



The ASACUSA Experiment

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On behalf on the ASACUSA collaboration

SSP 2022

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HIROSHIMA UNIVERSITY



東京大学
THE UNIVERSITY OF TOKYO

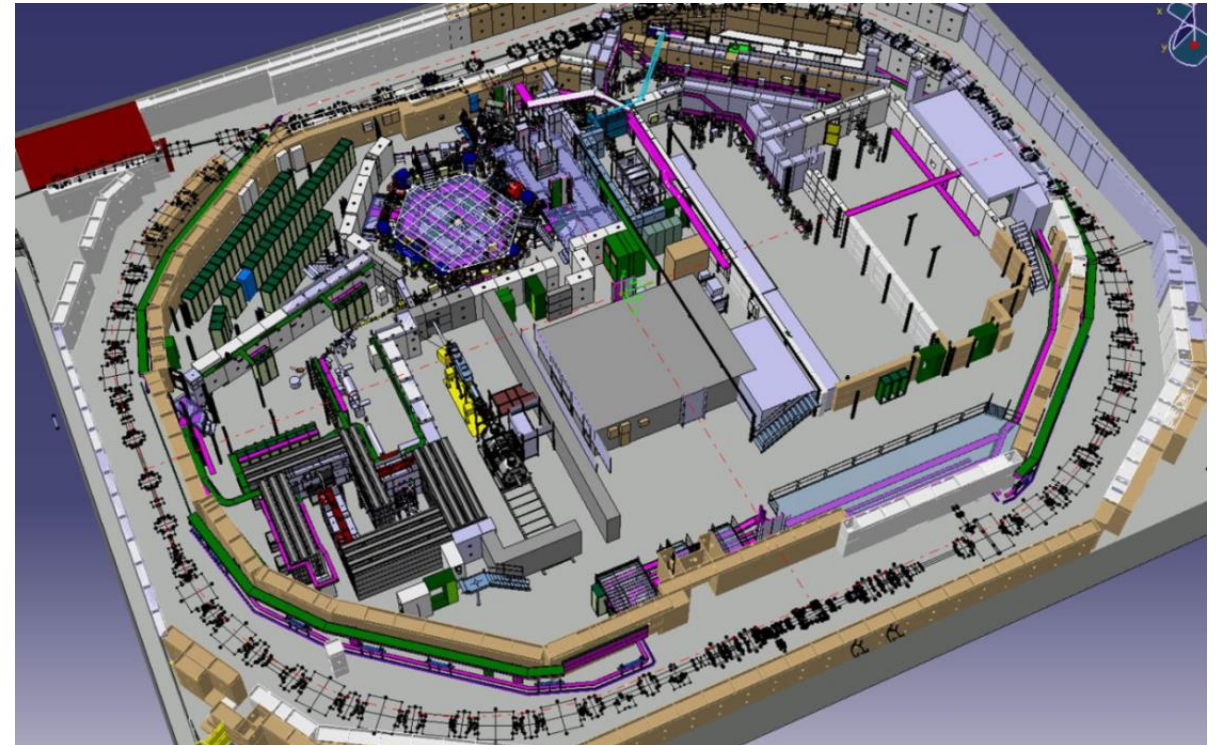
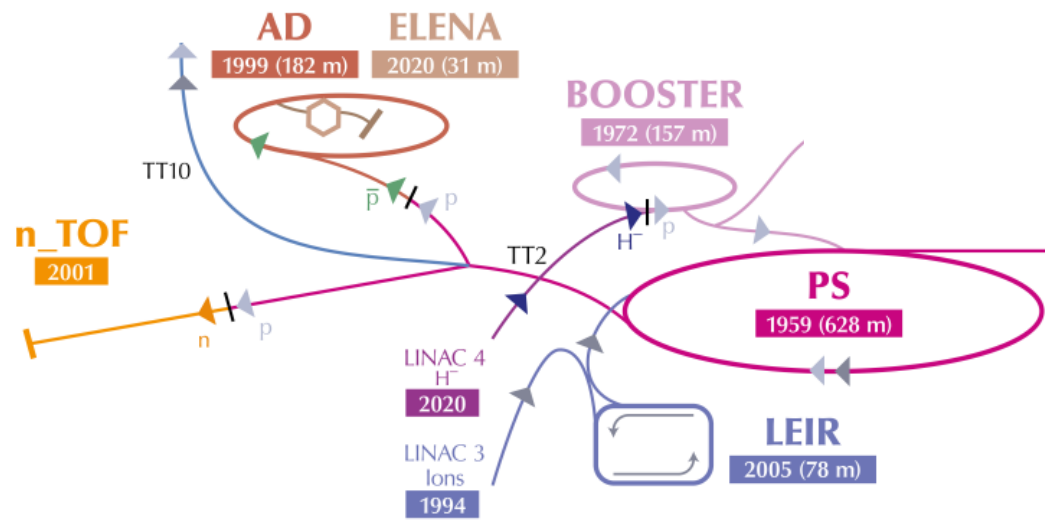
ASACUSA

Atomic
Spectroscopy
And
Collisions
Using
Slow
Antiprotons



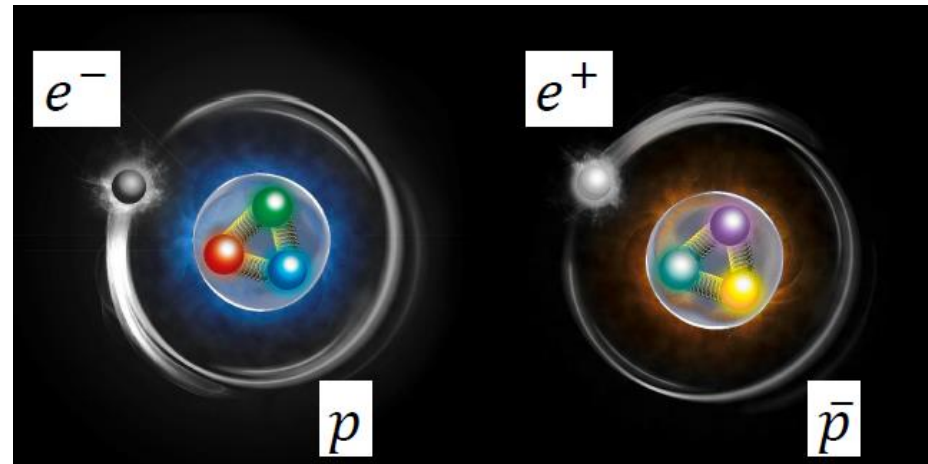
Asakusa in Tokyo

The ASACUSA-Cusp Experiment



What do we study

- The ASACUSA Cusp experiment focuses on producing a beam of antihydrogen for the measurement of the transition frequency ground state hyperfine splitting in a field free region.



Why antimatter ?

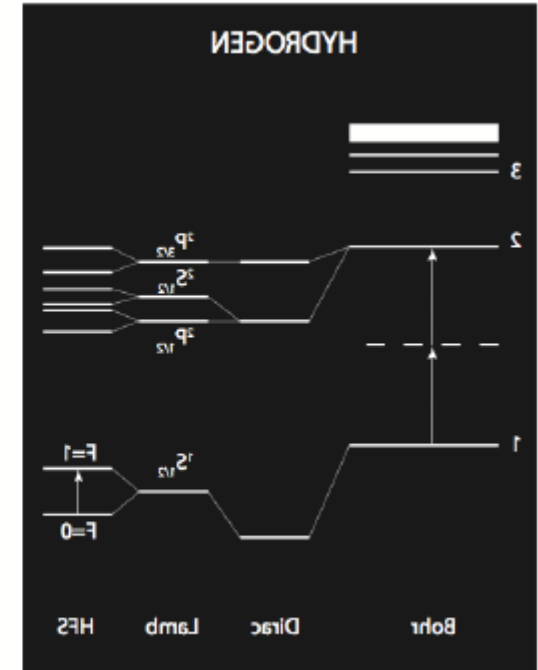
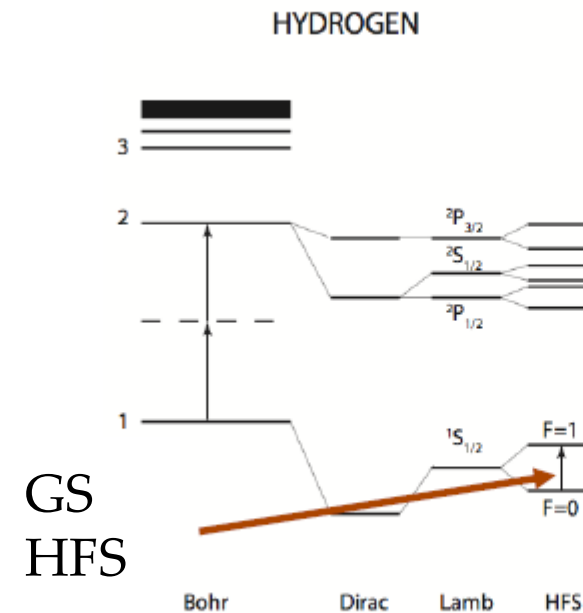
- The standard model of particle physics predicts matter-antimatter symmetry
 - Matter and antimatter created in equal quantities
 - Charge conjugation – Parity – Time reversal (CPT) unviolated
- Observed matter-antimatter asymmetry in nature

$$\eta = \frac{n_b - n_{\bar{b}}}{n_\gamma} \sim 6.1 \times 10^{-10}$$

- Where n_b and $n_{\bar{b}}$ are number densities of baryons and antibaryons and n_γ the number density of cosmic background radiation photons
- This violation points to physics beyond the standard model

Why the HFS of antihydrogen

- Antihydrogen is the simplest stable* purely antimatter system
 - Amenable to high precision spectroscopic investigation
- Hydrogen is the most well studied systems in the physical sciences
 - $\nu_{\text{HFS}} = 1,420,405.7513768(1) \text{ kHz}$



* If kept far enough away from any matter

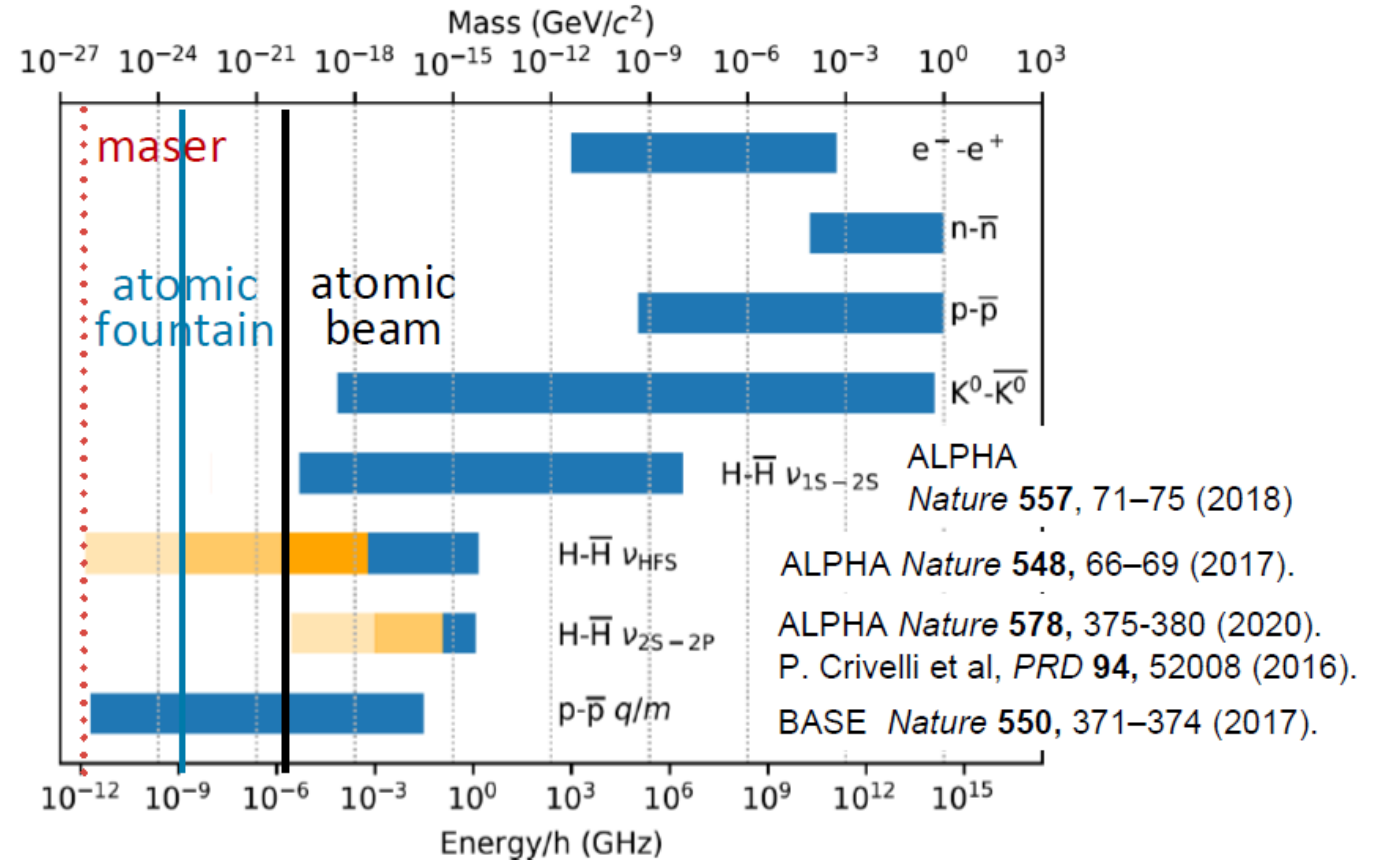
Karshenboim, Savely G. 'Precision Physics of Simple Atoms: QED Tests, Nuclear Structure and Fundamental Constants'. Physics Reports 422, no. 1 (1 December 2005): 1–63.
<https://doi.org/10.1016/j.physrep.2005.08.008>.

Why the HFS of antihydrogen

•Key :

- Right edge: value
- Bar length: relative precision
- Left edge: absolute sensitivity
 - Blue: measured
 - Orange: planned
 - Yellow: potentially measurable

See
<https://arxiv.org/2111.0456v2>
 for details

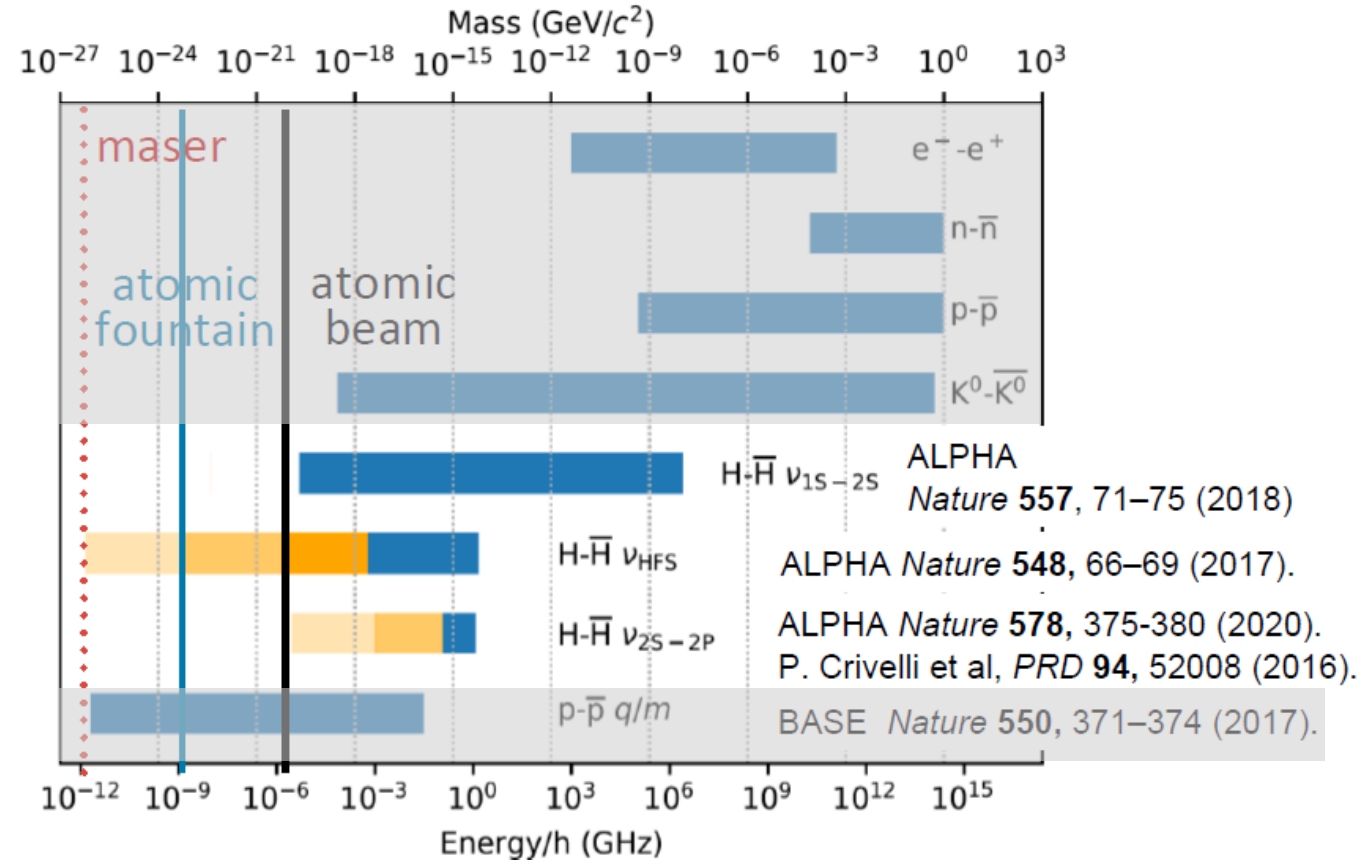


Why the HFS of antihydrogen

•Key :

