



Imperial College
London



Science and
Technology
Facilities Council

Simplified Likelihoods (at least one approach)

Nicholas Wardle

[Publication of statistical models: hands-on workshop](#)

Simplified likelihoods

10/11/2021

Common choices for searches at the LHC

General form* for our experimental likelihood
(for measurements, searches ...) is

$$L(\boldsymbol{\alpha}, \boldsymbol{\delta})\pi(\boldsymbol{\delta}) = \prod_{I=1}^P \text{Pr}\left(n_I^{\text{obs}} \mid n_I(\boldsymbol{\alpha}, \boldsymbol{\delta})\right)\pi(\boldsymbol{\delta})$$

Where $\boldsymbol{\alpha}$ are the “parameters of interest” (mass of a new hypothetical particle, cross-section for some new process ...) and $\boldsymbol{\delta}$ are the “nuisance parameters”.

[1] G. Cowan, K. Cranmer, E. Gross, O. Vitells [Eur.Phys.J.C71:1554,2011](https://arxiv.org/abs/1008.0442)

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$$\boldsymbol{\alpha} = \mu$$

At the LHC, the profiled likelihood ratio test statistic is the most common choice [1] \rightarrow one parameter of interest μ – common multiplier for total signal yield

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Sum over the signals / background contributions

$$n_I(\mu, \boldsymbol{\delta}) \rightarrow \mu \cdot \sum_{\text{sigs}} n_{s_k, I} + \sum_{\text{bkgs}} n_{b_k, I}(\boldsymbol{\delta}) \rightarrow \mu \cdot n_{s, I} + n_{b, I}(\boldsymbol{\delta})$$

$$\text{Pr}(n|\lambda) = \frac{\lambda^n}{n!} e^{-\lambda}$$

Often use *binned* likelihood $\rightarrow \text{Pr}(\cdot)$ are Poisson probabilities

[1] G. Cowan, K. Cranmer, E. Gross, O. Vitells [Eur.Phys.J.C71:1554,2011](https://arxiv.org/abs/1007.1719)

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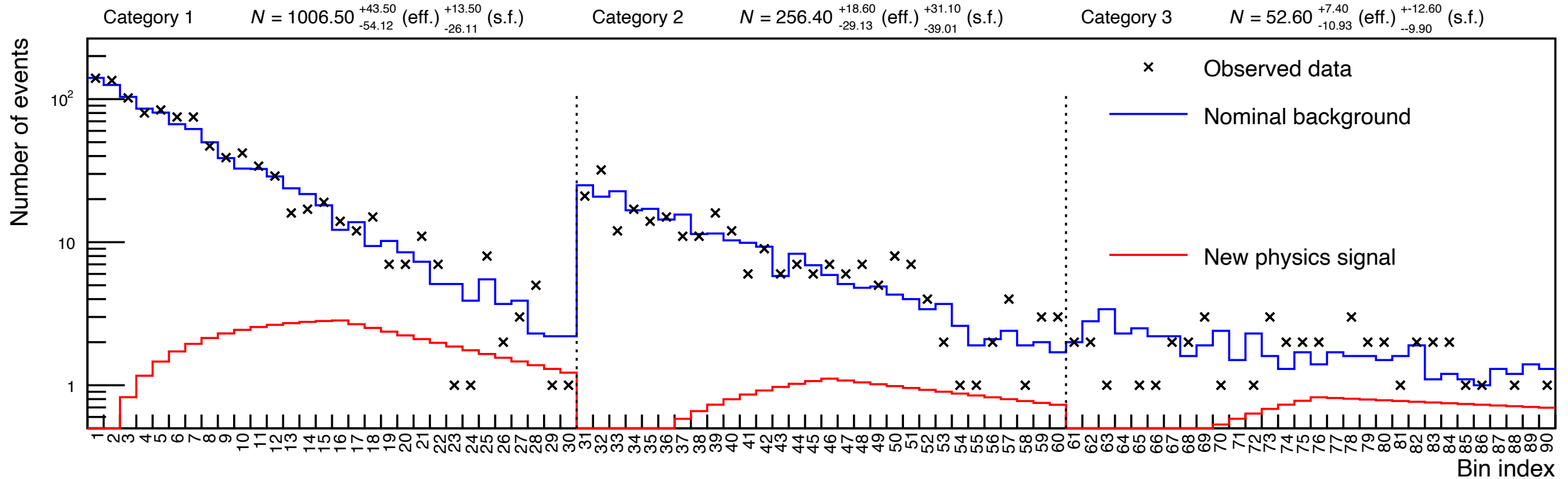
$$\pi(\boldsymbol{\delta})$$

Nuisance parameter priors and/or “in-situ” measurements of $\boldsymbol{\delta}$

[1] G. Cowan, K. Cranmer, E. Gross, O. Vitells [Eur.Phys.J.C71:1554,2011](https://arxiv.org/abs/1007.4730)

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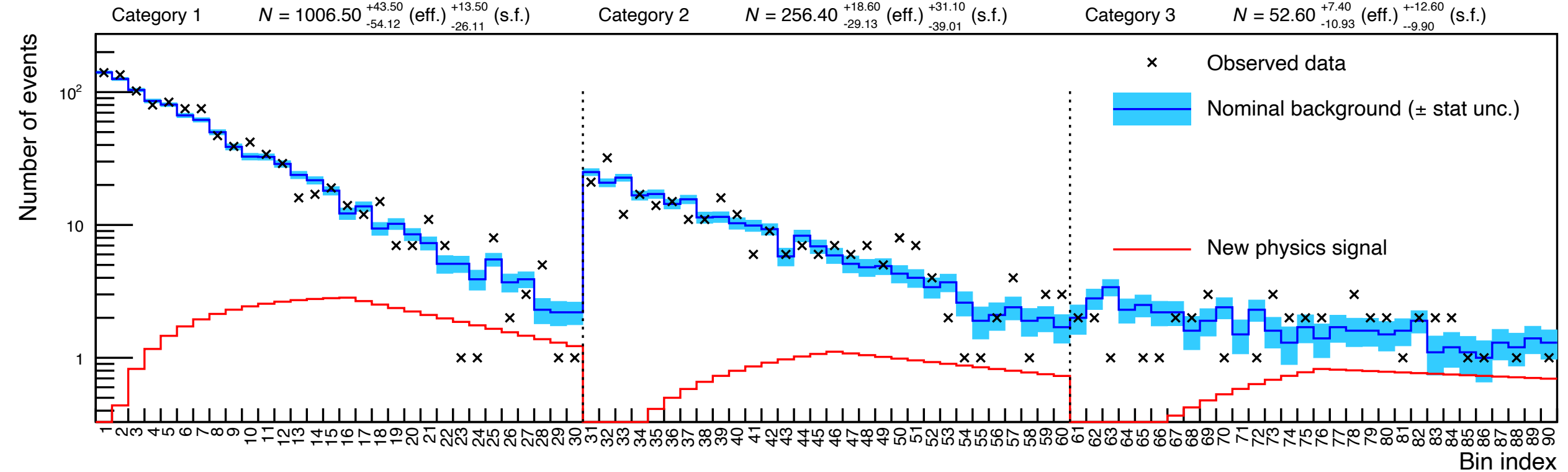
Toy search for new physics



Imagine a (rather simplified) model inspired by a typical search for some Supersymmetric particle or exotic signature.

