

# Forward Physics Facility (FPF): Neutrinos

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# Physics beyond colliders: FPF

## Proposal: Forward Physics Facility for the High Luminosity Era

### High energy collisions:

- collect particles along the beamline
- shielded by at least 100 m concrete and rock

Early suggestion by De Rujula & Ruckl (1984);  
Winter (1990), De Rujula, Fernandez, Gomez, NP  
B405 (1993); Vannucci (1993).

### Measurements and searches:

- neutrinos – weakly-interacting
- BSM particles – light, extremely weakly-interacting



# Physics beyond colliders: FPF

LHC Run 3: 480 m from ATLAS interaction point in existing tunnels TI12 and TI18

- FASER $\nu$  approved in December 2019
- FASER approved March 2019
- SND@LHC approved March 2021

## Measurements and searches:

- neutrinos – weakly-interacting
- BSM particles – light, extremely weakly-interacting

FPF: Bigger space, infrastructure



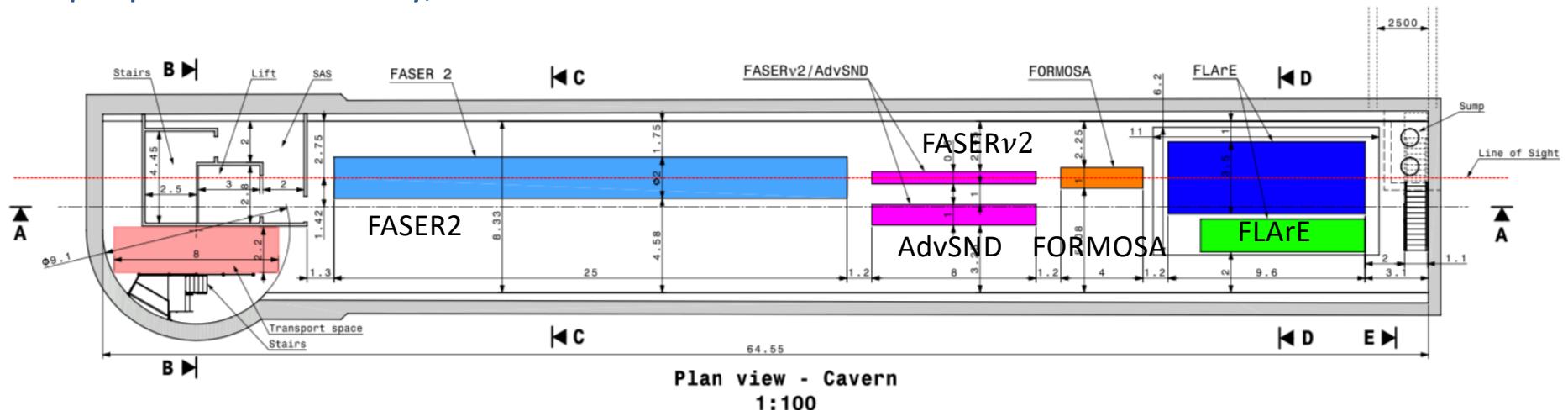
# Proposed Experiments

- **FLArE** – *neutrinos*, LArTPC
- FORMOSA – Forward MicroCharge Search, BSM search, plastic scintillator

In a purpose-built facility, would like this:

Talk by Jamie Boyd  
later today.

- FASER2 – BSM search, magnetized spectrometer
- FASER $\nu$ 2 – *neutrinos*, emulsion-based
- AdvSND (and AdvSND2) – *neutrinos*, calorimeters

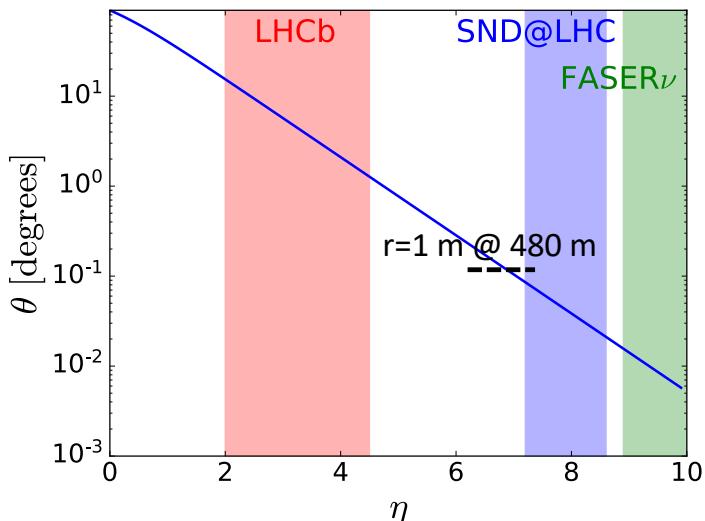


Short white paper 2109.10905

# Studies underway: Snowmass white papers

Short white paper - arXiv: 2109.10905, 80 authors, submitted to *Physics Reports*

Long white paper – in progress, J. Feng, F. Kling, J. Rojo, D. Soldin, MHR (editors)



How forward is FPF and  
Run 3 experiments?

