

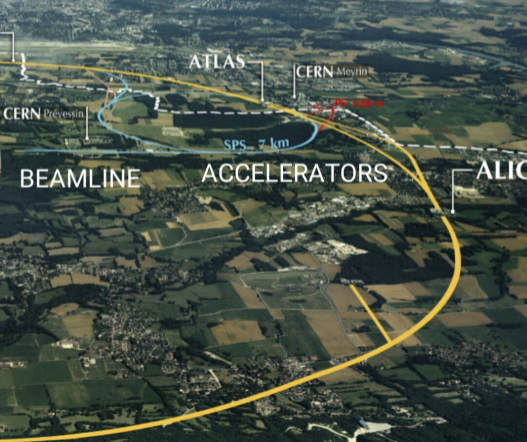
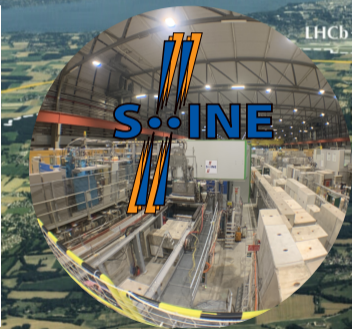
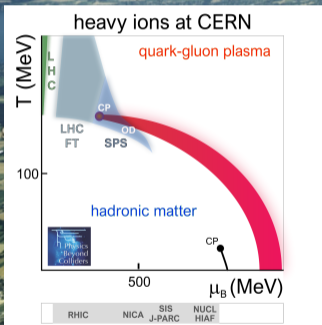


# NA61/SHINE long-term plans

Szymon Pulawski  
for the NA61/SHINE Collaboration

Detailed information is given in the [NA61/SHINE status report 2020](#)

# NA6I/SHINE - UNIQUE MULTIPURPOSE FACILITY: Hadron production in hadron-nucleus and nucleus-nucleus collisions at high energies



# NA61/SHINE 2021-2024 (2025?)



## Beam request submitted to SPSC

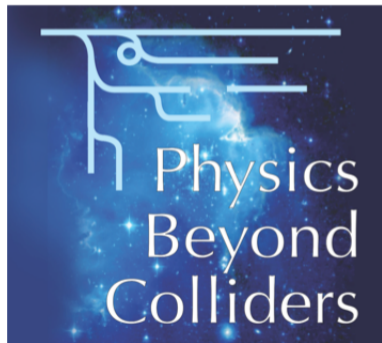
- **STRONG INTERACTIONS:** measurements of **charm hadron production** in **Pb+Pb collisions**
- **COSMIC-RAYS:** measurements of **nuclear fragmentation cross section**
- **NEUTRINO:** measurements of **hadron production induced by hadron beams**

CERN-SPSC-2021-027

# NA61++ 2027+



- Energy scan with light and medium mass ions to study the onset of fireball
- Measurements of heavy hadrons resonances in large statistics p+p interactions
- Measurements of hadron emission from the LBNF and HYPER-K replica targets
- Data for flux predictions in neutrino experiments using very low energy beams



CERN-SPSC-2020-023

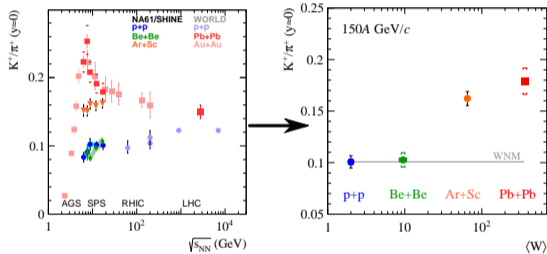


**Energy scan with  
light and medium mass  
ions to study the  
onset of (QGP) fireball**

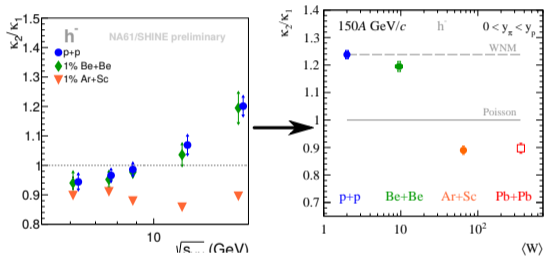
# Current NA61/SHINE results



Mean multiplicities:



