

Asymmetric dark matter via Leptogenesis

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Work in collaboration with M. Blennow,
B. Dasgupta and N. Rius

Research Interests

- Neutrino phenomenology
 - Measurement of unknown neutrino parameters
 - Phenomenology of models of neutrino masses
 - Non standard neutrino interactions
- Dark Energy
 - Extracting info on the DE eos from SN and BAO
- Dark Matter

Asymmetric DM: Motivation

Baryonic Matter

baryon mass $m_B \sim 938 \text{ MeV}$

matter-antimatter asymmetry:

$$\eta_B = \frac{n_B - n_{\bar{B}}}{n_\gamma} = \frac{n_B}{n_\gamma} = (6.19 \pm 0.15) \cdot 10^{-10} \quad \text{WMAP + BBN}$$

To give $\Omega_B = 0.0455$

Asymmetric DM: Motivation

WIMP Dark Matter

WIMP mass $m_{DM} \sim 1000 \text{ GeV}$

Thermal abundance of heavy weakly interacting particle gives abundance such that:

$$\Omega_{DM} = 0.227 \text{ "WIMP Miracle"}$$

$$\frac{\Omega_{DM}}{\Omega_B} \approx 5 \text{ but } \frac{n_{DM}}{n_B} \approx 10^{-3}$$

are DM and baryon abundances and masses related?

Asymmetric DM

If **DM** is a **Dirac** fermion it can have an accidental global symmetry, **DM number X**, just like **B** and **L** in the **SM**

A **DM asymmetry** will be stable, like the baryon asymmetry, no need to impose **R** or **KK** parity

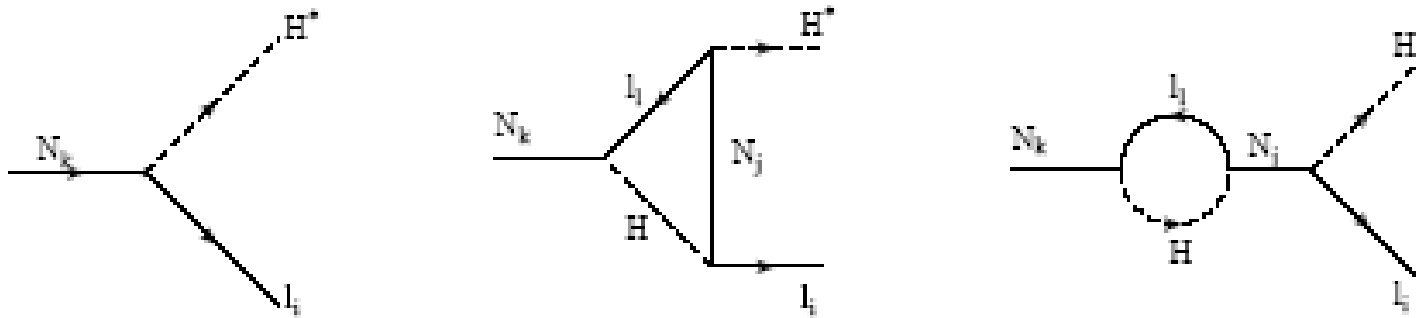
If the **DM** and baryon asymmetries have the same origin they can be similar, explaining the $\Omega_{\text{B}} - \Omega_{\text{DM}}$ coincidence

typically $m_{\text{DM}} \sim 5 \text{ GeV}$

Nussinov 1985; Barr, Chivukula and Farhi 1990; Kaplan 1992; Kuzmin 1997; Kusenko 1999; Kitano and Low 2004 and 2005; Hooper, March-Russell and West 2004; Farrar and Zaharijas 2004 and 2005; Agashe and Servant 2004; Cosme, Lopez Honorez and Tytgat 2005; Suematsu 2005; Banks, Echols and Jones 2006; Page 2007; Nardi, Sannino and Strumia 2009...

Asymmetric DM via Leptogenesis

Decay of **Majorana** right-handed neutrino N_R produces **L** asymmetry

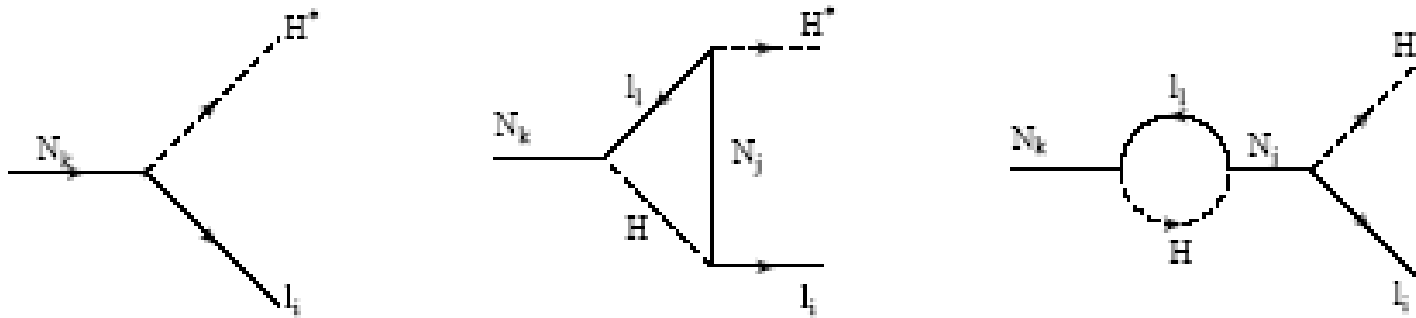


M. Fukugita and T. Yanagida 1986

SM sphaleron processes partially convert **L** into **B** conserving **B-L**

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Cant couple **5 GeV DM** to SM sphalerons we need **new** sphalerons