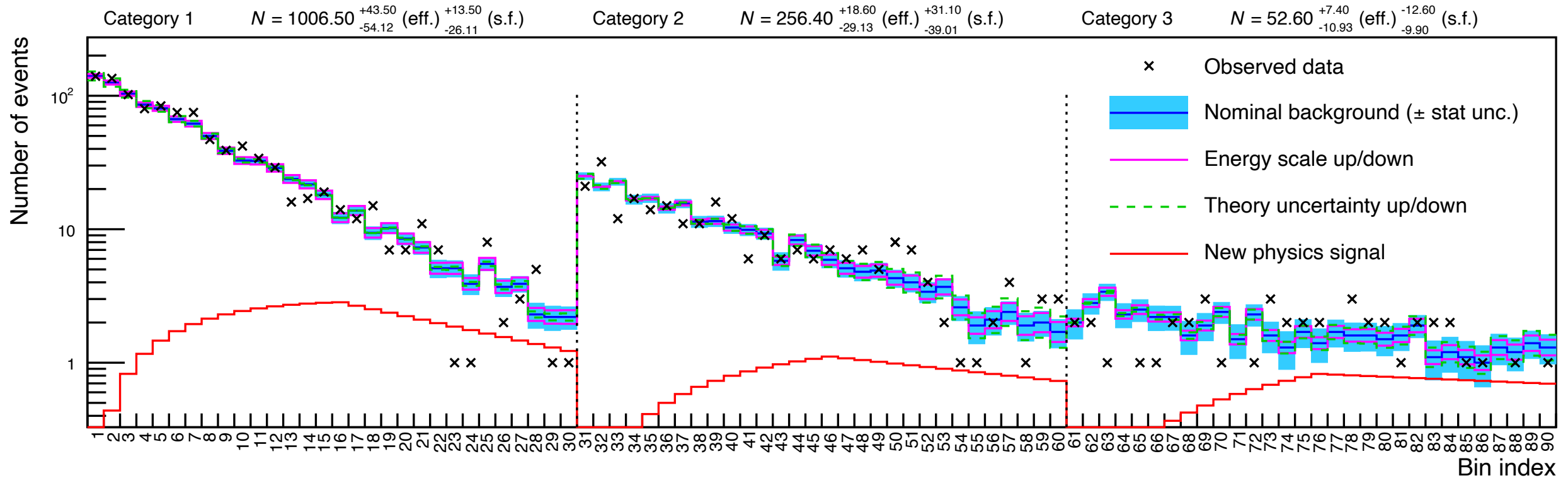
- There is a single source of background (can also think of this as the sum of all backgrounds)
- The data (observations) are divided into regions we have;
 - 3 categories for the data → each category has 30 bins
 - Increasing S/B with bin-number, within each category

Toy search for new physics



There are **two** uncertainties (labelled “efficiency” and “scale-factor”) on the background yields (N), and **each bin** has an uncertainty which is uncorrelated between bins (e.g this could be from limited Monte Carlo statistics used to estimate n_i)

Toy search for new physics



There are **two** uncertainties (labelled “efficiency” and “scale-factor”) on the background yields (N), and **each bin** has an uncertainty which is uncorrelated between bins (e.g this could be from limited Monte Carlo statistics used to estimate n_i)

Another **two** uncertainties correlated between bins (“energy scale” and “theory” uncertainty)

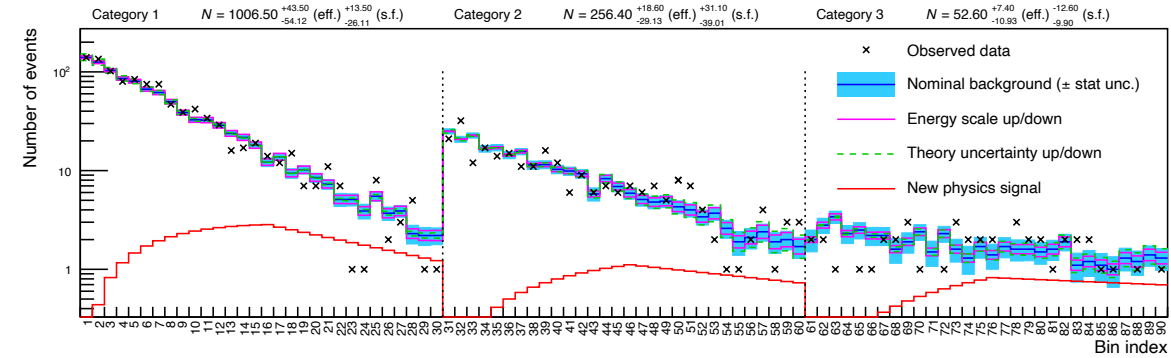
In total this means 94 nuisance parameters

Toy search for new physics

Think of the expected number of background events in a given bin I , as the fraction of events in that bin (f_I) multiplied by the total number of events (N)

δ are nuisance parameters representing **independent** sources of uncertainty (in our case 94 of them)

$$n_I(\delta) \equiv f_I(\delta)N(\delta)$$



$$N(\delta) = N^0 \cdot \prod_j (1 + K_j)^{\delta_j}$$

Uncertainties in the normalisation (N) typically follow log-normals

$$\frac{n_I(\delta)}{n_I^0} = \prod_j (1 + \epsilon_{Ij})^{\delta_j}$$

Similarly for un-correlated bin-by-bin uncertainties

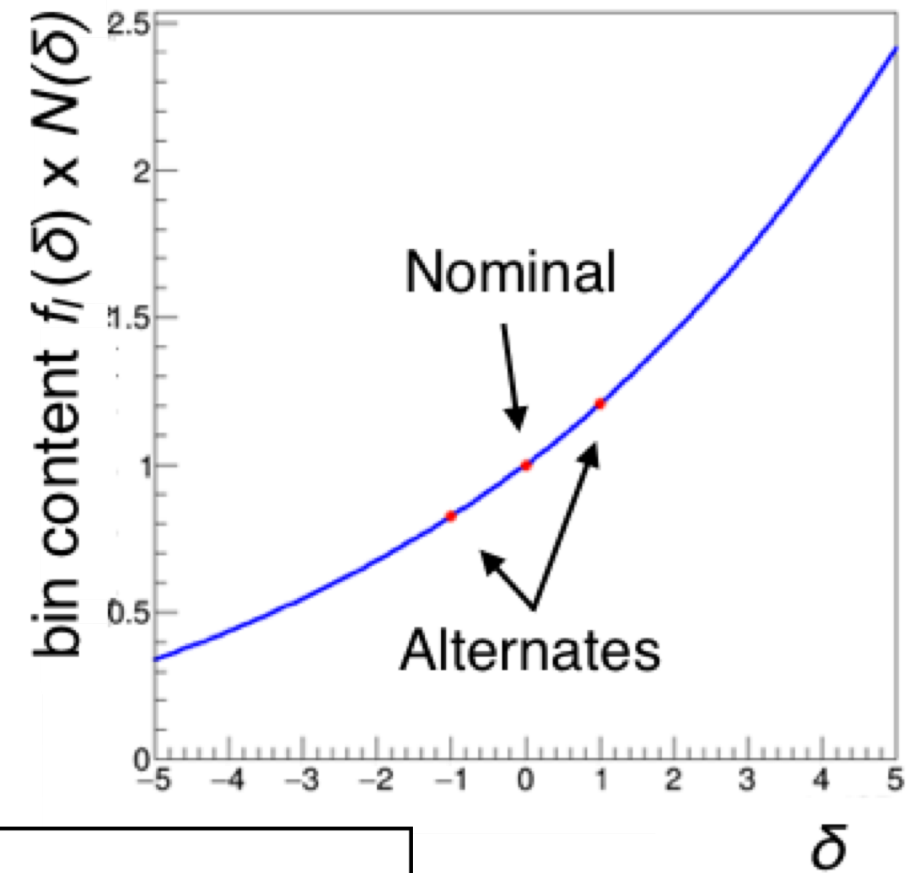
K_j and ϵ_{Ij} represent the relative size and direction of the uncertainty

Toy search for new physics

The effects of correlated systematic uncertainties on n_i are modelled using quadratic(linear) **interpo**(**extrapo**)lation function

$$f_I(\boldsymbol{\delta}) = f_I^0 \cdot \frac{1}{F(\boldsymbol{\delta})} \prod_j p_{Ij}(\delta_j)$$

$$F(\boldsymbol{\delta}) = \sum_I f_I(\boldsymbol{\delta})$$



$$p_{Ij}(\delta_j) = \begin{cases} \frac{1}{2}\delta_j(\delta_j - 1)\kappa_{Ij}^- - (\delta_j - 1)(\delta_j + 1) + \frac{1}{2}\delta_j(\delta_j + 1)\kappa_{Ij}^+ & \text{for } |\delta_j| < 1 \\ \left[\frac{1}{2}(3\kappa_{Ij}^+ + \kappa_{Ij}^-) - 2 \right] \delta_j - \frac{1}{2}(\kappa_{Ij}^+ + \kappa_{Ij}^-) + 2 & \text{for } \delta_j > 1 \\ \left[2 - \frac{1}{2}(3\kappa_{Ij}^- + \kappa_{Ij}^+) \right] \delta_j - \frac{1}{2}(\kappa_{Ij}^+ + \kappa_{Ij}^-) + 2 & \text{for } \delta_j < -1 \end{cases}$$

