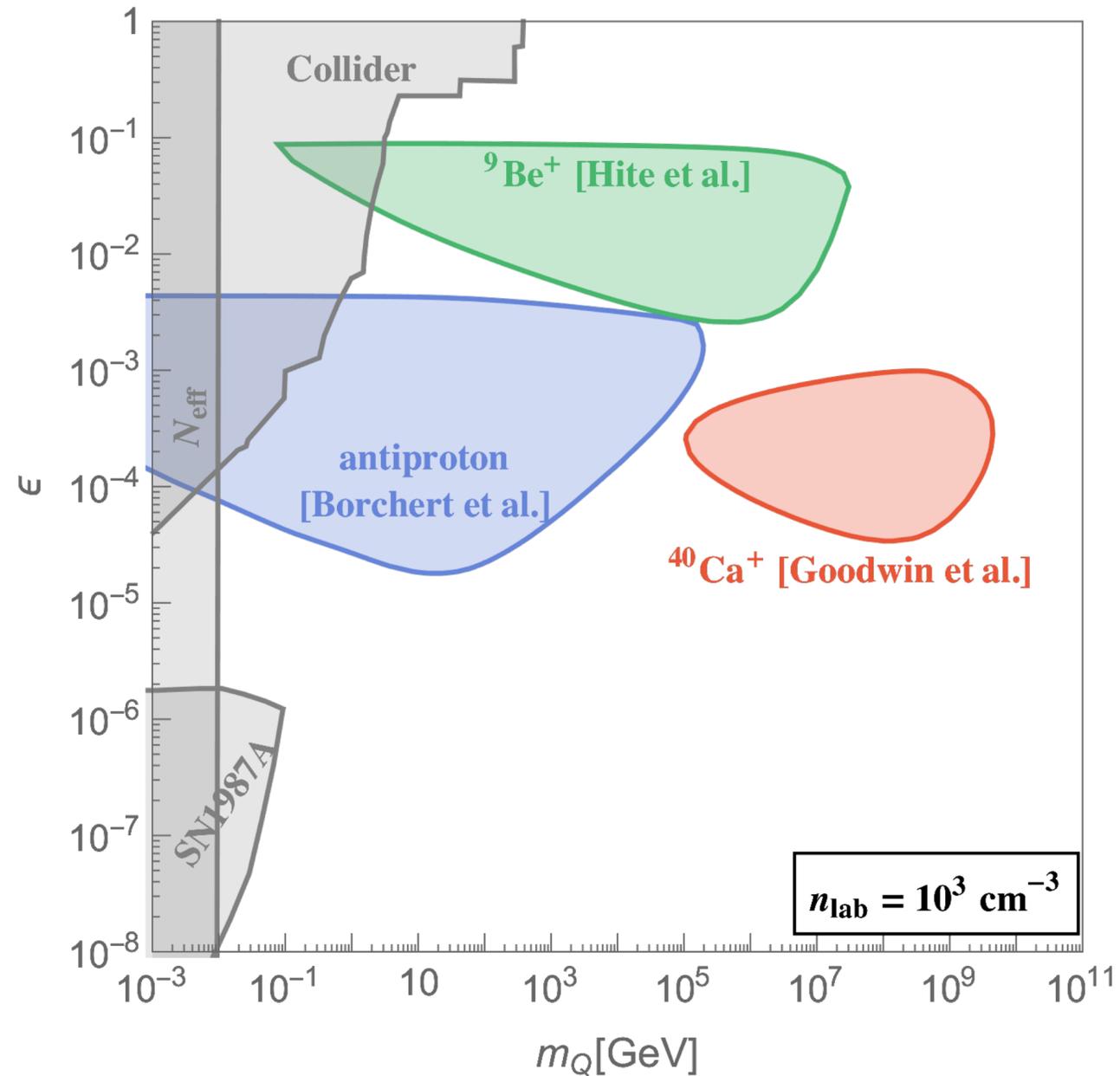
No reach for $\epsilon \gtrsim \frac{T_{\text{wall}}}{V_z}$

HEATING RATE

$$\frac{dE_{\text{dep}}}{dt} = \int E_{\text{dep}}(q^2) \frac{4\pi\alpha^2\epsilon^2}{v^2q^4} dq^2 \approx 10^{-6} \frac{\text{eV}}{\text{sec}} \epsilon^2 \frac{n_{\text{lab}}}{1/\text{cm}^3} \frac{\text{GeV}}{m_{\text{ion}}} \dots \gtrsim 10^{-10} \frac{\text{eV}}{\text{sec}}$$

TERRESTRIAL POPULATION CONSTRAINTS

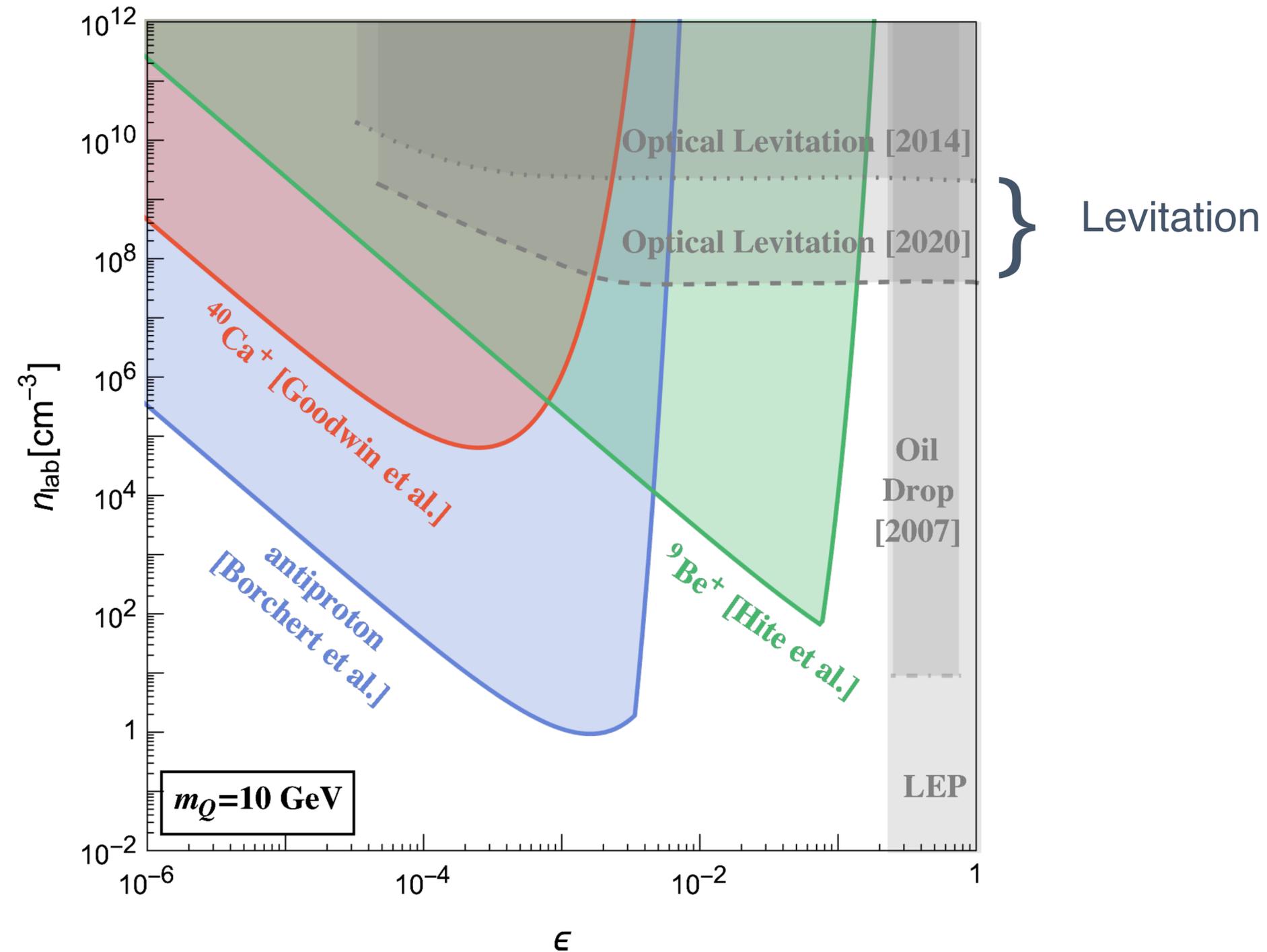
$$m_Q^{\min} = \frac{E_{\min}^2 m_T}{16 T_{\text{trap}} T_{\text{wall}}}$$