Fluctuations:



$$p+p \approx \text{Be+Be} \neq \text{Ar+Sc} \leq \text{Pb+Pb}$$

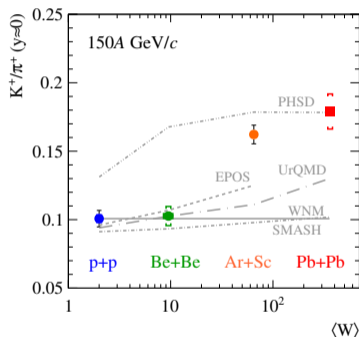
- Beginning of creation of strongly interacting matter with increasing nuclear mass number. Transition from **non-equilibrium** strings and resonances to **equilibrium** hadron gas or quark gluon plasma.

# Data vs models

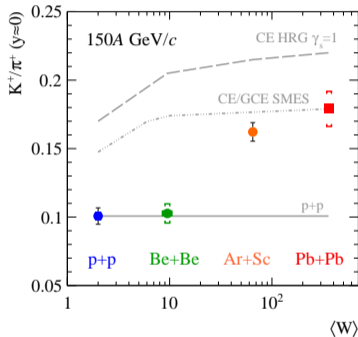


Multiplicities:  $K^+/\pi^+$

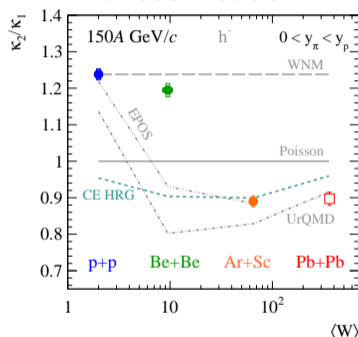
dynamical models



statistical models



Fluctuations:  
various models



- None of actual models describe NA61/SHINE results

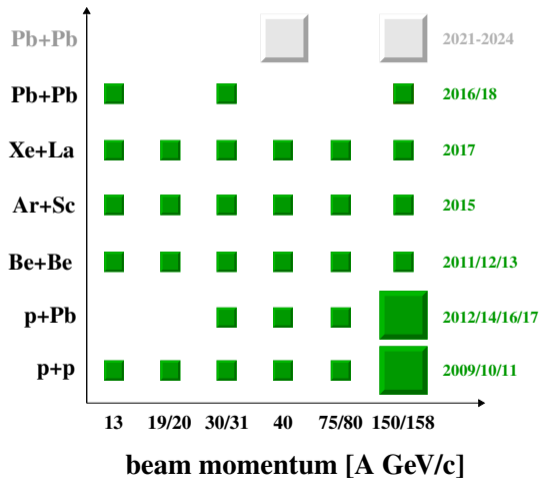
PHSD: Eur.Phys.J.A 56 (2020) 9, 223, arXiv:1908.00451 and private communication;  
 SMASH: J.Phys.G 47 (2020) 6, 065101 and private communication;  
 UrQMD and HRG: Phys. Rev. C99 (2019) 3, 034909  
 SMES: Acta Phys. Polon. B46 (2015) 10, 1991 - recalculated

p+p: Eur. Phys. J. C77 (2017) 10, 671  
 Be+Be: Eur. Phys. J. C81 (2021) 1, 73  
 Ar+Sc: NA61/SHINE preliminary  
 Pb+Pb: Phys. Rev. C66, 054902 (2002)  
 $\kappa_2/\kappa_1$ : NA61/SHINE preliminary