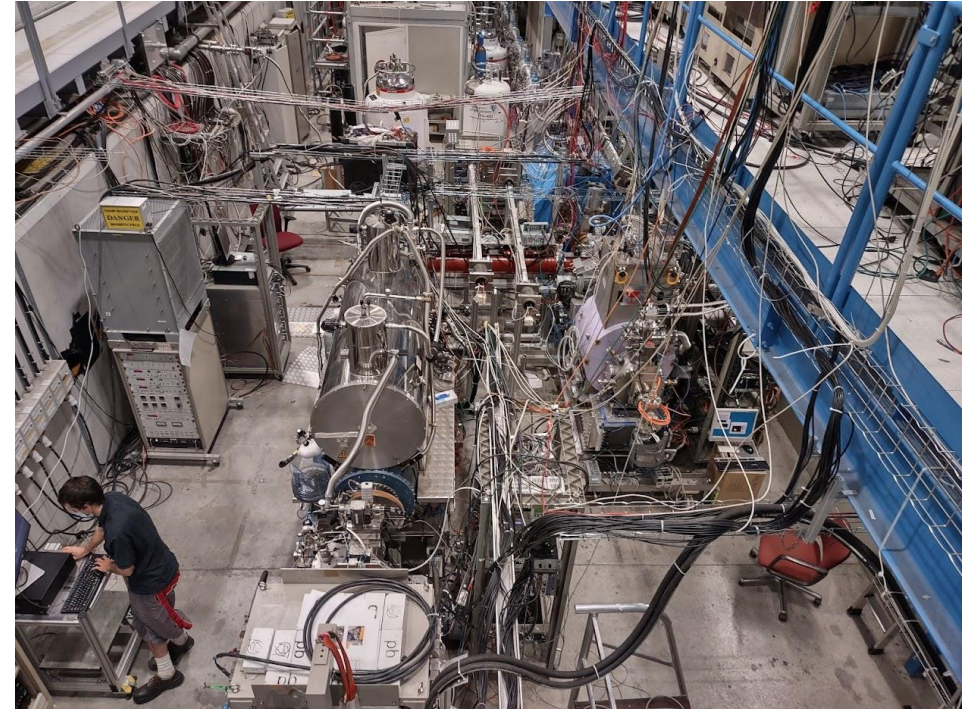
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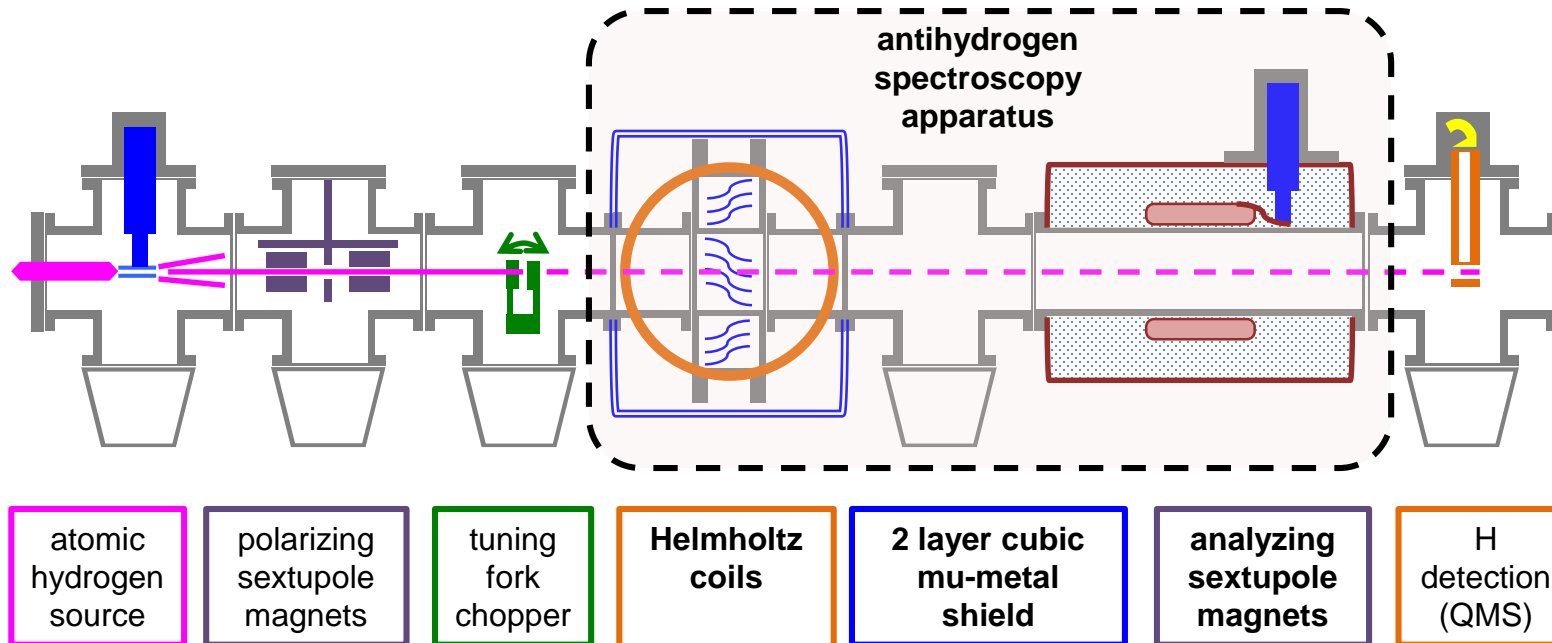
Overview

- The method
- Status of the experiment
- Work during LS2
- The beginning of the ELENA era



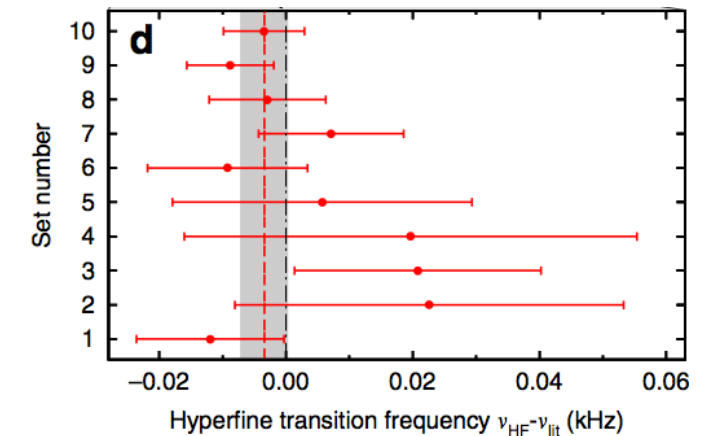
This talk will concentrate on the production and characterisation of an antihydrogen beam for the hyperfine spectroscopy measurement

Hydrogen Spectroscopy



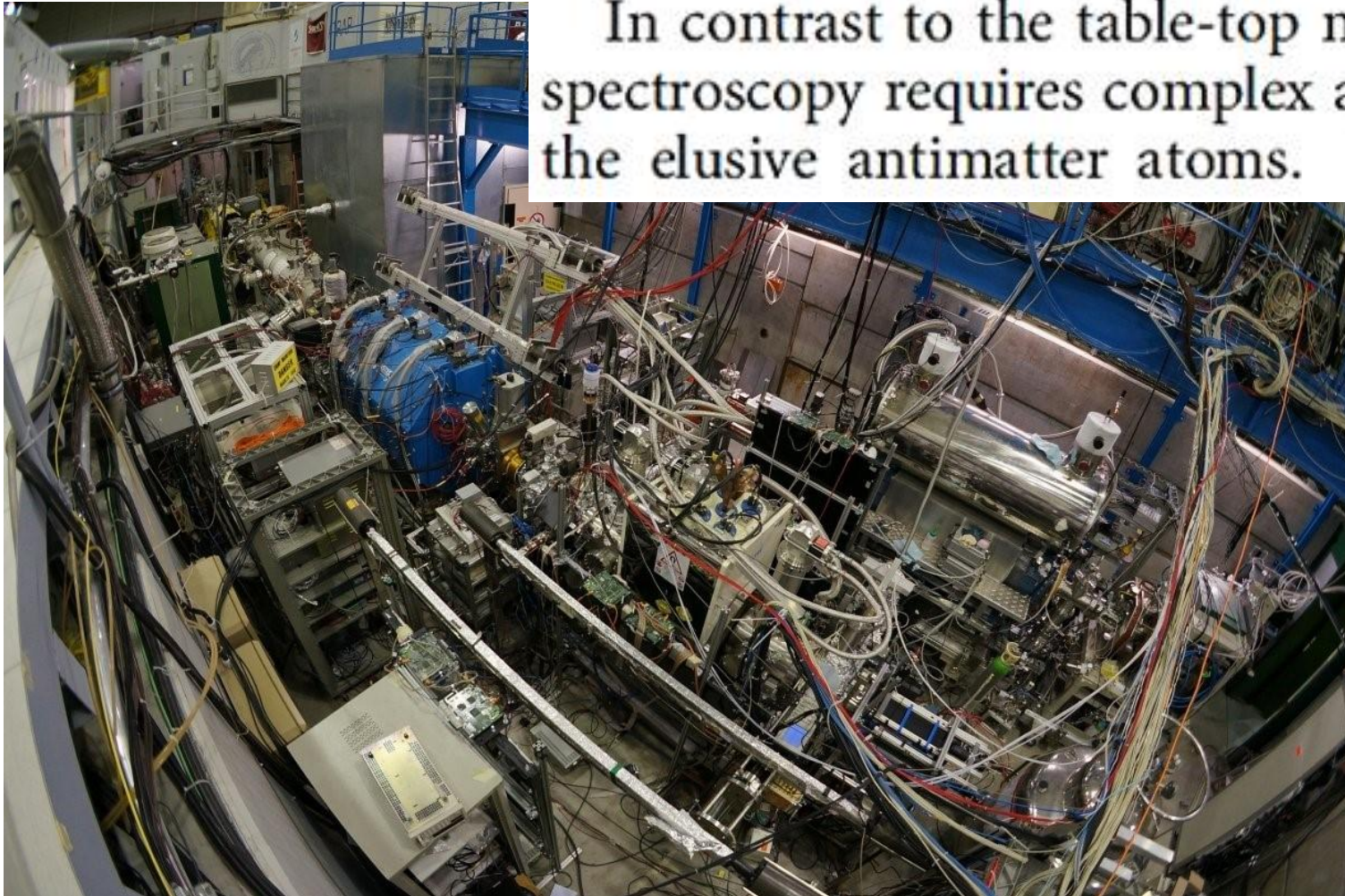
Main Results

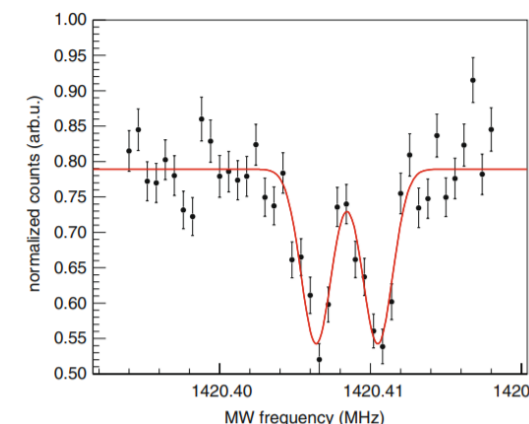
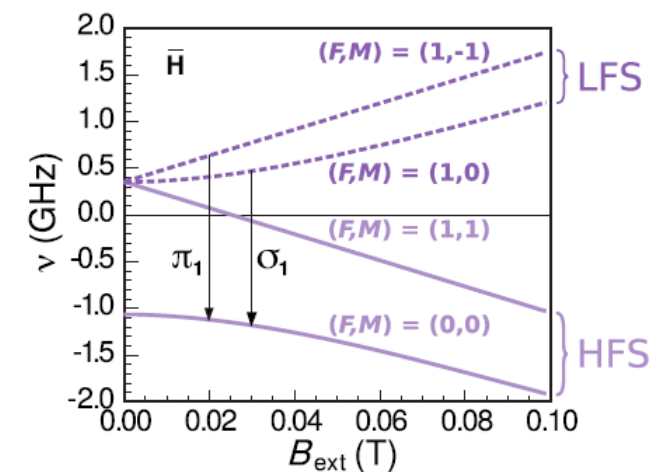
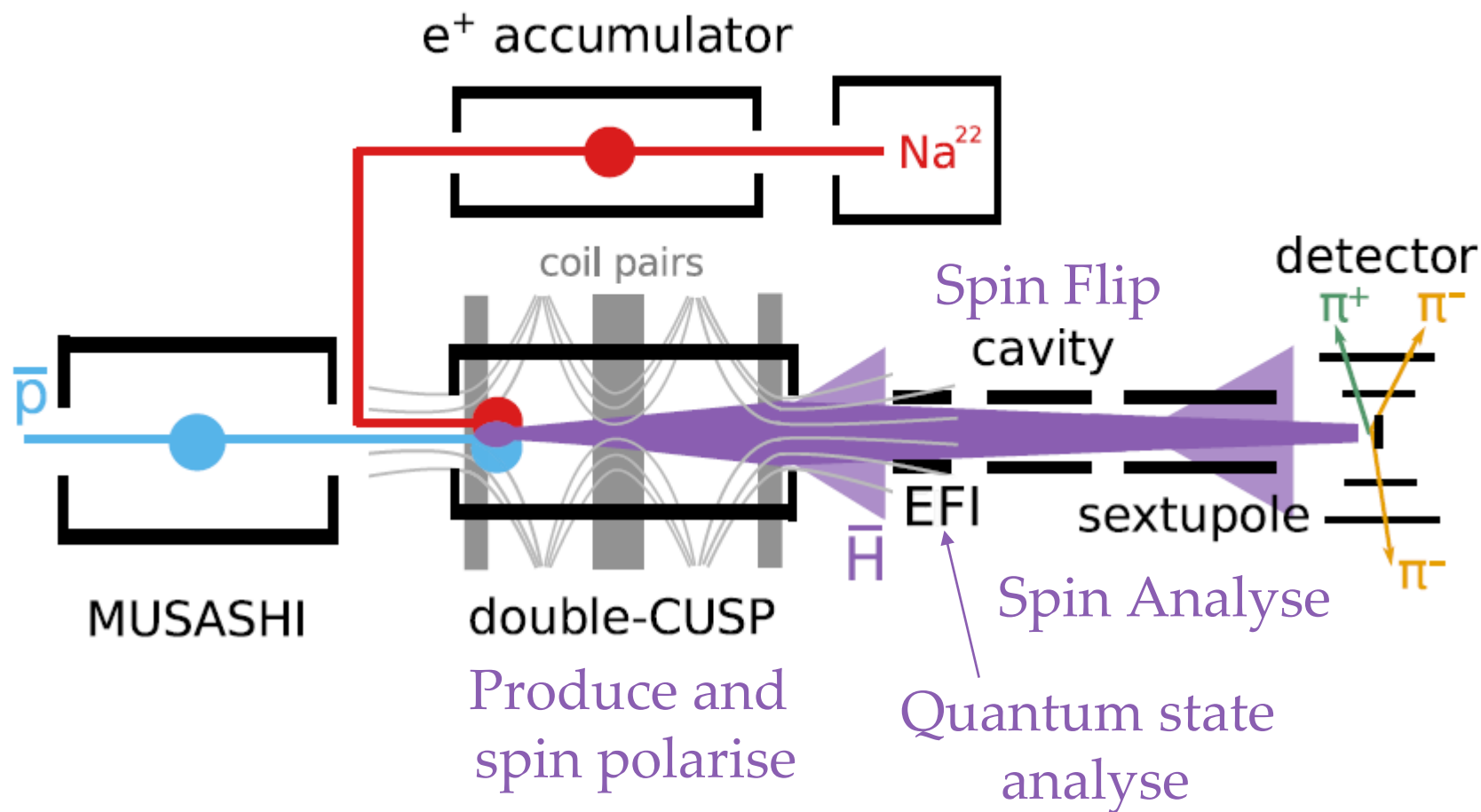
- line-shape understood in detail
- systematic effects? not on few ppb-level
- 8000 Hbar events on detector for ppm result (conservative estimate for σ -transition)



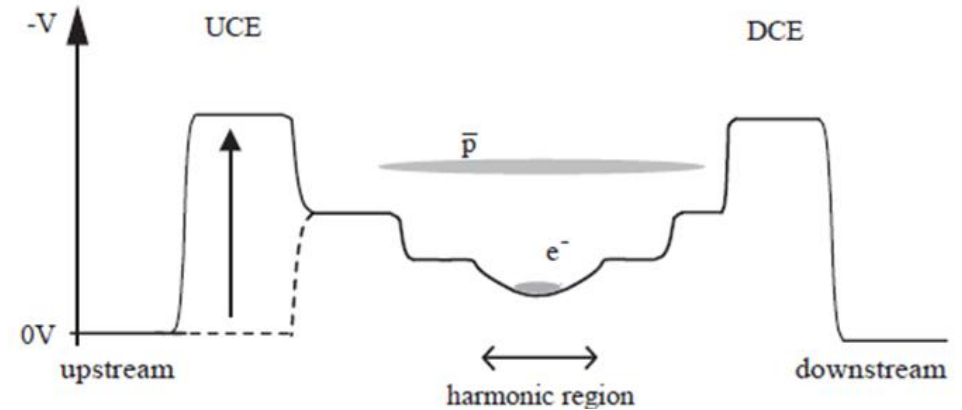
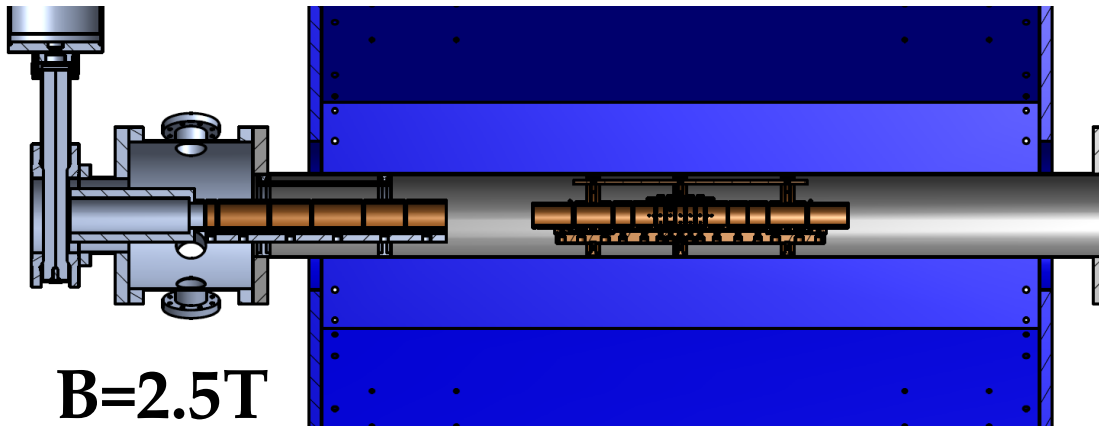
Nature Communications,
published on 12th June '17
DOI: 10.1038/ncomms15749

In contrast to the table-top matter experiments, antihydrogen spectroscopy requires complex apparatuses for the production of the elusive antimatter atoms.



