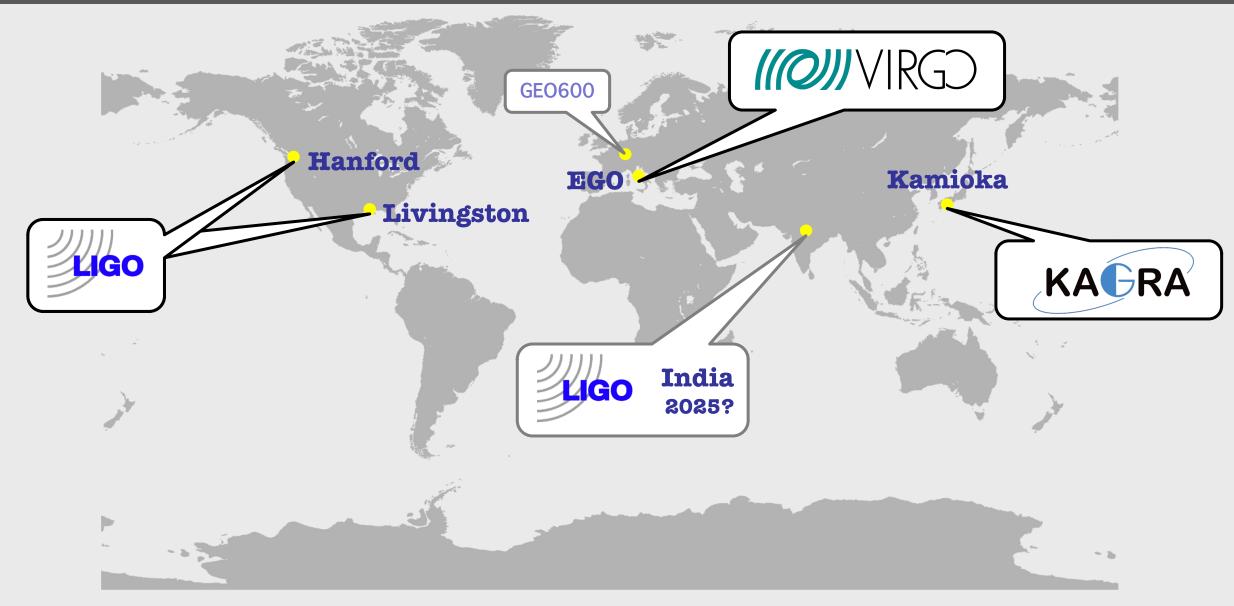
VIRGO & THE WLCG



European Gravitational Observatory, Cascina (Italy)

A WORLDWIDE NETWORK





THREE COMPUTING DOMAINS

On-site infrastructure

Plain old HTC (and some HPC)

Here's the fun

Online

- Data acquisition and pre-processing
- Instrument control
- Environmental monitoring
- ...

Offline

- Deep searches
- Offline parameter estimation
- Detector Characterization (DetChar)
- (Template bank generation)

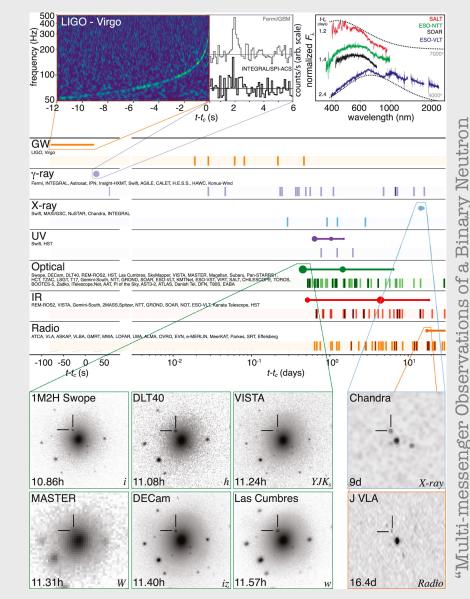
Low-latency

- Candidate search
- Sky localization
- LL parameter estimation
- Alert generation and distribution



