



# Status of the H.E.S.S. experiment

Operating Cherenkov telescopes 20 years after their installation





# The H.E.S.S. telescope array

- 5 telescopes with 960 / 1764 pixel cameras
- 3.5° – 5.0° field-of-view,  $\sim 0.1^\circ$  angular resolution
- 30 GeV – 100 TeV,  $\sim 15\%$  energy resolution

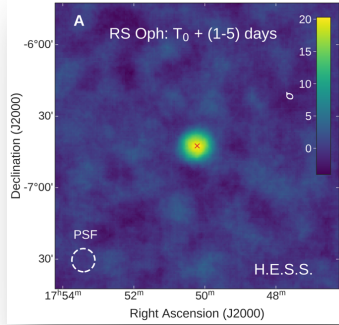
H.E.S.S. Collaboration:  $\sim 250$  members, at 38 institutes, in 13 countries

Stefan Ohm . RICH 2022 . Edinburgh . 09/2022



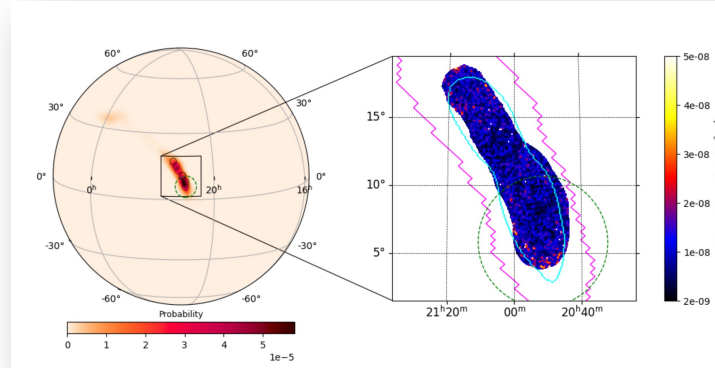
# Shift of science focus over the years

## 1<sup>st</sup> Galactic transient: Nova RS Oph

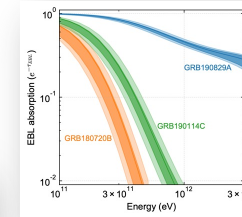


H.E.S.S. Collaboration: Science, 376, 6588 (2022) © DESY/H.E.S.S., Science Communication Lab

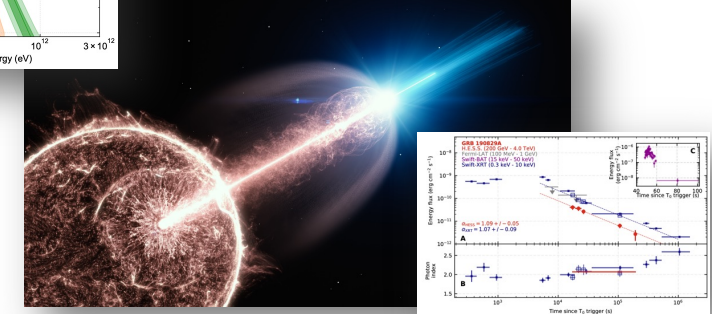
## Multi-messenger: GW follow-up



H.E.S.S. Collaboration: ApJ, 923, 109 (2021)

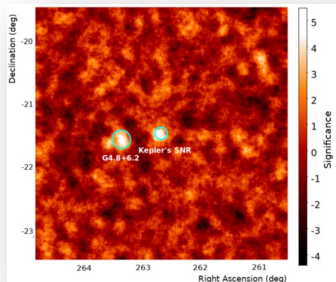


## Time-domain & multi-wavelength: e.g. GRBs



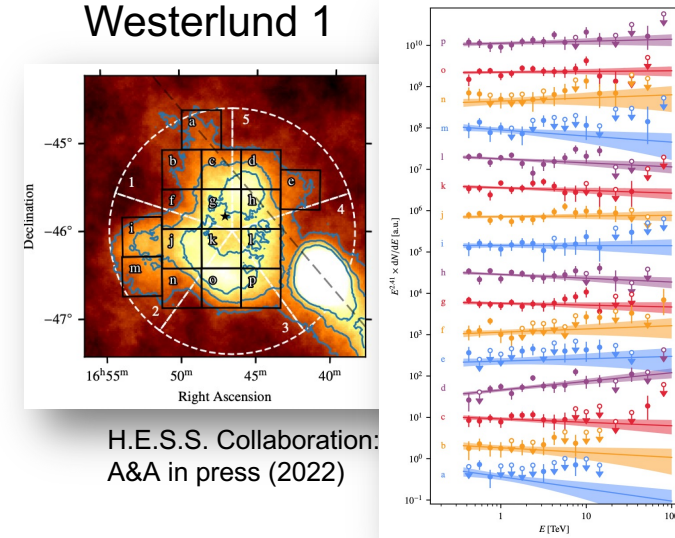
H.E.S.S. Collaboration: Science, 372, 1081 (2021) © DESY/H.E.S.S., Science Communication Lab

## Deep exposures: Kepler SNR

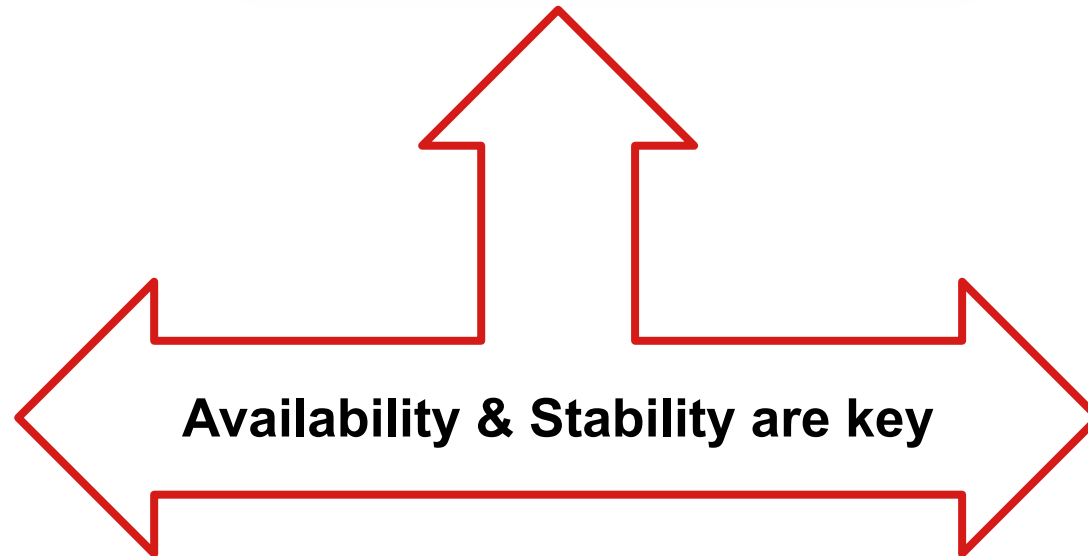


H.E.S.S. Collaboration: A&A, 662, A65 (2022)