Experimental likelihood

Now we can write the likelihood for this search as follows;

$$L(\mu, \boldsymbol{\delta})\pi(\boldsymbol{\delta}) = \prod_{I=1}^{90} P(n_I^{\text{obs}} | \mu \cdot n_{s,I} + n_{b,I}(\boldsymbol{\delta})) \cdot \prod_{j=1}^{94} e^{-\delta_j^2}$$

$$n_{b,I}(\boldsymbol{\delta}) = N_c^0 \cdot \prod_{k=1}^2 (1 + K_k)^{\delta_k} \cdot f_I^0 \cdot \frac{1}{F(\boldsymbol{\delta})} \prod_{j=3}^4 p_{I,j}(\delta_j) \cdot (1 + \epsilon_I)^{\delta_I}$$

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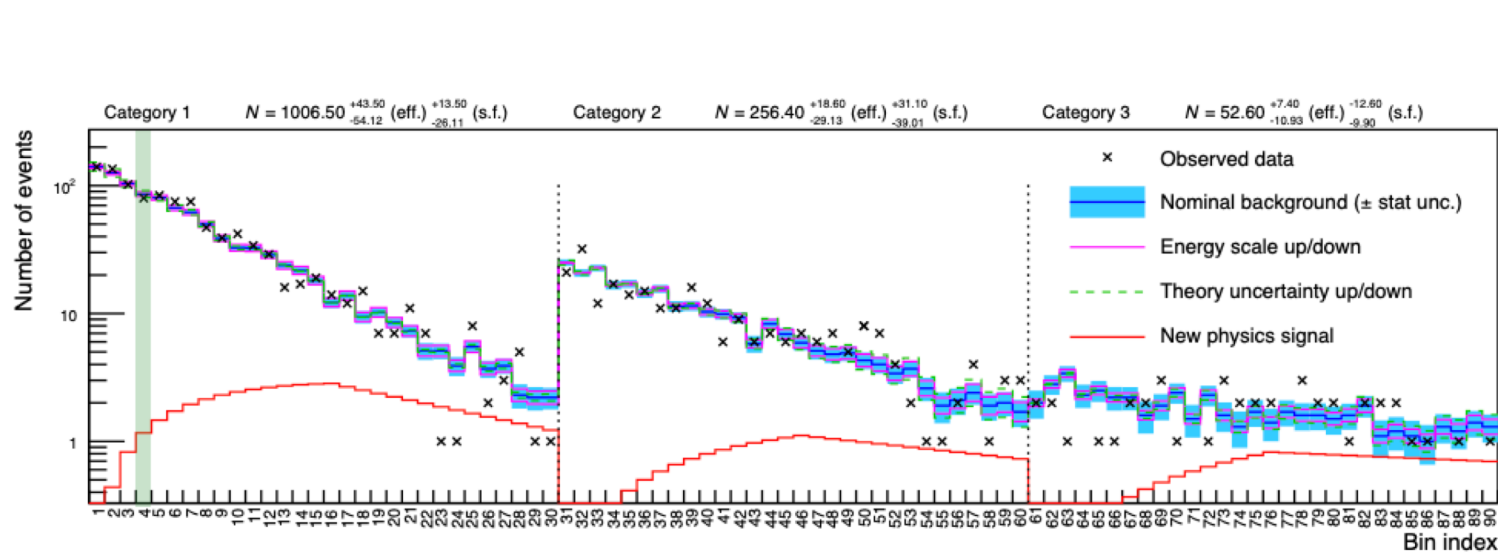
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Specifying these terms with this generic form means the full likelihood can be communicated as plain text!

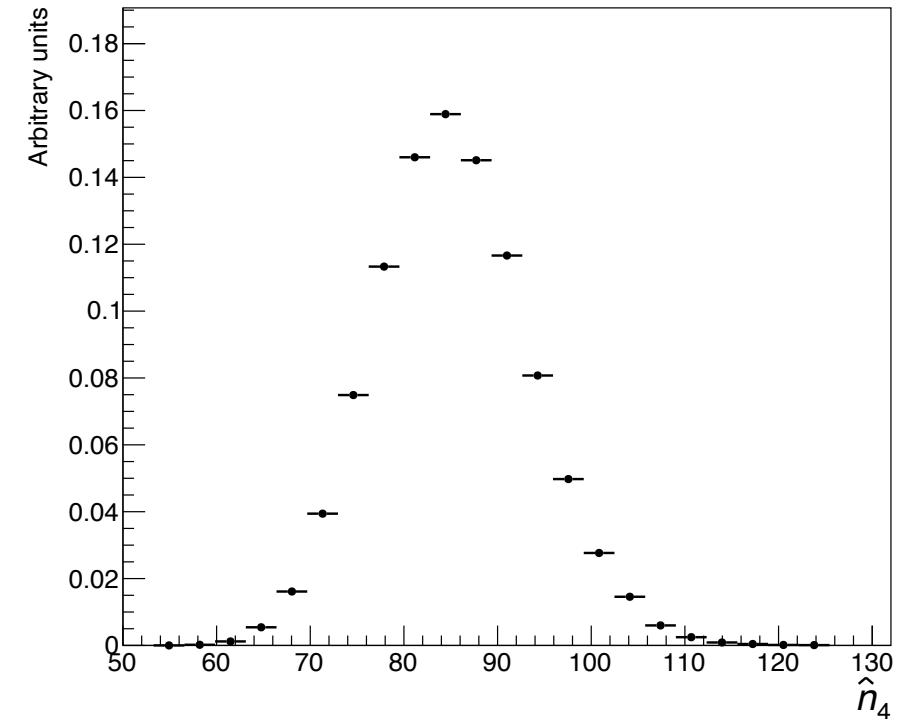
A lot of physicists' time working on an LHC search is spent on **these!**

Re-parameterize the backgrounds

We can generate pseudo-experiments for $n_{b,I}$ since we know $p(\boldsymbol{\delta}) := \pi(\boldsymbol{\delta}) \sim e^{-\frac{1}{2}\boldsymbol{\delta}\cdot\boldsymbol{\delta}}$
Use randomly sampled $\boldsymbol{\delta}'$ and $\hat{n}_I = n_{b,I}(\boldsymbol{\delta}')$ to determine the distribution of the backgrounds...

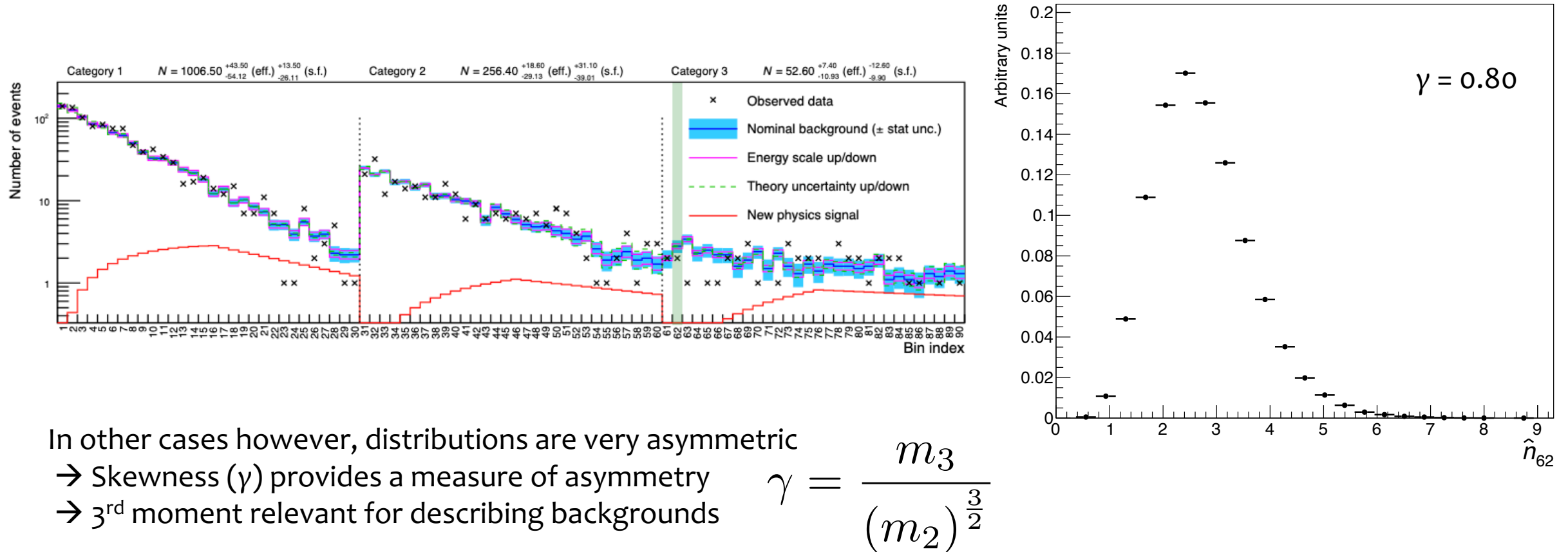


In some bins, distributions looks symmetric and Gaussian
→ can be described by 2 moments (mean and variance)