- A. Facility and Experiments
- B. Tools: MC and other codes
- C. BSM Physics
- D. QCD
- E. Neutrino physics
- F. Astroparticle physics

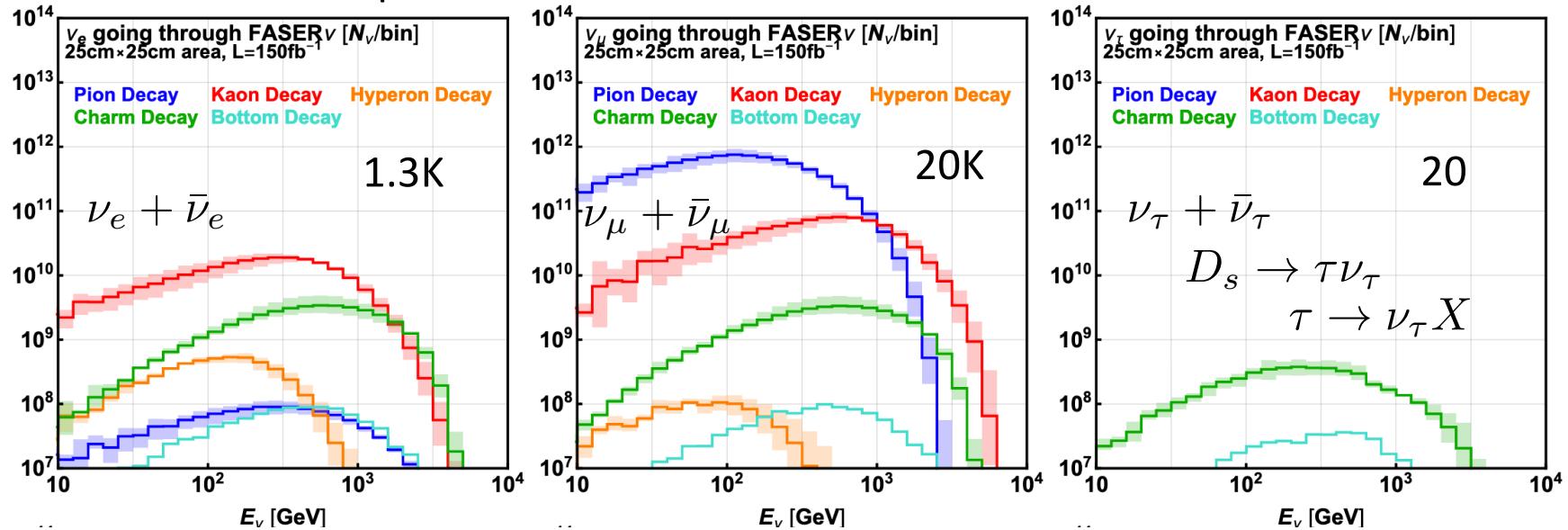
Snowmass Slack channel: #fpf-whitepaper  
4<sup>th</sup> FPF Workshop: January 31–February 1, 2022

# Neutrino production

- Charm production  $\sigma_{\bar{c}c}$  and fragmentation
- Charm hadron energy, decay distributions

Monte Carlo simulations of pion, kaon, charm production and decay – wider variation than shown here. Forward tuning ... and NLO+ QCD for charm hadron production.

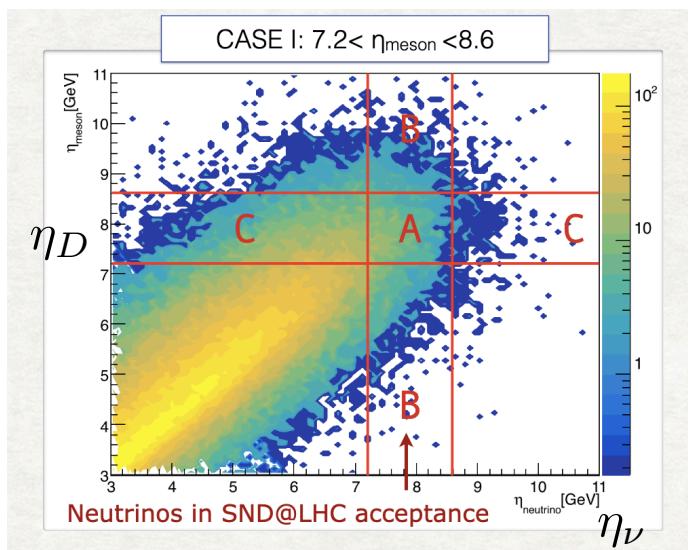
FASER $\nu$  example:



FASER Collab, Eur. J. Phys. C80 (2020) 81, 1908.02310. See also Kling, 2105.08270, Bai et al., JHEP 06 (2020) 032, 2002.03012

