STATUS OF NEGATIVE COUPLING MODIFIERS FOR EXTENDED HIGGS SECTORS

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INTRODUCTION

- The LHC shows that corrections from new physics appear to be small.
- This could be an artifact of the way we are accessing information. Lose information on the sign of couplings!
- In this work we focus on the ratio of coupling modifier

$$\lambda_{WZ} = \frac{\kappa_W}{\kappa_Z}.$$

- This observable probes custodial violation on the Higgs sector.
- Current CMS fits* gives a preference for **negative** λ_{WZ} .



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CUSTODIAL SYMMETRY IN BSM

- Custodial symmetry is an accidental global SU(2) × SU(2) symmetry of the Higgs sector.
- In the SM we have the $\lambda_{WZ} = 1$ and there is also the ρ parameter: $\rho = \frac{m_W^2}{c_W^2 m_Z^2}$.
- Naively one would think that both observables are connected in any model, which is not true.
- ρ is connected to the vacuum while λ_{WZ} is connected to the potential.
- We can have custodial symmetric vacuum with custodial violating potentials (Accidentally custodial models)!

CUSTODIAL SYMMETRY IN BSM

- If we assume that the vacuum preserves the symmetry, the custodial violation inside λ_{WZ} comes only from the diagonalization of the different multiplets:
- $h = R_0 \varphi_0^R + R_1 \psi_1^R + R_2 \psi_2^R + \cdots$
- The kappas have the form:
- $\kappa_V^h = R_0 \kappa_V^{doublet} + R_1 \kappa_V^{multiplet 1} + R_2 \kappa_V^{multiplet 2} + \cdots$
- If we want to explore if a region is excluded, we do not need to resolve the source of \vec{R} in terms of the potential, we can treat them as random variables and independent of the vevs!

ACCIDENTALLY CUSTODIAL TRIPLET

- The simplest model* that can generate negative λ_{WZ} involves two triplets: We have the usual Higgs doublet (φ⁺, φ⁰) with Y = 1, a complex triplet (χ⁺⁺, χ⁺, χ⁰) with Y = 2 and a real triplet (ξ⁺, ξ⁰, ξ⁻) with Y = 0.
- The EW symmetry breaking has the form:

$$\varphi^{0} = \frac{\nu_{\varphi}}{\sqrt{2}} + \frac{1}{\sqrt{2}}(\varphi^{0}_{R} + i\varphi^{0}_{I}), \xi^{0} = \nu_{\xi} + \xi^{0}_{R}, \chi^{0} = \nu_{\chi} + \frac{1}{\sqrt{2}}(\chi^{0}_{R} + i\chi^{0}_{I}), \xi^{0} = \nu_{\chi} + \xi^{0}_{R}, \chi^{0} = \nu_{\chi} + \frac{1}{\sqrt{2}}(\chi^{0}_{R} + i\chi^{0}_{I}), \xi^{0} = \nu_{\chi} + \xi^{0}_{R}, \chi^{0} = \nu_{\chi} + \frac{1}{\sqrt{2}}(\chi^{0}_{R} + i\chi^{0}_{I}), \xi^{0} = \nu_{\chi} + \xi^{0}_{R}, \chi^{0} = \nu_{\chi} + \frac{1}{\sqrt{2}}(\chi^{0}_{R} + i\chi^{0}_{I}), \xi^{0} = \nu_{\chi} + \xi^{0}_{R}, \chi^{0} = \nu_{\chi} + \frac{1}{\sqrt{2}}(\chi^{0}_{R} + i\chi^{0}_{I}), \xi^{0} = \nu_{\chi} + \xi^{0}_{R}, \chi^{0} = \nu_{\chi} + \frac{1}{\sqrt{2}}(\chi^{0}_{R} + i\chi^{0}_{I}), \xi^{0} = \nu_{\chi} + \xi^{0}_{R}, \chi^{0} = \nu_{\chi} + \frac{1}{\sqrt{2}}(\chi^{0}_{R} + i\chi^{0}_{I}), \xi^{0} = \nu_{\chi} + \xi^{0}_{R}, \chi^{0} = \nu_{\chi} + \frac{1}{\sqrt{2}}(\chi^{0}_{R} + i\chi^{0}_{I}), \xi^{0} = \nu_{\chi} + \frac{1}{\sqrt{2}}(\chi^{0}_{R} + i\chi^{0}), \xi^{0} = \nu_{\chi} + \frac{1}{\sqrt{$$

• The ρ parameter in this model is:

$$p = \frac{m_W^2}{c_W^2 m_Z^2} = \frac{\nu_{\varphi}^2 + 4\nu_{\chi}^2 + 4\nu_{\xi}^2}{\nu_{\varphi}^2 + 8\nu_{\chi}^2}$$