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In other cases however, distributions are very asymmetric

→ Skewness (γ) provides a measure of asymmetry

→ 3rd moment relevant for describing backgrounds

$$\gamma = \frac{m_3}{(m_2)^{\frac{3}{2}}}$$

Simplifying the likelihood?

For statistical (re-) interpretation purposes we eliminate nuisance parameters (δ)

→ We are mainly interested in profiled / marginalized likelihoods $L(\mu, \delta) \rightarrow L(\mu)$

Since the “backgrounds” are only dependent on the nuisance parameters, we can approximate in such a way that the profiled (or marginal) likelihood is preserved as follows [1];

1. Express $n_{b,l}$ as a simple expansion (quadratic) in terms of **combined nuisance parameters** ϑ_l

$$n_{b,I} \simeq a_I + b_I \theta_I + c_I \theta_I^2 \quad I = 1 \dots 90$$

[1] A. Buckley, M. Citron, S. Fichet, S. Kraml, W. Waltenberger, **NW** [J. High Energ. Phys. 2019, 64 \(2019\)](#)

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2. Re-parameterize likelihood in terms of μ^* and $\vartheta_l \rightarrow$ Need to derive $\pi(\vartheta)$!

$$L(\mu, \delta) \pi(\delta) \rightarrow L(\mu, \theta) \pi(\theta) = \prod_{I=1}^{P=90} P(n_I^{\text{obs}} | \mu \cdot n_{s,I} + a_I + b_I \theta_I + c_I \theta_I^2) \cdot \frac{1}{\sqrt{(2\pi)^P}} e^{-\frac{1}{2} \theta^T \rho^{-1} \theta}$$

$P(x|y)$ = Poisson probability as before

These are the same as the *full* likelihood

$$\rho_{I,J} = \rho_{J,I}$$

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Nearly done with the formulae...

Coefficients obtained by matching moments and appealing to CLT at NLO.

Coefficients a , b and c are determined from the first 3 central moments of the joint distributions of $n_{b,I}$ - Mean, covariance **and skew**

Solutions valid for $\frac{8(m_{2,II})^3}{(m_{3,I})^2} \geq 1$

$$c_I = -\text{sign}(m_{3,I}) \sqrt{2m_{2,II}} \cos \left(\frac{4\pi}{3} + \frac{1}{3} \arctan \left(\sqrt{8 \frac{m_{2,II}^3}{m_{3,I}^2} - 1} \right) \right)$$

$$b_I = \sqrt{m_{2,II} - 2c_I^2},$$

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$$\rho_{IJ} = \frac{1}{4c_I c_J} \left(\sqrt{(b_I b_J)^2 + 8c_I c_J m_{2,IJ}} - b_I b_J \right).$$

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Moments can be calculated analytically or (my preference) using pseudo experiments

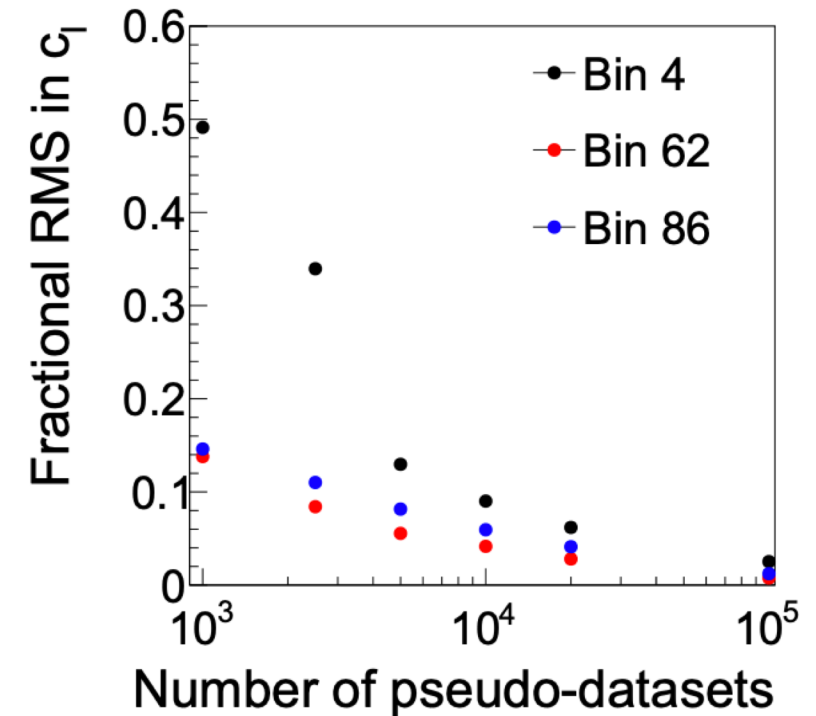
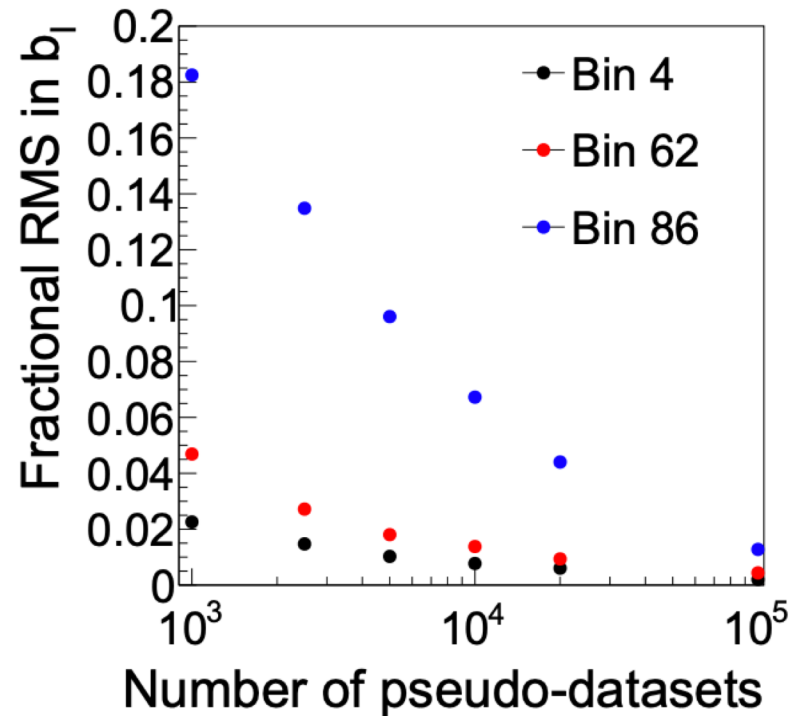
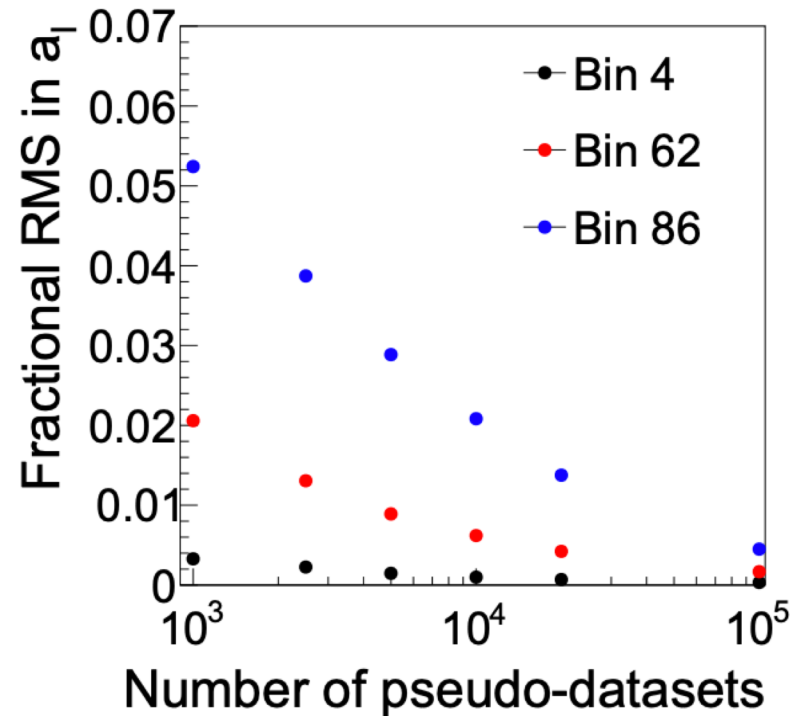
$$m_{1,I} = \mathbf{E}[\hat{n}_I]$$