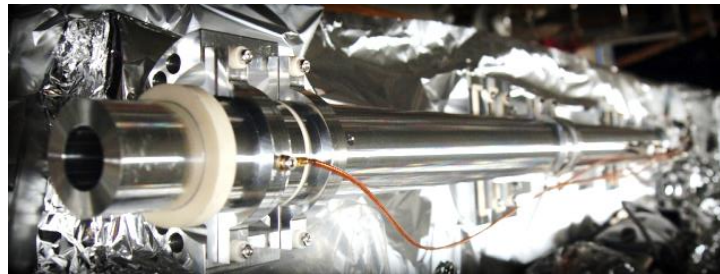
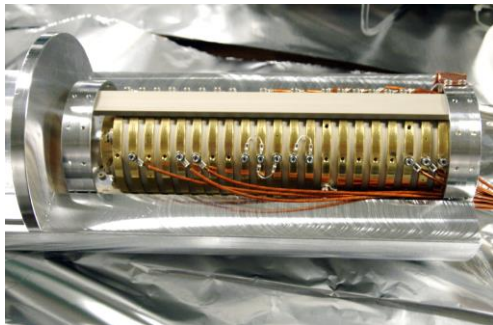
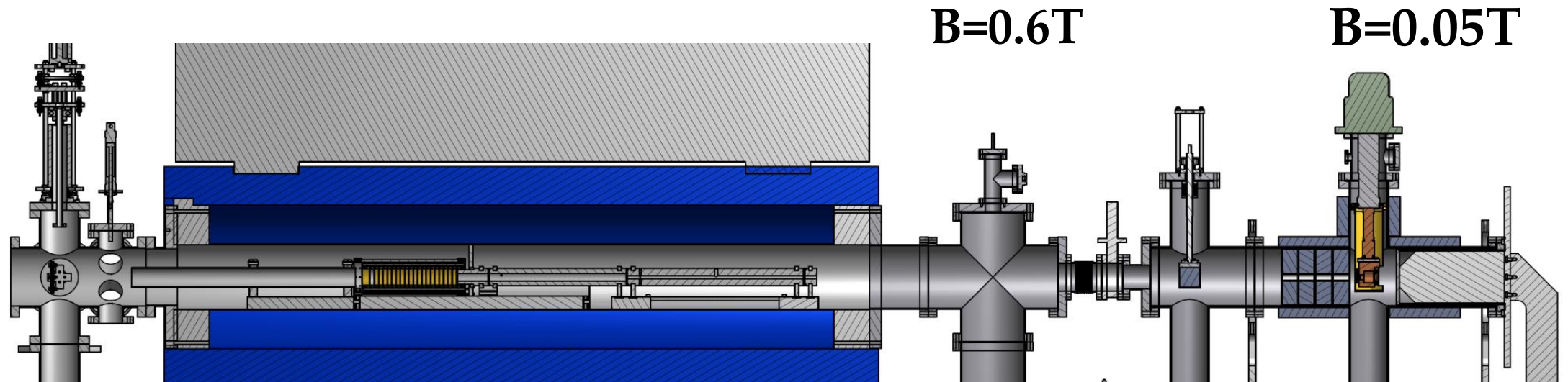


Antiproton Trap - MUSASHI

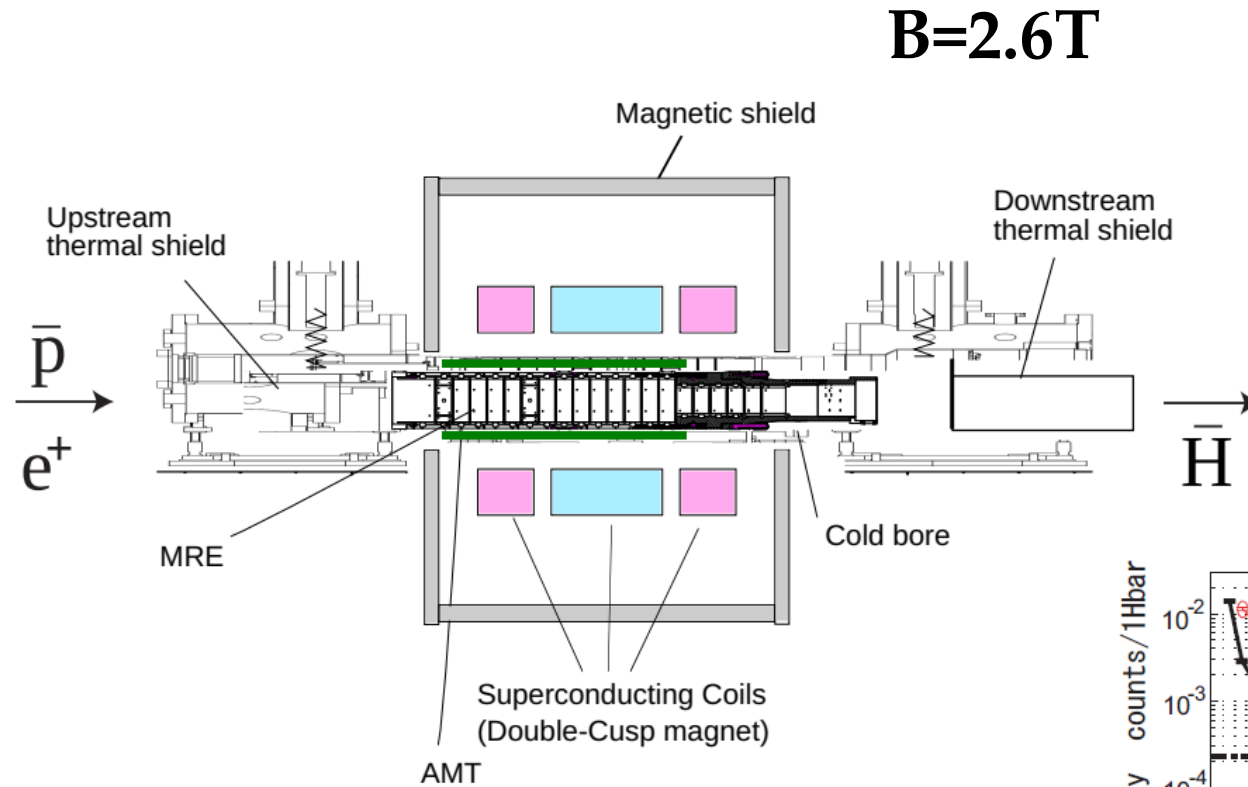
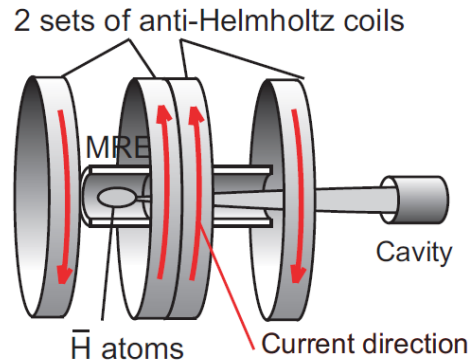


- Degraded foil reduces incoming antiproton energy to $\sim 10\text{keV}$
- Catching bias -12kV
- Antiprotons are cooled with $\sim 3 \times 10^8$ electrons
- 1-2 million antiprotons are trapped per AD cycle
- Antiprotons can be extracted with eV energies.

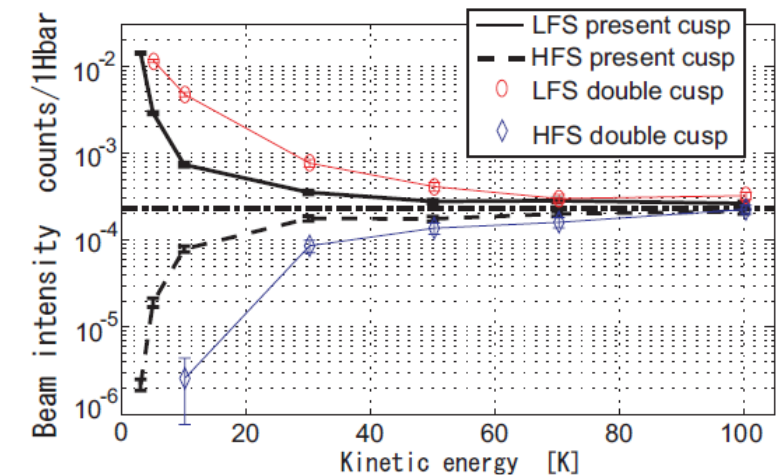
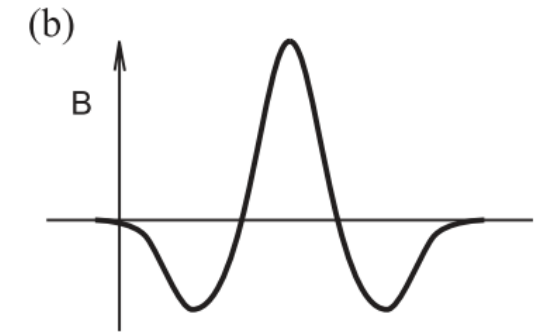
Positron Trap



Double Cusp Trap

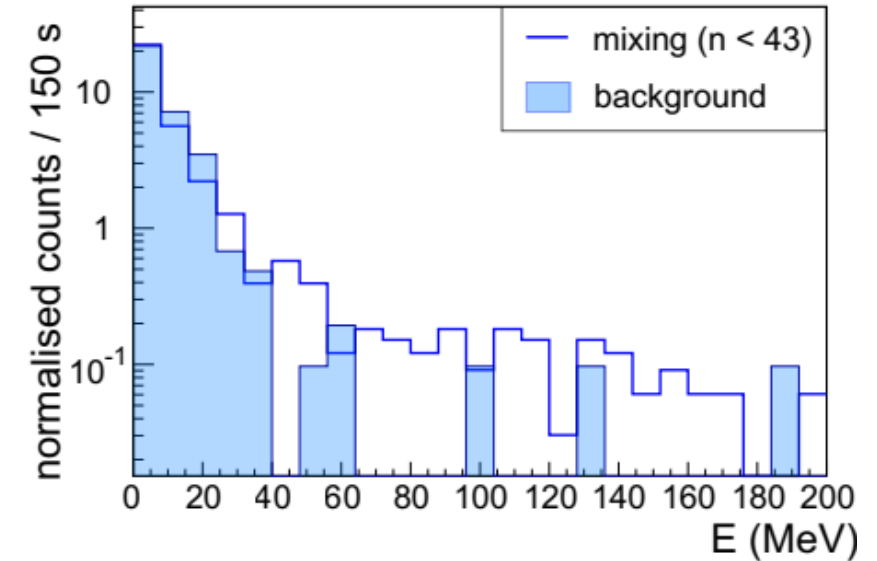
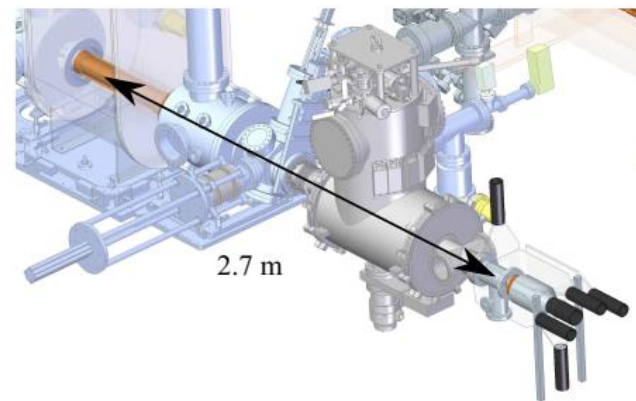
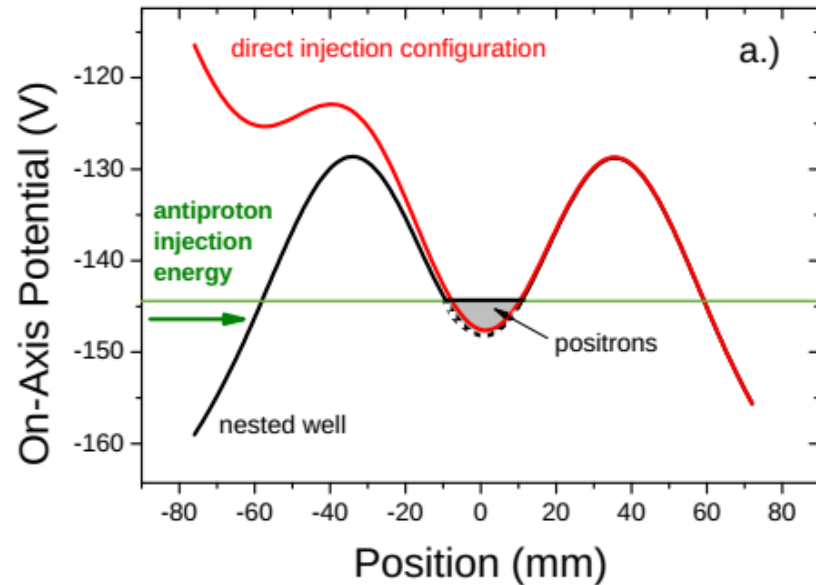


$B=2.6T$



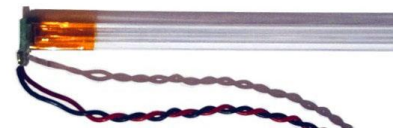
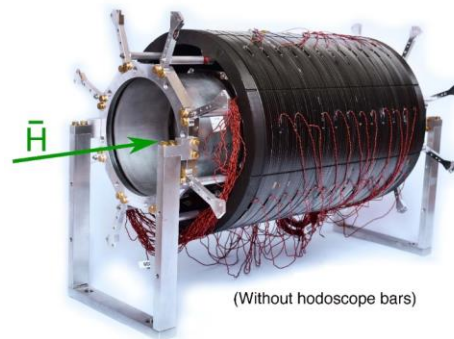
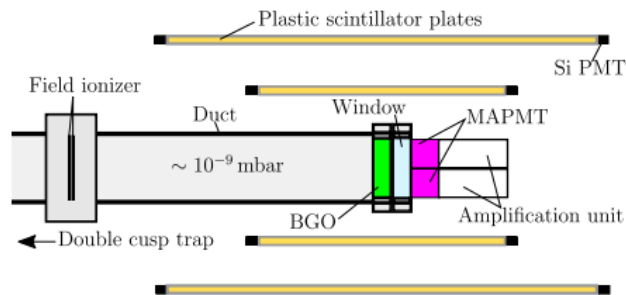
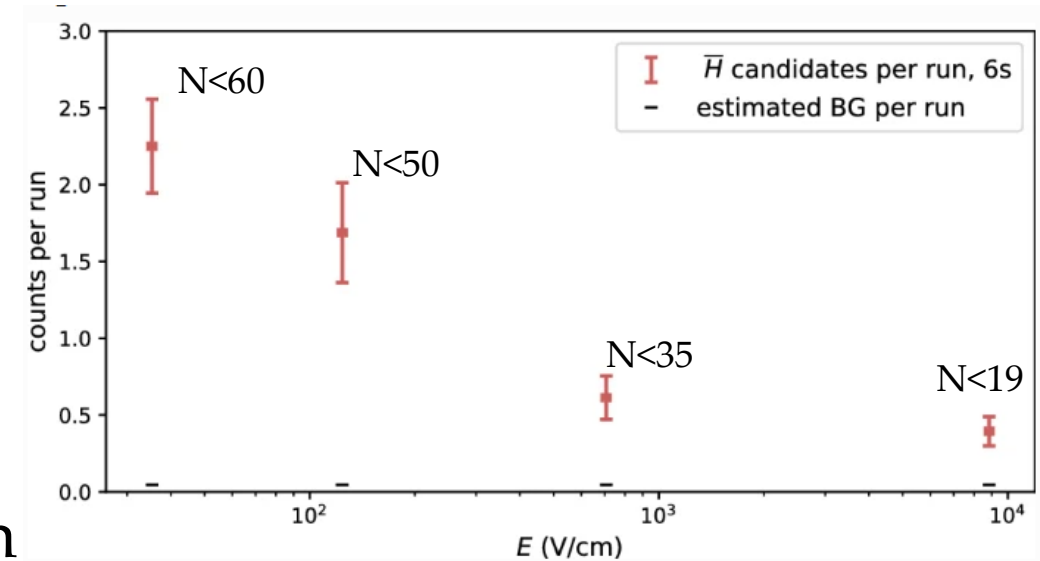
Radics, B., S. Ishikawa, N. Kuroda, D. J. Murtagh, Y. Nagata, M. Tajima, S. Van Gorp, et al. "Antihydrogen Synthesis in a Double-CUSP Trap towards Test of the CPT-Symmetry." *Hyperfine Interactions* 237, no. 1 (December 1, 2016): 156.

