

A novel window into Dirac vs. Majorana neutrinos using $CE\nu NS$

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Based on
2110.02233

in collaboration with

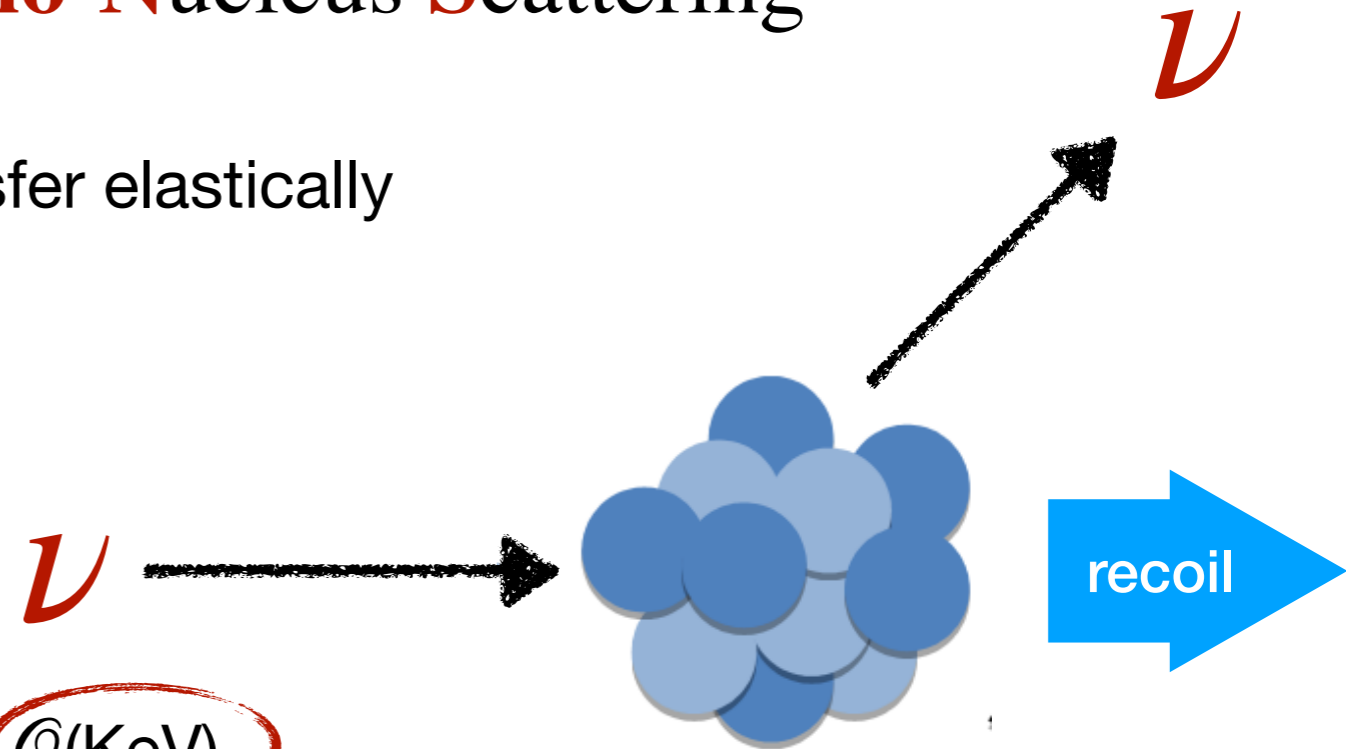
Patrick Bolton, Frank Deppisch, Kåre Fridell, Julia Harz, Suchita Kulkarni

Coherent **E**lastic **N**eutrino-**N**ucleus **S**cattering

Neutrino scatters with low momentum transfer elastically from entire nucleus

$$E_\nu \lesssim \frac{hc}{R_N} \sim \mathcal{O}(10 \text{ MeV}) \quad \text{for coherence}$$

The nuclear recoil energy: $E_r^{\text{max}} = \frac{2E_\nu^2}{M_A} \sim \mathcal{O}(\text{KeV})$



SM allowed process but hard to observe due to small nuclear recoil energy!

Akimov et al. Science 2017

First observation: Aug 2017 by

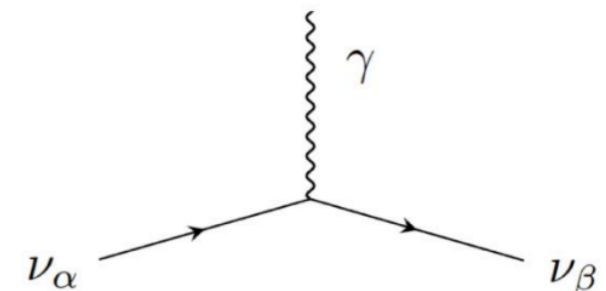
COHERENT Experiment

Threshold: **A few keV**

NUCLEUS Experiment @ CHOOZ

Threshold: **10's of eV !**

Sensitive to exotic ν interactions e.g. neutrino magnetic moment

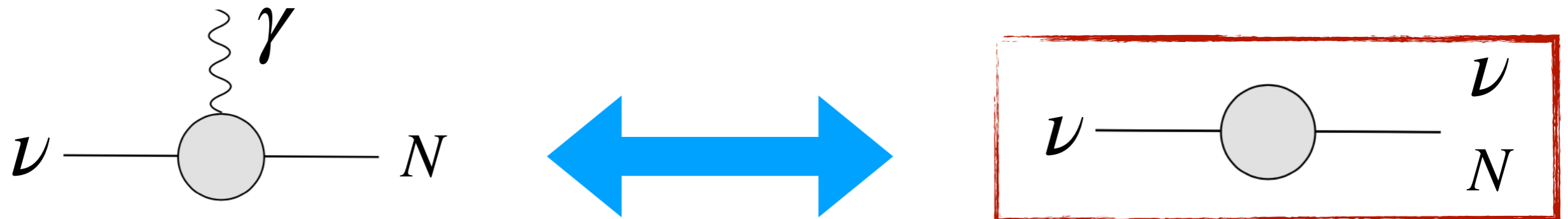


Other physics potentials by Danny and Valentina on Friday

Key ingredients

sterile neutrino + active-sterile transition magnetic moment

$$\mathcal{L} \supset \mu_{\nu N}^{\alpha} \bar{\nu}_{\alpha L} \sigma_{\mu\nu} P_R N F^{\mu\nu} + \mu_{N' N} \bar{N}' \sigma_{\mu\nu} P_R N F^{\mu\nu} + \text{h.c.}$$



correlation with active neutrino mass (more on this later!)

