

EFFECTS OF (SUB)NUCLEON STRUCTURE IN SMALL AND LARGE COLLISION SYSTEMS

BJÖRN SCHENKE, BROOKHAVEN NATIONAL LABORATORY

DECEMBER 6 2021 21ST ZIMÁNYI SCHOOL WINTER WORKSHOP ON HEAVY ION PHYSICS





SMALL SYSTEM ANISOTROPIC FLOW

MC-Glauber (-like) models describe v_n without the need for subnucleon structure



Bozek, Broniowski, PRC88 (2013) 014903

Shen, Paquet, Denicol, Jeon, Gale, PRC95 (2017) 014906 Also see: Kozlov, Luzum, Denicol, Jeon, Gale; Werner, Beicher, Guiot, Karpenko, Pierog; Romatschke; Kalaydzhyan, Shuryak, Zahed; Ghosh, Muhuri, Nayak, Varma; Qin, Mueller; Bozek, Broniowski, Torrieri; Habich, Miller, Romatschke, Xiang; T. Hirano, K. Kawaguchi, K. Murase; ... **BJÖRN SCHENKE**

CMS Coll. PLB724, 213-240 (2013)







SMALL SYSTEM ANISOTROPIC FLOW

IP-Glasma + Hydrodynamics underestimated the v_n in p+Pb collisions

CMS Coll. PLB724, 213-240 (2013)



B. Schenke, R. Venugopalan, Phys.Rev.Lett. 113 (2014) 102301

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Large (factor ~5) difference between model and experimental data









Proton position: thick circle; Nucleon positions in Pb: thin circles

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Proton position: thick circle; Nucleon positions in Pb: thin circles

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HOW SHOULD THE ENERGY DEPOSITION GO?

Energy (or entropy) deposition $\sim (T_A T_B)^q$ is preferred:

- behave as $\sim \sqrt{T_A T_B}$ and also prefers a nucleon substructure J.S. Moreland, J.E. Bernhard, and S.A. Bass, Phys. Rev. C 92 (2015) 011901 **JETSCAPE Collaboration, Phys.Rev.C 103 (2021) 5, 054904**
- AdS/CFT based calculations also result in such a relation P. Romatschke, J.D. Hogg, JHEP 04 (2013) 048

• IP-Glasma results in the initial energy density $\sim T_A T_B$

So, $T_A + T_B$ prescription with nucleons, that works for v_n in p+A, is generally disfavored

 $\sim T_A T_B$ for round proton leads to too small fluctuations (eccentricities)

 \rightarrow subnucleon fluctuations required

• Bayesian analysis: Trento model prefers initial transverse entropy density to

G. Nijs, W. van der Schee, U. Gürsoy, R. Snellings, Phys.Rev.Lett. 126 (2021), Phys.Rev.C 103 (2021) 5, 054909

• $T_A + T_B$ disfavored by centrality dependence of v_2 in A+A $\frac{G. Giacalone, J. Noronha-Hostler,}{J.-Y. Ollitrault, Phys. Rev. C 95, 054910 (2017)}$





CONSTRAINING NUCLEON SUBSTRUCTURE

Introduce a model for the nucleon substructure within IP-Glasma

$$T_{p}(\overrightarrow{b}_{\perp}) = \sum_{i=1}^{N_{q}} T_{q}(\overrightarrow{b}_{\perp} - \overrightarrow{b}_{\perp}^{i}) \quad T_{q}(\overrightarrow{b}_{\perp}) = \frac{1}{2\pi B_{q}}$$

$$\rightarrow \cdot$$

with b_{\perp}^{i} sampled from a Gaussian with parameter B_{ac}

Constrain parameters B_{qc} and B_{q} with HERA data

Exclusive diffractive J/\Psi production in e+p: Incoherent x-sec sensitive to fluctuations

H. Mäntysaari, B. Schenke, Phys. Rev. Lett. 117 (2016) 052301 Phys.Rev. D94 (2016) 034042 also see:

S. Schlichting, B. Schenke, Phys.Lett. B739 (2014) 313-319

- H. Mäntysaari, Rep. Prog. Phys. 83 082201 (2020)
- B. Schenke, Rep. Prog. Phys. 84 082301 (2021)



