





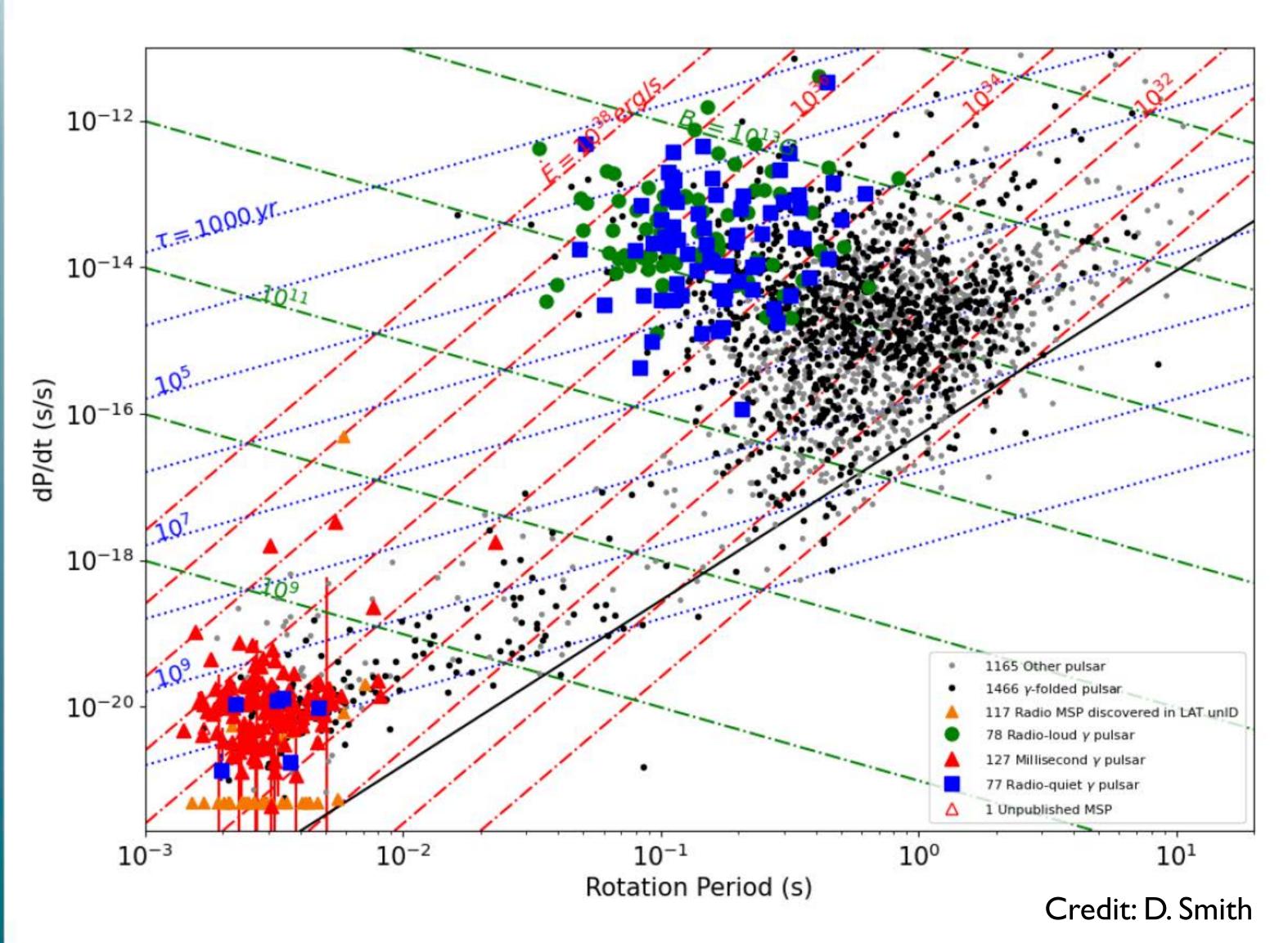
Pablo Saz Parkinson\*

Di Li, Pei Wang, Xian Hou, Colin Clark, David Smith, Matthew Kerr, and Massimiliano Razzano on behalf of the FAST team and the Fermi-LAT Collaboration

\*sazpark@ucsc.edu
UNIVERSITY OF CALIFORNIA



## Fermi LAT gamma-ray pulsars





#### **Summary Statistics**

Total number of pulsars: 278

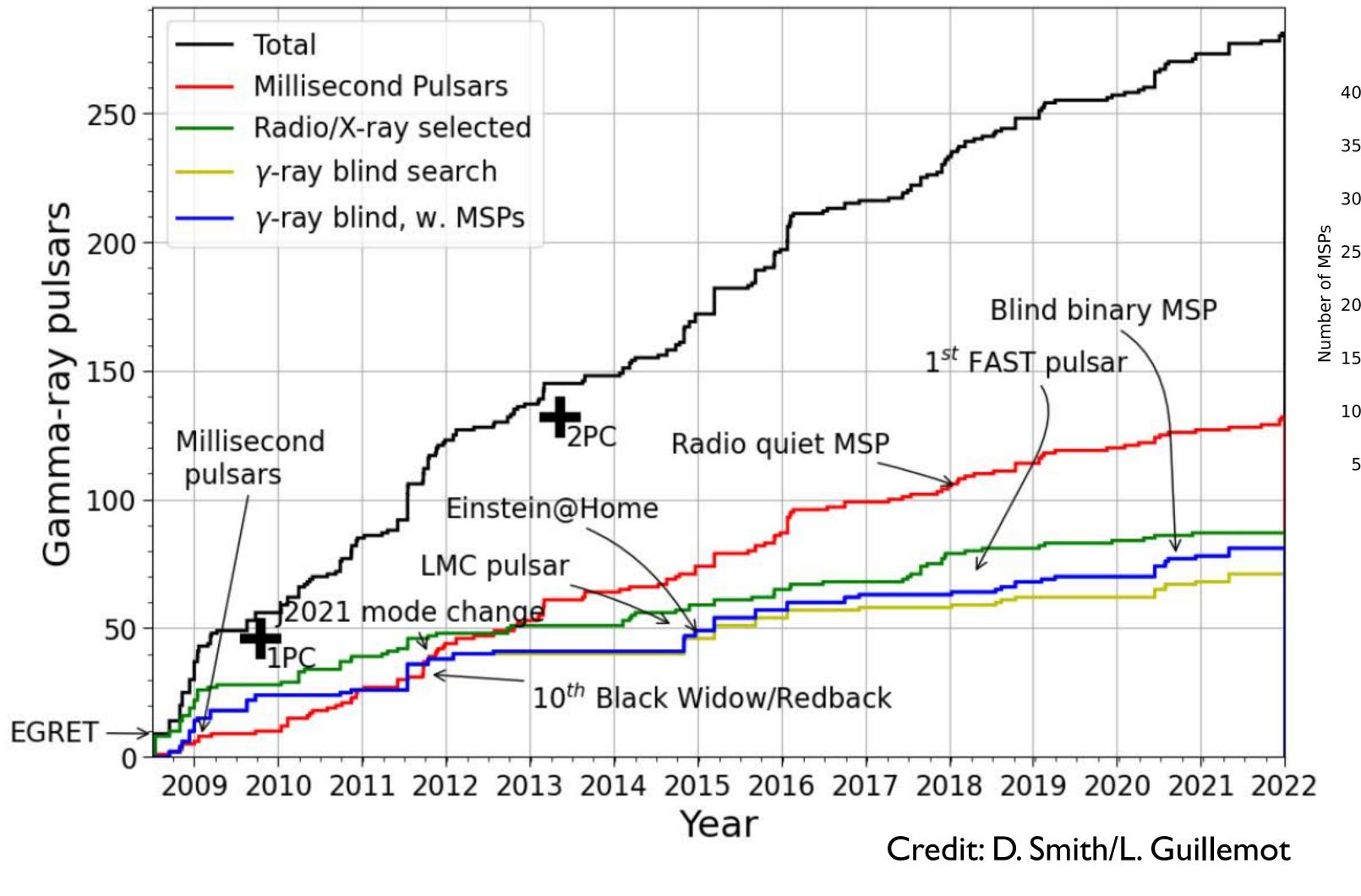
- Young, radio + X-ray selected: 73
- Young, gamma selected: 71
- MSP, radio selected: 118
- MSP, gamma selected: 10
- Found in searches of LAT unassociated sources: 145
- EGRET/COMPTEL pulsars: 7

