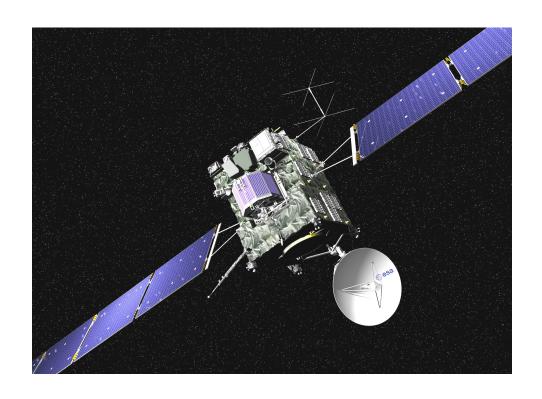


## Radiation Hardness Assurance (RHA)

Stephen Buchner Consultant

#### **Motivation**

- Several spacecraft have been affected failed due to TID and SEE, that have led to mission failure.
- The cost of a failed mission can be quite substantial – hundreds of millions of dollars.
- Adopt an approach to minimize the chance of failure.



#### Radiation-Hardness Assurance

- RHA is a method that ensures that the electronics and materials of a space system <u>perform according to their design specifications</u> during and after exposure to the space radiation environment.
- Mission requirements determine levels of radiation survivability:
  - Total Ionizing Dose (TID)
  - Single-Event Effects (SEEs)
  - Displacement Damage Dose (DDD)
- RHA deals with mission requirements, environment definition, radiation effects, part selection, part testing, spacecraft layout, radiation-tolerant design, worst-case analysis, and mitigation.
- RHA is aimed at reducing radiation-induced failures. It is not possible to eliminate risk, but it is possible to manage risk to make it acceptable.

## RHA is a Vital Activity

- RHA is just one of several activities that include modeling and testing thermal, mechanical, vacuum and electrical systems, to ensure that the mission will be successful.
- The following spacecraft requirements are impacted by radiation:
  - Reliability degree of confidence that the data are accurate
  - Availability probability that a system is operational when needed
  - Survivability probability that the spacecraft will continue operating properly during and after radiation exposure.
  - Maintainability can the spacecraft equipment be rapidly restored after suffering a radiation-induced outage

## Steps Involving RHA

- 1. A mission is proposed by scientists to an agency like NASA or ESA.
  - Study the sun, earth-observation, mission to other planets, etc
- 2. A set of requirements at various levels is established based on the mission goals, :
  - Downlink and uplink speeds
  - Data reliability
  - Data storage
  - Down time: eclipse, resetting inertial guidance, mission length, etc
  - Size, Weight and Power (SW&P) requirements
- 3. A radiation effects engineer (REE) is assigned to the project at the outset.
  - This should always be the case but is frequently not done due to budgetary constraints or lack of appreciation of the role of the REE.

## Steps Involving RHA

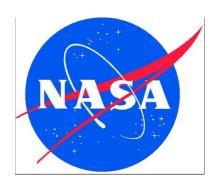
- 4. The radiation environment is established based on orbit, launch date, mission duration and shielding. The result are particle spectra to which the active parts will be exposed
- 5. Handle failure modes in parts due to radiation (TID, SEE and DDD) and calculate the part's survivability
- Parts are selected by designers for each subsystem that meet the operational requirements
- Selected parts are evaluated by REE regarding whether their TID, DD and SEE levels meet mission requirements. This is done by first scouring data bases

## Steps Involving RHA

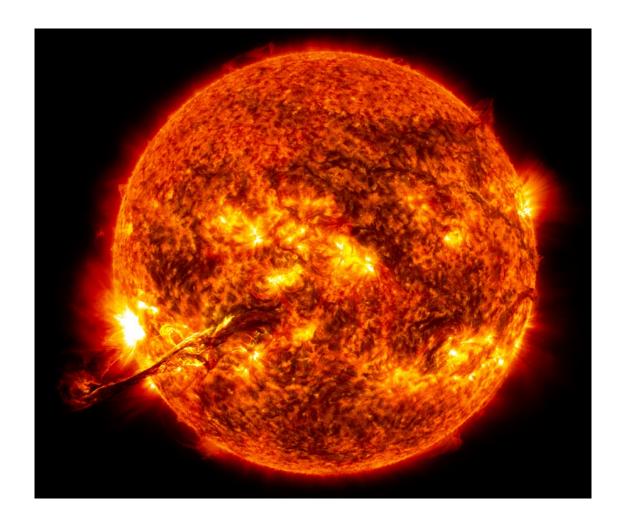
- 8. Radiation testing is performed on parts for which there is no radiation data. REE must write test plan, do the testing, and perform analysis of results
- 9. Mitigation is suggested for parts that don't meet requirements
- 10. Replacement parts are suggested in consultation with design engineer for those that do not qualify
- 11. Final approval is given when all parts have been qualified
- 12. Anomalies in space are tracked and analyzed for future reference

# Example of Radiation Hardness Assurance Solar Dynamics Observatory (SDO)

## Solar Dynamics Observatory

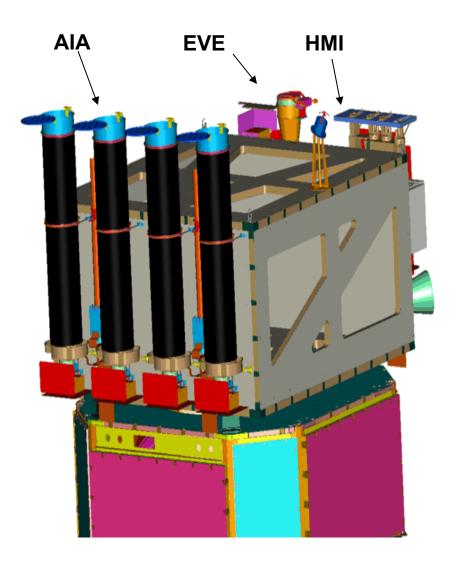


Launched 2/11/2010.