MULTIMESSENGER ASTRONOMY

- Only GW170817 so far...
- Public Alerts are "Triggers" for ground- and space-based EM observatories
- In O3 average latency was o(1/2 hour)
- Target latency for O4 is o(minute)
- However, for some events "early warning" alerts with negative latency are possible

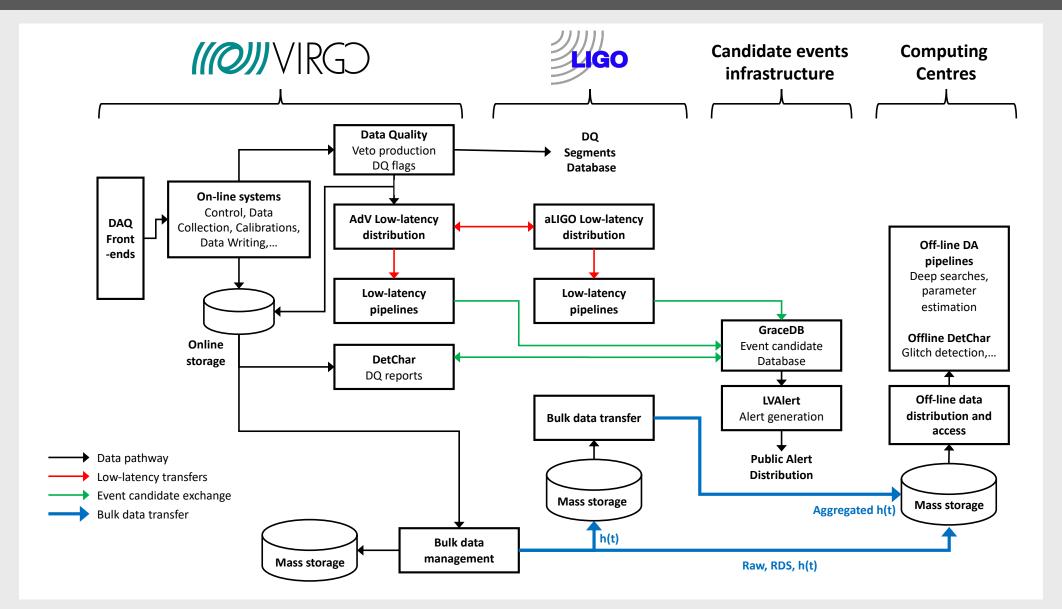


Star Merger"

B. P. Abbott *et al.* 2017 *ApJ* 848 L12



COMPLEX OVERALL DATA FLOWS







THE EU SIDE OF THE IGWN

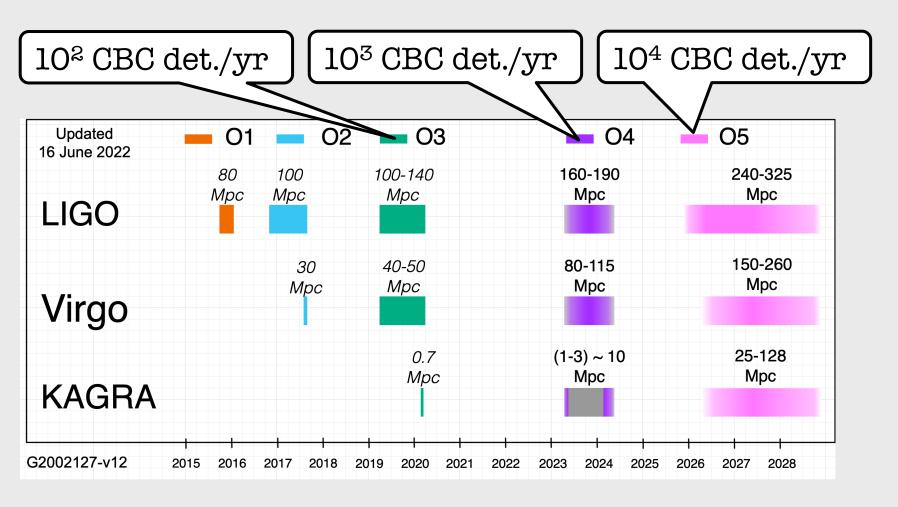




O Tier-2



LVK TIMELINE



Then:

- "Post-05" plans being prepared
- Projects for Thirdgeneration interferometers being proposed (Einstein Telescope, Cosmic Explorer)
- Growth in "computing size" (relatively) gradual



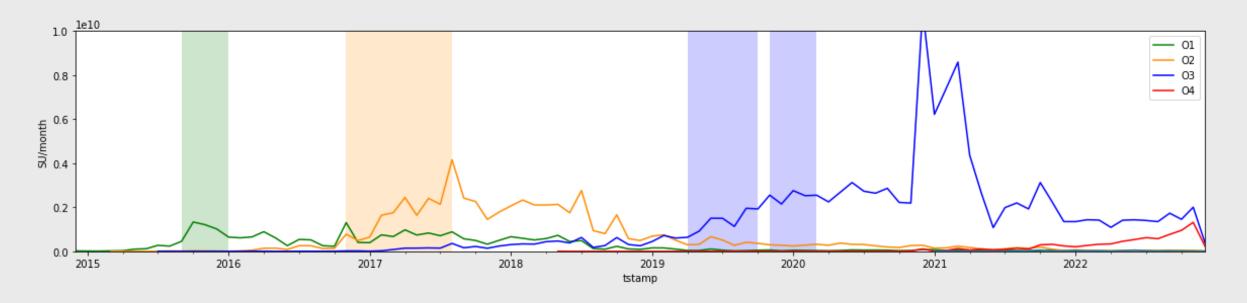
DATA AND MORE DATA

- Raw interferometer data don't grow much with increasing instrument sensitivity
 - In O3 we were writing o(1PB) per year of raw data per detector
 - h(t) (or "strain", the physics channel) + o(10⁵) control channels
 - Pre-processed data for final user analysis is more than 1 order of magnitude smaller
 - 1.5 × expected in 04 (more control channels!)
- What grows is the amount of useful scientific information embedded in the data
 - And the computing power needed to wring it out



OVERALL COMPUTING RESOURCES USAGE

- Including offline and low-latency
- Overall CPU for O3 was ≈ 7000 MHS06 Hours
- Large peaks after end of observation period (frequency-domain analyses)
- As expected, interest for older data wanes as new data become available
- Required computing power for 04 about $1.5 \div 2 \times 03$
- Overall: about 10% of an LHC experiment





THE MASTER PLAN

Need to define and deploy a common and sustainable GW computing environment

- Provide a uniform runtime environment for offline pipelines
- Full interoperability between Virgo, LIGO (and KAGRA)
- Provide scalability and move to distributed heterogeneous resources
- Adopt mainstream, widely used tools, leveraging upon HEP experience

Enter **IGWN** – the International Gravitational Waves observatories Network

• A coordination effort aimed at jointly discussing the computing policy, management, and architecture issues of LIGO, Virgo, and KAGRA.



THE MASTER PLAN CONT'D

So, for example...

- Move towards a fully distributed offline computing model
 - Based on HTCondor for workload management and CVMFS+OSDF for data distribution
 - Strong suport by OSG and HTCondor community, hope for support also by WLCG ©
- Gradually adopt Rucio for data management tasks
 - E.g. transfer of h(t) data to OSDF Origin server at UCLouvain for distribution
- Accept computing cycles from centres as in-kind contributions
 - Dicussion and update of MoUs underway



CONCLUSIONS

- GW computing started small but is coming of age
- IGWN is pushing the community towards "standard" (= HEP!) tools and architectures, most notably distributed computing
- We keep expanding our computing infratsructure to include more and more WLCG centres
- Einstein Telscope will be next player in the same timescale of HL-LHC
- Will need coordination with many WLCG sites, more practical to coordinate with WLCG itself!