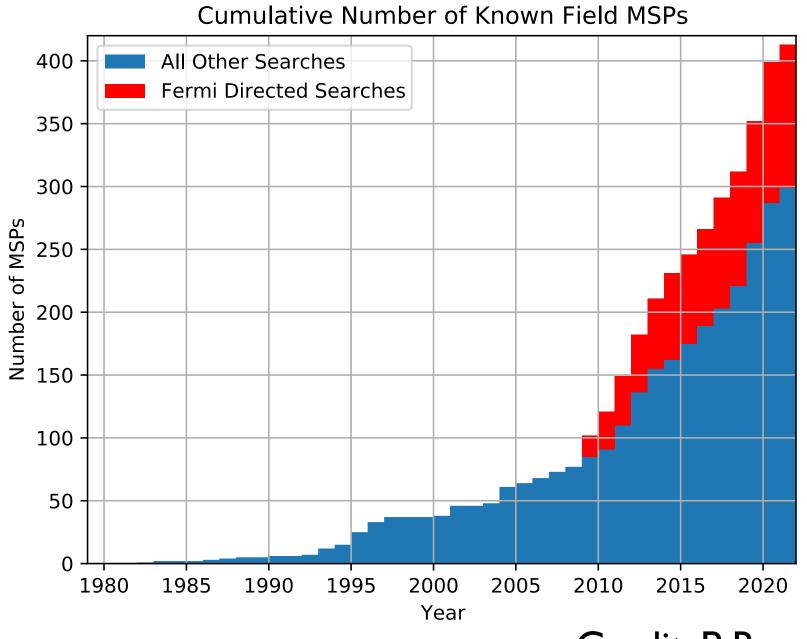
https://confluence.slac.stanford.edu/x/5Jl6Bg



## Fermi LAT gamma-ray pulsars

#### 3rd Fermi LAT Pulsar Catalog





Credit: P. Ray



#### FAST-Fermi-LAT Collaboration

1. LAT Team provides source candidates and ephemerides of known pulsars, and FAST conducts deep searches for pulsars (or radio pulsations, for known pulsars)

2. FAST team provides details on newly discovered pulsars, and the Fermi group searches for their highenergy counterparts