$$m_{2,IJ} = \mathbf{E}[(\hat{n}_I - \mathbf{E}[\hat{n}_I])(\hat{n}_J - \mathbf{E}[\hat{n}_J])]$$

$$m_{3,I} = \mathbf{E}[(\hat{n}_I - \mathbf{E}[\hat{n}_I])^3]$$

These quantities are the inputs needed to determine the **simplified likelihood**

Convergence of moment calculation with pseudo-data



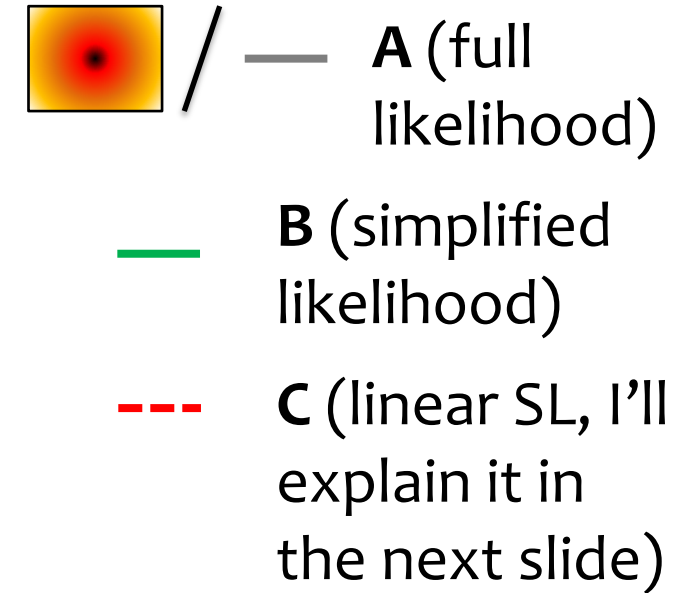
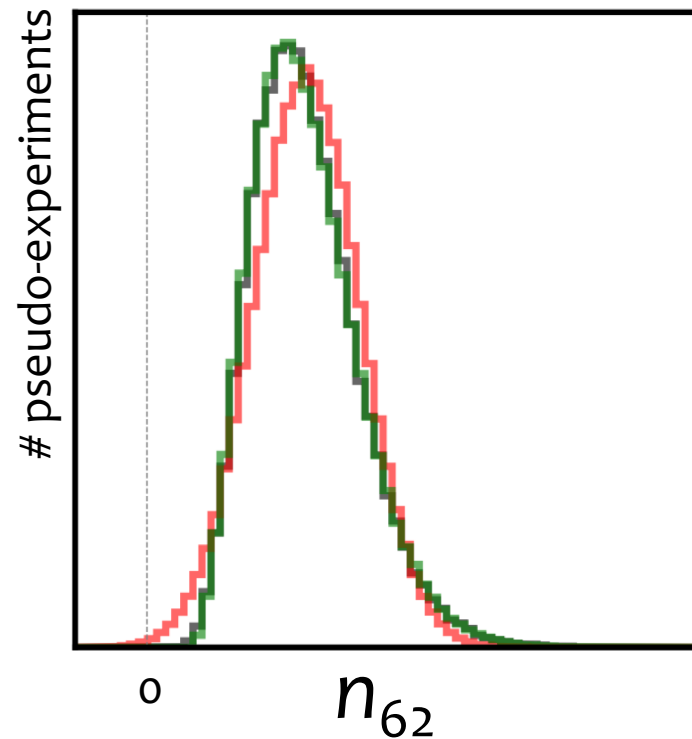
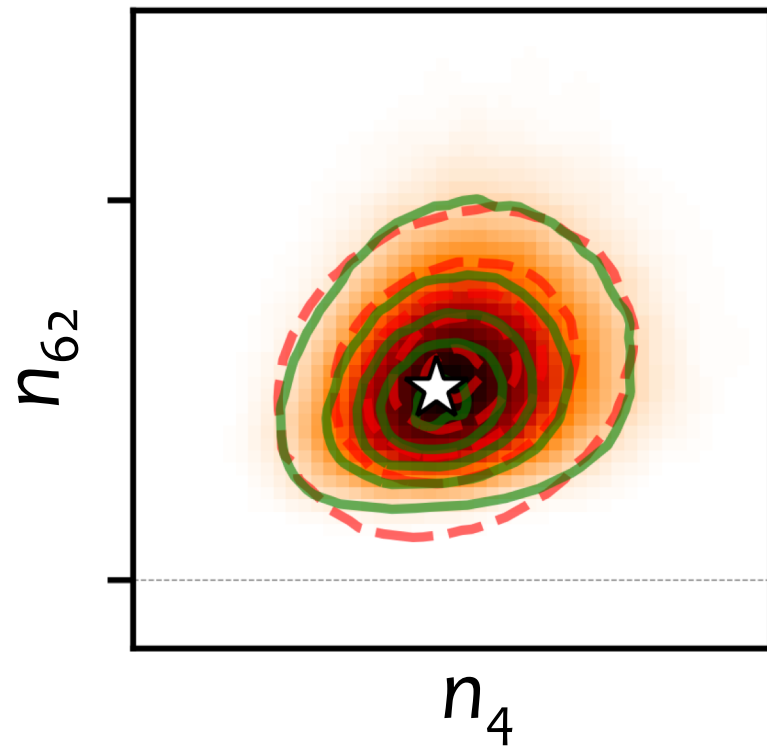
3rd Moment typically requires most toys to get accurate value, however this is mostly true when m_3 is small and therefore not so relevant!

How well does this approximate the distribution of n_I ?