# Onset of fireball

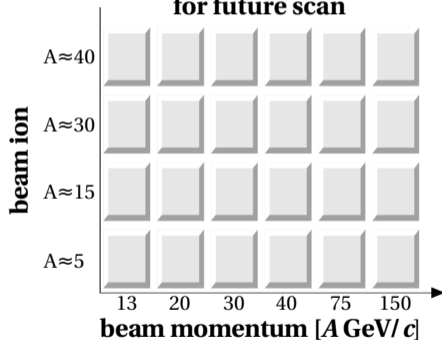


Current 2D scan



+

The very first idea for future scan

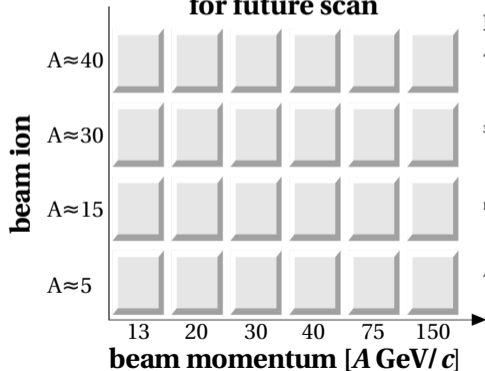




# Onset of fireball

- Ion beam request for the Run 4 period

The very first idea  
for future scan



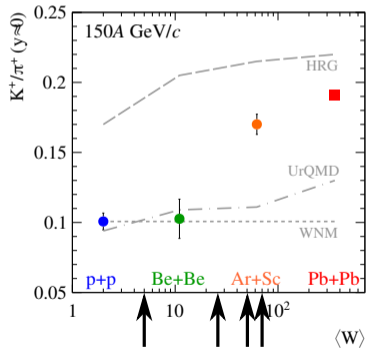
Example ion:

$^{40}\text{Ca}$  - synergy with  
Gamma Factory

$^{30}\text{P}$

$^{16}\text{O}$  - synergy with  
Cosmic-Ray LHC

$^4\text{He}$

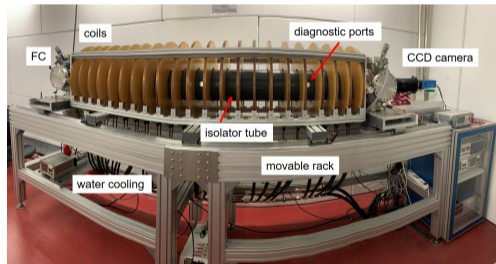


# Improvement of the beam-line



**Large emittance of the low-momentum (below 40A GeV/c) ion beams is always a factor reducing the data-taking efficiency**

- Ongoing efforts for improvement of the low energy ion beam emittance in SPS and in H2 beam-line in collaboration with BE-OP and BE-EA
- Ongoing efforts in investigating new innovative focusing techniques based on electrostatic lenses (“Gabor-lenses”)
- First tests are planned before LS3

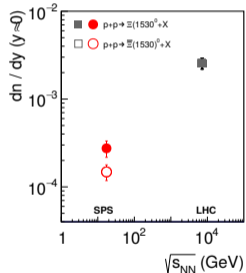
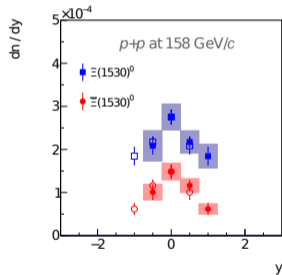
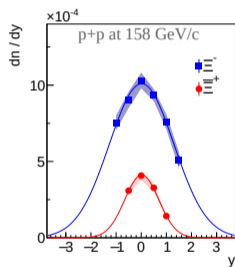




**Measurements of  
heavy hadrons and  
resonances in large  
statistics  $p+p$   
interactions**

# Multi-strange particles in NA61/SHINE

- Large statistics of  $p+p$  interactions at 158 GeV/c gathered in 2009-2011 allowed to obtain unique results on the production of multi-strange hadrons



- Study of heavy baryons and resonances requires a large-statistics data sample

- The only results on  $\Xi^0(1530)$  production in  $p + p$  in the SPS energy range
- Other existing results on  $\Xi^0(1530)$  production only by ALICE at  $\sqrt{s_{NN}} = 7$  TeV