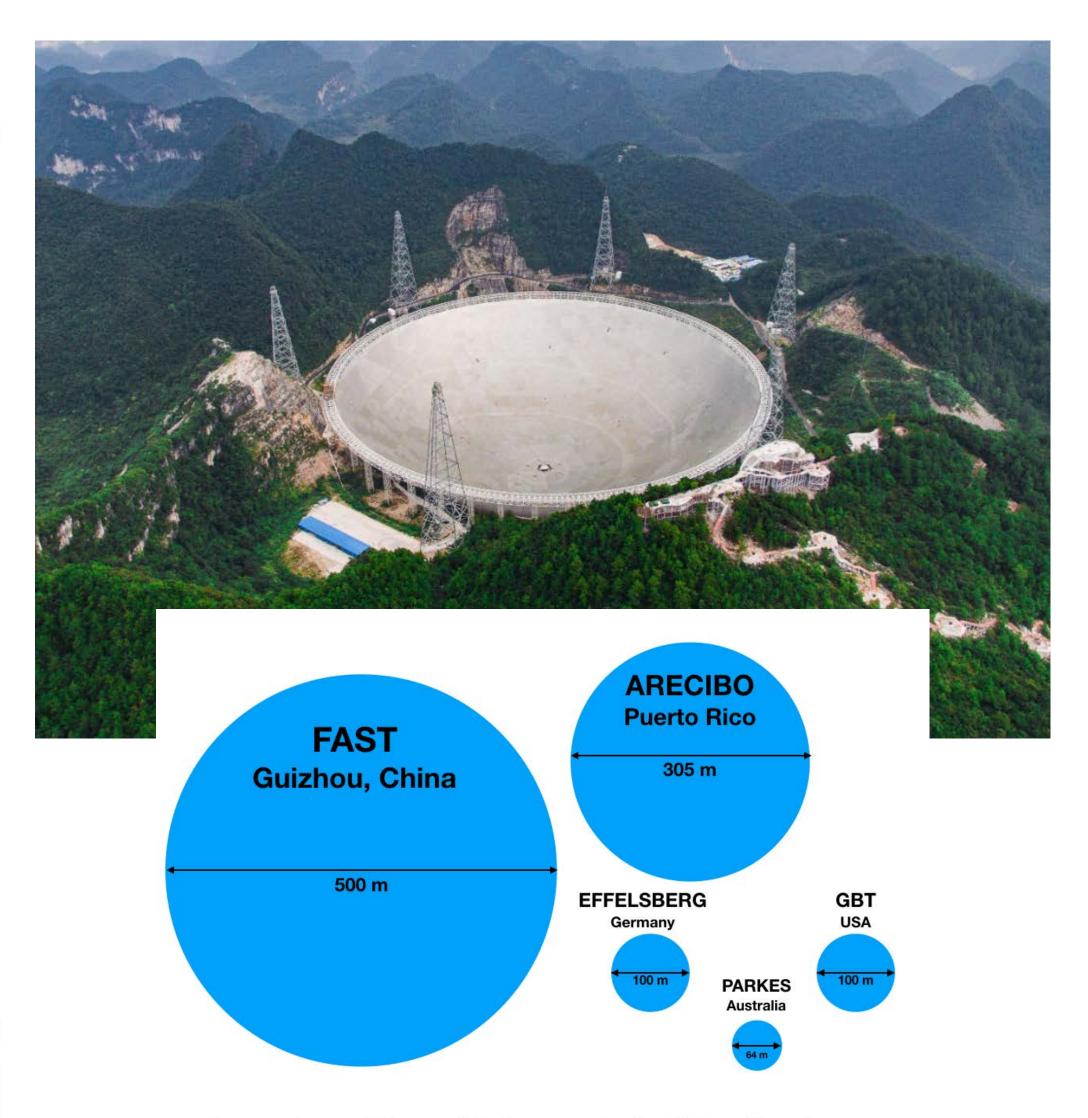




NAOC – SHAO – XAO – FermiLAT signed MoU 2017.12

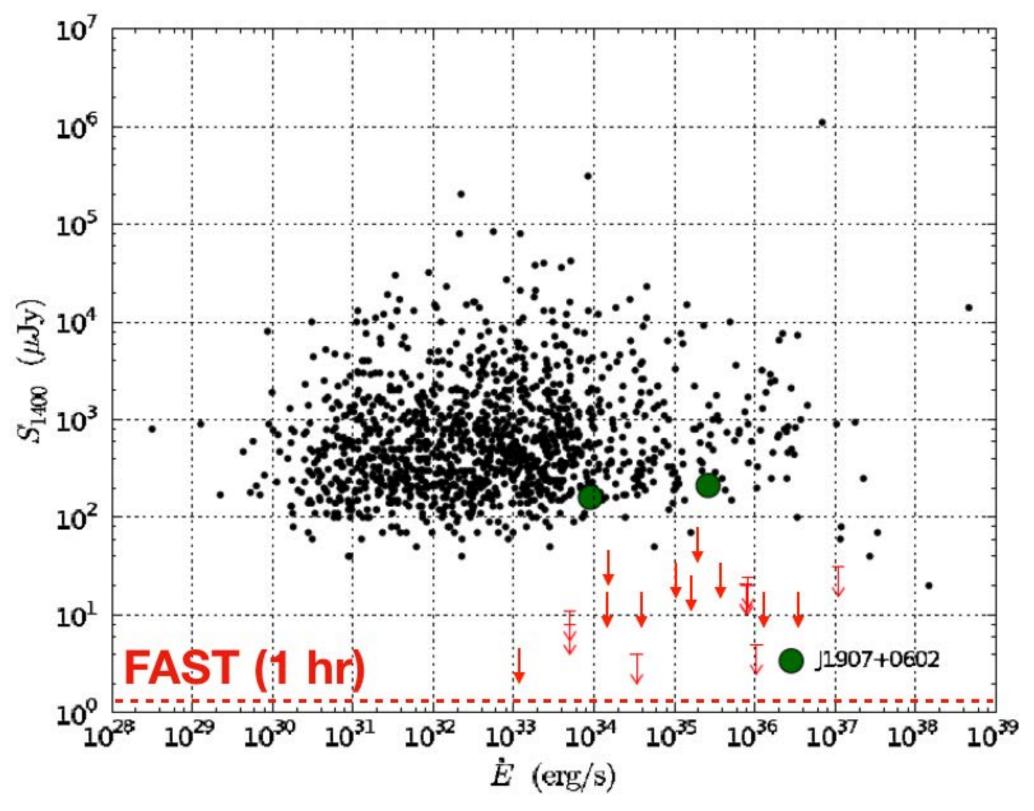


### Radio follow-up searches



Comparison of the world's largest single-dish radio telescopes (All members of the Fermi LAT Pulsar Search Consortium)

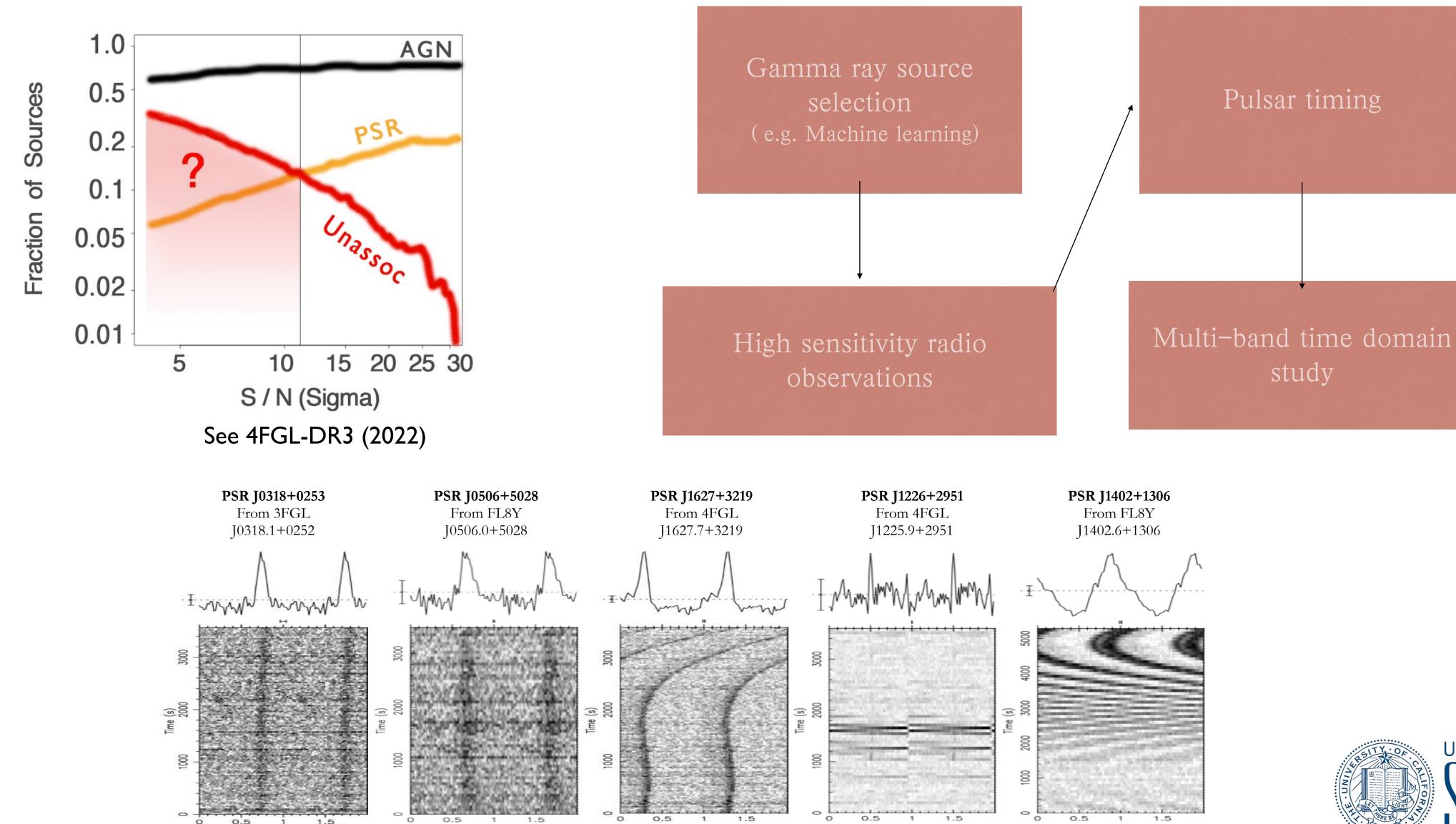
The Pulsar Search Consortium (PSC) was set up to search for radio pulsars in LAT sources/pulsars (Ray et al. 2012)