Detection of the first antihydrogen beam



Measurement of Quantum State Distribution

- The data analysis made use of machine learning techniques
 - Trained with \bar{p} annihilation
 - $\varepsilon_{\bar{p}} \sim 80\%$
 - False cosmic identification 0.25%
 - Cosmic rate $1.6\text{s}^{-1} \rightarrow 0.02$ per 5s mixing run



Status

- Beam observed
 - 0.006 s^{-1} ($n < 60$)
- Internal quantum number distribution measured
 - 0.001 s^{-1} ($n=1$) (4 h^{-1}) if $v=1000 \text{ ms}^{-1}$
- **How do we improve the beam intensity and the number of ground state atoms?**

How to make more Hbar?

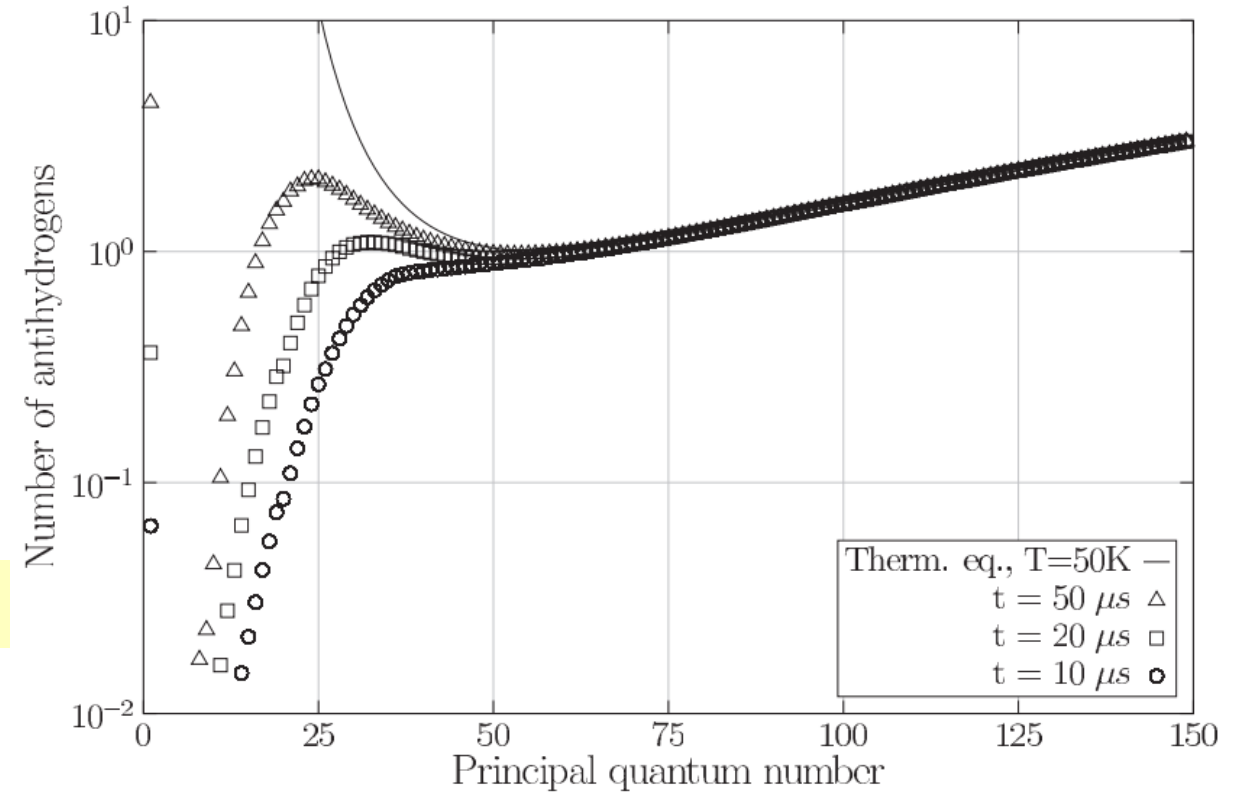
- Modelling of the formation and scattering of antihydrogen formation

$$\frac{dN(i)}{dt} = [C_{rr}(i) + C_{tbr}^+ n_e^+ n_p^+ - C_{ion}^+(i) n_e^+ N(i) + \sum_{j \neq i} [C_{col}^+(j, i) n_e^+ + C_{str}(j, i)] N(j) - N(i) \sum_{j \neq i} [C_{col}^+(i, j) n_e^+ + C_{str}(i, j)]]$$

Radiative recombination rate $\rightarrow C_{rr}(i)$
 Three Body recombination rate $\rightarrow C_{tbr}^+ n_e^+ n_p^+$
 Ionization Rate $\rightarrow C_{ion}^+(i) n_e^+ N(i)$
 Collisional (de)excitation Rate $\rightarrow C_{col}^+(j, i) n_e^+ + C_{str}(j, i)$
 Simulated/ Spontaneous Transition Rates $\rightarrow C_{col}^+(i, j) n_e^+ + C_{str}(i, j)$

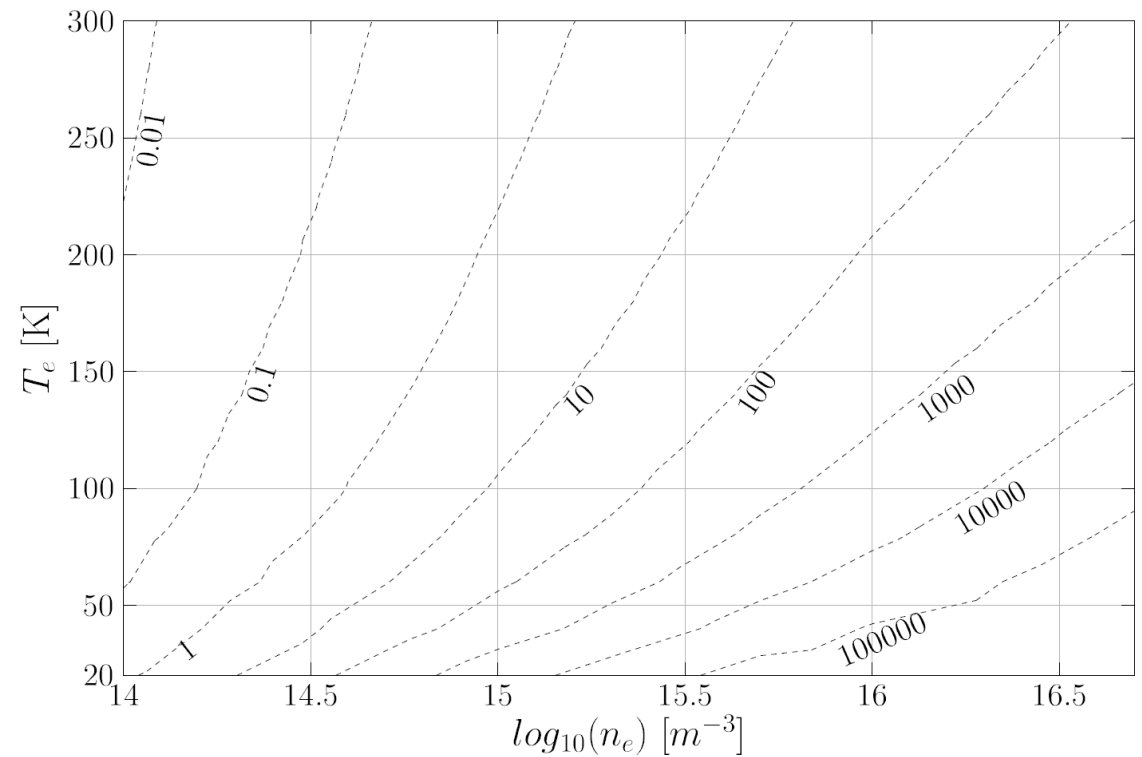
Antiproton Number $\rightarrow n_p^+$
 Positron density $\rightarrow n_e^+$
 Number of Antihydrogen in state $\rightarrow N(i)$

Populating State i (blue box)
 Depopulating State i (yellow box)



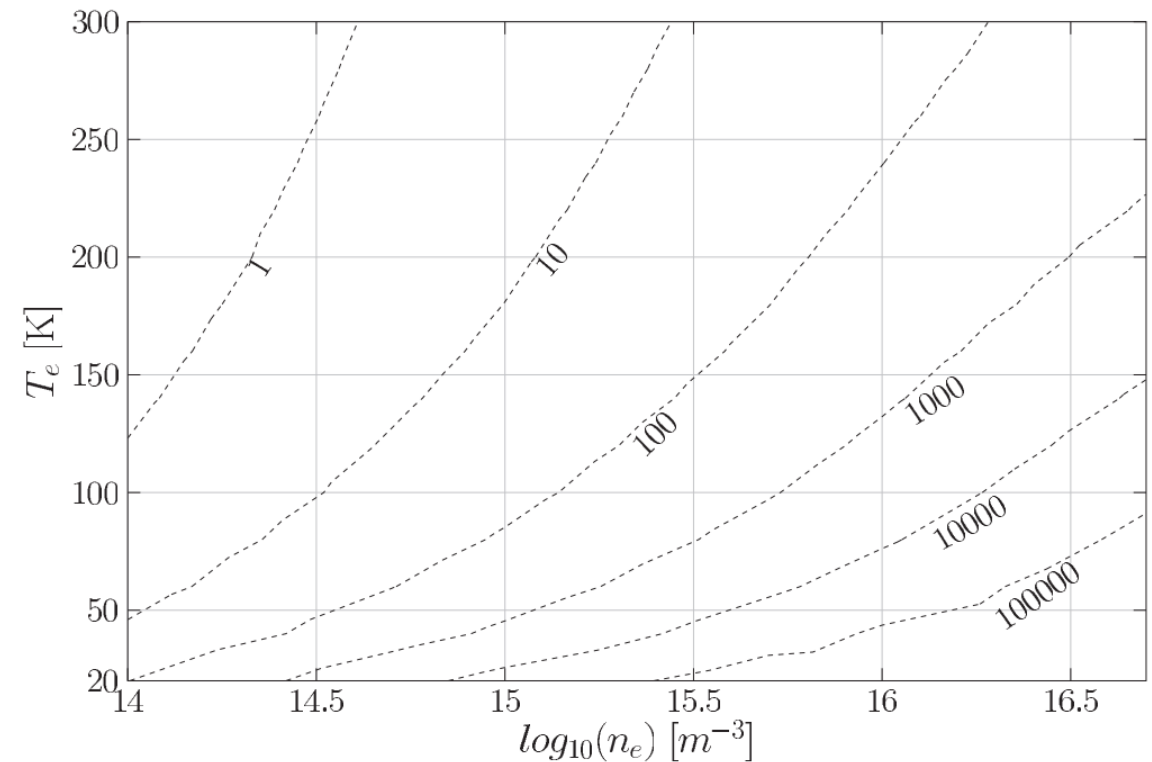
Plasma Temperature

- If the temperature of the positron plasma can be reduced from 200K to 20K we gain **two orders of magnitude** in antihydrogen production



Plasma Temperature

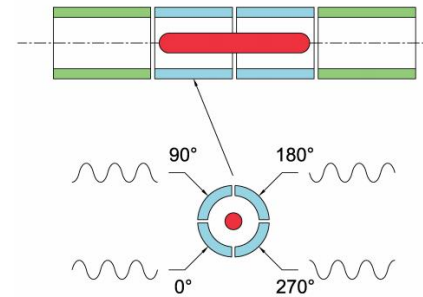
- If the temperature of the positron plasma can be reduced from 200K to 20K we gain **two orders of magnitude** in antihydrogen production
- If the positron plasma is colder then the antiprotons will be colder producing slower antihydrogen
 - **Longer to cascade → Lower n states!**



Control of electron/positron plasma properties

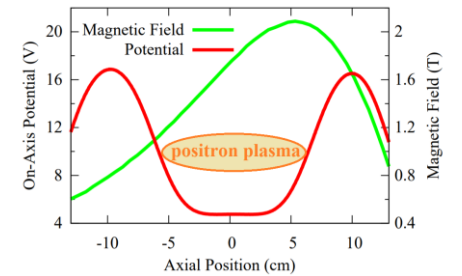
- Long Shutdown 2 (LS2) ran from December 2018 to August 2021 during this period no antiprotons were available
- A reproducible cold plasma was needed!
- What are we interested in ?
 - Temperature
 - Density
 - Number
 - Length
 - Radius