## Purpose of SDO Mission

- To study the energy sources in the sun during maximum in solar cycle:
  - The Helioseismic and Magnetic Imager (HMI) will gaze through the Sun at internal processes to help us understand the origins of solar weather.
  - The Extreme Ultraviolet Variability Experiment (EVE)
     will measure the solar extreme ultraviolet (EUV)
     irradiance to understand solar magnetic variations.
  - The Atmospheric Imaging Assembly (AIA) will study the solar coronal magnetic field and the plasma it holds to improve our understanding of how the Sun's atmospheric activity drives space weather.
- Needed congressional approval because of large cost \$850 million.
- Has been extremely successful and is still operational



## Mission Requirements – Radiation

#### 1. Mission launch date and duration (TID, SEE DDD):

- a) Launch date was February 2010 increased solar activity.
- b) 5-year mission (10-year option).
- c) Minimum on-board processing and maximum exposure time requires geosynchronous orbit over White Sands, New Mexico.

#### 2. Operation Requirement (SEE):

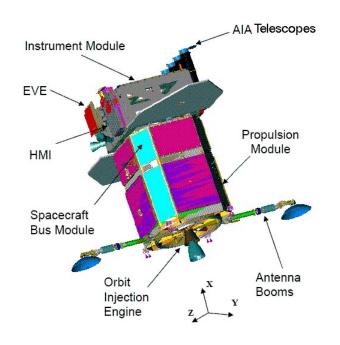
a) Must be operational 95% of the time (Down time = 2190 hours in 5 years).

#### 3. Data Requirement (SEE):

- a) Data downlink at 150 MBPS (250 DVDs per day).
- b) Data integrity must be 99.99% valid.

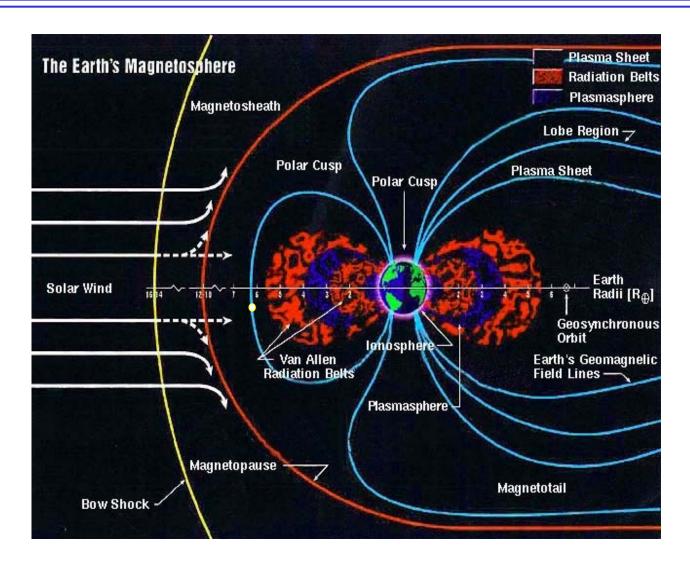
## 1. Establish the Environment

## **Establish Radiation Environment**



#### Geo is 5.45xR<sub>earth</sub>

- 1. Trapped electrons
- 2. Solar protons
- 3. Galactic cosmic rays

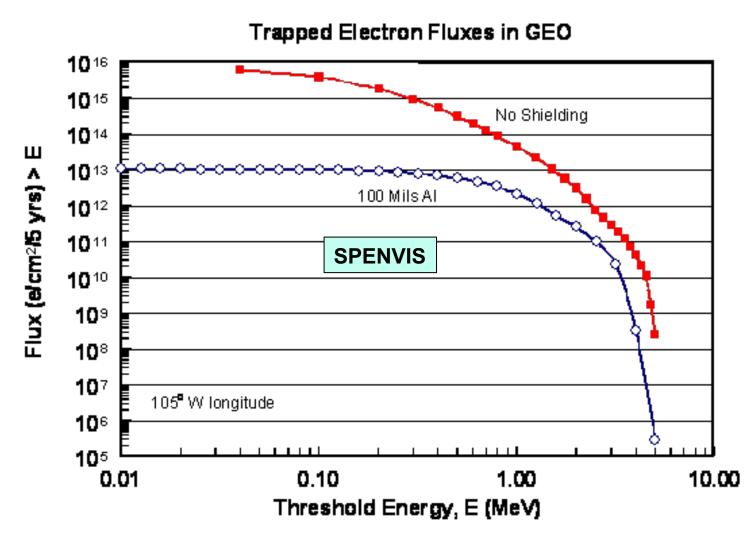


## Trapped Electron Flux at GEO

#### **Initial conditions:**

- 1. Orbit GEO
- 2. Launch date 2010 (solar cycle)
- 3. Mission duration 5-yr requirement /10-yr option
- 4. Shielding 200 mils

 Electrons are the main contributor to TID. No trapped protons.

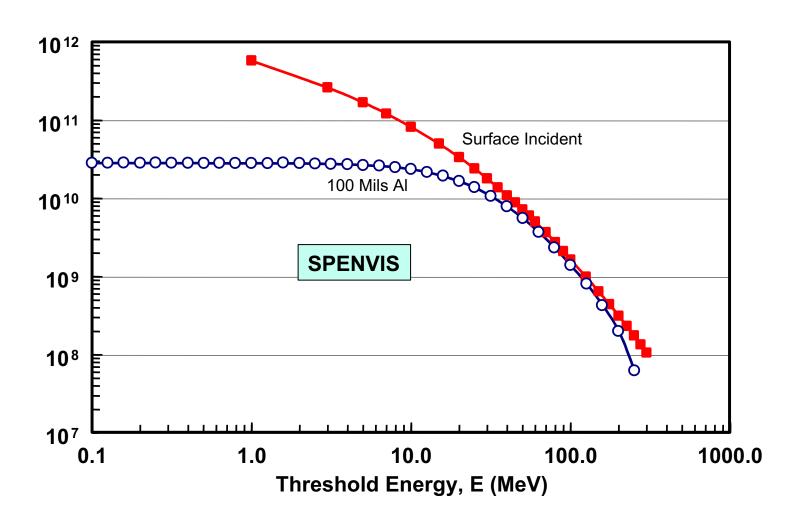


### Solar Proton Flux at GEO

- Solar protons are the main contributor to DDD, especially in optical imagers such as those on SDO.
- Also contribute to TID for thick shielding.

