"Particle Fever" – The movie

SUSY in Electroweak & Higgs Sector

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ATLAS/CMS Run 3 SUSY Workshop:

New Directions for SUSY Searches with LHC Run 3 Data

Monday Nov 15, 2021



"Particle Fever" – The movie



NOTORIOUS SUPERSYMMETRY



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Big Picture



What is Dark Matter?









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The Beloved Beautiful (& Unnatural

Standard Model



https://en.wikipedia.org/wiki/Elementary_particle

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3 generations of matter $SU(3)_{C} \times SU(2)_{L} \times U(1)_{Y}$

WHY????



The Beloved Beautiful (& Unnatural)

Standard Model



https://en.wikipedia.org/wiki/Elementary_particle



The Beloved Beautiful (& Unnatural)



Only *Left* handed fermions charged under the weak SM gauge group.

Fermion and gauge boson masses FORBIDDEN by symmetry.



Whatever gives rise to fundamental particle masses has to break electroweak symmetry (EWSB).

The Higgs Mechanism.

Spontaneous Breakdown of the symmetry: $SU(2)_L \times U(1)_Y \rightarrow U(1)_{EM}$

A scalar (Higgs) field is introduced. The Higgs field acquires a nonzero value to minimize its energy.



$$V(\phi) = -m^2 |\phi|^2 + \lambda |\phi|^4$$



Masses of fermions and gauge bosons proportional to their couplings to the Higgs field:

 $M_{z,w} = g_{z,w} v$ $m_t = h_t v$ $m_h^2 = \lambda v^2$

v = 246 GeV



$V(H) = m_H^2 H^{\dagger} H + \frac{\lambda}{2} \left(H^{\dagger} H \right)^2$



$$\left\langle H^{\dagger}H\right\rangle = -\frac{m_{H}^{2}}{\lambda}$$

Only know location of minimum (vev) Only know local curvature (mass)

$$Higgs MASS^2 = -4 mH^2$$

NEGATIVE mass parameter (m_{μ}^2) .

NO additional symmetry if Higgs massless.

NOTHING protects Higgs Mass: UNSTABLE under quantum corrections!!



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V_c

н



How do scalars interact with gauge bosons?

$$|D_{\mu}\phi|^{2} = (\partial_{\mu}\phi + ieA_{\mu}\phi)(\partial^{\mu}\phi^{*} - ieA^{\mu}\phi^{*})$$



>--->--

 $e^2 A^2 |\phi|^2$

 $\rightarrow h + v$

>-x

W/Z



×->-

W/Z

We have seen that the Higgs couples to W/Z, with approximately the right strength!!

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SM-Like Higgs

Higgs generates masses of the SM particles!

P. Higgs:

"My first paper was rejected because it was not relevant for phenomenology"







Still large uncertainties in couplings... but compatible with SM expectations.

Observed Higgs Production x Branching Ratios as a ratio to SM expectation



SUSY and the Higgs.



SM Higgs is a Doublet

- The Higgs *FIELD* is a two component weakly charged doublet.
- *h* is the neutral particle we think we have observed at the LHC: *h*₁₂₅
- *v* is the SM vev: 174 GeV.
- G^{+/-} and G⁰ are "eaten" by the W and Z gauge bosons to give them mass.

$$H_{SM} = \begin{bmatrix} G^{\pm} \\ \frac{1}{\sqrt{2}}(h + iG^{0}) + v \end{bmatrix}$$

But SUSY has MORE!!



But we SEE a SNI-like Higgs!



2 Higgs Doublet Model (2HDM).

$$\langle H_1 \rangle$$
, $\langle H_2 \rangle \rightarrow \langle H \rangle$, $\tan \beta$

In SUSY Need 2 Higgs doublets:

H_u –Couples only to up-type quarks
 H_d –Couples only to down-type quarks and leptons.

•
$$m_A \sim m_H$$

• $\tan \beta = v_u / v_d$





SM: Only 1 Higgs which then acquires a vev and leads to EWSB.