# Heavy particles-there is more

Baryon resonance particles

Λ particles		Σ particles		Ξ and Ω particles		Charmed particles		Bottom particles	
Λ	1/2 <sup>+</sup> ****	Σ <sup>+</sup>	1/2 <sup>+</sup> ****	Ξ <sup>0</sup>	1/2 <sup>+</sup> ****	Λ <sub>c</sub> <sup>+</sup>	1/2 <sup>+</sup> ****	Λ <sub>b</sub> <sup>0</sup>	1/2 <sup>+</sup> ****
Λ(1405)	1/2 <sup>-</sup> ****	Σ <sup>0</sup>	1/2 <sup>+</sup> ****	Ξ <sup>-</sup>	1/2 <sup>+</sup> ****	Λ <sub>c</sub> (2595) <sup>+</sup>	1/2 <sup>-</sup> ****	Λ <sub>b</sub> (5912) <sup>0</sup>	1/2 <sup>-</sup> ****
Λ(1520)	3/2 <sup>-</sup> ****	Σ <sup>-</sup>	1/2 <sup>+</sup> ****	Ξ(1530)	3/2 <sup>+</sup> ****	Λ <sub>c</sub> (2625) <sup>+</sup>	3/2 <sup>-</sup> ****	Λ <sub>b</sub> (5920) <sup>0</sup>	3/2 <sup>-</sup> ****
Λ(1600)	1/2 <sup>+</sup> ***	Σ(1385)	3/2 <sup>+</sup> ****	Ξ(1620)	*	Λ <sub>c</sub> (2765) <sup>+</sup>	*	Σ <sub>b</sub>	1/2 <sup>+</sup> ****
Λ(1670)	1/2 <sup>-</sup> ****	Σ(1480)	*	Ξ(1690)	***	Λ <sub>c</sub> (2880) <sup>+</sup>	3/2 <sup>+</sup> ****	Σ <sub>b</sub> <sup>-</sup>	3/2 <sup>+</sup> ****
Λ(1690)	3/2 <sup>-</sup> ****	Σ(1560)	**	Ξ(1820)	3/2 <sup>-</sup> ****	Λ <sub>c</sub> (2940) <sup>+</sup>	3/2 <sup>-</sup> ****	Ξ <sub>b</sub> <sup>0</sup> , Ξ <sub>b</sub> <sup>-</sup>	1/2 <sup>+</sup> ****
Λ(1710)	1/2 <sup>+</sup> *	Σ(1580)	3/2 <sup>-</sup> *	Ξ(1950)	***			Ξ <sub>b</sub> (5935) <sup>-</sup>	1/2 <sup>+</sup> ****
Λ(1800)	1/2 <sup>-</sup> ***	Σ(1620)	1/2 <sup>-</sup> *	Ξ(2030)	≥ 3/2 <sup>+</sup> ****	Σ <sub>c</sub> (2455)	1/2 <sup>+</sup> ****	Ξ <sub>b</sub> (5945) <sup>0</sup>	3/2 <sup>+</sup> ****
Λ(1810)	1/2 <sup>+</sup> ***	Σ(1660)	1/2 <sup>+</sup> ****	Ξ(2120)	*	Σ <sub>c</sub> (2520)	3/2 <sup>+</sup> ****	Ξ <sub>b</sub> (5955) <sup>-</sup>	3/2 <sup>+</sup> ****
Λ(1820)	5/2 <sup>+</sup> ****	Σ(1670)	3/2 <sup>-</sup> ****	Ξ(2250)	**	Σ <sub>c</sub> (2800)	***	Ω <sub>b</sub> <sup>-</sup>	1/2 <sup>+</sup> ****
Λ(1830)	5/2 <sup>-</sup> ****	Σ(1690)	**	Ξ(2370)	**			P <sub>c</sub> (4380) <sup>+</sup>	*
Λ(1890)	3/2 <sup>+</sup> ****	Σ(1730)	3/2 <sup>+</sup> *	Ξ(2500)	*	Ξ <sub>c</sub> <sup>0</sup>	1/2 <sup>+</sup> ****	P <sub>c</sub> (4450) <sup>+</sup>	*
Λ(2000)	*	Σ(1750)	1/2 <sup>-</sup> ***			Ξ <sub>c</sub> <sup>0</sup>	1/2 <sup>+</sup> ****		
Λ(2020)	7/2 <sup>+</sup> *	Σ(1770)	1/2 <sup>+</sup> *	Ω <sup>-</sup>	3/2 <sup>+</sup> ****	Ξ <sub>c</sub> <sup>+</sup>	1/2 <sup>+</sup> ****		
Λ(2050)	3/2 <sup>-</sup> *	Σ(1775)	5/2 <sup>-</sup> ****	Ω(2250) <sup>-</sup>	***	Ξ <sub>c</sub> <sup>0</sup>	1/2 <sup>+</sup> ****		
Λ(2100)	7/2 <sup>-</sup> ****	Σ(1840)	3/2 <sup>+</sup> *	Ω(2380) <sup>-</sup>	**	Ξ <sub>c</sub> (2645)	3/2 <sup>+</sup> ****		
Λ(2110)	5/2 <sup>+</sup> ***	Σ(1880)	1/2 <sup>+</sup> **	Ω(2470) <sup>-</sup>	**	Ξ <sub>c</sub> (2790)	1/2 <sup>-</sup> ****		