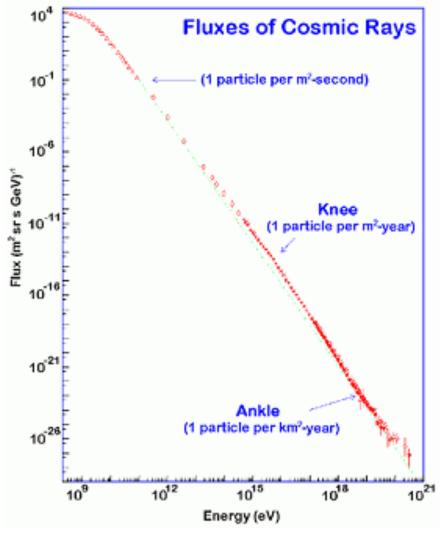
Flux (p/cm $^2$ /5 yrs) > E

 Also contribute to SEEs via nuclear interactions in sensitive parts.



## Cosmic Ray Flux - SEE

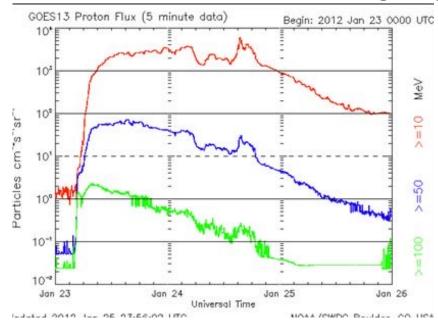
- Cosmic rays are the main contributor to SEEs.
- Negligible contributions to TID and DDD



**Varies with Solar Cycle** 

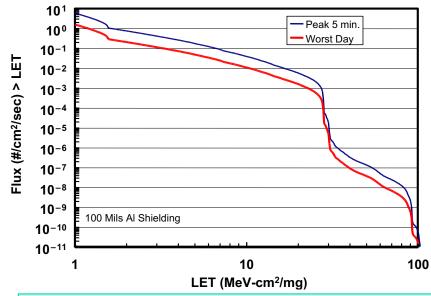
## Operate Through.....

#### Would like to continue taking images of the sun during a solar storm



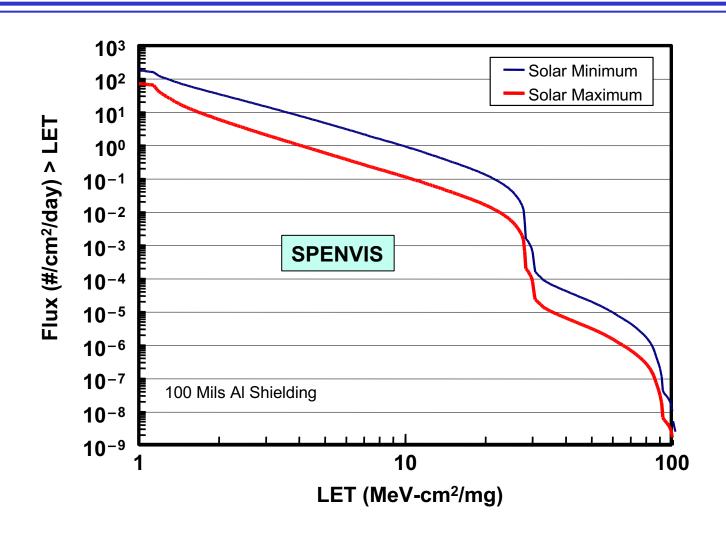
- Solar wind consists primarily of ionized hydrogen (electrons and protons) – 92% and helium – 8%, and trace amounts of heavier ions
- The wind varies continuously by a small amount.
- During a solar storm, flux observed to increase by 5 orders of magnitude



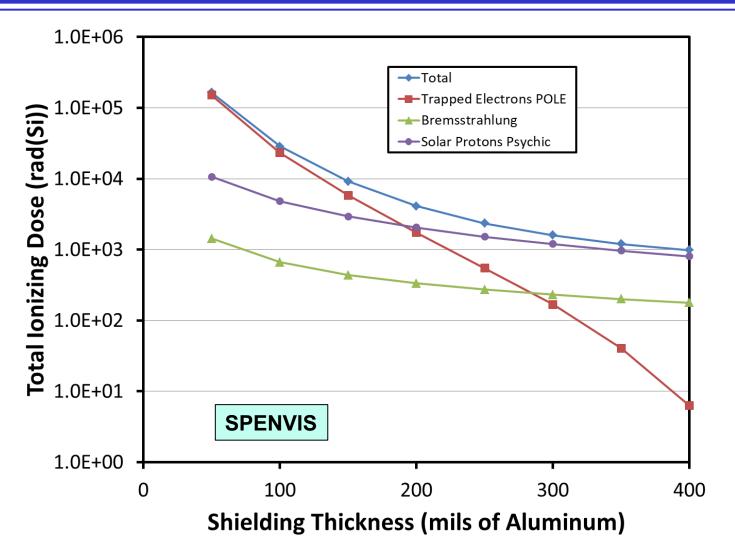


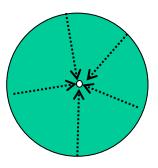
 Calculate spectra during peak 5 minutes, worst day, and worst week