Rotating Wall



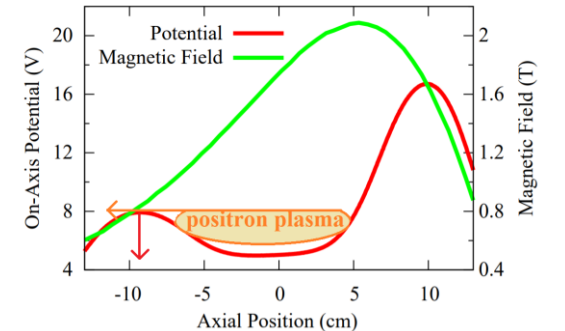
Radius

Electrical Potential



Length

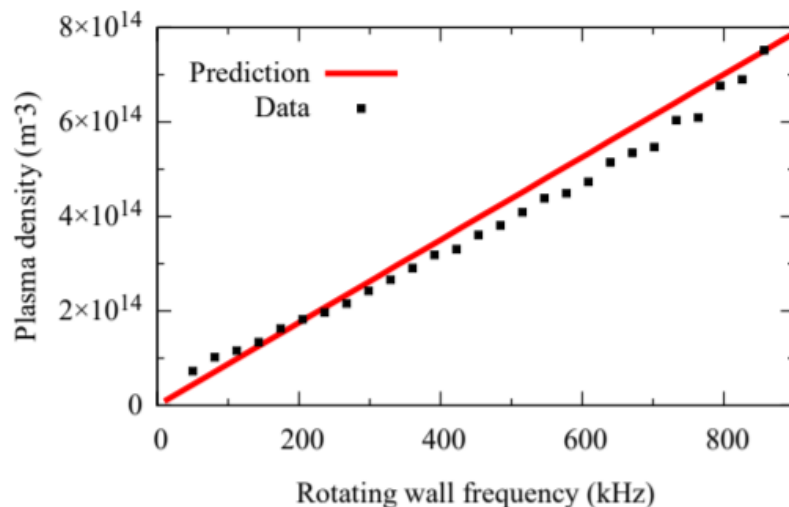
Evaporation



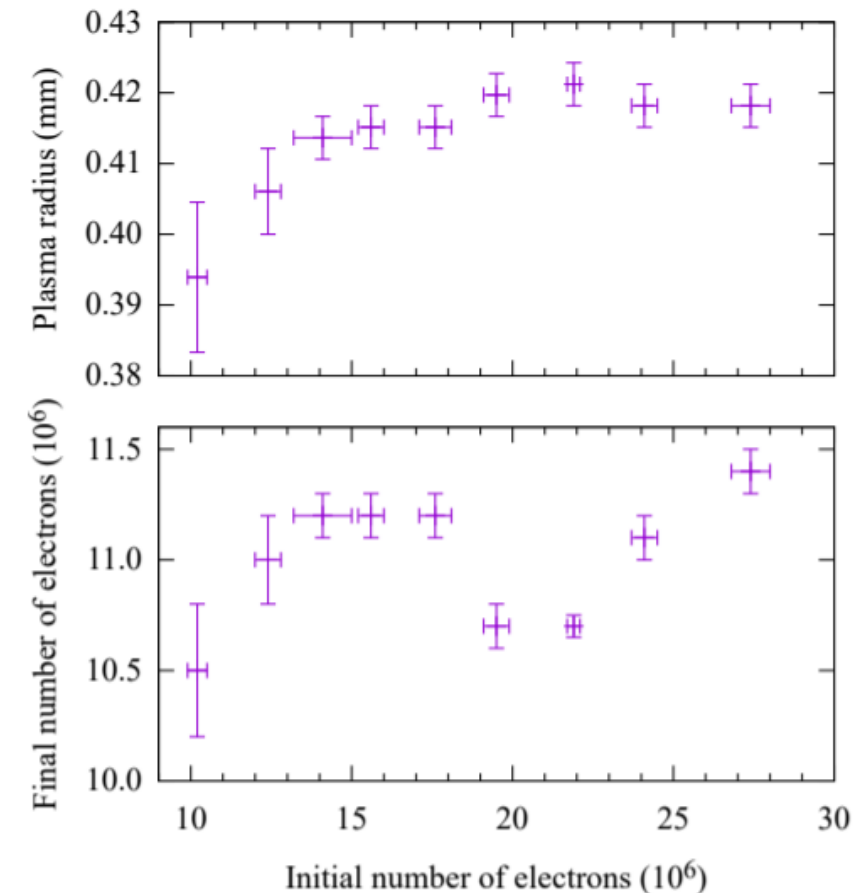
Number &
Temperature

Control of electron/positron plasma properties

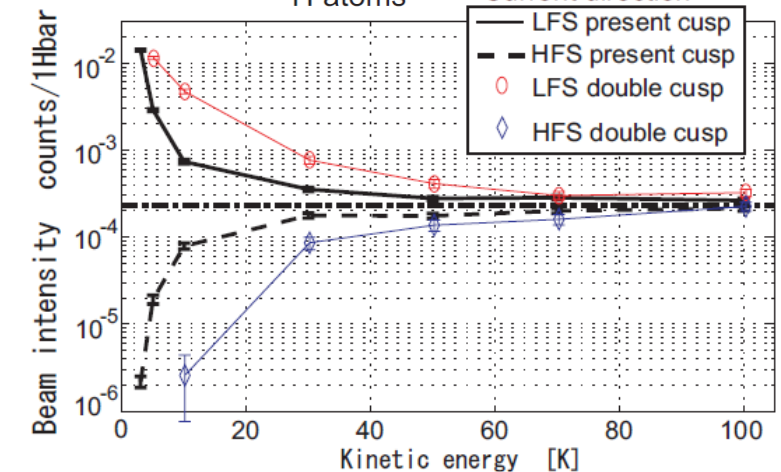
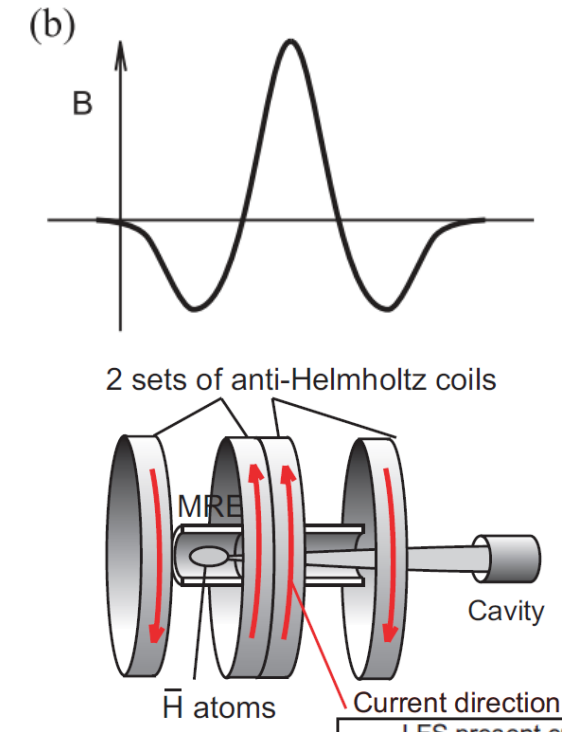
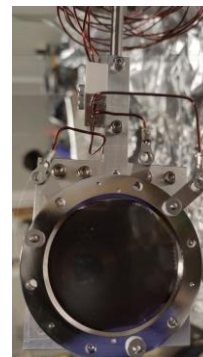
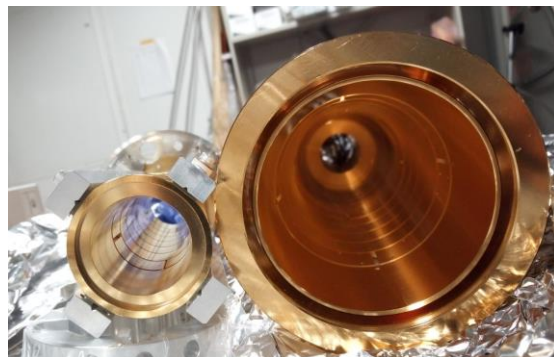
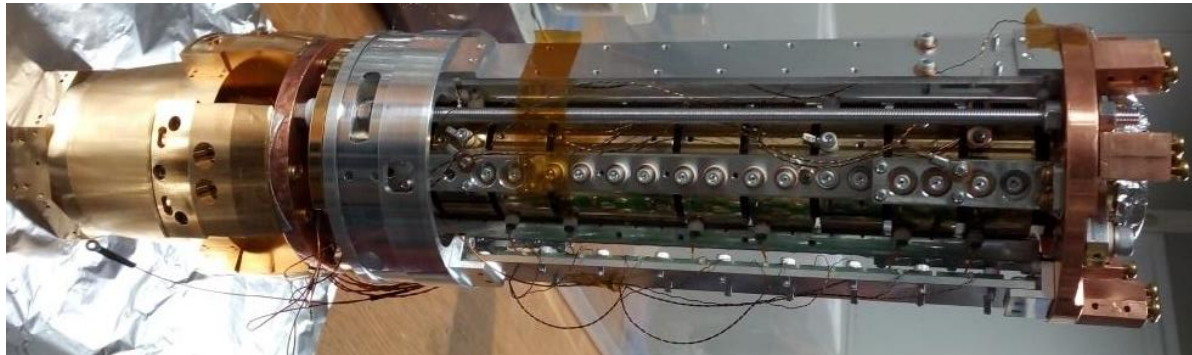
- To produce the same plasma each and every time : we fix a potential, apply strong drive rotating wall and evaporate → **SDR-EVC**



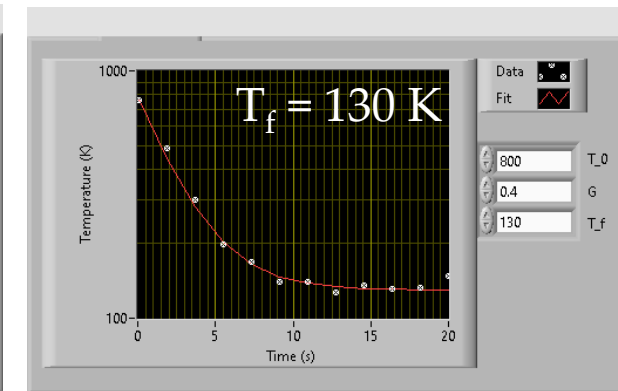
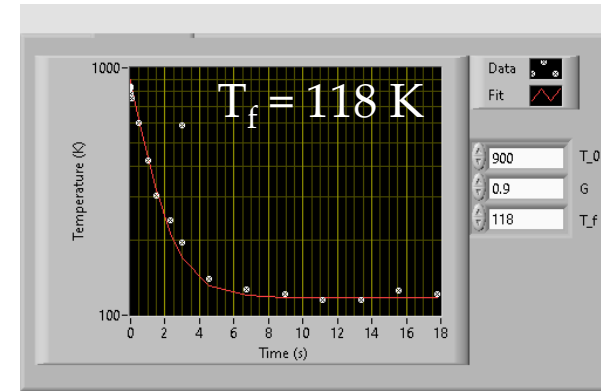
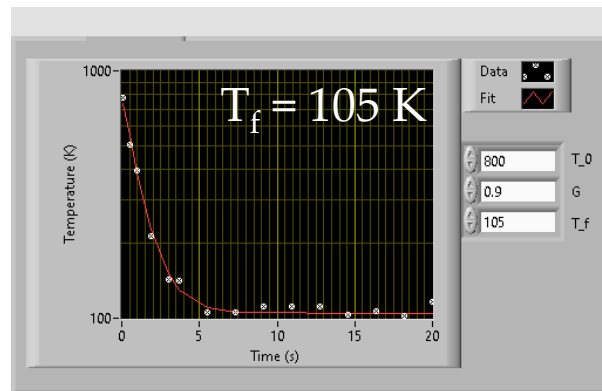
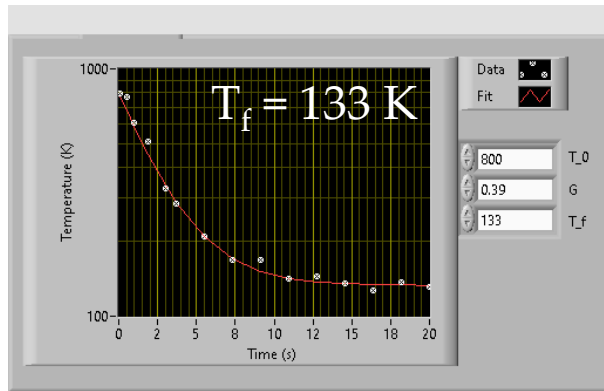
Property	Mean	SD
r_p (mm)	0.417	0.003
T (K)	360	30
N_f (10^6)	11.0	0.3



Double Cusp Trap



Plasma Cooling Rate

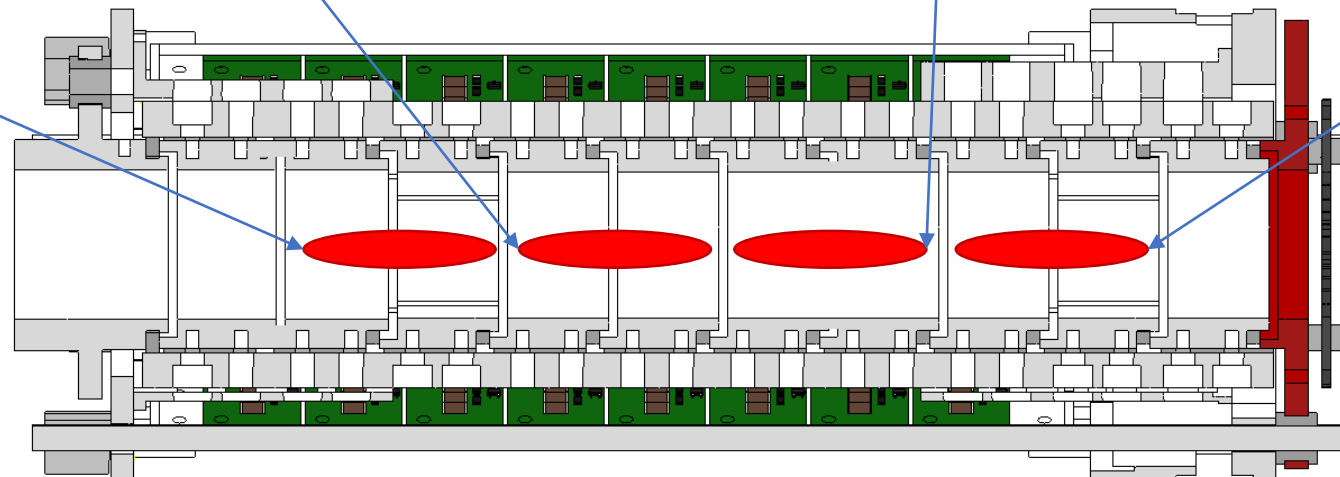


$$T_w = 35K$$

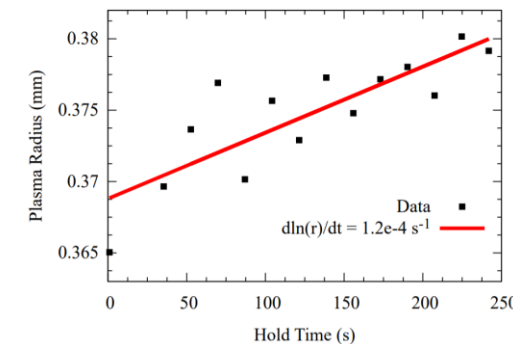
Heating Rate

$$T_f = T_w + H / G$$

Cyclotron Cooling Rate



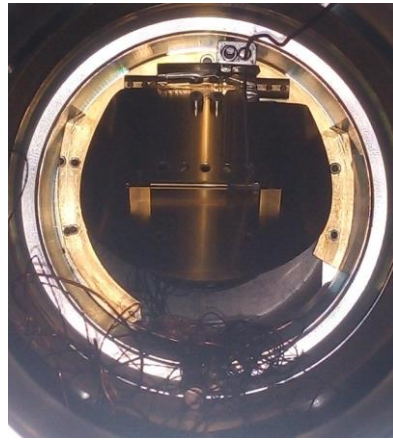
Name:	u13	u12	u11	u10	u9	u8	u7	u6	u5	u4	u3
Filter:	pulsed	slow	slow	none	slow	slow	slow	slow	none	slow	slow



Plasma cooling : radiation environment



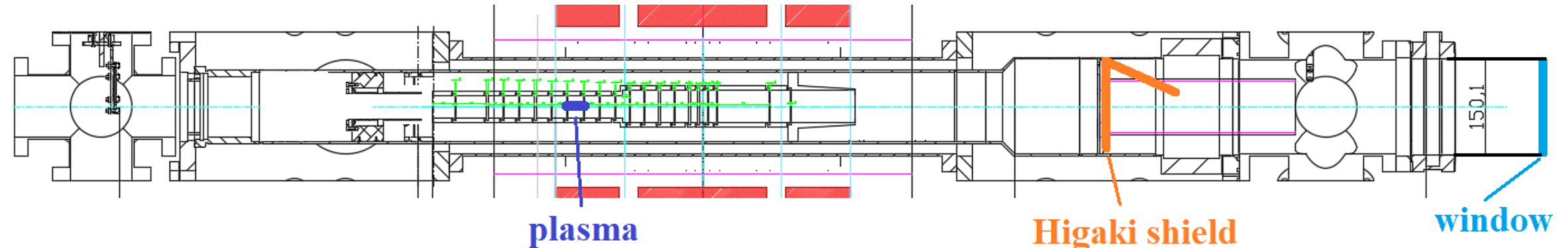
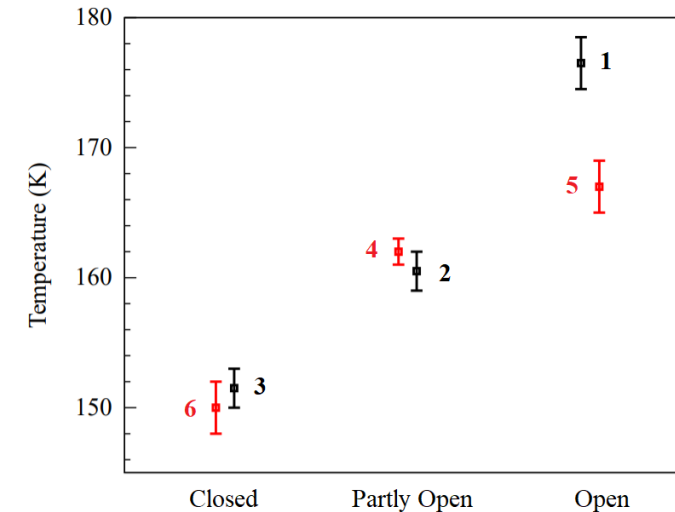
Closed



Partly Open

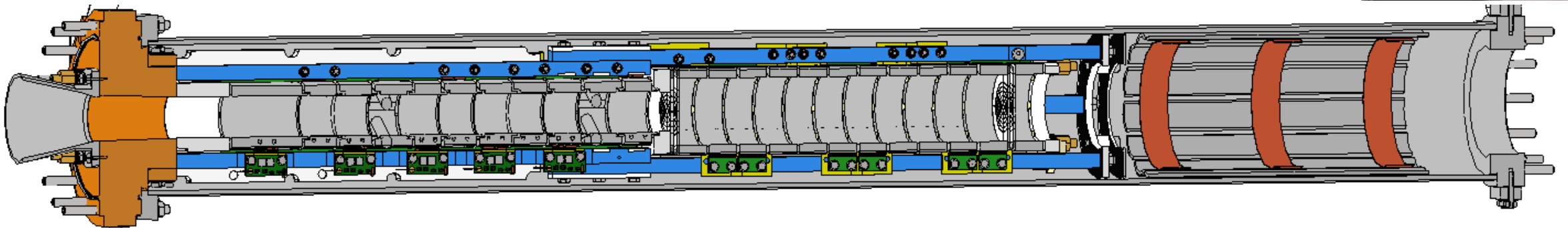
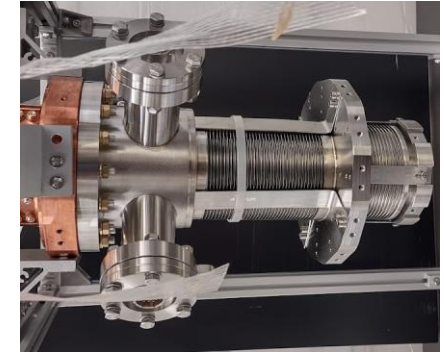
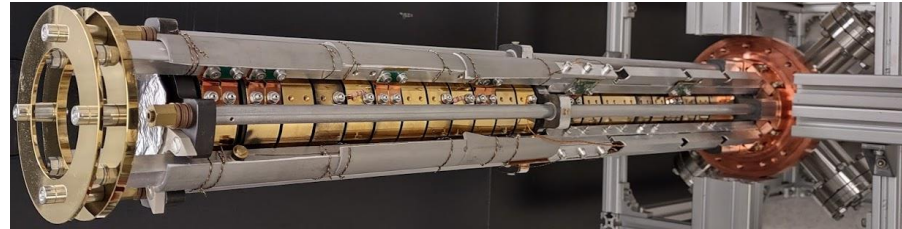


