

# On Perturbative Completions to the Neutrino Option: No-Go Constraints

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[2009.11813]

[2010.15428]

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10 May 2022 || Pheno Symposium 2022 || Pittsburgh, PA, USA

# Outline

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**The Neutrino Option and its phenomenology**



**UV completions & global symmetries**

**No-go constraints**



# The Neutrino Option: Type-I seesaw

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- The Type-I seesaw model (Minkowski '77, et al.) is perhaps the most popular mechanism for generating light neutrino masses:

$$\mathcal{L}_N = \frac{1}{2} (\bar{N}_p i \not{\partial} N_p - \bar{N}_p M_{pr} N_r) - [\bar{N}_p \omega_{p\beta} \tilde{H}^\dagger l_\beta + \text{h.c.}] \quad N_p = e^{i\theta_p/2} N_{R,p} + e^{-i\theta_p/2} (N_{R,p})^c$$

- Upon integrating out heavy sterile neutrinos  $N$ , one induces a contribution to the (Weinberg '79) operator of the dim-5 SM-EFT:

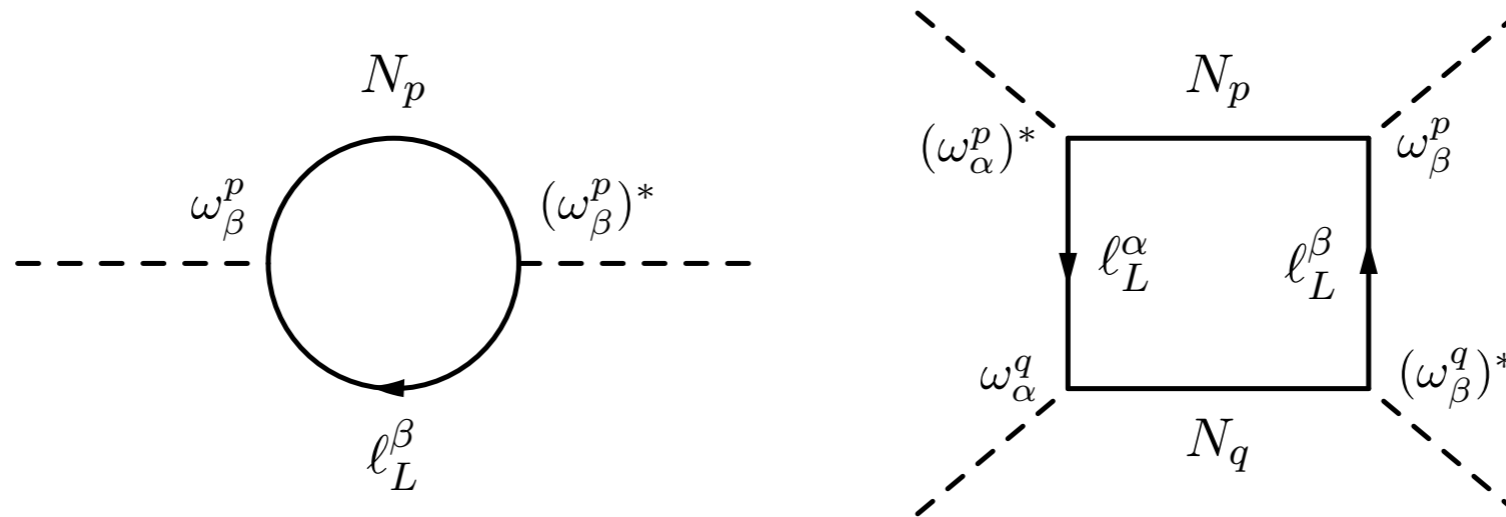
$$\mathcal{L}^{(5)} = \frac{c_{\alpha\beta}^{(5)}}{2} (l_\alpha^T \tilde{H}^*) C (\tilde{H}^\dagger l_\beta) + \text{h.c.} \quad c_{\alpha\beta}^{(5)} = (\omega^T M^{-1} \omega)_{\alpha\beta}$$