Adapted from Saz Parkinson et al. 2010



### The Fast-Fermi pipeline



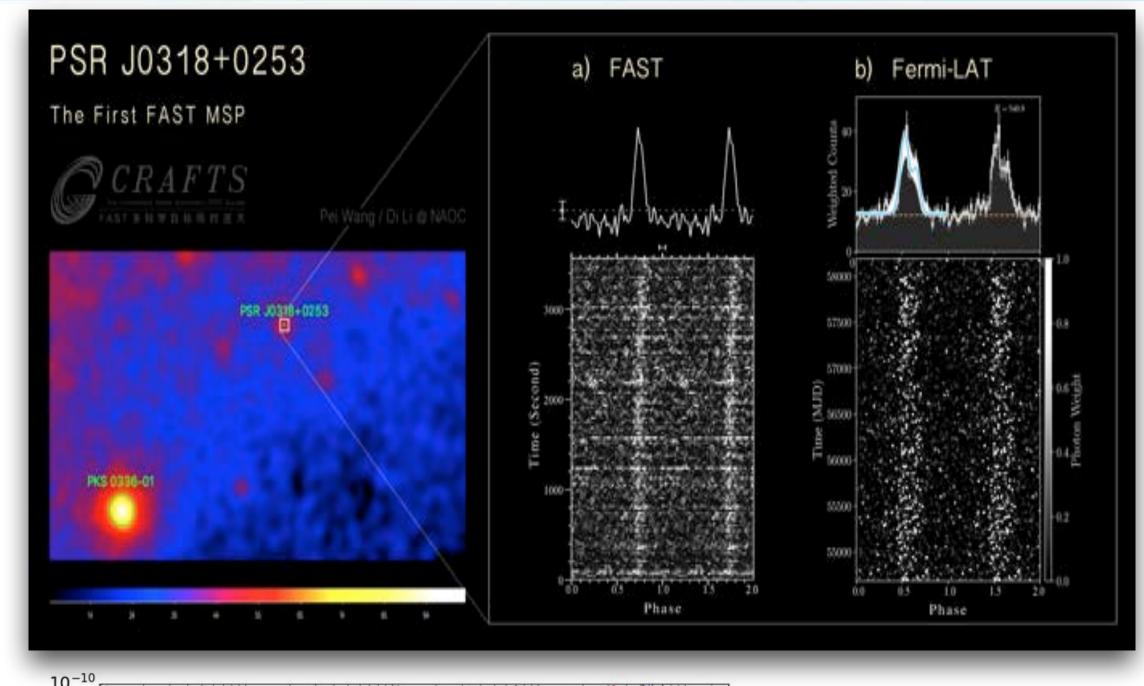
Phase

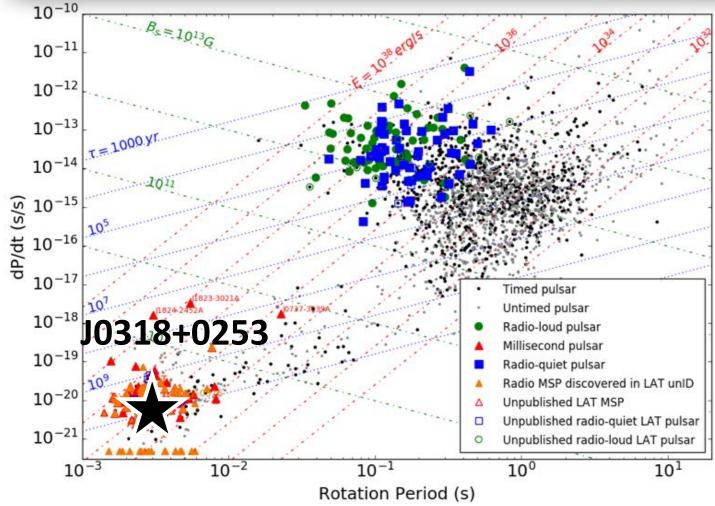
Phase

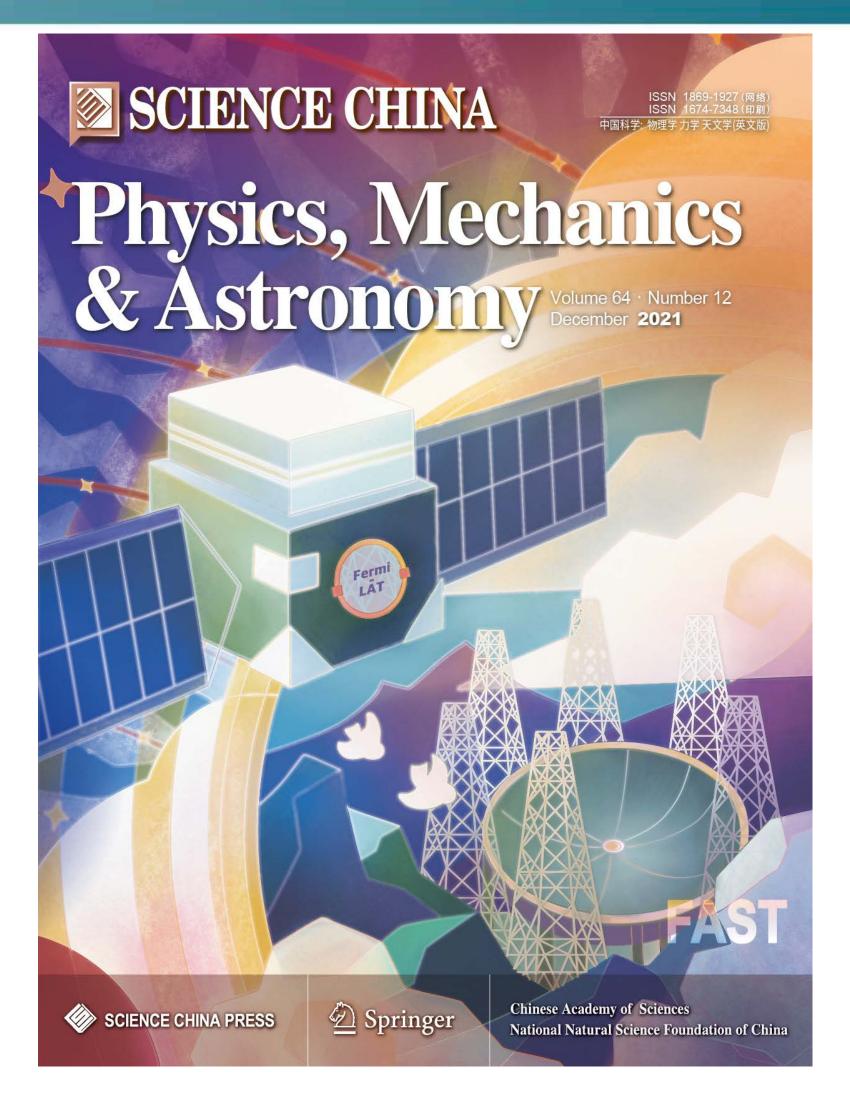


Phase

#### A Fermi/FAST radio-faint MSP







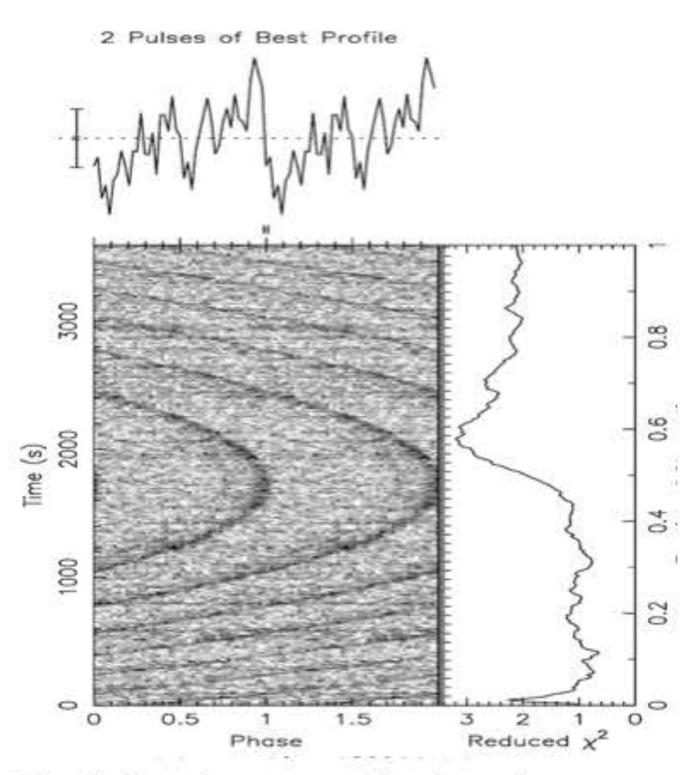
Wang et al., Science China Physics, Mechanics & Astronomy, Volume 64, Issue 12, article id.129562 (December 2021)