## Precision measurements: Westerlund 1

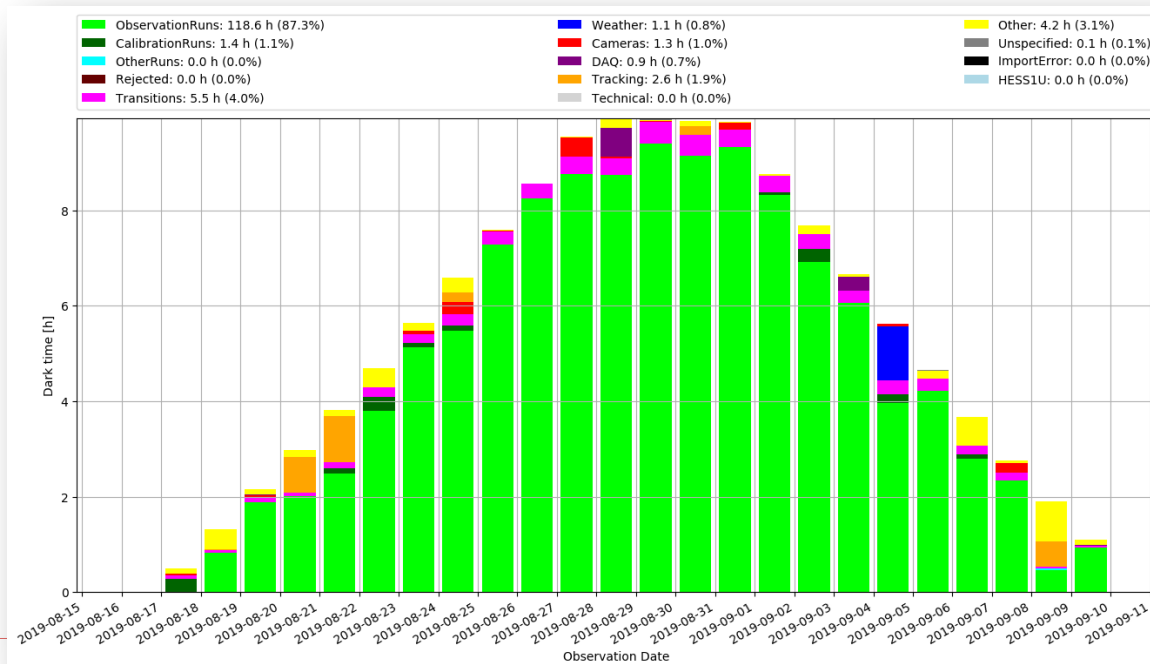


H.E.S.S. Collaboration: A&A in press (2022)

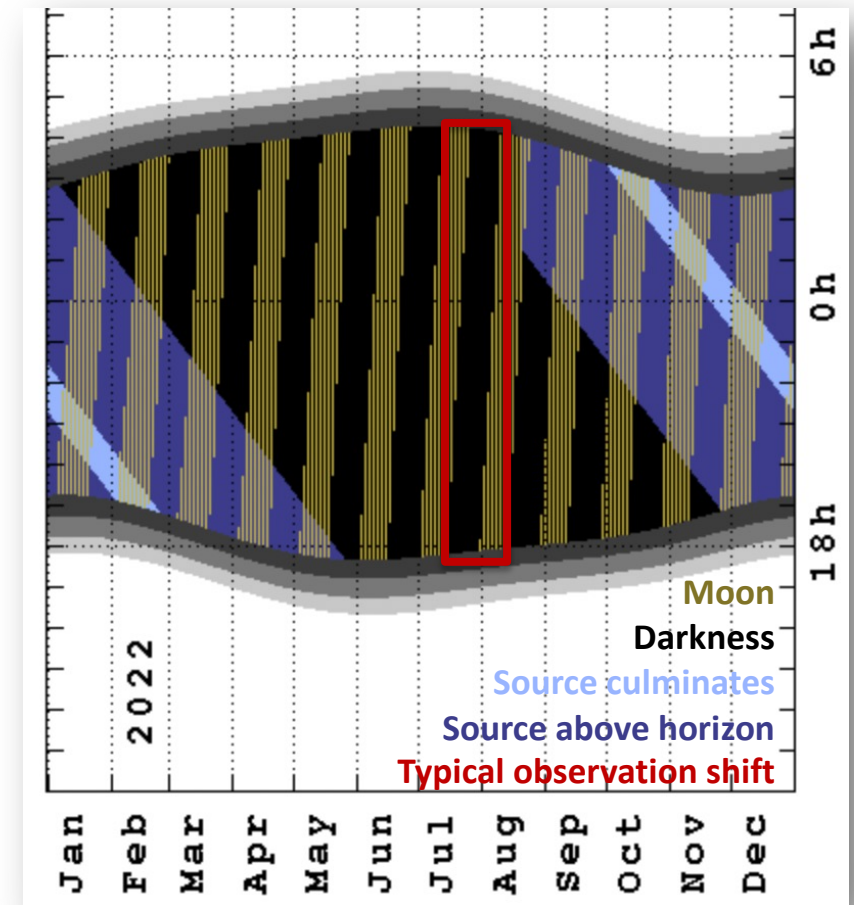


# How H.E.S.S. operates

- Observations in dark nights (~1000 hrs / 10% duty cycle)
- Typically spent on
  - Scientific observations
  - Calibration & technical observations
  - Transitions between sky objects
  - Downtime due to technical problems & maintenance, or
  - Downtime due to bad weather



