MMA: Our Three Pillars

Gravitational waves

Neutrinos









MMA Organization

Major Challenges

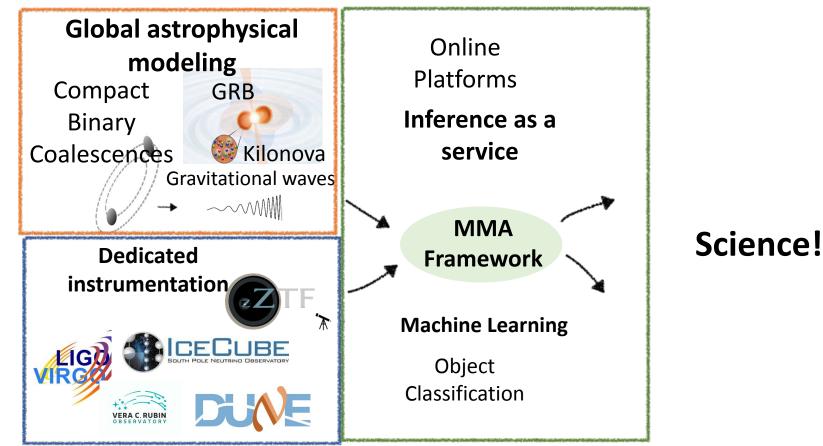
- Data rights: building open source data sets that A3D3 can work on
- Internal connections: disparate science areas with very different data sets (connected by similar science goals)
- Problem formulation: wide variety of data sets and physical intuition
- Connection with stakeholders: useful algorithm development all the way to deployment requires connections outside of A3D3

Organizational Discussion points

- How do we want to organize meetings (and balance internal collaboration ML meetings vs. A3D3 activities)?
- Do we need a regular "A3D3 MMA" meeting?
- How should we prioritize personpower? Relative to stakeholder deadlines presumably.

Scientific Discussion points

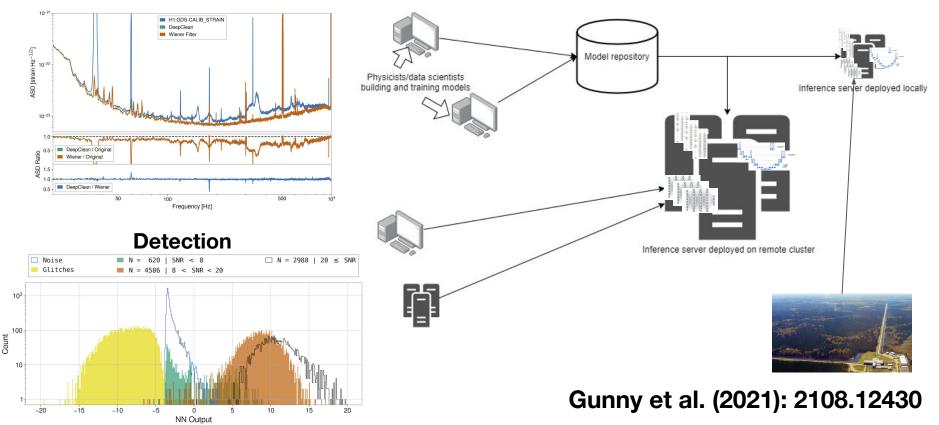
- What is the role of open data from experiments? Can we produce MMA Monte Carlo datasets for joint (end-to-end) analyses (example case studies: (1) CC-SN2023A in M31, (2) GW221109 at 100Mpc)?
- How can we work towards identifying off-the-shelf products/approaches that can be expanded to other areas: e.g., IaaS, image recognition?
- Can we identify a common conceptual path to establish the potential/merit for transition from traditional algorithms/computing to ML/HW acceleration? Establishing that path requires interaction with all groups in A3D3.



Multi-Messenger science includes representation from many of the most interesting experiments today.

