



**Colorado  
State**  
University

First laboratory bounds on ultra-light dark photons from precision atomic spectroscopy

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**PHENOMENOLOGY 2022 SYMPOSIUM**  
**May 9th - 11th 2022.**

**Based on work with Joshua Berger**  
**( To Appear )**

# Overview

- 1. Ultra-light dark photons.**
- 2. Search for dark photons using high precision atomic spectroscopy.**

# Introduction: Ultra-light dark photon (ULDP) dark matter

**Extend Standard Model by a massive U(1) gauge boson**

$$\mathcal{L} = \mathcal{L}_{\text{SM}} - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \frac{1}{4} F'_{\mu\nu} F'^{\mu\nu} + m^2 A'_\mu A'^\mu - \frac{e}{(1 + \epsilon)^2} (A_\mu + \epsilon A'_\mu) J_{\text{EM}}^\mu,$$

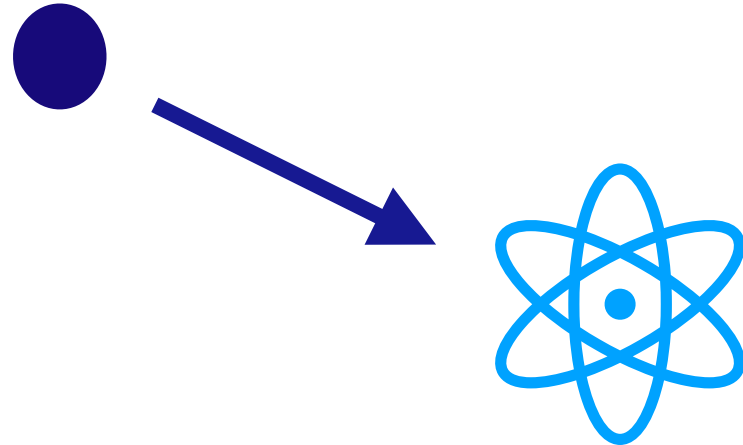
**“Dark” Electromagnetism with a massive photon**

**Stable DM candidate for  $m \ll 2m_e$**

**Ultra-light regime,  $m \simeq \mathcal{O}(10^{-21})$  eV, motivated by puzzles in galactic structure formation simulations**

# Introduction: WIMP vs ULDM

## WIMP



**Low number density**

**Single particle deposits small amount of energy to atom or nucleus**

**Low threshold detectors**

## ULDM



**High number density**

**Single particle picture not applicable: coherent source**

**Energy deposit too small to trigger even low threshold detectors**



# Detecting ULDP using High Precision Atomic Spectroscopy

**Need new detection techniques that exploit coherent nature of ULDP**

**Can think of ULDP as a background Electromagnetic Field**

$$\mathbf{E} \simeq \sqrt{2\rho_{DM}}\epsilon \sin\left(\frac{m_{\gamma'} c^2 t}{\hbar} + \phi_0\right) \hat{\mathbf{n}} \quad \left. \vphantom{\mathbf{E}} \right\}$$

**Time  
varying  
Stark  
shift in  
atoms**

**Time  
varying  
Zeeman  
shift in  
atoms**

$$\left\{ \mathbf{B} \simeq v\sqrt{2\rho_{DM}} \epsilon \sin\left(\frac{m_{\gamma'} c^2 t}{\hbar} + \phi_0\right) \hat{\mathbf{n}}' \right.$$