- Upon EWSB, this then describes light, LH Majorana neutrinos in accord with data:

$$\mathcal{L}^{(5)} \supset -\frac{m_{\nu,k}}{2} \overline{\nu_L^{lc,k}} \nu_L^{lk} + \text{h.c.} \quad m_{\nu,k} = -\frac{v^2}{2} (U^T)_{k\alpha} c_{5,\alpha\beta} U_{\beta k}$$

# The Neutrino Option: basic idea

- But integrating out heavy  $N$  does more than just induce light neutrinos...



$$V(H) = -\frac{m_{h0}^2 + \Delta m_h^2}{2} H^\dagger H + (\lambda_0 + \Delta\lambda) (H^\dagger H)^2 \quad \longrightarrow \quad \Delta m_h^2 = M_p^2 \frac{|\omega_p|^2}{8\pi^2}$$

- This is either **(A)** a direct manifestation of the EW hierarchy problem (Vissani 1998),
- or **(B)** a route to a minimal solution of the EW hierarchy problem (Brivio, Trott 2017)!  
This scenario is the so-called **Neutrino Option**.

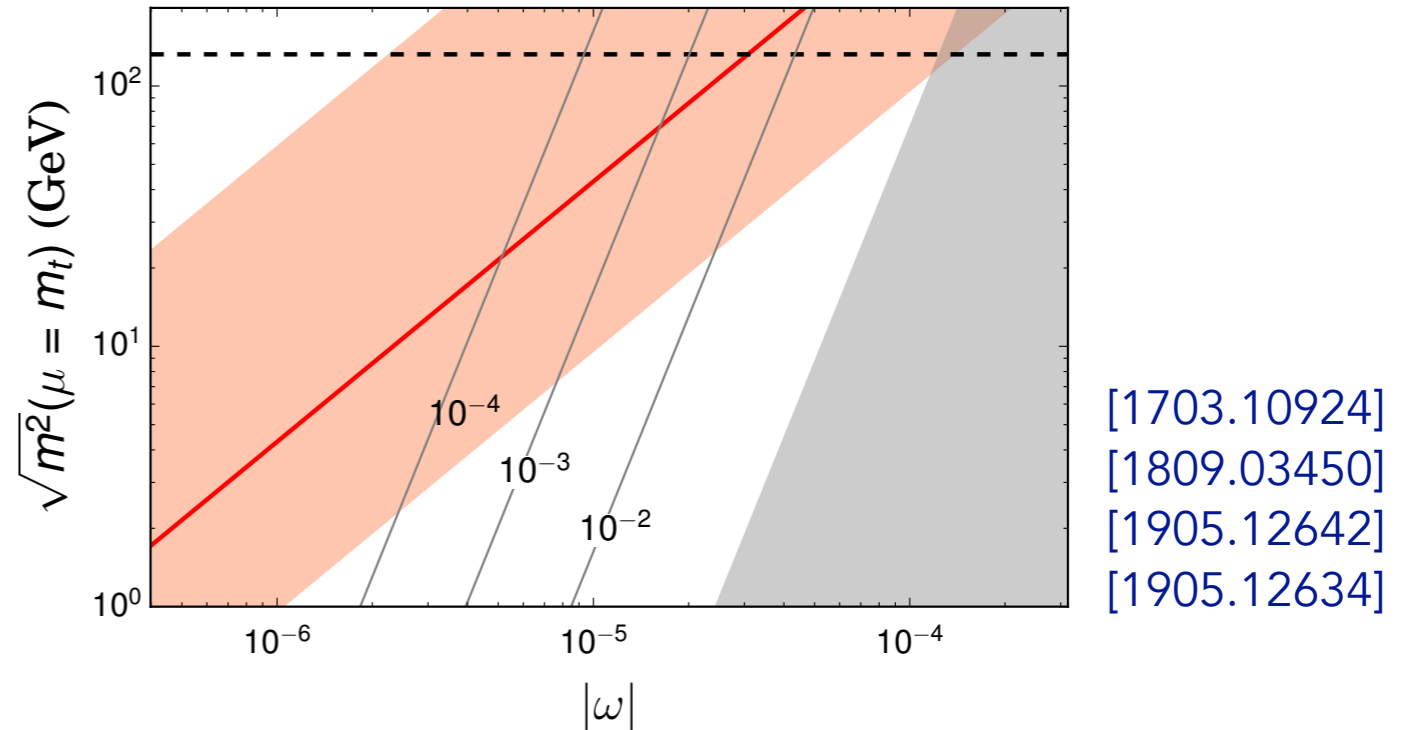
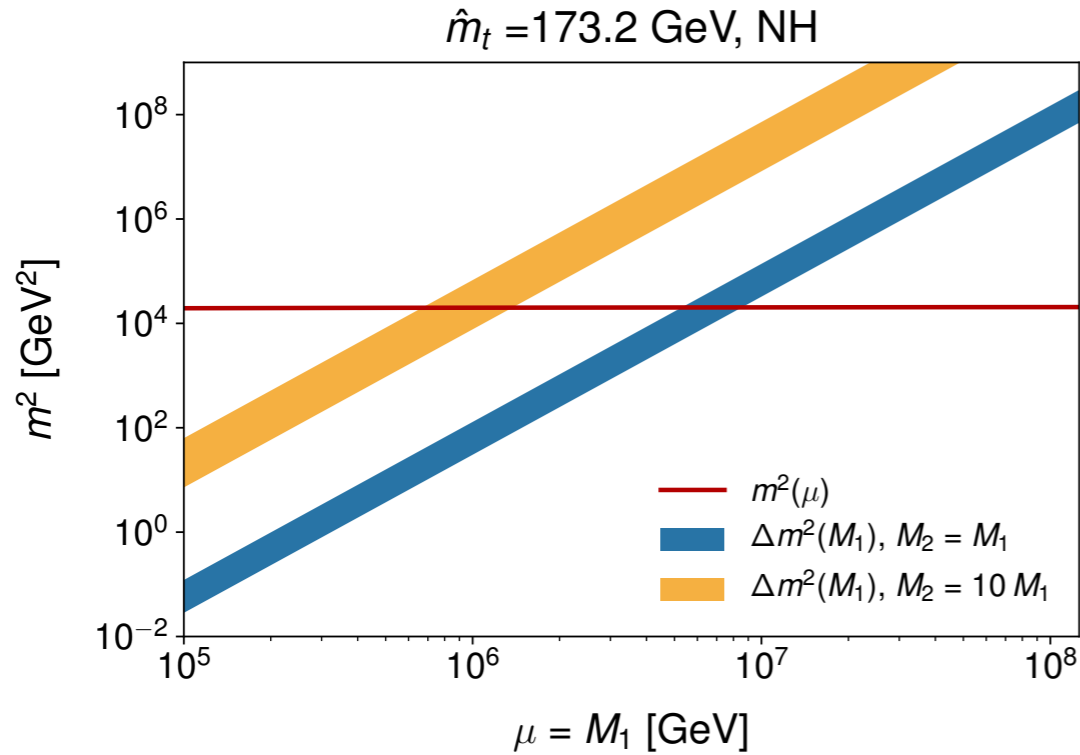
$$m_\nu \sim \frac{\omega_p^2 \bar{v}_T^2}{M_p}, \quad m_h \sim \frac{\omega_p M_p}{4\pi}, \quad \bar{v}_T \sim \frac{\omega_p M_p}{4\sqrt{2}\pi\sqrt{\lambda}}$$

[1703.10924]

[1809.03450]



# The Neutrino Option: phenomenology



Higgs Mass

Neutrino Masses

$$\overline{\mathbf{125.10 \pm 0.14}} \quad + \quad \frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2} \quad \frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2} \quad + \quad \text{Other precision inputs (e.g. top mass) \& RGE} =$$