$$m_Q^{\max} = \frac{16 m_T T_{\text{trap}} T_{\text{wall}}}{E_{\min}^2}$$

arXiv: 2108.05283
 HR with
 D. Budker,
 C. Smorra,
 P. Graham,
 F. Schmidt-Kaler,
 S. Ulmer

TERRESTRIAL POPULATION CONSTRAINTS

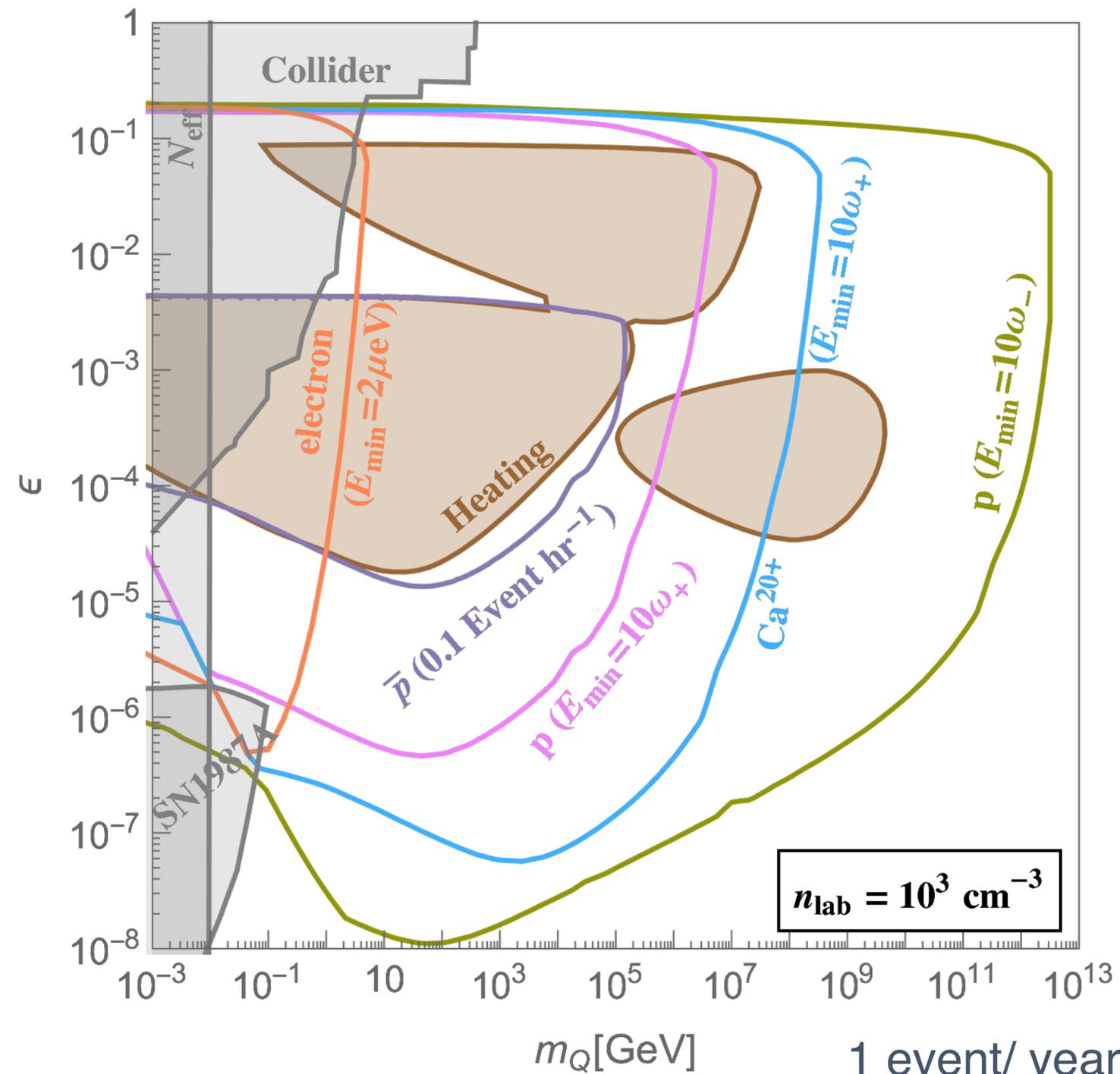


arXiv: 2108.05283
 HR with
 D. Budker,
 C. Smorra,
 P. Graham,
 F. Schmidt-Kaler,
 S. Ulmer