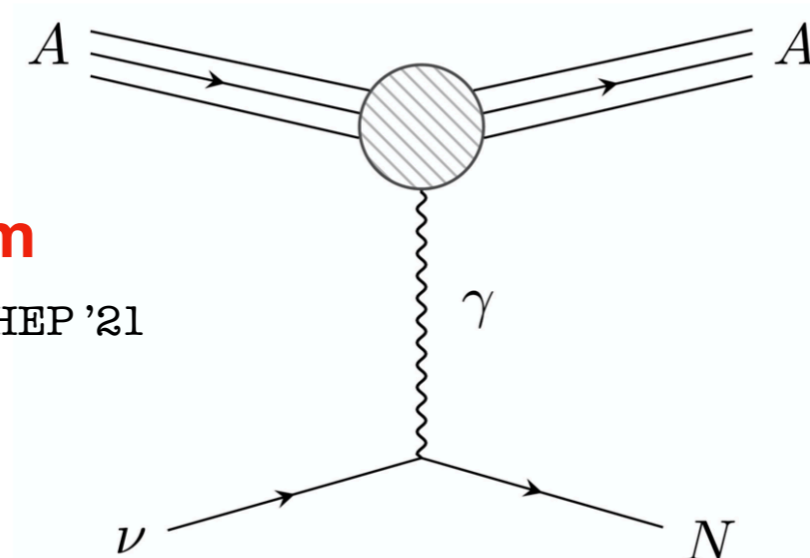
Primakoff upscattering process $\nu A \rightarrow NA$

$$\frac{d\sigma}{dt} \propto \mu_{\nu N}^2 \quad \longrightarrow \quad \text{limit on the transition mag. mom}$$

Miranda, Papoulias, Sanders, Tortola, Valle JHEP '21

cannot give information about the nature of N

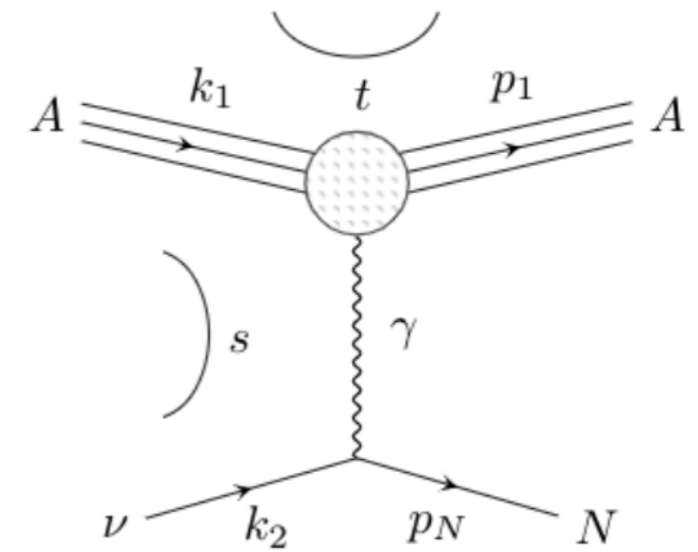
But what if the N decays to $N \rightarrow \nu\gamma$?



The kinematics and differential distribution is more involved than simple upscattering

phase space

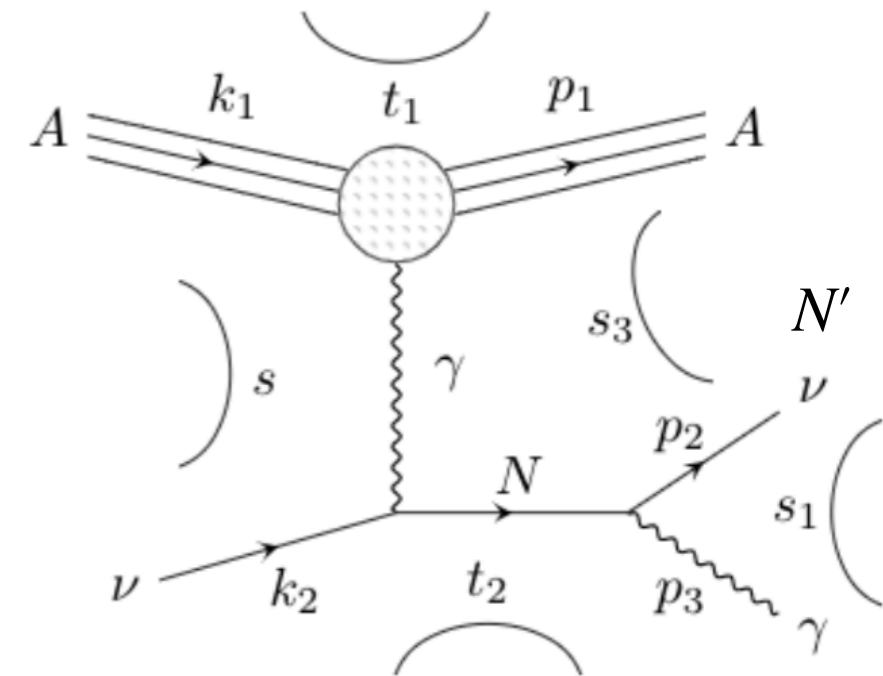
$$d^5\sigma_{\nu_\alpha A \rightarrow \nu_\beta A \gamma} = \frac{1}{2(s - m_A^2)} \frac{1}{2} \sum_{\text{spins}} |\mathcal{M}_{\nu_\alpha A \rightarrow \nu_\beta A \gamma}^{\text{D(M)}}|^2 d\Phi_3$$