Open



Double Cusp Trap

Ceramic “bracelet” to absorb cyclotron radiation from the plasma

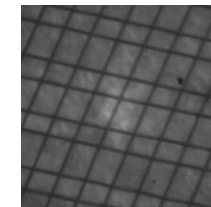


High transparency copper mesh to reflect incoming microwaves

0.25 mm pitch

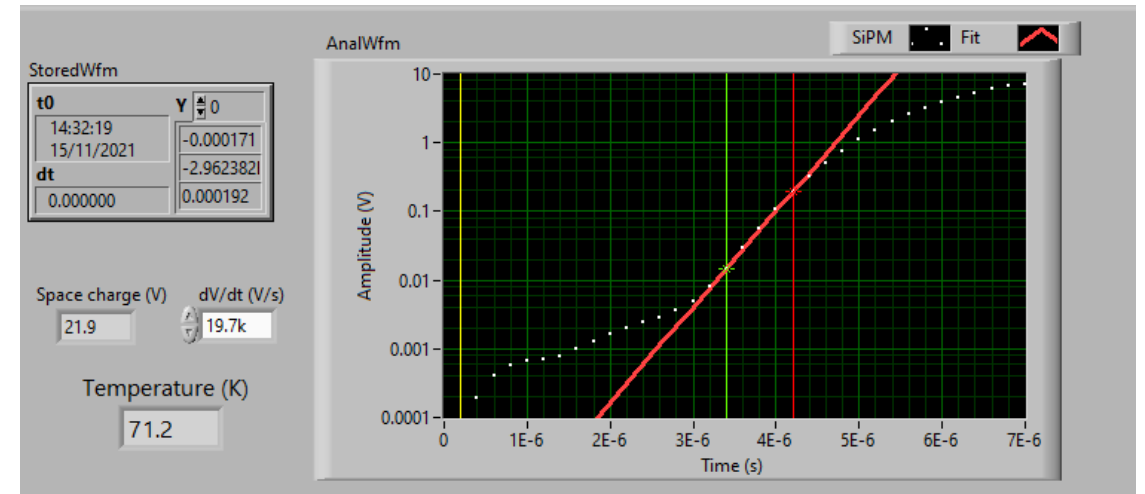
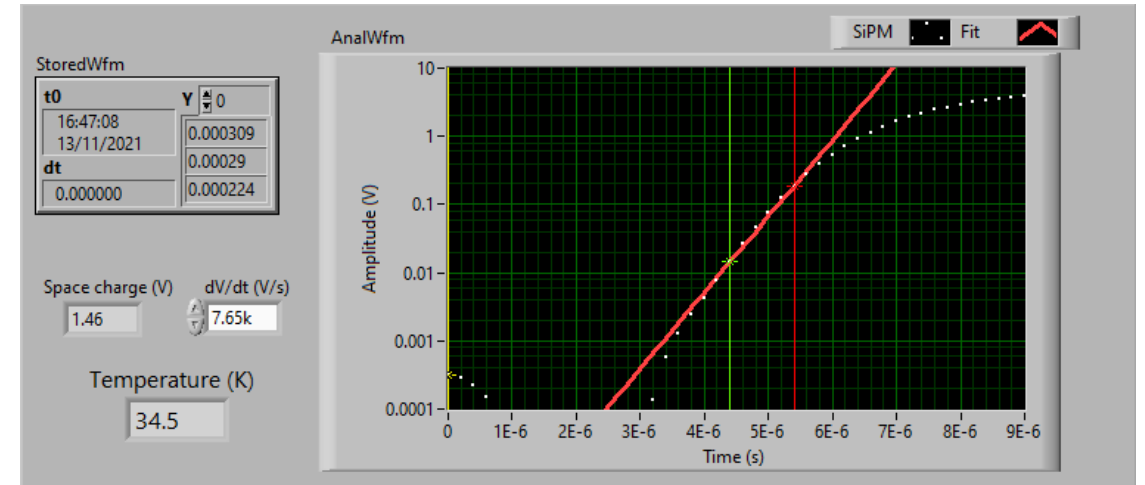
0.03 mm wire diameter

>20 dB attenuation at 60 GHz



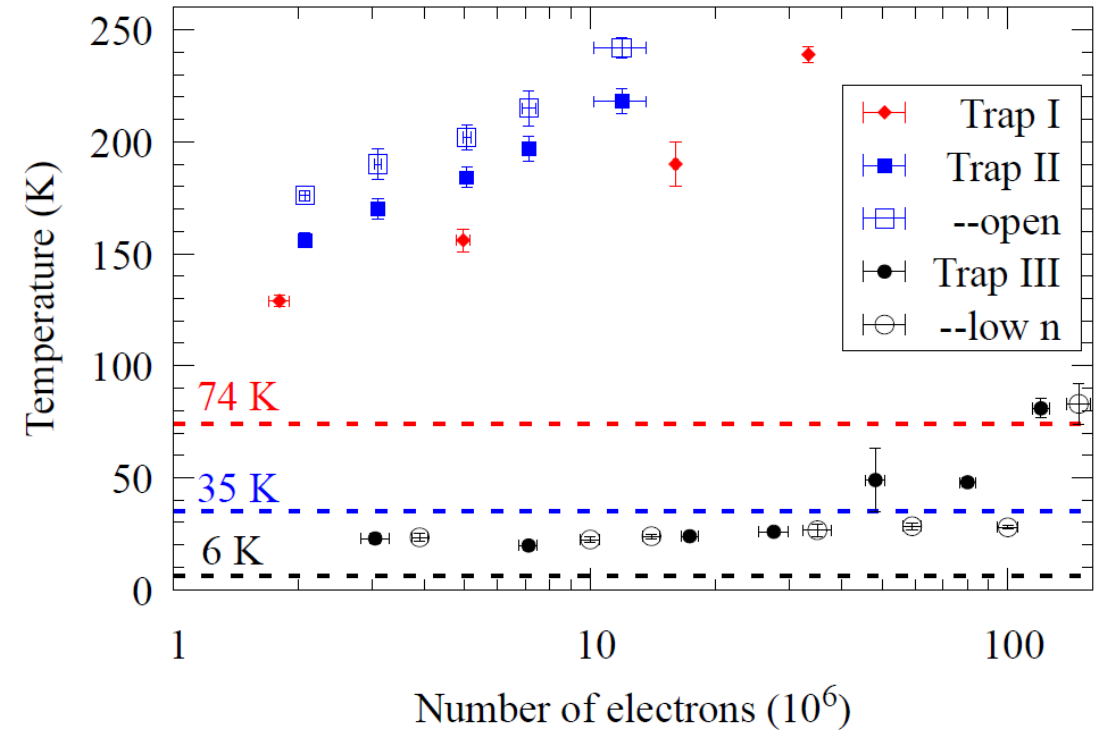
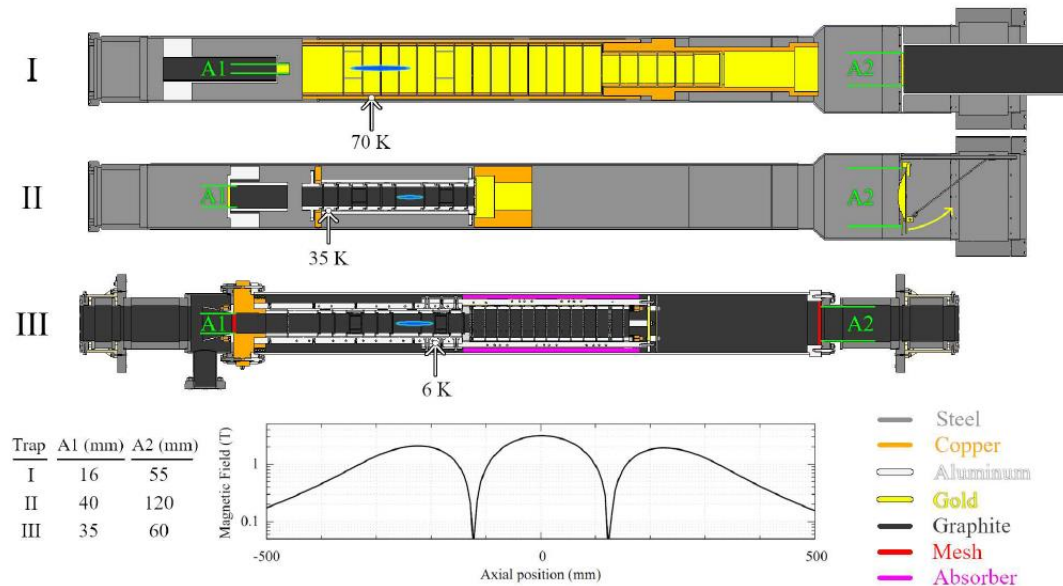
Double Cusp Trap

- First results
 - Electrode temperature 5-6 K
 - Approximately 10 K colder than the previous design
 - Plasma Temperatures
 - 3-4 million electron 35 K (previously 150 K)
 - 60 million electrons 70 K



Double Cusp Trap

- Electrode temperature 5-6 K
- Approximately 10 K colder than the previous design

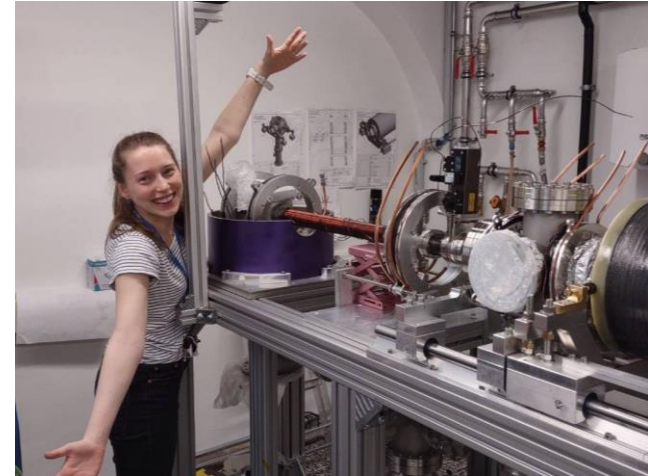
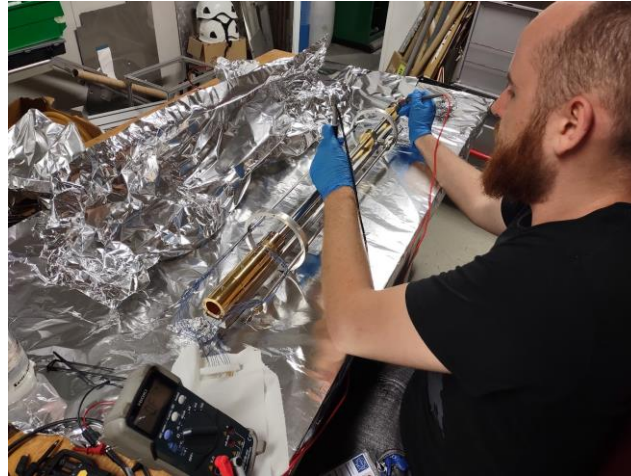


Outlook

- The ASACUSA Cusp collaboration successfully detected a very weak and highly excited beam of antihydrogen 2.7 meters from their mixing region
- To increase the brightness and produce a ground state beam, colder mixing plasmas were required
- Thus far we have succeeded in producing cold electron plasmas in the new mixing trap
- This year we are working to repeat this with positrons and the cold positrons mixed with antiprotons
- There have been many more developments which I haven't had time to talk about
- We hope the future is cold (plasmas) and bright (beams)

Acknowledgements

- The SMI positron group
 - Dr Eric Hunter
 - Andreas Lanz
 - Alina Weiser



Full Collaboration

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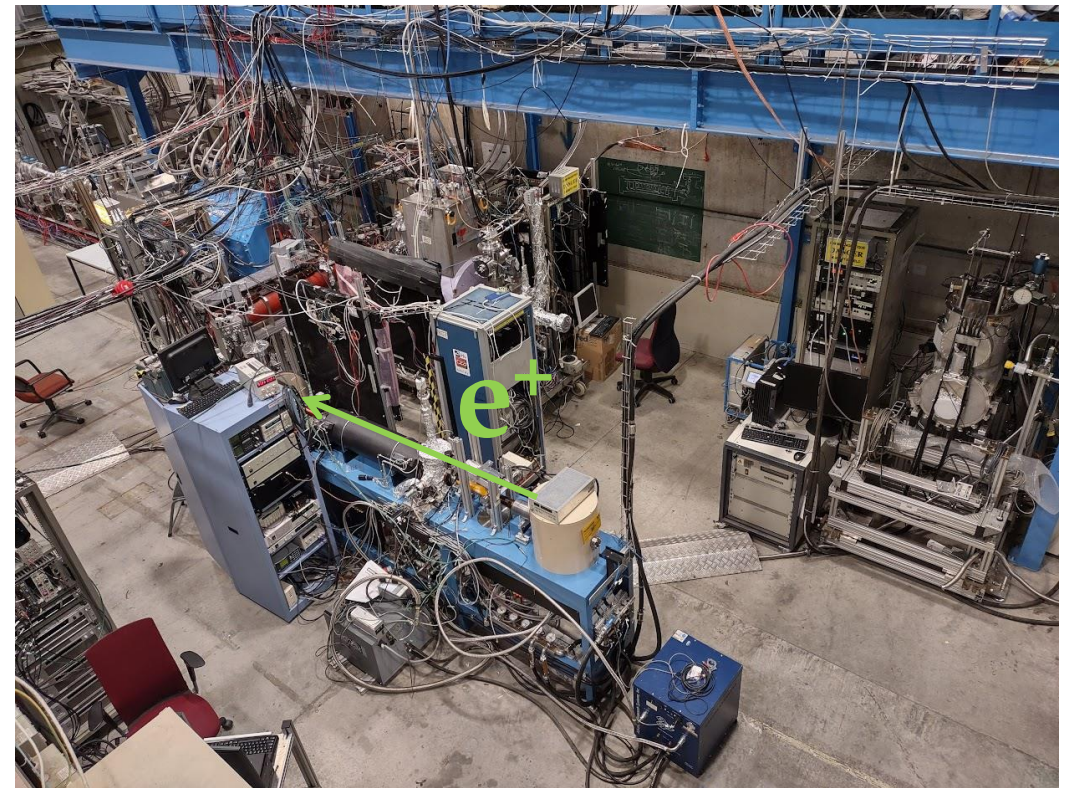
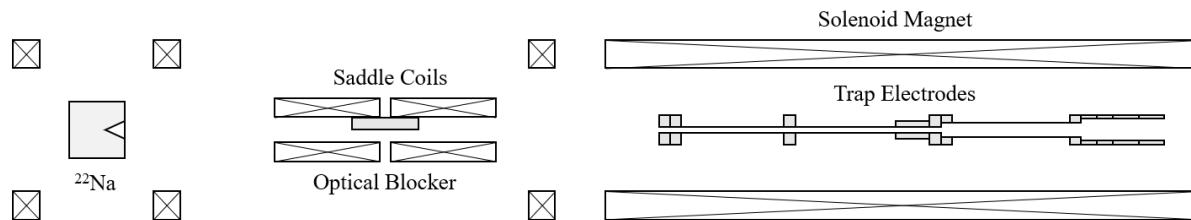
¹³⁾RIKEN Nishina Center for Accelerator-Based Science, 351-0198 Saitama,



Backup

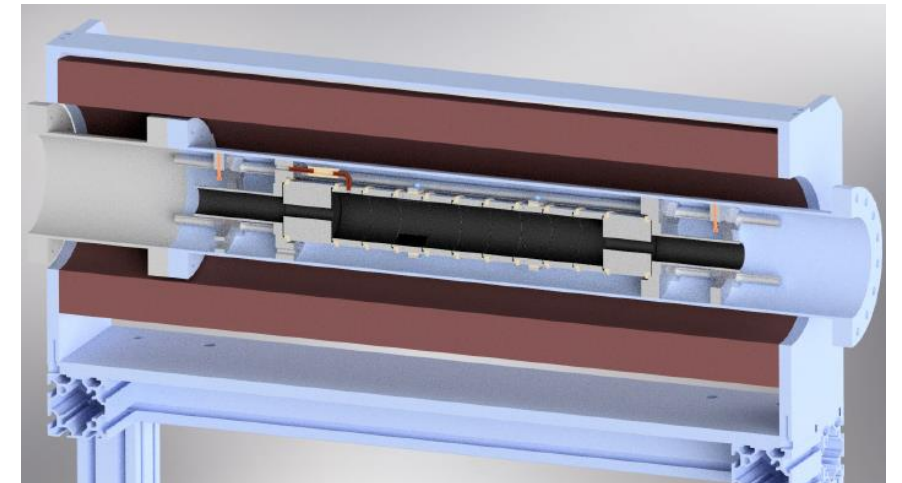
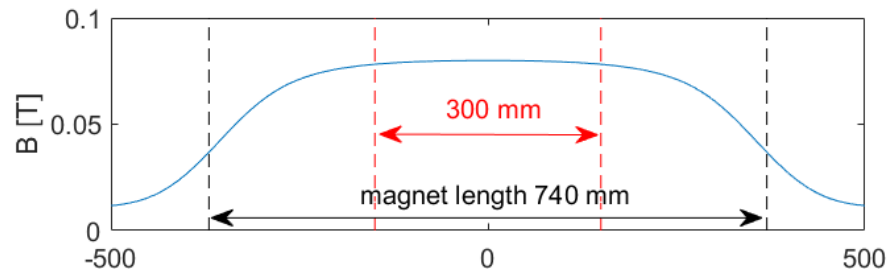
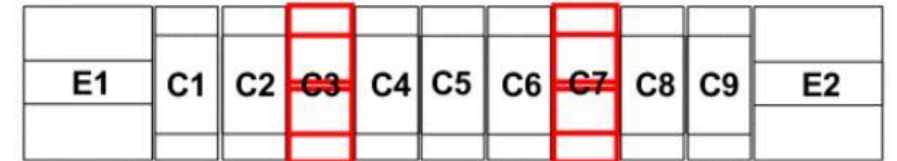
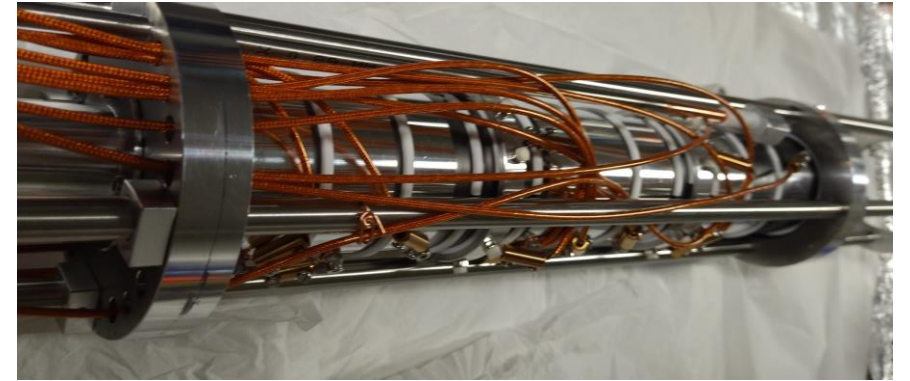
Positron Trap Replaced

The existing positron trap was due to retire in 2021 but due to a lack of personnel in 2020/1 the new FPS trap was installed in 2022 and is currently being developed.