## Solar Wind Affects GCR Flux – at GEO



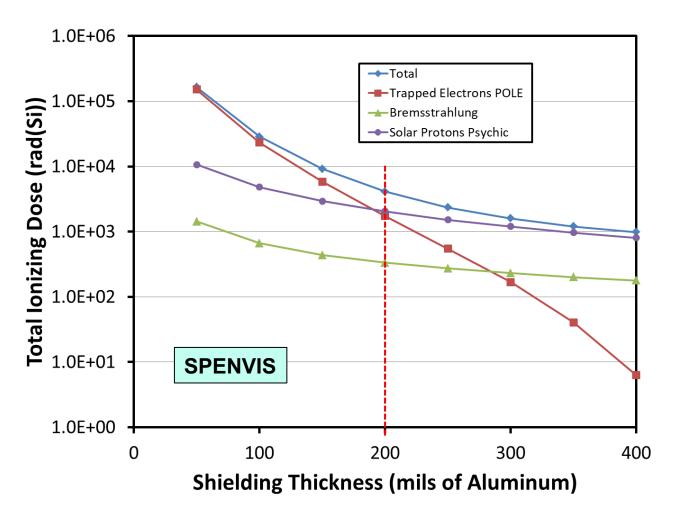
## Dose-Depth Curves – 1 Year at GEO





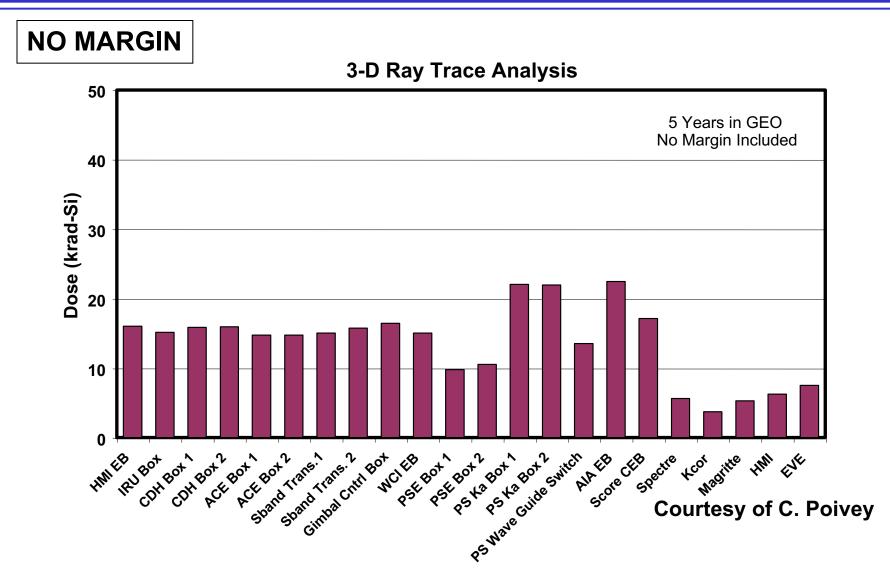
- Dose at the center of an aluminum sphere.
- Calculation done before structure of spacecraft finalized.
- For more accurate estimation of dose, use a program like NOVICE
- At low shielding thickness, dose dominated by trapped electrons, and at large shielding thickness, dose dominated by solar protons

## **Initial TID Level**



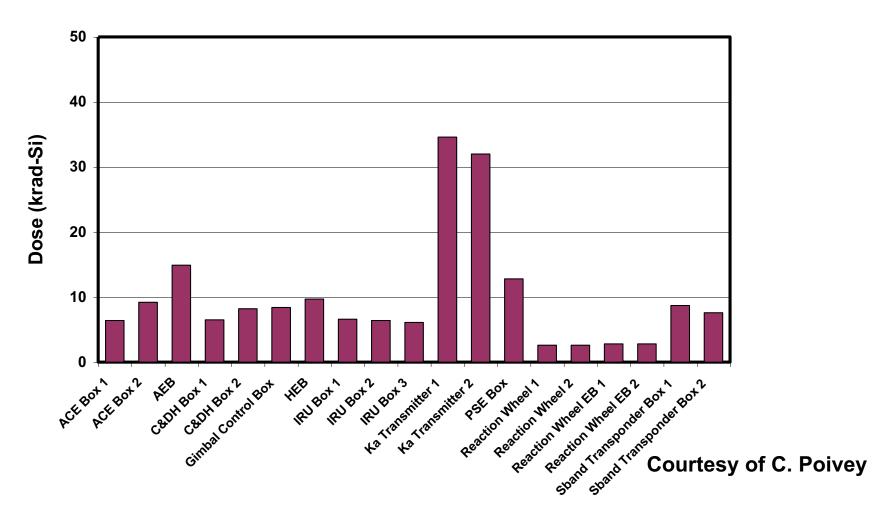
TID = 40 krad(Si) over 5 years including margin of 2x

## Final TID Levels



### Final TID Levels

#### MARGIN OF 2 USING ACCURATE SPACECRAFT MODEL and NOVICE



## 2. Design Engineer Proposes Parts and Radiation Effects Engineer Evaluates them

#### Parts are Selected

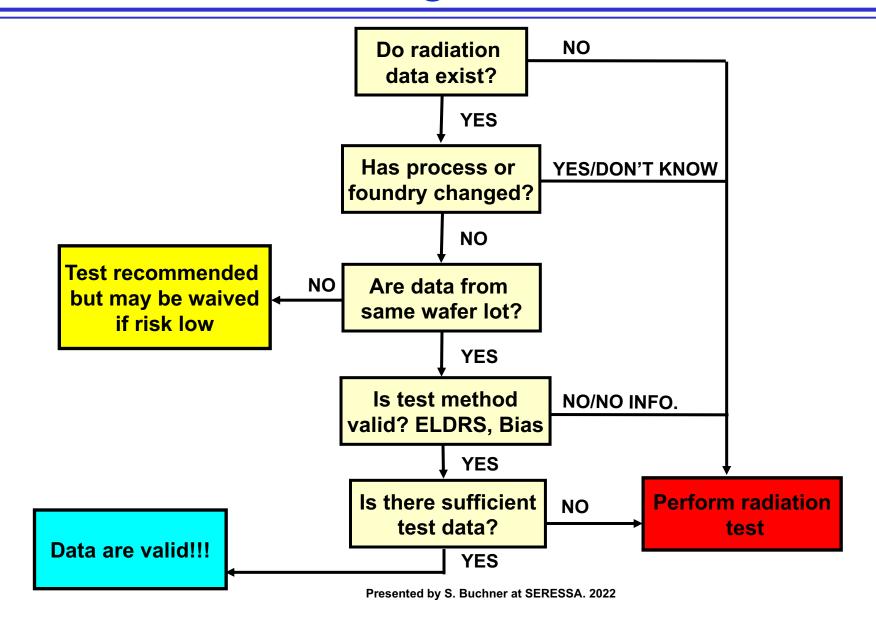
#### The Design Engineer

- a) Provides a list of proposed parts, hopefully with radiation effects in mind.
- b) Must be done in a *timely fashion* due to long lead times when ordering some parts