## Visibility of objects throughout the year

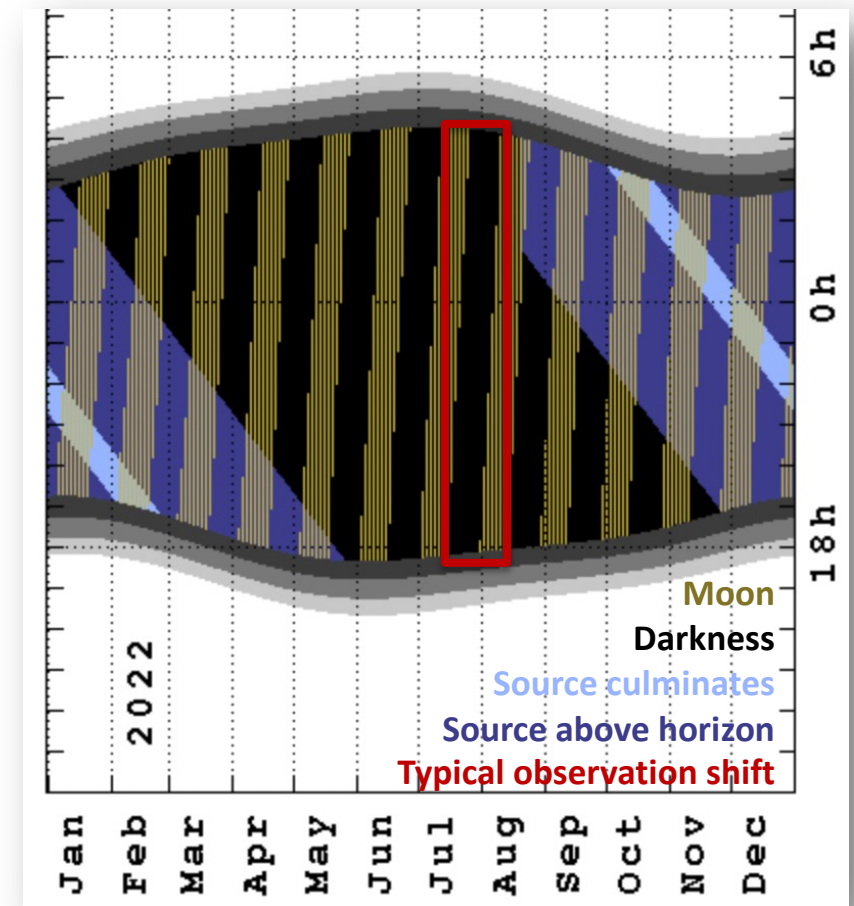




## How H.E.S.S. operates, and where one can improve

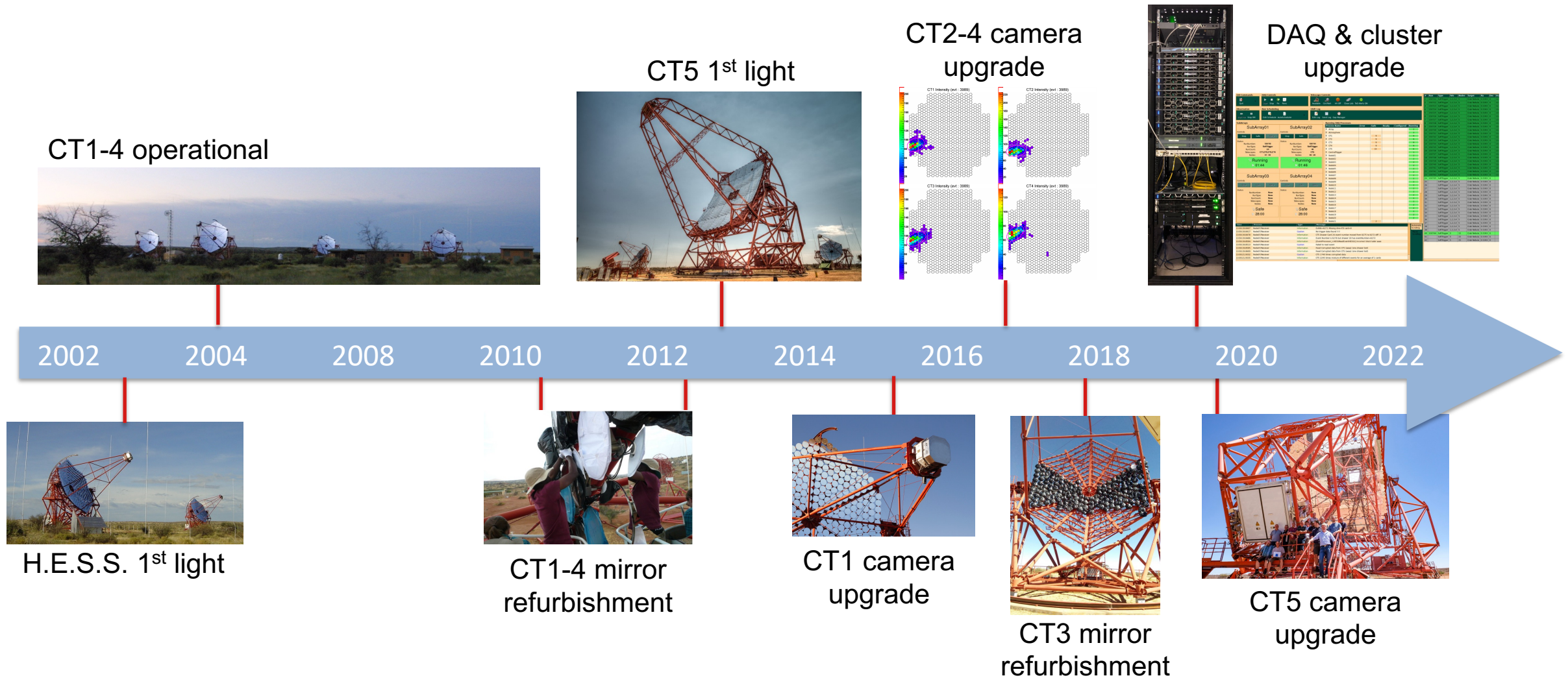
- Observations in dark nights (~1000 hrs / 10% duty cycle)
- Typically spent on
  - Scientific observations
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  - Transitions between sky objects
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  - Downtime due to bad weather
- Potential to increase observation time
  - Observe under moderate moonlight
  - Observe under “twilight”
  - Reduce downtime due to technical problems & maintenance
  - Optimise calibration & technical observations
- Regular hardware maintenance and upgrades to maintain e.g. optical throughput of the system