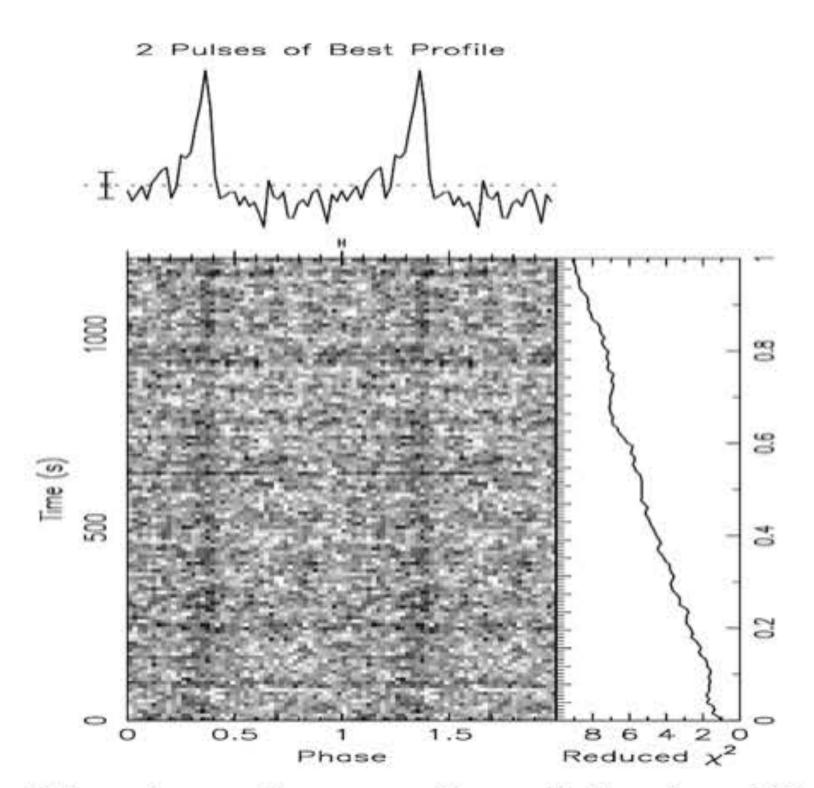
SHANTA CALIFORNIA

### 4FGL J1627.7+3219 aka PSR J1627+3219

Spin period is 2.2 milliseconds, an estimated distance of about 4.4 kpc, and as potentially one of the shortest orbital period binary systems (4.2 hours).



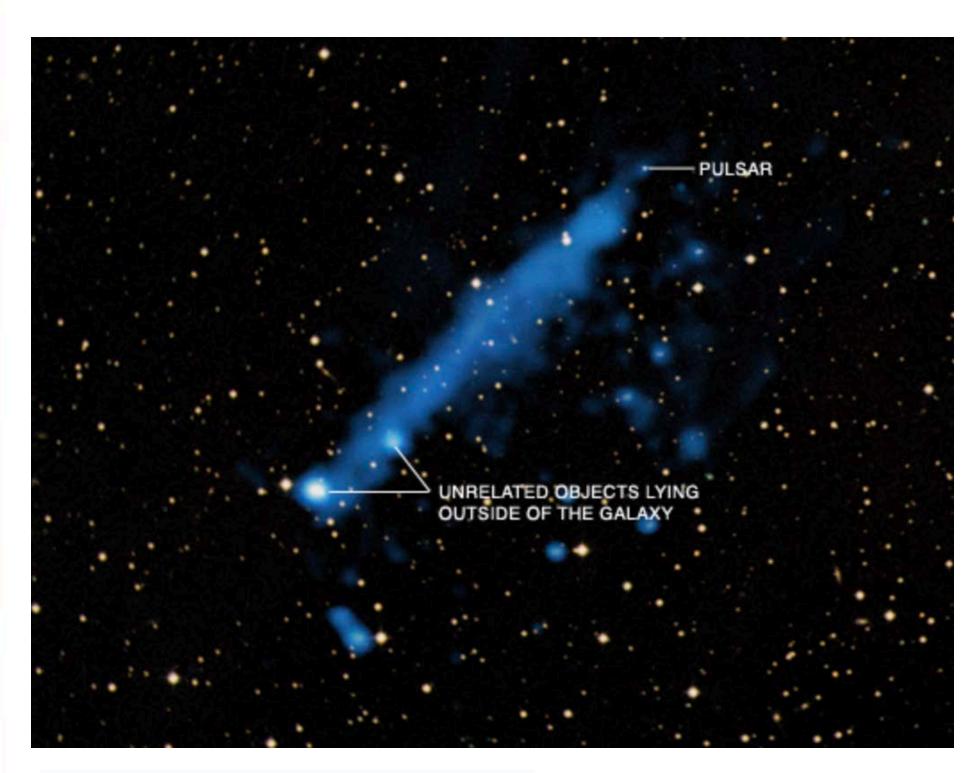
Orbital modulation



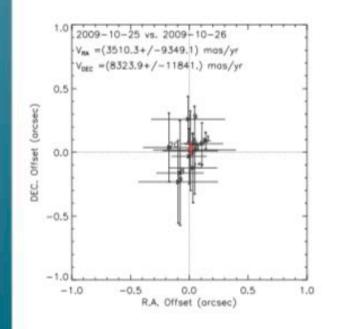
Elimination of orbital effects

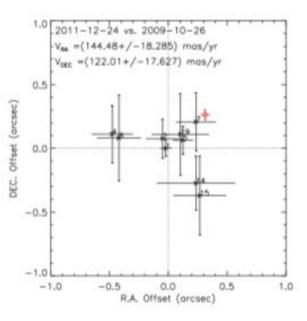


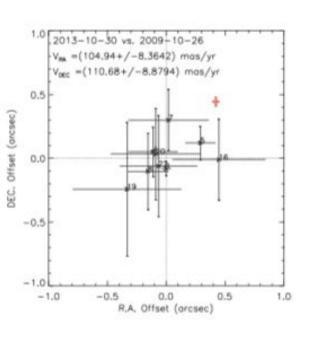
#### PSR J0357+3205: Nearest or fastest?