# Charm production

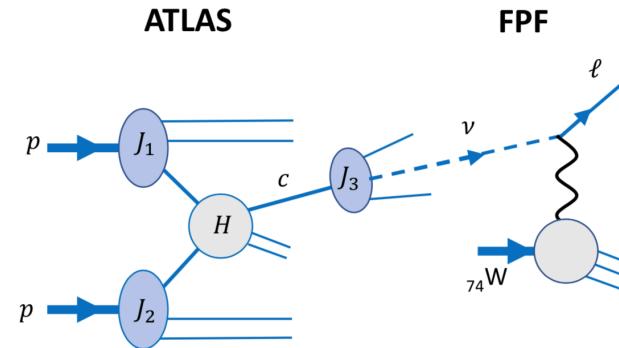
Forward neutrino  $\eta_\nu$  correlated with  $y_c$ .  
 meson=charm meson



A. Di Crescenzo for SND@LHC, 3<sup>rd</sup> FPF Meeting

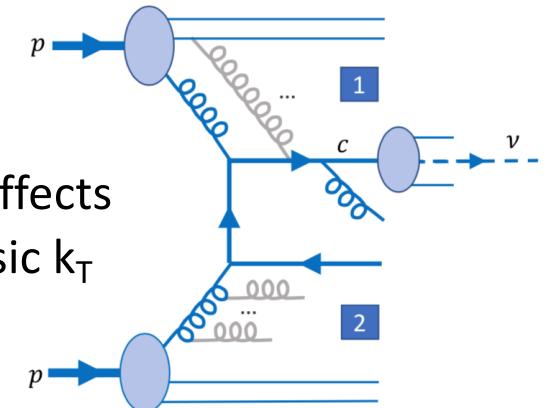
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## Many opportunities:

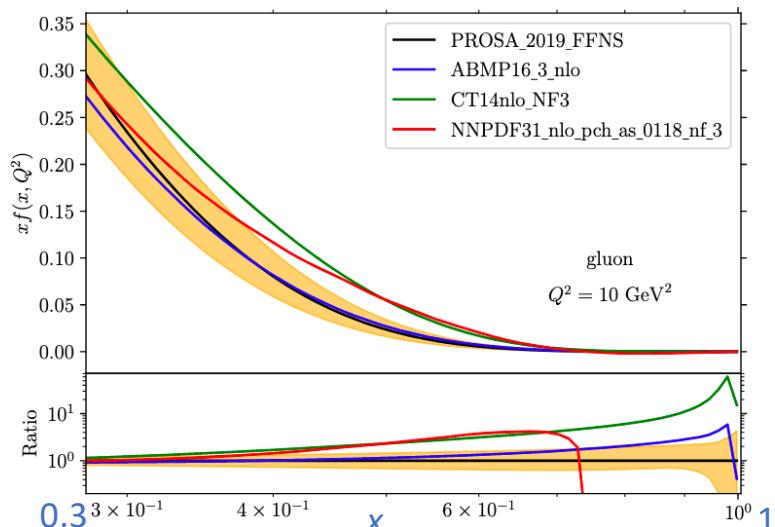
- PDFs, small  $x$  and large  $x$
- PDFs, intrinsic charm
- Fragmentation, spectator effects
- Higher order effects, intrinsic  $k_T$