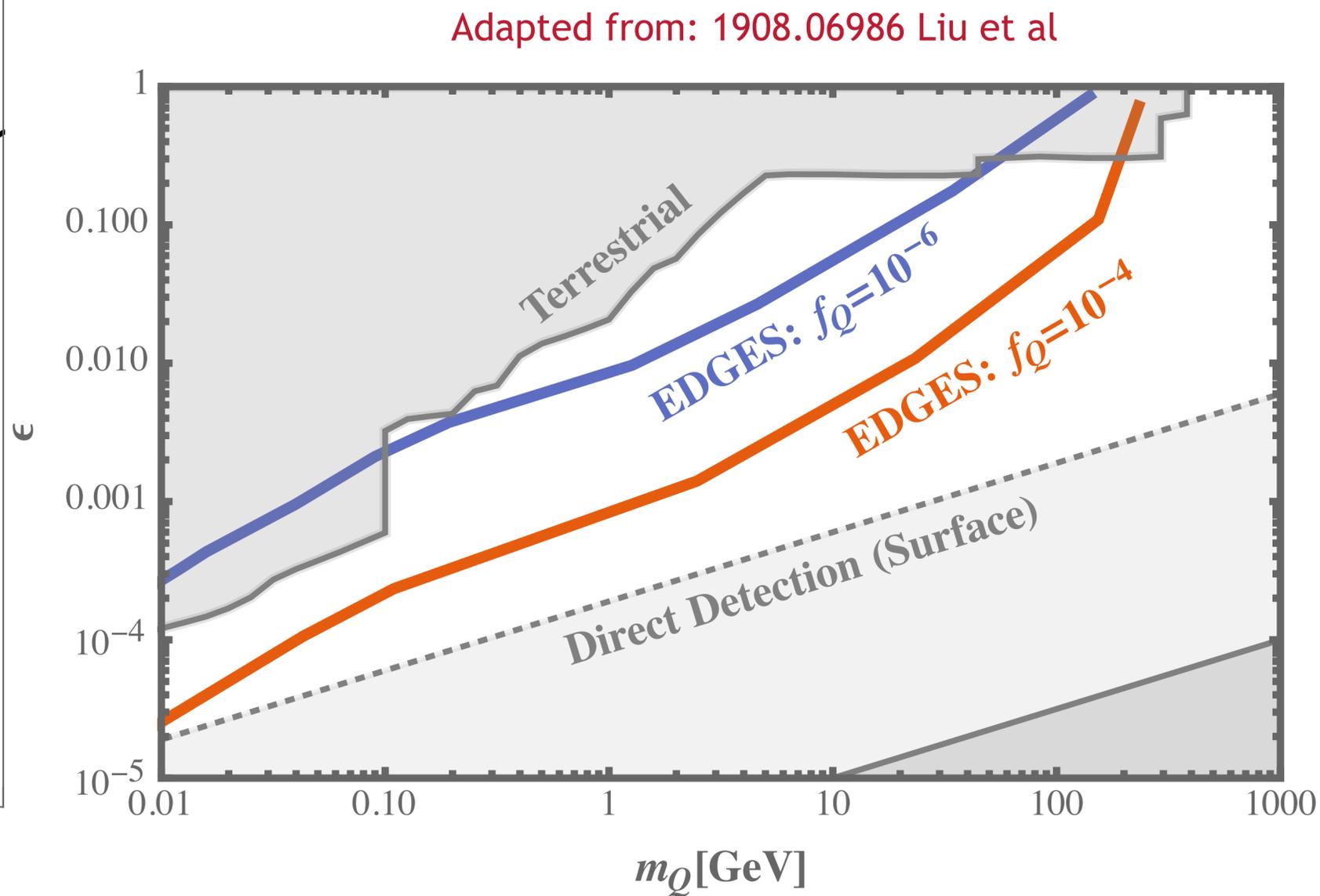
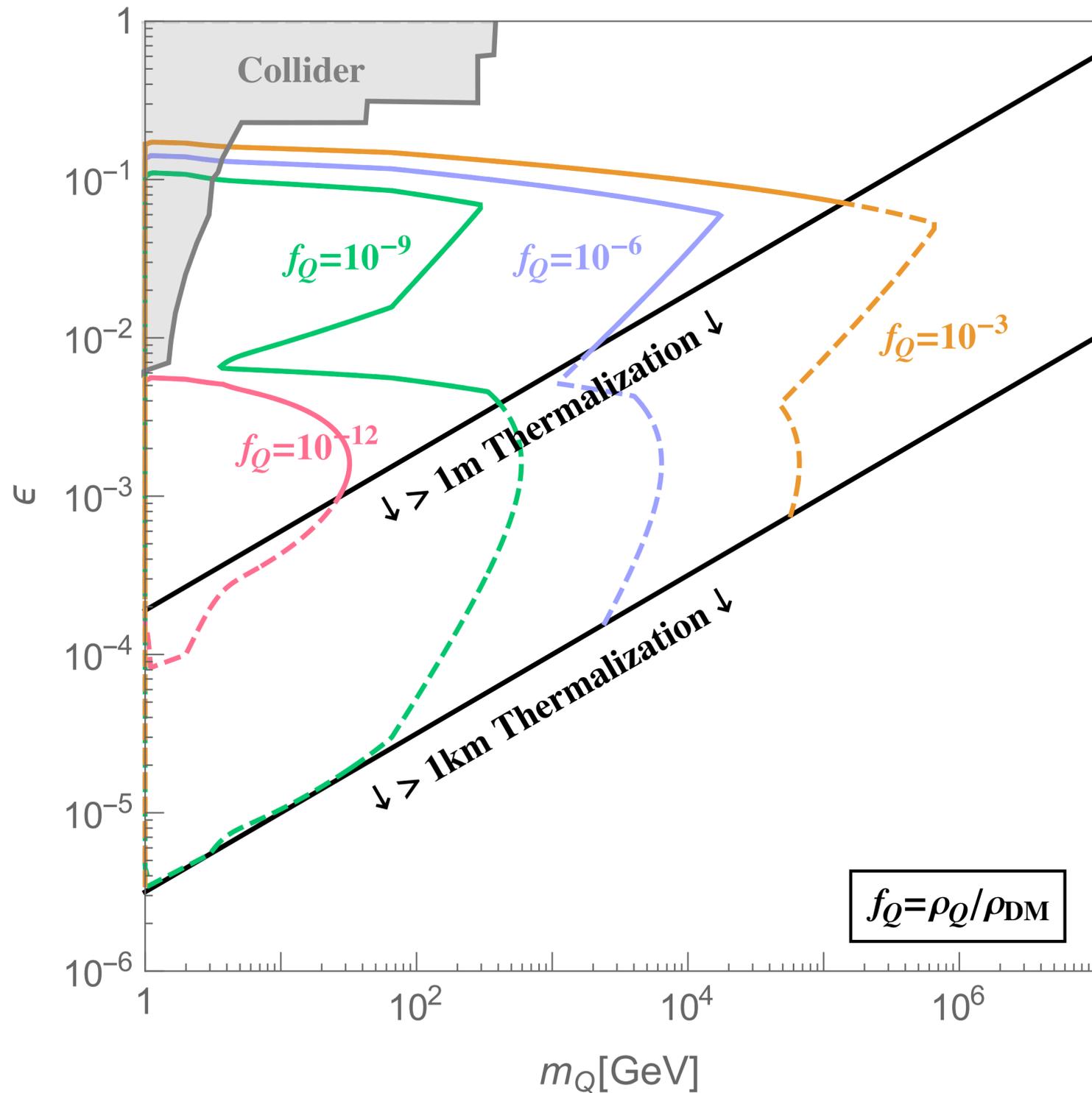
POSSIBLE IMPROVEMENTS

- ◆ Single Event Measurement instead of cumulative heating rate
- ◆ Requires high cadence data and energy resolution
- ◆ Highly Charged Ions
- ◆ Using electron targets - less momentum transfer for same energy transfer