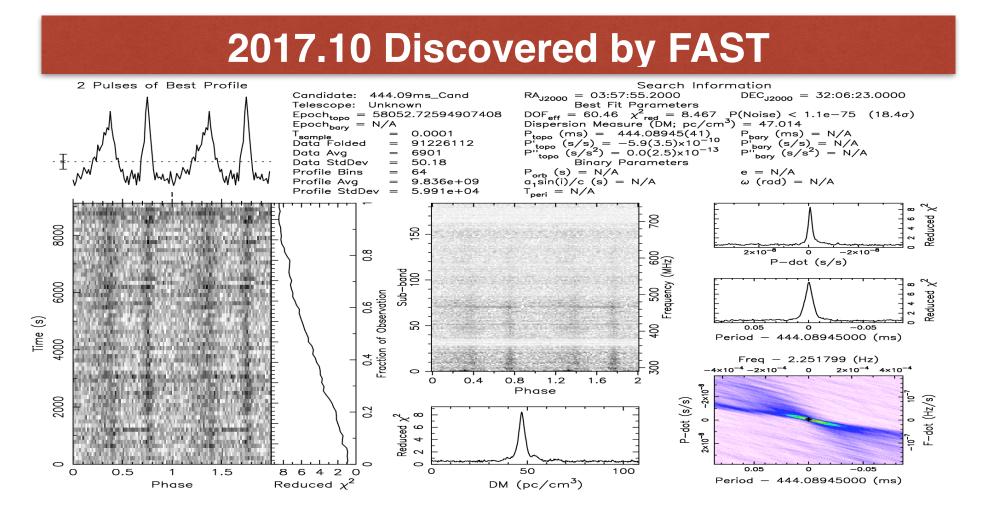
Credit: X-ray: NASA/CXC/IUSS/A.De Luca et al; Optical: DSS