(transverse momentum transfer)²



EFFECT OF SUBSTRUCTURE

B. Schenke, Rep. Prog. Phys. 84 082301 (2021)









EFFECT OF SUBSTRUCTURE

Substructure improves the description of anisotropic flow in p+A collisions with IP-Glasma



B. Schenke, Rep. Prog. Phys. 84 082301 (2021) BJÖRN SCHENKE





TRENTO ALSO NEEDS SUBSTRUCTURE The same holds true in Trento p=0 (~ $(T_A T_B)^{1/2}$)

Figure from G. Giacalone







FIREBALL SHAPE ~ PROTON SHAPE

The shape of the overlap region in p+A collisions resembles the proton's shape

Color map: Energy density distribution (arbitrary units)

Contour lines: Shape of the projectile proton (quantified using a measure of the gluon density in the proton)

B. Schenke, Rep. Prog. Phys. 84 082301 (2021)

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• Exactly match $T^{\mu\nu}$ when switching from one part to the next

B. Schenke, P. Tribedy, and R. Venugopalan, Phys. Rev. Lett. 108, 252301 (2012) B. Schenke, S. Jeon, and C. Gale, Phys. Rev. Lett. 106, 042301 (2011) S. A. Bass et al., Prog. Part. Nucl. Phys. 41, 255 (1998); M. Bleicher et al., J. Phys. G25, 1859 (1999)

Described in **detail in**

B. Schenke, C. Shen, P. Tribedy, Phys. Rev. C 102 (2020) 4, 044905 **"Running the gamut** of high energy nuclear collisions"

The term gamut was adopted from the field of **music**, where in middle age Latin "gamut" meant the entire range of musical notes of which musical melodies are composed



TRANSPORT COEFFICIENTS B. Schenke, C. Shen, P. Tribedy, Phys. Rev. C 102 (2020) 4, 044905

Transport coefficients:

Shear viscosity: $\eta/s = 0.12$

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SOME HYBRID MODEL RESULTS B. Schenke, C. Shen, P. Tribedy, Phys. Rev. C 102 (2020) 4, 044905



Parameters constrained only by HERA data (proton shape) and Au+Au data at RHIC **Underestimate** v_n in p+p and v_3 in p+Pb

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DO WE AGREE ON THE NUCLEON SIZE?

We constrained the nucleon and hot spot size from e+p collisions at HERA.

We assumed a 2D Gaussian:

Similar values for w used in the past to describe heavy ion collisions see e.g. B. Schenke, S. Jeon, C. Gale, Phys.Rev.C 85 (2012) 024901 J. E. Bernhard, J. S. Moreland, S. A. Bass, J. Liu and U. Heinz, Phys. Rev. C 94, no.2, 024907 (2016)

But some newer Bayesian analyses of heavy ion data find much larger values

J. E. Bernhard, J. S. Moreland and S. A. Bass, Nature Phys. 15, no.11, 1113-1117 (2019) $W = 0.96 \, \mathrm{fm}$

D. Everett et al. [JETSCAPE], Phys. Rev. Lett. 126, no.24, 242301 (2021)

G. Nijs, W. van der Schee, U. Gürsoy, R. Snellings, Phys.Rev.Lett. 126 (2021) 20, 202301 $_Wpprox 0.94\,{
m fm}$ **BJÖRN SCHENKE**



We found a width $w = 0.4 \, \text{fm}$ (we also use subnucleon hot spots with width $w_q = 0.11 \, \text{fm}$)

 $w = 0.9 - 1.1 \, \text{fm}$





DIFFERENT *w* **BUT DATA IS WELL DESCRIBED**



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DO WE AGREE ON THE NUCLEON SIZE? NO

Trento16

[Bass, Bernhard, Moreland 1605.03954]

Trento18

[Bass, Bernhard, Moreland 1808.02106]

Trento19

[Bass, Bernhard, Moreland Nature Phys. 15 (2019)]

IP-Glasma

[Schenke, Shen, Tribedy 2005.14682]

JETSCAPE

[JETSCAPE Collaboration 2011.01430, 2010.03928]

Trajectum

[Nijs, van der Schee, Gürsoy, Snellings 2010.15130, 2010.15134]

Figure by G. Giacalone







DOWEAGREE ON THE NUCLEON SIZE? NO



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The fits trade a large w for smaller viscosities (especially bulk)

Can we pin down w and viscosities individually?

