





# Hadronic Interactions at High and Low Energies

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23rd MCnet meeting

### This talk

This talk will be a review on the things I did during my PhD. I will present my perspectives on a few phenomena, results of what I did, and talk about some things that I did not have time to do.

- Hadronic rescattering
- Exotic hadrons
- Cosmic rays

#### Outline

#### Hadronic rescattering

Exotic hadrons

Cosmic rays

### How I view rescattering



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### Features of rescattering

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- To some extent it can also give rise to jet quenching.
- Resonances produced in rescattering can be hard to observe. Some resonances can be suppressed.

#### Multiplicities - pp vs. PbPb @ 5.02 TeV

Rescattering is implemented 2 so multiplicity will increase.

*n* processes, but not *n* 2,



"Rule of thumb": Mul ti parton Interactions: pTORef = 2.345 Particle composition is still changed, e.g. by  $p\bar{p}$  + 0 -

### Flow and jet quenching

Phenomena such as collective flow has been observed, and are usually attributed to the formation of a *quark-gluon plasma* (QGP)



These phenomena have also been observed in pA and pp collisions. Open question: can this be explained by the QGP model?

#### Flow - PbPb @ 5.02 TeV



(Data from arXiv:1903.01790)

- Very good description at high multiplicities, where there is more rescattering activity
- Other effects like ropes and shoving should also contribute, so the result with only rescattering should be below data

### Jets /AA - PbPb @ 2.76 TeV

 $I_{AA}$  is the PbPb/pp ratio of associated particle yield per trigger 8 GeV  $< \rho$  , trig < 15 GeV, 4 GeV  $< \rho$  , assoc  $< \rho$  , trig

### Jets /AA - PbPb @ 2.76 TeV



(Data from arXiv:1110.0121)

NB:  $\rho$  spectrum is also modified by other mechanisms. Would be interesting to study in more detail.

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### Resonance formation and signal suppression

I did not had time to study resonance signal suppression, but this would be an interesting thing to study.

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  - Experiments indicate that K signals are suppressed by rescattering, but signals are not. This puts limits on the duration of rescattering [arXi v: 1910. 14419]. How does this compare to rescattering in Pythia?

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- Experiments indicate that *K* signals are suppressed by rescattering, but signals are not. This puts limits on the duration of rescattering [arXi v: 1910. 14419]. How does this compare to rescattering in Pythia?
- On the other side of the coin, rescattering produces new resonances, but they can be difficult to detect. For example, in the process K K K, the K mass spectrum is unchanged.

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- To what extent does rescattering lead to jet quenching?
- To what extent does rescattering lead to resonance signal suppression?

#### Outline

Hadronic rescattering

#### Exotic hadrons

Cosmic rays

#### Motivation: what is the nature of exotic hadrons?



#### Bag model or molecular state?

In 2015, LHCb observed two peaks when studying  $_{b}$  J/  $\rho K^{-}$  (J  $\mu^{+}\mu^{-}$ ) decays, designated  $P_{c}^{+}$ (4380) and  $P_{c}^{+}$ (4450).

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In our work, we also studied the  $_{C1}(3872)$  tetraquark.

# My work

The objective of my work was to create a framework for studying exotic hadrons in Pythia. Specifically, we studied these pentaquarks and the  $_{c1}(3872)$  tetraquark.

We implemented pentaquark production them through  $_{b}$   $P_{c}^{+}$   $K^{-}$  decays, as well as exotic hadron resonances in rescattering, e.g.  $_{c}^{+} \bar{D}^{0}$   $P_{c}^{+}$  (4312) or  $D^{0} \bar{D}^{-0}$   $_{c1}$  (3872).

#### Results - the $P_c^+$ pentaquarks



- Rescattering cross section is somewhat model-dependent.
- Rate of production from  $\begin{array}{cc} 0 \\ b \end{array}$   $P_c^+ K^-$  decays is model dependent, specifically on  $P_c^+$  pJ/ branching ratio.

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### Results - experimental signal

Recall that rescattering resonances are hard to observe directly. The experimental pentaquark signal comes from the  $P_c^+$  pJ/  $p\mu^+\mu^-$  channel  $(p_{-,p} > 1 \text{ GeV and } p_{-,\mu} > 0.5 \text{ GeV, and } 2 < _{-p,\mu} < 5)$ 

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	P <sub>c</sub> <sup>+</sup> (4312)	P <sub>c</sub> <sup>+</sup> (4440)	P <sub>c</sub> <sup>+</sup> (4457)
model 1 model 2	$1 \times 10^{-6} \ 5 \times 10^{-5}$	$1 \times 10^{-6} \ 3 \times 10^{-5}$	$5 \times 10^{-7}$ $4 \times 10^{-5}$

Ratio of signal to background

- Contributions are too small to be observed at LHCb
- But this study shows that there is an order of magnitude difference between the two models

### Results - the $_{C1}(3872)$ tetraquark

For  $_{c1}(3872)$ , we measured the production cross sections in pp collisions at  $\overline{s} = 7$  TeV. (  $_{pp}$  90 mb)

- $\mbox{ }$  The inclusive production cross section has been measured by LHCb1, and found to be  $5.4~\pm~1.3~\pm~0.8$  nb.
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  0.04 nb, which is 1 % of the total production cross section.

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While the overall contribution is small, future LHCb data may make it possible to separate cross sections by production mechanism. This could provide insights on how to model tetraquark formation.

<sup>1</sup>Eur. Phys. J. C (2013) 73:2421

### Conclusions

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### Conclusions

- Overall, rescattering does not seem to be a major production mechanism for exotic hadrons.
- Rescattering will probably have a negligible effect on  $P_c^+$ pentaquark formation, but future data from LHCb can provide insights on how to model <sub>c1</sub>(3872).
- But this framework is just the beginning the framework can be extended to include other exotic hadrons, or other particles such as deuterons.

#### Outline

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Image credit: A. Chantelauze, S. Staffi, L. Bret

# Objectives

Programs such as  $CORSIKA^2$  are used to simulate hadronic cascades. These programs need models for generic hadron-nucleon collisions.

<sup>2</sup>arXi v: 1808. 08226, arXi v: 1902. 02822 (www. i ap. ki t. edu/corsi ka/)

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To do this, we need to total and partial cross section at perturbative energies, and describing PDFs for the relevant hadron species.

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#### Hadronic cross sections

For total cross sections, we use the Donnachie-Landshoff model:

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$$n_{\rm e} = n_d + n_u + 0.6n_s + 0.2n_c + 0.07n_b$$

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The Y coefficients are more complicated, and I won't go into technical details here.

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#### Parton distribution functions

Based on the ansatz by Glück and Reya,<sup>3</sup>

$$f(x, Q_0^2 = 0.26 \text{ GeV}^2) = N x^a (1 - x)^b$$

and evolve to higher scales using the QCDNUM program.

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The atmosphere is made of nuclei, not nucleons. Angantyr is not equipped to flexibly handle variable energies.

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### Outlook

With this work, also high energy interactions are available for important hadron-nucleon combinations (yay!)

There is some room for improvement in Angantyr:

- Variable energies are not yet supported
- There is no detailed handling of nuclear remnants
- Collisions below 100 GeV are not modelled accurately

Despite these shortcomings, Pythia is around the level of state-of-the-art models used by CORSIKA.

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- First steps towards studying exotic hadrons with MC generators (negative result: rescattering is probably not that significant production mechanism)
- With these general hadronic interactions, Pythia offers a plugin to hadronic cascade simulators