PROJECTIONS



LIMITS ON DARK MATTER



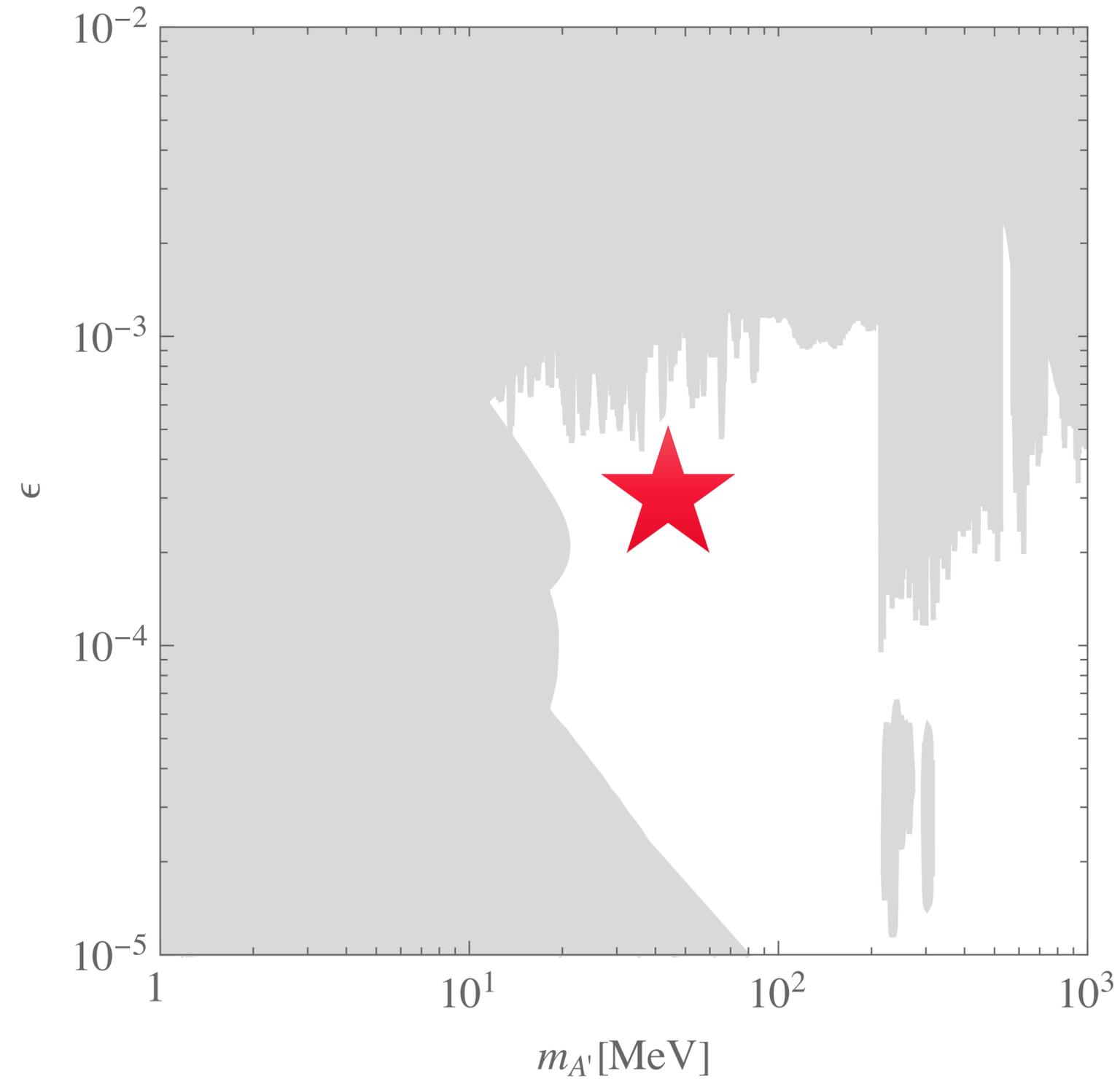
OUTLOOK

- ◆ Repeating experiment in deep mine
- ◆ Collective excitations in Ion lattices
- ◆ Accumulating mCPs in an electric field bottle

OUTLINE

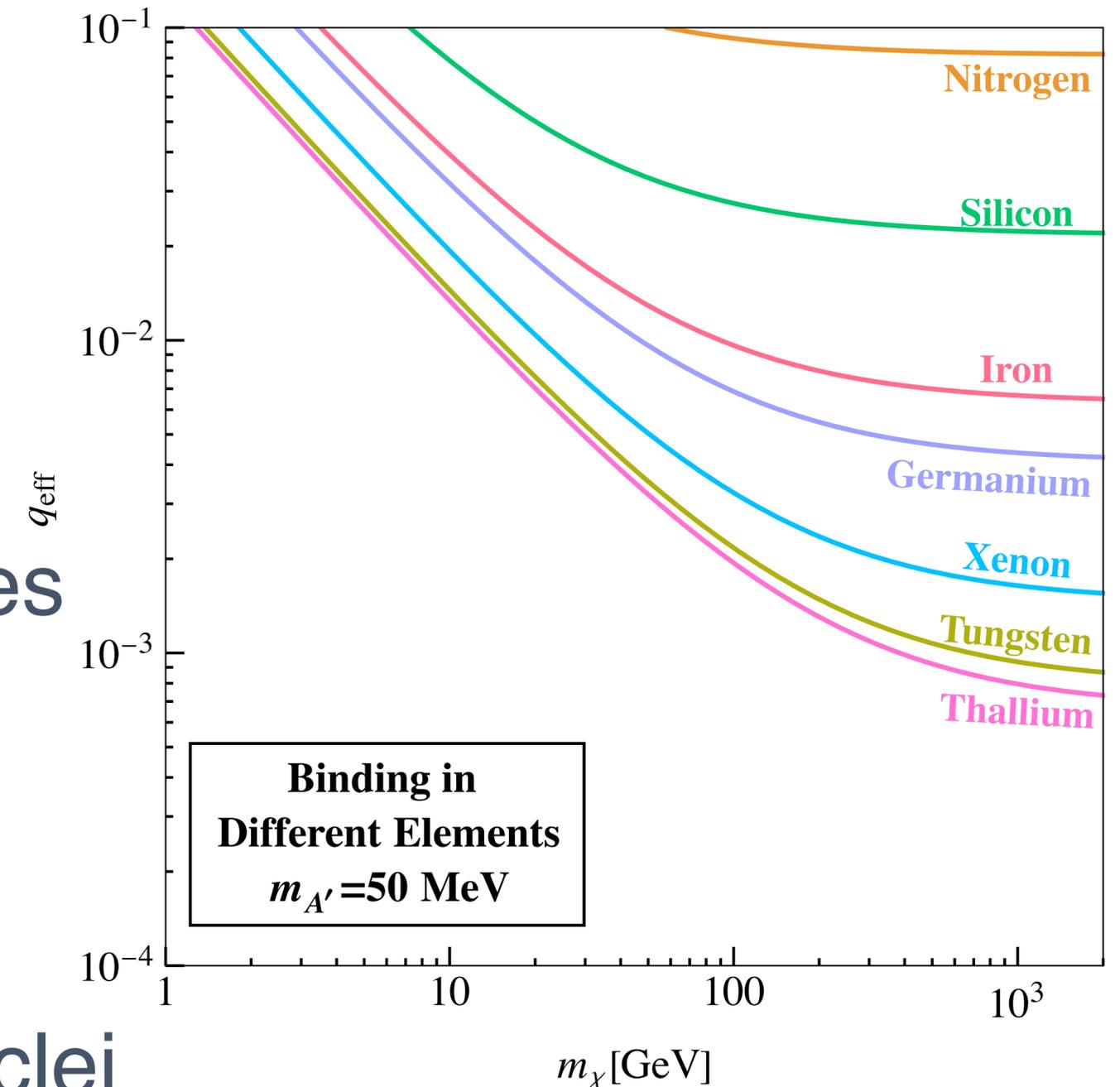
- ◆ Strongly Interacting Dark Relics & Traffic Jams
- ◆ Milli-charge Particle Detection with Ion Traps
- ◆ Dark-Standard Model bound states for a heavy dark photon

HEAVY DARK PHOTON MEDIATOR



MEDIATOR FOR BOUND STATES?