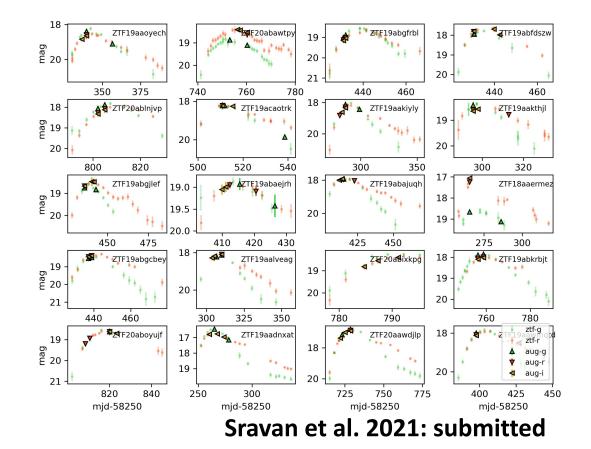
Pillar 1 Project: Gravitational-wave Inference as a Service

Data Cleaning

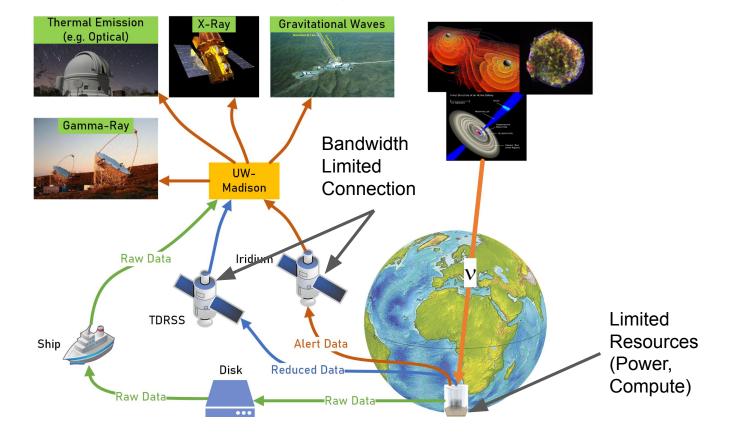


Pillar 2 Project: Optimal Light curve Augmentation

- Data will be insufficient for full science inference without additional follow-up
- Need to perform value-driven follow-up
- Broadly an optimal real-time resource allocation problem

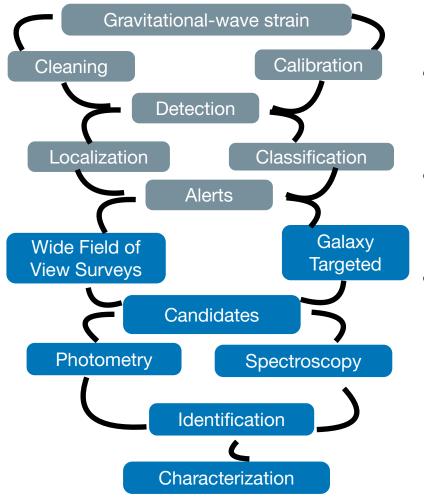


Pillar 3 Project: MMA at the Edge



Project 3: MMA at the Edge

- Want to send best results to community ASAP
- Resources limited at experiment
 - Bandwidth to Northern Hemisphere
 - Limited computational resources Power constraint
 - Can not communicate with partners easily
 - Future detectors cannot transmit all their data to the surface
- Time critical and precise reconstruction
 - Need to transmit fast and accurate directional pointing
 - Data set needs to be analyzed on-site
 - ML recos O(1s) vs. traditional recos O(hours days) FPGA or GPU acceleration -Goal 2024/25
 - Fast unfolding of ML inputs (waveforms) FPGA or GPU acceleration Goal 2023/4
 - Push waveform unfolding down to the sensors FPGA Goal 2025/6 (new detector design for Gen2 is ongoing)



- HERMES (2108.12430): Saleem/Will, Alec/Dylan/Jeff
 - DeepClean (2005.06534)
 - BBHNet
- Optimal Light curve augmentation: Tyler/Mouza, Ari
 - NMMA (2002.11355)
 - ORACLEs (2003.08943)
 - SkyPortal (skyportal.io)

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