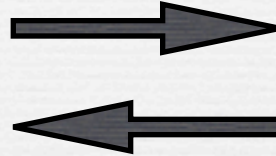


**New TeV scale  
physics**



**Cosmology**

cosmological probes of  
the EW symmetry  
breaking mechanism

- baryogenesis
- dark matter

Géraldine SERVANT  
CERN-TH & IPhT CEA Saclay



# Baryon asymmetry and the EW scale

1) nucleation and expansion of bubbles of broken phase

2) CP violation at phase interface responsible for mechanism of charge separation

3) In symmetric phase,  $\langle \Phi \rangle = 0$ , very active sphalerons convert chiral asymmetry into baryon asymmetry

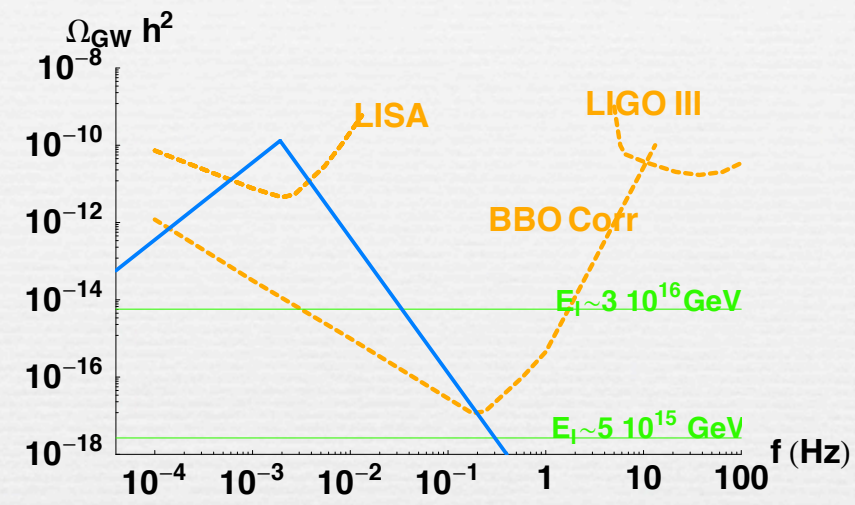
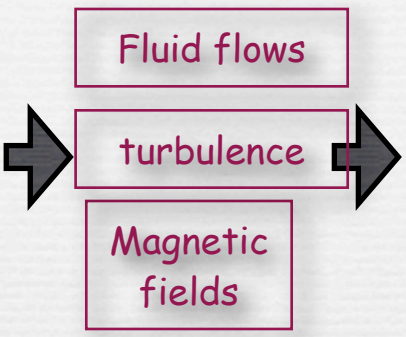
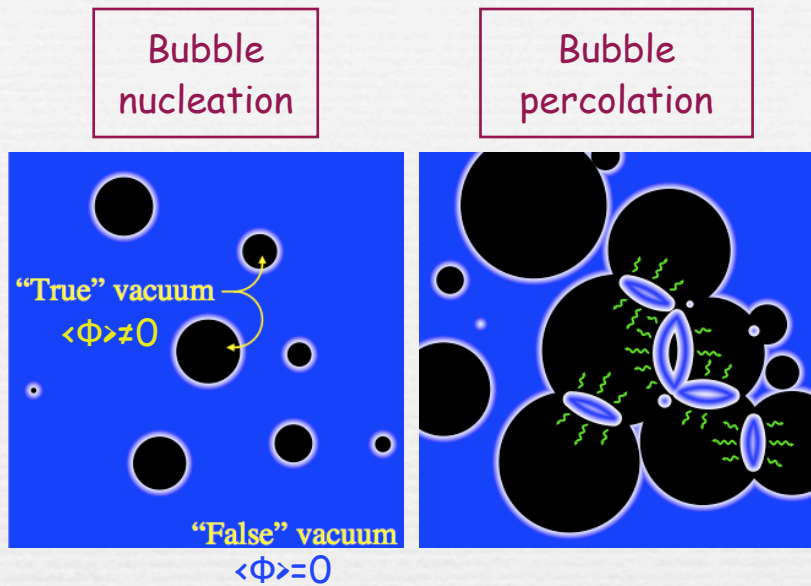
broken phase  
 $\langle \Phi \rangle \neq 0$   
Baryon number  
is frozen