This is what we want!

Lighter (h) is 125 GeV SM-like Higgs.

Additional states can exist!

Additional States can be light!

Haber and Gunion, '03, M. Carena, I. Low, N.R.S. & C. Wagner, '13, A. Delgado, G. Nardini & M. Quiros, '13, N. Craig, J. Galloway & S. Thomas,'13, P. Dev, A. Pilaftsis '14, M. Carena, H. Haber, I. Low, N.R.S. & C. Wagner '14 & '15 etc....

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 $<H_{d} > = v \cos \beta$ $<H_{u} > = v \sin \beta$ $\Rightarrow <H_{SM} > = v$ $<H_{NSM} > = 0$

SM-like HIGGS

ALIGNMENT



Minimal Supersymmetric SM (MSSM).

 $(\mathbf{\dot{U}})$

Standard particles



SUSY particles





For every fermion there is a boson of equal mass and couplings and visa versa.

's'particles and 'inos

No new couplings.

SUSY has to be broken.



Soft masses: Parametrize our Ignorance.

Higgs Mass = 125 GeV.

$$V = m_{ij}^2 \Phi_i^{\dagger} \Phi_j + \lambda_i \Phi_j^{\dagger} \Phi_k \Phi_l^{\dagger} \Phi_m$$

H. Haber and J. Gunion, '03

Quartics without quantum corrections related only to SM couplings.

Higgs mass bounded by m_z at tree-level.

91 **#** 125

Need large radiative corrections.





Stop + top Quartic Contributions!



$$\mathbf{M}_{\tilde{t}}^{2} = \begin{pmatrix} \mathbf{m}_{Q}^{2} + \mathbf{m}_{t}^{2} + \mathbf{D}_{L} & \mathbf{m}_{t} \mathbf{X}_{t} \\ \mathbf{m}_{t} \mathbf{X}_{t} & \mathbf{m}_{U}^{2} + \mathbf{m}_{t}^{2} + \mathbf{D}_{R} \end{pmatrix}$$

 $X_t = A_t - \mu / \tan \beta \rightarrow \text{LR}$ stop mixing

$$m_h^2 \simeq M_Z^2 \cos^2 2\beta + \frac{3}{4\pi^2} \frac{m_t^4}{v^2} \left[\frac{1}{2} \tilde{X}_t + t + \frac{1}{16\pi^2} \left(\frac{3}{2} \frac{m_t^2}{v^2} - 32\pi\alpha_3 \right) \left(\tilde{X}_t t + t^2 \right) \right]$$

$$t = \log(M_{SUSY}^2 / m_t^2) \qquad \tilde{X}_t = \frac{2X_t^2}{M_{SUSY}^2} \left(1 - \frac{X_t^2}{12M_{SUSY}^2}\right)$$

Carena, Espinosa, Quiros, Wagner, '95,96



Fine Tuned or Not??

Lighter Stops demand large splitting between left- and right-handed stops.







$$\cos(\theta - \alpha) = 0$$

$$-\frac{\sin lpha}{\cos eta} = \sin(eta - lpha) - \tan eta \cos(eta - lpha)$$

Approximately for t_{β} larger than a few, and $m_{H} > 200$ GeV:

$$\cos(\alpha - \beta) \simeq \frac{1}{\tan\beta \left(m_H^2 - m_h^2\right)} \left[m_h^2 + m_Z^2 + \frac{3m_t^4}{16\pi^2 v^2} \frac{\mu A_t}{m_Q^2} \left(1 - \frac{A_t^2}{6m_Q^2} \right) \tan\beta \right]$$

If $\mu A_t / m_0^2$ small (-> decoupling): Obtain • $\cos(\alpha-\beta) t_{\beta} \sim Const / (m_{H}^2 - m_{h}^2)$ ALIGNMENT for small tan β If $\mu A_t / m_0^2$ relevant: • Can have $cos(\alpha-\beta) \approx 0$