Short white paper 2109.10905

# PDFs at large $x$ and small $x$

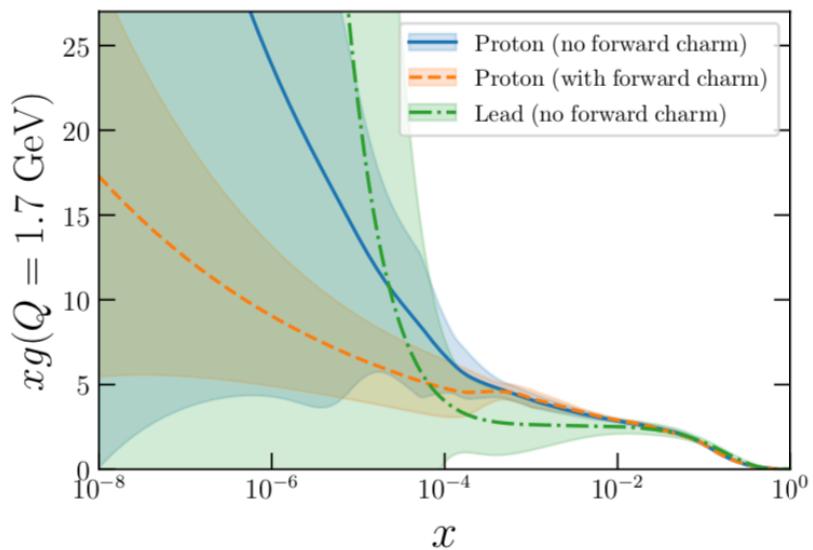
Large  $x$  effects at high energy



Impacts HE neutrino flux, Bai et al., in preparation

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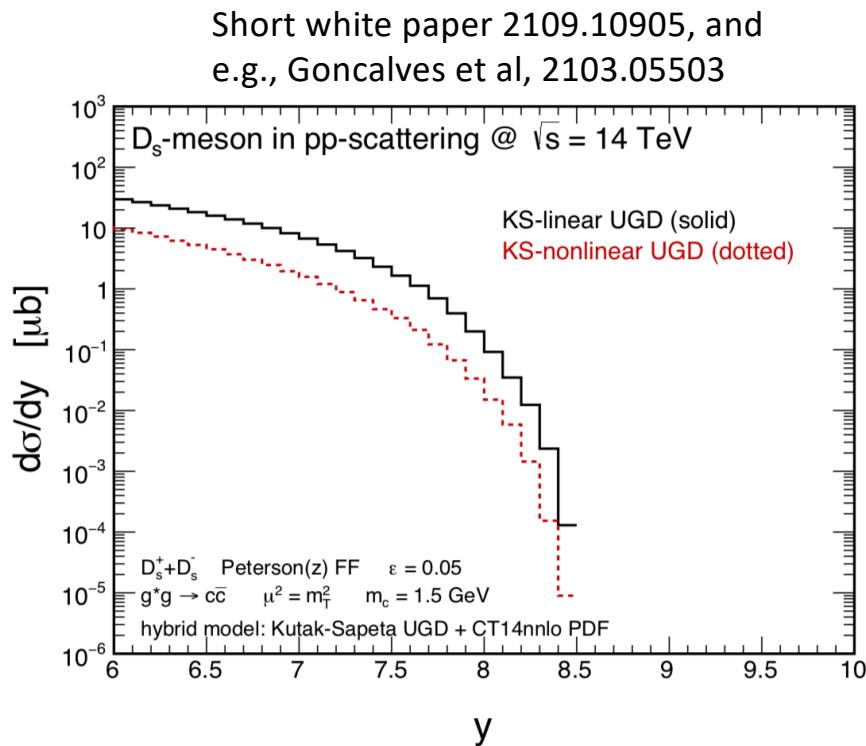
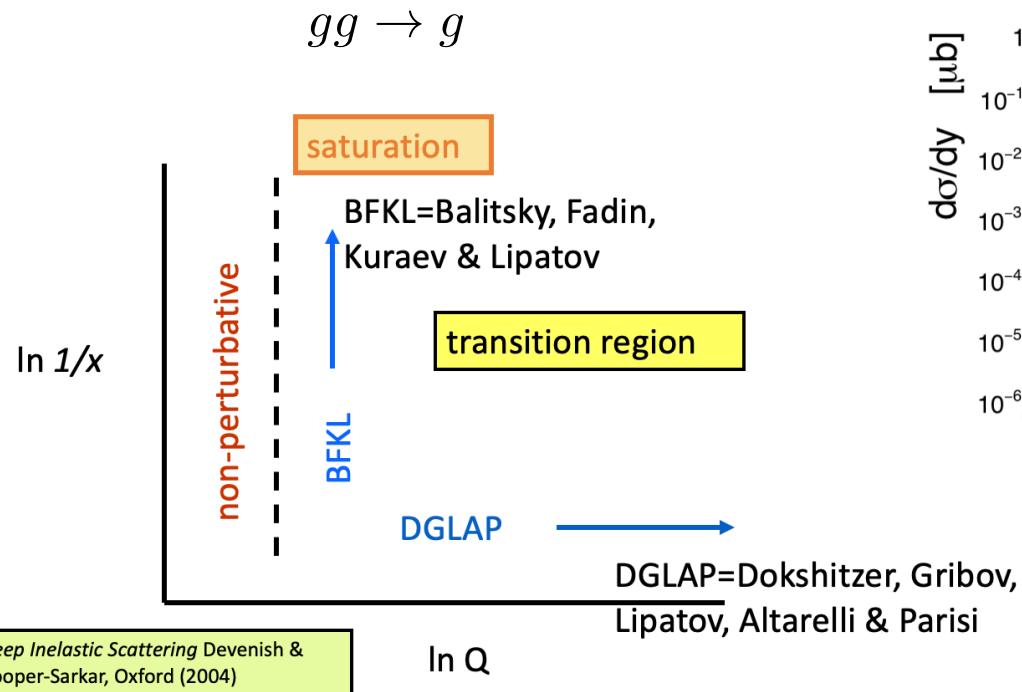
Low  $x$  uncertainty, NNPDF3.1 with/  
 without LHCb charm at 5, 7, 13 TeV.  
 [See also Lead (green).]



Short white paper 2109.10905, Gauld&Rojo, PRL 118 (2017)  
 072001, Zenaiev et al (PROSA), Eur. Phys. J C75 (2015) 396.