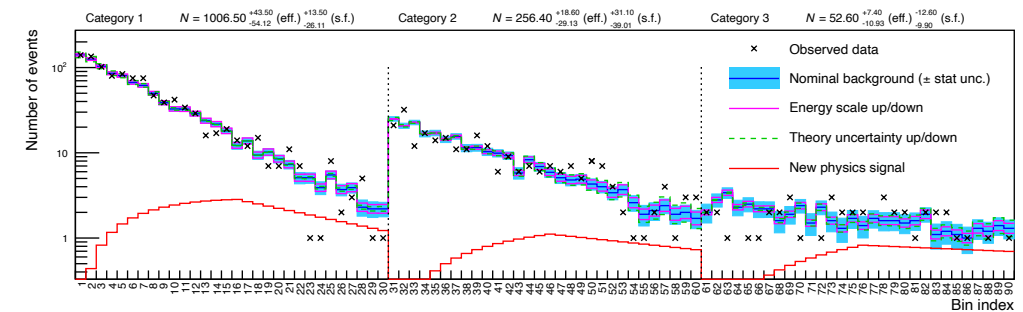
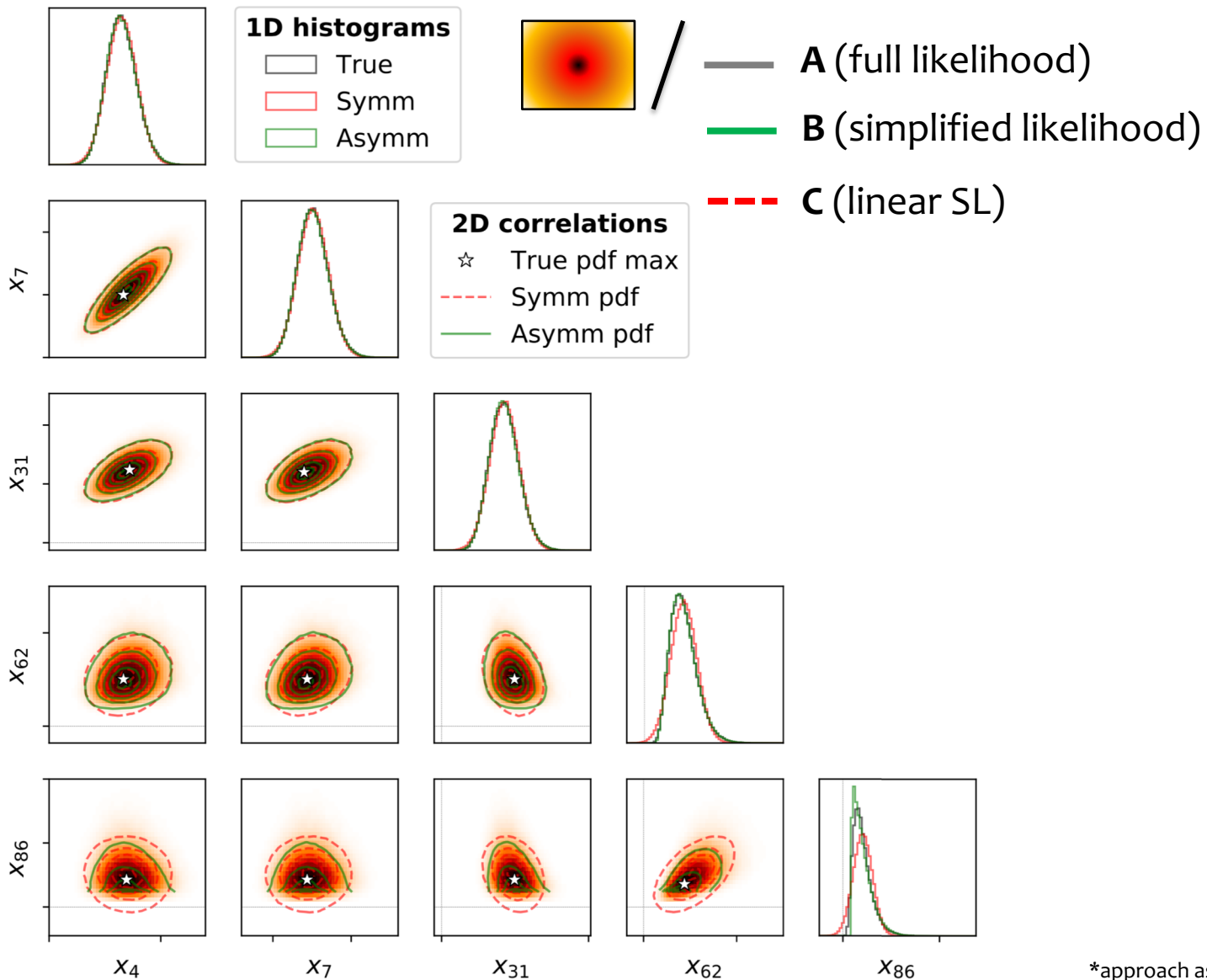
We can compare the distribution of \hat{n}_I obtained in the pseudo-data from

A. $\hat{n}_I = n_{b,I}(\boldsymbol{\delta}')$ generating from $p(\boldsymbol{\delta}) := \pi(\boldsymbol{\delta}) \sim e^{-\frac{1}{2}\boldsymbol{\delta} \cdot \boldsymbol{\delta}}$

B. $\hat{n}_I = n_{b,I}(\boldsymbol{\theta}'_I)$ generating from $p(\boldsymbol{\theta}) \sim e^{-\frac{1}{2}\boldsymbol{\theta}^T \boldsymbol{\rho}^{-1} \boldsymbol{\theta}}$

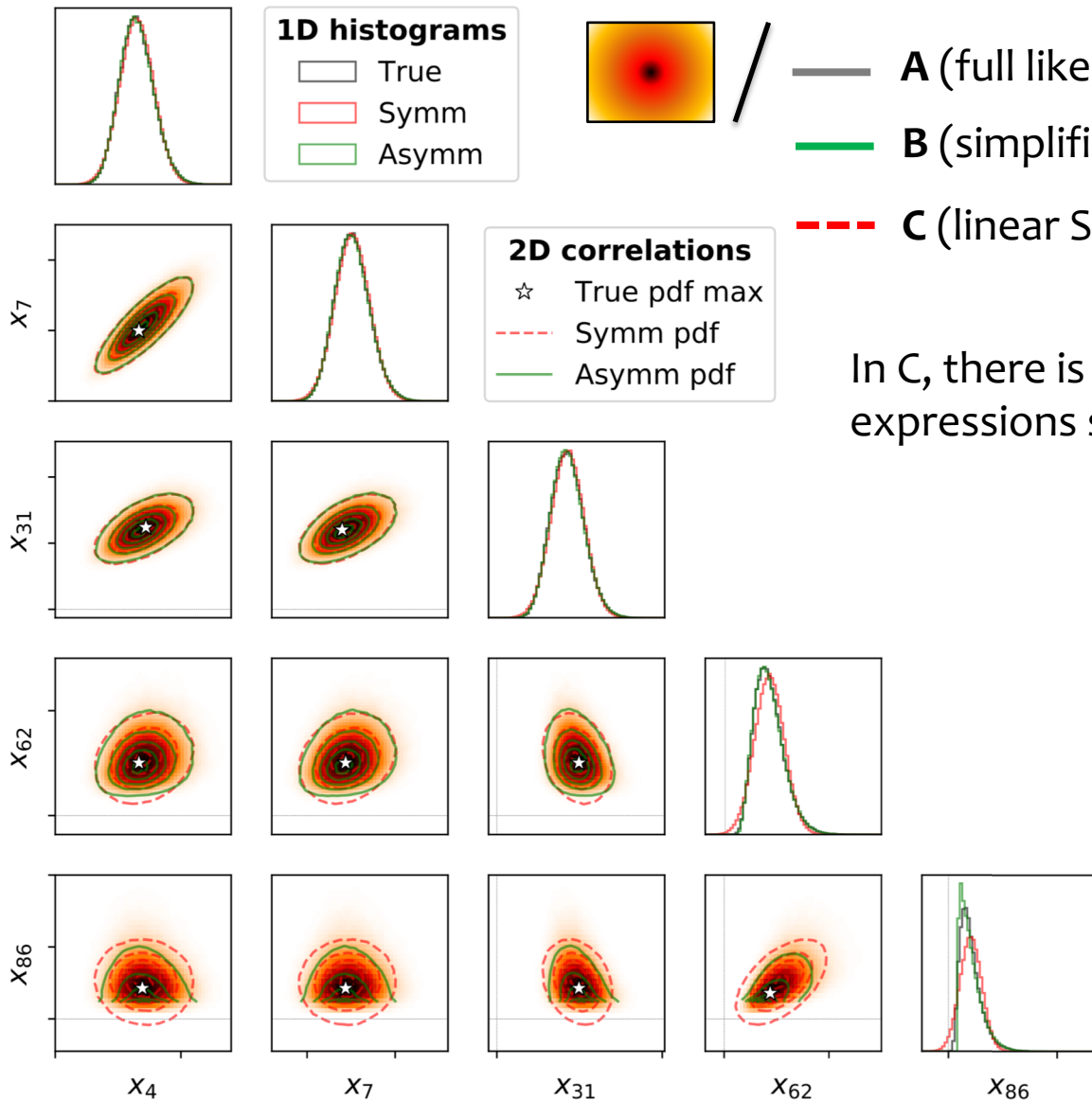


How well does this approximate the distribution of n_l ?



*approach as in [CMS-NOTE-2017-001](#), and K. Cranmer, S. Kreiss, D. López-Val, T. Plehn, [PhysRevD 91 054032](#)

How well does this approximate the distribution of n_I ?