INDEPENDENT OF M_{A}

 $\Rightarrow LARGE \quad A_t/m_Q \\ \Rightarrow LARGE \quad \mu / m_0$

Would make it difficult to test low m_A and t_B via Higgs coupling measurements

Relevant parameters for Higgs Sector: A_{t}/m_{O} , μ/m_{O} , m_{A} , t_{B} Only Log dependence on squark masses, m_0 (relevant for $m_h^{\sim}125$ GeV)



Heavy Higgs Searches: A / H $\rightarrow \tau \tau$ Precision Higgs Physics: $h \rightarrow W W/Z Z$

- σ (bbH/A+ggH/A) × BR(H/A $\rightarrow \tau \tau$) (8 TeV)

--- σ (bbh+ggh) × BR(h \rightarrow WW)/SM



Higgs Mass = 125 GeV.

$$V = m_{ij}^2 \Phi_i^{\dagger} \Phi_j + \lambda_i \Phi_j^{\dagger} \Phi_k \Phi_l^{\dagger} \Phi_m$$

H. Haber and J. Gunion, '03

Quartics without quantum corrections related only to SM couplings.

Higgs mass bounded by m_z at tree-level.



91 **#** 125

Need large radiative corrections. ...Or something else?



$W = \lambda S H_u H_d + \frac{\kappa}{2} S^3$

- 2 Doublets (H_u, H_d) + Singlet (S)
- Singlet couples only to Higgs Sector.
- vevs: (H_u , $H_{d,}$, S) =(v_u , v_d , $v_s = \mu / \lambda$)
- 3 CP-Even Higgs bosons:
 - Mixing between all three (H_u , H_d and S).
- 2 CP-Odd Higgs bosons:
 - Mixtures of "MSSM" m_A and singlet.
- Charged Higgs bosons
- Singlino mass: 2 κ μ/λ

Something else: NMSSM

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 $-\mathcal{L}_{\text{soft}} = \lambda A_{\lambda} S H_u H_d + \frac{1}{2} \kappa A_{\kappa} S^3$

MA

M. Carena, H. Haber, I. Low, N.R.S., C. Wagner, '15



Alignment (No-Mixing):

$$m_h^2 \simeq \lambda^2 \frac{v^2}{2} \sin^2 2\beta + M_Z^2 \cos^2 2\beta + \Delta_{\tilde{t}}$$

 $\Delta_{\tilde{t}} = -\cos 2\beta (m_h^2 - M_Z^2)$

Well Known

- 125 GeV Higgs
 - Tree-level contribution to Higgs mass from λ.
 - $\lambda \sim 0.65-0.7$
- Low tanβ
- Light Stops

NMSSM-125 GeV Higgs Naturally!





Not so well known:

• Excellent Alignment (very little mixing)

M. Carena, H. Haber, I. Low, N.R.S., C. Wagner, '15 1.0 $\frac{m_h^2 - M_Z^2 c_{2\beta}}{v^2 s_\beta^2}$ $\lambda_{
m alt}^2 =$ 0.9 mine = 125±3 GeV ~ 0.8 0.7 $m_h = 125 \pm 3 \text{ GeV}$ 0.6 1.5 2.0 2.53.0 1.0tan β



Originate dynamically from Higgs compositeness at the GUT scale?

SM-Like Higgs Naturally!

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N. Coyle, C. Wagner, '19



Singlet Alignment

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SM - like h₁₂₅: Mass Correlations!

S. Baum, N.R.S, K. Freese, '19



NMSSMTools scan with consistent h_{125} pheno



So ...

HOW ABOUT NEUTRALINO WIMPS FOR Neutral Superpartners of Higgs, Z, Y

DARK MATTER?



Charginos/Neutralinos & Sleptons...