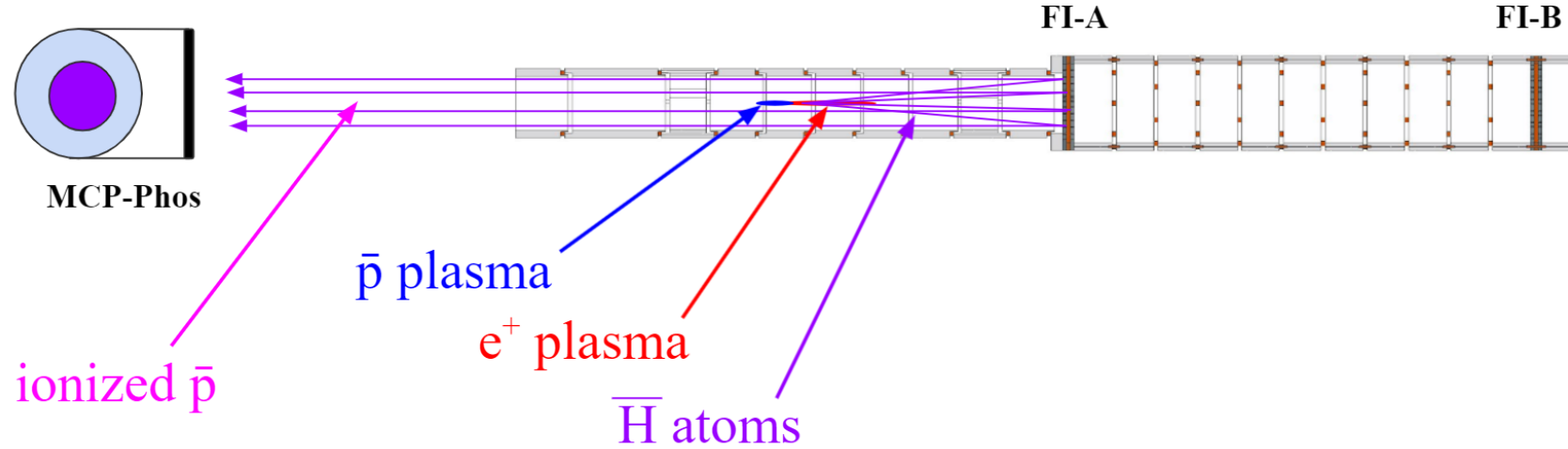


Positron Accumulator

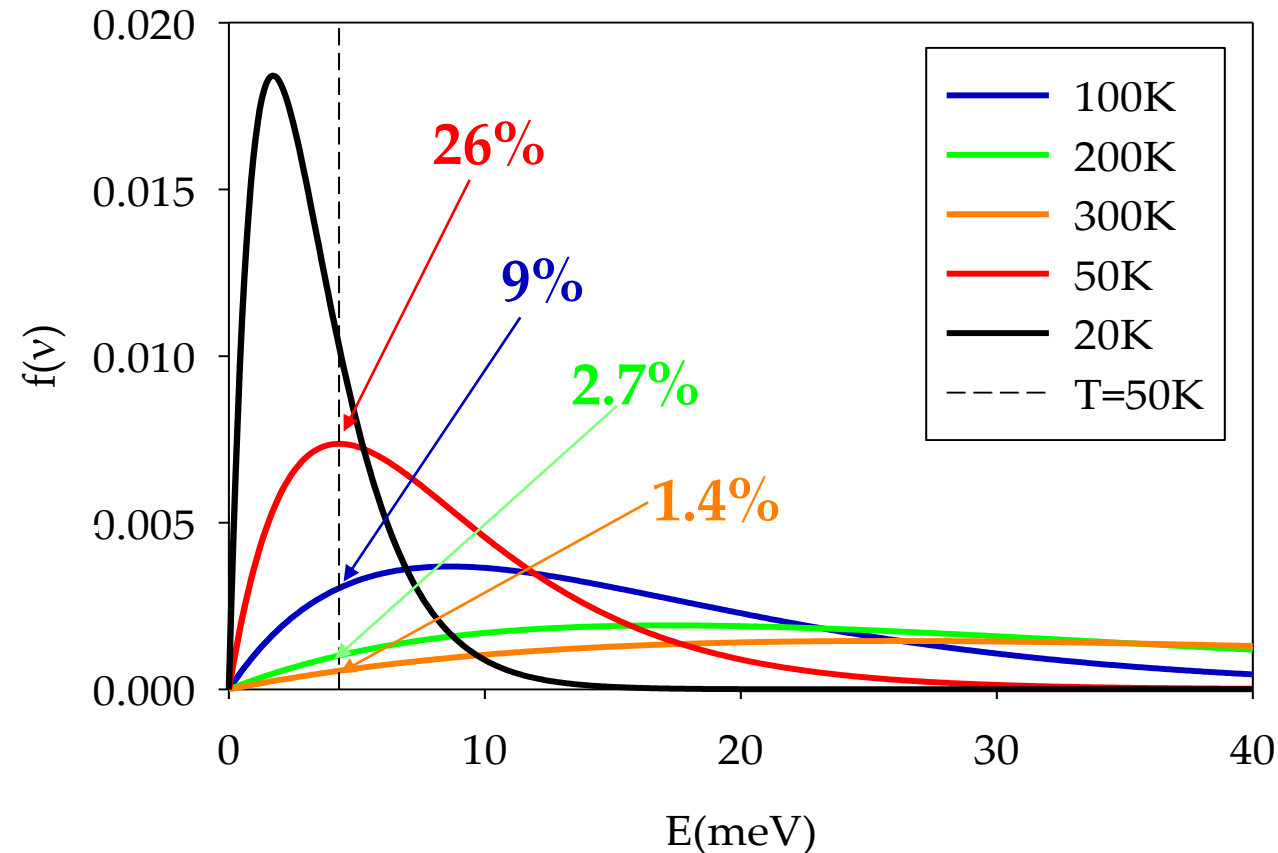
- As the FPS trap has a lifetime of just 2s a new positron accumulator was built in 2021
- At present this device has been installed in the first coil of the positron transfer line and is being commissioned at this moment



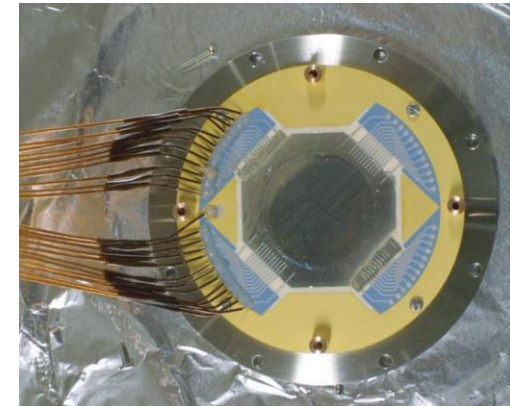
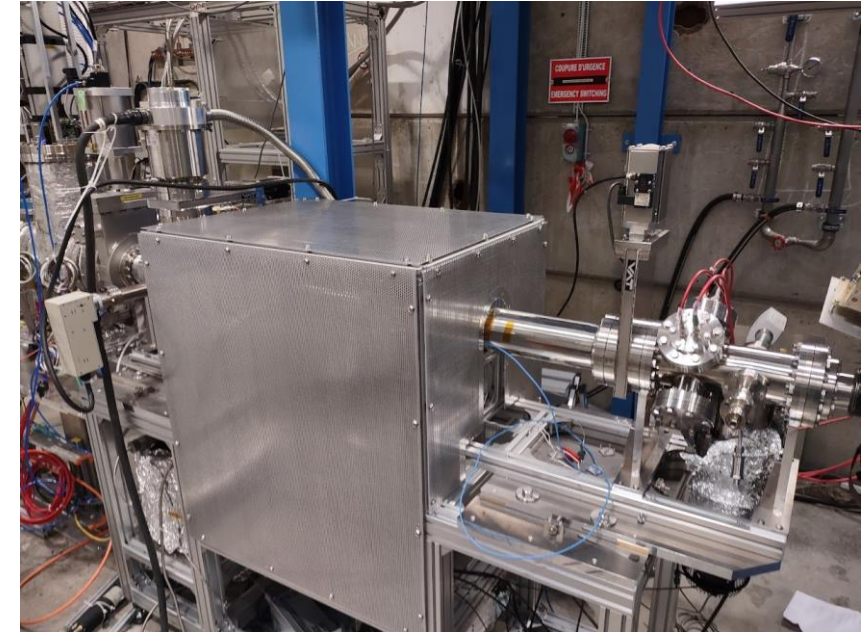
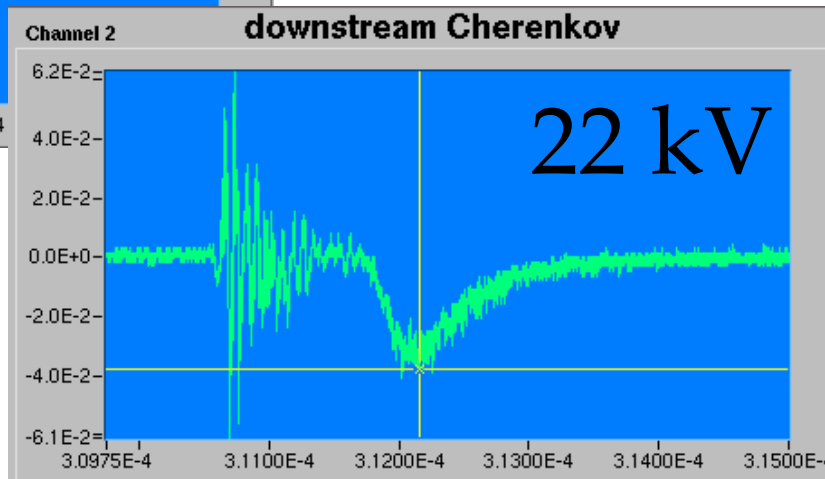
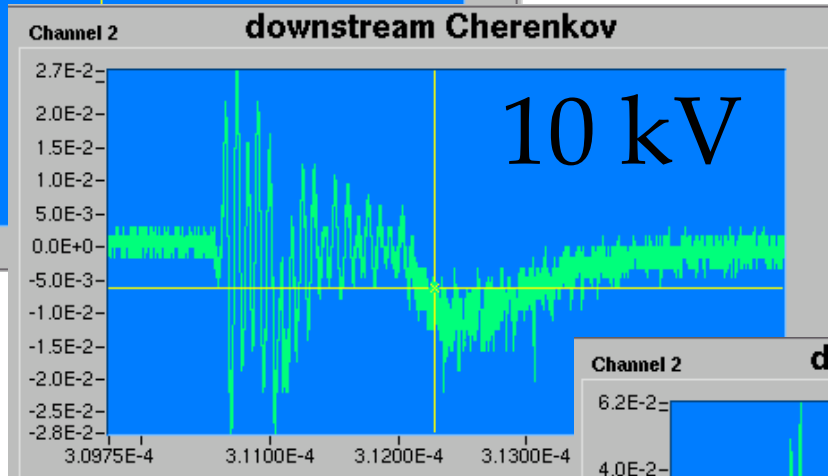
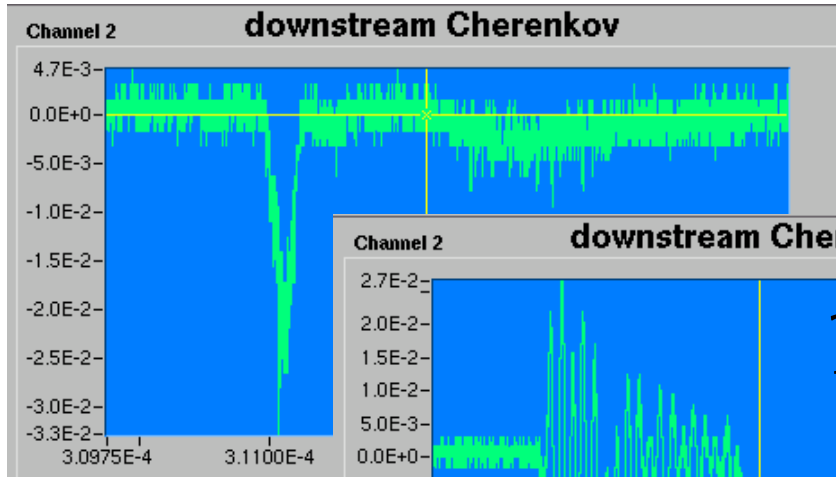
Field Ionizer



Why start with cold positrons?

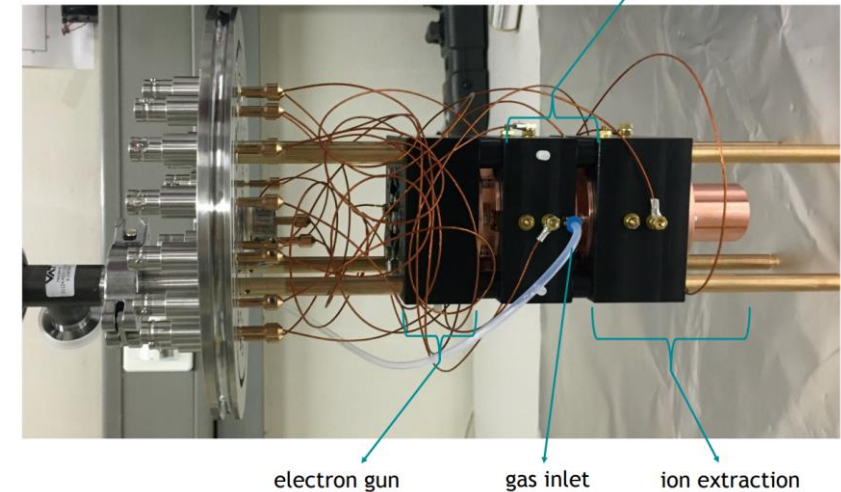
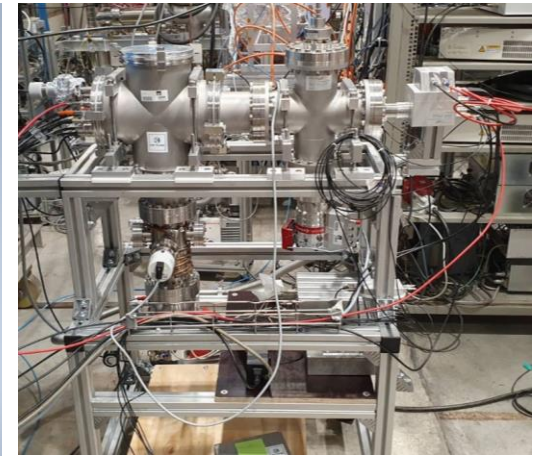
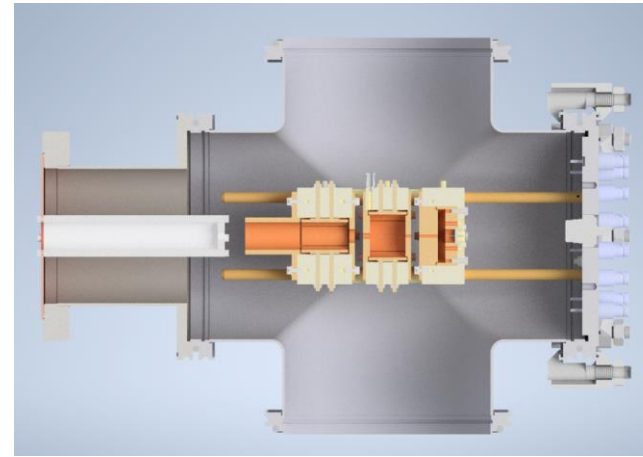


Drift Tube



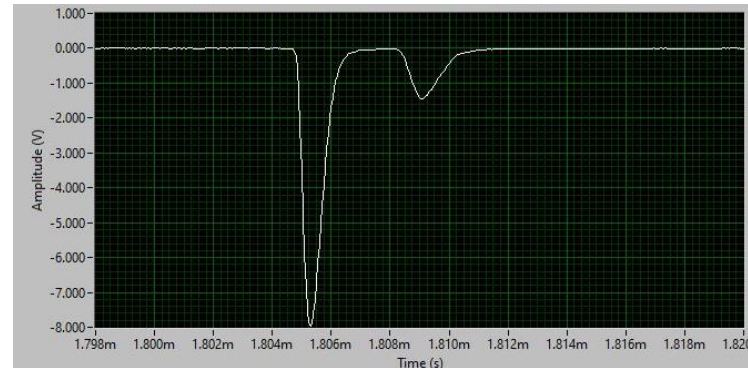
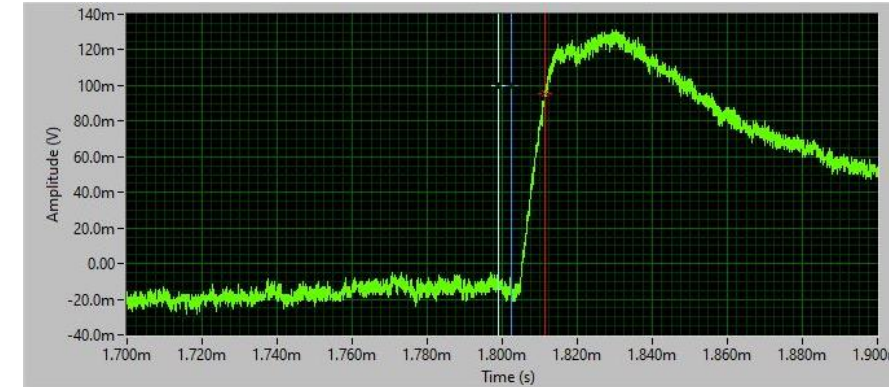
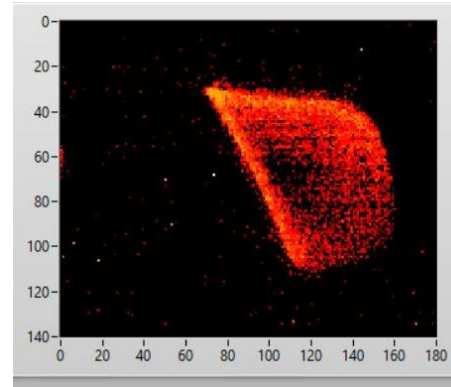
Proton Source