Amplitude depends on: Dirac vs. Majorana mediator

$$i\mathcal{M}_{\text{Dirac}} \propto [\bar{u}_{\nu_\beta} \sigma_{\lambda\xi} P_R \epsilon^{\lambda*} p_3^\xi i(\not{p}_N + m_N) \sigma_{\mu\rho} P_L q^\rho u_{\nu_\alpha}]$$

$$i\mathcal{M}_{\text{Maj}} \propto [\bar{u}_{\nu_\beta} \sigma_{\lambda\xi} (P_R - P_L) \epsilon^{\lambda*} p_3^\xi i(\not{p}_N + m_N) \sigma_{\mu\rho} P_L q^\rho u_{\nu_\alpha}]$$



Integrate over the free phase space parameters

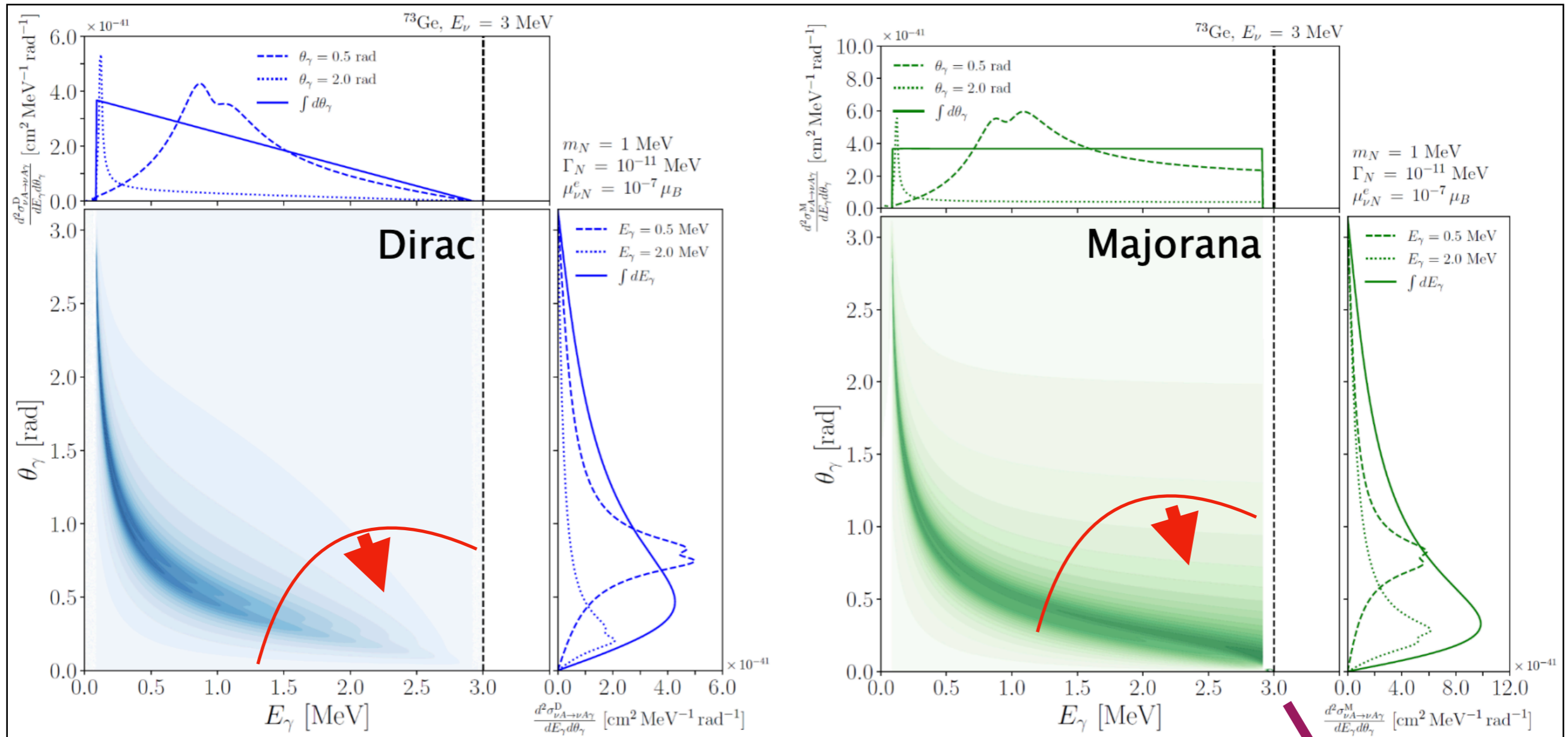
=> differential distributions in desired lab frame variables

$$\frac{d^2\sigma_{\nu_\alpha A \rightarrow \nu_\beta A \gamma}^{\text{D(M)}}}{dE_\gamma d\theta_\gamma} \quad \frac{d\sigma_{\nu_\alpha A \rightarrow \nu_\beta A \gamma}^{\text{D(M)}}}{dE_R}$$

$$\mathcal{B}_{N \rightarrow X \gamma} P_{\text{decay}} = \frac{\Gamma_{N \rightarrow X \gamma}}{\Gamma_N^{\text{inv}} + \Gamma_{N \rightarrow X \gamma}} \left[1 - \exp\left(-\frac{L_{\text{det}}(\Gamma_N^{\text{inv}} + \Gamma_{N \rightarrow X \gamma})}{\beta\gamma}\right) \right]$$

Details in the “shine”: Dirac vs Majorana sterile

Benchmark choices: $E_\nu = 3 \text{ MeV}$ $m_N = 1 \text{ MeV}$ $\mu_{\nu N}^e = 3 \times 10^{-7} \mu_B$