Asymmetric DM via Leptogenesis

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With masses similar to the SM baryons

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$SU(2)_H$ is a horizontal chiral symmetry that provides new sphalerons

$SU(2)_H$ doublets:

$$\begin{pmatrix} \mu \\ e \end{pmatrix}_R \quad \begin{pmatrix} s \\ d \end{pmatrix}_R \quad \begin{pmatrix} c \\ u \end{pmatrix}_R \quad \begin{pmatrix} x_2 \\ x_1 \end{pmatrix}_R$$

N_R and x_L are singlets to prevent the **Witten** anomaly

Asymmetric DM via Leptogenesis

N_R is a gauge singlet \rightarrow Seesaw model and L generation in its decay

$SU(2)_L$ sphalerons violate B, L and X in the direction: $\Delta B = \Delta L, \Delta X = 0$

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Symmetric DM component from thermal freeze out form

“dark mesons” below the $SU(3)_{DC}$ phase transition which decay into

SM particles via the $SU(2)_H$ interaction

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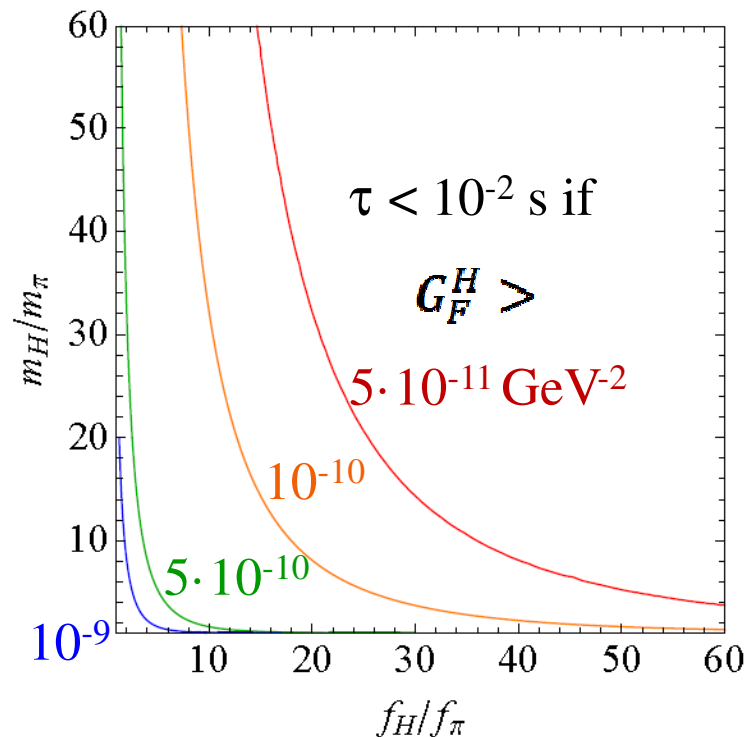
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Too weak to decay before **BBN**, need to break the symmetry in stages or to couple mainly to **2nd** and **3rd** generations