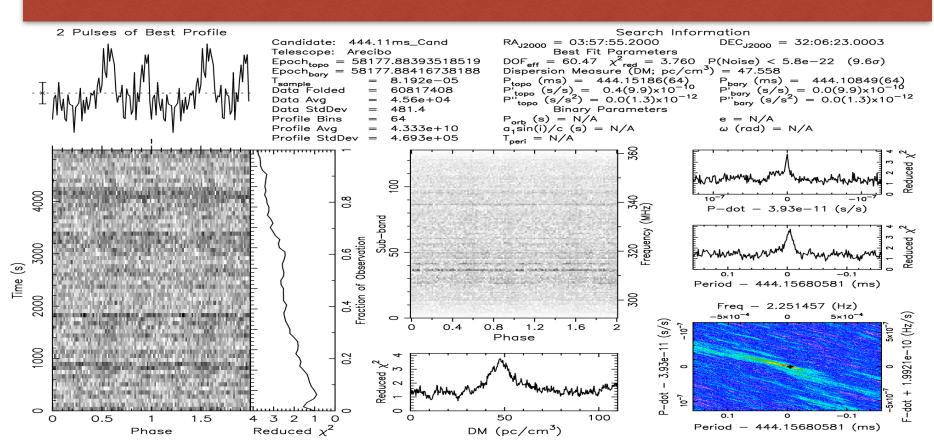




De Luca, A. et al. 2013

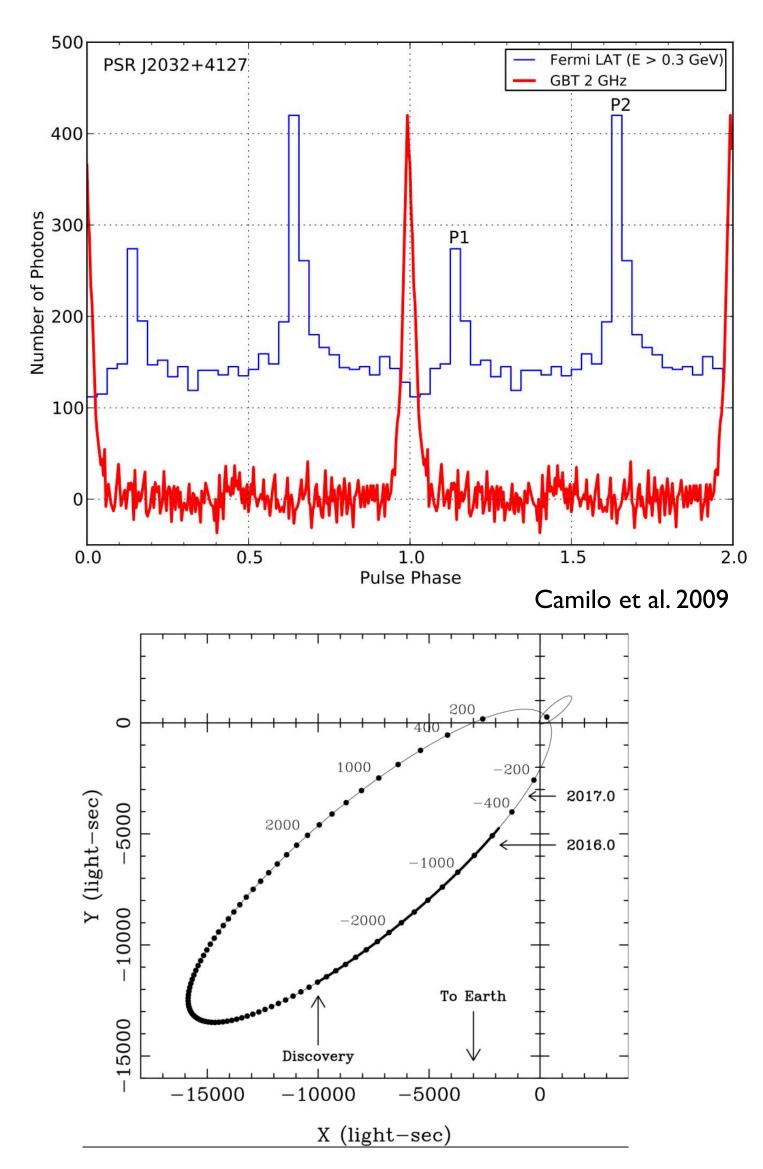


#### 2018.3.1 Confirmed by Arecibo DDT

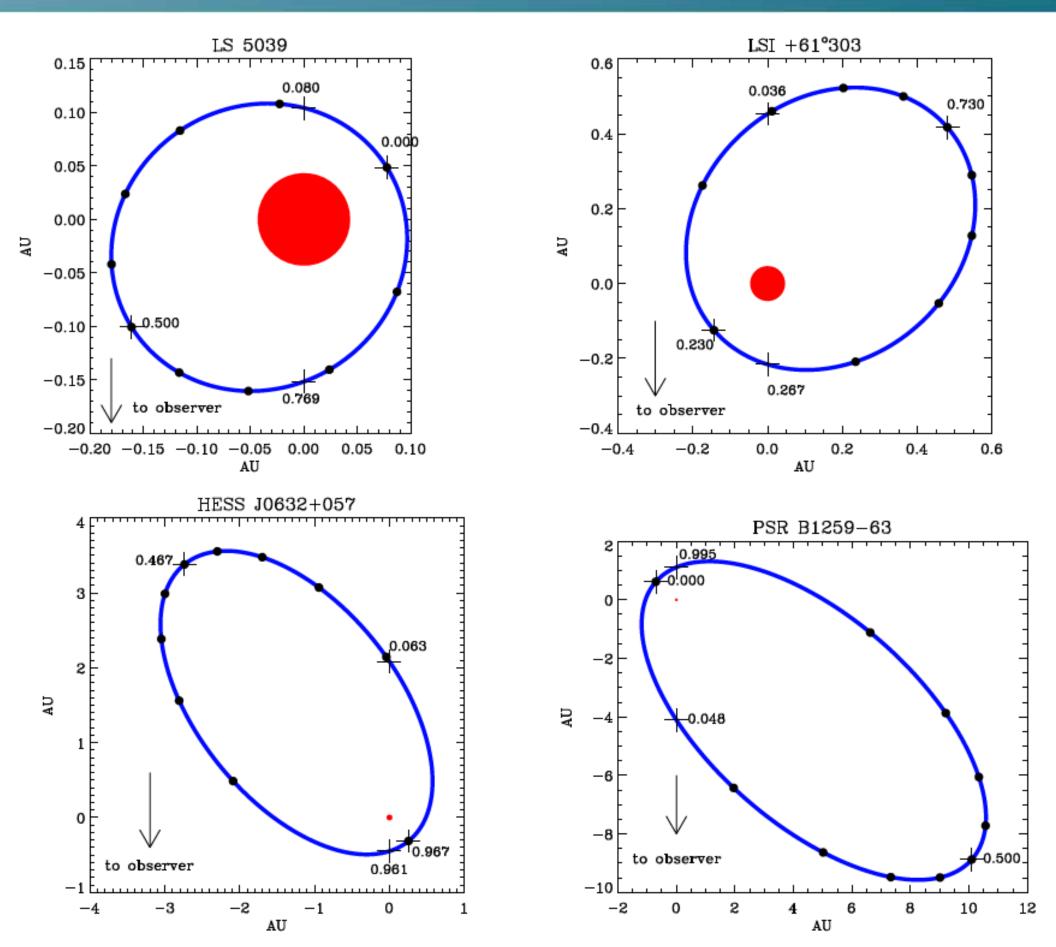




## Young pulsars in gamma-ray binaries?



Lyne et al. 2015



Gamma-ray binaries: typically consist of a massive star in orbit with a compact object, often a pulsar.

See: Nieder et al. 2020



## FAST discovery of a pulsar in LS I + 61 303

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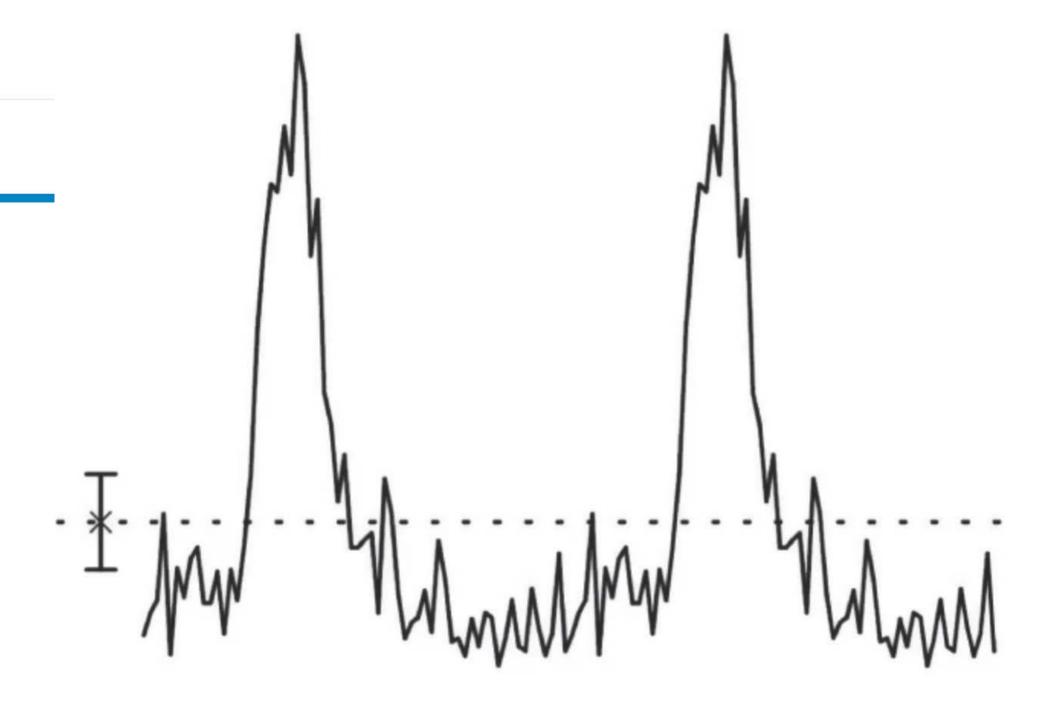
Letter Published: 17 March 2022

# Radio pulsations from a neutron star within the gamma-ray binary LS I +61° 303

Shan-Shan Weng ☑, Lei Qian, Bo-Jun Wang, D. F. Torres ☑, A. Papitto, Peng Jiang, Renxin Xu, Jian Li, Jing-Zhi Yan, Qing-Zhong Liu, Ming-Yu Ge & Qi-Rong Yuan

Nature Astronomy 6, 698–702 (2022) Cite this article

747 Accesses 5 Citations 16 Altmetric Metrics



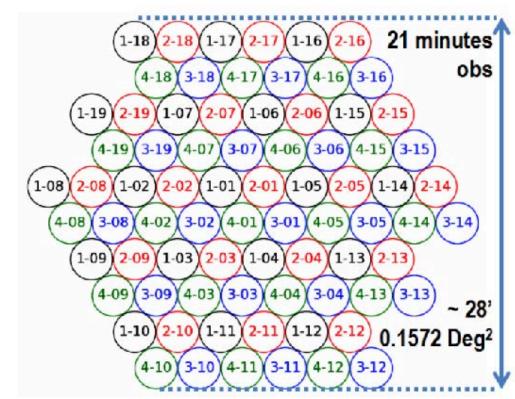
Two cycles of integrated pulse profile with the best-fitted spin period (P=269.15508 ms). The dotted line and the asterisk mark the mean flux level of pulsed emission, and the standard deviation of the pulse profile is plotted with the error bar.