(2020 PDG Average) 7.42<sup>+0.21</sup><sub>-0.20</sub> 2.514<sup>+0.028</sup><sub>-0.027</sub> (2020 NuFIT)

$$M \lesssim 10^4 \text{ TeV} \simeq 10 \text{ PeV} \quad \left| \omega \right| \simeq \frac{\text{TeV}}{M} \quad + \quad \text{Leptogenesis} \quad M \gtrsim 1 \text{ PeV}$$

# What about the Majorana scale?!

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Hierarchy Problem \_\_\_\_\_



Neutrino Masses \_\_\_\_\_



CP Violation / Leptogenesis -



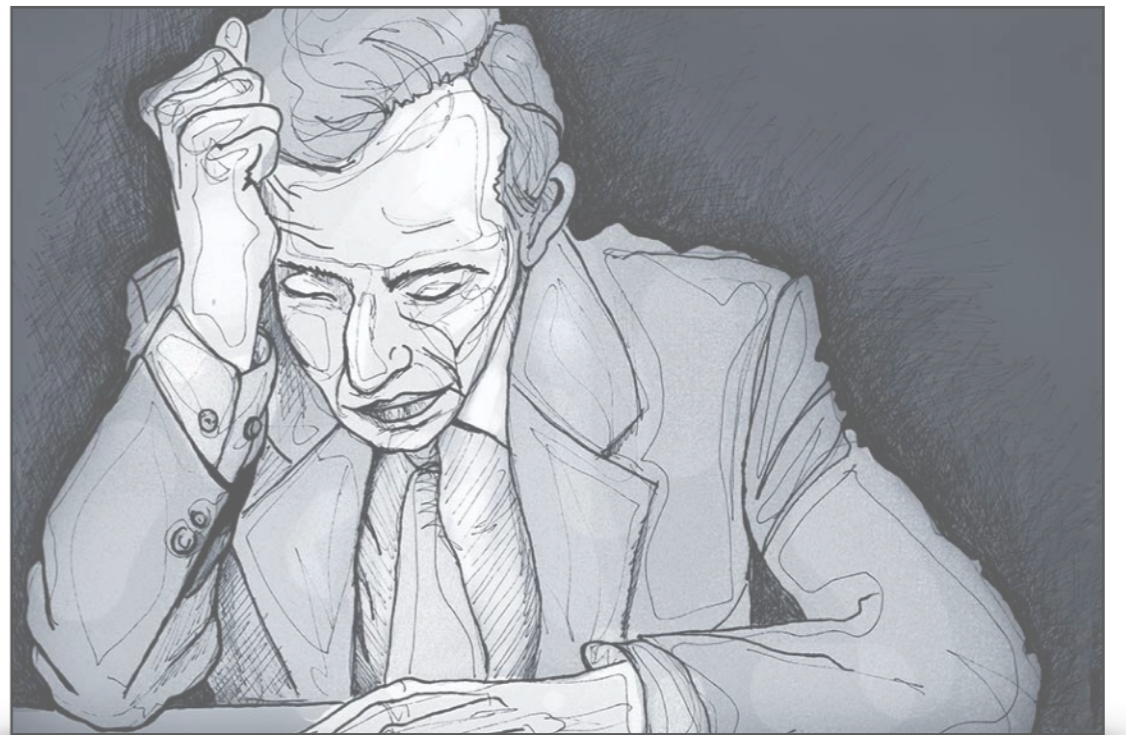
Dark Matter \_\_\_\_\_



Flavour Problem \_\_\_\_\_



...



K.Steegmans

- The Neutrino Option is consistent with a number of the outstanding issues of BSM physics!
- **BUT**, an explanation for the required PeV scale Majorana neutrinos is needed...

What are the minimal options for UV-completing the Neutrino Option?

# UV completions: the minimal requirements

In general, a successful UV completion to the Neutrino Option will

1. generate at least two non-zero sterile neutrino masses.
2. not introduce additional large threshold corrections to the Higgs mass, from beyond the Majorana mass sector.
3. not generate unsuppressed EFT terms proportional to  $(\bar{N}N)(H^\dagger H)$ .
4. not spoil the RGE of Higgs and neutrino parameters via new states.
5. not introduce fine-tuning of parameter space.