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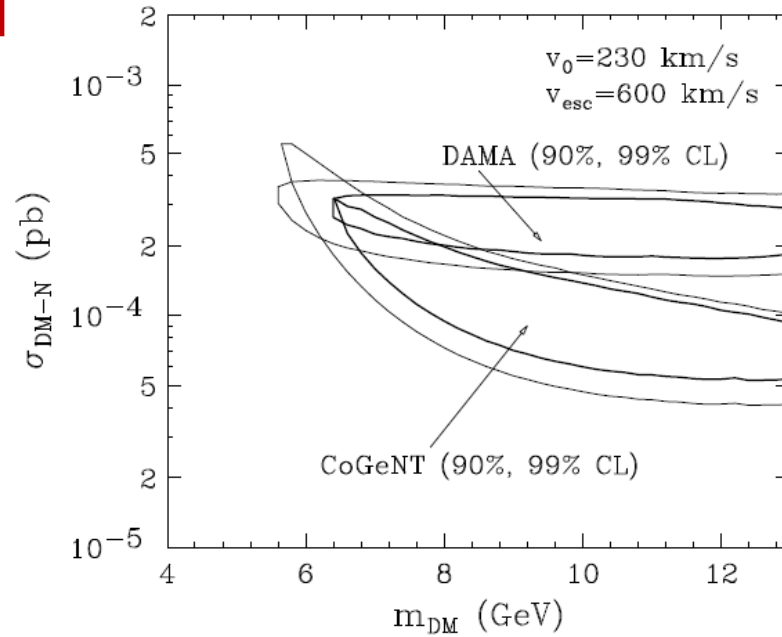
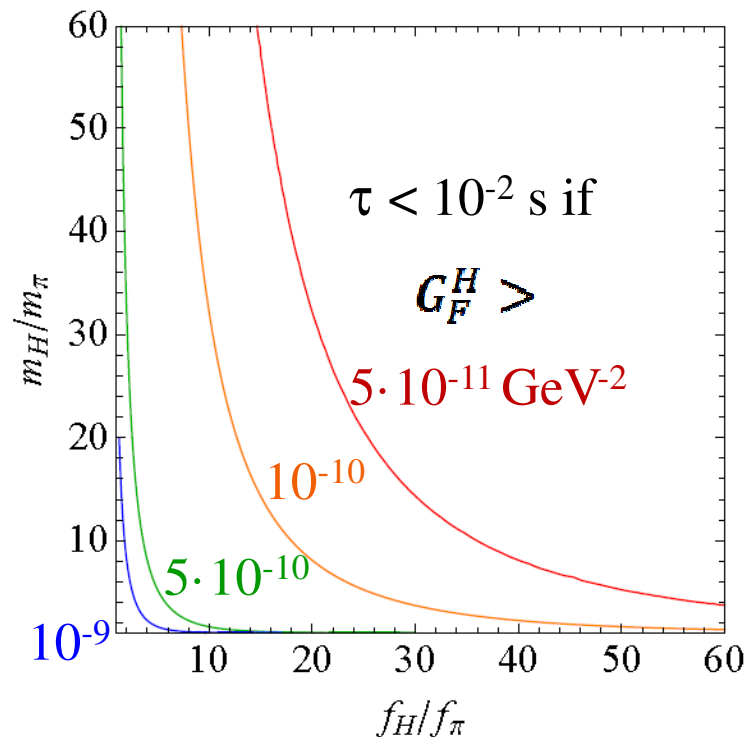
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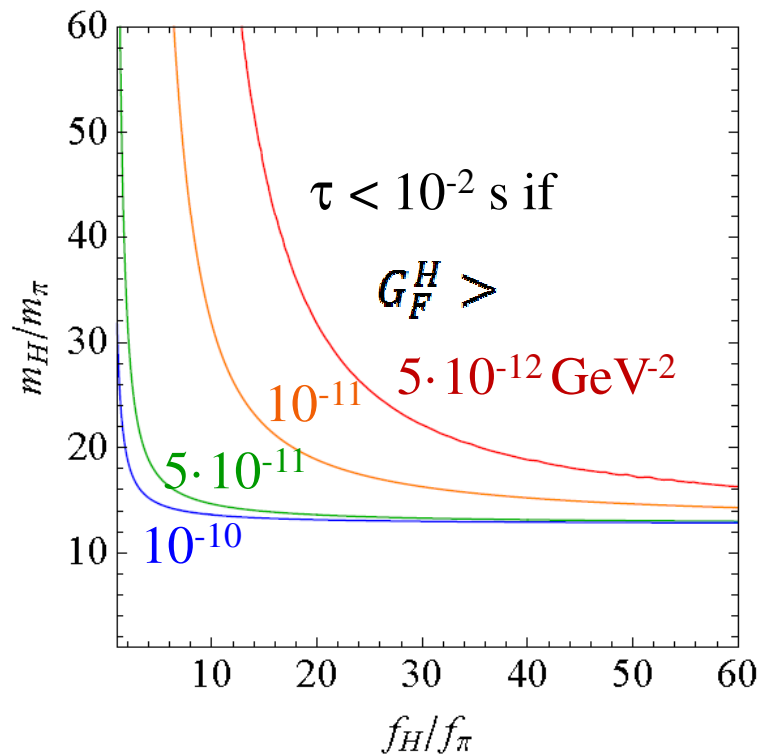


D. Hooper et al
arXiv:1007.1005

Requires $G_F^H \approx 10^{-7} \text{ GeV}^{-2}$
tension with **LEP** bounds

Coupling to 2nd and 3rd generations

If $SU(2)_H$ interaction involves mainly 2nd and 3rd generations constraints are weaker



Also provides a new source of mixing and CP violation in the B_s system that can accommodate the Tevatron dimuon anomaly with

$$G_F^H \approx 7 \cdot 10^{-11} \text{ GeV}^{-2}$$

S. C. Park, J. Shu, K. Wang and T. Yanagida arXiv:1008.4445

Conclusions and Outlook

- Extending the SM with N_R and DM fermions + $SU(2)_H \times SU(3)_{DC}$ induces asymmetric DM via leptogenesis
 - DM is stable without additional parities
 - DM mass and abundance similar to baryons
- A flavour-conserving Z' remnant of $SU(2)_H$ can have low mass and lead to signals at colliders or direct detection experiments
- If $SU(2)_H$ couples mainly to the 2nd and 3rd generations it can accommodate the Tevatron dimuon anomaly
- Can the $SU(2)_H$ symmetry breaking help with the flavour puzzle?