Visibility of objects throughout the year





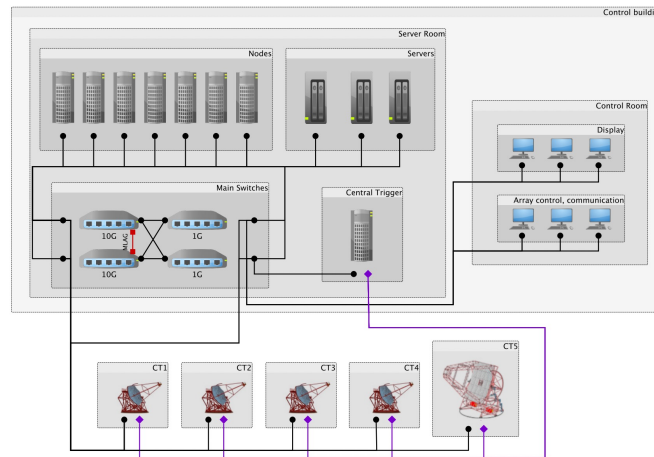
# Upgrades maintain & improve performance and stability



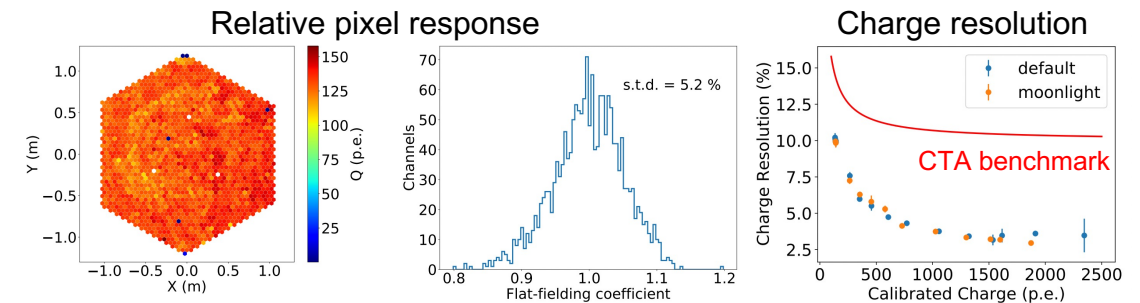
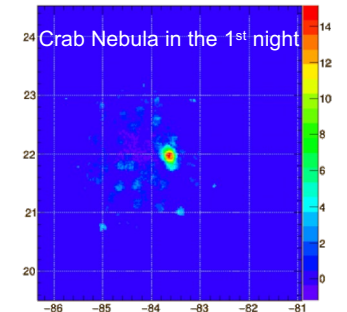


# Hardware changes in the first HESS extension

- Data Acquisition System (DAQ) upgrade
  - Upgrade of main DAQ hard- and software on- and off-site in 2019
  - Preparations for new CT5 camera and increased computing needs
  - Preparation for continued HESS operations
  - Enhanced functionality for cluster setup and monitoring, ++
- CT5 camera upgrade
  - Installation of CTA MST prototype camera in CT5 in October 2019
  - Very stable operation since ~3 years (>99% uptime)
  - Performance meeting CTA requirements



S.J. Zhu, et al., ICRC 2021



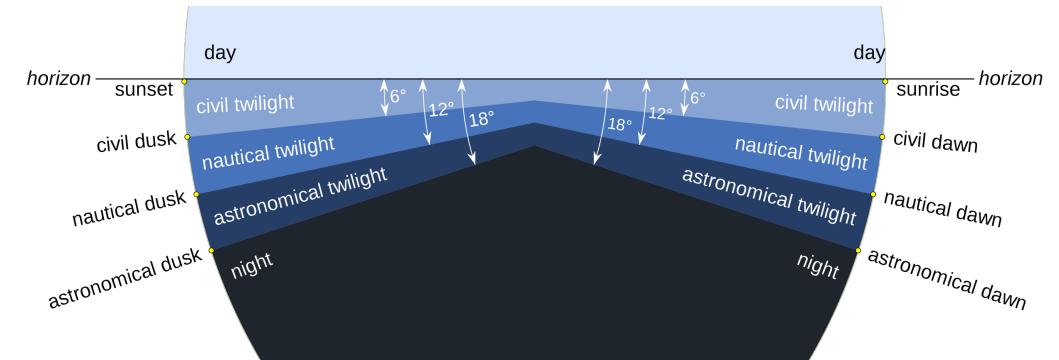
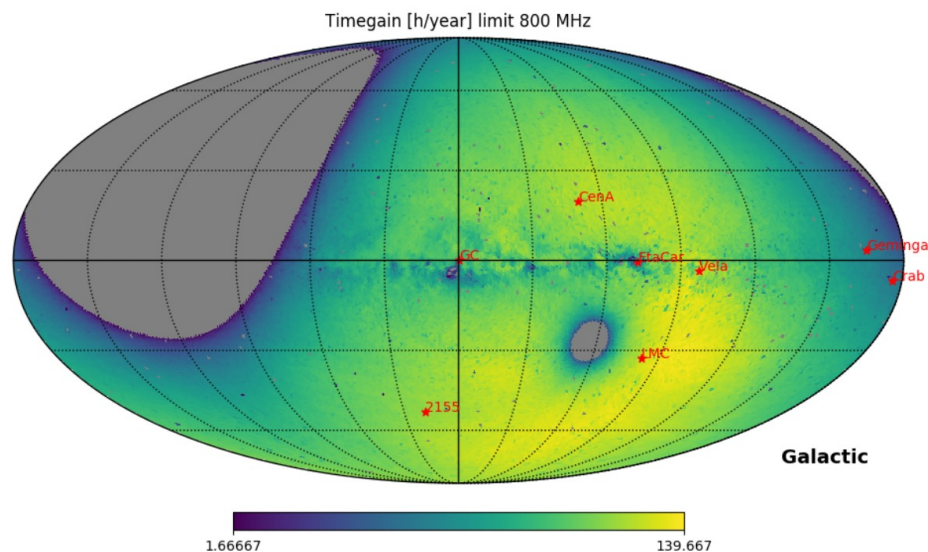
B. Bi, et al., ICRC 2021