# Generic symmetric perturbations

[2010.15428]

- Consider a minimal scenario where a single UV Majorana scale is given by GUT or Plankian dynamics. E.G. :

UV mechanism plausibly flavor-blind, see e.g. [9205230]

$$M = \frac{M_{UV}}{3} \begin{pmatrix} 1 & 1 & 1 \\ 1 & 1 & 1 \\ 1 & 1 & 1 \end{pmatrix} \longrightarrow M = \begin{pmatrix} 0 & & \\ & 0 & \\ & & M_{UV} \end{pmatrix}$$

- Consider generic perturbations about symmetric matrix elements:

$$U_0^f = \begin{pmatrix} a & a & a \\ a & a & a \\ a & a & a \end{pmatrix} \longrightarrow \boxed{U_0 + U = U_0 + \sum_k \alpha_k \cdot \chi^k}$$

Induced by some perturbative UV mechanism...

$$\lambda_{U_0^f} \in \{3a, 0, 0\}$$

$$\chi^{(1,2,3)}|_{3D} \in \left\{ \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix} \right\}$$

$$\chi^{(4,5,6)}|_{3D} \in \left\{ \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \begin{pmatrix} 0 & 0 & 1 \\ 0 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix}, \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} \right\}$$

- Both diagonal and off-diagonal perturbations can generate additional non-zero eigenvalues:

$$U_0^f + \alpha_k \chi^k|_{k=1,2,3} \implies \lambda_{U_1^f} \in \left\{ 0, \frac{1}{2} \left( 3a + \alpha_k \pm \sqrt{9a^2 - 2a\alpha_k + \alpha_k^2} \right) \right\}$$

$$U_0^f + \alpha_k \chi^k|_{k=4,5,6} \implies \lambda_{U_1^f} \in \left\{ -\alpha_k, \frac{1}{2} \left( 3a + \alpha_k \pm \sqrt{9a^2 + 2a\alpha_k + \alpha_k^2} \right) \right\}$$

- GOAL:** identify physical mechanism for generating perturbations to M(UV).
- BUT:** we must simultaneously **DE**-couple  $M_{UV}$  from Higgs threshold corrections...

# Global symmetries of the Type-I seesaw

[2010.15428]

- Before looking into these, let's return to the Type-I seesaw Lagrangian:

$$\mathcal{L}_N = \frac{1}{2} (\bar{N}_p i \not{\partial} N_p - \bar{N}_p M_{pr} N_r) - [\bar{N}_p \omega_{p\beta} \tilde{H}^\dagger l_\beta + \text{h.c.}]$$

- This exhibits a number of global (continuous & discrete) symmetries, depending on the number of sterile N generations (n):

## Kinetic Terms

$$U(1)_{N,3} \times SU(3)_N$$

$$U(1)_{1,3} \times SU(3)_1$$



## Mass Term

$$U(1)_{N,2} \times SU(2)_N \times \mathbb{Z}_2$$

$$U(1)_N \times \mathbb{Z}_2 \times \mathbb{Z}_2$$

$$\mathbb{Z}_2 \times \mathbb{Z}_2$$

[Lam 2007]

## Yukawa Terms

$$(n = 1)$$

$$(n = 2)$$

$$(n = 3)$$



$$U(1)_{1+N} \supset U(1)_{1,3} \times U(1)_{N,3}$$

- Augment the (a priori) accidental  $\mathbb{Z}_2$  symmetry to a UV fixed-point *natural* symmetry of the UV Lagrangian:

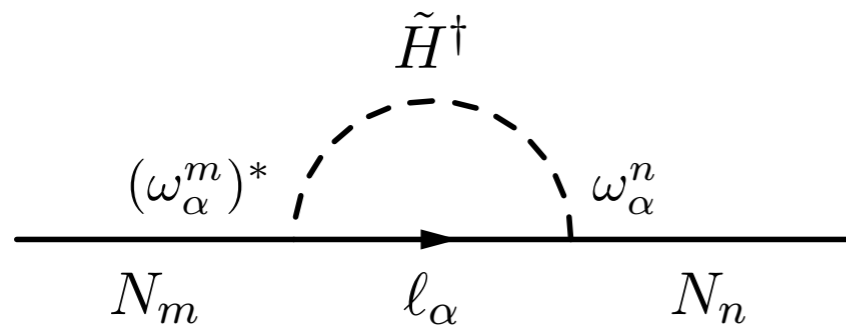
$$N_p \longrightarrow T_{pr} N_r, \quad \text{with} \quad T_{pr} = \text{diag}(1, 1, -1)$$

$$\begin{aligned} \bar{N} \omega \tilde{H}^\dagger l &\stackrel{!}{=} \bar{N} T^\dagger \omega \tilde{H}^\dagger l \\ \Rightarrow T^\dagger \omega &= \omega \Rightarrow \omega_{3\beta} \equiv 0 \end{aligned}$$

- Finally, recall that non-zero M eigenvalues correspond to  $\Delta L = 2$  violation.

# One-loop perturbative corrections

- The relevant one-loop diagram to consider in the Type-I seesaw is given by:



???

$$\delta_M^{(1)} = \frac{|\omega|^2}{16\pi^2} M_{UV}$$

$$\delta_M^{(1)} \simeq \text{PeV} \quad \text{for} \quad \begin{cases} |\omega| \simeq 10^{-4}, & (M_{UV} \simeq M_{GUT}) \\ |\omega| \simeq 10^{-5.5}, & (M_{UV} \simeq M_{Pl}) \end{cases}$$

- Unfortunately, neither **threshold corrections** nor **RGE** can induce non-zero masses from initially massless N at one-loop, in minimal setup:

$$\Delta L = 2 \quad \times$$

- Option (1):** consider  $n = 2$  UV mass scales. Then corrections go as  $\omega\omega^\dagger$ .

$$\mathbb{Z}_2 \times \mathbb{Z}_2 \quad \times$$

- Allowing for soft  $\mathbb{Z}_2$  breaking, the RGE at one-loop reads:

$$16\pi^2 \mu \frac{dM}{d\mu} = (\omega\omega^\dagger) M + M (\omega\omega^\dagger)^T \equiv \mathcal{R}, \quad \longrightarrow \quad \begin{aligned} M_p(\mu) &= \gamma_p(\mu, \mu_0) M_p(\mu_0), \\ \gamma_p(\mu, \mu_0) &\sim 1 + \frac{\omega^2}{16\pi^2} \ln \left[ \frac{\mu}{\mu_0} \right] \end{aligned} \quad \longrightarrow \quad \gamma_p(\text{PeV}, M_{UV}) = \epsilon \simeq \frac{\text{PeV}}{M_{UV}} \sim 10^{-10} - 10^{-13}$$