$\rho_{DM} \simeq 0.3 \text{ GeV cm}^{-3}$  DM energy density

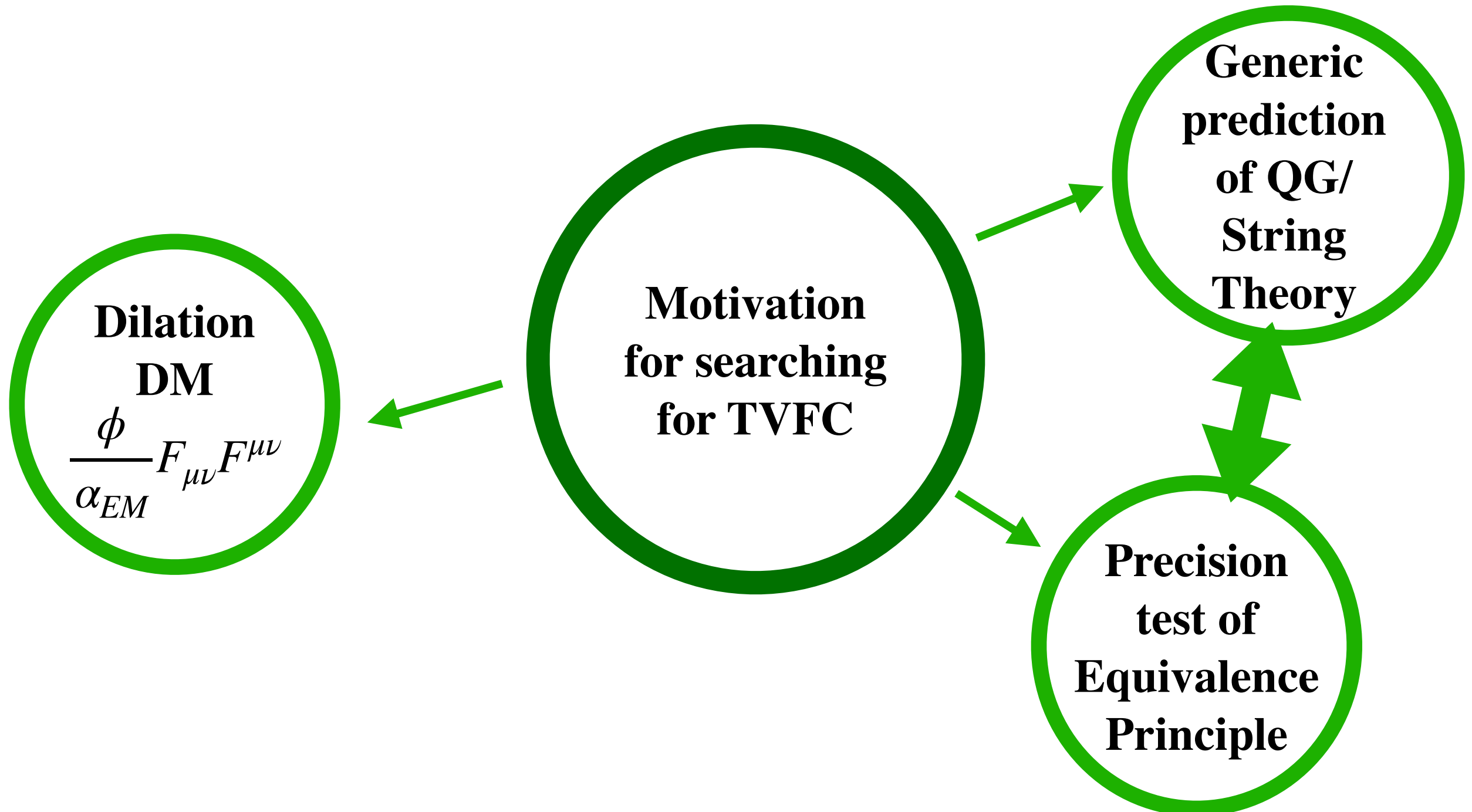
$v \simeq 10^{-3} c$  DM velocity

# Detecting ULDP using High Precision Atomic Spectroscopy

Time varying shift in atomic frequencies



Time variation of fundamental constants (TVFC)



# Detecting ULDP using High Precision Atomic Spectroscopy

Three atomic clocks at NIST/JILA Boulder.