Chirality Flux  
in front of the wall

Electroweak baryogenesis mechanism relies on a first-order phase transition

wall velocity is a crucial quantity

Stochastic background of gravitational radiation



$$\Omega_{GW} \sim \kappa^2$$

fraction  $\kappa$  of vacuum energy density  $\epsilon$  converted into kinetic energy

$$\kappa = \frac{3}{\epsilon \xi_w^3} \int w(\xi) v^2 \gamma^2 \xi^2 d\xi$$

fluid velocity

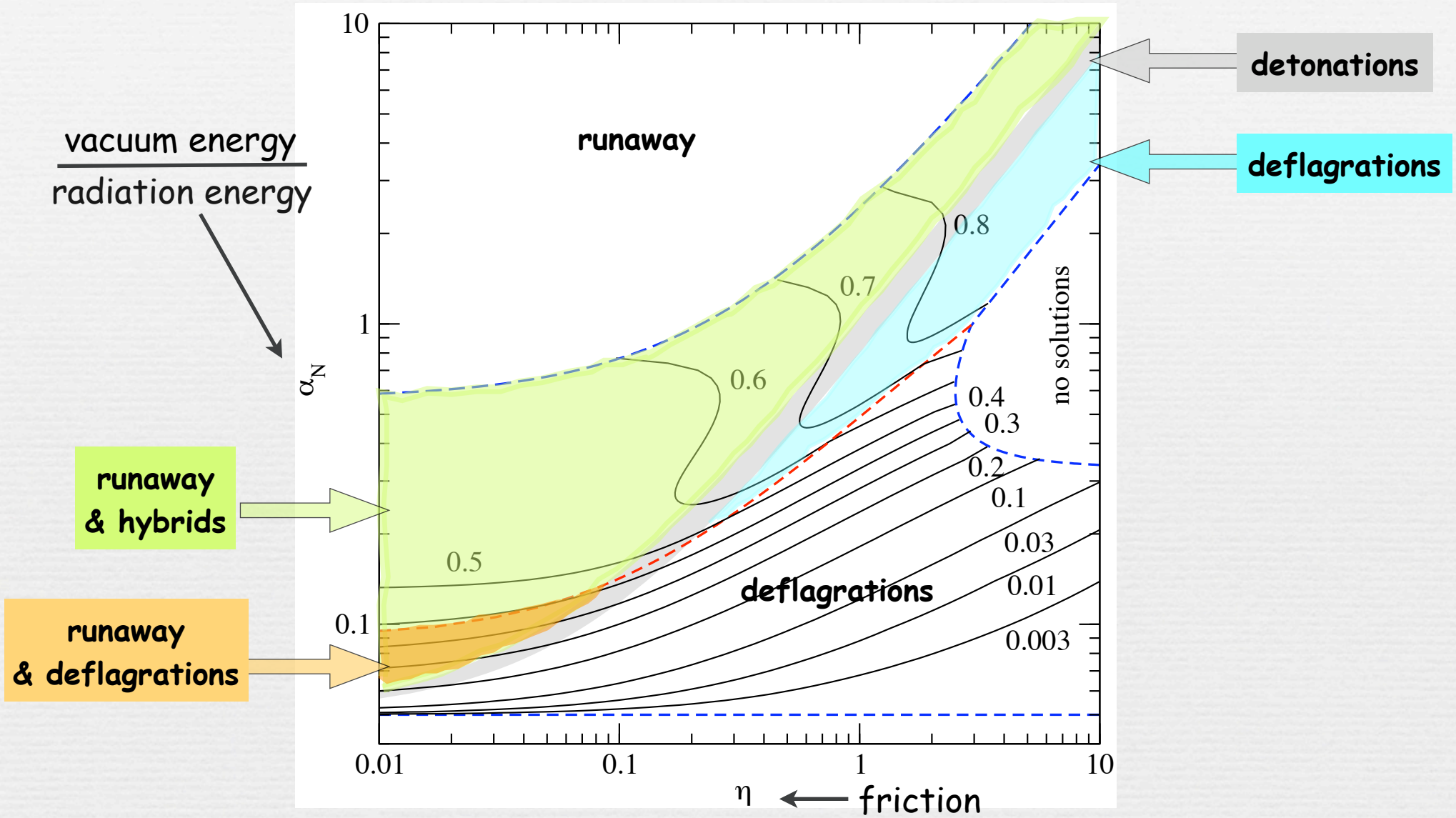
wall velocity

$$\rightarrow \Omega_{GW} \sim v^4$$



# Model-independent $\kappa$ contours

Espinosa, Konstandin, No, Servant'10



$$\eta_{\text{SM}} \sim 10^{-3}$$

$$\eta_{\text{MSSM}} \sim 10^{-2}$$

$$v \sim 0.05 - 0.1$$

Baryogenesis without ~~B~~ nor ~~L~~ nor ~~CPT~~

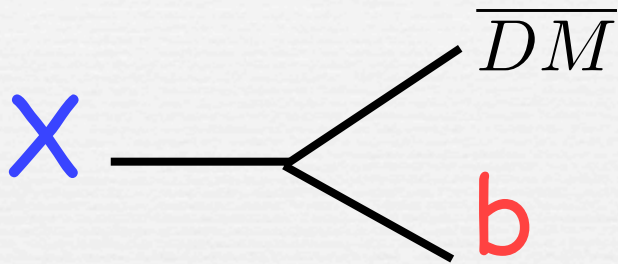
Possible if dark matter carries baryon number

Farrar-Zaharijas hep-ph/0406281

Agashe-Servant hep-ph/0411254

In a universe where baryon number is a good symmetry, Dark matter would store the overall negative baryonic charge which is missing in the visible quark sector

asymmetry between  $b$  and  $\bar{b}$  is created via the out-of-equilibrium and CP-violating decay :



$$Q_{\text{DM}}(n_{\overline{DM}} - n_{\text{DM}}) = Q_b(n_b - n_{\bar{b}})$$

out-of equilibrium and CP violating decay of  $X$  sequesters the anti baryon number in the dark sector, thus leaving a baryon excess in the visible sector