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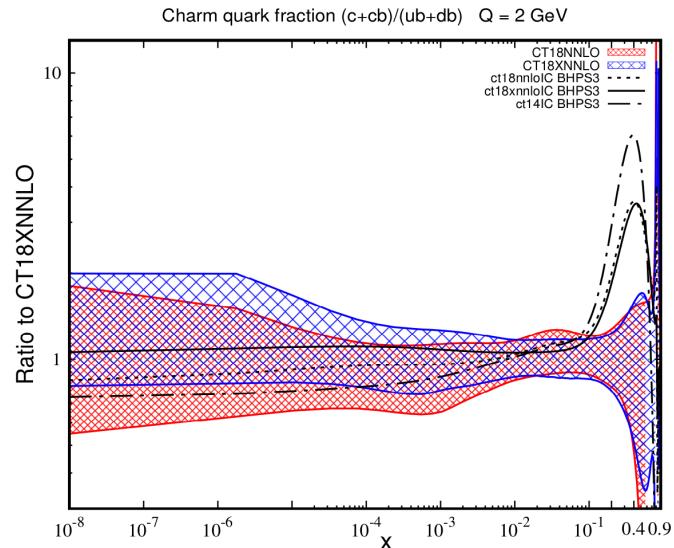
# Gluon saturation



Example: BFKL resummation,, small  $x k_T$  dependent unintegrated DF, with/wo saturation, large  $x$  collinear PDF

# Intrinsic charm

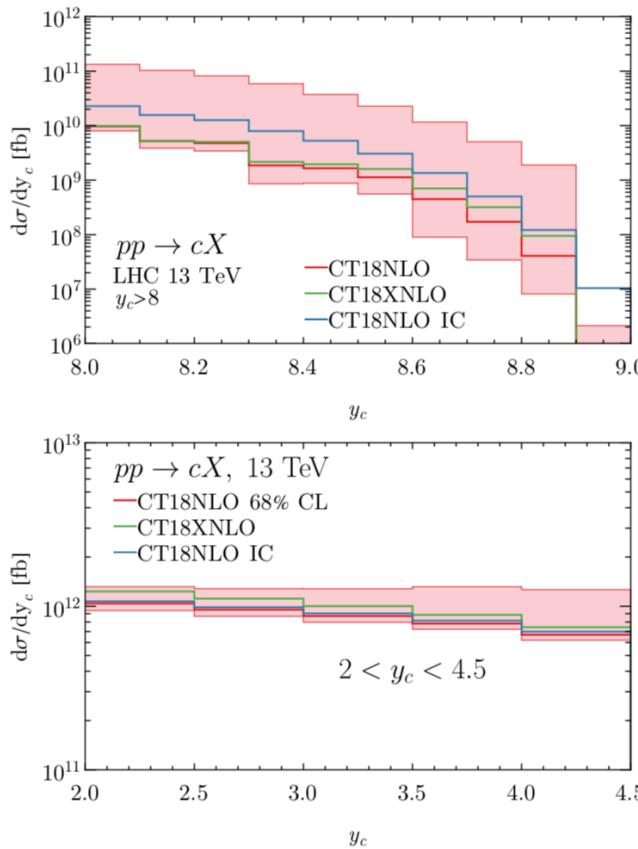
## charm quark fraction



Short white paper 2109.10905 and Hou et al,  
PRD 103 (2021) 014013

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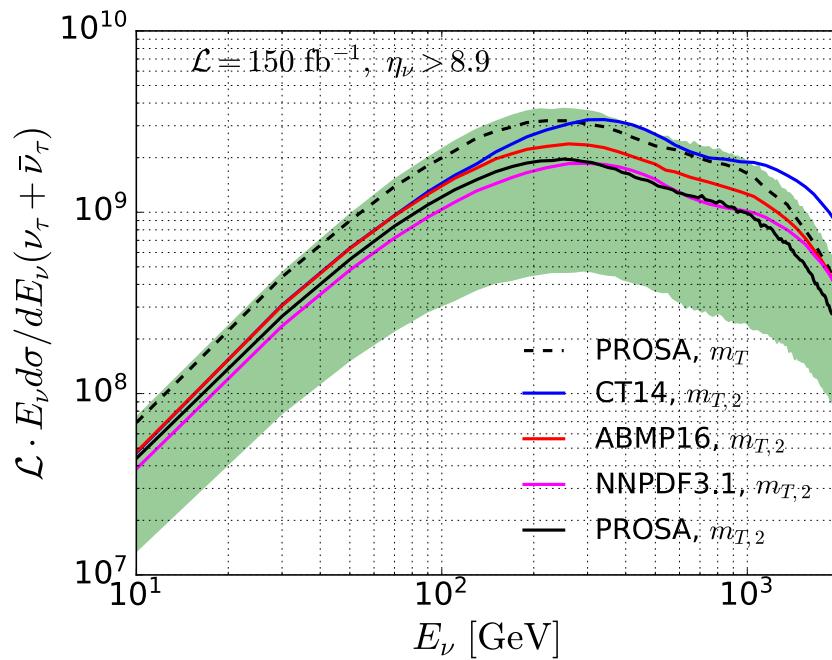
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Relations between different rapidity regions.

- CT18NLO
- CT18XNLO (tuned  $\mu_F(x, Q^2)$  in DIS)
- CT18NLO IC (intrinsic charm)

# $\nu_\tau + \bar{\nu}_\tau$ , collinear NLO QCD production of $D_s$



Bai et al., in preparation. See also,  
 Bai et al., JHEP 06 (2020) 032.