- Many poorly known meson and baryon states.
- SPS energies well suited for their measurements

**$K^*(892)$**

$$I(J^P) = \frac{1}{2}(1^-)$$

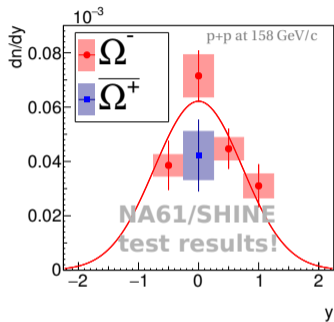
895.50 ± 0.92 ± 2.6

48.8 ± 1.8 ± 2.0

<sup>9</sup> ADUSZKIEW...20A NA61 158 pp

<sup>8</sup> ADUSZKIEW...20A NA61 158 pp

# Large-statistics data sample needed



## Assuming

- average SPS super cycle: 40 s flat top: 4.8 s
- typical data taking efficiencies summarized in [NA61/SHINE Addendum 2018](#), Sec 9.3
- 1 kHz read-out rate

**We could achieve 600M p+p interactions over a month of the data taking**

- Study of  $\Omega$  already suffers from having only  $\sim 60$ M events



## Measurements for flux predictions in neutrino experiments

# Measurements of hadron production for neutrino and cosmic-ray physics



- **NA61/SHINE will continue to support needs of neutrino and cosmic ray communities after LS3 in the scope of:**
  - ▶ hadron yields measurements with use of LBNF and Hyper-Kamiokande replica targets
  - ▶ hadron production using low-energy hadron beams (1-20 GeV/c), where existing hadron production data is sparse, but high statistics of high-quality data is needed
- **Addendum to NA61/SHINE proposal submitted to SPSC**  
[CERN-SPSC-2021-028](#)



# Needs for measurements employing low-E beams

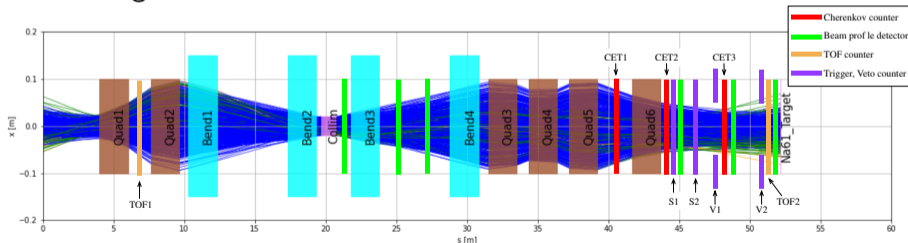
- **Accelerator-based neutrino experiments:** to reduce the leading uncertainty on neutrino flux prediction (that is hadron production)  
**Requirement:** [T2K/Hyper-K] low-E secondary hadron interactions measurements (1-10 GeV/c pions, kaons, protons); [SBN] p+Be at 8 GeV/c
- **Atmospheric neutrino experiments:** to understand low-E hadron production for sub-GeV neutrino flux (0.1-1 GeV neutrinos)  
**Requirement:** low-E proton beam (1-20 GeV/c) on nitrogen or carbon target
- **Spallation neutron source neutrino experiments:** to understand pion production rate from mercury target  
**Requirement:** [JSNS2 at J-PARC] 3 GeV/c proton on a mercury target; [SNS at ORNL] 1 GeV/c proton on a mercury target

# Very Low Energy beam-branch idea



NA61/SHINE with the strong support of BE-EA experts is advancing on studies for the implementation of a **Very Low Energy** (VLE) beam-line still before LS3:

- ▶ simulation of an intermediate (secondary) target for low-E beam production
- ▶ investigation of various options for particle identification in the beam-line
- ▶ investigation on beam-line designs aiming to maximize acceptance, minimize beam-line length and cost



CERN-SPSC-2021-028

# Milestones



- **Now:**

- ▶ main focus on completion of the upgrade and preparation to the approved physics programme
- ▶ gathering the ideas for 2027+ physics programme
- ▶ initial studies within working groups
- ▶ discussion of the collaboration format (NA61/SHINE or a new collaboration)

- **2022:**

- ▶ open workshop on NA61/SHINE future

- **2023:**

- ▶ formulation of a letter of intend for 2027+ physics programme



**Thank you**

NA61/SHINE would like to thank the CERN EP, BE, HSE and EN Departments for support

# Summary

- **Status and plans for 2021-2024:**

- ▶ The detector upgrade is ongoing.
- ▶ Readiness for physics data taking in spring 2022
- ▶ Beam request under discussion with SPSC

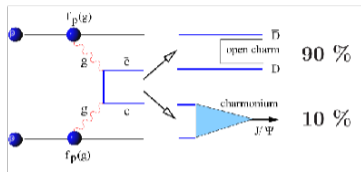
- **Ideas for 2027+:**

- ▶ **Study of the onset of fireball - energy scan with low and medium mass ions**
- ▶ **Measurements of heavy hadrons and resonances in large statistics  $p+p$  interactions**
- ▶ **Measurements with anti-proton beams**
- ▶ **Hadron emission study using replica targets from the LBNF and HYPER-K**
- ▶ **Measurements to improve quality of flux predictions in neutrino experiments using very low energy beams**

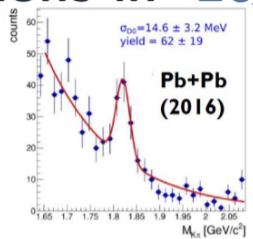


**Additional slides**

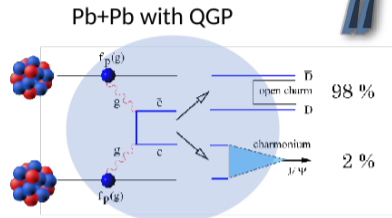
# Strong interactions in 2021-2024



Open charm and  $J/\psi$  production within Matsui-Satz model [PL B178 416]



NA61/SHINE pilot measurements



- What is the mechanism of open charm production?
- How does the onset of deconfinement impact open charm production?
- How does the formation of quark-gluon plasma impact  $J/\psi$  production?

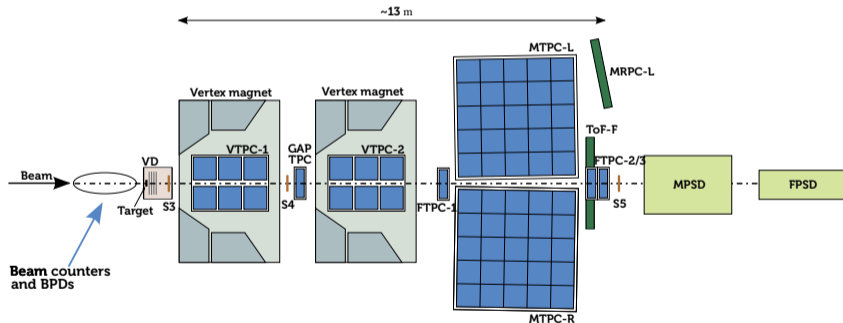
Medium reduces probability of  $J/\psi$  production

$$P(c\bar{c} \rightarrow J/\psi) \equiv \frac{\langle J/\psi \rangle}{\langle c\bar{c} \rangle} \equiv \frac{\sigma_{J/\psi}}{\sigma_{c\bar{c}}}$$

$$P_{\text{vacuum}}(c\bar{c} \rightarrow J/\psi) > P_{\text{medium}}(c\bar{c} \rightarrow J/\psi)$$

# NA61/SHINE's status

- Work on the upgrade is ongoing. Full readiness of the NA61/SHINE detector foreseen for spring 2022..



- First physics data taking in 2022 - beam request submitted to SPSC.