SM Gauge Bosons, Higgs (2HDM in SUSY) and leptons Superpartners		Bunch of weakly coupled scalars, neutral Majorana fermions, lightest stable due to R-parity: WIMPI								
B: SU(2) singlet		Bino M ₁ g ₁								
{W ⁰ , W ¹ , W ² }: SU(2) triplet		Winos M ₂ g ₂ 1 neutral, 2 charged								
{H _u H _d }: SU(2) doublets		Higgsinos μ Yukawas etc 2 neutral, 2 charged								
{e, μ, τ}		Selectron, smuon, staus M _{µL,R} etc SM couplings								
ACCNA, A noutral "Noutralings" mixtures of interaction states (Also 2 shored "Charging										

MSSM: 4 neutral "Neutralinos", mixtures of interaction states (Also 2 charged "Charginos" mixtures of wino and Higgsinos). (Also Singlino for NMSSM) $\chi = N_{11}\tilde{B} + N_{12}\tilde{W} + N_{13}\tilde{H}_d + N_{14}\tilde{H}_u$

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Thermal Relic?



What sets the abundance of the Dark Matter observed?

Annihilations try to maintain thermal equilibrium.

Universe is Expanding!

At some point a DM particle can't "find" another DM particle to annihilate with: FREEZE-OUT.

LARGER annihilation rate means LOWER number density.

Hooper, 09

WIMPs

Weakly Interacting Massive Particles



The WIMP Miracle.

$$\sigma = \frac{"\alpha"^2}{m^2}$$

Weak scale couplings and masses (100 GeV)!

$\langle \sigma v \rangle \Longrightarrow \text{observed } \Omega_{DM}$

NATURALLY!





SIDD + Ωh^2 ??





SENON1T arXiv:1805.12562

m_x ~ few 100 GeV Break the Connection! Multiple mediators for destructive interference Co-annihilation/Resonance ALL consistent with extended Higgs sectors!!



Relic Density: Annihilations

Light Binos or Singlinos (SM singlets) excellent candidates for Dark Matter.



Cheung, Papucci, Sanford, NRS, Zurek '14, Freese Lopez, NRS, s-channel ResonSheleya '15, Carena, Osborne, NRS, Wagner '19, etc... DM close to half of mediating particle mass. Constrained but possible for h_{125} , Z, & additional Higgs bosons.

t-channel annihilation with light stau. Happens in a natural way at large *tan*β. NRS, Pierce, Freese'1



Relic Density: Co - Annihilations

When DM can annihilate against other rapidly annihilating particles.

Ibarra, Pierce, NRS, Vogl '15, Pierce, NRS, Vogl '18, Baum, Carena, NRS, Wagner '18, '21 etc

Mass difference of Dark Matter with the other weak scale weakly interacting particles must be of the order of a few tens of GeV.

Naturally leads to compressed spectrum -> reduced sensitivity at the LHC.

Some relevant channels in the case of sleptons or Winos (too light Higginos/ small μ leads to large SD cross sections).





Relic Density: Co - Annihilations

S. Baum, M. Carena, N.R.S, C. Wagner, '17

Thermal production via co-annihilation of Bino with singlino or resonance



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$$\sigma_p^{SI} \propto \frac{m_Z^4}{\mu^4} \left[2(m_{\tilde{\chi}_1^0} + 2\mu/\tan\beta) \frac{1}{m_h^2} + \mu \tan\beta \frac{1}{m_H^2} + (m_{\tilde{\chi}_1^0} + \mu \tan\beta/2) \frac{1}{m_{\tilde{Q}_1}^2} \right]^2$$

$$(2(m_{\tilde{\chi}_1^0} + 2\frac{\mu}{\tan\beta}) \frac{1}{m_h^2} \simeq -\mu \tan\beta \left(\frac{1}{m_H^2} + \frac{1}{2m_{\tilde{Q}_2}^2} \right) \begin{array}{c} \mu \times m_{\tilde{\chi}^0} < 0\\ m_{\tilde{\chi}^0} \simeq M_1 \end{array} \right]^2$$