Clear distinction for $E_\gamma > E_\nu/2$

Coincidence for BG rejection: realistic @ experiments like NUCLEUS

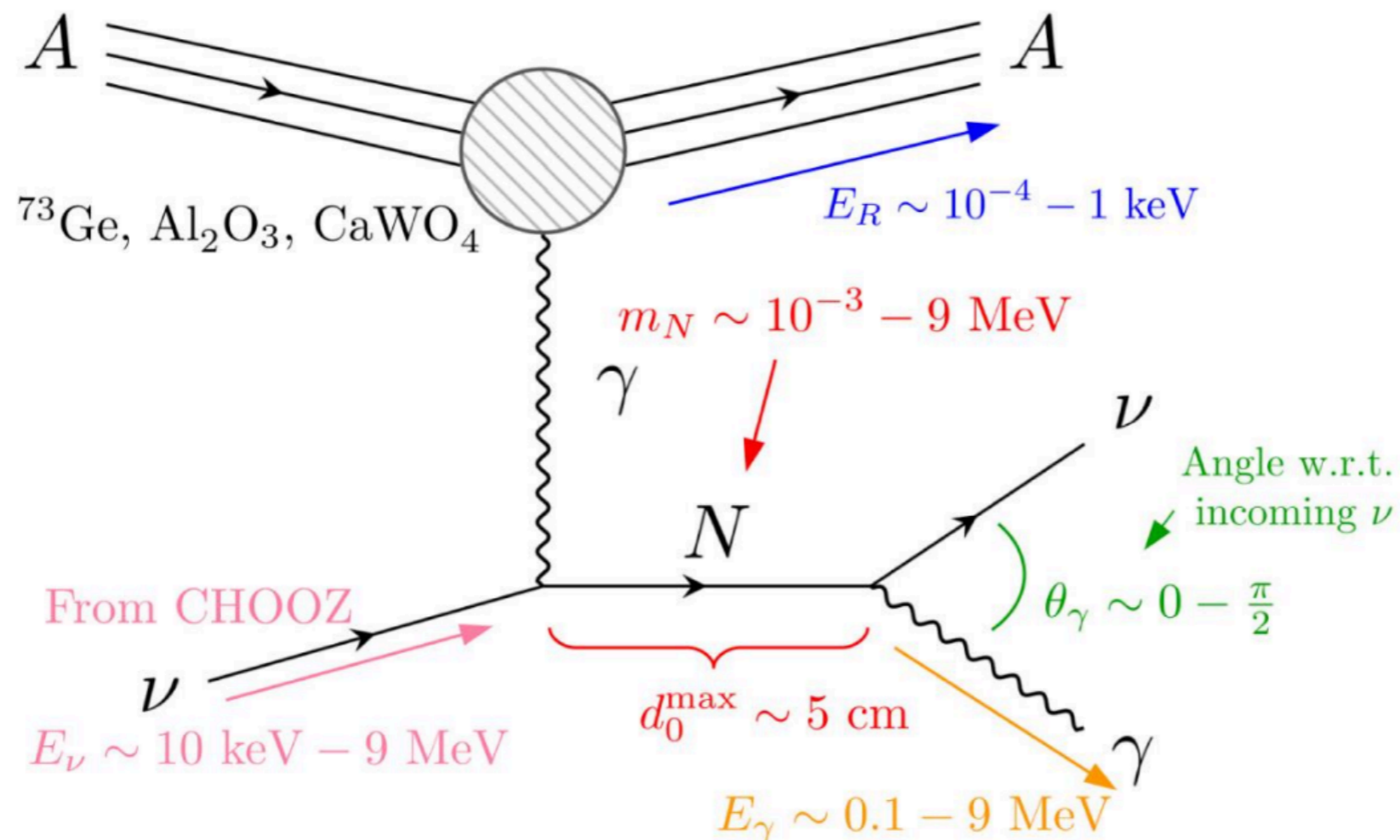
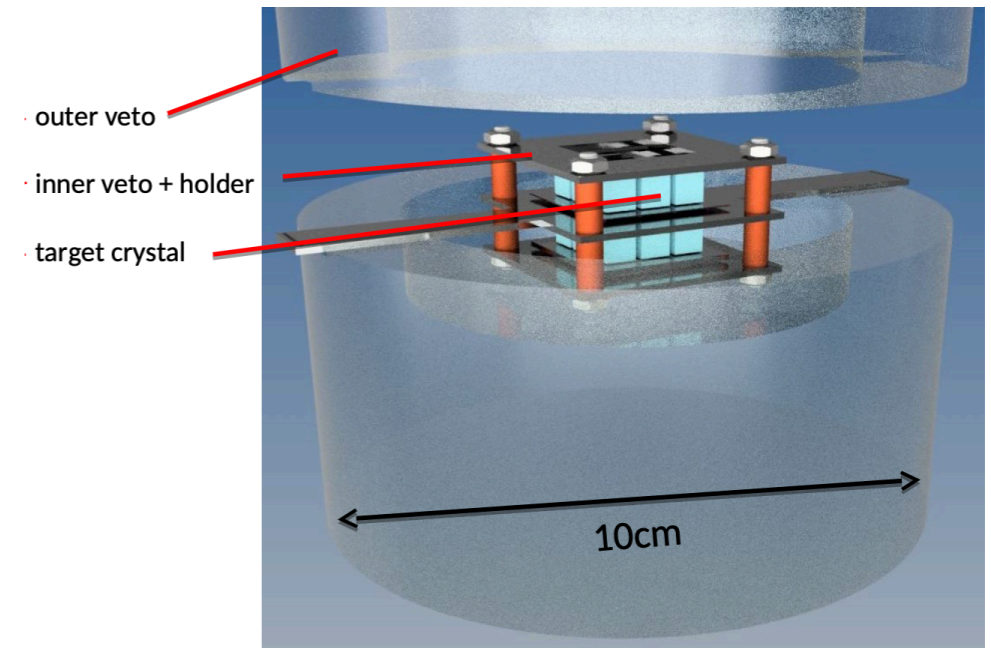
E_ν