- Collinear parton model with Gaussian  $k_T$  smearing.
  - Scale dependence: green band around central black curve.
  - Based on PROSA PDFs and scale:
- $$m_{T,2}^2 = p_T^2 + 4m_c^2 \quad \langle k_T \rangle = 0.7 \text{ GeV}$$
- Tie forward physics to LHCb charm measurements.

# Numbers of neutrinos

Example: predictions from 2 leading-order generators. Still work to be done....

Detector			Interactions at FPF			
Name	Mass	Coverage	CC $\nu_e + \bar{\nu}_e$	CC $\nu_\mu + \bar{\nu}_\mu$	CC $\nu_\tau + \bar{\nu}_\tau$	NC
FASER $\nu$ 2	20 tonnes	$\eta \gtrsim 8.5$	178k / 668k	943k / 1.4M	2.3k / 20k	408k / 857k
FLArE	10 tonnes	$\eta \gtrsim 7.5$	36k / 113k	203k / 268k	1.5k / 4k	89k / 157k
AdvSND1	2 tonnes	$7.2 \lesssim \eta \lesssim 9.2$	6.5k / 20k	41k / 53k	190 / 754	17k / 29k
AdvSND2	2 tonnes	$\eta \sim 5$	29 / 14	48 / 29	2.6 / 0.9	32 / 17

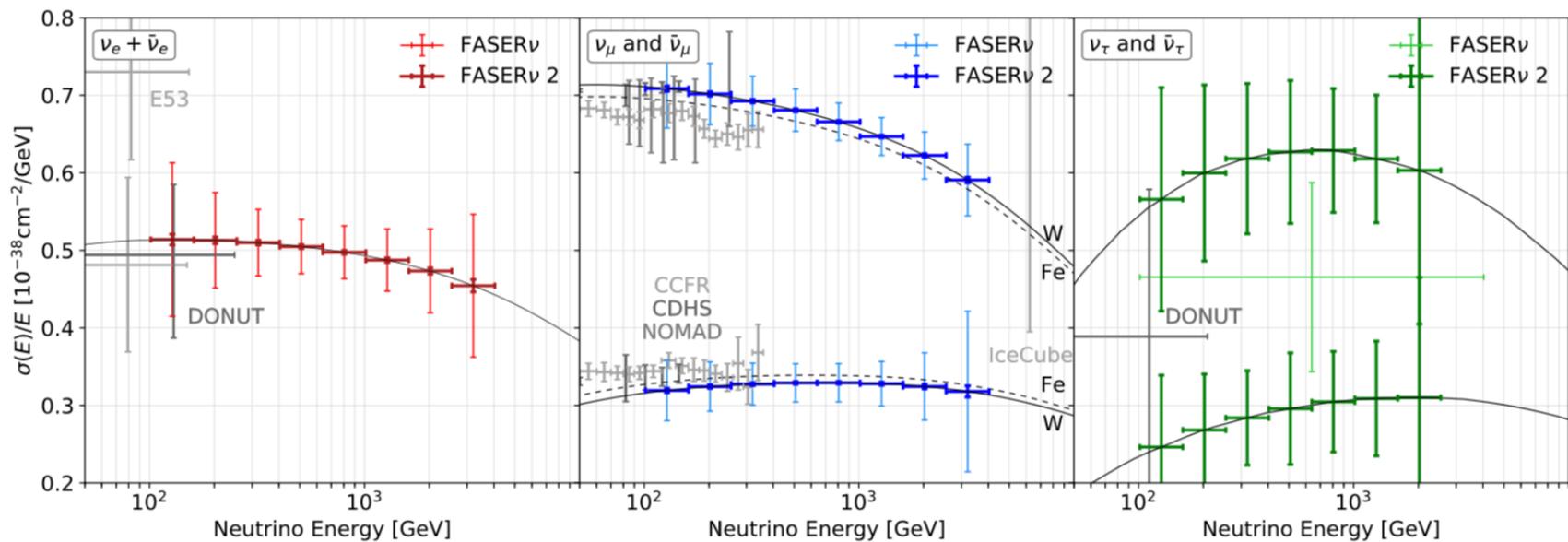
TABLE III. The estimated number of neutrino interactions as obtained using two different event generators, Sibyll 2.3d and DPMJET 3.2017, for FPF experiments located 620 m downstream of the ATLAS IP at the HL-LHC with 14 TeV  $pp$  collisions and an integrated luminosity of  $\mathcal{L} = 3 \text{ ab}^{-1}$ .

Short white paper 2109.10905

# DIS $\nu/\bar{\nu}$ cross sections

- New energy regimes
- Test lepton universality

Estimated statistical uncertainties for CC cross sections, 150/fb and 3/ab.  
Systematic uncertainties under study.