$$(2(m_{\tilde{\chi}_1^0} + 2\frac{\mu}{\tan\beta}) \frac{1}{m_h^2} \simeq -\mu \tan\beta \left(\frac{1}{m_H^2} + \frac{1}{2m_{\tilde{Q}_2}^2} \right) \begin{array}{c} \mu \times m_{\tilde{\chi}^0} < 0\\ m_{\tilde{\chi}^0} \simeq M_1 \end{array} \right]^2$$

$$(2(m_{\tilde{\chi}_1^0} + 2\frac{\mu}{\tan\beta}) \frac{1}{m_h^2} \simeq -\mu \tan\beta \left(\frac{1}{m_H^2} + \frac{1}{2m_{\tilde{Q}_2}^2} \right) \begin{array}{c} \mu \times m_{\tilde{\chi}^0} < 0\\ m_{\tilde{\chi}^0} \simeq M_1 \end{array} \right)^2$$

$$(2(m_{\tilde{\chi}_1^0} + 2\frac{\mu}{m_h^2}) \frac{1}{m_h^2} \simeq -\mu \tan\beta \left(\frac{1}{m_H^2} + \frac{1}{2m_{\tilde{Q}_2}^2} \right) \begin{array}{c} \mu \times m_{\tilde{\chi}^0} < 0\\ m_{\tilde{\chi}^0} \simeq M_1 \end{array} \right)^2$$

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$$(2(m_{\tilde{\chi}^0} + 2\frac{\mu}{m_h^2}) \frac{1}{m_h^2} \simeq -\mu \tan\beta \left(\frac{1}{m_H^2} + \frac{1}{2m_{\tilde{Q}_2}^2} \right) \begin{array}{c} \mu \times m_{\tilde{\chi}^0} \approx M_1 \end{array} \right)^2$$

$$(2(m_{\tilde{\chi}^0} + 2\frac{\mu}{m_h^2}) \frac{1}{m_h^2} \approx -\mu \tan\beta \left(\frac{1}{m_h^2} + \frac{1}{2m_{\tilde{Q}_2}^2} \right) \left(\frac{1}{m_{\tilde{\chi}^0}} + \frac{1}{m_h^2} + \frac{1}{$$

 $\sigma^{
m SD} \propto rac{m_Z^4}{\mu^4} \cos^2(2eta)$

NMSSM: similar, but different signs.

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BUT ISN'T SUSY ALREADY RULED OUT BELOW THE TEV SCALE ??



Strong Physics

Guinos:

If decay directly to 3rd generation squarks, gluinos must be heavier than about 1.5 to 2.2 TeV

Cascade decays into intermediate chargino/neutralino states and compressed spectrum present the weakest limits, and the bound falls short of 2 TeV for non-compressed spectrum. Bound of 2.2 TeV in the most extreme case.

Hard to evade the TeV bound.

Stops:

Higgs mass implies stops masses heavier than ~ 1 TeV

Combining all searches, in the simplest decay scenarios, it is hard to avoid the constraints 600 GeV - 1.2 TeV for stops.

We are just starting to explore the mass region suggested by the Higgs mass!



All that Weak stuff?

Situation here is far less well defined than in the strongly interacting sector.

Sleptons, in particular staus, are only weakly constraint beyond the LEP limits.

Winos as NLSP's are the strongest constrained particles.

Sensitivities in the search for these particles will increase only at **high luminosities**, but bounds on Higgsinos will remain weak.

In general, a SUSY scenario with large cascade decays with light electroweakinos is the most natural one and the least constrained so far.



LHC?

- Interesting cascade decays of the additional Higgs bosons in the NMSSM - Eg: $H_{NSM} \rightarrow h_{S} H_{SM}$ or $H_{NSM} \rightarrow a_{S} Z$ $H_{NSM} \rightarrow WW/ZZ/H_{SM} H_{SM}$ or $A \rightarrow Z$ H_{SM} suppressed due to alignment
- Can have visible and invisible decays of h_s/a_s



LHC All Searches: 3000 fb⁻¹ NMSSM Higgs Sector



Expect ~50% scanned region with $m_A < 1$ TeV probed.