- ◆ $q_{\text{eff}} = \epsilon \sqrt{\frac{\alpha_D}{\alpha}}$
- ◆ Bohr Radius $R_B^{-1} \approx Z_N \alpha q_{\text{eff}} \mu$
- ◆ If $R_B \gtrsim m_{A'}^{-1}$ then no bound states
- ◆ For some q_{eff} binding only with heavy SM nuclei
- ◆ WIMP DM Detectors = heavy nuclei



TRACKING THRU OVERBURDEN

- ◆ Large $q_{\text{eff}} \implies$ Large local overabundances
- ◆ No binding in atmosphere with Nitrogen/Oxygen
- ◆ No binding in crust upto Iron
- ◆ Makes its way into Xenon tank
- ◆ Binds with Xenon - releasing E_B

BINDING IN XENON

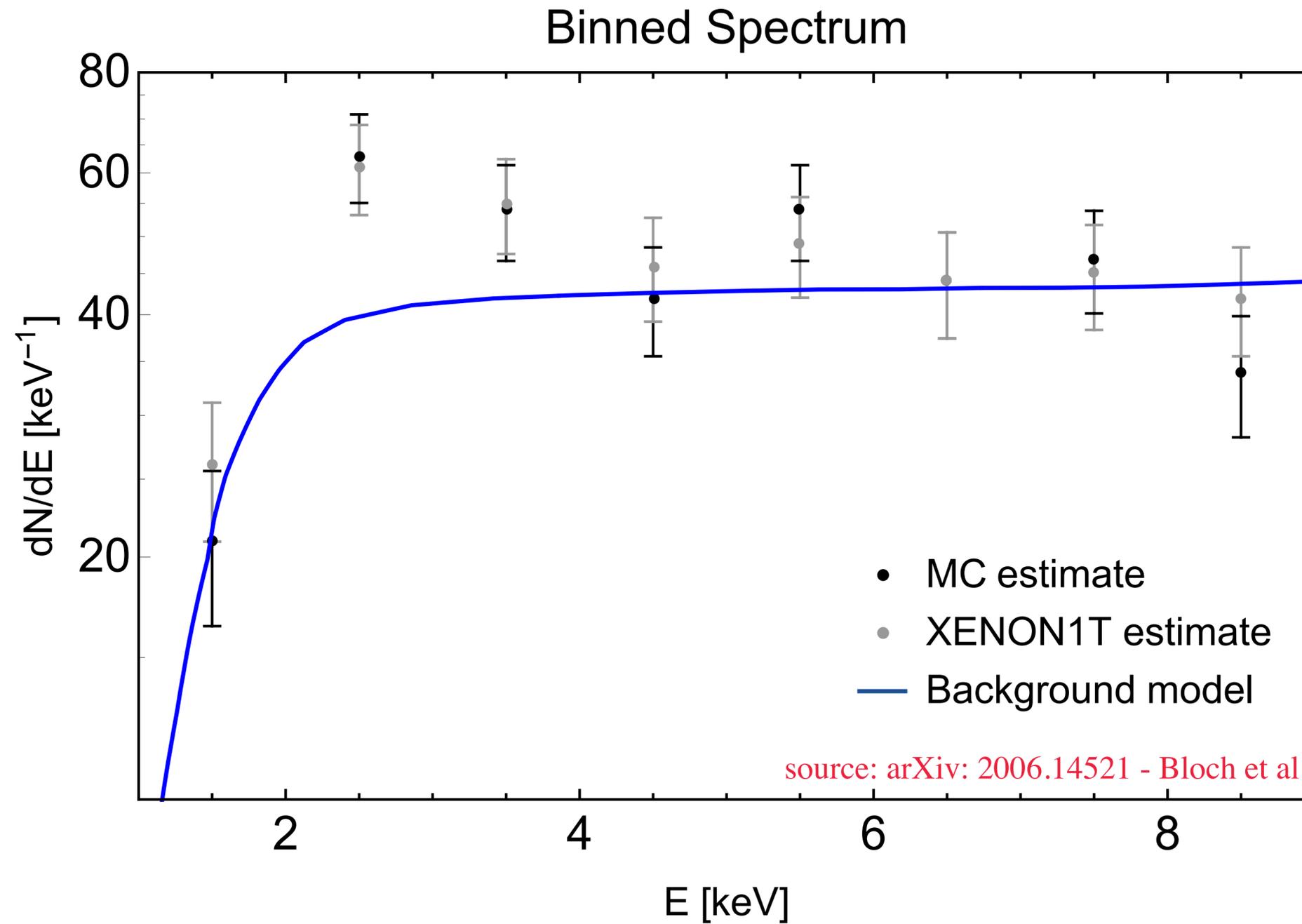
- ◆ Binds with Xenon - releasing E_B

$$E_B \sim \boxed{10 \text{ keV}} \times \left(\frac{q_{\text{eff}}}{10^{-3}} \right)^2 \left(\frac{Z}{54} \right)^2 \left(\frac{\mu}{122 \text{ GeV}} \right)$$

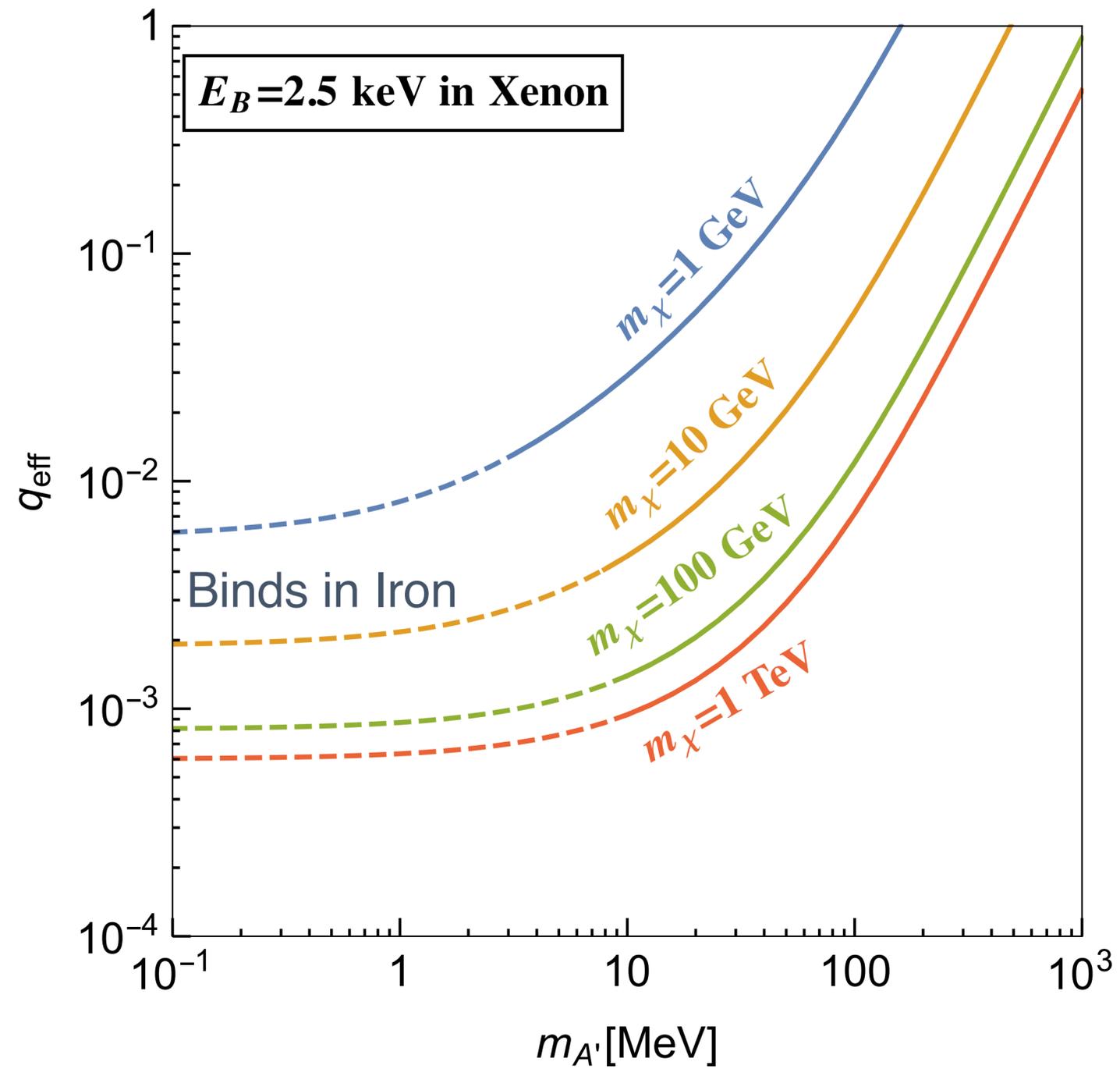
- ◆ Usually KE (scattering) / Mass (absorption)
- ◆ New source of energy
- ◆ Monochromatic electron ejection