# Extending observation time through moonlight and twilight observations

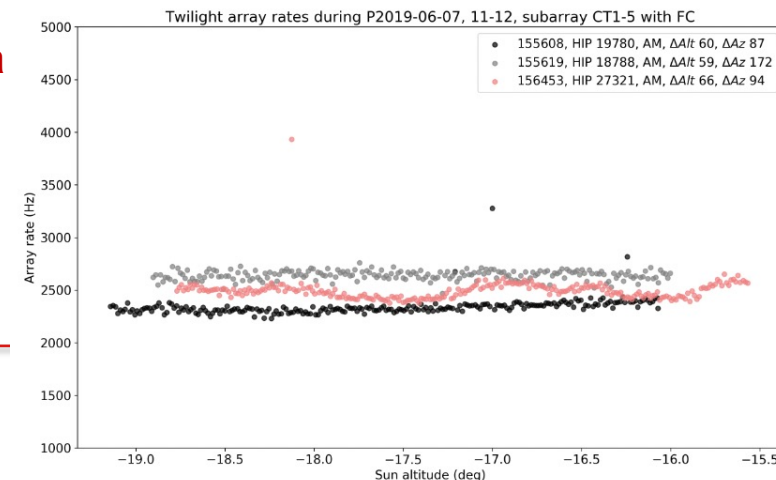
- Moonlight observations
  - Initial tests of moon observations in 2019
  - Hardware settings fixed since 03/2020
  - Full implementation since 01/2021 including scheduling, transient follow-up system, etc.
- Current moon settings
  - Max. illuminated moon fraction: 40%
  - Target-moon separation:  $45^\circ - 145^\circ$
  - Max. Night-sky background (NSB):  $3.5 \times \text{NSB}_{\text{dark}}$

⇒ >250 hours/yr extra



- “Darkness” definition
  - Historically, observations conducted in astronomical darkness (sun elevation:  $-18^\circ$ )
  - Extensive tests to start/end observations in astronomical twilight
  - Darkness definition changed to sun el. of  $-16^\circ$

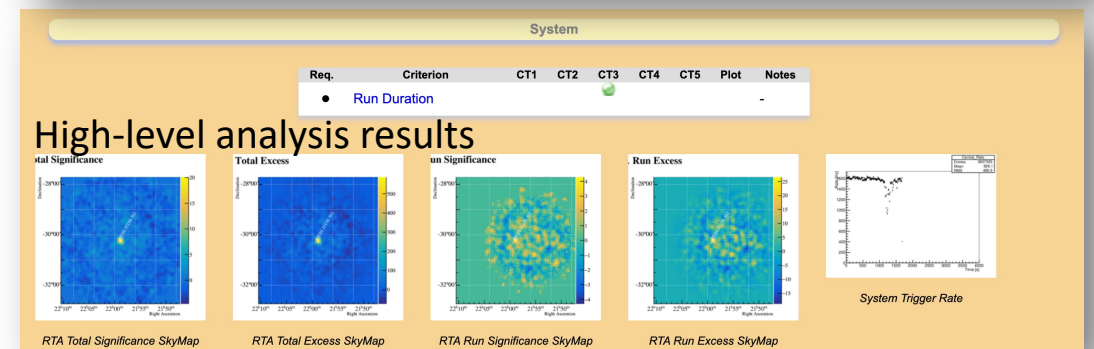
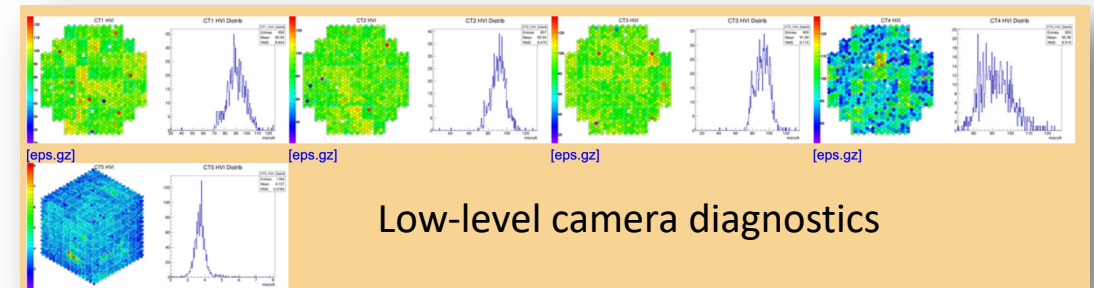
⇒ ~70 hours/yr extra





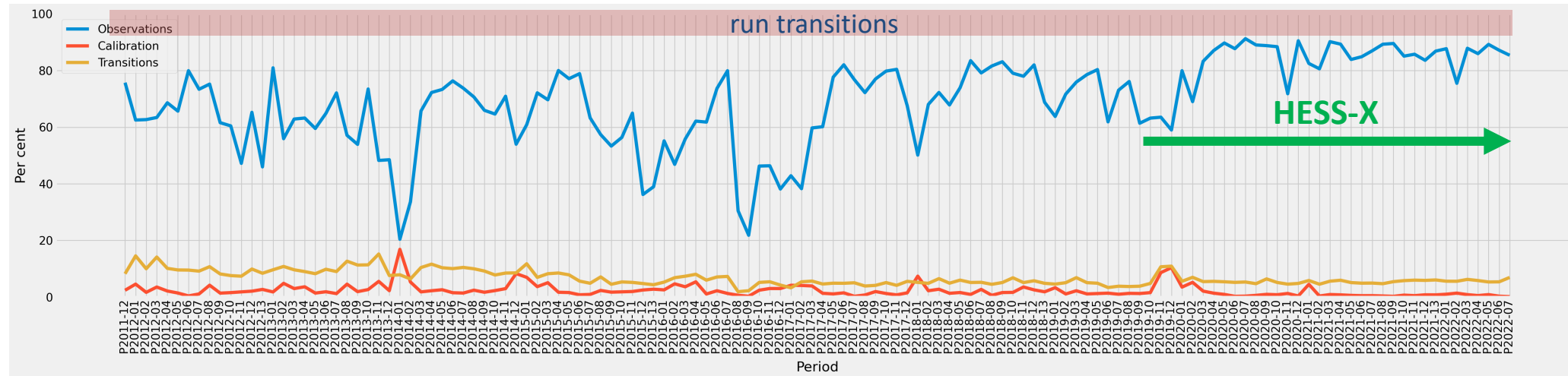
# Extending observation time through changes in procedures