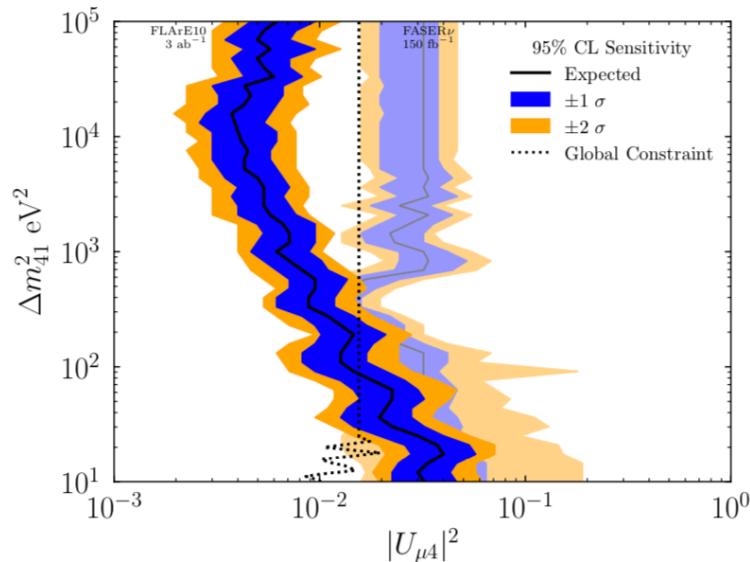


Short white paper 2109.10905

# BSM physics – oscillation into sterile neutrinos

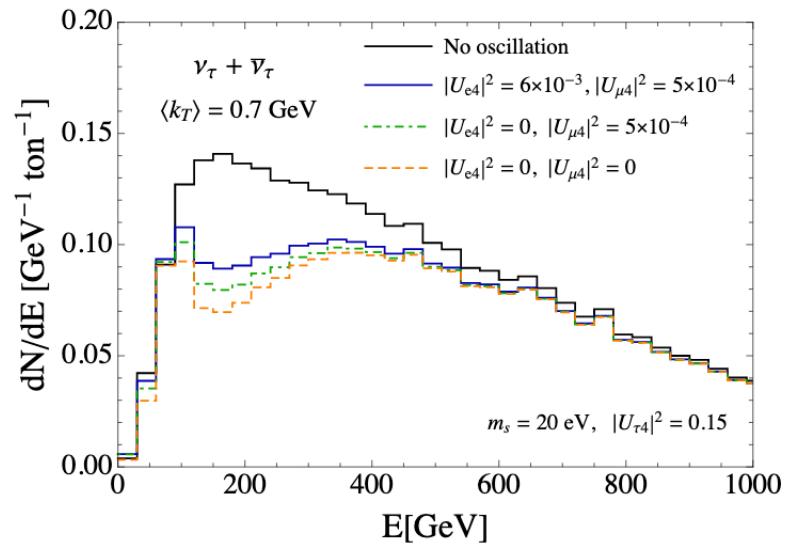
Disappearance, spectral distortions

FASER $\nu$  Run 3 @150/fb, FLARE10 @3/ab



Short white paper 2109.10905

$\eta_\nu > 6.9, 3/\text{ab}$

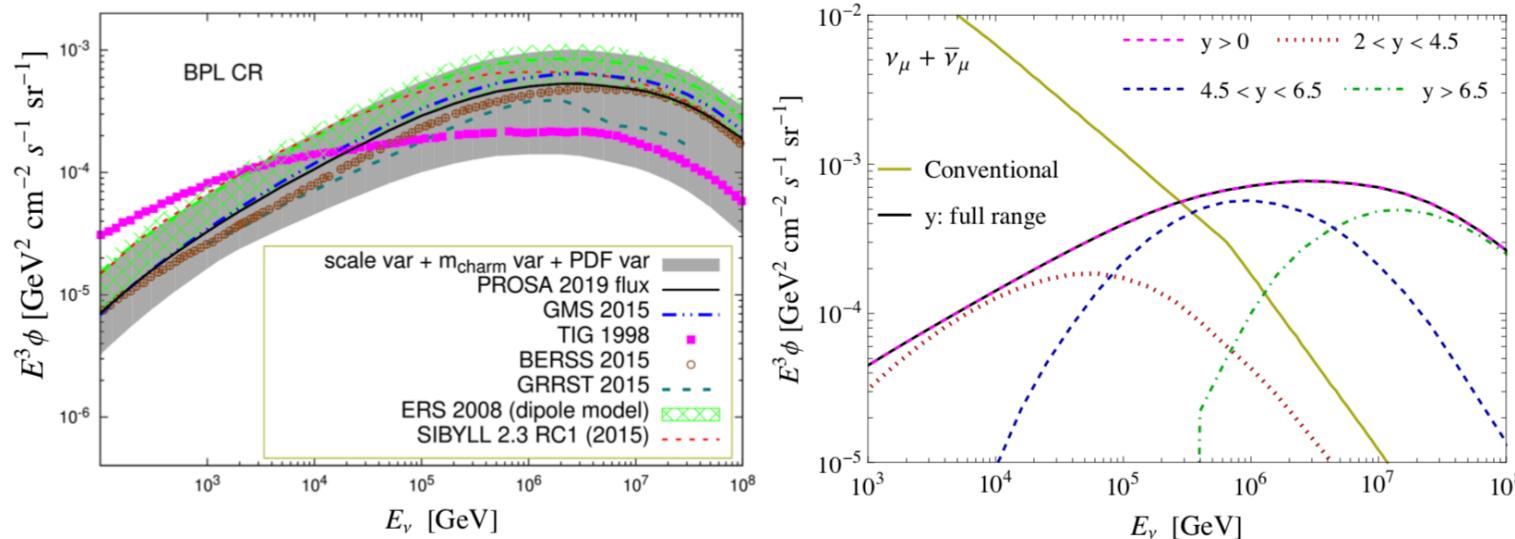


Bai et al., JHEP 06 (2020) 032

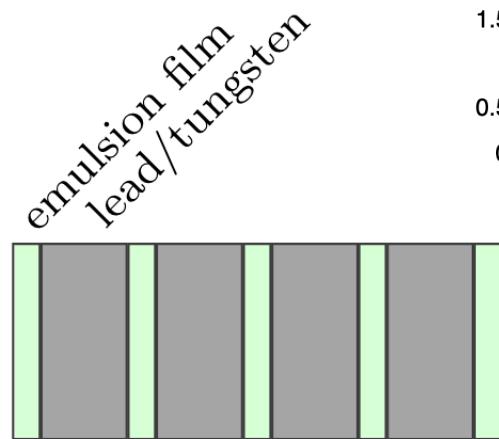
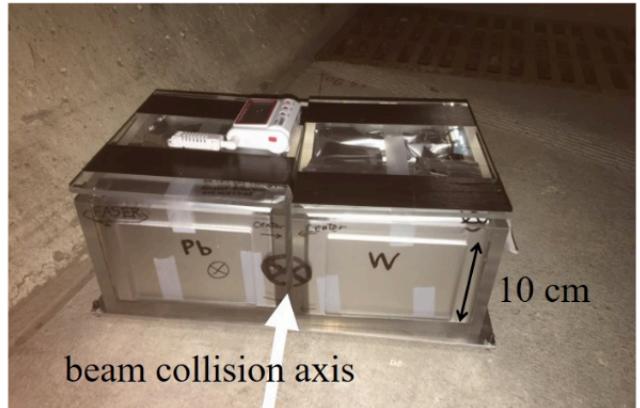
# Astroparticle physics

Natural connection to the prompt atmospheric neutrino flux (neutrinos from charm) and hadronic physics in cosmic ray air showers.

Neutrinos from charm:



# So far – the FASER suitcase

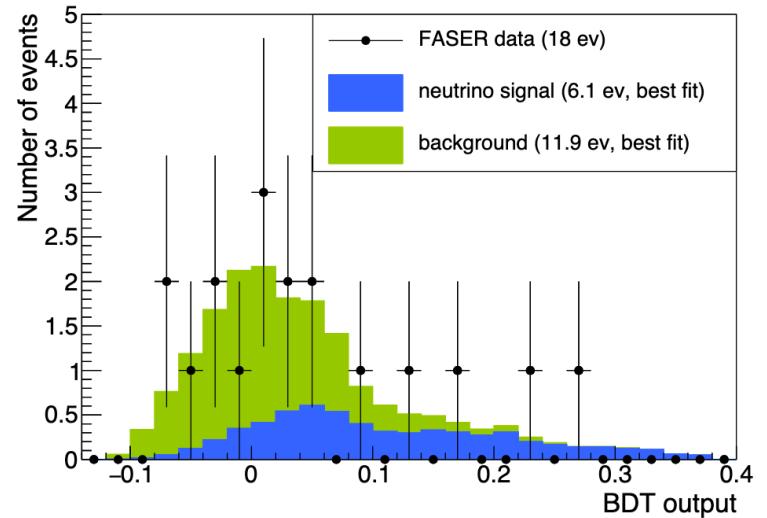


Pilot emulsion detector (spare parts of NA65/DsTau)  
29 kg with 2018 4-week run with 12.2/fb at 13 TeV.

FASER Collab, Abreu, Phys. Rev. D 104 (2021) 091101, 2105.06197

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. . . Beginning of the era of neutrino measurements in the LHC forward region.

# Summary

- ❑ FASER, SND@LHC start soon. Tens to a hundred  $\nu_\tau + \bar{\nu}_\tau$  events depending on detector and production model. Hundreds to thousands for  $\nu_e + \bar{\nu}_e$  and more for  $\nu_\mu + \bar{\nu}_\mu$  events.
- ❑ Lots of work underway to better understand the forward physics environment.
- ❑ There are many opportunities that combine neutrino physics, QCD, BSM physics and astroparticle physics in Forward Physics Facility experiments.
  - PDFs in neutrino production and interactions, e.g.: "extreme" x values, role of small x resummation, nuclear PDFs in neutrino scattering, strange distribution.
  - Tests of the SM & searches for BSM physics, e.g.: lepton flavor universality, oscillations to sterile neutrinos, NSI's, neutrino magnetic moments.
  - Connections to astroparticle physics, e.g.: prompt lepton flux, hadronic interaction models, forward particle production at high energies.