XENON S₁-S₂ SIGNAL

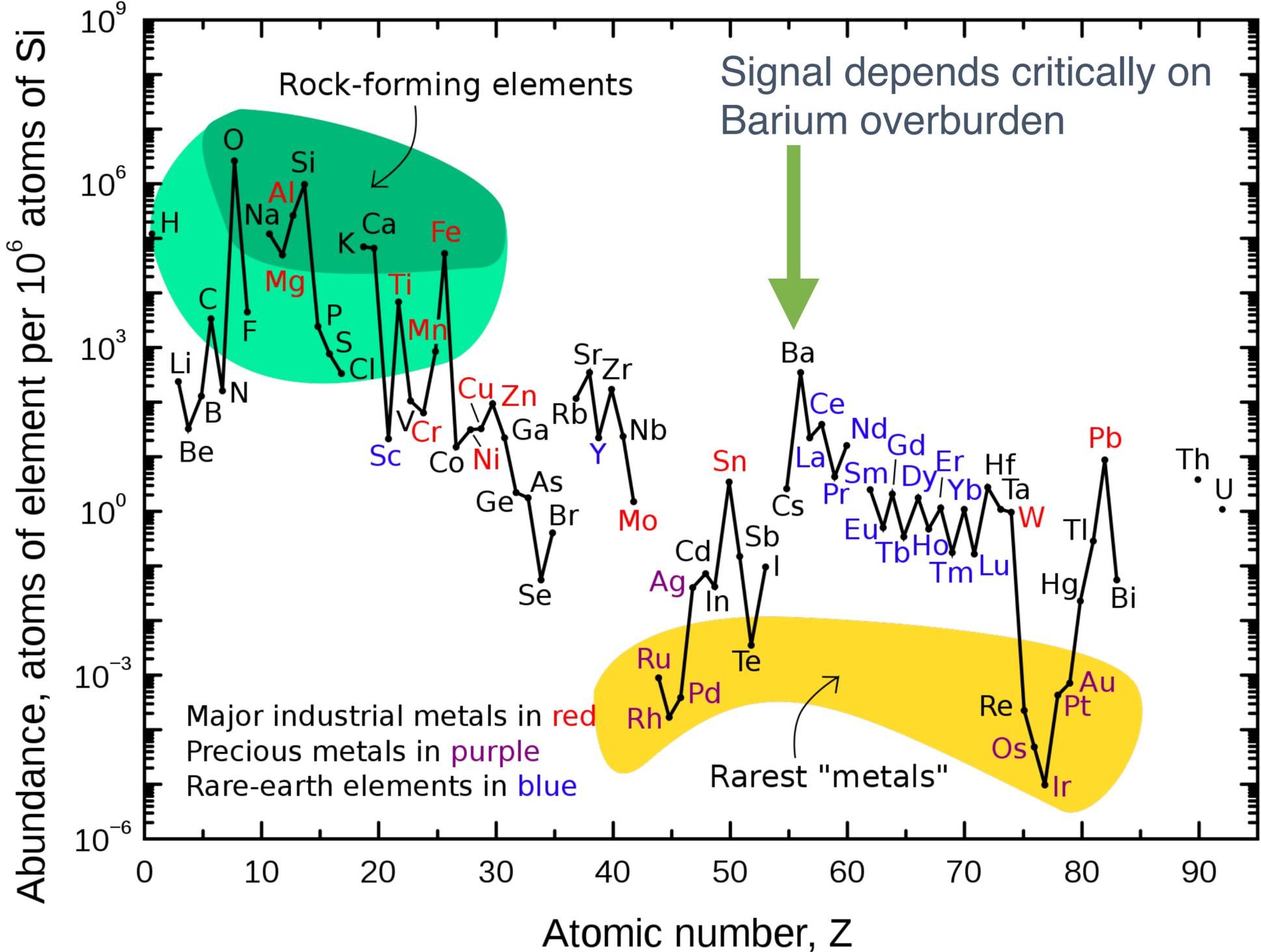
arXiv: 2006.09721v3 - Xenon1T collaboration