- Established professional operators during Corona
- Supported by shifters from partner institutes
- Established remote control room at DESY Zeuthen; more locations coming
- Documentation, error policy, guidelines
- Consolidated calibration strategy
- Day shifters at partner institutes monitor data quality and long-term instrument behaviour  
=> Training, fast problem identification

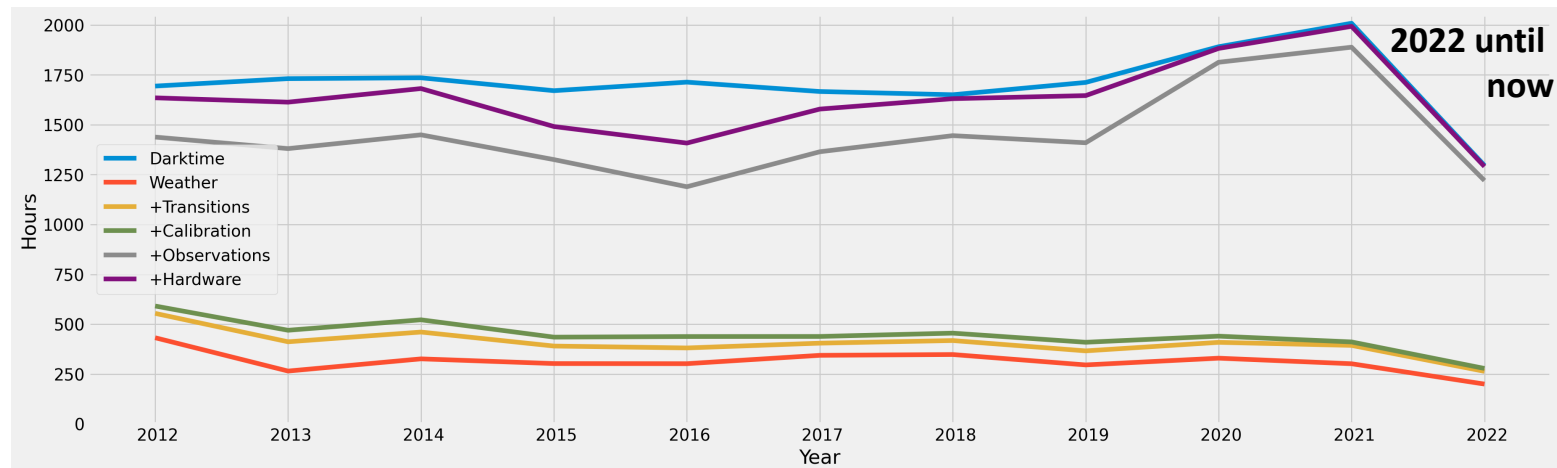




# Data taking efficiency: long-term and in 1<sup>st</sup> H.E.S.S. extension

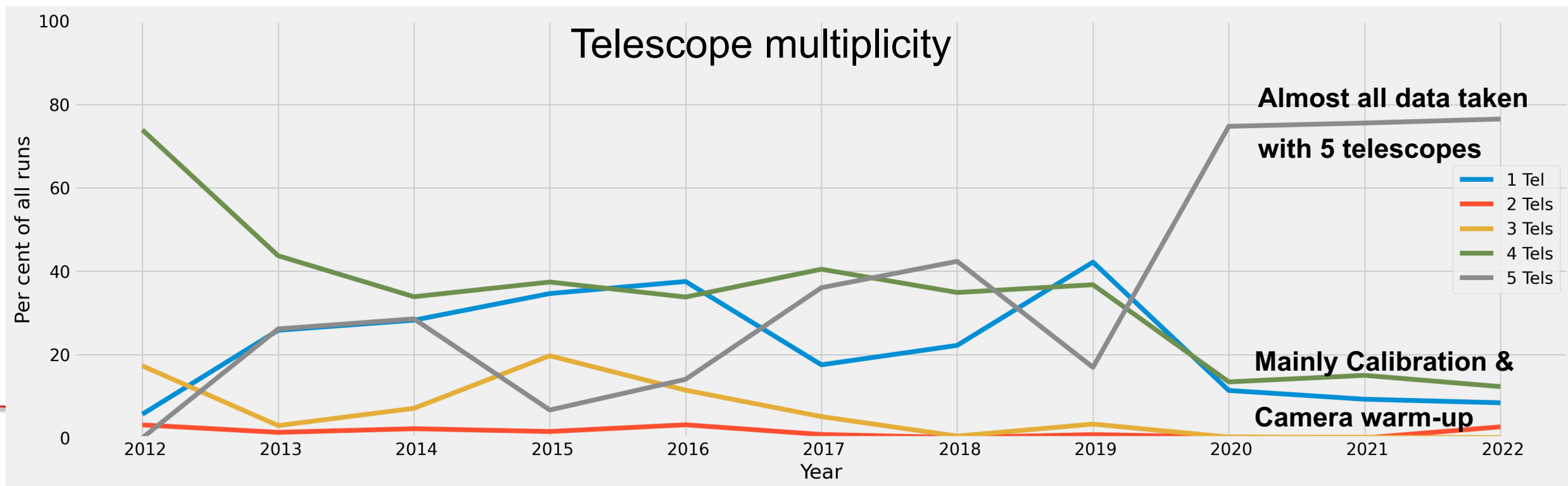
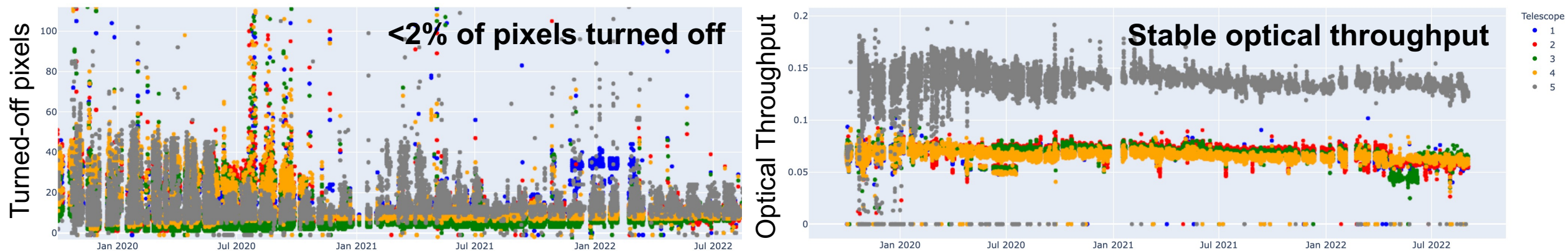