Atom	Transition	Energy (eV)
Al <sup>+</sup>	$3s^2 \ ^1S_0 \text{ — } 3s3p \ ^3P_0$	4.643
Sr	$5s^2 \ ^1S_0 \text{ — } 5s5p \ ^3P_0$	1.776
Yb	$4f^{14}6s^2 \ ^1S_0 \text{ — } 4f^{14}6s6p \ ^3P_0$	2.145

## MOTIVATIONS

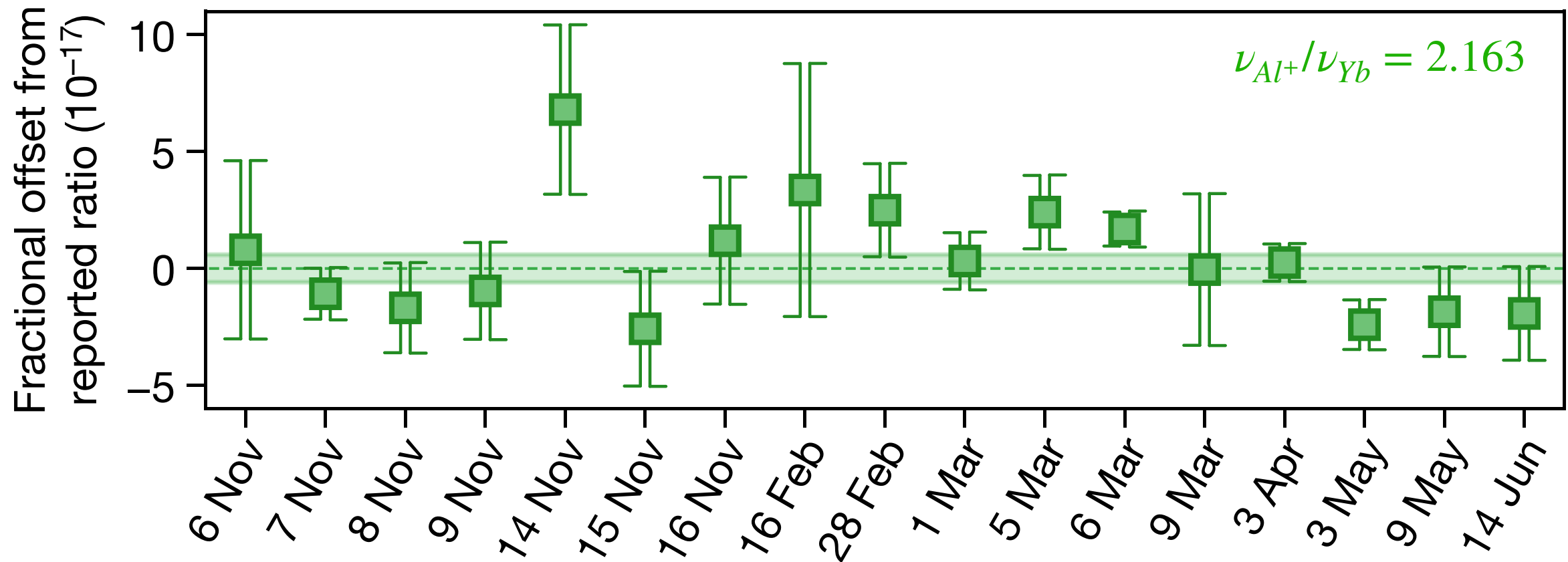
**Better fractional uncertainty**

$^1S_0 \rightarrow ^3P_0$  **relatively insensitive to external B and E fields. No first order Zeeman shift. Better background control**