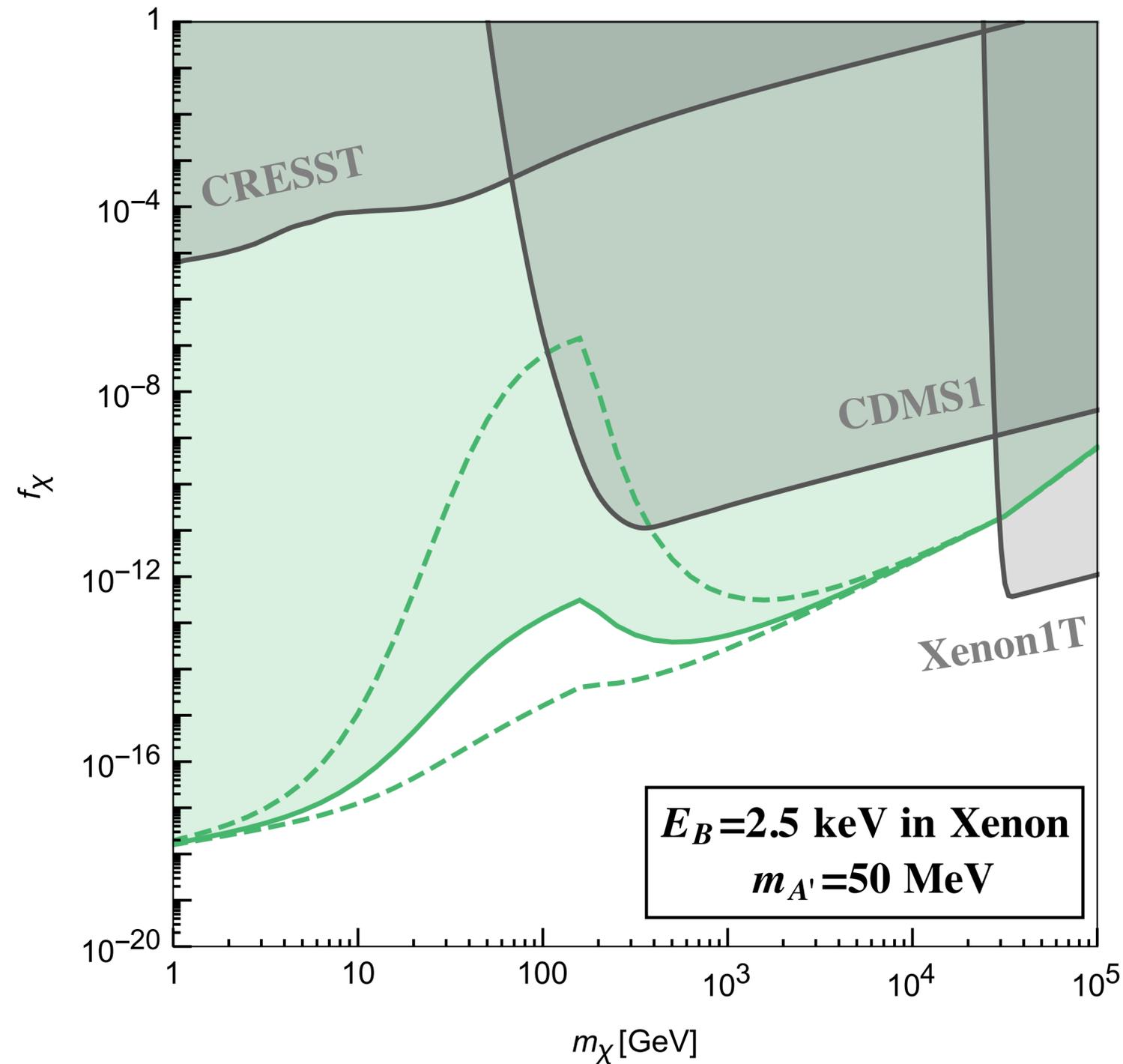
$E_B = 2.5 \text{ keV}$ in Xenon



HEAVY ELEMENTS IN OVERBURDEN

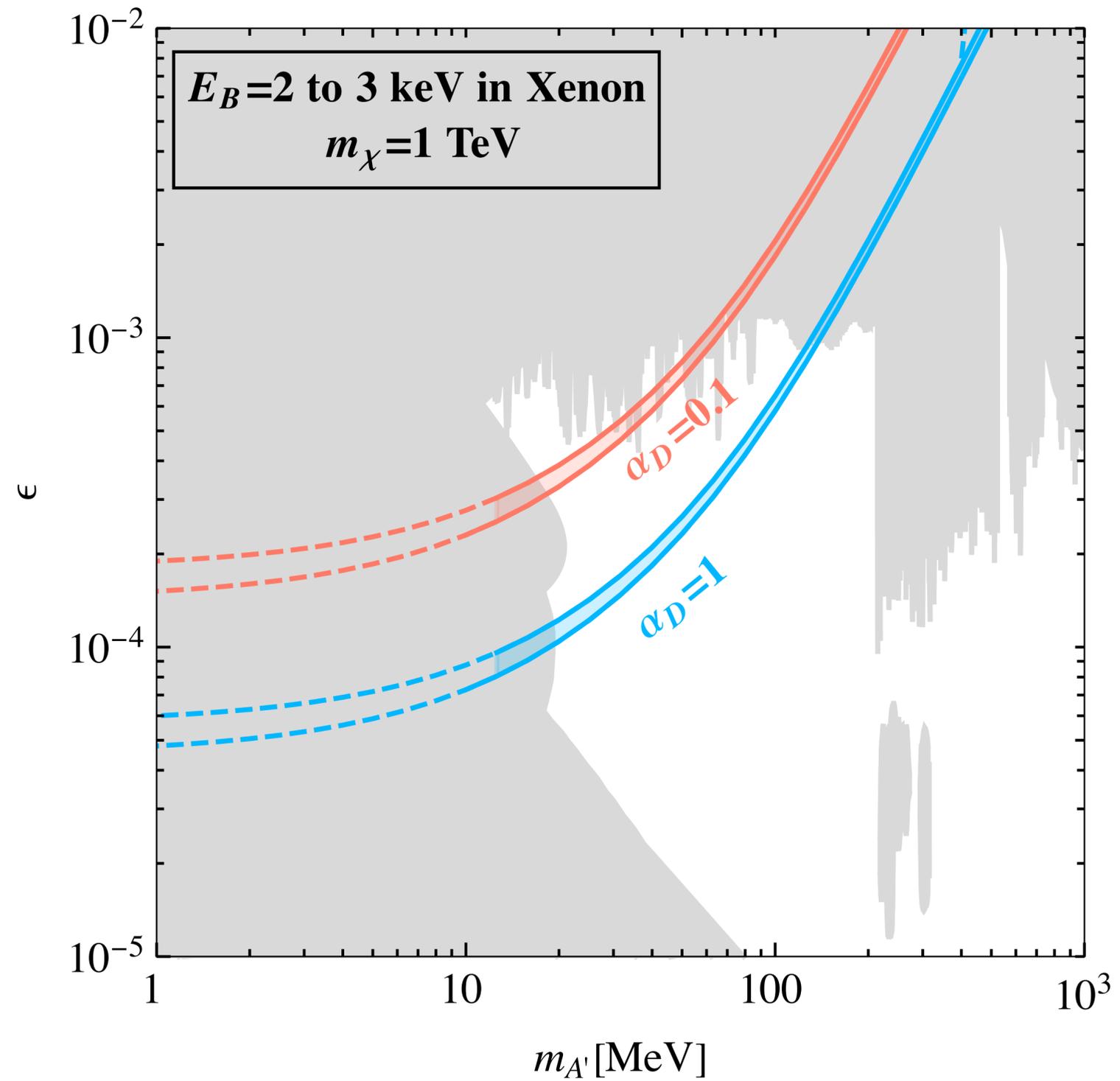


PARAMETER SPACE



- ◆ $n_{\text{Ba}} \in (0.7 - 6) \times 10^{18} \text{cm}^{-3}$
- ◆ Extremely small DM fractions explain anomaly
- ◆ Limits set everywhere above

DARK PHOTON PARAMETER SPACE



TRAFFIC JAM DETECTION

MODEL	DETECTION STRATEGY	arXiv
Contact Interactions/ QCD	1) Isomeric Tantalum 2) Uncollider 3) Bolometers	1911.07865 Forthcoming Forthcoming
Dark Photon Mediated	Bound states in Xenon	2110.06217
Millicharge Particles	1) Electrostatic Accelerator 2) Ion Traps	2012.03957 2108.05283
B-L mediated blobs	1) Neutron Bottle	2008.06061