~90% if improvements by 1 order of magnitude wrt our projections, & 2 orders of magnitude in conventional search channels wrt current limits.

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Generic NMSSM: Explore 2HDM + S!!

S. Baum & N.R.S., '18



If either m_a or $m_h > 350$ GeV, can decay predominantly to tt. Could be hidden in h_{125} tt?



Max Misalignment Scenario: $H \rightarrow h_{125} h_{125}$

Compare to CMS/ATLAS limits (2019) in *bbbb, bbyy* final states.

S. Baum & N.R.S., '19



Gray shaded regions excluded by $(H \rightarrow h_{125}h_{125})$ searches. Black dashed lines label cross sections a factor two smaller, and hence may be probed with 300 fb⁻¹ of data. Red [orange] shaded area denote regions ruled out by current (H \rightarrow WW) [(H \rightarrow ZZ)] LHC searches



MUCH MORE to Explore!

Stops and gluinos > TeV

Higgs & Electroweak Sector Tangled

DM + (g-2) (S. Heinemeyer talk) + Alignment + Higgs Mass (CPV?!)

Experimental Simplified Model Analysis Limits

Naïve Reinterpretation to Exclude ~~ imes

Bottle neck: Data + Search strategies for compressed spectra

NOT Energy!



Snowmass - EF08: Model Specific Explorations

Convenors: Elliot Lipeles, Nausheen Shah, Jim Hirschauer

 We are looking for people to take on specific (but manageable) responsibility for individual EF08 summary plots for the Snowmass report (current list of plots below). The work would involve:

1) Reviewing corresponding European Strategy plot and documenting source of all numbers/lines in plots. We hope to have this done in the next few weeks. Specifically, we'd like to fill out spreadsheets that look like this:

Machine			ES Report						Snowmass			
	Energ			Figure/Ta ble in ES	Number/ Line in			Comment on Number		Update		
Collider	y	L (ab-1)	Presented Number/Line	Report	figure	Reference	Fig	Source	Comment method	Priority	Group(s)	Status
								LHC fig, 2 sigma at				
HL-LHC	14	6?	at 3 ab-1, to light quarks	8.6	3.2 TeV	https://arxiv.org/pdf/1812.07831.pdf	2.1.1	mX=0	Delphes			
			at 3 ab-1, to light quarks					LHC fig, 2 sigma on				
			compressed	8.6	1.5 TeV	https://arxiv.org/pdf/1812.07831.pdf	2.1.1	diagonal	Delphes			
			at 3 ab-1, to light top quarks	8.6		https://arxiv.org/pdf/1812.07831.pdf	2.1.6		CMS official			
			at 3 ab-1, to s top quarks	8.6		unclear?						
FCC-hh	75	30	at 30 ab-1, to light quarks	8.6		https://link.springer.com/article/10.1140%2Fepjc%2Fs1 0052-019-6904-3	9.1		FCC			
			at 30 ab-1, to light quarks compressed	8.6		"extrapolated from FCC-hh prospects"						
			at 30 ab-1, to s top quarks	8.6		https://link.springer.com/article/10.1140%2Fepjc%2Fs1 0052-019-6904-3	9.1		FCC			
FCC-hh	100	30										
FCC-hh	150	30										

https://docs.google.com/spreadsheets/d/1KxxcOpF2PgrAe8p_3QV D910hwGHrYqRFmAIG3xEWSao/edit?usp=sharing

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More Machine (e+e-, muon, e-h...)

Snowmass - EF08: Model Specific Explorations

Convenors: Elliot Lipeles, Nausheen Shah, Jim Hirschauer

• The work would involve (con't...):

2) Identify new numbers/lines to include. Similarly in the next few weeks.

 Identify changes to plots large or small with a target to have big picture changes defined by end of calendar year.