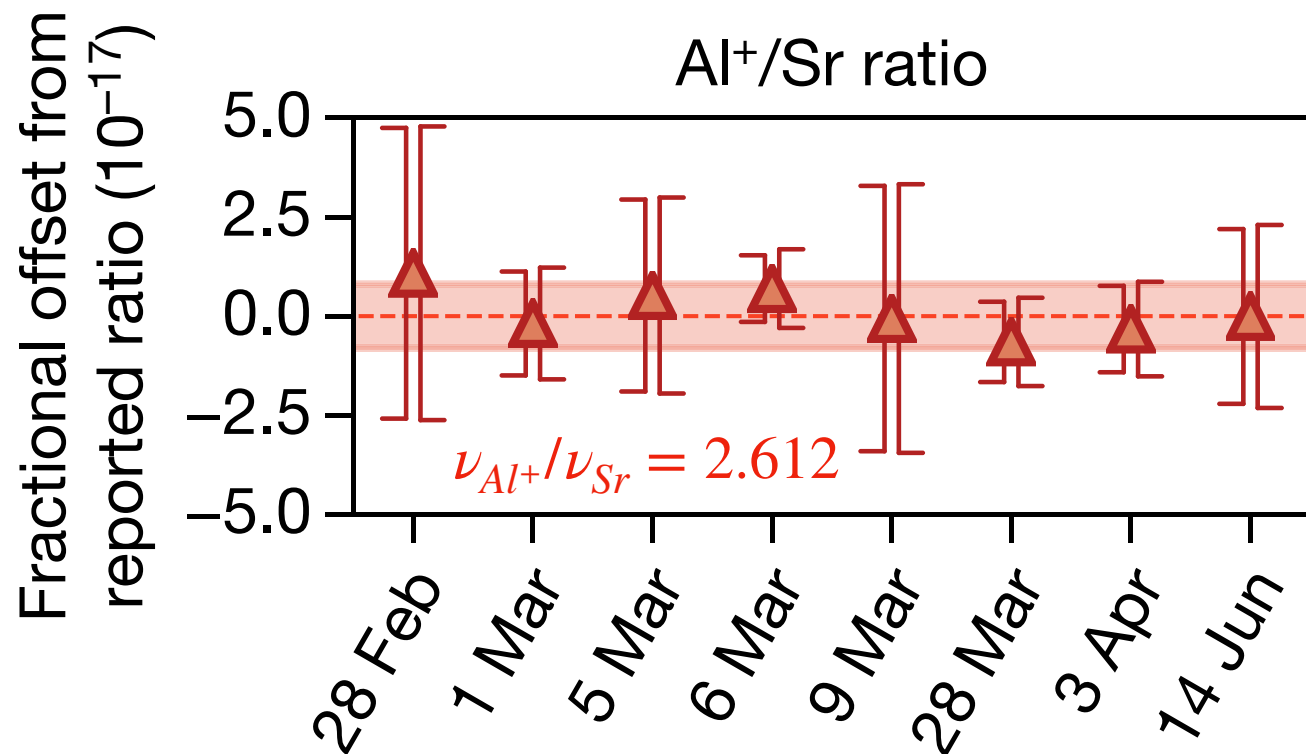
# Detecting ULDP using High Precision Atomic Spectroscopy

Nature 591, 564–569 (2021)