SUMMARY

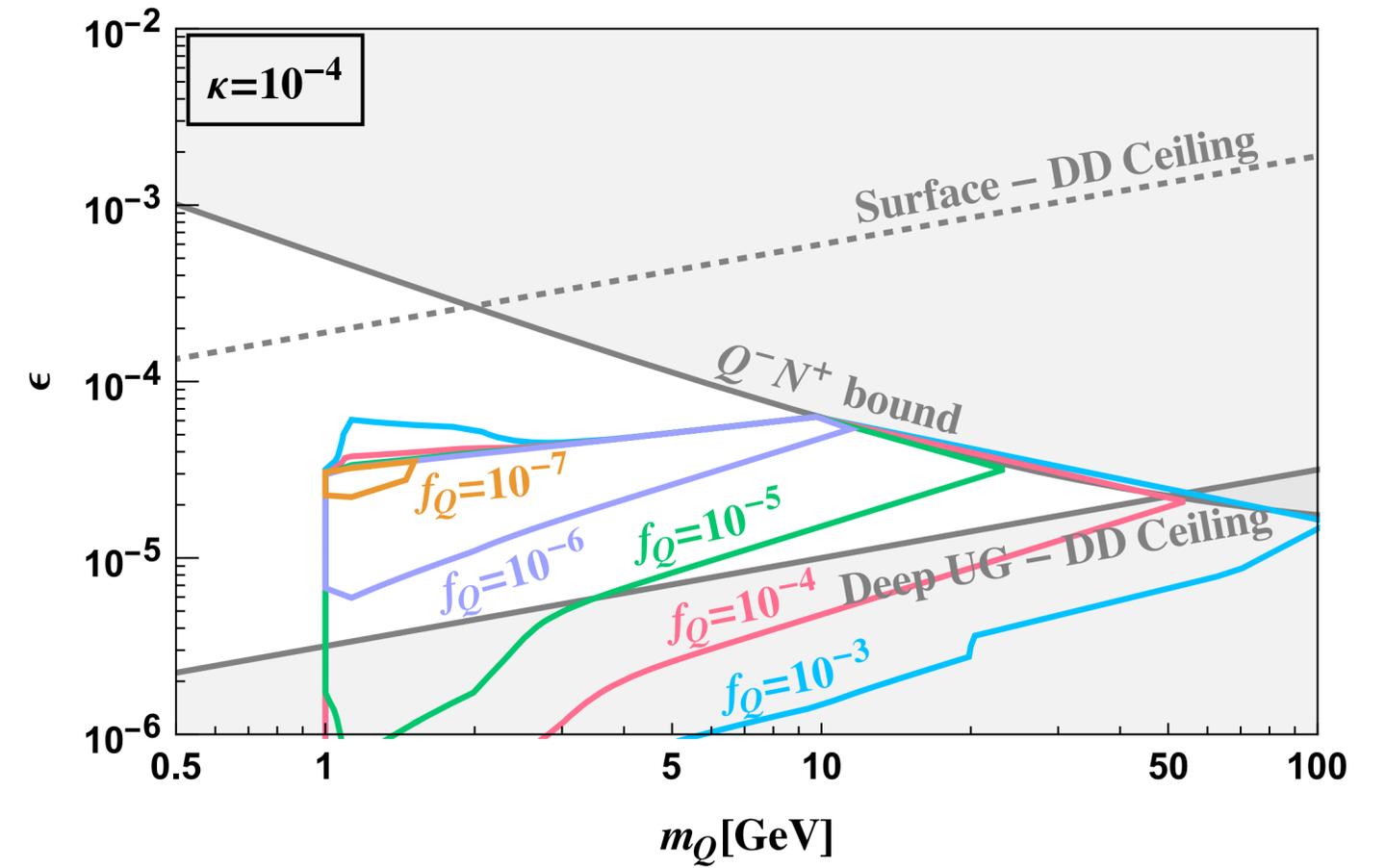
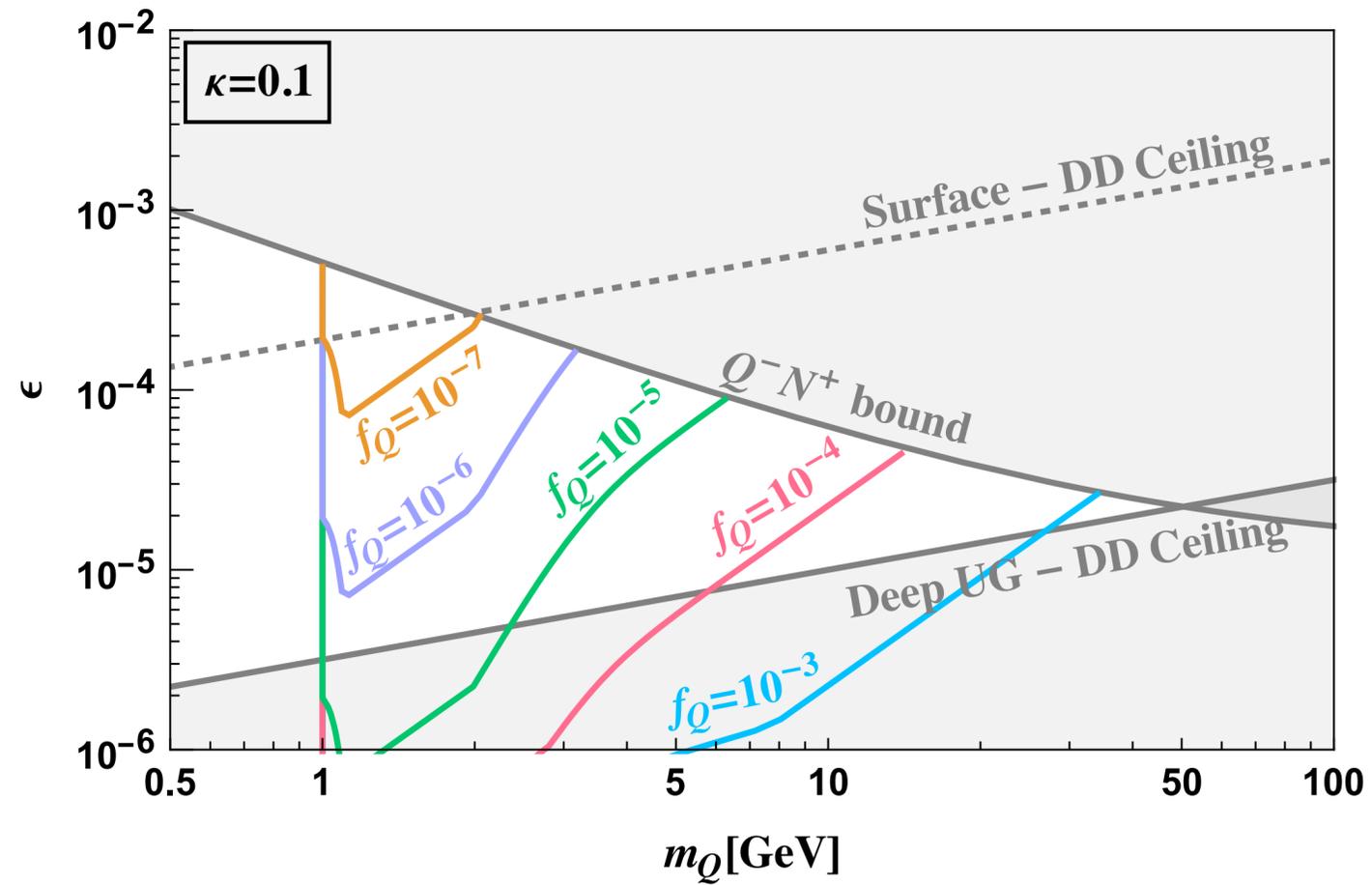
- ◆ Small abundances of stable dark relics are natural
- ◆ Large SM couplings \implies captured on earth
- ◆ Huge terrestrial densities due to traffic jam
- ◆ Model 1: mCPs can be detected with ion traps
- ◆ Model 2: heavy dark photon mediated, form bound states in heavy element detectors.

BACKUP

SIGNALS

SYMMETRIC POPULATION

ANNIHILATIONS IN SUPER-K



CAPTURE CROSS-SECTION

$$\sigma_B v_{\text{rel}} \approx \frac{4\pi}{9} \frac{(Z\alpha m_e)^5}{(2\mu E_B)^{7/2}} \left(\frac{\mu}{m_N}\right)^4 F_\chi^2 F_e^2.$$

$$\sigma_B v_{\text{rel}} \simeq 5 \times 10^{-35} \text{cm}^2 \times \left(\frac{\mu}{\text{GeV}}\right)^{0.55}$$

MEASUREMENT

◆ $\nu_+, \nu_-, \nu_z \approx \text{MHz} \approx 4\text{neV} \approx 50\mu\text{K}$

◆ Strong inhomogeneous magnetic field B_2

$$\text{◆ } \Delta\nu_z(n_+, n_-, m_s) = \frac{h\nu_+}{4\pi^2 m_p \nu_z} \frac{B_2}{B_0} \left[\left(n_+ + \frac{1}{2} \right) + \frac{\nu_-}{\nu_+} \left(n_- + \frac{1}{2} \right) + \frac{g_p m_s}{2} \right]$$

◆ $\Delta\nu_z$ measured with image current detection to detect Δn_+