 In 2022: track results needed, collect results in appropriate formats, and ultimately assemble the plots.

Of course, we'd also like contributions to the specific results identified in (3).

Current list of plots essentially the same as the European Strategy (Suggestions welcome for additions):

- Gluino
- Light squarks
- Stop squark
- High-Delta M electroweak (not necessary Wino-Bino, to be discussed)
- Compressed electroweak

We could potentially sub-divide the task for a plot for example into hadron and lepton colliders if you are interested.

If you want to volunteer to coordinate a plot or set of lines in a plot (e.g. all e+e-), let us know: lipeles@sas.upenn.edu, jhirsch@fnal.gov, nausheen.shah@wayne.edu



Thank You!



What are the right questions?

SM & DM beautiful & EMPIRICAL...

Electroweak sector heart of the puzzle.

SUSY (may?) be the answer! LHC: Strong constraints on strong sector...

LHC: Electroweak sector -> Need more data!

Higgs/Electroweak???

Data + Theory: Where to look next!

Absence of Evidence != Evidence of Absence Flavors of SUSY -> Wildly different Pheno -> Much to explore!





BACK UP SLIDES



Top quark contribution:



Fine Tuning?

$$egin{aligned} \left(\Delta M_{H}^{2}
ight)_{t} &= -iN_{c}(-ih_{t})^{2}\intrac{d^{4}k}{(2\pi)^{4}}\operatorname{Tr}\,\left(rac{i}{k}-m_{t}rac{i}{k}-m_{t}
ight) \ &\sim -rac{N_{c}(h_{t})^{2}}{4\pi^{2}}\Lambda^{2} rac{1}{\Lambda\sim M_{P}\sim10^{19}\,\,\mathrm{GeV}} -10^{37}\,\,\mathrm{GeV}^{2} \end{aligned}$$

Complex scalar contributions:

$${\cal L}=-\lambda_r H^2 \phi_r^\dagger \phi_r$$

$$H \xrightarrow{\phi_r} H \qquad H \xrightarrow{\phi_r} H \qquad H \xrightarrow{\phi_r} H$$

$$\left(\Delta M_{H}^{2}
ight)_{\phi_{m{r}}}\simrac{\lambda_{r}N_{c}}{8\pi^{2}}\Lambda^{2}$$

Need to CANCEL LARGE contributions to produce physical Higgs mass of 125 GeV!!



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Top quark contribution:



Fine Tuning: **SUSY Cancelations!**

Complex scalar (STOP) contributions :

$${\cal L}=-\lambda_r H^2 \phi_r^\dagger \phi_r$$

$$H \xrightarrow{\phi_r} H H H \xrightarrow{\phi_r} H$$

$$\left(\Delta M_{H}^{2}
ight)_{\phi_{T}}\simrac{\lambda_{r}N_{c}}{8\pi^{2}}\Lambda^{2}$$

Quadratic divergences CANCEL if EQUALITY of couplings.

Residual (mass dependent) logarithmic contributions remain



- Interaction basis: (H_u, H_d, S)
 - H_u: Couples only to up-type fermions
 - H_d: Couples only to down-type fermions
 - S: Only couples to Higgs



CP-Even Higgs Bases



- Interaction basis: (H_u, H_d, S)
 - H_u: Couples only to up-type fermions
 - H_d: Couples only to down-type fermions
 - S: Only couples to Higgs

$$= v_{u}$$
$$= v_{d}$$
$$t_{\beta} = v_{u}/v_{d}$$
$$~~= \mu/\lambda~~$$

Only SM state couples to WW or ZZ!!