Figure by G. Giacalone







CORRELATION OF $[p_T]$ WITH v_2 P. Bozek, Phys. Rev. C 93, 044908 (2016); B. Schenke, C. Shen, D. Teaney, Phys. Rev. C 102, 034905 (2020)

The correlation of $[p_T]$ and v_n fluctuations can help!

$$\mathbf{Define}\,\hat{\rho}(v_n^2, [p_T]) = \frac{\langle \hat{\delta}v_n^2\,\hat{\delta}[p_T]\rangle}{\langle (\hat{\delta}v_n^2)^2 \rangle \langle (\hat{\delta}[p_T])^2 \rangle}$$

$$\begin{split} \delta O &\equiv O - \langle O \rangle \\ \hat{\delta} O &\equiv \delta O - \frac{\langle \delta O \delta N \rangle}{\sigma_N^2} \delta N \text{ is the } \\ \sigma_N^2 & \text{A.C.} \end{split}$$

e variation of O at fixed multiplicity

Olszewski, W. Broniowski, Phys. Rev. C96, 054903 (2017)



DEPENDENCE OF ρ **CORRELATOR ON** w

G. Giacalone, B. Schenke, C. Shen, arXiv:2111.02908; ALICE Collaboration, arXiv:2111.06106 ATLAS-CONF-2021-001. https://cds.cern.ch/record/2748818?ln=en



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DEPENDENCE OF ρ **CORRELATOR ON** w

G. Giacalone, B. Schenke, C. Shen, arXiv:2111.02908



[JETSCAPE Collaboration **2011.01430**, **2010.03928**] [Nijs, van der Schee, Gürsoy, Snellings **2010.15130**, **2010.15134**]

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[Giacalone, Schenke, Shen **2005.14682**]



SUMMARY

- Energy deposition $(T_A T_B)^q$ is preferred (by data and theoretical consideration)
- With such energy deposition subnucleon fluctuations are needed to describe v_n in p+A
- Models disagree on nucleon size in A+A collisions
- The correlation of mean transverse momentum and elliptic flow can distinguish the models
- This observable should be included in future Bayesian analyses





CODES-ALL PUBLIC

- be downloaded from https://github.com/chunshen1987/iEBE-MUSIC/releases
- The official code repository of the IP-Glasma initial conditions is https://github.com/schenke/ipglasma. Here we used v1.0: https://github.com/schenke/ipglasma/releases
- MUSIC is the numerical implementation of (3+1)D relativistic viscous hydrodynamic simulations for high energy heavy-ion collisions. Its official website is http://www.physics.mcgill.ca/music. This work uses the version 3.0 of MUSIC, which can be downloaded from https://github.com/MUSIC-fluid/MUSIC/releases
- which can be downloaded from https://github.com/chunshen1987/iSS/releases
- We use the official UrQMD v3.4 and set it up to run as the afterburner mode https://bitbucket.org/Chunshen1987/urqmd_afterburner/src/master/
- transport models. This work uses v1.0 of the code which can be downloaded from https://github.com/chunshen1987/hadronic_afterburner_toolkit/releases
- The NEOS equation of state v0.11: https://sites.google.com/view/qcdneos/

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The iEBE-MUSIC framework integrates individual physical models which describe different stages of relativistic heavy-ion collisions. This work uses v1.0 of this framework, which can

The iSS code package is an open-source particle sampler based on the Cooper-Frye freeze-out prescription. It converts fluid cells to particle samples. This work uses v1.0 of the iSS,

The hadronic afterburner toolkit is a code package which performs analysis of particle spectra, flow observables, and their correlations using the outputs from hadronic