• $v_{\xi} = v_{\chi} \rightarrow$ vacuum custodially symmetric

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ACCIDENTALLY CUSTODIAL TRIPLET

- $h = R_0 \varphi_R^0 + R_1 \chi_R^0 + R_2 \xi_R^0$
- $\lambda_{WZ} = \frac{2\sqrt{2}\nu_{\chi}R_1 + 4\nu_{\chi}R_2 + \nu_{\varphi}R_0}{4\sqrt{2}\nu_{\chi}R_1 + \nu_{\varphi}R_0}$
- Assume $\lambda_{WZ} = -1$, write the other observables:

$$\lambda_{fZ} = rac{\kappa_f}{\kappa_Z}$$
 and $\kappa_{fZ} = rac{\kappa_f \kappa_Z}{\kappa_h}$

• Then, solve for \vec{R} , if we assume $\lambda_{fZ} = \pm 1$ we get:

$$\kappa_{fZ} = R_0 \frac{\nu_{\varphi}}{\nu}$$
 for $\lambda_{fZ} = 1$
 $\kappa_{fZ} = -1.39 R_0 \frac{\nu_{\varphi}}{\nu}$ for $\lambda_{fZ} = -1$

GENERALIZATION FOR DIFFERENT MULTIPELTS

- Any model with a custodial limit can contribute to λ_{WZ} while avoiding the ρ parameter.
- The particle content of those models can be constructed from the generalized Georgi-Machacek models*, which we can break down into SU(2) x U(1) quantum numbers:
- AC triplets: one field with (1,2) and one with (1,0)
- AC quartets: one field with (3/2,3) and one with (3/2,1)
- AC pentets: one field with (2,4), one with (2,2) and one with (2,0)
- AC sextets: one field with (5/2,5), one with (5/2,3) and one with (5/2,1)

GENERALIZATION FOR DIFFERENT MULTIPELTS

- Perturbative unitarity constraints the number of allowed models, assuming one doublet we can have 4487 possibilities.
- From these combinations, we can only have at most one AC sextet, four AC pentets, 23 AC quartets, or 145 AC triplets.
- The models that we study here are AC triplet, AC quadruplet, AC pentet, AC sexplet, 2 AC triplets, AC pentet + AC sexplet, and two AC pentet + AC sexplet. Additionally, we also explore the case with general vev's for each of these models.
- We also explore the limiting cases of four AC pentets, 23 AC quartets, and 145 AC triplets.

SCANS



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SCANS



More multiplets is not better. Same behavior as the largest multiplet.

SCANS



General vevs also does not help. ρ parameter constraint is too strong, even at 5 σ . Same behaviour as the situation with only the largest multiplet

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CONCLUSION

- In this work, we studied the current status of negative coupling modifiers in the extended Higgs sector, with the focus on the observable λ_{WZ} .
- We present the analysis for the simplest case of AC triplets, and then we show how to generalize the procedure to different multiplets.
- The possibility of exploring this wide range of models lies in the fact that the coupling modifiers, in the end, depend only on the diagonalization matrix and the vevs.
- Under the analysis done we can see that all the models with one or more AC multiplets studied here are excluded under ATLAS results at 99.7%CL (99.5%CL for GEN sextet).
- What does this mean for negative λ_{WZ} ?
- We can then say that this region of parameter space is heavily disfavored for any weakly coupling extended scalar sector.
- In contrast, if the measured value for CMS stays to be negative and different experiments confirm this, we would not be able to describe the new physics using the current methods.
- It may be that new physics is hiding in plain sight, after all, only future experiments can tell.

BACKUP



BACKUP

Observables used:

$$\kappa_{fZ} = \frac{\kappa_f \kappa_Z}{\kappa_h}$$

$$\lambda_{WZ} = \frac{\kappa_W}{\kappa_Z}$$

$$\lambda_{fZ} = \frac{\kappa_f}{\kappa_Z}$$

$$\kappa_h = \sqrt{0.75\kappa_f^2 + 0.22\kappa_W^2 + 0.03\kappa_Z^2}$$

ATLAS FIT: $(\lambda_{fZ}, \kappa_{fZ}) = (0.99, 0.98)$

$$COV = \begin{pmatrix} 0.0093 & -0.00054 \\ -0.00054 & 0.0020 \end{pmatrix}$$

CMS FIT (UNCORRELATED): $\kappa_{fZ} = 1.03 \pm 0.09$ $\lambda_{fZ} = 1.10 \pm 0.11$ $\lambda_{WZ} = -1.13^{+0.10}_{-0.11}$

BACKUP

Coupling modifiers for different multiplets:

DOUBLET:

$$\kappa_f^{(\frac{1}{2},1)} = \frac{\nu_{(\frac{1}{2},1)}}{\nu}, \, \kappa_W^{(\frac{1}{2},1)} = \frac{\nu_{(\frac{1}{2},1)}}{\nu} \, \kappa_Z^{(\frac{1}{2},1)} = \frac{\nu_{(\frac{1}{2},1)}}{\nu}$$

AC TRIPLET:

$$\kappa_f^{(1,2)} = 0$$
, $\kappa_W^{(1,2)} = 2\sqrt{2} \frac{\nu_{(1,2)}}{\nu} \kappa_Z^{(1,2)} = 4\sqrt{2} \frac{\nu_{(1,2)}}{\nu}$
 $\kappa_f^{(1,0)} = 0$, $\kappa_W^{(1,0)} = 4 \frac{\nu_{(1,0)}}{\nu} \kappa_Z^{(1,0)} = 0$

AC QUARTET:

$$\kappa_{f}^{(\frac{3}{2},3)} = 0, \ \kappa_{W}^{(\frac{3}{2},3)} = 3\sqrt{2} \frac{\nu_{(\frac{3}{2},3)}}{\nu} \kappa_{Z}^{(\frac{3}{2},3)} = 9\sqrt{2} \frac{\nu_{(\frac{3}{2},3)}}{\nu} \kappa_{f}^{(\frac{3}{2},1)} = 0, \ \kappa_{W}^{(\frac{3}{2},1)} = 7\sqrt{2} \frac{\nu_{(\frac{3}{2},1)}}{\nu} \kappa_{Z}^{(\frac{3}{2},1)} = \sqrt{2} \frac{\nu_{(\frac{3}{2},3)}}{\nu}$$

AC PENTET:

$$\kappa_f^{(2,4)} = 0, \, \kappa_W^{(2,4)} = 4\sqrt{2} \frac{\nu_{(2,4)}}{\nu} \kappa_Z^{(2,4)} = 16\sqrt{2} \frac{\nu_{(2,4)}}{\nu} \kappa_f^{(2,2)} = 0, \, \kappa_W^{(2,2)} = 10\sqrt{2} \frac{\nu_{(2,2)}}{\nu} \kappa_Z^{(2,2)} = 4\sqrt{2} \frac{\nu_{(2,2)}}{\nu} \kappa_f^{(2,0)} = 0, \, \kappa_W^{(2,0)} = 12 \frac{\nu_{(2,0)}}{\nu} \kappa_Z^{(2,0)} = 0$$

AC SEXTET:

$$\begin{split} \kappa_{f}^{(\frac{5}{2},5)} &= 0 \ , \ \kappa_{W}^{(\frac{5}{2},5)} = 5\sqrt{2} \frac{\nu_{(\frac{5}{2},5)}}{\nu} \kappa_{Z}^{(\frac{5}{2},5)} = 25\sqrt{2} \frac{\nu_{(\frac{5}{2},5)}}{\nu} \\ \kappa_{f}^{(\frac{5}{2},3)} &= 0 \ , \ \kappa_{W}^{(\frac{5}{2},3)} = 13\sqrt{2} \frac{\nu_{(\frac{5}{2},3)}}{\nu} \kappa_{Z}^{(\frac{5}{2},3)} = 9\sqrt{2} \frac{\nu_{(\frac{5}{2},3)}}{\nu} \\ \kappa_{f}^{(\frac{5}{2},1)} &= 0 \ , \ \kappa_{W}^{(\frac{5}{2},1)} = 17 \frac{\nu_{(\frac{5}{2},1)}}{\nu} \kappa_{Z}^{(\frac{5}{2},1)} = \sqrt{2} \frac{\nu_{(\frac{5}{2},3)}}{\nu} \end{split}$$

BACKUP

N copies limit:

$$\begin{split} \kappa_{V}^{h} &= R_{0} \kappa_{V}^{doublet} + \left(R_{1} \kappa_{V}^{mult \, 1} + R_{2} \kappa_{V}^{mult \, 2} \right) + \left(R_{3} \kappa_{V}^{mult \, 1} + R_{4} \kappa_{V}^{mult \, 2} \right) + \dots = \\ &= R_{0} \kappa_{V}^{doublet} + \left(R_{1} + R_{3} + \dots \right) \kappa_{V}^{mult \, 1} + \left(R_{2} + R_{4} + \dots \right) \kappa_{V}^{mult \, 2} = \\ &= R_{0} \kappa_{V}^{doublet} + \widetilde{R_{1}} \kappa_{V}^{mult \, 1} + R_{2} \widetilde{\kappa_{V}^{mult \, 2}} \end{split}$$

Cauchy inequality:

$$R_0^2 + \frac{\widetilde{R_1}^2}{N} + \frac{\widetilde{R_2}^2}{N} \le 1$$