- Data taking efficiency during 1<sup>st</sup> H.E.S.S. extension at very high level (~90%, close to optimal)
- 5-telescope availability
  - 2020: 92%; 2021: 88%; 2022: 87%
- From 2012 – 2019: ~900 – 1000 hrs/yr
- In 1<sup>st</sup> H.E.S.S. extension
  - Moonlight observations: +250 hours
  - Hardware availability: +100 – 150 hours
  - Darkness definition: +70 hours
  - Procedures: +50 hours
  - Calibration optimization: +15 hours





# Long-term data quality and stability in 1<sup>st</sup> H.E.S.S. extension

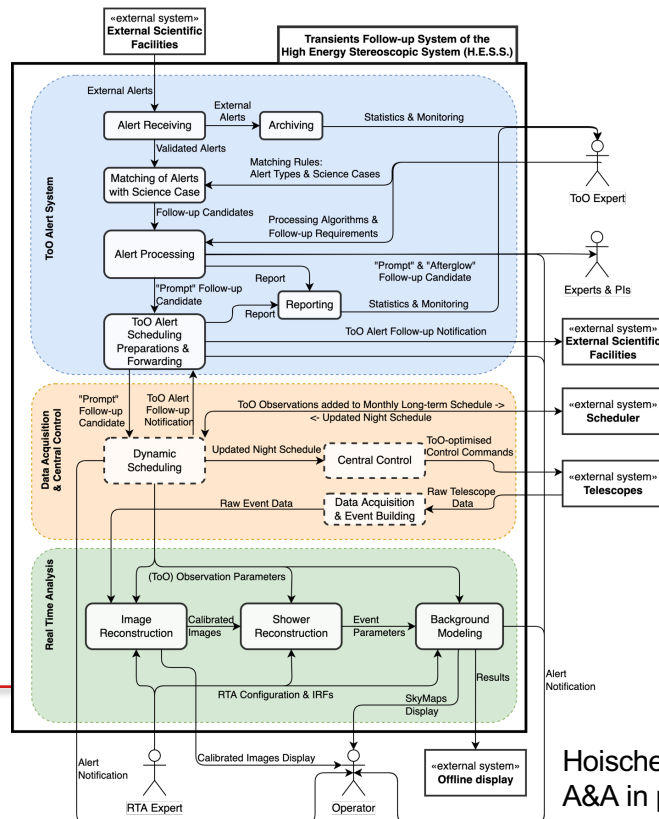




# H.E.S.S. in the context of time-domain astrophysics

- Time-domain and multi-messenger astronomy needs
  - End-to-end strategy for follow-up
  - Real-time analysis (RTA) to guide follow-ups
  - Fast data transfer for final analysis
  - Next-day analysis to alert astrophysics community
- H.E.S.S. on-site analysis capabilities
  - RTA tailored to different source types (**latency ~2 min**)
  - Fast and performant on-site calibration and next-day analysis (**latency ~8 hours**)
- Data transfer
  - Internet connection supports additional moonlight data + 2 modes of charge integration data for CT1-4
  - Implemented data transfer via internet in 2019 (**latency ~12 hours**)

## H.E.S.S. Transients System



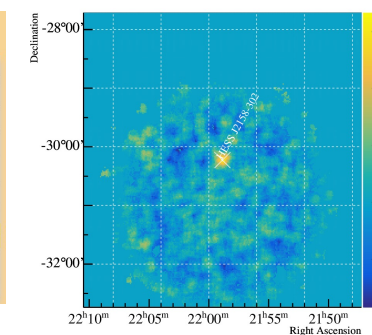
⇒ Final analysis results available in ~24 hours

### RTA night view

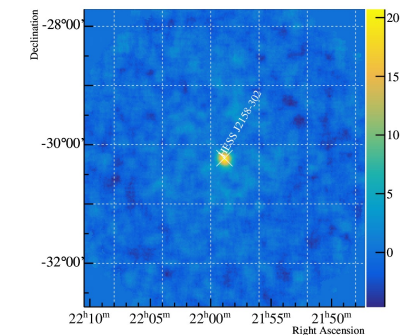
Preliminary Nightly Analysis Results (for runs that finished processing)  
see Results View for longer-term monitoring of sources

Target	Analysis	config	Runs	LiveTime	Signif ( $\sigma$ )	$\sigma/\sqrt{t}$	Excess	$\langle zen \rangle$	$\langle offs \rangle$
	RTA	RingBg std_zeta_hybrid	3	01:26	2.69	2.25	25.65	31°	0.06°
	RTA	RingBg std_zeta_hybrid	2	00:57	-0.64	-0.65	-5.91	13°	0.05°
♁	RTA	RingBg std_zeta_hybrid	1	00:28	2.27	3.30	15.83	5°	0.03°
♁	RTA	RingBg std_zeta_hybrid	1	00:22	0.98	1.62	5.52	7°	0.03°
	RTA	RingBg std_zeta_hybrid	2	00:30	1.10	1.56	5.70	21°	0.05°
	RTA	RingBg std_zeta_hybrid	2	00:57	1.04	1.06	4.61	34°	0.53°
	RTA	RingBg std_zeta_hybrid	1	00:29	<b>4.40</b>	<b>6.32</b>	<b>27.11</b>	<b>4°</b>	<b>0.35°</b>
	RTA	RingBg std_zeta_hybrid	4	01:55	2.08	1.50	10.42	34°	0.66°
	RTA	RingBg std_zeta_mono	6	02:45	-1.45	-0.87	-9.60	44°	0.49°