#### The Radiation Effects Engineer

- Evaluates proposed parts to determine whether they comply with the mission by consulting data bases
- b) Checks to see whether there are *radiation-hardened versions* of the parts available
- c) If not, orders sufficient parts from same wafer or lot/date code for radiation testing.
- d) Suggests a different part

## **Evaluating Parts for TID**



#### Parts are Evaluated

#### 1. Guaranteed radiation hard:

a) If a part is purchased from a vendor on the QML (Qualified Manufacturer's List), and the guaranteed radiation specifications meet those of the mission, then accept.

#### 2. If there are data available on the part:

- a) Was the data taken according to specifications?
- b) Is it from same wafer or lot/date code?
- c) Was the data taken less than 5 years prior?
- d) Do you trust the organization taking the data?
- e) Does the part meet specifications with appropriate design margins?

#### 3. If there are no data available:

- a) Test for TID, DDD and SEE.
- b) Consult with the design engineer about a replacement part.

## Testing is Performed

#### 1. Sufficient parts are purchased:

- a) Need sufficient parts for radiation testing, destructive physical analysis and usage requirements – cost could be an issue
- b) At least 12 for TID, 3 for SEE

#### 2. TID Testing is carried out:

- a) Gamma-ray testing at a 60Co source
- b) High dose rate vs low dose rate
- c) Biased vs unbiased

#### 3. Displacement damage in optoelectronic devices:

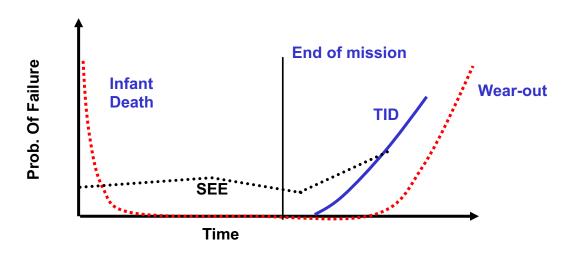
- a) Proton or electron accelerator or neutron reactor flux and energy
- b) Radioactive after exposure

#### 4. Single-event effects:

- a) Proton beam
- b) Heavy ion beam
- c) Pulsed laser light

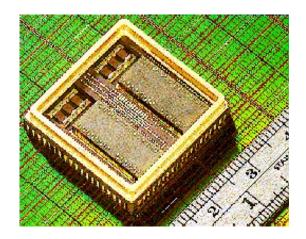
#### **Observations**

- 1. Most non-radiation failures follow "U-shaped" failure probability.
- 2. Radiation failure:
  - TID should occur after end of mission
  - Destructive SEEs should not occur
- 3. Probability of SEE varies with environment. An SEE can occur at any time, even if the probability is low.



#### **Observations**

- There are two types of ionizing dose failure
  - Parametric failure (increases in leakage current, slower operation) in some cases the part can still be used
  - Functional failure dead
- Non-destructive, non-critical SEE rates based on budgeted down time that includes:
  - Eclipses,
  - Instrument calibration,
  - Antenna handover,
  - Momentum shedding,
  - RADIATION
- Destructive SEEs (SEL, SEB, SEGR) should have a vanishingly low probability
- Use of LOT/DATE code does not guarantee all parts are the same, especially for COTS



Stacked devices and hybrids can present a unique challenge for review and test

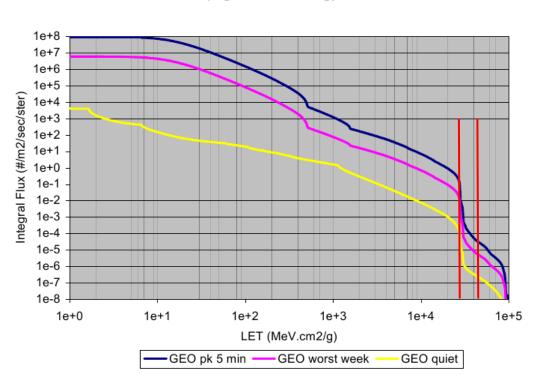
## 3. Determine SEE Requirements Error Rates

## SEE Requirements for SDO

#### Single-Event Minimum LET

- Non-Destructive (Suggested LET<sub>0</sub> > 36 MeV.cm<sup>2</sup>/mg)
  - Single Event Upset (SEU),
  - Single Event Transient (SET),
  - Single Event Functional Interrupt (SEFI).
- Destructive (Suggested LET<sub>0</sub> > 80 MeV.cm<sup>2</sup>/mg)
  - Single Event Latchup (SEL)
  - Single Event Burnout (SEB)
  - Single Event Gate Rupture (SEGR)

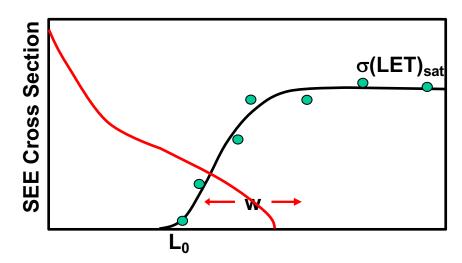
#### CREME96 GCR LET Spectra (for 1AU) (1 g/cm<sup>2</sup> shielding)



## For Error-Rate Calculation σ(LET)

#### If the error rate is required:

- ☐ Obtain cross-section vs LET from **literature**
- $\Box$  If no data are available perform accelerator testing to obtain  $\sigma(LET)$
- ☐ Fit the data with a Weibull curve to extract out four parameters ( $L_0$ , w, s,  $\sigma(LET)_{sat}$ )



