Searching for Dark Matter with High-energy neutrinos





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Dark Matter Evidence

 Big Bang Nucleosynthesis/ Cosmic microwave background

Galaxy Clusters



• Galaxies/ local dynamics

Dark Matter Identification

Direct Detection





Collider searches





Indirect detection





Indirect detection



Particle Physics Astrophysics/detector

$$\frac{dF}{dE} = \frac{1}{4\pi} \frac{\Gamma}{m_{\chi}} \frac{dN}{dE} \int d\Omega \int d\ell \rho_{\chi}[r(\ell)]$$

Neutrino astronomy









lceCube





IceCube 2018

Neutrino detection in IceCube

Tracks

- ν_{μ} -CC interactions + 17% ν_{τ} -CC
- Good angular resolution, bad energy resolution
- Can also count through-going events



- electromagnetic or hadronic showers
- All NC + ν_e -CC + 83% ν_{τ} -CC

- Up going events are preferred
 - Earth Shielding



IceCube data

- Two components?
 - not significant yet
- Origin of most neutrinos are still unknown





Dark Matter interpretation

 Decaying dark matter can contribute to the neutrino events







Neutrino detection in IceCube

- Tracks
 - ν_{μ} -CC interactions + 17% ν_{τ} -CC
 - Good angular resolution, bad energy resolution
 - Can also count through-going events

- cascades
 - electromagnetic or hadronic showers
 - All NC + ν_e -CC + 83% ν_{τ} -CC
 - Cherenkov light scattering in ice => bad angular resolution
- Up going events are preferred
 - Earth Shielding
 - IceCube cannot see the galactic centre!



KM3NeT

 Next-generation 1-gigaton neutrino detector in the Mediterranean







on

k matter advantage

"Downward sensitivity"



Fig. 1. Field of view of a neutrino telescope at the South Pole (top) and in the Mediterranean (bottom), given in galactic coordinates. A 2π -downward sensitivity is assumed; the grey regions are then invisible. Indicated are the positions of some candidate neutrino sources.



Distan

10⁷ ;∖ve£]_v

10⁷ [GeV]



 $E_{\rm true} \, [{\rm GeV}]$

KM3NeT letter of intent: 1601.07459

KM3NeT dark matter sensitivity

- Cascade contributes significantly the sensitivity!
- Fully test the IceCube dark matter interpretation



The Large High Altitude Air Shower Observatory (LHAASO)

- WCDA
 - 4x HAWC
- KM2A (electron + muon detectors)
 - Unprecedented PeV gamma-ray sensitivity
- WFCTA
 - Air Cherenkov telescope for cosmic ray study





The Large High Altitude Air Shower Observatory (LHAASO)

• LHAASO field of view (Galactic coordinate)



LHAASO white paper 1905.02773

The Large High Altitude Air Shower Observatory (LHAASO)

point source sensitivity



LHAASO white paper 1905.02773

LHAASO DM searches

+ Large field of view

+ 100% duty factor

+ Good enough energy and angular resolution

Dark Matter searches with LHAASO

- Ando, *Chianese*, Fiorillo, Miele + KCYN + LHAASO Collab. (*Zhe Li*)
 soon
- Both prompt and secondary gamma-ray included



Dark Matter decay searches with 1/2- KM2A

- Significant improvements in limits
- In tension with DM interpretation of IceCube data





Chianese ICRC 2021

Sun as a dark matter detector

- DM-nucleon scattering
 - Dark matter capture in the Sun
- -> Dark matter annihilation in the centre of the Sun







Solar WIMP search

- Best limit on SD cross sections
 - Hard Channels

- Both scattering and Annihilation!
- How far can neutrino telescopes reach?



The Sun as a cosmic-ray beam dump

- Solar atmospheric neutrinos
 - background to dark matter search
 - difficult to eliminate
- Solar atmospheric gamma rays
 - Detected
 - can help understand solar atm. neutrinos

CR protons Hadronic

Solar gamma rays are not understood



KCYN+ 1508.06276 Tang + 1804.06846 Linden+ 1803.05436 HAWC+ 1808.05620 Linden+ 2012.04654

- Observation higher than Model (Green: Seckel et al 1991)
 - only partially remedied in Mazziotta et al (2001.09933) and Li et ar (2009.03888)
- Hard spectral index (close to $E^{-2.2}$, harder than CR index)
- Large time variations with solar activity
- non-trivial solar surface morphology

=> Magnetic Fields strongly affect CR interactions in the surface!





Solar atmospheric neutrino 'Floor'

 KCYN+ 1703.10280, also Arguelles+ 1703.07798 and Edsjo+ 1704.02892

 Once suspected Solar atm. neutrino is detected. => hard to find weaker DM signals

- No reliable predictions
 - Magnetic field effects are complicated
 - Understand gamma rays first



Solar atmospheric neutrino search with IceCube

- Box not yet opened.
- Updated sensitivity presented in ICRC2021
- Based on non-magnetic models
- Detection could be soon



Summary

 Active scene in search for dark matter with neutrinos

- Dark matter decay search with high-energy neutrinos
 - Strong tests from LHAASO soon



- Dark matter annihilation searches from the Sun
 - Event detection could be soon
 - Need better understandings of the Sun to quantify background.