— **A** (full likelihood)
 — **B** (simplified likelihood)
 - - - **C** (linear SL)

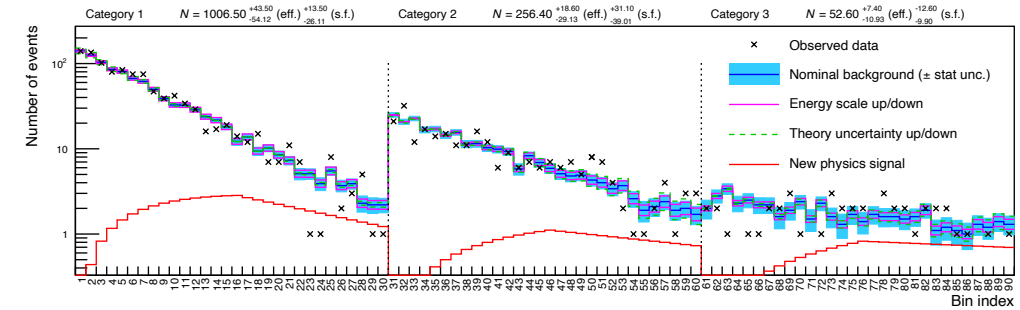
In C, there is a further simplification that $m_{3,I}$ is 0. In this case, the expressions simplify to*

$$n_{b,I}(\theta_I) = A_I + B_I \theta_I$$

$$p(\boldsymbol{\theta}) \sim e^{-\frac{1}{2} \boldsymbol{\theta}^T \mathbf{v}^{-1} \boldsymbol{\theta}}$$

$$A_I = m_{1,I}, \quad B_I = m_{2,II}, \quad v_{IJ} = m_{2,IJ}$$

When $m_{3,I}/(m_{2,II})^{\frac{3}{2}}$ (the skew) is small, the linear approximation is fairly good, as expected.



*approach as in [CMS-NOTE-2017-001](#), and K. Cranmer, S. Kreiss, D. López-Val, T. Plehn, [PhysRevD 91 054032](#)

Get to the punchline already Nick ...

Eliminating nuisance parameters (δ or θ) indicates how *accurately* we can reproduce statistical interpretations.

e.g. the profiled likelihood ratio test-statistic* is used to set limits on new physics processes at the LHC

$$t_\mu = -2 \ln \frac{L_S^{\max}(\mu)}{L_S^{\max}}$$

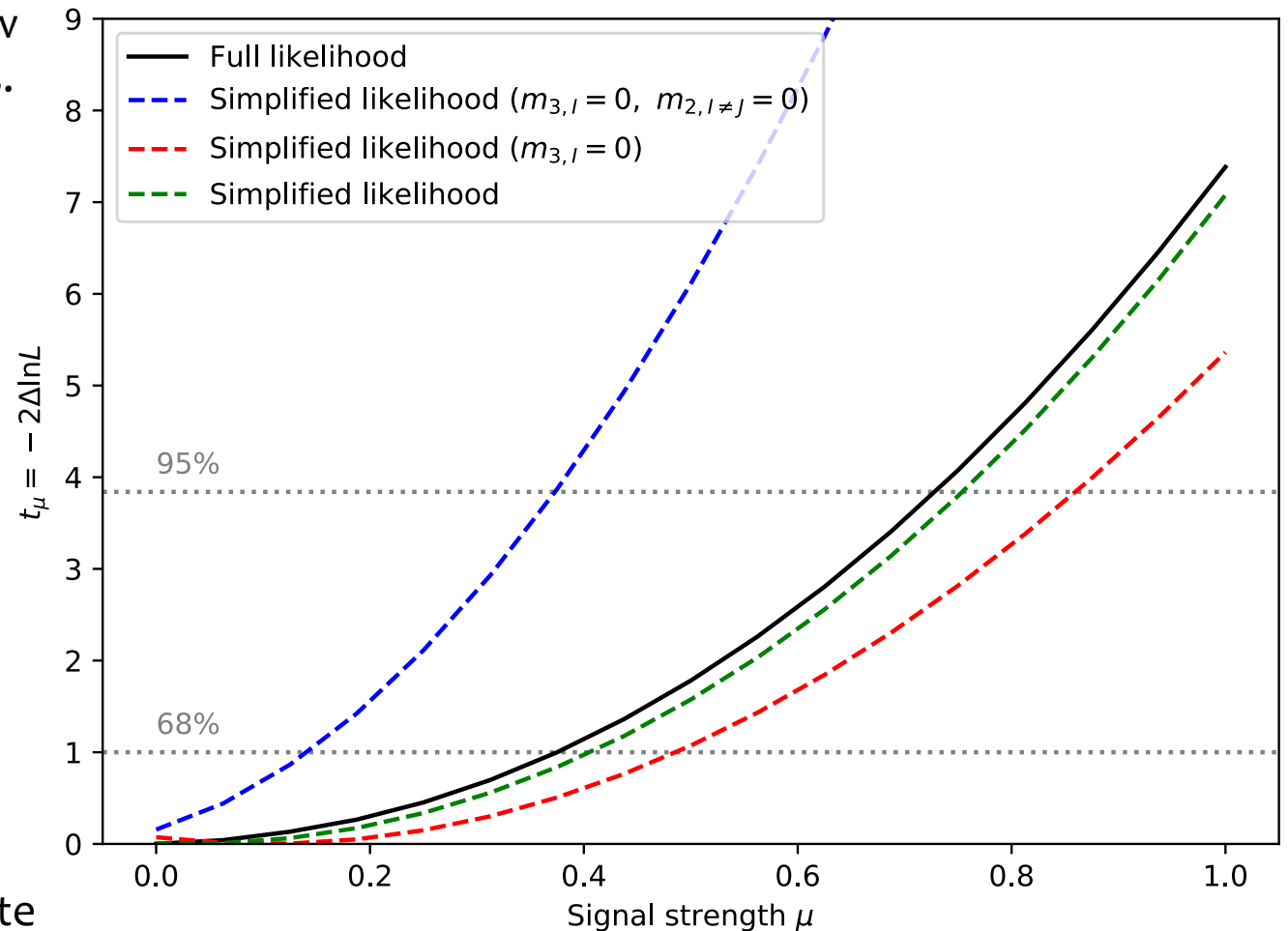
$$L_S^{\max}(\mu) = \max_{\theta_I} \{L_S(\mu, \theta)\}$$



Inputs for toy search
uploaded to [HepData](#)



Public scipy-based code to calculate
SL coefficients and run statistical tests
on [GitLab](#)



*No reason why we couldn't have marginalised the likelihood to compare Bayesian posterior distributions instead of profiling.

Discussion

Can we implement this in phHF simplification routines ?

Some things to mull over

- One only needs to calculate moments in different signal region bins : use MC (as we do in CMS) or propagate directly and use logL derivatives?

- Signal region vs control regions : For simplification, assume only interested in signal region (control data summarized also in covariance/skews)

- If using CRs and not including in procedure, ideally use post-fit estimates for generating the toys (include CRs in fit but not SRs to avoid double counting!)

Backup slides

Simplified likelihood log-likelihood

$$\ln(L_S(\mu, \boldsymbol{\theta})\pi(\boldsymbol{\theta})) = \sum_I^P \left[n_I^{\text{obs}} \ln(\mu n_{s,I} + n_{b,I}(\boldsymbol{\theta})) - (\mu n_{s,I} + n_{b,I}(\boldsymbol{\theta})) - n_I^{\text{obs}}! \right] - \frac{1}{2} \boldsymbol{\theta}^T \boldsymbol{\rho}^{-1} \boldsymbol{\theta} - \frac{P}{2} \ln 2\pi \quad (\text{B.1})$$

$$\frac{\partial \ln L_S}{\partial \mu} = \sum_I^P \left(\frac{n_I^{\text{obs}}}{\mu n_{s,I} + n_{b,I}(\boldsymbol{\theta})} - 1 \right) \cdot n_{s,I} \quad (\text{B.2})$$

$$\frac{\partial \ln L_S}{\partial \theta_A} = \left(\frac{n_A^{\text{obs}}}{\mu n_{s,A} + n_{b,A}(\boldsymbol{\theta})} - 1 \right) \cdot (b_A + 2c_A \theta_A) - \sum_I^P \rho_{AI}^{-1} \theta_I, \quad (\text{B.3})$$