For a realistic experiment: integrated over the flux + sterile decay width from detector dimensions

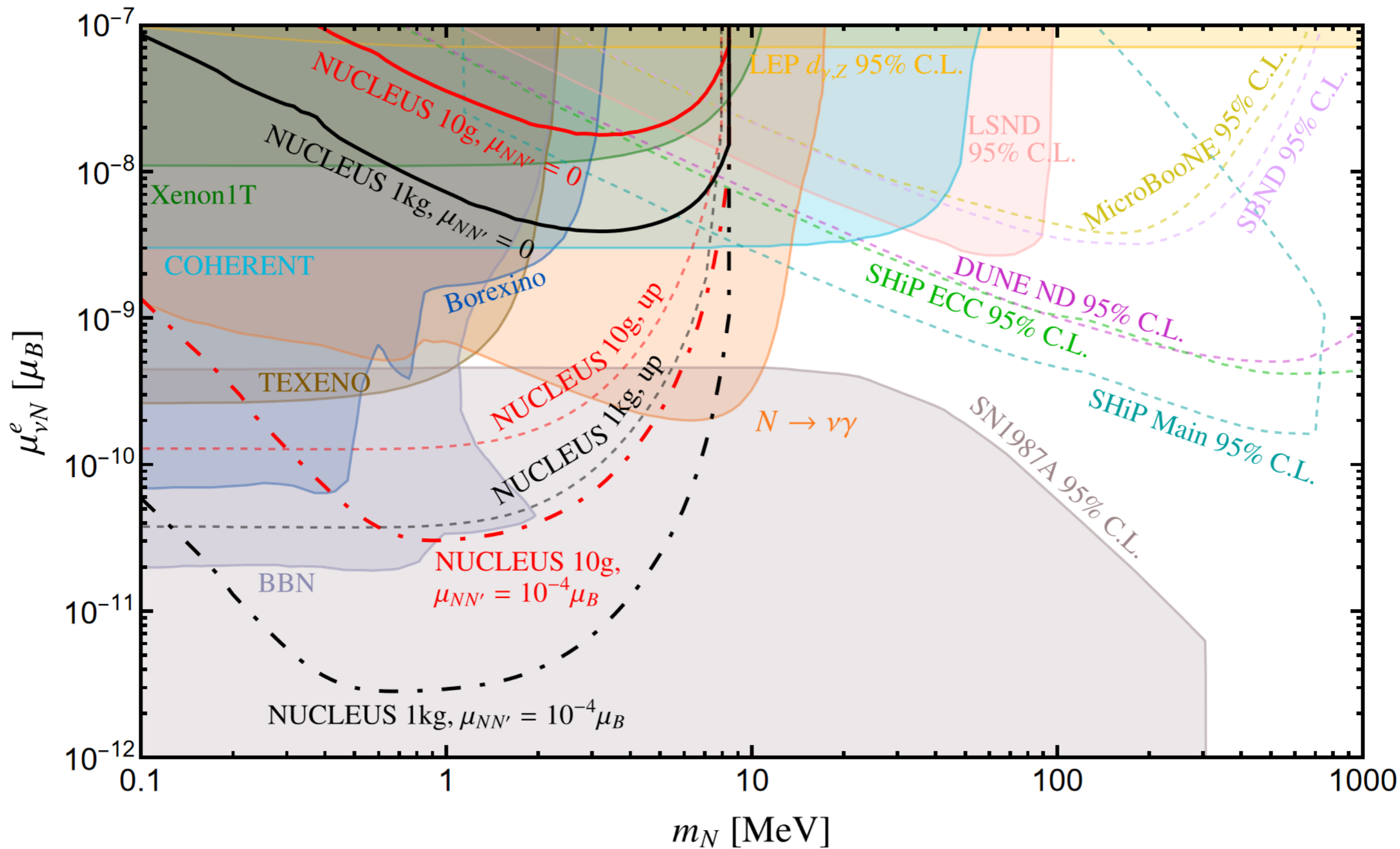
Taking NUCLEUS @ CHOOZ as a case study

- 10g $Al_2O_3/CaWO_4$ for NUCLEUS-Phase 1: $L \sim 5$ cm
- Future 1kg Ge upgrade possibility: $L \sim 25$ cm
- Sensitivity to photon energy: 1 keV to 10 MeV
@ Cryogenic outer veto

energy resolution: 50-100 keV @MeV energies



Sensitivity @ NUCLEUS

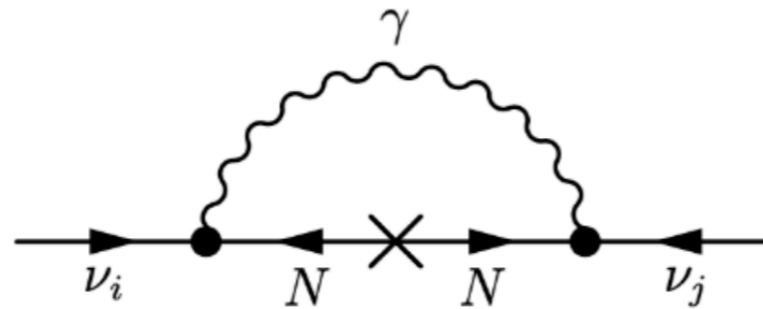


$\nu A \rightarrow \nu A \gamma$ distribution consistent with
Dirac/Majorana mediator => ?