**Extreme fine-tuning required!  $\times$**

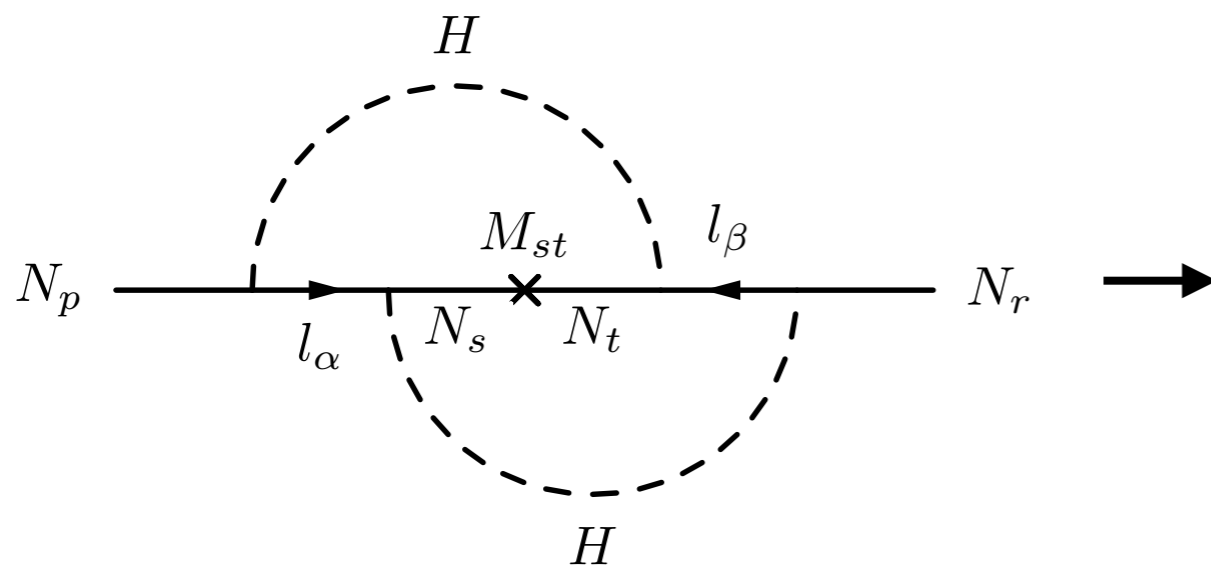


# Multi-loop corrections?

[1802.09997]  
[2006.13584]  
[2010.15428]

- **Option (2):** consider two/multi-loop perturbative generation with  $n=1$ :

## Multi-loop corrections?



$$\delta_{M,pr}^{(2)} \sim \frac{(\omega\omega^\dagger)_{p3}(\omega\omega^\dagger)_{r3}}{256\pi^4} M_{UV}$$

$$\delta_M^{(2)} \neq 0 \text{ only if } \omega_{3\beta} \neq 0$$

**Higgs mass unprotected! ❌**

$$\omega\omega^\dagger M, \quad \omega\omega^\dagger M(\omega\omega^\dagger)^T, \quad (\omega\omega^\dagger)(\omega\omega^\dagger)M$$

$$\mathbb{Z}_2 \text{ ❌}$$

- Even if we allow for soft  $\mathbb{Z}_2$  breaking, the desired mass scales are too low for the PeV scales required. ❌
- Such scales may be interesting, however, for light sterile neutrino phenomenology...
- Multi-loop corrections with new states may also be interesting...

# New BSM states?

[2010.15428]

- Allow a generic new boson (vector or scalar) in a minimal extension to the seesaw.

**New boson?**

$$\frac{(\rho_S^2 + \rho_P^2) M_{33}}{64\pi^2} \left[ 1 + \frac{1}{M_{33}^2 - m_\chi^2} \left( M_{33}^2 \log \frac{\mu^2}{M_{33}^2} - m_\chi^2 \log \frac{\mu^2}{m_\chi^2} \right) \right]$$

$$\frac{3 \rho_V^2 M_{33}^5}{16\pi^2 \Lambda^4} \left[ 1 + \log \frac{\mu^2}{M_{33}^2} \right]$$

$\Delta L$  and  $Z_2$  safe, but...
 
 $\chi\chi HH$   
 $\chi NN$

**Scalar masses unprotected! ❌**

$\frac{1}{2} \bar{N} [\rho_S \chi_S + i \rho_P \gamma_5 \chi_P] N$   
 $\frac{1}{2\Lambda^2} \bar{N} [\rho_V \gamma_\mu \gamma_5] N \partial_\nu \chi_V^{\nu\mu}$

- For states with non-trivial SM gauge #s, things are no less complicated. MSSM ?? [1006.1092] [1603.04993] yields no new L violation and leptoquarks, e.g., only induce N masses at two-loops!
- Others (Brdar et al.) have considered adding new states, in realizing a **conformal NO**. Strong dynamics, dark matter, and gravity waves explored. [1807.11490] [1905.12634] [1810.12306] [2007.04367] [2006.02960]