Analytic simplified likelihood coefficients

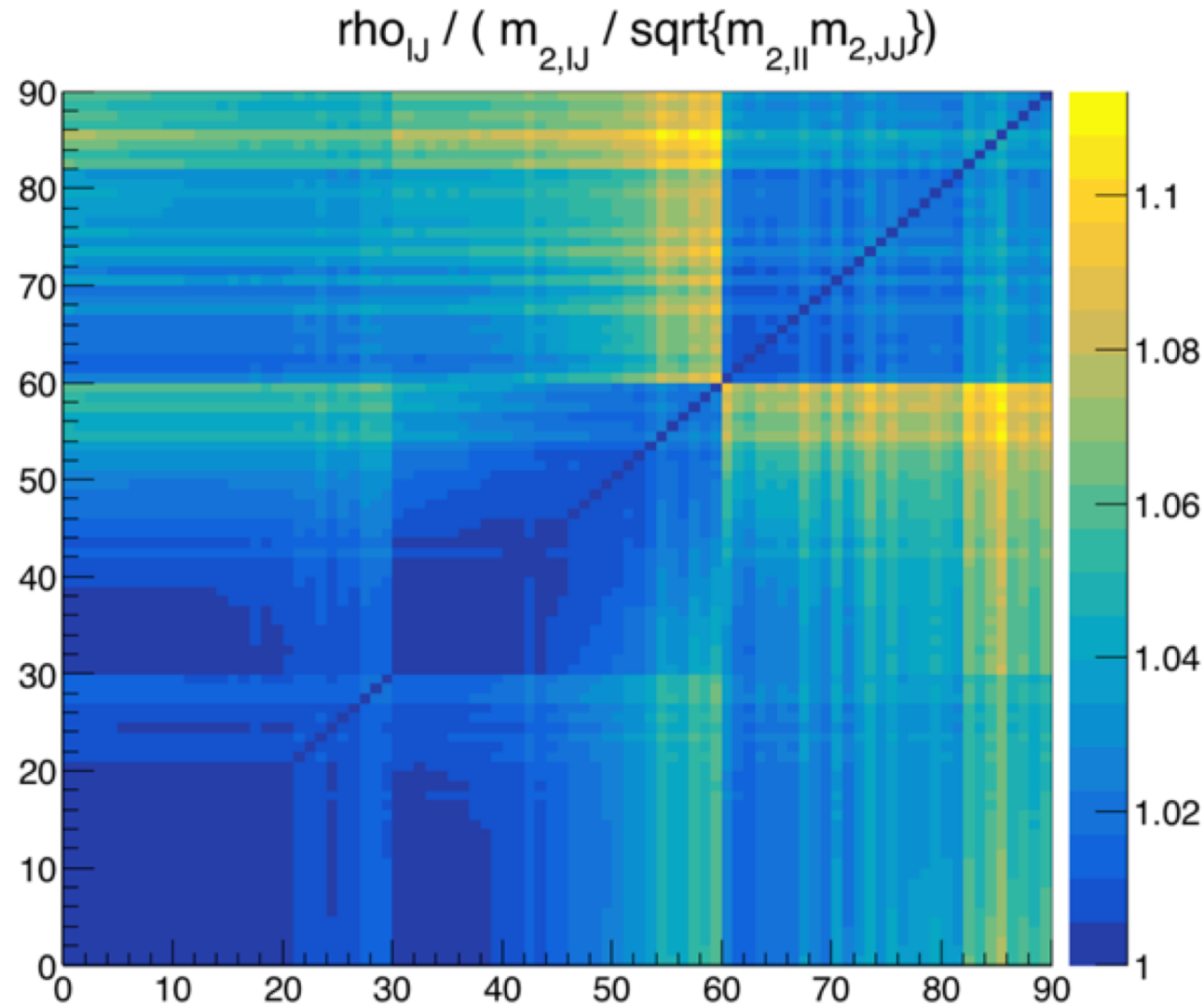
$$a_I = n_I^0 \left(1 + \text{tr } \Delta_{2,I} - \frac{1}{6} \sum_{i=1}^N \gamma_i (\Delta_{1,I,i})^3 + O(\Delta^4) \right),$$

$$b_I = a_I \left(\Delta_{1,I}^T \cdot \Delta_{1,I} + 2 \sum_{i=1}^N \gamma_i \Delta_{1,I,i} \Delta_{2,I,i} + O(\Delta^4) \right)^{1/2},$$

$$\rho_{IJ} = \frac{a_I a_J}{b_I b_J} \left(\Delta_{1,I}^T \cdot \Delta_{1,J} + \sum_{i=1}^N \gamma_i (\Delta_{1,I,i} \Delta_{2,J,i} + \Delta_{1,J,i} \Delta_{2,I,i}) \right) + O(\Delta^4),$$

$$c_I = \frac{a_I}{6} \sum_{i=1}^N \gamma_i (\Delta_{1,i})^3 + O(\Delta^4),$$

Corrections to correlations



NSL definition of correlation modified due to skew term

Ratio of ρ_{IJ} to linear correlation shows up to 15% correction in toy model

SL approximation for a log-normal

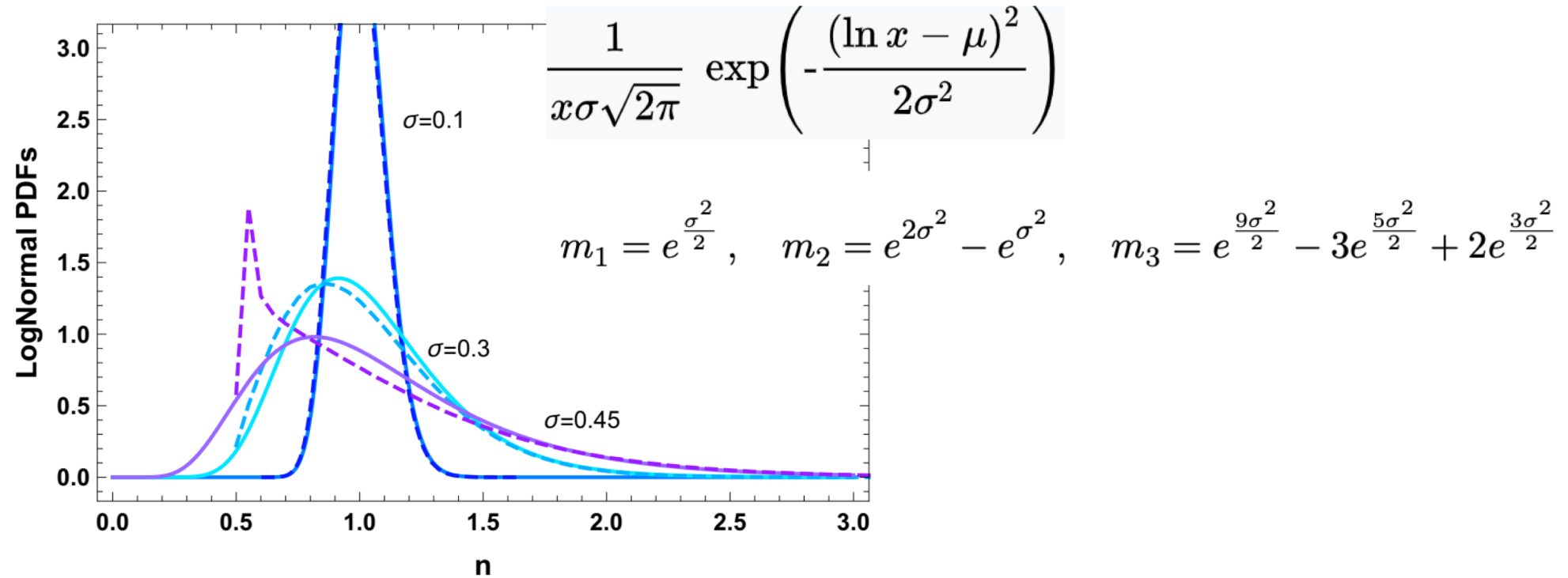


Figure 1. The log normal PDFs and corresponding normal approximations for $\sigma = 0.1, 0.3$ and 0.45 are shown in blue, cyan and purple respectively. Solid curves show the true distributions, dashed curves show the approximate distributions.