#### **Linear Energy Transfer**

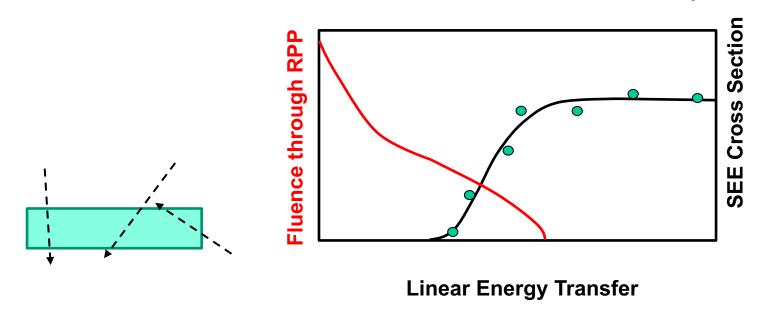
#### **Integral Weibull Curve**

$$\sigma(LET) = \sigma(LET)_{sat} \left( 1 - \exp\left( -\left( \left( \frac{(L - L_0)}{W} \right)^s \right) \right) \right)$$

## Obtain σ(LET) for Error Rate

#### • If the error rate is required:

- □ Use a program, such as SPENVIS, to calculate error rate using integral particle fluence *f*(L,φ,cos(Θ)) and dimensions of sensitive volume (x,y,z).
- ☐ Determine whether error rate meets requirements



## Conditions on LET<sub>th</sub>

#### • <u>LET<sub>th</sub> > 80</u>

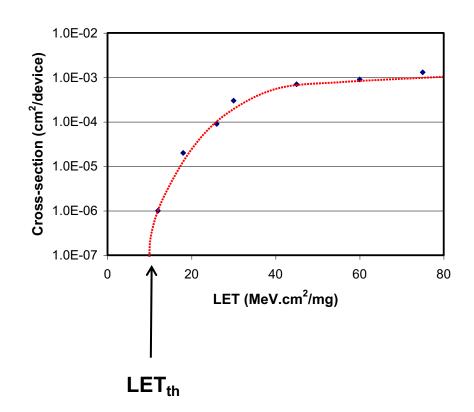
SEE risk negligible, no further analysis needed

#### • 80 > LET<sub>th</sub> > 15

 SEE risk moderate, heavyion induced SEE rates must be calculated.

#### • <u>15 > LET<sub>th</sub></u>

 SEE risk high, heavy ion and proton induced SEE effects and rates must be calculated.



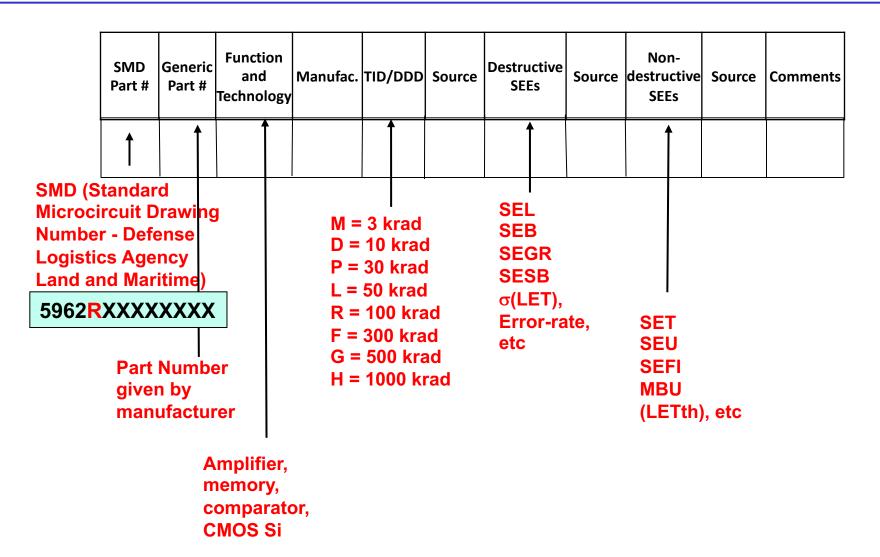
## 4. Examples of Parts Scrubbing

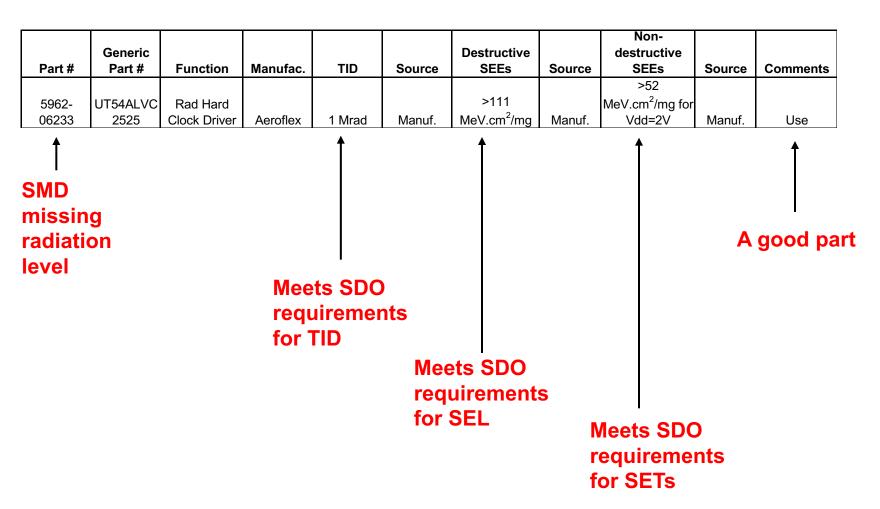
Total number of active parts can run into the hundreds