"Extended" Higgs basis: (H_{NSM}, H_{SM}, S)

- H_{NSM} : (down, up, V) = ($y_d t_\beta$, y_u / t_β .
- H_{SM} : (down, up, V) = (y_d , y_u , g_{hVV})

CP-Even Higgs Bases



- Interaction basis: (H_u, H_d, S)
 - H_u: Couples only to up-type fermions
 - H_d: Couples only to down-type fermions
 - S: Only couples to Higgs

-
$$H_{NSM}$$
: (down, up, V) = ($y_d t_\beta$, y_u / t_β , 0)

$$- H_{SM}: (down, up, V) = (y_d, y_u, g_{hVV})$$

$$H^{i} = S^{i}_{NSM} H_{NSM} + S^{i}_{SM} H_{SM} + S^{i}_{S} H_{S}$$

CP-Even Higgs Bases

$$= v_{u}$$
$$= v_{d}$$
$$t_{\beta} = v_{u}/v_{d}$$
$$~~= \mu/\lambda~~$$

- Interaction basis: (H_u, H_d, S)
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- "Extended" Higgs basis: (H_{NSM}, H_{SM}, S)
 - H_{NSM} : (down, up, V) = ($y_d t_\beta$, y_u / t_β , 0)

$$- H_{SM}: (down, up, V) = (y_d, y_u, g_{hVV})$$

• Mass basis: $(H_3, H_2, H_1) \longrightarrow (H, h125, h)$ $H_i = S_i^{NSM} H_{NSM} + S_i^{SM} H^{SM} + S_i^{S} H_S$

CP-Even Higgs Bases

CP-odd mix similarly: $A_i = P_i^{NSM} A_{NSM} + P_i^{S} A_{S}$



Current Bounds from Direct Dark Matter Detection

Current Limits

 $1 \text{ pb} = 10^{-36} \text{ cm}^2, \qquad 1 \text{ zb} = 10^{-45} \text{ cm}^2$



Spin Independent Interactions

Spin Dependent Interactions



NMSSM: SI Direct Detection

 $\Omega h^2 = 0.12 \pm 50 \%, B$ SI Direct Detection cross section suppressed by: Cheung, Papucci, Sanford, NRS, Zurek, `14; Huang, Wagner, 2014, etc. 10^{-10} i = hs 10^{-11} Proximity to h_{125} blind spot 10^{-12} Destructive interference between CP-even $[\mathrm{qd}]_{\mathrm{IS}^{d}} 10^{-13}$ Higgs mass eigenstates 10^{-14} S. Baum, M. Carena, N.R.S, C. Wagner, '17 10^{-15} 10^{-10} 10^{-16} -10^{-11} -10^{-13} $\pm 10^{-15}$ 10^{-15} 10^{-8} $(\overline{\sigma_p^{\mathrm{SI}}})_i$ [pb] $\Omega h^2 = 0.12 \pm 50 \%, \, \tilde{S}$ 10^{-10} 10^{-10} i = H<u>a</u> 10⁻¹² 10^{-11} $\widetilde{H}^{\dagger} + \widetilde{S}$ 10^{-12} $b^{15}b^{a}10^{-14}$ \widetilde{B} $[qd]_{IS}^{\ \ d} 10^{-1}$ \widetilde{S} 10^{-16} $\widetilde{B}+\widetilde{H}+\widetilde{S}$ 10^{-1} $\widetilde{B} + \widetilde{H}$ 10^{-18} 10^{-15} $\widetilde{B} + \widetilde{S}$ 10^{-20} 10^{-9} 10^{-11} 2 -1 $m_{\chi}/\mu\sin 2\beta$

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Stau searches: Bounds depend on mixing!



Weak limit at this point, we are now exploring limits beyond LEP. Observe that this assumes degenerate stau masses.



Slepton Searches



Assuming all leptons are degenerate, bound can be as large as 700 GeV. Bounds are significantly relaxed if mass difference between sleptons and neutralinos is smaller than ~100 GeV.

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Current Electroweakino Mass Bounds Wino NLSP, BR = 1



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But also... The Squarks!

In evaluating these bounds the squarks have been decoupled. **But** cross section depends on the squark masses due to a *t* and *u* channel contribution

Liu, McGinnis, Wang, Wagner arXiv:2008.11847

The resulting cross sections may differ by factors of a few!