**Achieving simple perturbations to M(UV) is not so easy, while realizing the Neutrino Option!**



# Non-perturbative speculations

- What if the Majorana scale is not generated perturbatively, or through a new scalar VEV? What if it is instead **non-perturbative**?

$$m_s e^{-U} \overline{N} N$$

$m_s$ : fundamental (UV) scale

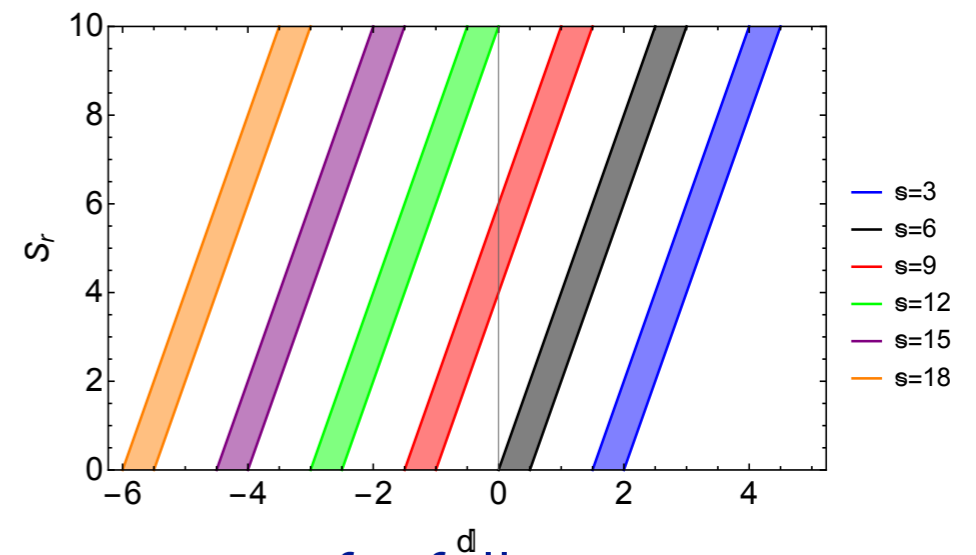
$U$ : suppression factor associated to non-perturbative dynamics

- Such a scenario exists in certain stringy orientifold compactifications:

$$M = 2m_s \sum_r e^{-S_r} \text{diag} \left( d_1^{(r)}, d_2^{(r)}, d_3^{(r)} \right) \cdot \begin{pmatrix} 1 & 1 & 1 \\ 1 & 1 & 1 \\ 1 & 1 & 1 \end{pmatrix} \cdot \text{diag} \left( d_1^{(r)}, d_2^{(r)}, d_3^{(r)} \right)$$

$$s \equiv \log_{10} m_s, \quad d \equiv \log_{10} d_a^{(r)}, \quad p \equiv \log_{10} M_p$$

$$s + 2d - \left( \frac{1}{2} S_r + p \right) \approx 0$$



- However**, this scenario is not yet predictive, nor am I aware of a fully consistent string theory that can generate the N mass term without additional excitations that destabilize the Higgs mass...???
- Other scenarios might exist, e.g. through gravitational condensation (see e.g. Barenboim et al.) ???

# Summary and outlook

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- The **Neutrino Option** represents an elegant and minimal approach to solving the EW hierarchy problem alongside the neutrino mass (and potentially more) problem(s)!
- However, **minimal perturbative explanations** for the origin of the Majorana mass scale required for Neutrino Option phenomenology seem **limited** due to **global (discrete) symmetries** (Lepton Number  $\times$   $Z_2$  Mass).
- Other, less-minimal frameworks which introduce **new states** may be viable, e.g. the Conformal Neutrino Option of Brdar et al..
- **Non-perturbative** mechanisms explaining the origin of the Majorana scale may exist! Further formal analysis is required here.
- Neutrino-Option-inspired **phenomenology** of sterile neutrinos possible — cf. ongoing work with Michael Trott (not discussed in this talk)...

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**THANK YOU!**