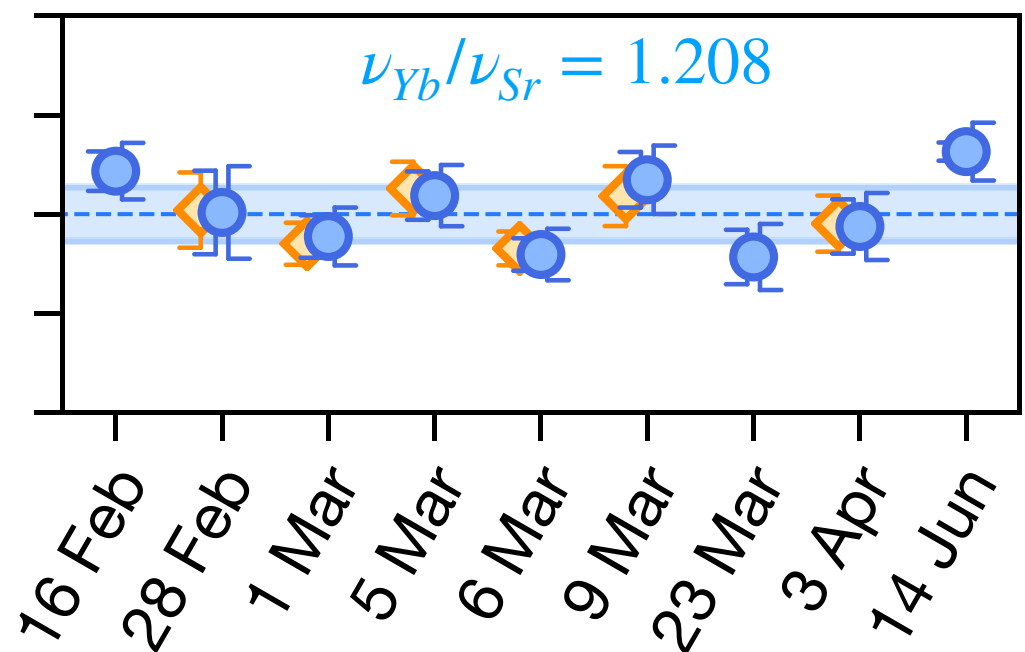
Al<sup>+</sup>/Yb ratio



Al<sup>+</sup>/Sr ratio

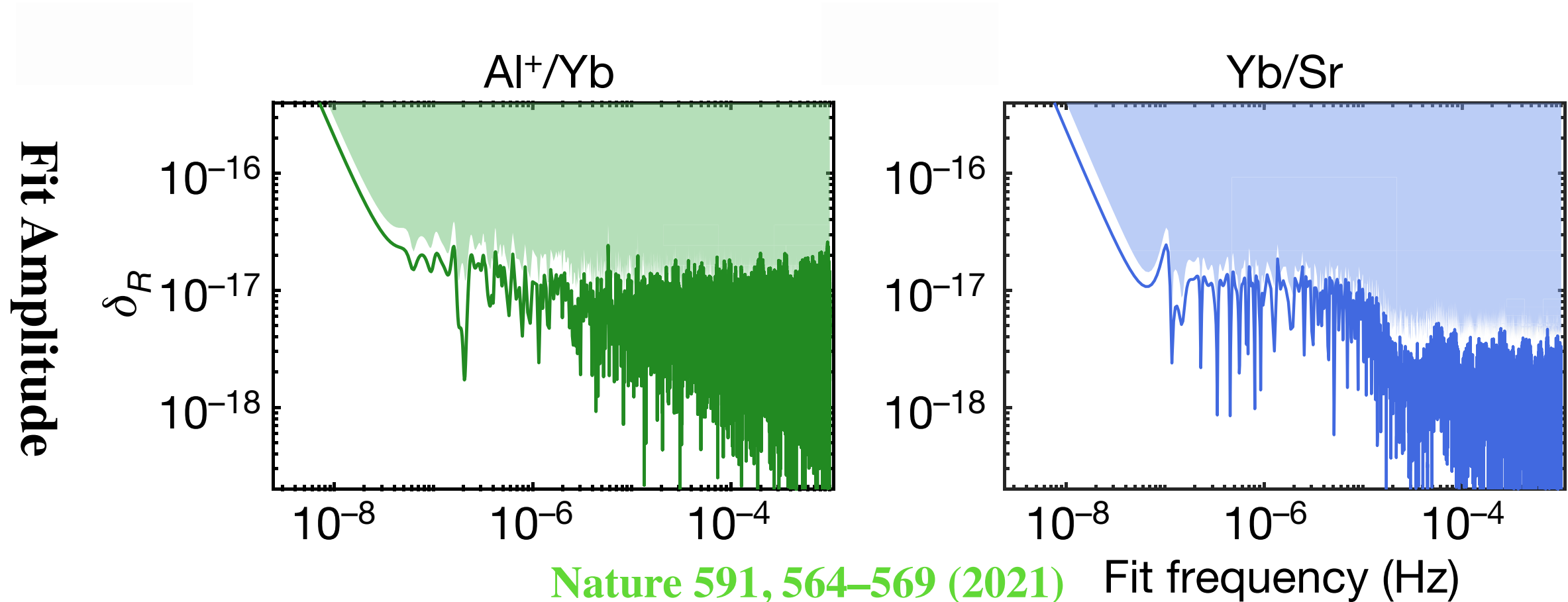


Yb/Sr ratio



# Detecting ULDP using High Precision Atomic Spectroscopy

Fit a sinusoidal signal to time series ratio data. Previously done for dilation  
DM



Fit Frequency: dark photon mass  $\sim \sin^2(mt + \phi_0)$

$$\delta_R = 2\pi\epsilon^2\rho_{\text{DM}}a_0^3\left(\frac{\Delta\alpha_{D,1}}{\hbar\omega_1} - \frac{\Delta\alpha_{D,2}}{\hbar\omega_2}\right)$$

# Preliminary Results