### RTA run significance



### RTA integrated significance



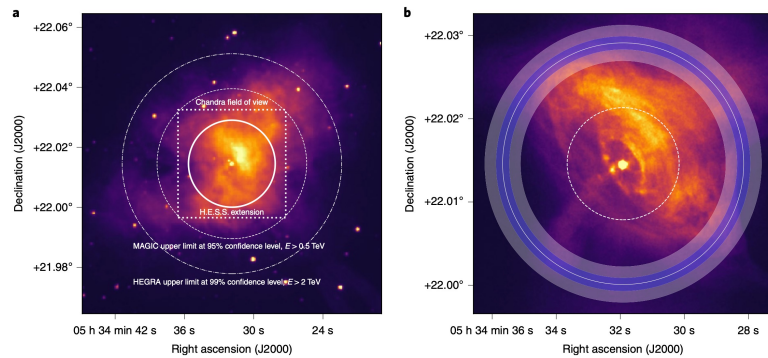


# Instrument stability and further improvements

- Continuous developments to further improve data and analysis products
- Data/MC validation chain (Leuschner et al., Gamma 2022)
- Run-wise simulations (Holler et al., APh, 123, 102491, 2020)
  - Provide instrument response functions tailored to observation and hardware
  - Application to extension measurement of e.g. Crab Nebula and Centaurus A

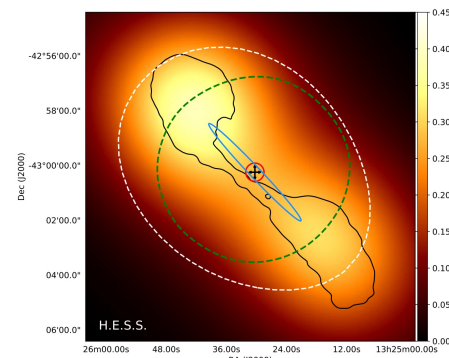
- Atmosphere monitoring and correction scheme
  - AERONET, radiosonde & aircraft data provide measure of aerosol content and atmosphere transparency
  - Correction reduces systematic errors and makes archival data available that was not considered before

Crab Nebula

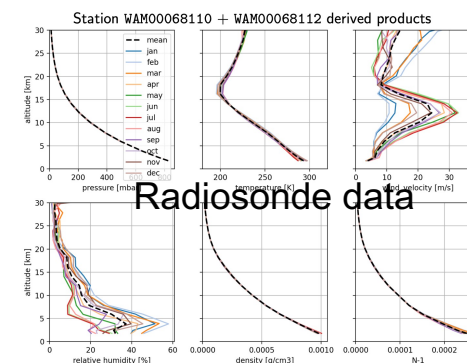
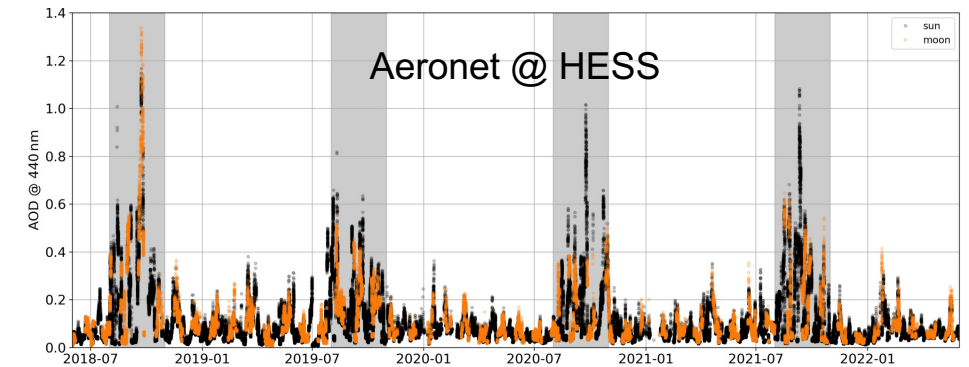


H.E.S.S. Collaboration (2020), Nature Astronomy, 4, 167

Cen A

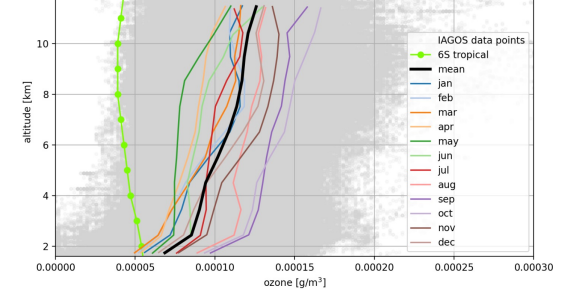


H.E.S.S. Collaboration (2020), Nature, 582, 7812



Radiosonde data

Ozone data from aircraft



Holch et al., AtmoHEAD conference (2022)

Stefan Ohm . RICH 2022 . Edinburgh . 09/2022

## Summary

- H.E.S.S. observed throughout the pandemic thanks to dedicated personnel on- and off-site
- More data (from 10% – 15% duty cycle)
- Better data (more homogenous data)
- Exciting new discoveries
  
- 2<sup>nd</sup> extension phase starting in October 2022
  - Aim at maintaining high uptime & efficiency
  - Operate CTA hardware for extended periods of time (e.g. Cherenkov cameras)
  - Further engage people in instrument and data quality => train future CTA personnel
  - Knowledge transfer to CTA via e.g. public technical notes
  - Increase recognition of individuals for technical work (cf. JENAS survey)

⇒ H.E.S.S. is operating better than ever and is prepared for continued observations!

The screenshot shows the Zenodo website interface. At the top, there is a search bar, 'Upload' and 'Communities' buttons, and 'Log in' and 'Sign up' links. The main heading is 'The H.E.S.S. gamma-ray experiment'. Below this, there is a 'Recent uploads' section with a search bar and a 'View' button. A green 'New upload' button is visible on the right. The main content area displays a dataset entry: 'Observation of the gamma-ray binary HESS J0632+057 with the H.E.S.S., MAGIC, and VERITAS telescopes - data release'. It includes the date 'August 4, 2021 (0.1.0)', 'Dataset', and 'Open Access' tags. The description mentions 'The VERITAS, MAGIC, H.E.S.S. Collaborations;' and 'Observation of the gamma-ray binary HESS J0632+057 with the H.E.S.S., MAGIC, and VERITAS telescopes - data release'. It states that the results of gamma-ray observations of the binary system HESS J0632+057 were collected during 450 hours over 15 years, between 2004 and 2019, with the H.E.S.S., MAGIC, and VERITAS telescopes. The upload date is 'Uploaded on August 9, 2021'. On the right side, there is a 'Community' section with the H.E.S.S. logo.