Insights into the

- (i) Dirac vs. Majorana nature**
- (ii) mass mechanism
of active neutrinos**

For a distribution consistent with Majorana N



Majorana N + a transition dipole moment \Rightarrow
majorana nature of active neutrino

$$m_\nu^{\text{Maj}} \sim \mu_{\nu N}^2 \frac{\Lambda^2 m_N^{\text{Maj}}}{16\pi^2}$$

Complimentary to $0\nu\beta\beta$ decay and LNV rare meson decays

For a distribution consistent with Dirac N

no conclusive statement can be made about the nature of ν !

$\nu A \rightarrow \nu A \gamma$ distribution consistent with Majorana sterile

Transition magnetic moment via loop diagram with heavy NP (Λ)=>

Dirac mass term $m_{\nu N} \bar{\nu}_L N_R$

$$\frac{\mu_{\nu N}}{\mu_B} \approx \frac{m_e \delta m_{\nu N}}{\Lambda^2}$$

Magill, Plestid, Pospelov, Tsai '18



SM+sterile states => transition magnetic moment via loop diagrams with charged leptons

$$\frac{|\mu_{\nu N}|}{\mu_B} = \frac{3m_{\nu N}m_e}{16\pi^2} \frac{G_F}{\sqrt{2}} \sim 10^{-13} \left(\frac{m_{\nu N}}{1 \text{ MeV}} \right)$$

Pal '81, Shrock '82

too tight to explain a signal for radiative $CE\nu NS$ for canonical type I seesaw !

preferred scenarios: **Symmetry driven models** (“unnatural” cancellation with tree level mass term)

Basis of independent operators at d=6

Bell et al. PRL 2005

$$\mathcal{O}_1^{(6)} = g_1 \bar{L} \tilde{H} \sigma^{\mu\nu} N_R B_{\mu\nu}$$

$$\mathcal{O}_2^{(6)} = g_2 \bar{L} \tau^a \tilde{H} \sigma^{\mu\nu} N_R W_{\mu\nu}^a$$

$$\mathcal{O}_3^{(6)} = \bar{L} \tilde{H} N_R (H^\dagger H)$$

$$\frac{\mu_{\nu N}}{\mu_B} = -16\sqrt{2} \left(\frac{m_e v}{\Lambda^2} \right) \left[C_1^{(6)}(v) + C_2^{(6)}(v) \right]$$

$$\delta m_{\nu N} = -C_3^{(6)}(v) \frac{v^3}{2\sqrt{2}\Lambda^2}$$

operator mixing=>

$$\frac{|\mu_{\nu N}|}{\mu_B} \sim 10^{-15} \left(\frac{\delta m_{\nu N}}{1 \text{ eV}} \right) \text{ for } \Lambda = 1 \text{ TeV}$$

Again: non-trivial mechanism/symmetry needed to get large mag. mom. without blowing up ν mass

Concluding remarks



- Experiments e.g. NUCLEUS (and COHERENT) will be able to detect high energy γ
 - ⇒ Search for $\nu A \rightarrow \nu A \gamma$
 - ⇒ Coincidence signal with low background
 - ⇒ Constrains on active-sterile/sterile-sterile transition magnetic moment

- If there exists N with large active-sterile magnetic moment : ⇒

Try to measure outgoing E_γ and θ_γ
distributions can distinguish Dirac vs. Majorana nature of N
Insights into active neutrino mass mechanisms

Thank you for your attention!

LNV: Sterile neutrinos

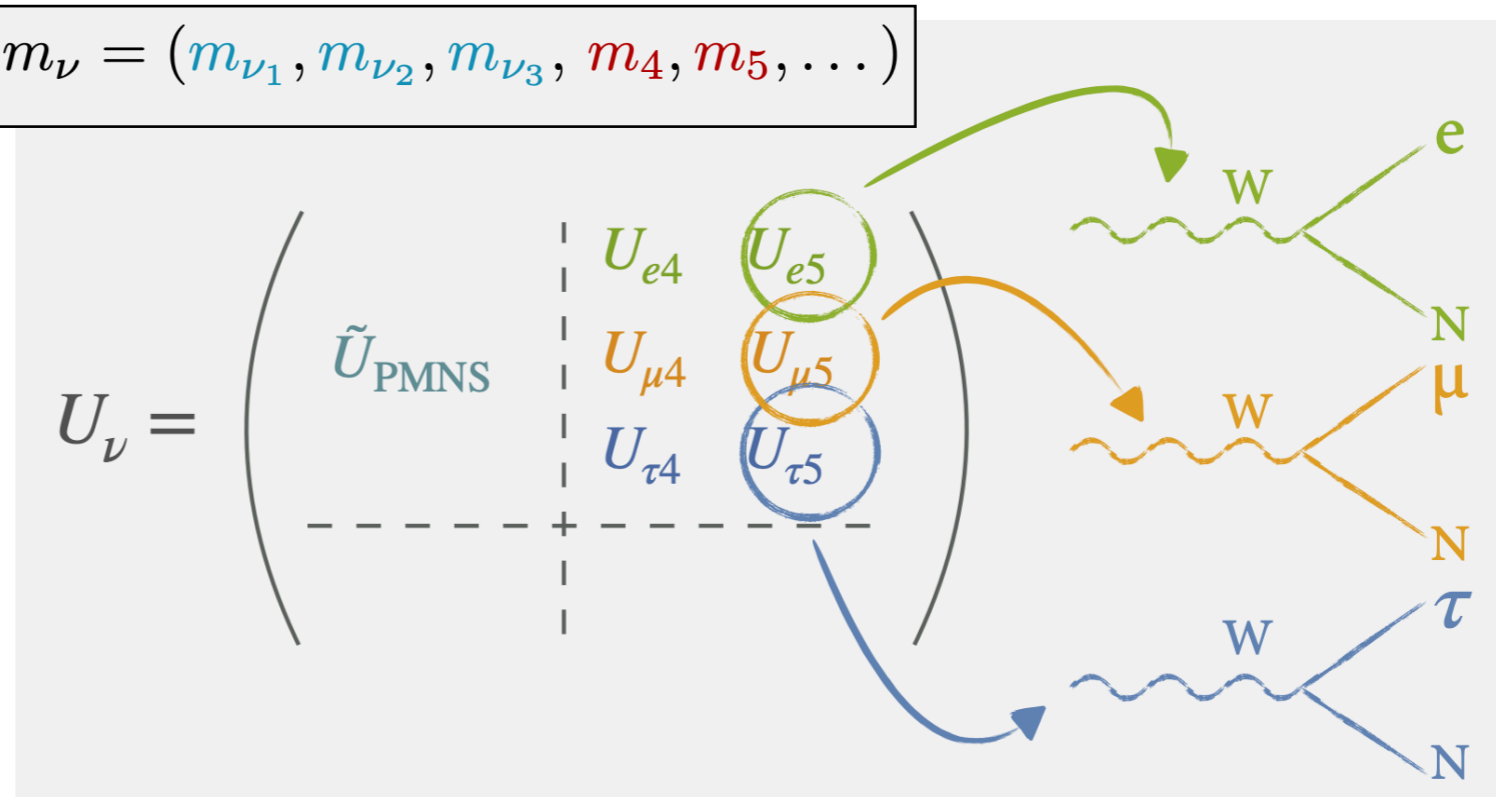
“Sterile”: uncharged under SM gauge group $SU(3)_c \times SU(2)_L \times U(1)_Y$