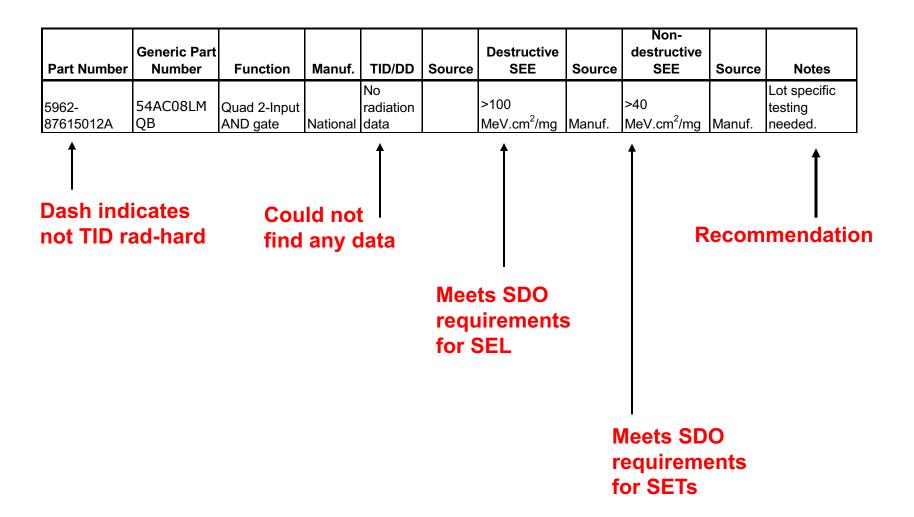


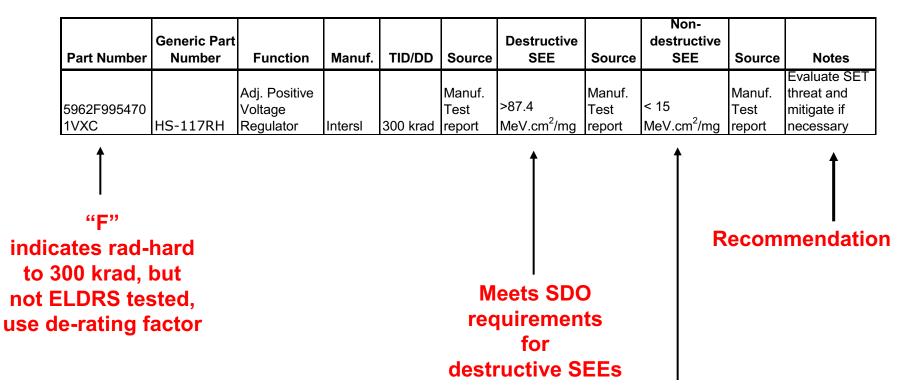
#### Sources of Radiation Data

- IEEE Transactions on Nuclear Science
  - https://ieeexplore.ieee.org/Xplore/home.jsp
- 2. NSREC Data Workshop Proceedings also published by IEEE
  - https://ieeexplore.ieee.org/Xplore/home.jsp
- RADECS Data Workshop Conference Proceedings
- 4. NASA JPL Radiation Effects Database
  - https://parts.jpl.nasa.gov/radiation-effects/
- 5. NASA GSFC Radiation Effects & Analysis
  - https://radhome.gsfc.nasa.gov/radhome/RadDataBase/RadDataBase.html
- Defense Logistics Agency Land and Maritime (DLA)
  - https://landandmaritimeapps.dla.mil/offices/doc\_control/Resources.aspx
- 7. European Space Agency (ESA) European Space Components Information Exchange System (ESCIES).
  - https://escies.org/webdocument/showArticle?id=227&groupid=6
- PMPedia
  - https://pmpedia.space/
- Manufacturer's data sheets on the WWW

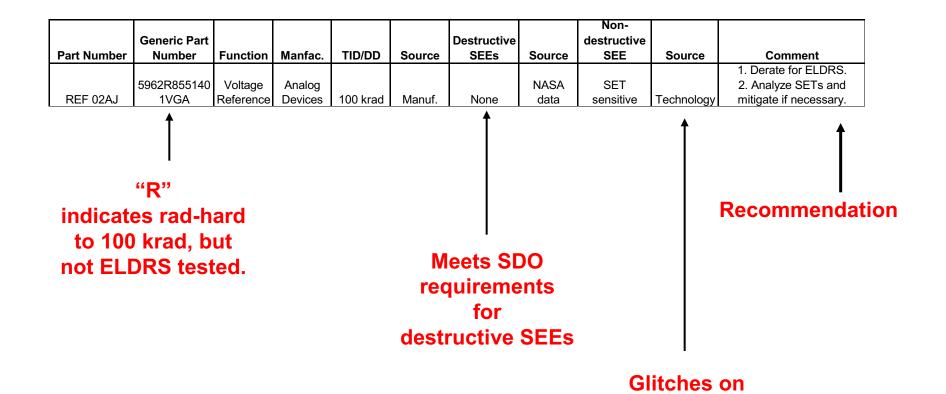




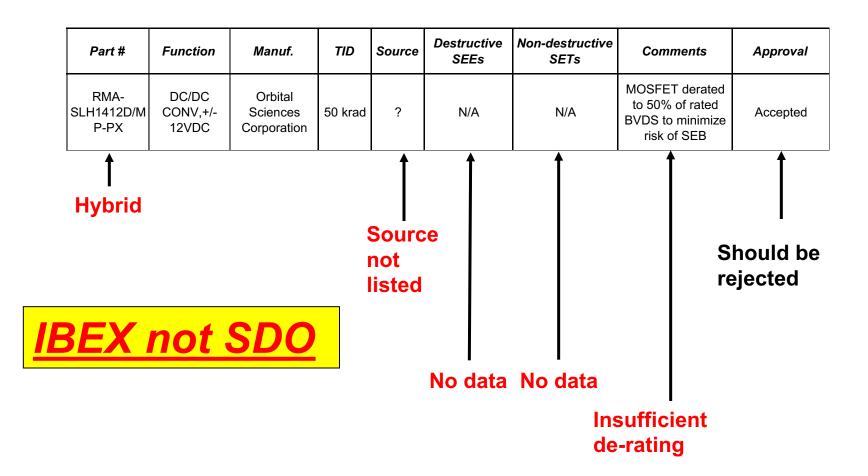




Does not meet SDO requirements for SETs



output. Must know amplitude and width



# Final Approval is Given

• If the parts meet radiation specifications, the radiation effects engineer approves.