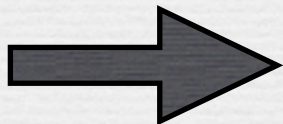
If efficient annihilation between  $DM$  and  $\overline{DM}$ , and  $b$  and  $\bar{b}$

$$\rho_{\text{DM}} = m_{\text{DM}} n_{\overline{DM}} \approx 6\rho_b \rightarrow m_{\text{DM}} \approx 6 \frac{Q_{\text{DM}}}{Q_b} \text{ GeV}$$

A unified explanation for DM and baryogenesis

$$\Omega_b \approx \frac{1}{6} \Omega_m$$

turns out to be quite natural in warped GUT models...



GUT baryogenesis at the TeV scale !



In parallel, also interested in collider phenomenology, in particular:

## Search for new physics in $t\bar{t}+\bar{X}$ production at the LHC

→  $t\bar{t}$

with Degrande, Gérard, Grojean, Maltoni, '10

→  $t\bar{t}WW, t\bar{t}W$

from  $p\bar{p} \rightarrow T\bar{T}, T\bar{t}$  where  $T \rightarrow tW$

with Contino '08

→  $t\bar{t} + \cancel{E}_T$

from  $p\bar{p} \rightarrow T\bar{T}$  where  $T \rightarrow t \text{ DM}$

with Mahbubani, in progress

→  $t\bar{t}t\bar{t}$

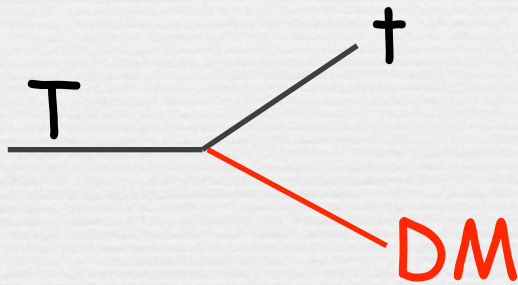
with Gauthier, in progress

# $Z_2$ versus $Z_3$ Dark Matter

Agashe et al, 1003.0899  
Mahbubani-Servant, in progress.

Most Dark Matter models rely on a  $Z_2$  symmetry. However, other symmetries can stabilize dark matter. Can the nature of the underlying symmetry be tested?

$Z_2$



$Z_3$  (+1=-2)