Mostly interacts through active-sterile mixing (we will discuss an exception in detail)

3+n steriles masses:

$$m_\nu = (m_{\nu_1}, m_{\nu_2}, m_{\nu_3}, m_4, m_5, \dots)$$

3+n unitary mixing matrix:



Sterile masses and mixings as free parameters: give different seesaw limits

$$\mathcal{L}_\nu = \mathcal{L}_{\text{old}} - \lambda_{\alpha i} L^\alpha H N^i - \sum_{i=1}^3 \frac{M_i}{2} N^i N^i + H.c. \quad \mu = \lambda \langle H^0 \rangle$$

$M = 0$: 3 Dirac pairs

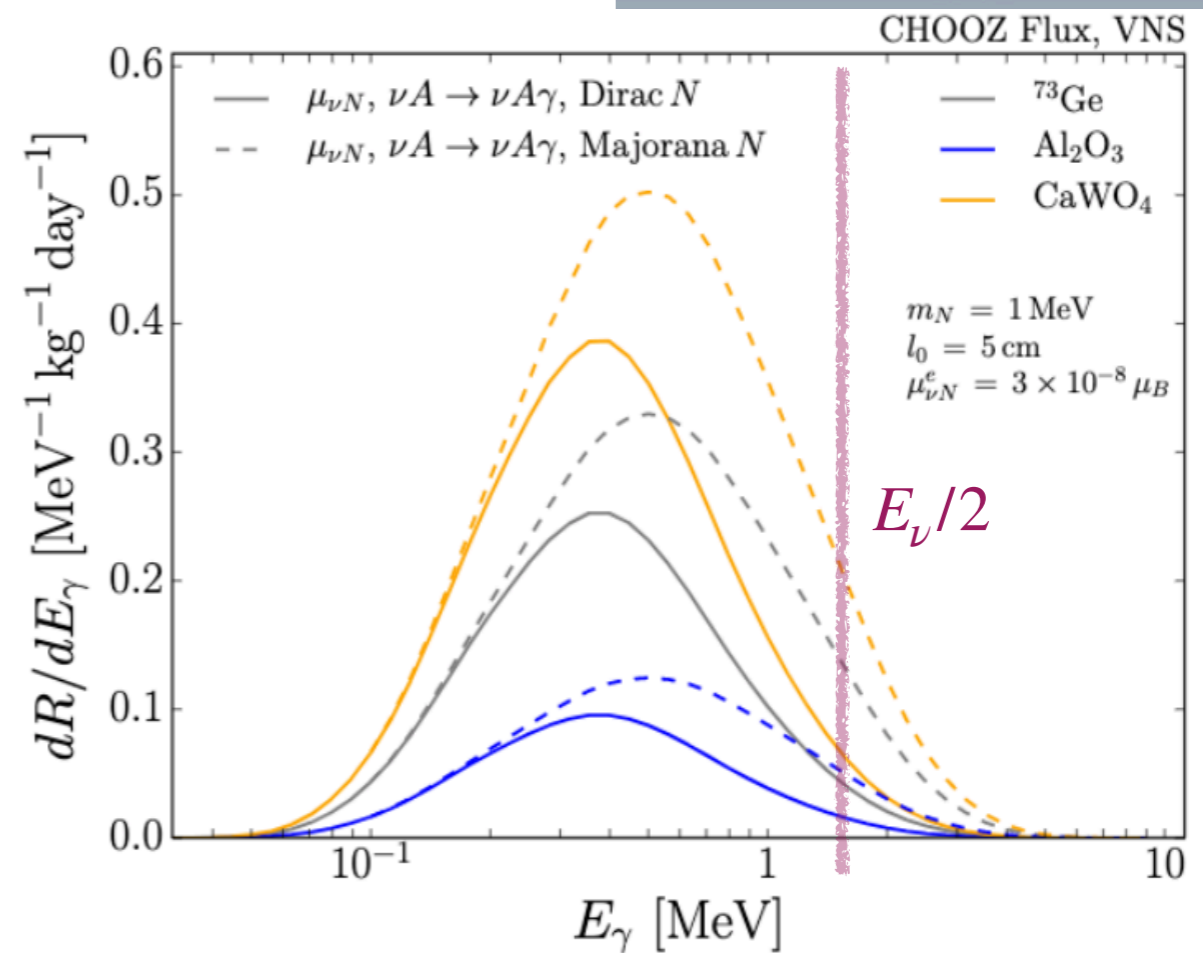
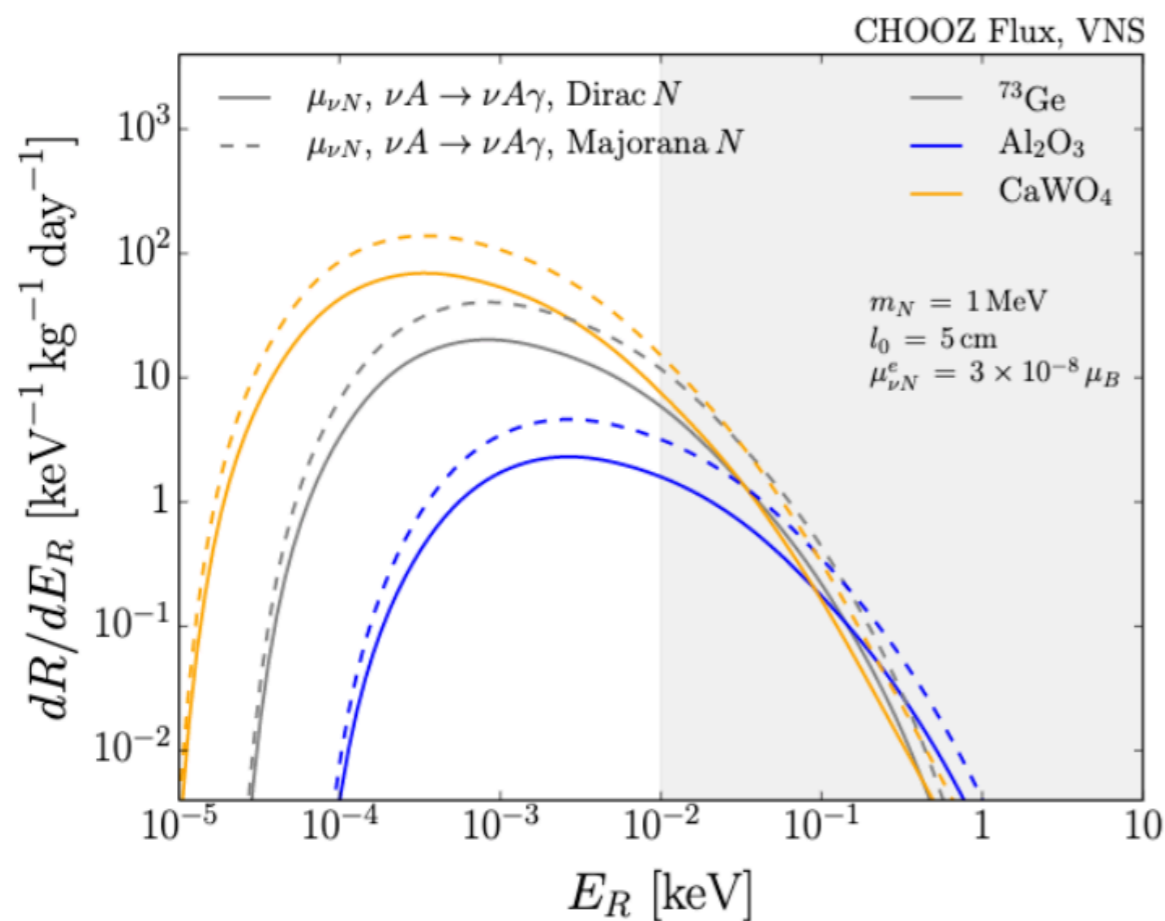
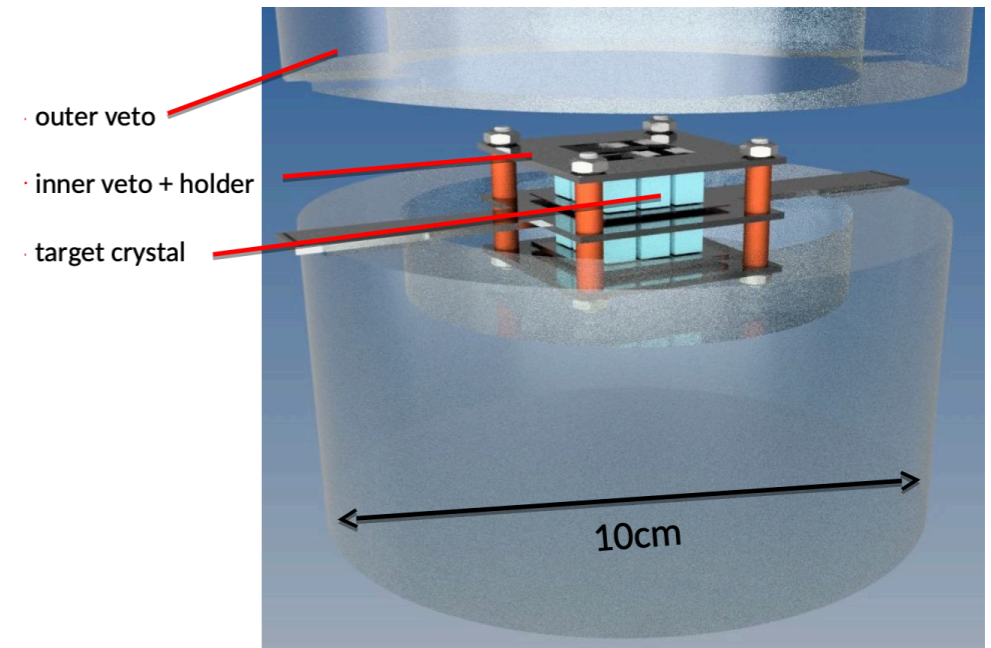
$M \sim \mu$: 6 similar masses + very large mixing

$M \gg \mu$: type I seesaw

$M \ll \mu$: quasi-Dirac + maximal mixing

NUCLEUS @ CHOOZ as a case study

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Dirac case photon energy distribution falls off very quickly after $E_\gamma > E_\nu/2$

NUCLEUS can provide an energy resolution: 50-100 keV @MeV energies