- A novel proton source was developed at SMI
- Protons were produced via electron impact ionization of H_2 gas
- This process typically creates more H_2^+ and H_3^+ ions than protons
- The protons were 'filtered' by trapping the ions in the gas cell while applying a RW drive to 4 split electrodes

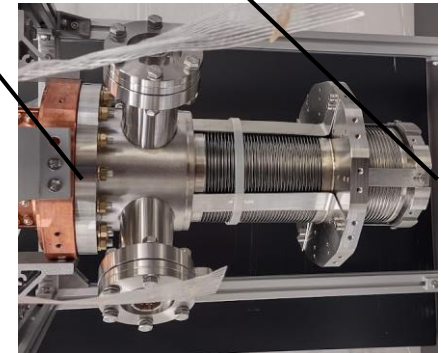
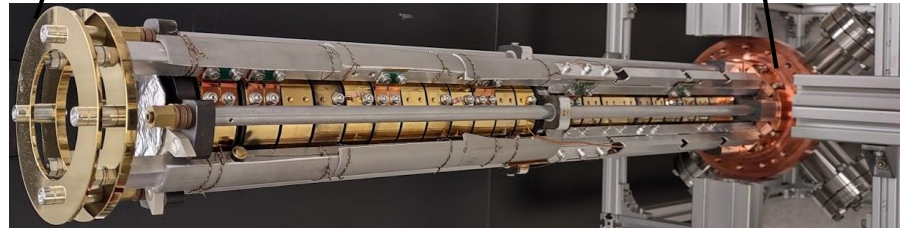
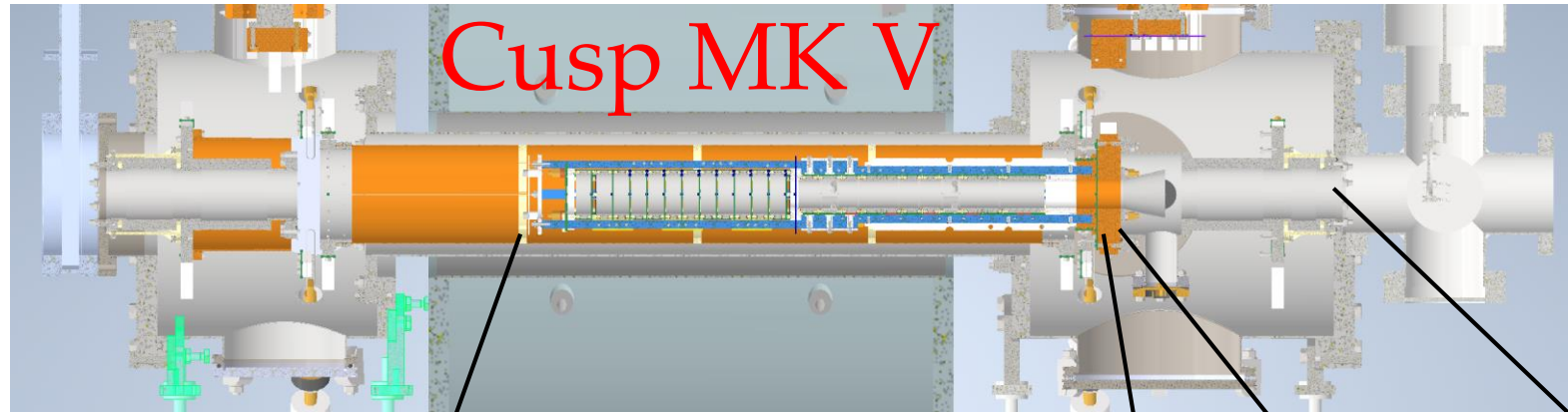


Proton Source

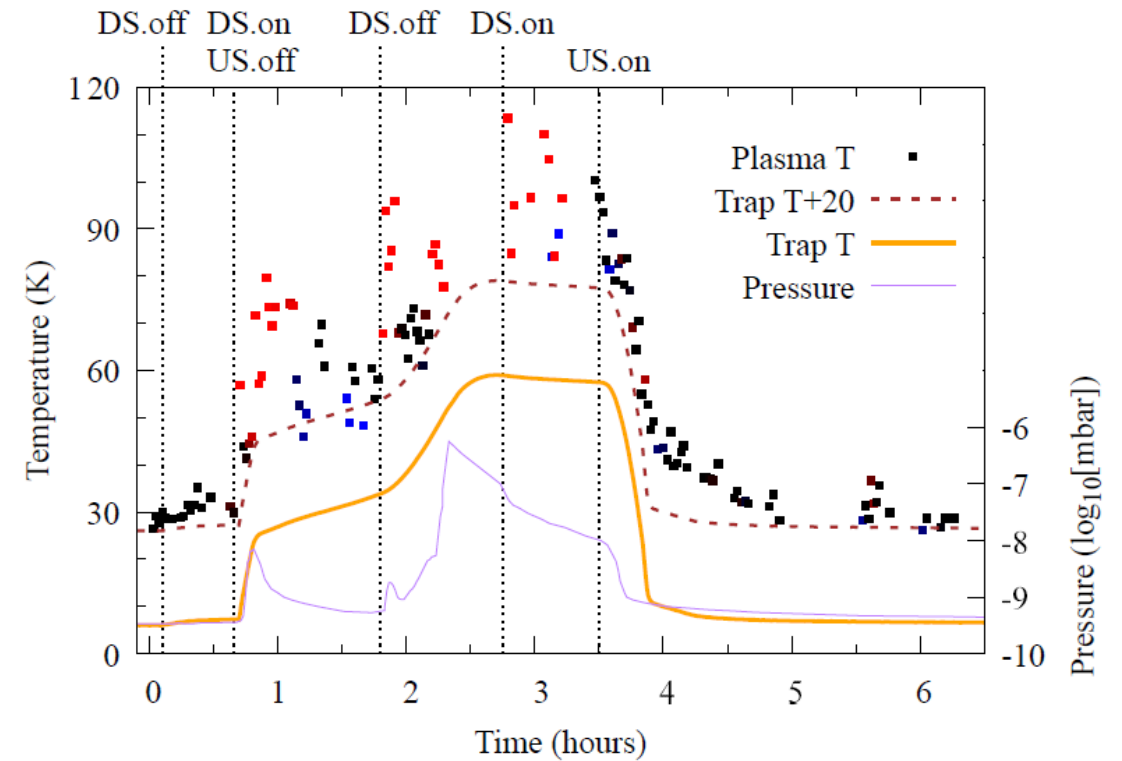
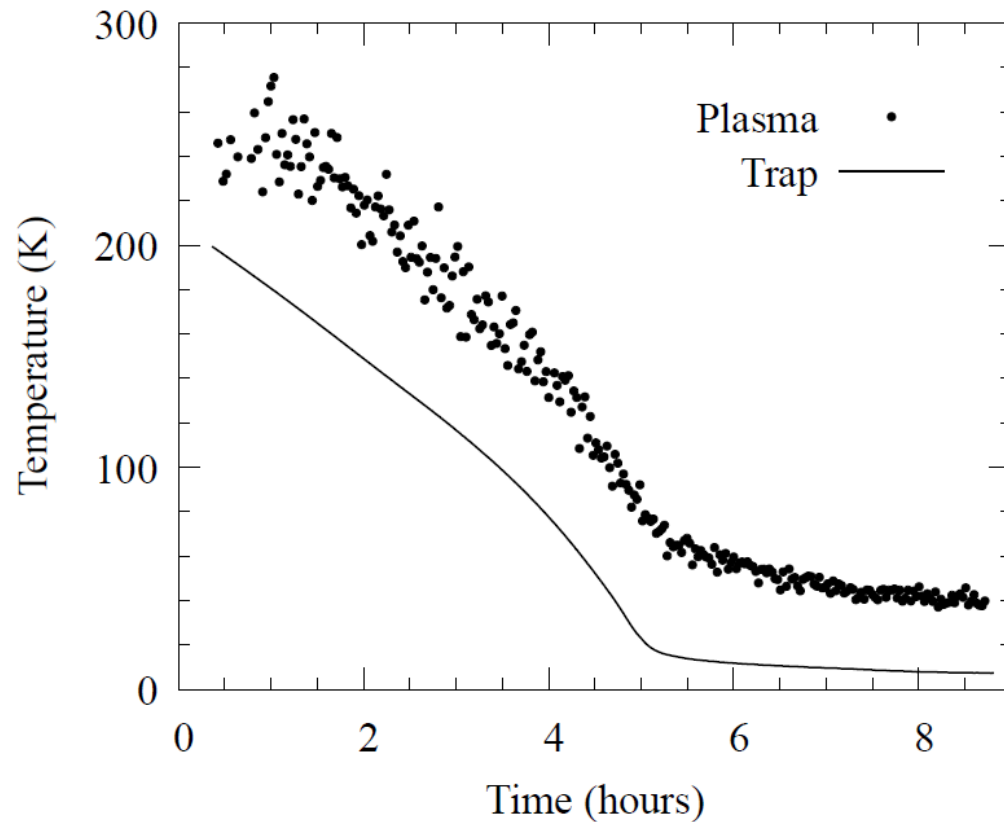
- The proton source was connected to the Cusp Mk IV trap
- We were able to accumulate ~1 million protons in the Cusp within a few seconds
- Unfortunately time ran out for work with protons so it was not possible to try to produce Ry-H



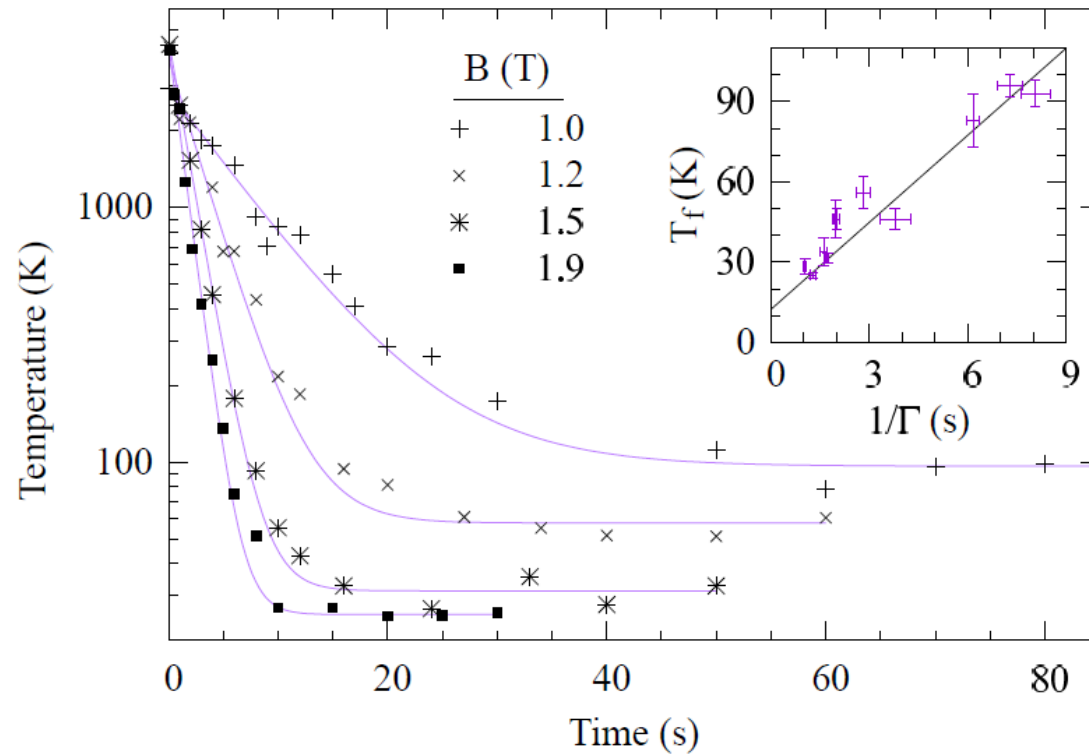
Double Cusp Trap



Pressure and Trap Temperature Plasma Temperature Dependence

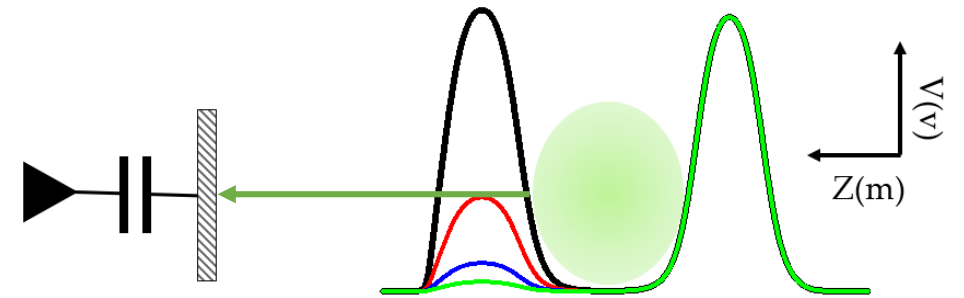


Cooling time Vs B



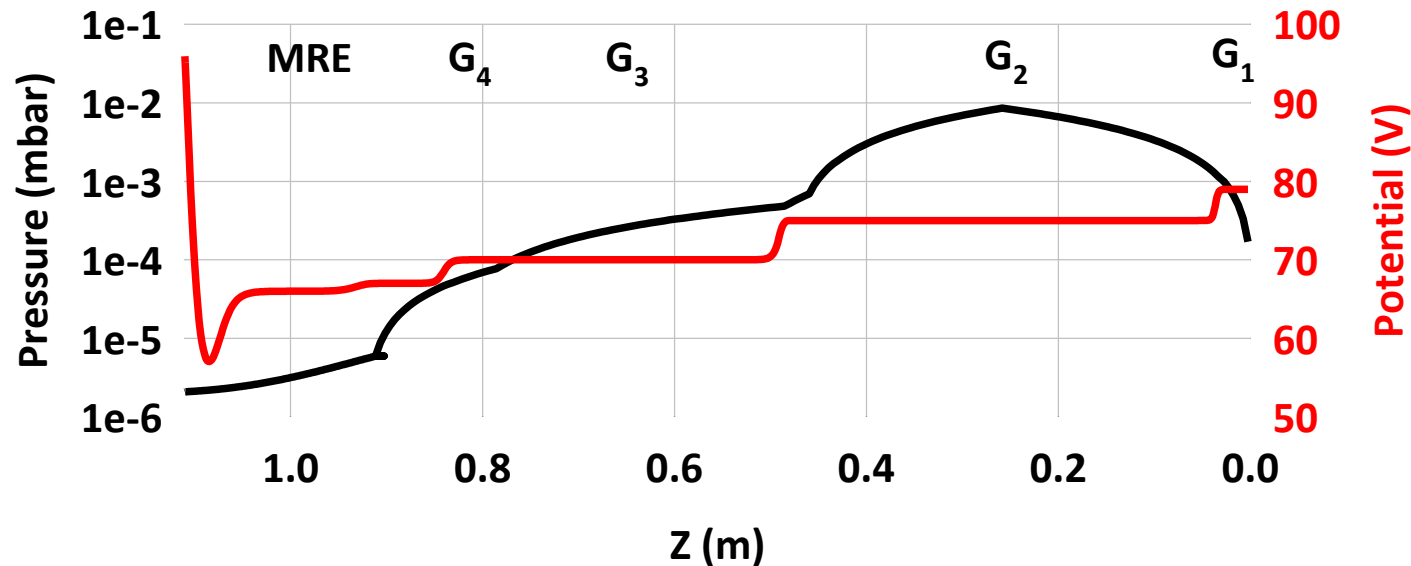
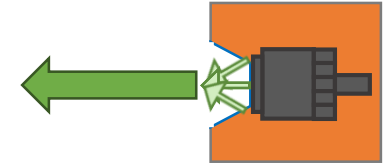
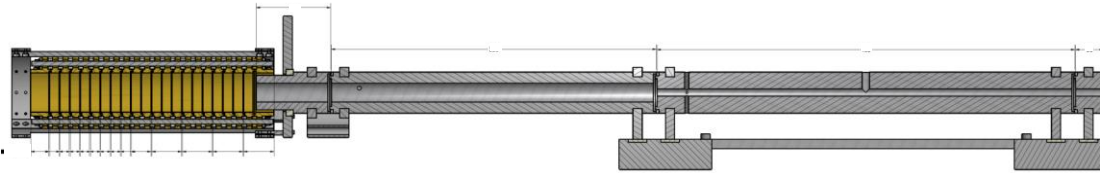
Plasma Temperature

- To measure the plasma temperature the potential barrier is slowly lowered while measuring the current delivered to an MCP (either with a charge amp or using the light from a phosphor screen and SiPM)



Combination $k_b T$ + space charge determines the dN/dV

Positron Trap



• Sodium-22

- Half-life: 2.6 Years
- Supplier: iThemba labs
- Strength: up to 50mCi (1.85GBq)

• Neon

- $5 \times 10^{-3} \leq \varepsilon_m \leq 2.5 \times 10^{-2}$
- Band Gap Insulator
- Require cryogenic equipment