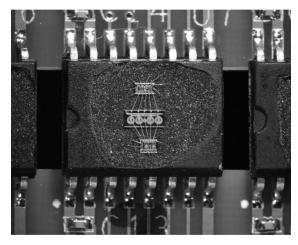


## Innovative Approaches to Testing

# Screening Parts Using SEL

#### Replace opto-isolators to save power

- MIS Mission uses 75 isolators
- Power consumption if opto-isolators are used is 10 W
- Replace with galvano-isolators to reduce power to 2 W
- Parts selected were:
  - Analog Devices: ADuM1410/12
  - Texas Instruments: ISO7240
  - NVE: IL515 and IL715
- These are COTS parts that need TID testing
- Used pulsed laser to check for SEL as an initial screen

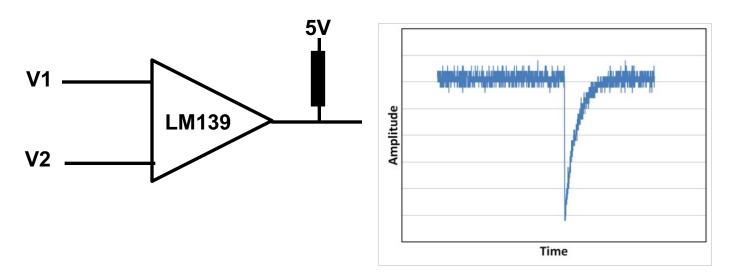


Presented by S. Buchner at SERESSA. 2022

## Single Event Test – Worst Case

#### Use a laser to measure worst-case SETs

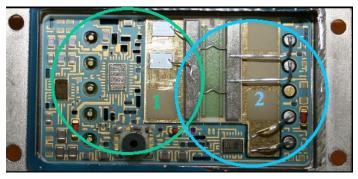
- Linear devices, such as op-amps, voltage regulators, and comparators give rise to analog SETs that depend on specific configuration.
- Cannot retest a part for each application because of time and expense.
- Pulsed laser can provide worst-case transients, i.e., in orbit, the SETs won't be worse.



# **Example of Unexpected Results**

- Solid State Power Controller (SSPC) from DDC (RP-21005DO-601P)
  - DDC replaced FET from Signetics with non rad-hard FET from IR.
  - Parts engineer suspicious and asked for testing.
  - Heavy-ion testing at Texas A&M revealed the presence of SETs causing the SSPC to switch off.
  - Pulsed laser testing revealed that the ASIC was sensitive to SETs, and that large SETs caused the SSPC to switch off.
  - Previous SEE testing by GSFC of ASIC at Brookhaven revealed no SETs.
  - Replaced DDC SSPC with Micropac SSPC
  - SEE testing successful at TAMU

Problem attributed to short range of ions at Brookhaven National Laboratory

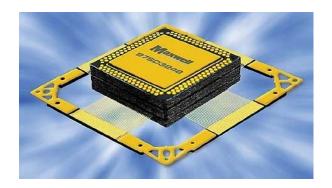


# Example of Mitigation on SDO

# SDRAM (Maxwell/Elpida) used as a temporary buffer to store data from all three telescopes prior to down-linking.

#### • SDRAM Requirement

- SDRAM suffers from SEFIs due to ion strikes to control circuitry.
- Mitigate SEFIs by rewriting registers frequently.
- At temperatures above 42 C, cannot write to SDRAM.
- Determined it was due to a timing issue in rewriting registers.
- New mitigation involves triple-voting three SDRAMs.



## Mitigation based on SEE Rate

#### Non-destructive SEEs

- If LET<sub>th</sub> for non-destructive SEEs is below 36 MeV.cm<sup>2</sup>/mg.
  - Mitigate if critical (e.g., majority vote, EDAC, filters)
  - Add watchdog timer
  - Replace if critical and cannot mitigate
  - Accept if non-critical (e.g., housekeeping)

#### Destructive SEEs

- If LET<sub>th</sub> for destructive SEEs is below 80 MeV.cm<sup>2</sup>/mg.
  - Mitigate (e.g., latchup protection circuit)
  - De-rate (COTS Power MOSFETs have  $V_{sd}$  de-rated to 35%, rad-hard Power MOSFETs to 60%)
  - Replace part if cannot mitigate

(Sometimes have no other choice but to accept part.)

## **TID Mitigation**

#### Shielding

- Use positioning judiciously to provide shielding to most sensitive devices. – ray trace analysis
- Localized shielding has little weight penalty

#### <u>Derating</u>

Operate the device at a lower voltage or a lower frequency

#### Conservative Circuit Design

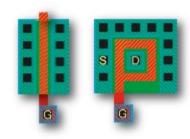
Accept a part that will fail parametric requirements but not operational

#### Extra cold spares

Unpowered devices will not suffer TID degradation (except ELDRS)

#### Radiation hardening by design

Use rad-hard by design parts to avoid rad-hard by process



# Some Thoughts

- There can be hundreds of different active parts on a spacecraft that have to meet requirements for radiation tolerance.
- Radiation effects engineers spends 95% of their time on 5% of the parts, such as FPGAs, Processors, ADCs, etc
- Generally, are not concerned with TID and SEE in resistors, inductors and capacitors.
- Many manufacturers claim a part is radiation-hard if the part has TID immunity. They completely ignore SEE.
- Lag time between deciding to test part and receiving part from manufacturer can be up to a year. Obsolescence is an issue.

# More Thoughts

- Linear bipolars must be tested for ELDRS using low dose rates with gamma rays. Testing takes a lot of time.
- CMOS parts should always be checked for Single Event Latchup.
- Some parts are expensive to test \$100K per part. May have to modify test protocol.

### **Final Points**

- The RHA approach is based on risk management and not on risk avoidance
- The RHA process is not confined to the part level, but includes
  - Spacecraft layout
  - System/subsystem/circuit design
  - System requirements and system operations
- RHA should be taken into account in the early phases of a program, including the proposal and feasibility analysis phases.