### FAST surveys

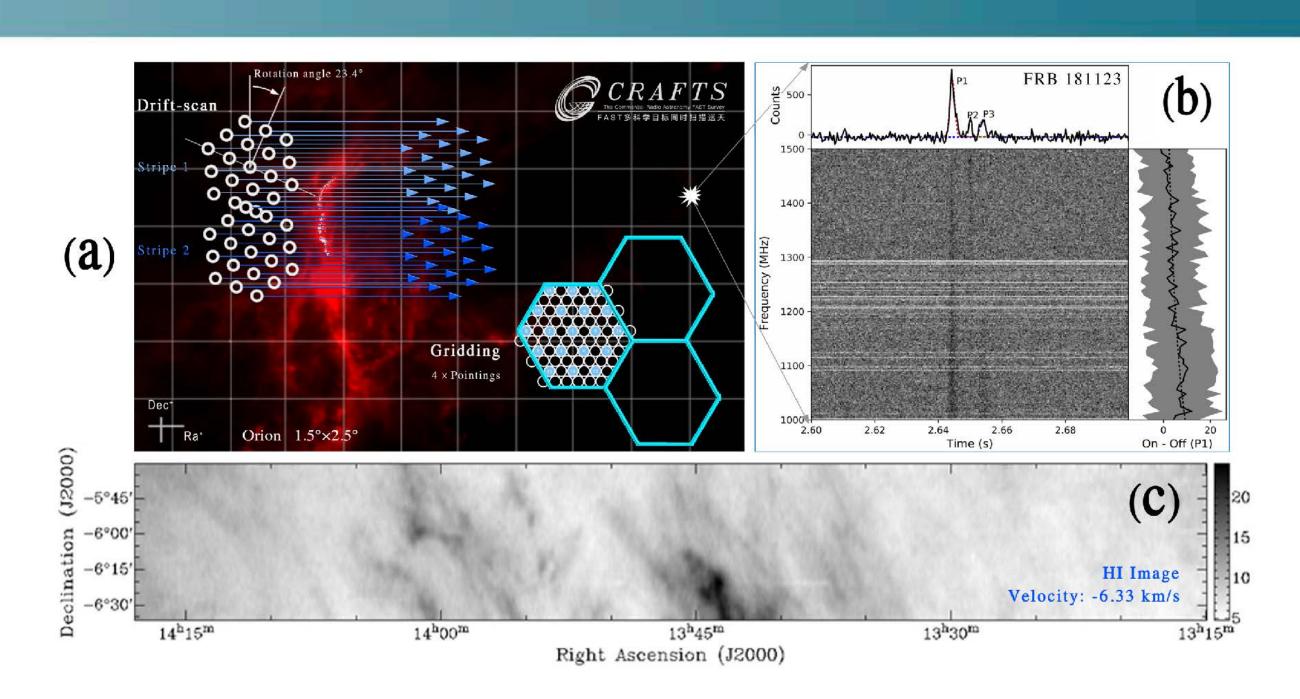


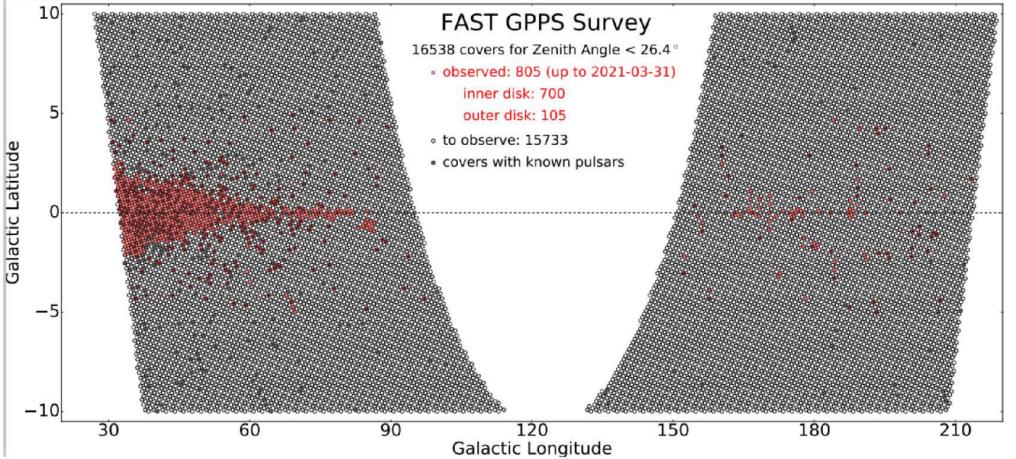
http://groups.bao.ac.cn/ism/CRAFTS/



**GPPS** 

http://zmtt.bao.ac.cn/GPPS/





#### FAST-CRAFTS Pulsar Survey

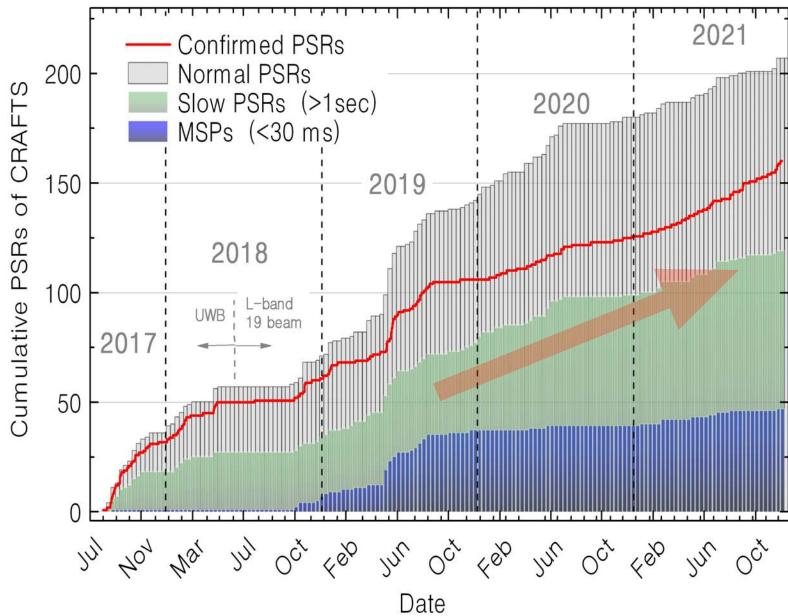
- Only 15% of FAST sky coverage completed
- A large number of high-value sources need to be followed up with observations
- FAST-based, combined with international multi-equipment to accelerate the pulsar timing process

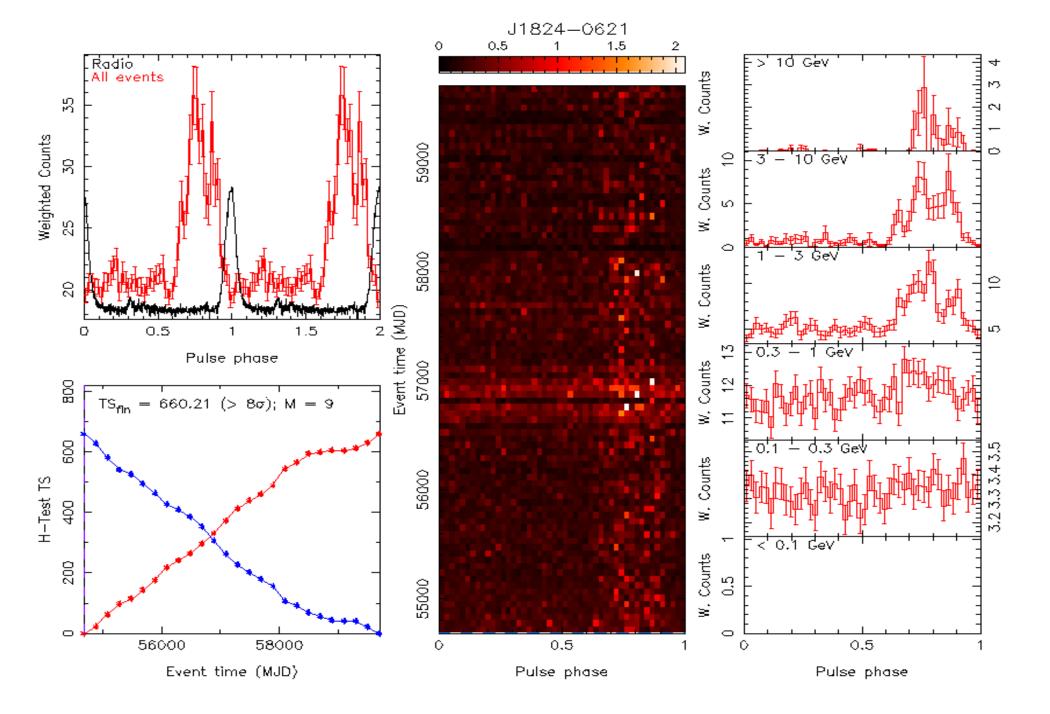
**>40** MSPs

>20 Binary systems (including 1 DNS)

**~30** RRATs

Confirmed 170 pulsar discoveries!





Successful detection of gamma-ray counterparts
PSR J1824-0621









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